

# Solutions to Pathria's Statistical Mechanics

## Chapter 3

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March 14, 2016

**Problem 3.1**

**Problem 3.2**

**Problem 3.3**

**Problem 3.4**

**Problem 3.5**

**Problem 3.6**

**Problem 3.7**

**Problem 3.11**

Suppose  $pV^n = C$ , so the work done is

$$\Delta W = \int_{V_1}^{V_2} \frac{C}{V^n} dV = \frac{C}{n-1} (V_2^{1-n} - V_1^{1-n}) \quad (1)$$

The energy difference is given by

$$\Delta U = p_2 V_2 - p_1 V_1 = C(V_2^{1-n} - V_1^{1-n}) \quad (2)$$

Therefore, the heat absorbed is

$$\Delta Q = C \frac{n-2}{n-1} (V_2^{1-n} - V_1^{1-n}) \quad (3)$$

**Problem 3.21**

(a) Classically, the harmonic equation of motion leads to  $x = A \sin \omega t$ . As a result, the kinetic energy and potential energy will be  $m\omega^2 A^2 \cos^2 \omega t / 2$  and  $m\omega^2 A^2 \sin^2 \omega t / 2$  respectively. Average them it's easy to see that  $\bar{K} = \bar{U} = m\omega^2 A^2 / 4$ .

Quantum-mechanically,  $\psi = \sum_n c_n \psi_n$  where  $\psi_n$  is the  $n$ -th Hermitian polynomial. Using the recursive relations, we have

$$\bar{K} = -\frac{\hbar^2}{2m} \sum_n |c_n|^2 \int \psi^* \frac{d^2}{dx^2} \psi dx = \sum_n |c_n|^2 \frac{\hbar\omega(2n+1)}{4} = \frac{1}{2} \sum_n |c_n|^2 E_n \quad (4)$$

$$\bar{U} = \frac{m\omega^2}{2} \sum_n |c_n|^2 \int \psi^* x^2 \psi dx = \sum_n |c_n|^2 \frac{\hbar\omega(2n+1)}{4} = \frac{1}{2} \sum_n |c_n|^2 E_n \quad (5)$$

(b) In Bohr-sommerfeld model, a quantized orbits are hypothesized, namely  $m_e v r = n\hbar$ . In the  $n$ -th orbit, the total energy is  $E_n = -Z^2 k^2 e^4 m_e / 2\hbar^2 n^2$ . The radius of which is  $r_n = n^2 \hbar^2 / Z k e^2 m_e$ . By a naive calculation  $\bar{U} = -Z^2 k^2 e^4 m_e / \hbar^2 n^2$  and  $\bar{T} = Z^2 k^2 e^4 m_e / 2\hbar^2 n^2$ .

In the Schroedinger hydrogen atom,  $\psi_{nlm} = R_{nl}(r)Y_{lm}(\theta, \phi)$ . The kinetic energy is given by

$$\begin{aligned} \bar{T} &= \frac{\hbar^2}{2m} \int \psi_{nlm}^* \left( \frac{d^2}{dr^2} + \frac{2}{r} \frac{d}{dr} - \frac{l(l+1)}{r^2} \right) \psi_{nlm} r^2 \sin \theta dr d\theta d\phi \\ &= \frac{\hbar^2}{2m} \int R_{nl}(r) \left( \frac{1}{n^2 a^2} \right) R_{nl}(r) r^2 dr \\ &= \frac{e^2}{2an^2} \end{aligned} \quad (6)$$

so  $\bar{U} = -e^2/an^2$ .  $a$  is the Bohr radius.

(c) This is also a central force case. The results are quite identical to (b).

### Problem 3.31

“Partition function” for single particle is

$$Q_1 = 1 + e^{-\varepsilon/kT}. \quad (7)$$

So a list of quantities can be obtained:

$$Q_N = (1 + e^{-\varepsilon/kT})^N \quad (8)$$

$$A = -NkT \ln(1 + e^{-\varepsilon/kT}) \quad (9)$$

$$\mu = -kT \ln(1 + e^{-\varepsilon/kT}) \quad (10)$$

$$p = 0 \quad (11)$$

$$S = Nk \ln(1 + e^{-\varepsilon/kT}) + \frac{N\varepsilon}{T} \frac{e^{-\varepsilon/kT}}{1 + e^{-\varepsilon/kT}} \quad (12)$$

$$U = N\varepsilon \frac{e^{-\varepsilon/kT}}{1 + e^{-\varepsilon/kT}} \quad (13)$$

$$C_p = C_V = \frac{N\varepsilon^2 e^{-\varepsilon/kT}}{kT^2 (1 + e^{-\varepsilon/kT})^2} \quad (14)$$

This specific heat is sometimes referred to *Schottky anomaly*.

### Problem 3.41

The equilibrium temperature will be positive, since the energy of the whole system is not bounded from above. This case is a bit like the spin and lattice case. For the subsystem of spins, its energy is bounded from above, so it is possible to attain a negative temperature. While the subsystem of lattice, i.e. ideal gas in this problem, only has positive temperature. The whole system doesn't have a energy limit, so the temperature will only be positive. And energy may flow from the spin subsystem to the ideal gas.