

$$V(k_0) = \sum_{t=0}^{\infty} [\beta^t \ln(1 - \alpha\beta) + \beta^t \alpha \ln k_t]$$

$$= \ln(1 - \alpha\beta) \sum_{t=0}^{\infty} \beta^t + \alpha \sum_{t=0}^{\infty} \beta^t \ln k_0$$

$$= \frac{\alpha}{1 - \alpha\beta} \ln k_0 + \frac{\ln(1 - \alpha\beta)}{1 - \beta} + \alpha \ln(\alpha\beta) \sum_{t=0}^{\infty} \left[\frac{\beta^t}{1 - \alpha} - \frac{(\alpha\beta)^t}{1 - \alpha} \right]$$

$$= \frac{\alpha}{1 - \alpha\beta} \ln k_0 + \frac{\ln(1 - \alpha\beta)}{1 - \beta} + \frac{\alpha\beta}{(1 - \beta)(1 - \alpha\beta)} \ln(\alpha\beta)$$



$$\text{右边} = \max \{u(f(k) - y) + \beta V(y)\}$$

Do not ask what it is. Ask what you can say about it.

$$= \ln(k^\alpha - \alpha\beta k^\alpha) + \beta \left[\frac{\alpha}{1 - \alpha\beta} \ln \alpha\beta k^\alpha + A \right]$$

$$= \ln(k^\alpha - \alpha\beta k^\alpha) + \beta \left[\frac{\alpha}{1 - \alpha\beta} \ln \alpha\beta k^\alpha + A \right]$$

$$= \ln(1 - \alpha\beta) + \alpha \ln k + \beta \left[\frac{\alpha}{1 - \alpha\beta} [\ln \alpha\beta + \alpha \ln k] + k \right]$$

$$= \alpha \ln k + \frac{\alpha\beta}{1 - \alpha\beta} \alpha \ln k + \ln(1 - \alpha\beta) + \frac{\alpha\beta}{1 - \alpha\beta} \ln \alpha\beta + \beta A$$

$$= \frac{\alpha}{1 - \alpha\beta} \ln k + \ln(1 - \alpha\beta) + \frac{\alpha\beta}{1 - \alpha\beta} \ln \alpha\beta + \beta A$$

$$= \frac{\alpha}{1 - \alpha\beta} \ln k + (1 - \beta)A + \beta A$$

$$= \frac{\alpha}{1 - \alpha\beta} \ln k + A$$

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所以，左边 = 右边，证毕。

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Chapter 1

Special relativity



1.1 The principle of relativity

First, we assume there is an upper limit of velocity of propagation of interaction c . Second, we assume that inertial reference frame are all the same in describing the law of physics. Then, we can find the invariant intervals when transforming from one inertial reference frame to another, $ds^2 = -c^2 dt^2 + dx^2 + dy^2 + dz^2$. (In the following, we assume that $c = 1$.) This transformation is called Lorentz transformation, which can be written as

$$\bar{x}^\mu = \Lambda^\mu{}_\nu x^\nu$$

The invariant symbol of the vector representation of Lorentz transformation is $\eta^{\mu\nu}$

$$\Lambda^\mu{}_\rho \Lambda^\nu{}_\sigma \eta^{\rho\sigma} = \eta^{\mu\nu},$$

where,

$$\eta^{\mu\nu} \equiv \begin{bmatrix} -1 & & & \\ & +1 & & \\ & & +1 & \\ & & & +1 \end{bmatrix}$$

The inverse matrix of $\eta^{\mu\nu}$ is

$$\eta_{\mu\nu} = \begin{bmatrix} -1 & & & \\ & +1 & & \\ & & +1 & \\ & & & +1 \end{bmatrix}$$

We can use $\eta^{\mu\nu}$ and its inverse $\eta_{\mu\nu}$ to raise and lower vector indices,

$$x_\mu \equiv \eta_{\mu\nu} x^\nu$$

And we can verify the following equations,

$$\Lambda^\mu{}_\rho \Lambda_\nu{}^\rho = \delta_\mu^\nu$$

$$x^\mu = \eta^{\mu\nu} x_\nu$$

$$\bar{x}_\mu = \Lambda_\mu{}^\nu x_\nu$$

$$\Lambda_\mu{}^\rho \Lambda_\nu{}^\sigma \eta_{\rho\sigma} = \eta_{\mu\nu},$$

In a special case when the new reference frame move along $\hat{1}$ direction with velocity β , we have

$$\bar{x}^0 = \gamma x^0 - \gamma \beta x^1$$

$$\bar{x}^1 = -\gamma \beta x^0 + \gamma x^1$$

Some physical quantity will behave like a tensor (vector, scalar) when transforming form one inertial frame to another. For example,

scalar proper time: $d\tau$, mass: m , electrical charge e

vector four velocity: $u^\mu = \frac{dx^\mu}{d\tau}$, four momentum: $p^\mu = m u^\mu$, four acceleration: $a^\mu = \frac{du^\mu}{d\tau}$, four force: $f^\mu = m a^\mu$.

We can also define the corresponding three vector.

three velocity : $\hat{v}^i = \frac{dx^i}{dt}$

$$u^0 = \gamma_v, u^i = \gamma \hat{v}^i$$

When the new reference frame move along $\hat{1}$ direction with velocity β , we have

$$\bar{\hat{v}}^1 = \frac{\hat{v}^1 - \beta}{1 - \hat{v}^1 \beta}$$

$$\bar{\hat{v}}^2 = \frac{\hat{v}^2}{\gamma(1 - \hat{v}^2 \beta)}$$

$$\bar{\hat{v}}^3 = \frac{\hat{v}^3}{\gamma(1 - \hat{v}^3 \beta)}$$

three momentum : $\hat{p}^i = p^i$

three acceleration : $\hat{a}^i = \frac{dv^i}{dt}$

three force : $\hat{f}^i = \frac{d\hat{p}^i}{dt}$

$$f^i = \gamma_v \hat{f}^i$$

1.2 Relativistic Mechanics

For a free particle, we have

$$\frac{dp^\mu}{d\tau} = 0$$

It can be formulated in several ways



Lagrangian formulation

$$S = -m \int_a^b d\tau, \quad \delta x^\mu(a) = \delta x^\mu(b) = 0$$

$$\delta S = 0 \Rightarrow m \frac{du^\mu}{d\tau} = 0$$

Hamiltonian formulation

$$S = -m \int_{t_1}^{t_2} \sqrt{1 - \dot{x}_i \dot{x}^i} dt$$

$$L = -m \sqrt{1 - \dot{x}_i \dot{x}^i}$$

$$\pi^i = \frac{\partial L}{\partial \dot{x}_i} = \gamma m \eta^{ij} \dot{x}_j$$

$$H = \pi^i \dot{x}_i - L = \gamma m = \sqrt{m^2 + \pi^i \pi_i}$$

So, the Hamiltonian equations are

$$\dot{\pi}^i = 0, \quad \dot{x}_i = \eta_{ij} \frac{\pi^j}{\sqrt{m^2 + \pi^k \pi_k}}$$

Hamiltonian-Jacobi equation

$$H = -\frac{\partial S}{\partial t}, \quad \pi^i = \frac{\partial S}{\partial x_i}$$

If we define $p^0 = H$, $p^i = \pi^i$, then we can verify that $p^\mu = \frac{\partial S}{\partial x_\mu}$. So, p^μ is a vector under Lorentz transformation. The Hamiltonian-Jacobi equation can be written as

$$\left(\frac{\partial S}{\partial t}\right)^2 = m^2 + \left(\frac{\partial S}{\partial x}\right)^2 + \left(\frac{\partial S}{\partial y}\right)^2 + \left(\frac{\partial S}{\partial z}\right)^2$$

For a non-free particle, we have the revised newton's second law:

$$f^\mu = \frac{dp^\mu}{d\tau}$$

The formula is the definition of the four force. It can also be written in three vector as

$$\hat{f}^i = \gamma_v m \hat{a}^i + \gamma_v^3 (\hat{a}^j \hat{v}_j) m \hat{v}^i$$

If the system consists of more than one particles interacting with each other. We can derive the conservation laws from the symmetry.



Translational symmetry and conservation of momentum

$$\bar{x}^\mu = x^\mu + \delta x^\mu$$

$$\delta S = \sum m u_\mu \delta x^\mu|_a^b = 0$$

$\sum p^\mu$ is conserved.

Rotational symmetry and conservation of angular momentum

$$\bar{x}^\mu = x^\mu + x_\nu \delta \Omega^{\mu\nu}$$

$$\delta S = \sum m u^\mu x^\nu \delta \Omega_{\mu\nu}|_a^b = 0$$

$\sum M^{\mu\nu}$ is conserved, where $M^{\mu\nu} = x^\mu p^\nu - x^\nu p^\mu$.

1.3 Relativistic Scattering

1.3.1 Distribution function

The number of particles in the region $\mathbf{r} + d\mathbf{r}$ and $\mathbf{p} + d\mathbf{p}$ is $f(\mathbf{p}, \mathbf{r}) dp_x dp_y dp_z dx dy dz$. Then $f(\mathbf{p}, \mathbf{r})$ is called distribution function.

We first determine the properties of the "volume element" $dp_x dp_y dp_z$, with respect to Lorentz transformations. If we introduce a four-dimensional coordinate system, on whose axes are marked the components of the four-momentum of a particle, then $dp_x dp_y dp_z$, can be considered as the zeroth component of an element of the hypersurface defined by the equation $p^\mu p_\mu + m^2 = 0$. The element of hypersurface is a four-vector directed along the normal to the hypersurface; in our case the direction of the normal obviously coincides with the direction of the four-vector p^μ . From this it follows that the ratio

$$\frac{dp_x dp_y dp_z}{E}$$

is an invariant quantity, since it is the ratio of corresponding components of two parallel four-vectors.

Then, we notice that $dV dt$ is invariant under Lorentz transformation and $dt = \frac{E}{m} d\tau$. So we can infer that

$$dx dv dz E$$

is an invariant quantity. Putting all together, we know the phase volume

$$dp_x dp_y dp_z dx dy dz$$

is an invariant volume. So, we have

$$f(\mathbf{r}, \mathbf{p}) = f'(\mathbf{r}', \mathbf{p}')$$

in coordinate transformation.



1.3.2 Invariant cross section

Recall the definition of cross section

$$\sigma \equiv \frac{\text{Number of Events per target}}{\text{Time} \times \text{Incident Flux}}$$

Here, the incident flux and time are measured in the frame of target particle.

Suppose that we have two colliding beams; we denote by n_1 and n_2 the particle densities in them and by \mathbf{v}_1 and \mathbf{v}_2 the velocities of the particles. In the reference system in which particle 2 is at rest, we are dealing with the collision of the beam of particles 1 with a stationary target. Then according to the usual definition of the cross-section σ , the number of collisions occurring in volume dV in time dt is

$$dN = \sigma v_{rel} n_1 n_2 dV dt$$

,where v_{rel} is the velocity of particle 1 in the rest system of particle 2 (which is just the definition of the relative velocity of two particles in relativistic mechanics).

The number dN is by its very nature an invariant quantity. We would like to express it in a form which is applicable in any reference system:

$$dN = A n_1 n_2 dV dt$$

where A is a number to be determined, for which we know that its value in the rest frame of one of the particles is $v_{rel}\sigma$. We shall always mean by σ precisely the cross-section in the rest frame of one of the particles, i.e. by definition, an invariant quantity. From its definition, the relative velocity v_{rel} is also invariant. The product $dVdt$ is an invariant. Therefore the product $A n_1 n_2$ must also be an invariant. The law of transformation of the particle density n is

$$n = \frac{n_0}{\sqrt{1 - v^2}} = n_0 E / m$$

so we can construct A in an arbitrary frame as

$$A = -\sigma v_{rel} \frac{p_1^\mu p_{2\mu}}{E_1 E_2}$$

Note that

$$-p_1^\mu p_{2\mu} = \frac{m_1}{\sqrt{1 - v_{rel}^2}} m_2 = m_1 m_2 \frac{1 - \mathbf{v}_1 \cdot \mathbf{v}_2}{\sqrt{(1 - \mathbf{v}_1^2) \cdot (1 - \mathbf{v}_2^2)}}$$

we can get the following expression for v_{rel} :

$$v_{rel} = \frac{\sqrt{(\mathbf{v}_1 - \mathbf{v}_2)^2 - (\mathbf{v}_1 \times \mathbf{v}_2)^2}}{1 - \mathbf{v}_1 \cdot \mathbf{v}_2}$$

Finally, we have

$$dN = \sigma \sqrt{(\mathbf{v}_1 - \mathbf{v}_2)^2 - (\mathbf{v}_1 \times \mathbf{v}_2)^2} n_1 n_2 dV dt$$

If the velocities \mathbf{v}_1 and \mathbf{v}_2 are collinear, then we have

$$dN = \sigma |\mathbf{v}_1 - \mathbf{v}_2| n_1 n_2 dV dt$$

If we have only one target, then

$$dN = \sigma |\mathbf{v}_1 - \mathbf{v}_2| n_1 dt$$



1.3.3 Elastic scattering between two particles

$$(E_1, \mathbf{p}_1, E_2, \mathbf{p}_2) \rightarrow (E'_1, \mathbf{p}'_1, E'_2, \mathbf{p}'_2)$$

In lab frame (L frame),

$$E_2 = m_2 \quad \mathbf{p}_2 = 0$$

By the conservation of momentum, we can derive that

$$\cos \theta_1 = \frac{E'_1(E_1 + m_2) - E_1 m_2 - m_1^2}{p_1 p'_1}$$

$$\cos \theta_2 = \frac{(E_1 + m_2)(E'_2 - m_2)}{p_1 p'_2}$$

Here $\theta_1(\theta_2)$ is the angle between $\mathbf{p}'_1(\mathbf{p}'_2)$ with \mathbf{p}_1 . In a special case where $m_1 = 0$, we have

$$E'_1 = \frac{m_2}{1 - \cos \theta_1 + \frac{m_2}{E_1}}$$

Suppose in the center of mass frame (C frame), the scattering angle is χ , then we can derive that

$$E'_1 = E_1 - \Delta E \quad E'_2 = m_2 + \Delta E$$

Here,

$$\Delta E = \frac{m_2(E_1^2 - m_1^2)}{m_1^2 + m_2^2 + 2m_2 E_1} (1 - \cos \chi)$$

Next, suppose in L frame, let $x = p'_1 \cos \theta_1, y = p'_1 \sin \theta_1$, we can get that

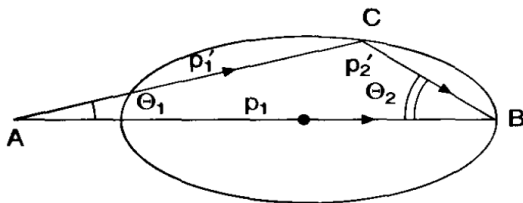
$$\frac{(x - c)^2}{a^2} + \frac{y^2}{b^2} = 1$$

Here,

$$a = \frac{p_1(E_1 m_2 + m_2^2)}{m_1^2 + m_2^2 + 2m_2 E_1} = \frac{m_2 V}{\sqrt{1 - V^2}} \quad b = \frac{m_2 p_1}{\sqrt{m_1^2 + m_2^2 + 2m_2 E_1}} = \frac{a}{\sqrt{1 - V^2}} \quad c = \frac{p_1(E_1 m_2 + m_1^2)}{m_1^2 + m_2^2 + 2m_2 E_1}$$

Here $V = \frac{p_1}{E_1 + m_2}$ is the velocity of particle 2 before scattering in C frame. And it is easy to see that $a + c = p_1$. The result above can be represented by a picture.

(a) $m_1 > m_2$



(b) $m_1 < m_2$

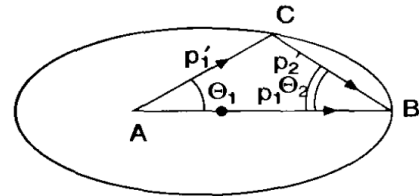


Figure 1.1: Relativistic scattering



Chapter 2

Classical field theory



2.1 Lagrangian formulation

$$S = \int \mathcal{L}(\phi_a, \dot{\phi}_a, \nabla \phi_a) d^4x, \quad \delta\phi_a|_{\Sigma} = 0$$

$$\delta S = 0 \Rightarrow \partial_{\mu} \left(\frac{\partial \mathcal{L}}{\partial(\partial_{\mu} \phi_a)} \right) - \frac{\partial \mathcal{L}}{\partial \phi_a} = 0$$

Locality of the theory

There are no terms in the Lagrangian coupling $\phi(\mathbf{x}, t)$ directly to $\phi(\mathbf{y}, t)$ with $\mathbf{x} \neq \mathbf{y}$. The closest we get for the \mathbf{x} label is coupling between $\phi(\mathbf{x}, t)$ and $\phi(\mathbf{x} + \delta\mathbf{x}, t)$ through the gradient term $\nabla \phi$.

Lorentz invariance

Scalar fields:

$$\bar{\phi}(x) = \phi(\Lambda^{-1}x)$$

Vector fields:

$$\bar{A}^{\mu}(x) = \Lambda^{\mu}_{\nu} A^{\nu}(\Lambda^{-1}x)$$

$$\bar{A}_{\mu}(x) = (\Lambda^{-1})^{\nu}_{\mu} A_{\nu}(\Lambda^{-1}x) = \Lambda_{\mu}^{\nu} A_{\nu}(\Lambda^{-1}x)$$

$$\overline{\partial_{\mu} \phi}(x) = (\Lambda^{-1})^{\nu}_{\mu} \partial_{\nu} \phi(\Lambda^{-1}x) = \Lambda_{\mu}^{\nu} \partial_{\nu} \phi(\Lambda^{-1}x)$$

Lagrangian is a scalar, or more loosely, action is invariant under Lorentz transformation.

2.2 Symmetry and conservation law

Theorem 2.1 Noether's theorem

Every continuous symmetry of the Lagrangian gives rise to a conserved current $j^\mu(x)$ such that the equation of motion imply $\partial_\mu j^\mu = 0$. Suppose that the infinitesimal transformation is

$$\phi_a \rightarrow \phi_a + \delta\phi_a$$

$$\mathcal{L} \rightarrow \mathcal{L} + \delta\mathcal{L}$$

and if $\delta\mathcal{L} = \partial_\mu K^\mu$, we can get

$$j^\mu = -\frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi_a)}\delta\phi_a + K^\mu$$

space-time translation

$$\bar{x} = x - a$$

$$j^\mu = a_\nu T^{\mu\nu}$$

$$T^{\mu\nu} \equiv -\frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi_a)}\partial^\nu\phi_a + \eta^{\mu\nu}\mathcal{L}$$

If we define $P^\mu \equiv \int T^{0\mu} d^3x$, then we have the law of momentum conservation:

$$\frac{dP^\mu}{dt} = 0$$

Lorentz Transformation

$$\bar{x}^\mu = x^\mu + \delta\omega^\mu{}_\nu x^\nu$$

The infinitesimal Lorentz transformation can be written as $I + \delta\omega^\mu{}_\nu$

$$\delta\omega^\mu{}_\nu = \begin{bmatrix} 0 & \beta_1 & \beta_2 & \beta_3 \\ \beta_1 & 0 & -\theta_3 & \theta_2 \\ \beta_2 & \theta_3 & 0 & -\theta_1 \\ \beta_3 & -\theta_2 & \theta_1 & 0 \end{bmatrix}$$

This time, we assume that

$$\bar{\phi}_a(x) = \mathcal{S}_a{}^b \phi_b(\Lambda^{-1}x)$$

In the limit of infinitesimal Lorentz transformation, we have

$$\mathcal{S}_a{}^b = \delta_a{}^b + \frac{1}{2}\delta\omega_{\alpha\beta}(\Sigma^{\alpha\beta})_a{}^b$$

$$j^\mu = \frac{1}{2}M^{\mu\nu\rho}\delta\omega_{\nu\rho}$$



$$M^{\mu\nu\rho} \equiv x^\nu T^{\mu\rho} - x^\rho T^{\mu\nu} - \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi_a)} (\Sigma^{\nu\rho})_a{}^b \phi_b$$

If we define $M^{\nu\rho} \equiv \int M^{0\nu\rho} d^3x$, then we have the law of angular momentum conservation:

$$\frac{dM^{\nu\rho}}{dt} = 0$$

2.3 Functional derivatives

Definition 2.1 Functional derivatives

Given a manifold M representing (continuous/smooth) functions ρ (with certain boundary conditions etc.), and a functional F defined as

$$F: M \rightarrow \mathbb{R} \quad \text{or} \quad F: M \rightarrow \mathbb{C},$$

the functional derivative of $F[\rho]$, denoted $\frac{\delta F}{\delta \rho}$, is defined by

$$\begin{aligned} \int \frac{\delta F}{\delta \rho}(x) \phi(x) dx &= \lim_{\epsilon \rightarrow 0} \frac{F[\rho + \epsilon \phi] - F[\rho]}{\epsilon} \\ &= \left[\frac{d}{d\epsilon} F[\rho + \epsilon \phi] \right]_{\epsilon=0}, \end{aligned}$$

where ϕ is an arbitrary function. The quantity $\epsilon \phi$ is called the variation of ρ .

Like the derivative of a function, the functional derivative satisfies the following properties, where $F[\rho]$ and $G[\rho]$ are functionals:

Linearity:

$$\frac{\delta(\lambda F + \mu G)[\rho]}{\delta \rho(x)} = \lambda \frac{\delta F[\rho]}{\delta \rho(x)} + \mu \frac{\delta G[\rho]}{\delta \rho(x)},$$

where λ, μ are constants.

Product rule:

$$\frac{\delta(FG)[\rho]}{\delta \rho(x)} = \frac{\delta F[\rho]}{\delta \rho(x)} G[\rho] + F[\rho] \frac{\delta G[\rho]}{\delta \rho(x)},$$

Chain rules:

If F is a functional and G an operator, then

$$\frac{\delta F[G[\rho]]}{\delta \rho}(y) = \int dx \frac{\delta F[G[\rho]]}{\delta G[\rho]}(x) \cdot \frac{\delta G[\rho]}{\delta \rho}(x, y)$$

If G is an ordinary differentiable function g , then this reduces to

$$\frac{\delta F[g(\rho)]}{\delta \rho}(y) = \frac{\delta F[g(\rho)]}{\delta g[\rho]}(y) \frac{dg(\rho)}{d\rho}(y)$$



Proposition 2.1 Properties of functional derivatives

$$\frac{\delta F}{\delta \rho}(y) = \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon} \{F[\rho(x) + \epsilon \delta(x - y)] - F[\rho(x)]\}$$

$$\frac{\delta f(x)}{\delta f(y)} = \delta(x - y)$$

$$\frac{\delta}{\delta f(y)} \int g(f(x)) dx = g'(f(y))$$

$$\frac{\delta f'(x)}{\delta f(y)} = \frac{d}{dx} \delta(x - y)$$

$$\frac{\delta}{\delta f(y)} \int g(f'(x)) dx = -\frac{d}{dy} g'(f'(y))$$

**2.4 Hamiltonian formulation**

$$\pi^a(x) = \frac{\partial \mathcal{L}}{\partial \dot{\phi}_a}$$

$$\mathcal{H}(\phi_a, \nabla \phi_a, \pi^a) = \pi^a \dot{\phi}_a - \mathcal{L}$$

$$H = \int \mathcal{H} d^3x$$

Now, we can get the Hamilton equation form $\delta S = 0$,

$$\dot{\phi}_a(\mathbf{x}, t) = \frac{\delta}{\delta \pi^a(\mathbf{x}, t)} H = \frac{\partial \mathcal{H}}{\partial \pi^a}$$

$$\dot{\pi}^a(\mathbf{x}, t) = -\frac{\delta}{\delta \phi_a(\mathbf{x}, t)} H = -\frac{\partial \mathcal{H}}{\partial \phi_a} + \left(\frac{\partial \mathcal{H}}{\partial \phi_{a,b}} \right)_{,b}$$

2.4.1 Poisson bracket

First, we define that

$$[\phi_a(\mathbf{x}), \phi_b(\mathbf{y})] \equiv [\pi^a(\mathbf{x}), \pi^b(\mathbf{y})] \equiv 0$$

$$[\phi_a(\mathbf{x}), \pi^b(\mathbf{y})] \equiv \delta_a^b \delta(\mathbf{x} - \mathbf{y})$$

then, we demand that the bracket operation has the same properties as the Poisson bracket in classical mechanics. And we also demand that

$$[\partial_x A(\mathbf{x}), B(\mathbf{y})] = \partial_x [A(\mathbf{x}), B(\mathbf{y})]$$

and

$$\left[\int d^3x A(\mathbf{x}), B(\mathbf{y}) \right] = \int d^3x [A(\mathbf{x}), B(\mathbf{y})]$$



As a result, we can verify that

$$[W[\phi(\mathbf{x}), \pi(\mathbf{x})], Z[\phi(\mathbf{x}), \pi(\mathbf{x})]] = \int d^3x \left\{ \frac{\delta W}{\delta \phi(\mathbf{x})} \frac{\delta Z}{\delta \pi(\mathbf{x})} - \frac{\delta W}{\delta \pi(\mathbf{x})} \frac{\delta Z}{\delta \phi(\mathbf{x})} \right\}$$

Especially,

$$[\phi_a(\mathbf{x}), H] = \frac{\delta}{\delta \pi^a(\mathbf{x})} H, \quad [\pi^a(\mathbf{x}), H] = -\frac{\delta}{\delta \phi_a(\mathbf{x})} H$$

So, the Hamilton equation can be written as

$$\dot{\phi}_a = [\phi_a, H], \quad \dot{\pi}^a = [\pi^a, H]$$

Further more, we can prove that

$$\frac{dO(\phi, \pi, t)}{dt} = [O, H] + \frac{\partial O}{\partial t}$$

and

$$\frac{d[A, B]}{dt} = [A, \frac{dB}{dt}] + [\frac{dA}{dt}, B]$$

2.4.2 Momentum

It is easy to verify that

$$P^0 = H, \quad P^i = \int -\pi^a \partial^i \phi_a d^3x$$

And we can get the Poisson bracket

$$\begin{aligned} [\phi_a, P^\mu] &= -\partial^\mu \phi_a \\ [\pi^a, P^\mu] &= -\partial^\mu \pi^a \\ [P^\mu, P^\nu] &= 0 \end{aligned}$$

2.4.3 Angular momentum

It is easy to verify that

$$M^{\mu\nu} = \int (x^\mu T^{0\nu} - x^\nu T^{0\mu} - \pi^a (\Sigma^{\mu\nu})_a^b \phi_b) d^3x$$

We define that

$$\begin{aligned} M_L^{\mu\nu} &\equiv \int (x^\mu T^{0\nu} - x^\nu T^{0\mu}) d^3x & M_S^{\mu\nu} &\equiv \int (-\pi^a (\Sigma^{\mu\nu})_a^b \phi_b) d^3x \\ (L^{\mu\nu})_a^b &\equiv -(x^\mu \partial^\nu - x^\nu \partial^\mu) \delta_a^b & (S^{\mu\nu})_a^b &\equiv -(\Sigma^{\mu\nu})_a^b \end{aligned}$$

So, we can get the Poisson bracket

$$\begin{aligned} [\phi_a, M_L^{\mu\nu}] &= (L^{\mu\nu})_a^b \phi_b & [\phi_a, M_S^{\mu\nu}] &= (S^{\mu\nu})_a^b \phi_b \\ [\pi^a, M_L^{\mu\nu}] &= (L^{\mu\nu})_b^a \pi^b & [\pi^a, M_S^{\mu\nu}] &= -(S^{\mu\nu})_b^a \pi^b \end{aligned}$$



Because $\frac{dM^{\mu\nu}}{dt} = 0$, $M^{\mu\nu}$ can commute with $\frac{d}{dt}$, so

$$[[\phi(x), M^{\mu\nu}], M^{\rho\sigma}] = (L^{\mu\nu} + S^{\mu\nu})(L^{\rho\sigma} + S^{\rho\sigma})\phi(x)$$

At last, we can get the Poisson bracket

$$[\phi(x), [M^{\mu\nu}, M^{\rho\sigma}]] = (L^{\mu\nu}L^{\rho\sigma} - L^{\rho\sigma}L^{\mu\nu} + S^{\mu\nu}S^{\rho\sigma} - S^{\rho\sigma}S^{\mu\nu})\phi(x)$$

Since we can prove that

$$L^{\mu\nu}L^{\rho\sigma} - L^{\rho\sigma}L^{\mu\nu} = -\eta^{\nu\rho}L^{\mu\sigma} + \eta^{\sigma\mu}L^{\rho\nu} + \eta^{\mu\rho}L^{\nu\sigma} - \eta^{\sigma\nu}L^{\rho\mu}$$

if we demand that

$$S^{\mu\nu}S^{\rho\sigma} - S^{\rho\sigma}S^{\mu\nu} = -\eta^{\nu\rho}S^{\mu\sigma} + \eta^{\sigma\mu}S^{\rho\nu} + \eta^{\mu\rho}S^{\nu\sigma} - \eta^{\sigma\nu}S^{\rho\mu}$$

We will get the Poisson bracket of the $M^{\mu\nu}$,

$$[M^{\mu\nu}, M^{\rho\sigma}] = -\eta^{\nu\rho}M^{\mu\sigma} + \eta^{\sigma\mu}M^{\rho\nu} + \eta^{\mu\rho}M^{\nu\sigma} - \eta^{\sigma\nu}M^{\rho\mu}$$

up to the possibility of a term on the right-hand side that commutes with $\phi(x)$ and its derivatives.

We now define $J_i \equiv \frac{1}{2}\epsilon_{ijk}M^{jk}$ and $K_i \equiv M^{i0}$, so we have

$$M^{\mu\nu} = \begin{bmatrix} 0 & -K_1 & -K_2 & -K_3 \\ K_1 & 0 & J_3 & -J_2 \\ K_2 & -J_3 & 0 & J_1 \\ K_3 & J_2 & -J_1 & 0 \end{bmatrix} \quad \left(\delta\omega_{\mu\nu} = \begin{bmatrix} 0 & -\beta_1 & -\beta_2 & -\beta_3 \\ \beta_1 & 0 & -\theta_3 & \theta_2 \\ \beta_2 & \theta_3 & 0 & -\theta_1 \\ \beta_3 & -\theta_2 & \theta_1 & 0 \end{bmatrix} \right)$$

the Poisson bracket can be written as

$$\begin{aligned} [J_i, J_j] &= \epsilon_{ijk}J_k \\ [J_i, K_j] &= \epsilon_{ijk}K_k \\ [K_i, K_j] &= -\epsilon_{ijk}J_k \end{aligned}$$

We can use the similar method to derive that

$$[P^\mu, M^{\rho\sigma}] = \eta^{\mu\sigma}P^\rho - \eta^{\mu\rho}P^\sigma$$

It can also be written as

$$\begin{aligned} [J_i, H] &= 0 \\ [J_i, P_j] &= \epsilon_{ijk}P_k \\ [K_i, H] &= P_i \\ [K_i, P_j] &= \delta_{ij}H \end{aligned}$$

At last, we define $L_i \equiv \frac{1}{2}\epsilon_{ijk}M_L^{jk}$ and $S_i \equiv \frac{1}{2}\epsilon_{ijk}M_S^{jk}$ we can demonstrate that

$$\begin{aligned} [L_i, S_j] &= 0 \\ [S_i, P_j] &= 0 \\ [L_i, P_j] &= \epsilon_{ijk}P_k \end{aligned}$$



Chapter 3

Classical Electrodynamics



3.1 The formulation of classical electrodynamics

3.1.1 Maxwell equations and Lorentz force

The Lagrangian of the EM field A^μ when coupling with current is

$$\mathcal{L} = -\frac{1}{4}F^{\mu\nu}F_{\mu\nu} + j^\mu A_\mu$$

Here

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \quad j^\mu = \rho_{e0}u^\mu$$

ρ_{e0} is the charge density measured in the frame of charge. Then we can derive the Euler-Lagrange equation of EM field:

$$\partial_\nu F^{\mu\nu} = j^\mu$$

And we can derive the charge conservation equation

$$\partial_\mu j^\mu = 0$$

For a charged particle moving in EM field with orbit $x^\mu(\tau)$, we have

$$j^\mu = e_a \delta(\mathbf{r} - \mathbf{r}_a(t)) \sqrt{1 - \mathbf{v}_a^2} u_a^\mu$$

Then we can derive that

$$\int dV dt j^\mu A_\mu = \int dx^\mu e A_\mu$$

So, the action for a charged particle when coupling with EM field is

$$S = -m \int d\tau + e \int dx^\mu A_\mu$$

We can derive the Euler-Lagrangian equation of the particle:

$$ma_\mu = eF_{\mu\nu}u^\nu$$

The Hamiltonian formulation of electrodynamics will be discussed in detail in the Hamiltonian formulation of general relativity and canonical quantization formulation of quantum electrodynamics.

We now define the electric field and magnetic field as follows,

$$E^i \equiv F^{0i} = -\dot{A}^i - \nabla^i A^0 \quad B^i \equiv \epsilon_{ijk} \nabla^j A^k$$

We also define $\rho_e \equiv j^0$ and $J^i \equiv j^i$, so the field equation can then be written

$$\begin{aligned}\nabla \times \mathbf{B} &= \frac{\partial \mathbf{E}}{\partial t} + \mathbf{J} \\ \nabla \times \mathbf{E} &= -\frac{\partial \mathbf{B}}{\partial t} \\ \nabla \cdot \mathbf{E} &= \rho_e \\ \nabla \cdot \mathbf{B} &= 0\end{aligned}$$

That is the so-called Maxwell equation.

The equation of motion of the charged particle can be written as

$$\frac{d\mathbf{p}}{dt} = e(\mathbf{E} + \mathbf{v} \times \mathbf{B}) \quad \frac{d\mathcal{E}}{dt} = e\mathbf{E} \cdot \mathbf{v}$$

That is the so-called Lorentz force.

We note that the Field A cannot be completely determined by Maxwell and Lorentz equation. If we make the transformation $A_\mu \rightarrow A_\mu + \partial_\mu \xi(x)$, \mathcal{L} and $F^{\mu\nu}$ would be invariant, and the Maxwell and Lorentz equation is still valid. This arbitrariness of ξ is called gauge invariance. The topic will be discussed in detail in QED.

3.1.2 Lorentz transformation of fields

In the new coordinate system $\bar{x}^\mu = \Lambda^\mu_\nu x^\nu$, we have

$$\bar{A}^\mu(\bar{x}) = \Lambda^\mu_\nu A^\nu(\Lambda^{-1}\bar{x})$$

In a special case when the new reference frame move along $\hat{1}$ direction with velocity β , we have

$$\bar{x}^0 = \gamma x^0 - \gamma\beta x^1 \quad \bar{x}^1 = -\gamma\beta x^0 + \gamma x^1$$

So, we have

$$\bar{A}^0 = \gamma A^0 - \gamma\beta A^1 \quad \bar{A}^1 = -\gamma\beta A^0 + \gamma A^1$$

We can further derive that

$$\begin{aligned}\bar{E}_1 &= E_1 & \bar{E}_2 &= \gamma E_2 - \gamma\beta B_3 & \bar{E}_3 &= \gamma E_3 + \gamma\beta B_2 \\ \bar{B}_1 &= B_1 & \bar{B}_2 &= \gamma B_2 + \gamma\beta E_3 & \bar{B}_3 &= \gamma B_3 - \gamma\beta E_2\end{aligned}$$

It can be written in the three vector form as

$$\begin{aligned}\bar{\mathbf{E}} &= \gamma(\mathbf{E}_\perp + \boldsymbol{\beta} \times \mathbf{B}) + \mathbf{E}_\parallel \\ \bar{\mathbf{B}} &= \gamma(\mathbf{B}_\perp - \boldsymbol{\beta} \times \mathbf{E}) + \mathbf{B}_\parallel\end{aligned}$$

If $\beta \ll 1$, neglect the higher order of β^2 , we have

$$\begin{aligned}\bar{\mathbf{E}} &= \mathbf{E} + \boldsymbol{\beta} \times \mathbf{B} \\ \bar{\mathbf{B}} &= \mathbf{B} - \boldsymbol{\beta} \times \mathbf{E}\end{aligned}$$

We also note that $F_{\mu\nu}F^{\mu\nu}$ and $\epsilon_{\mu\nu\rho\sigma}F^{\mu\nu}F^{\rho\sigma}$ is invariant under Lorentz transformation. It can be represented by the electric and magnetic field as

$$E^2 - B^2 = \text{inv} \quad \mathbf{E} \cdot \mathbf{B} = \text{inv}$$



3.1.3 Energy-momentum tensor

For a free EM field, the energy-momentum tensor is

$$T_f^{\mu\nu} \equiv -\frac{\partial \mathcal{L}}{\partial(\partial_\mu A_\rho)} \partial^\nu A_\rho + \eta^{\mu\nu} \mathcal{L} = \partial^\nu A^\rho F^\mu_\rho - \frac{1}{4} \eta^{\mu\nu} F_{\rho\sigma} F^{\rho\sigma}$$

We note that the energy-momentum tensor defined above is not symmetric, so we define a modified energy-momentum tensor by adding a term $-\partial^\rho A^\nu F^\mu_\rho$, so

$$T_e^{\mu\nu} = F^{\nu\rho} F^\mu_\rho - \frac{1}{4} \eta^{\mu\nu} F_{\rho\sigma} F^{\rho\sigma}$$

Note that for free EM field, $\partial^\rho A^\nu F^\mu_\rho = \partial^\rho (A^\nu F^\mu_\rho)$. So, we can get that

$$\partial_\mu T_f^{\mu\nu} = 0 \quad P_f'^\mu = P_f^\mu$$

So, from now on, we will use T_f' as the energy-momentum tensor of EM field and omit the prime of T_f' for simplicity. The momentum of the free EM field is

$$P_f^0 = \int dV \frac{1}{2} (\mathbf{E}^2 + \mathbf{B}^2) \equiv \int dV w \quad P_f^i = \int dV \mathbf{E} \times \mathbf{B} \equiv \int dV \mathbf{S}$$

If there also exists charged particles in the system, i.e. the source of EM field, we must also include the energy-momentum tensor of the particles to get the right conservation equation. The energy-momentum tensor of particles is defined as

$$T_p^{\mu\nu} \equiv \sum_a m_a \delta(\mathbf{r} - \mathbf{r}_a) \sqrt{1 - \mathbf{v}_a^2} u_a^\mu u_a^\nu$$

In this definition, we have

$$P_p^0 = \sum_a \frac{m_a}{\sqrt{1 - \mathbf{v}_a^2}} \quad \mathbf{P}_p = \sum_a \frac{m_a}{\sqrt{1 - \mathbf{v}_a^2}} \mathbf{v}_a$$

So, our definition is consistent with the definition of four momentum in previous chapter. Recall that

$$j_e^\mu = \sum_a e_a \delta(\mathbf{r} - \mathbf{r}_a) \sqrt{1 - \mathbf{v}_a^2} u_a^\mu$$

We can define the mass current as

$$j_m^\mu \equiv \sum_a m_a \delta(\mathbf{r} - \mathbf{r}_a) \sqrt{1 - \mathbf{v}_a^2} u_a^\mu$$

If there is no particle creation and annihilation, and the mass of the particle is constant during the motion, we would have

$$\partial_\mu j_{ma}^\mu = 0$$

Recall the Lorentz force equation $ma_\mu = eF_{\mu\nu} u^\nu$, it can be rewritten as

$$\rho_{m0a} \frac{du_\mu}{d\tau} = F_{\mu\nu} j_{ea}^\nu$$



Here, $\rho_{m0a} \equiv m_a \delta(\mathbf{r} - \mathbf{r}_a) \sqrt{1 - \mathbf{v}_a^2}$ is the mass density measured in the rest frame of the particle. Then, on the one hand, we have

$$\partial_\mu T_p^{\mu\nu} = \sum_a \rho_{m0a} u^\mu \frac{\partial u_a^\nu}{\partial x^\mu} = \sum_a \rho_{m0a} \frac{du_a^\nu}{d\tau} = F^\nu{}_\mu j_e^\mu$$

On the other hand, we can derive that

$$\partial_\mu T_f^{\mu\nu} = -F^\nu{}_\mu j_e^\mu$$

by implementing Maxwell equation. Define

$$T^{\mu\nu} \equiv T_f^{\mu\nu} + T_p^{\mu\nu}$$

We have

$$\partial_\mu T^{\mu\nu} = 0$$

We define the Maxwell stress tensor as

$$f^{ij} \equiv -T_f^{ij} = E^i E^j + B^i B^j - w \delta^{ij}$$

Then, we can write down the conservation law of energy and momentum as

$$\begin{aligned} \frac{d}{dt} \left(P_p^0 + \int w dV \right) &= - \oint \mathbf{S} \cdot d\boldsymbol{\sigma} \\ \frac{d}{dt} \left(\mathbf{P}_p + \int \mathbf{S} dV \right) &= \oint \mathbf{f} \cdot d\boldsymbol{\sigma} \end{aligned}$$

We assume there is no particle cross the boundary.

At last, because $T^{\mu\nu} = T^{\nu\mu}$, we have

$$\partial_\mu (x^\nu T^{\mu\rho} - x^\rho T^{\mu\nu}) = 0$$

It is easy to write down the conservation law of angular momentum as

$$\frac{d}{dt} \left(\mathbf{L}_p + \int \mathbf{r} \times \mathbf{S} dV \right) = \oint \mathbf{r} \times \mathbf{f} \cdot d\boldsymbol{\sigma}$$

Here,

$$\mathbf{L}_p = \sum_a \frac{m_a}{\sqrt{1 - \mathbf{v}_a^2}} \mathbf{r}_a \times \mathbf{v}_a$$

3.1.4 Charged particles in a given EM field

Now we suppose the EM field is given, i.e. we will neglect the EM field generated by the test charged particles. Then we have

$$S = \int_{t_1}^{t_2} (-\sqrt{1 - v^2} m + e \mathbf{A} \cdot \mathbf{v} - e\phi) dt$$

So,

$$L = -\sqrt{1 - v^2} m + e \mathbf{A} \cdot \mathbf{v} - e\phi$$



The canonical momentum is

$$\boldsymbol{\pi} = \frac{\partial L}{\partial \mathbf{v}} = \gamma m \mathbf{v} + e \mathbf{A}$$

The Hamiltonian is therefore

$$H = \boldsymbol{\pi} \cdot \mathbf{v} - L = \gamma m + e\phi = \sqrt{m^2 + (\boldsymbol{\pi} - e\mathbf{A})^2} + e\phi$$

If $v \ll 1$, we would have

$$L = \frac{mv^2}{2} + e\mathbf{A} \cdot \mathbf{v} - e\phi \quad \boldsymbol{\pi} = m\mathbf{v} + e\mathbf{A} \quad H = \frac{(\boldsymbol{\pi} - e\mathbf{A})^2}{2m} + e\phi$$

If the EM field is constant in time, we have

$$\nabla \times \mathbf{E} = 0$$

we would choose the gauge that

$$\dot{\mathbf{A}} = 0 \quad \mathbf{E} = -\nabla\phi$$

Because $\frac{\partial L}{\partial t} = 0$, we can know that $\gamma m + e\phi$ is a constant.

Now we would list some special case.

Motion in a uniform and constant electric field

Suppose the the direction of electric field is \hat{x} , the orbit is in the $x - y$ plane. So, the equation of motion is

$$\dot{p}_x = eE \quad \dot{p}_y = 0$$

The final solution is

$$x = \frac{1}{eE} \sqrt{\mathcal{E}_0^2 + (eEt)^2} \quad y = \frac{p_0}{eE} \operatorname{arcsinh} \frac{eEt}{\mathcal{E}_0}$$

Here, we assume when $t = 0$, $p_x = 0, p_y = p_0$. The orbit function is

$$x = \frac{\mathcal{E}_0}{eE} \cosh \frac{eEy}{p_0}$$

Motion in a uniform and constant magnetic field

Suppose the direction of magnetic field is \hat{z} . Note that particle's kinetic energy $\mathcal{E} = \gamma m$ is constant if there are no electric field, we can derive the equation of motion

$$\dot{v}_x = \omega v_y \quad \dot{v}_y = -\omega v_x \quad \dot{v}_z = 0$$

Here, $\omega = \frac{eB}{\gamma m}$. The final solution is

$$x = x_0 + r \sin(\omega t + \alpha) \quad y = y_0 + r \cos(\omega t + \alpha) \quad z = z_0 + v_{0z}t$$

Here, $x_0, y_0, z_0, r, \alpha, v_{0z}$ should be determined by initial condition.



Motion in a uniform and constant EM field

We only focus on the case that the velocity of particle is much smaller the velocity of light. Suppose the direction of magnetic field is \hat{z} , the direction of electric field is within $y - z$ plane. The equation of motion is

$$m\ddot{x} = eB\dot{y} \quad m\ddot{y} = eE_y - eB\dot{x} \quad m\ddot{z} = eE_z$$

The solution is

$$\dot{x} = a \cos \omega t + \frac{E_y}{B} \quad \dot{y} = -a \sin \omega t \quad \dot{z} = v_{0z} + \frac{eE_z}{m}t$$

Here, $\omega = \frac{eB}{m}$, a, v_{0z} is determined by initial condition. As we suppose that $v \ll 1$ is satisfied, we must have that

$$a \ll 1 \quad v_{0z} \ll 1 \quad \frac{eE_z t}{m} \ll 1 \quad E_y \ll B$$

3.2 Constant electromagnetic field

3.2.1 Coulomb' law

For constant electric field, the Maxwell equation take the form

$$\nabla \cdot \mathbf{E} = \rho_e \quad \nabla \times \mathbf{E} = 0$$

So, we have

$$\mathbf{E} = -\nabla \phi \quad \nabla^2 \phi = -\rho_e$$

The solution is

$$\phi(\mathbf{r}) = \int \frac{\rho_e(\mathbf{r}')}{4\pi|\mathbf{r} - \mathbf{r}'|} dV'$$

If $\rho_e(\mathbf{r}') = Q\delta(\mathbf{r}')$, we have

$$\phi(\mathbf{r}) = \frac{Q}{4\pi|\mathbf{r}|} \quad \mathbf{E}(\mathbf{r}) = \frac{Q\mathbf{r}}{4\pi|\mathbf{r}|^3}$$

For a system of static charged particles, the total energy is

$$U = \frac{1}{2} \int E^2 dV = \frac{1}{2} \int \rho \phi dV = \frac{1}{2} \sum e_a \phi_a + \frac{1}{2} \sum e_a \Phi_a$$

Here, ϕ_a is the electric potential at the point where e_a is located, produced by e_a itself, while Φ_a is the potential produced by other charges. It is obvious that $U_{self} = \frac{1}{2} e_a \phi_a$ is infinite, indicating that classical electrodynamics is no more valid in small distance. This problem will be solved in quantum electrodynamics: the mass of charged particle we measured is already renormalized to include the electromagnetic self energy. So, actually, we have

$$U = \frac{1}{2} \int E^2 dV - U_{self} = \frac{1}{2} \sum_{a \neq b} \frac{e_a e_b}{4\pi R_{ab}}$$



If the charged particle is moving with a constant velocity, we can derive the electric field it produced by Lorentz transformation, the final result is that

$$\mathbf{E} = \frac{e\mathbf{R}}{4\pi R^3} \frac{1 - V^2}{(1 - V^2 \sin^2 \theta)^{3/2}} \quad \mathbf{B} = \mathbf{V} \times \mathbf{E}$$

Here, \mathbf{R} is the vector point from the particle to the point we measure the electric field, and θ is the angle between \mathbf{V} and \mathbf{R} . If $V \sim 1$, the electric field will be concentrated in the direction perpendicular to the \mathbf{V} . If $V \ll 1$, we have

$$\mathbf{E} = \frac{e\mathbf{R}}{4\pi R^3} \quad \mathbf{B} = \frac{e\mathbf{V} \times \mathbf{R}}{4\pi R^3}$$

3.2.2 Multipole moments

For a system of charged particles, the potential it produced at \mathbf{R} is

$$\phi = \sum_a \frac{e_a}{4\pi |\mathbf{R} - \mathbf{r}_a|}$$

If $R \gg r_a$, we can expand the equation around $r_a = 0$. Generally, we have

$$\frac{1}{|\mathbf{R} - \mathbf{r}|} = \sum_{l=0}^{\infty} \sum_{m=-l}^l \frac{r^l}{R^{l+1}} \frac{4\pi}{2l+1} Y_{lm}^*(\Theta, \Phi) Y_{lm}(\theta, \phi)$$

So, we have

$$\phi = \sum \phi^{(l)}$$

Here,

$$\begin{aligned} \phi^{(l)} &= \frac{1}{4\pi R^{l+1}} \sum_{m=-l}^l \sqrt{\frac{4\pi}{2l+1}} Q_m^{(l)} Y_{lm}^*(\Theta, \Phi) \\ Q_m^{(l)} &= \sum_a e_a r_a^l \sqrt{\frac{4\pi}{2l+1}} Y_{lm}(\theta_a, \phi_a) \end{aligned}$$

Special case:

$$\phi^{(0)} = \frac{Q}{4\pi R} \quad E^{(0)} = \frac{Q\mathbf{n}}{4\pi R^2} \quad Q = \sum_a e_a$$

$$\phi^{(1)} = \frac{\mathbf{d} \cdot \mathbf{n}}{4\pi R^2} \quad E^{(1)} = \frac{3(\mathbf{d} \cdot \mathbf{n})\mathbf{n} - \mathbf{d}}{4\pi R^3} \quad \mathbf{d} = \sum_a e_a \mathbf{r}_a$$

$$\phi^{(2)} = \frac{\mathbf{n} \cdot \mathbf{D} \cdot \mathbf{n}}{8\pi R^3} \quad E^{(2)} = \frac{5(\mathbf{n} \cdot \mathbf{D} \cdot \mathbf{n})\mathbf{n} - (\mathbf{n} \cdot \mathbf{D} + \mathbf{D} \cdot \mathbf{n})}{8\pi R^4} \quad D_{ij} = \sum e(3x_i x_j - r^2 \delta_{ij})$$

For a system of charged particles in the electric field $\phi(\mathbf{r})$, if all the particles is near the $r = 0$, we can make the expansion

$$\phi(\mathbf{r}) = \sum_{l=0}^{\infty} r^l \sum_{m=-l}^{m=l} a_{lm} \sqrt{\frac{4\pi}{2l+1}} Y_{lm}(\theta, \phi)$$



So

$$U = \sum_{l=0}^{\infty} U^{(l)} \quad U^{(l)} = \sum_{m=-l}^l a_{lm} Q_m^{(l)}$$

Special case:

$$\begin{aligned} U^{(0)} &= Q\phi_0 & F^{(0)} &= Q\mathbf{E}_0 \\ U^{(1)} &= -\mathbf{d} \cdot \mathbf{E}_0 & F^{(1)} &= \mathbf{d} \cdot \nabla \mathbf{E}_0 & M^{(1)} &= \mathbf{d} \times \mathbf{E}_0 \\ U^{(2)} &= -\frac{1}{6} \mathbf{D} \cdot \nabla \mathbf{E}_0 & F^{(2)} &= \frac{1}{6} \mathbf{D} \cdot \nabla \nabla \mathbf{E}_0 & M^{(2)} &= \frac{1}{3} \nabla \cdot (\mathbf{D} \times \mathbf{E}_0) \end{aligned}$$

3.2.3 Biot-Savart law

Let us consider the magnetic field produced by charges which perform a finite motion, in which the particles are always within a finite region of space and the momenta also always remain finite. Consider the time average magnetic field $\bar{\mathbf{B}}$, produced by the charges; this field will now be a function only of the coordinates and not of the time. We take the time average of the Maxwell equations

$$\nabla \cdot \mathbf{B} = 0 \quad \nabla \times \mathbf{B} = \frac{\partial \mathbf{E}}{\partial t} + \mathbf{j}$$

Note that the average value of the derivative $\partial \mathbf{E} / \partial t$, like the derivative of any quantity which varies over a finite range, is zero. We can get

$$\nabla \cdot \bar{\mathbf{B}} = 0 \quad \nabla \times \bar{\mathbf{B}} = \bar{\mathbf{j}}$$

Recall that $\mathbf{B} = \nabla \times \mathbf{A}$. We impose the gauge condition $\nabla \cdot \mathbf{A} = 0$, we have

$$\nabla^2 \bar{\mathbf{A}} = -\bar{\mathbf{j}}$$

The solution is

$$\bar{\mathbf{A}}(\mathbf{r}) = \frac{1}{4\pi} \int \frac{\bar{\mathbf{j}}(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|} dV' = \frac{1}{4\pi} \sum \frac{\overline{e_a \mathbf{v}_a}}{|\mathbf{r} - \mathbf{r}_a|}$$

And the magnetic field is

$$\bar{\mathbf{B}}(\mathbf{r}) = \frac{1}{4\pi} \int \frac{\bar{\mathbf{j}}(\mathbf{r}') \times \mathbf{R}}{R^3} dV' \quad \mathbf{R} = \mathbf{r} - \mathbf{r}'$$

3.2.4 Magnetic moment

For a system of charged particles, the potential it produced at \mathbf{R} is

$$\bar{\mathbf{A}} = \frac{1}{4\pi} \sum \frac{\overline{e_a \mathbf{v}_a}}{|\mathbf{R} - \mathbf{r}_a|}$$

If $R \gg r_a$, we can expand the equation around $r_a = 0$ the first order,

$$4\pi \bar{\mathbf{A}} = \frac{1}{R} \sum e \bar{\mathbf{v}} - \sum e \mathbf{v} \left(\mathbf{r} \cdot \nabla \frac{1}{R} \right)$$



Firstly,

$$\sum e\bar{\mathbf{v}} = \overline{\frac{d}{dt} \sum e\mathbf{r}} = 0$$

Secondly,

$$-\sum e\mathbf{v} \left(\mathbf{r} \cdot \nabla \frac{1}{R} \right) = \frac{1}{R^3} \sum e\mathbf{v}(\mathbf{r} \cdot \mathbf{R})$$

Note that

$$\sum e\mathbf{v}(\mathbf{r} \cdot \mathbf{R}) = \frac{1}{2} \frac{d}{dt} e\mathbf{r}(\mathbf{r} \cdot \mathbf{R}) + \frac{1}{2} \left(\sum e\mathbf{r} \times \mathbf{v} \right) \times \mathbf{R}$$

Define the magnetic moment as

$$\mathbf{m} \equiv \frac{1}{2} \left(\sum e\mathbf{r} \times \mathbf{v} \right)$$

We can get

$$\bar{\mathbf{A}} = \frac{\bar{\mathbf{m}} \times \mathbf{R}}{4\pi R^3} \quad \bar{\mathbf{B}} = \frac{3\mathbf{n}(\bar{\mathbf{m}} \cdot \mathbf{n}) - \bar{\mathbf{m}}}{4\pi R^3}$$

If all the particles have the same charge-mass ratio, and the velocity of all the particles is much smaller than that of light, we have

$$\mathbf{m} = \frac{e}{2m} \sum m\mathbf{r} \times \mathbf{v} = \frac{e}{2m} \mathbf{M}$$

Let us consider a system of charges in an external constant uniform magnetic field. The time average of the force acting on the system,

$$\bar{\mathbf{F}} = \sum e\bar{\mathbf{v}} \times \bar{\mathbf{B}} = \overline{\frac{d}{dt} \sum e\mathbf{r} \times \mathbf{B}} = 0$$

The average value of the moment of the forces is

$$\bar{\mathbf{K}} = \sum e\mathbf{r} \times (\mathbf{v} \times \mathbf{B})$$

We can derive that

$$\bar{\mathbf{K}} = \bar{\mathbf{m}} \times \mathbf{B}$$

Let us consider the change in the average angular momentum $\bar{\mathbf{M}}$ of the system. According to a well-known equation of mechanics, the derivative of \mathbf{M} is equal to the moment \mathbf{K} of the forces acting on the system. We therefore have

$$\frac{d\bar{\mathbf{M}}}{dt} = \bar{\mathbf{m}} \times \mathbf{B}$$

If the charge-mass ratio is the same for all particles of the system, the angular momentum and magnetic moment are proportional to one another, and we find:

$$\frac{d\bar{\mathbf{M}}}{dt} = -\boldsymbol{\Omega} \times \bar{\mathbf{M}} \quad \boldsymbol{\Omega} = \frac{e}{2m} \mathbf{B}$$

This equation states that the vector $\bar{\mathbf{M}}$ rotates with angular velocity $-\boldsymbol{\Omega}$ around the direction of the field, while its absolute magnitude and the angle which it makes with this direction remain fixed. This motion is called the Larmor precession. Let us consider the Lagrangian for a system of charges in an external constant uniform magnetic field, it contains the additional term

$$L_H = \sum e\mathbf{A} = \sum e(\mathbf{B} \times \mathbf{r}) \cdot \mathbf{v} = \sum e(\mathbf{r} \times \mathbf{v}) \cdot \mathbf{B} = \mathbf{m} \cdot \mathbf{B}$$



3.3 Electromagnetic waves

3.3.1 Electromagnetic waves

electromagnetic fields occurring in vacuum in the absence of charges are called electromagnetic waves. We choose the Coulomb's Gauge, i.e.

$$\phi = 0 \quad \nabla \cdot \mathbf{A} = 0$$

So,

$$\mathbf{E} = -\frac{\partial \mathbf{A}}{\partial t} \quad \mathbf{B} = \nabla \times \mathbf{A}$$

Then, from Maxwell equation we can derive that

$$\nabla^2 \mathbf{A} - \frac{\partial^2 \mathbf{A}}{\partial t^2} = 0$$

This is the equation which determines the potentials of electromagnetic waves. It is called the d'Alembert equation, or the wave equation. We can verify that the electric and magnetic fields \mathbf{E} and \mathbf{H} satisfy the same wave equation.

We consider the special case of electromagnetic waves in which the fields depend only on one coordinate, say x . Such waves are said to be plane. In this case the equation for the field becomes

$$\frac{\partial^2 f}{\partial t^2} - \frac{\partial^2 f}{\partial x^2} = 0$$

The solution is

$$f(t, x) = f_1(t - x) + f_2(t + x)$$

$f_1(t - x)$ represents a plane wave moving in the positive direction along the x axis. $f_2(t + x)$ represents a plane wave moving in the negative direction along the x axis. The Coulomb's gauge would imply that $A_x = 0$, and we can obtain

$$\mathbf{E} = -\mathbf{A}' \quad \mathbf{B} = -\mathbf{n} \times \mathbf{A}' = \mathbf{n} \times \mathbf{E}$$

where the prime denotes differentiation with respect to $t - x$ and \mathbf{n} is a unit vector along the direction of propagation of the wave. We see that the electric and magnetic fields \mathbf{E} and \mathbf{B} of a plane wave are directed perpendicular to the direction of propagation of the wave. For this reason, electromagnetic waves are said to be transverse.

The energy density and flux of the plane waves are

$$W = \mathbf{E}^2 \quad \mathbf{S} = W\mathbf{n}$$

3.3.2 Monochromatic wave

A very important special case of electromagnetic waves is a wave in which the field is a simply periodic function of the time. Such a wave is said to be monochromatic. All quantities (potentials, field components) in a monochromatic wave depend on the time through a factor



of the form $\cos(\omega t + a)$. The quantity ω is called the cyclic frequency of the wave (we shall simply call it the frequency). For the monochromatic wave, the wave equation becomes

$$\frac{\partial^2 f}{\partial x^2} + \omega^2 f = 0$$

The vector potential of such a wave is most conveniently written as the real part of a complex expression

$$\mathbf{A} = \text{Re} \{ \mathbf{A}_0 e^{i(\mathbf{k} \cdot \mathbf{r} - \omega t)} \} \quad \mathbf{k} = \omega \mathbf{n}$$

The electric and magnetic field are

$$\mathbf{E} = i\mathbf{k}\mathbf{A} \quad \mathbf{B} = i\mathbf{k} \times \mathbf{A}$$

And we can verify that (ω, \mathbf{k}) transform like a four-vector.

Generally, the electric field can be written as

$$E_y = A \cos(\phi) \quad E_z = B \cos(\phi + \delta) \quad \phi = kx - \omega t \quad -\pi < \delta \leq \pi$$

The end of the vector \mathbf{E} in $y - z$ plane will form an ellipse. The magnitudes of the semiaxes of the polarized ellipse are

$$|\sqrt{A^2 + B^2 + 2AB \sin \delta} \pm \sqrt{A^2 + B^2 - 2AB \sin \delta}|$$

The angle θ between the major axis and y axis satisfies the equation

$$\tan 2\theta = \frac{2AB \cos \delta}{A^2 - B^2}$$

If $-\frac{\pi}{2} < \delta < \frac{\pi}{2}$, major axis is in the second and forth quadrant. If $\delta > \frac{\pi}{2}$ or $\delta < -\frac{\pi}{2}$, major axis is in the first and third quadrant. If $\delta = \pm \frac{\pi}{2}$ and $A > B$, the major axis is y axis. If $\delta = \pm \frac{\pi}{2}$ and $A < B$, the major axis is z axis. If $\delta = \pm \frac{\pi}{2}$ and $A = B$, the ellipse becomes a circle.

If $0 < \delta < \pi$, the rotation is positive in the direction of x axis (left handed). If $-\pi < \delta < 0$, the rotation is negative in the direction of x axis (right handed). If $\delta = 0, \pi$, the ellipse becomes a line.

Any field is expandable in a Fourier integral containing a continuous or discrete distribution of different frequencies. Such an expansion has the form

$$f(t) = \int_{-\infty}^{\infty} f_{\omega} e^{-i\omega t} \frac{d\omega}{2\pi}$$

where the Fourier components are given in terms of the function $f(t)$ by the integrals

$$f_{\omega} = \int_{-\infty}^{\infty} f(t) e^{i\omega t} dt$$

Because $f(t)$ must be real, so

$$f_{-\omega} = f_{\omega}^*$$

The total intensity of the wave is

$$\int_{-\infty}^{\infty} f^2 dt = \int_{-\infty}^{\infty} |f_{\omega}|^2 \frac{d\omega}{2\pi} = 2 \int_0^{\infty} |f_{\omega}|^2 \frac{d\omega}{2\pi}$$



3.3.3 Partially polarized light

Every monochromatic wave is necessarily polarized. However we usually have to deal with waves which are only approximately monochromatic, and which contain frequencies in a small interval $\Delta\omega$. We consider such a wave, and let ω be some average frequency for it. Then its field at a fixed point in space can be written in the form

$$\mathbf{E}_0(t)e^{-i\omega t}$$

where the complex amplitude \mathbf{E}_0 is some slowly varying function of the time. Since \mathbf{E}_0 determines the polarization of the wave, this means that at each point of the wave, its polarization changes with time, such a wave is said to be partially polarized.

The polarization properties of electromagnetic waves are observed experimentally by passing the light to be investigated through various bodies and then observing the intensity of the transmitted light. From the mathematical point of view this means that we draw conclusions concerning the polarization properties of the light from the values of certain quadratic functions of its field. Here of course we are considering the time averages of such functions.

Quadratic functions of the field are made up of terms proportional to the products $E_\alpha E_\beta$, $E_\alpha^* E_\beta^*$ or $E_\alpha^* E_\beta$. Products of the form $E_\alpha E_\beta$ and $E_\alpha^* E_\beta^*$ contain the rapidly oscillating factors $e^{-i2\omega t}$ and will give zero when the time average is taken. Thus we see that the polarization properties of the light are completely characterized by the tensor

$$J_{\alpha\beta} = E_{0\alpha} E_{0\beta}^*$$

The trace of the tensor

$$J \equiv J_{\alpha\alpha} = \mathbf{E}_0 \mathbf{E}_0^*$$

determines the intensity of the wave, as measured by the energy flux density. To eliminate this quantity which is not directly related to the polarization properties, we introduce the tensor

$$\rho_{\alpha\beta} = \frac{J_{\alpha\beta}}{J}$$

we call it the polarization tensor.

Generally, the polarization tensor is expressed as

$$\rho = \frac{1}{2} \begin{bmatrix} 1 + p_3 & p_1 - ip_2 \\ p_1 + ip_2 & 1 - p_3 \end{bmatrix}$$

If we introduce the Pauli matrix, i.e.

$$\sigma_1 = \begin{bmatrix} 0 & 1 \\ 1 & 0 \end{bmatrix} \quad \sigma_2 = \begin{bmatrix} 0 & -i \\ i & 0 \end{bmatrix} \quad \sigma_3 = \begin{bmatrix} 1 & 0 \\ 0 & -1 \end{bmatrix}$$

The ρ can be expressed as

$$\frac{1}{2}(1 - P)I + \frac{1}{2}P(I + \mathbf{n} \cdot \boldsymbol{\sigma})$$

Here,

$$P = \sqrt{p_1^2 + p_2^2 + p_3^2} \quad \mathbf{n} = \left(\frac{p_1}{P}, \frac{p_2}{P}, \frac{p_3}{P}\right)$$



For a totally polarized light with polarization state $|E\rangle = (\cos \frac{\theta}{2} e^{-i\frac{\phi}{2}}, \sin \frac{\theta}{2} e^{i\frac{\phi}{2}})$, the polarization tensor $\rho = |E\rangle\langle E|$. We can verify that

$$P = 1 \quad \mathbf{n} = (\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta)$$

So, for an arbitrary light,

$$\rho = (1 - P)\rho_n + P\rho_p \quad \rho_n = \frac{1}{2}I \quad \rho_p = \frac{1}{2}(I + \mathbf{n} \cdot \boldsymbol{\sigma})$$

Thus, we call P the Polarization degree.

Suppose there is a polarizing filter, which allow the light with polarization state

$$|D\rangle = (\cos \frac{\theta}{2} e^{-i\frac{\phi}{2}}, \sin \frac{\theta}{2} e^{i\frac{\phi}{2}})$$

to pass totally. If a light with polarization tensor ρ pass through the device, the relative intensity will become

$$\langle D|\rho|D\rangle = \frac{1}{2} + \frac{1}{2}\mathbf{p} \cdot \mathbf{m}$$

Here, $\mathbf{p} = (p_1, p_2, p_3)$, $\mathbf{m} = (\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta)$

In optics, the Stokes vectors are defined as The Stokes parameters are defined by

$$\begin{aligned} I &\equiv \langle E_x^2 \rangle + \langle E_y^2 \rangle \\ &= \langle E_a^2 \rangle + \langle E_b^2 \rangle \\ &= \langle E_l^2 \rangle + \langle E_r^2 \rangle \\ Q &\equiv \langle E_x^2 \rangle - \langle E_y^2 \rangle \\ U &\equiv \langle E_a^2 \rangle - \langle E_b^2 \rangle \\ V &\equiv \langle E_l^2 \rangle - \langle E_r^2 \rangle \end{aligned}$$

where the subscripts refer to three different bases of the space of Jones vectors: the standard Cartesian basis \hat{x}, \hat{y} , a Cartesian basis rotated by 45° \hat{a}, \hat{b} , and a circular basis \hat{l}, \hat{r} . The symbols $\langle \cdot \rangle$ represent expectation values. The light can be viewed as a random variable taking values in the space C^2 of Jones vectors (E_1, E_2) . It is easy to verify that

$$Q = Ip_3 \quad U = Ip_2 \quad V = Ip_1$$

