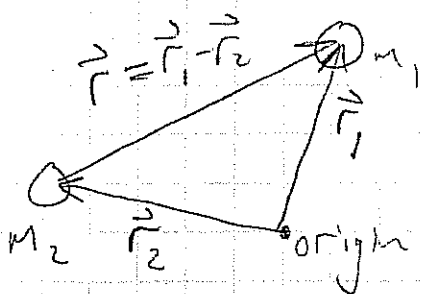


CH08 - Two-body central force problem

$$\vec{F}_{\text{by } 2 \text{ on } 1} = -\frac{GM_1 M_2}{r^2} \hat{r}$$

$\vec{r} = \vec{r}_1 - \vec{r}_2$ is the relative position of m_1

$$r = |\vec{r}| = |\vec{r}_1 - \vec{r}_2| \quad \text{and} \quad \hat{r} = \frac{\vec{r}_1 - \vec{r}_2}{|\vec{r}_1 - \vec{r}_2|}$$

$$U(\vec{r}_1, \vec{r}_2) = - \int \vec{F}_{\text{by } 2 \text{ on } 1} \cdot d\vec{r}$$

$$= - \int \frac{-GM_1 M_2}{r^2} dr$$

$$= GM_1 M_2 \int \frac{1}{r^2} dr = -\frac{GM_1 M_2}{r} + C$$

Choose C so that $U(r=\infty) = 0$, then $C = 0$

$$U = -\frac{GM_1 M_2}{r} \quad \text{where } r = |\vec{r}_1 - \vec{r}_2|$$

So, U is only a function of r . $\xrightarrow{\text{write}} U(r)$

$$K = \frac{1}{2} m_1 v_1^2 + \frac{1}{2} m_2 v_2^2$$

$$K = \frac{1}{2} m_1 \dot{\vec{r}}_1^2 + \frac{1}{2} m_2 \dot{\vec{r}}_2^2$$

The Lagrangian is

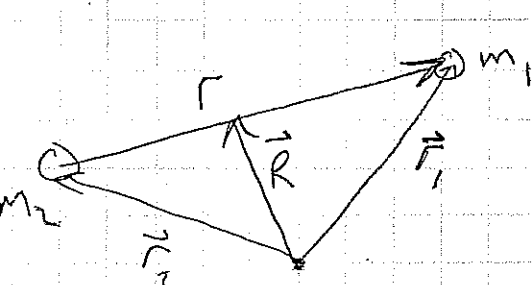
$$L = K - U$$

$$= \frac{1}{2} m_1 \dot{\vec{r}}_1^2 + \frac{1}{2} m_2 \dot{\vec{r}}_2^2 + \frac{GM_1 M_2}{r}$$

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The Lagrangian is a function of six coordinates (3 for \vec{r}_1 and 3 for \vec{r}_2). It is possible to simplify this (especially for N bodies) by using the CM frame and relative frame.

The center of mass is



$$\vec{R} = \frac{m_1 \vec{r}_1 + m_2 \vec{r}_2}{m_1 + m_2} \Rightarrow \text{weighted average}$$

$$M = m_1 + m_2$$

$$\begin{aligned} \vec{P}_{\text{sys}} &= \vec{p}_1 + \vec{p}_2 \\ &= m_1 \vec{v}_1 + m_2 \vec{v}_2 \\ &= (m_1 + m_2) \dot{\vec{R}} \\ &= M \dot{\vec{R}} \end{aligned}$$

1. $\vec{F}_{\text{net}} = 0$, then \vec{P}_{sys} is constant and $\dot{\vec{R}}$ is constant.

Kinetic energy of the system is $K = \frac{1}{2} m_1 v_1^2 + \frac{1}{2} m_2 v_2^2$

$$\vec{r} = \vec{r}_1 - \vec{r}_2$$

$$\dot{\vec{r}} = \dot{\vec{r}}_1 - \dot{\vec{r}}_2$$

$$\vec{R} = \frac{m_1 \vec{r}_1 + m_2 \vec{r}_2}{M}$$

$$\dot{\vec{R}} = \frac{m_1}{M} \dot{\vec{r}}_1 + \frac{m_2}{M} \dot{\vec{r}}_2$$

$$= \frac{1}{2} m_1 \dot{\vec{r}}_1^2 + \frac{1}{2} m_2 \dot{\vec{r}}_2^2$$

Solve for $\dot{\vec{r}}_1$ and $\dot{\vec{r}}_2$ in terms of $\dot{\vec{r}}$ and $\dot{\vec{R}}$

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$$\begin{aligned}\dot{\vec{r}}_1 &= \dot{\vec{r}} - \dot{\vec{r}}_2 \\ &= \dot{\vec{r}} - \left(\frac{M\dot{\vec{r}} - m_1\dot{\vec{r}}_1}{m_2} \right)\end{aligned}$$

$$\dot{\vec{r}}_1 = \dot{\vec{r}} - \frac{M}{m_2}\dot{\vec{r}} + \frac{m_1}{m_2}\dot{\vec{r}}_1$$

$$\dot{\vec{r}}_1 \left(1 - \frac{m_1}{m_2}\right) = \dot{\vec{r}} - \frac{M}{m_2}\dot{\vec{r}}$$

$$\dot{\vec{r}}_1 = \frac{\dot{\vec{r}} - \frac{M}{m_2}\dot{\vec{r}}}{\left(1 - \frac{m_1}{m_2}\right)}$$

messy

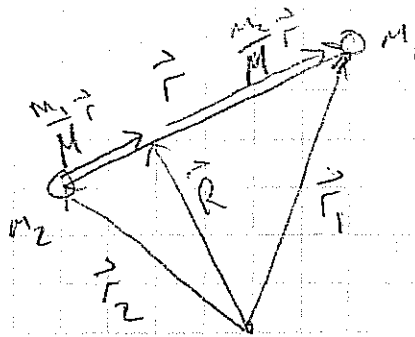
use geometric approach

$$\dot{\vec{r}}_2 + \frac{m_1}{m_2}\dot{\vec{r}} = \dot{\vec{r}}$$

$$\dot{\vec{r}}_2 = \dot{\vec{r}} - \frac{m_1}{M}\dot{\vec{r}}$$

$$\dot{\vec{r}} + \frac{m_2}{m_1}\dot{\vec{r}} = \dot{\vec{r}}_1$$

$$\dot{\vec{r}}_1 = \dot{\vec{r}} + \frac{m_2}{M}\dot{\vec{r}}$$



$$K = \frac{1}{2} m_1 \dot{\vec{r}}_1^2 + \frac{1}{2} m_2 \dot{\vec{r}}_2^2$$

$$= \frac{1}{2} m_1 \left(\dot{\vec{r}} + \frac{m_2}{M} \dot{\vec{r}} \right)^2 + \frac{1}{2} m_2 \left(\dot{\vec{r}} - \frac{m_1}{M} \dot{\vec{r}} \right)^2$$

$$= \frac{1}{2} m_1 \left[\dot{\vec{r}}^2 + \frac{m_2^2}{M^2} \dot{\vec{r}}^2 + 2 \dot{\vec{r}} \dot{\vec{r}} \frac{m_2}{M} \right] + \frac{1}{2} m_2 \left[\dot{\vec{r}}^2 + \frac{m_1^2}{M^2} \dot{\vec{r}}^2 - 2 \dot{\vec{r}} \dot{\vec{r}} \frac{m_1}{M} \right]$$

$$= \frac{1}{2} m_1 \dot{\vec{r}}^2 + \frac{1}{2} m_1 \frac{m_2^2}{M^2} \dot{\vec{r}}^2 + \frac{1}{2} m_2 \dot{\vec{r}}^2 + \frac{1}{2} m_2 \frac{m_1^2}{M^2} \dot{\vec{r}}^2$$

Cancel

$$= \frac{1}{2} (m_1 + m_2) \dot{\vec{r}}^2 + \frac{1}{2} \frac{(m_1 + m_2)(m_1 m_2)}{M} \dot{\vec{r}}^2$$

$$K = \frac{1}{2} M \dot{\vec{r}}^2 + \frac{1}{2} \left(\frac{m_1 m_2}{m_1 + m_2} \right) \dot{\vec{r}}^2$$

has units of mass, called reduced mass

$$\mu = \frac{m_1 m_2}{m_1 + m_2}$$

$$\mu = \frac{m_1 m_2}{m_1 + m_2} = \frac{m_1 m_2}{m_1 (1 + \frac{m_2}{m_1})} = \frac{m_2}{1 + \frac{m_2}{m_1}} = \frac{m_1}{1 + \frac{m_1}{m_2}} \quad \text{or}$$

$$\left. \begin{array}{ll} \text{if } m_2 > m_1, & \mu < m_2 \\ \text{if } m_1 > m_2, & \mu < m_1 \\ \text{if } m_1 = m_2, & \mu = \frac{m_1^2}{2m_1} = \frac{m_1}{2} \end{array} \right\} \mu < m_1 \text{ and } \mu < m_2 \text{ regardless of the masses.}$$

Thus why it is called the reduced mass

$$K_{\text{trans}} = \frac{1}{2} M \dot{R}^2 \quad K_{\text{rel}} = \frac{1}{2} \mu \dot{r}^2$$

If $K_{\text{trans}} = 0$ (center-of-mass position is constant) then

$$K = K_{\text{rel}} = \frac{1}{2} \mu \dot{r}^2 \quad \text{Then } \dots$$

we can treat the system as if it's a single particle of mass μ with position \vec{r} (relative to \vec{r}_c) and velocity $\dot{\vec{r}}$.

The Lagrangian is $L = K - U$

$$L = \underbrace{\frac{1}{2} M \dot{R}^2}_{L_{\text{trans}}} + \underbrace{\left(\frac{1}{2} \mu \dot{r}^2 - U(r) \right)}_{L_{\text{rel}}}$$

The translational term has no associated potential energy because there are no external forces. The relative term is the Lagrangian for a single particle moving in a central force field.

Generalized coordinates are \vec{R} and \vec{r} .

\vec{R} :

$$\frac{\partial L}{\partial R_x} = \frac{d}{dt} \frac{\partial R}{\partial R_x}$$

$$0 = \frac{d}{dt} (M \dot{R}_x) = M \ddot{R}_x \quad \text{also} \quad 0 = M \ddot{R}_y \quad \text{and} \quad 0 = M \ddot{R}_z$$

$$\text{so} \quad 0 = M \ddot{\vec{R}} \quad \text{and} \quad \dot{\vec{R}} = \text{constant}$$

Newt. 2nd law tells us the same thing since the net external force is zero.

In other words, the momentum is conserved.

Use the CM reference frame

In the CM frame, $\dot{\vec{R}} = 0$. Then,

$$L = L_{\text{rel}} = \frac{1}{2} \mu \dot{\vec{r}}^2 - U(r)$$

The system can be thought of as a single particle, mass μ and velocity $\dot{\vec{r}}$ and position \vec{r} .

Angular Momentum

$$\vec{L} = \vec{r} \times \vec{p} \quad \text{so}$$

$$\vec{L}_{\text{tot}} = \vec{L}_1 + \vec{L}_2 = \vec{r}_1 \times \vec{p}_1 + \vec{r}_2 \times \vec{p}_2$$

$$= \vec{r}_1 \times \dot{\vec{r}}_1 m_1 + \vec{r}_2 \times \dot{\vec{r}}_2 m_2$$

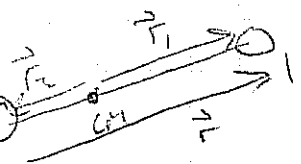
$$= \frac{m_2}{M} \vec{r} \times \frac{m_1 m_2}{M} \dot{\vec{r}} + \left(\frac{m_1}{M} \right) \vec{r} \times \left(\frac{-m_1 m_2}{M} \dot{\vec{r}} \right)$$

$$= \frac{m_1 m_2}{M^2} (m_2 \vec{r} \times \dot{\vec{r}} + m_1 \vec{r} \times \dot{\vec{r}})$$

$$= \frac{m_1 m_2}{M^2} (m_1 + m_2) \vec{r} \times \dot{\vec{r}}$$

$$\vec{L}_{\text{tot}} = \mu \vec{r} \times \dot{\vec{r}} = \boxed{\vec{r} \times \mu \dot{\vec{r}}}$$

Same as particle of mass μ , velocity $\dot{\vec{r}}$

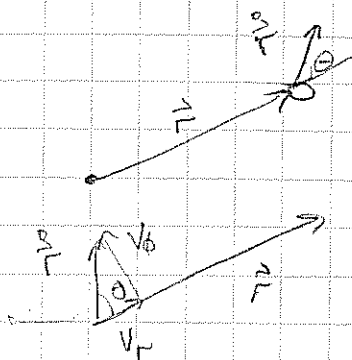


$$\vec{r}_1 = \frac{m_2}{M} \vec{r}$$

$$\vec{r}_2 = -\frac{m_1}{M} \vec{r}$$

$$|\vec{L}| = |\vec{r}| \mu |\dot{\vec{r}}| \sin \theta \quad \dot{\vec{r}} = \langle V_r, V_\phi \rangle \quad (\text{polar coordinates})$$

$$\text{so } |\dot{\vec{r}}|^2 = V_r^2 + V_\phi^2 = \dot{r}^2 + (r\dot{\phi})^2$$



$$\sin \phi = \frac{V_\phi}{|\dot{\vec{r}}|}$$

$$\text{so } |\vec{L}| = r \mu V_\phi = r \mu r \dot{\phi}$$

$$|\vec{L}| = r^2 \mu \dot{\phi} \equiv \ell \quad \text{for reduced mass particle in our model}$$

The Lagrangian is

$$L = L_{\text{rel}} = \frac{1}{2} \mu \dot{\vec{r}}^2 - U$$

$$= \frac{1}{2} \mu (\dot{r}^2 + (r\dot{\phi})^2) - U(r) = \frac{1}{2} \mu \dot{r}^2 + \frac{1}{2} \mu r^2 \dot{\phi}^2 - U(r)$$

$$\frac{\partial L}{\partial \phi} = \frac{d}{dt} \frac{\partial L}{\partial \dot{\phi}}$$

$$0 = \frac{d}{dt} \frac{\partial L}{\partial \dot{\phi}}$$

$$\text{so } \frac{\partial L}{\partial \dot{\phi}} = \text{constant}$$

$$\frac{1}{2} \mu 2 r^2 \dot{\phi} = \text{const}$$

$$\mu r^2 \dot{\phi} = \text{const}$$

This means that $\ell = \mu r^2 \dot{\phi}$ is constant, exactly as expected since $\vec{r}_{\text{net}} = 0$

$$\frac{\partial L}{\partial r} = \frac{d}{dt} \frac{\partial L}{\partial \dot{r}}$$

$$\mu r \dot{\phi}^2 - \frac{\partial U}{\partial r} = \frac{d}{dt} (\mu \dot{r})$$

$$\mu r \dot{\phi}^2 - \frac{\partial U}{\partial r} \equiv \mu \ddot{r}$$

$$\nabla U = -F_r$$

$$\mu \ddot{r} = -\nabla U + \mu r \dot{\phi}^2$$

due to $F_{\text{grav}}(r)$

outward radial force call it a "centrifugal force"

$$F_{\text{cf}} = \mu r \dot{\phi}^2$$

The centrifugal force term is

$$F_{cf} = \mu r \dot{\phi}^2 \left(\frac{r}{r} \right) = \mu r^2 \dot{\phi} \left(\frac{\dot{\phi}}{r} \right) = \frac{l \dot{\phi}}{r} \left(\frac{r^2 \mu}{r^2 \mu} \right) = \frac{l^2}{\mu r^3}$$

Write a potential energy function using

$$U_{cf} = - \int F_{cf} dr = - \int \frac{l^2}{\mu r^3} dr = \frac{l^2}{\mu} \frac{r^{-2}}{2} = \frac{l^2}{2\mu r^2}$$

Thus, $U = U(r) + U_{cf}(r)$

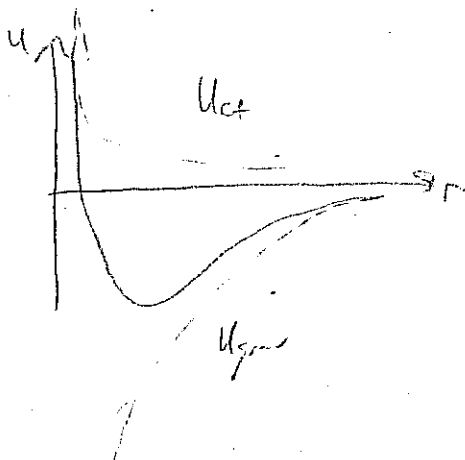
$$\text{Then, } \mu \ddot{r} = - \frac{dU}{dr} + - \frac{dU_{cf}}{dr}$$

$$\mu \ddot{r} = - \frac{d}{dr} (U + U_{cf})$$

$$\mu \ddot{r} = - \frac{dU_{eff}}{dr}$$

It's a 1-D problem with $U_{eff} = U + U_{cf}$

$$U_{eff} = - \frac{GM_1 m_2}{r} + \frac{l^2}{2\mu r^2}$$



CH08 Continued...

Lagrangian: $L = L_{\text{rel}}$ (in CM frame $L_{\text{trans}} = 0$)

$$L_{\text{rel}} = \frac{1}{2} \mu \dot{r}^2 + \frac{1}{2} \mu r^2 \dot{\phi}^2 - U(r)$$

$$\therefore \boxed{\mu r^2 \dot{\phi} = \text{const}}$$

$$\therefore \mu \ddot{r} = -\nabla U + \mu r \dot{\phi}^2$$

$$\boxed{\mu \ddot{r} = -\nabla U_{\text{eff}}(r)} \quad \text{where } U_{\text{eff}} = U + U_{\text{ct}}$$

$$= -\frac{Gm_1 m_2}{r} + \frac{l^2}{2\mu r^2}$$

$$\mu \dot{r} \ddot{r} = \frac{d}{dt} \left(\frac{1}{2} \mu \dot{r}^2 \right) \quad (\text{Verify this!})$$

$$\rightarrow = 2 \left(\frac{1}{2} \right) \mu \dot{r} = \mu \ddot{r} \checkmark$$

$$\text{so } \mu \dot{r} \ddot{r} = -\frac{d}{dr} U_{\text{eff}}(r) \dot{r}$$

$$\frac{d}{dt} \left(\frac{1}{2} \mu \dot{r}^2 \right) = -\frac{dU_{\text{eff}}}{dt}$$

$$\frac{d}{dt} \left(\frac{1}{2} \mu \dot{r}^2 + U_{\text{eff}} \right) = 0$$

$$\frac{1}{2} \mu \dot{r}^2 + U_{\text{eff}} \text{ IS constant!}$$

all this the total energy $\boxed{E = \frac{1}{2} \mu \dot{r}^2 + U_{\text{eff}}}$ of our single particle model.

summary

Treat system as a single particle with one generalized coord.

$$L = \frac{1}{2} \mu \dot{r}^2 - U_{\text{eff}}, \quad U_{\text{eff}} = -\frac{Gm_1 m_2}{r} + \frac{l^2}{2\mu r^2}$$

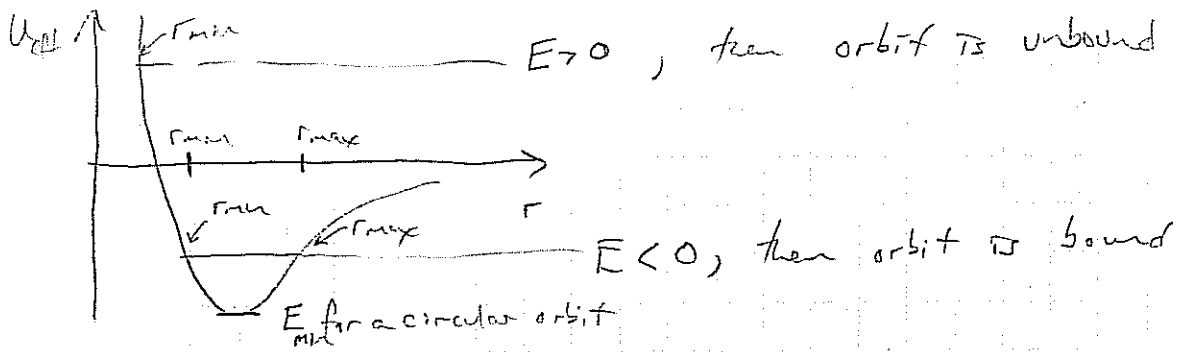
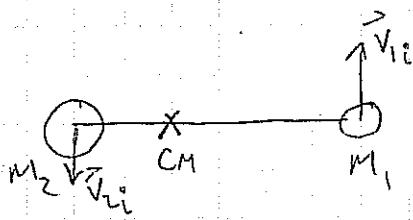
const

changes accordingly to $\mu \ddot{r} = -\nabla U_{\text{eff}}$, $l = r^2 \mu \dot{\phi} = \text{constant}$, $E = \frac{1}{2} \mu \dot{r}^2 + U_{\text{eff}}$

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$$E = \frac{1}{2} \mu \dot{r}^2 - U_{\text{eff}}$$

Example

$$L = r_{1i} m_1 v_{1i} + r_{2i} m_2 v_{2i}$$

$$\dot{\phi} = \frac{L}{r^2 \mu}$$

$$\phi = \frac{d\theta}{dt}$$

$$\mu \ddot{r} = -\frac{dU_{\text{eff}}}{dr} \quad U_{\text{eff}} = -\frac{G m_1 m_2}{r^2} + \frac{L^2}{2 \mu r^2} \quad K = \frac{1}{2} \mu \dot{r}^2$$

$$\mu \ddot{r} = -\frac{G m_1 m_2}{r^2} + \frac{L^2}{\mu r^3}$$

$$x = r \cos \phi$$

$$y = r \sin \phi$$

$$M_1 = 2 \times 10^3 \text{ kg}$$

$$m_2 = 2 m_1$$

$$\vec{v}_{1i} = \langle 0, 2 \times 10^4, 0 \rangle \frac{\text{m}}{\text{s}}$$

$$\vec{v}_{2i} = -\frac{m_1}{m_2} \vec{v}_{1i} \text{ so that } \vec{v}_{\text{cm}} = 0$$

$$\vec{L}_1 = \vec{r}_{1i} \times \vec{p}_{1i}$$

$$L_1 = r_{1i} m_1 v_{1i}$$

$$\vec{L}_2 = \vec{r}_{2i} \times \vec{p}_{2i}$$

$$\vec{L}_2 = r_{2i} m_2 v_{2i}$$

$$\vec{r}_1 = \langle +1.5 \text{ell}, 0, 0 \rangle$$

$$\vec{r}_2 = -\frac{m_1}{m_2} \vec{r}_1 \text{ so } \vec{r}_c$$

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Witnessed & Understood by me, _____

Date _____

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