

EAS 2025

# Discovering optimal analytic dark matter halo profiles with deep RL

Wassim Tenachi

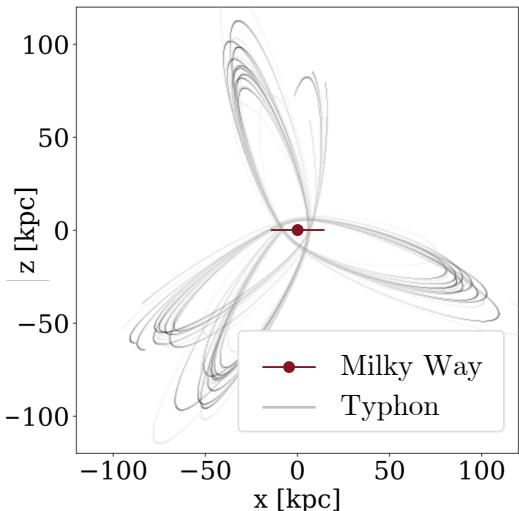
With Yashar Hezaveh, Laurence Perault-Levasseur, Jonathan Freundlich, Jenna Karcheski & Rodrigo Ibata



# Background

# Background

Detection of a dwarf galaxy remnant from the outer halo passing through the solar neighborhood

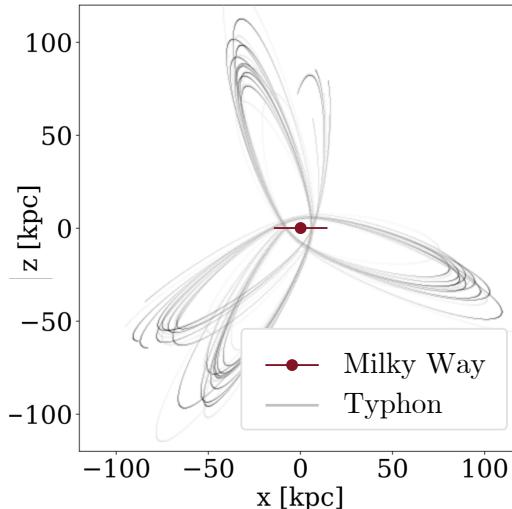


Representation of  
Typhon (abyss) in Greek  
mythology

[Tenachi et al 2022]

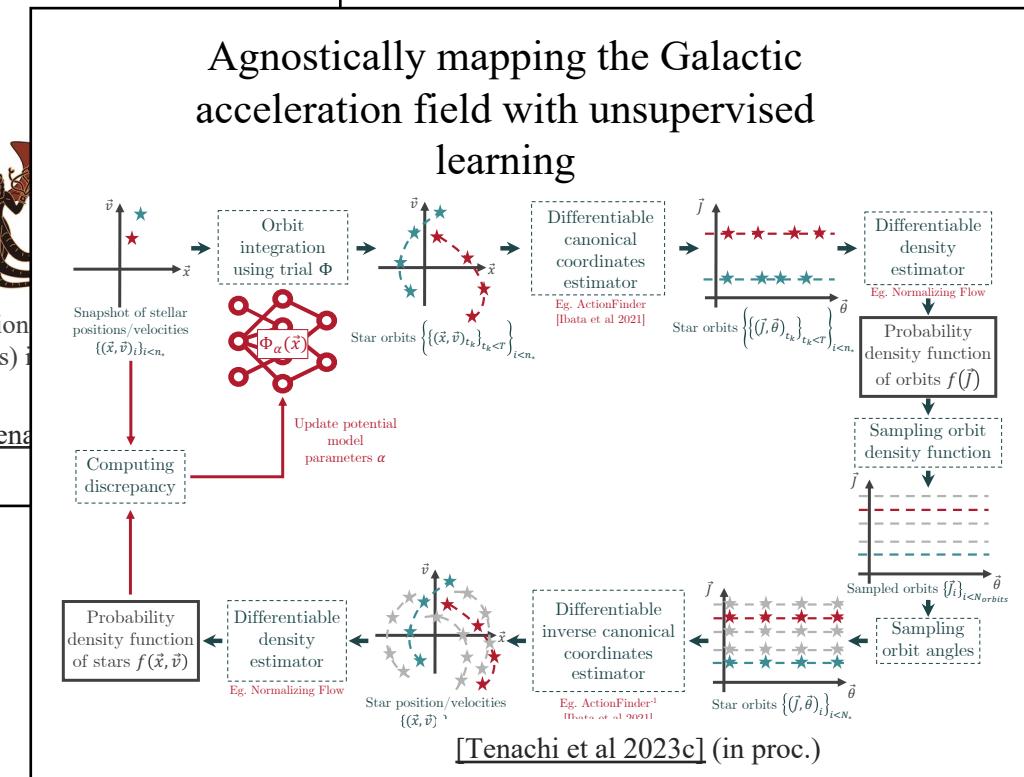
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Detection of a dwarf galaxy remnant from the outer halo passing through the solar neighborhood



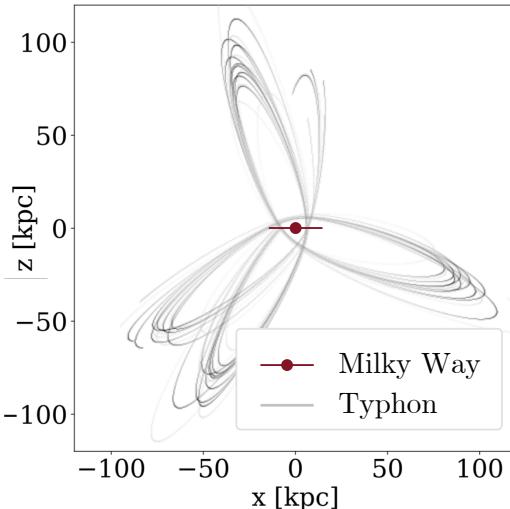
Typhon (abyss) in mythology

[Tenachi et al. 2023c]



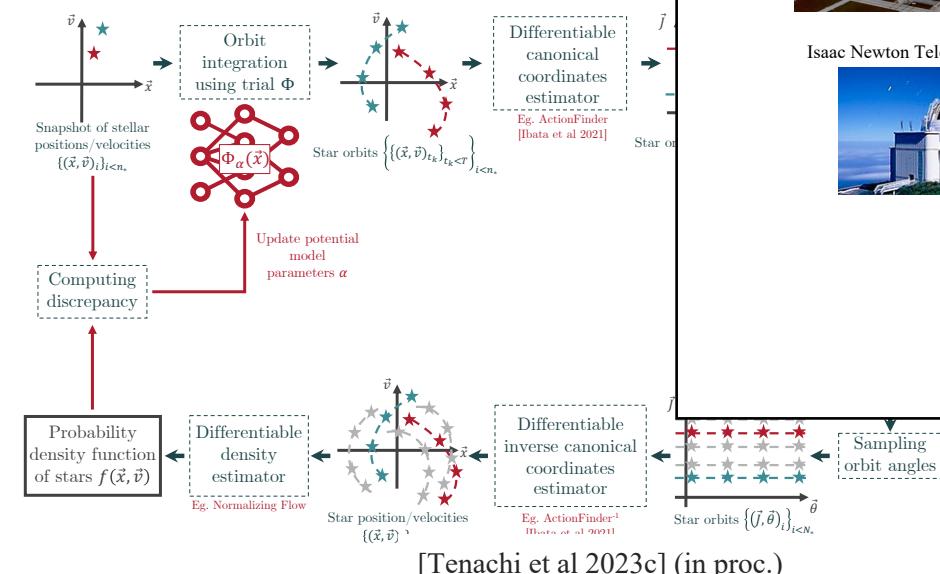
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Detection of a dwarf galaxy remnant from the outer halo passing through the solar neighborhood

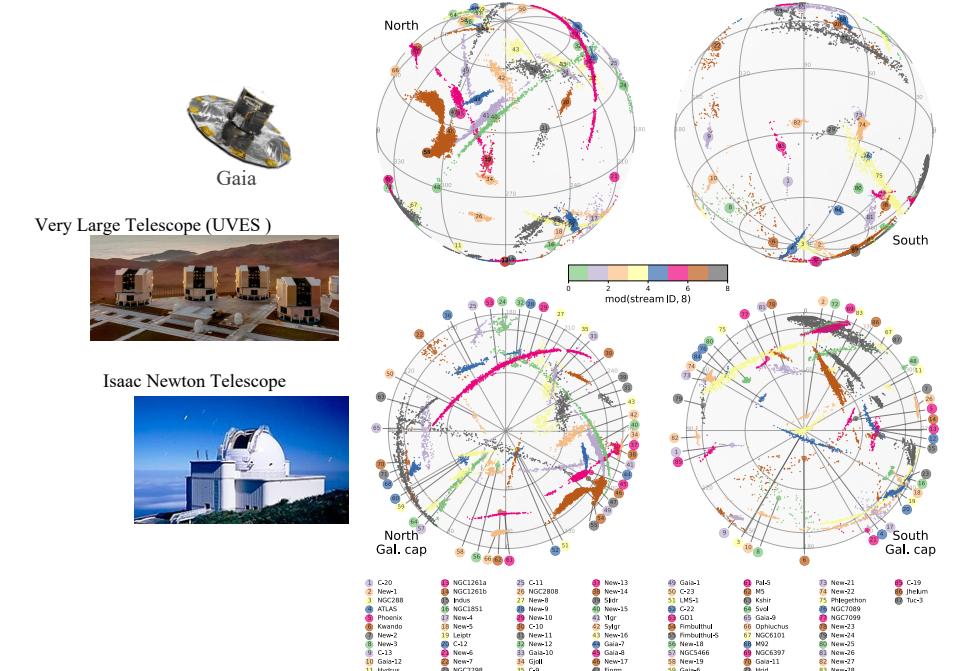


Typhon (abyss) in mythology  
Presentation of Typhon

## Agnostically mapping the Galactic acceleration field with unsupervised learning

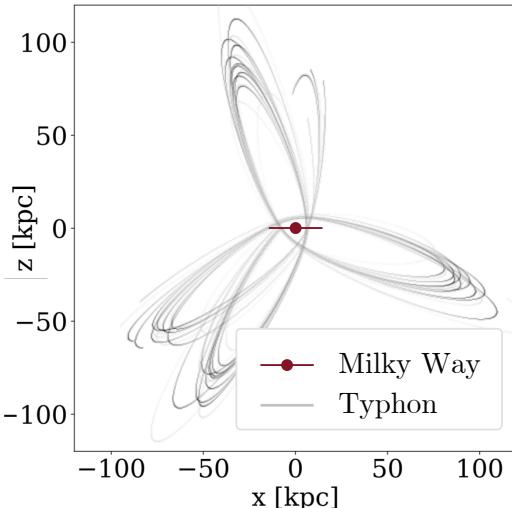


## Constraining the dark matter distribution of our galaxy



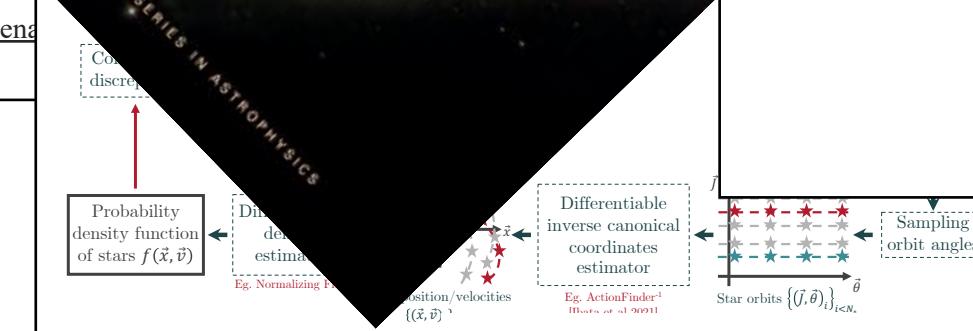
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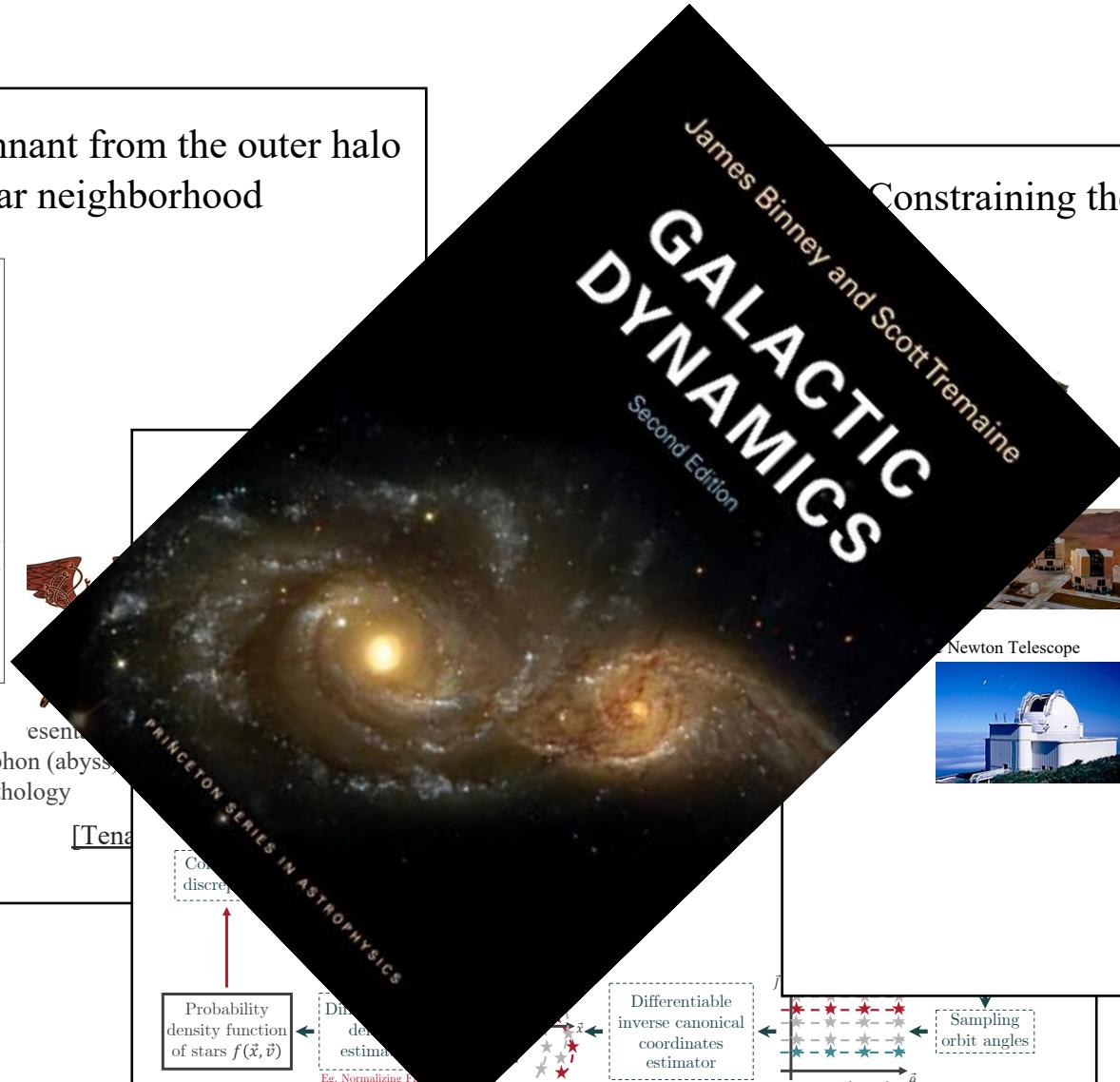


Typhon (ab  
mythology)

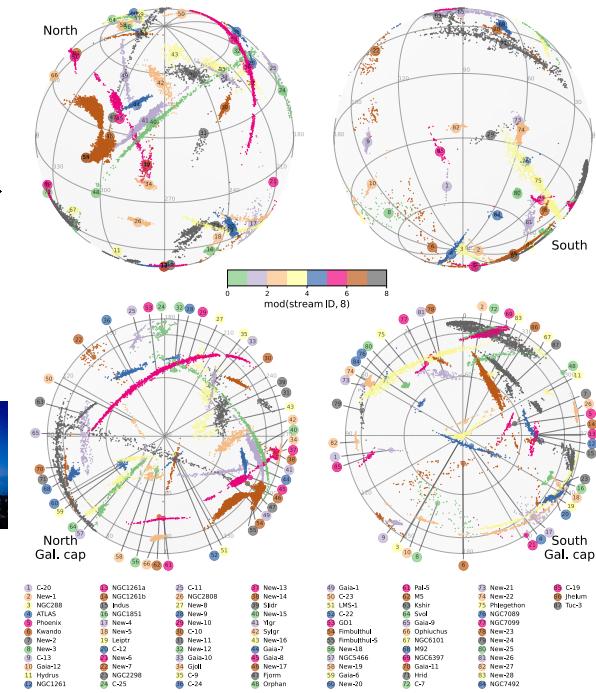
Ten



[Tenachi et al 2023c] (in proc.)

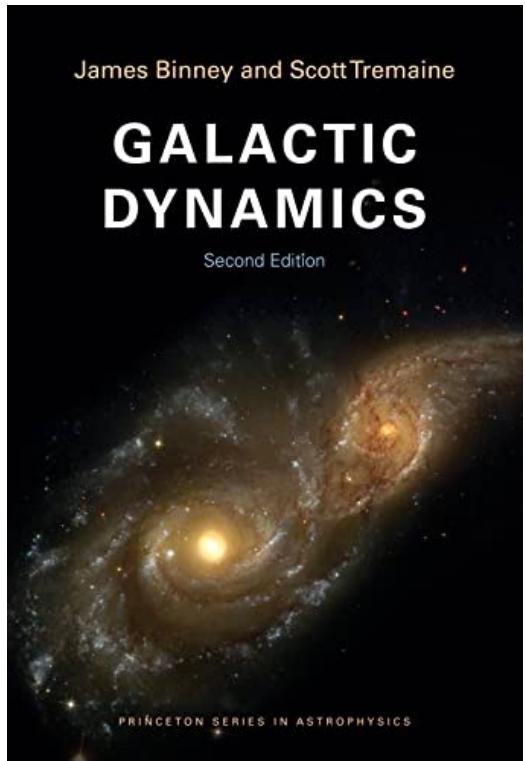


# Constraining the dark matter distribution of our galaxy



[Ibata et al 2023] (incl. WT)

# A zoo of DM profiles

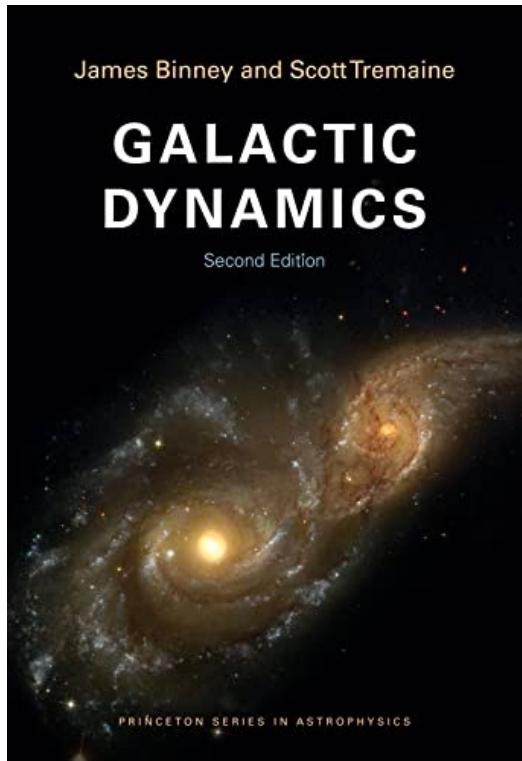


# A zoo of DM profiles

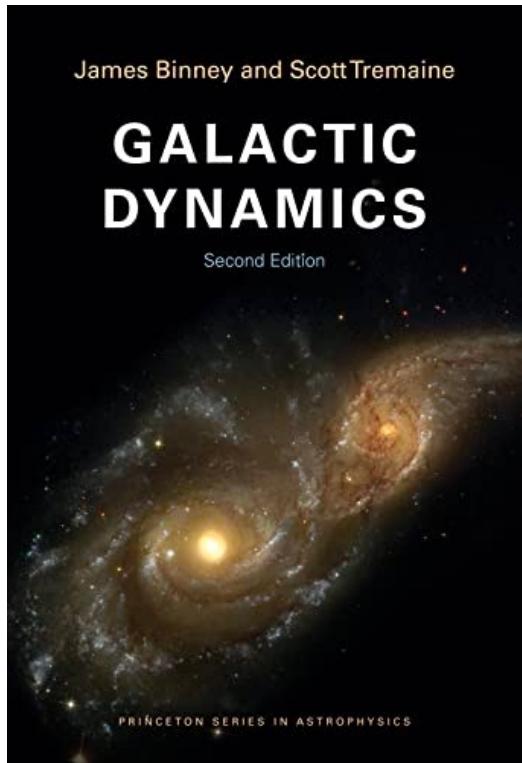
$$\text{NFW} \quad \rho(r) = \frac{\rho_0}{\left(\frac{r}{r_s}\right) \left(1 + \frac{r}{r_s}\right)^2}$$

$$\text{SuperNFW} \quad \rho(r) = \frac{\rho_0}{\left(\frac{r}{r_s}\right) \left(1 + \frac{r}{r_s}\right)^{5/2}}$$

(cusps)



# A zoo of DM profiles



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$$\text{Lucky13} \quad \rho(r) = \frac{\rho_0}{\left(1 + \frac{r}{r_s}\right)^3}$$

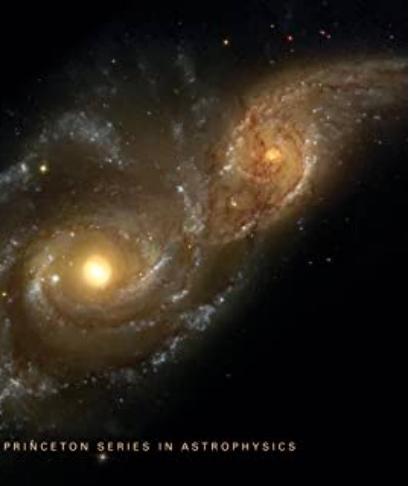
(cores)

# A zoo of DM profiles

James Binney and Scott Tremaine

## GALACTIC DYNAMICS

Second Edition



$$\text{NFW} \quad \rho(r) = \frac{\rho_0}{\left(\frac{r}{r_s}\right) \left(1 + \frac{r}{r_s}\right)^2}$$

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(cusps)

$$\text{Einasto} \quad \rho(r) = \rho_0 \exp\left(-\frac{2}{\alpha} \left[\left(\frac{r}{r_s}\right)^\alpha - 1\right]\right)$$

$$\text{pISO} \quad \rho(r) = \frac{\rho_0}{1 + \left(\frac{r}{r_s}\right)^2}$$

$$\text{gNFW} \quad \rho(r) = \frac{\rho_0}{\left(\frac{r}{r_s}\right)^\alpha \left(1 + \left(\frac{r}{r_s}\right)\right)^{3-\alpha}}$$

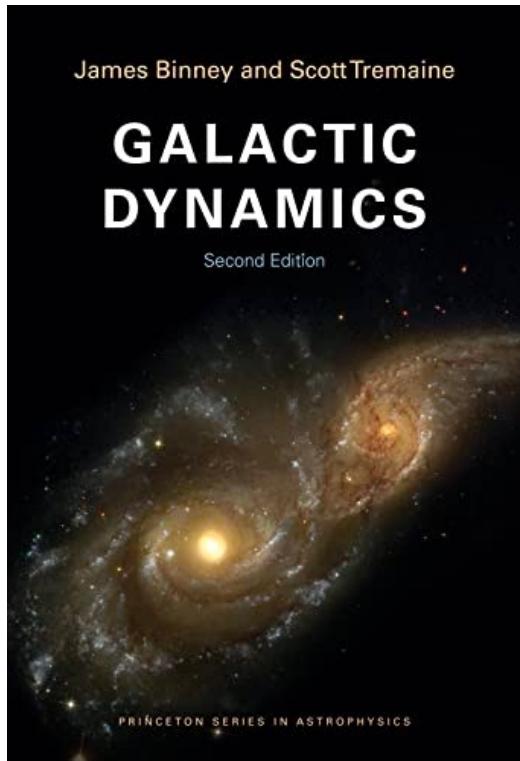
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(variable inner slope)

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# A zoo of DM profiles



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(cores)

→ Let's explore the space of analytical equations for DM profiles

# Symbolic regression (SR)

$(x, y)$  data

$x_1$	$x_2$	$y$
0.75582	0.25850	0.02674
0.36786	0.42401	0.06278
0.69507	0.38057	0.74014
0.96493	0.33398	0.81558
0.07139	0.16604	0.07716
0.86413	0.41952	0.87872
0.18012	0.40581	0.63637

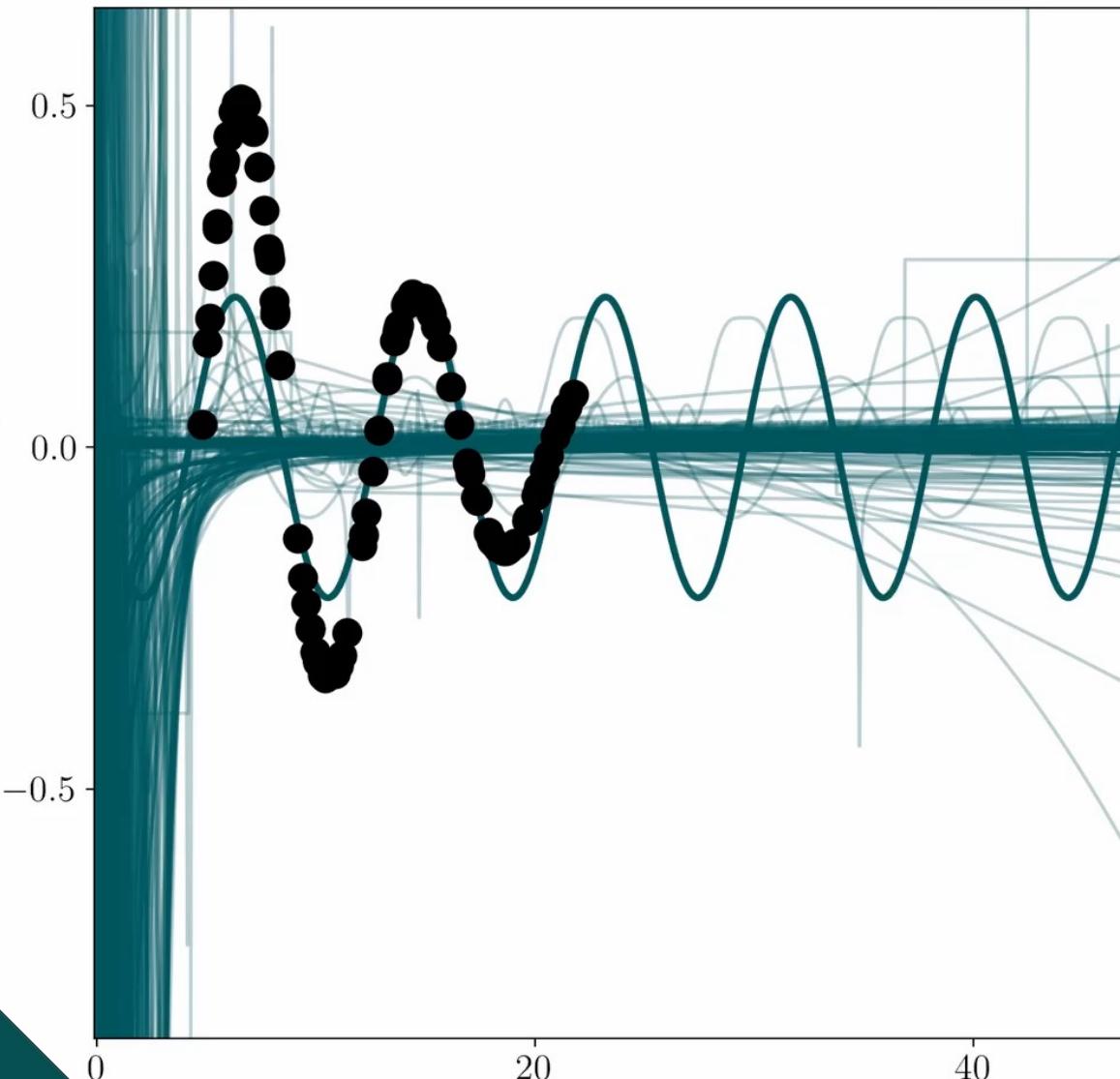


$f: \mathbb{R}^n \rightarrow \mathbb{R}$  such that  
 $y = f(\mathbf{x})$

# Symbolic regression (SR)

# Φ-SO

Physical Symbolic Optimization.



Best fit:

$$x(t) = \sin(\alpha t) \cos(\phi^2)$$

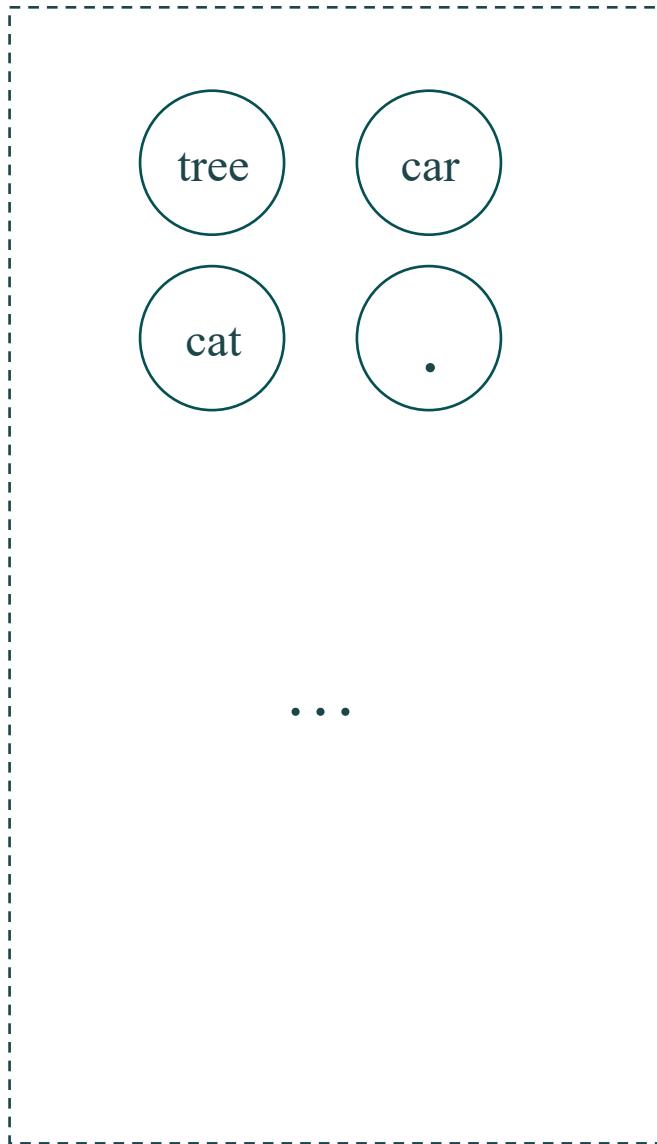
Trying:

$$x(t) = \sin(\alpha t) \cos(\phi^2)$$

<https://youtu.be/wubzMkoTUY>

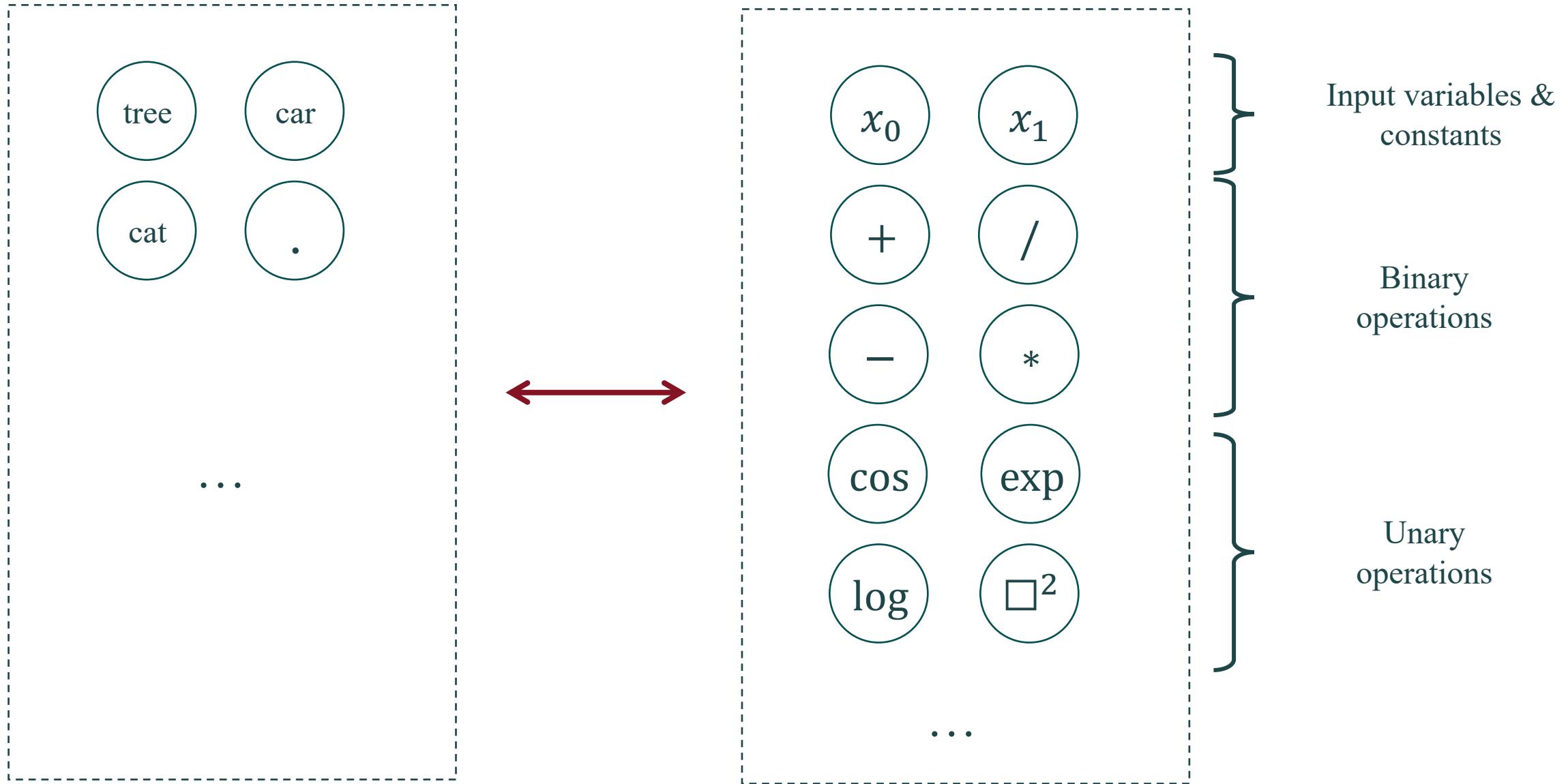
# Language processing for symbolic mathematics

$\Phi$ -SO : Embedding

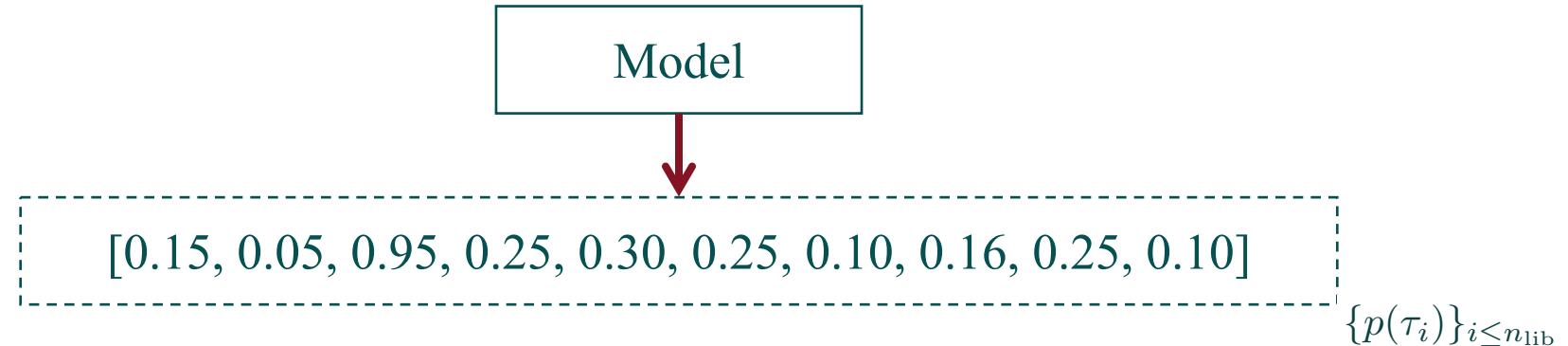


# Language processing for symbolic mathematics

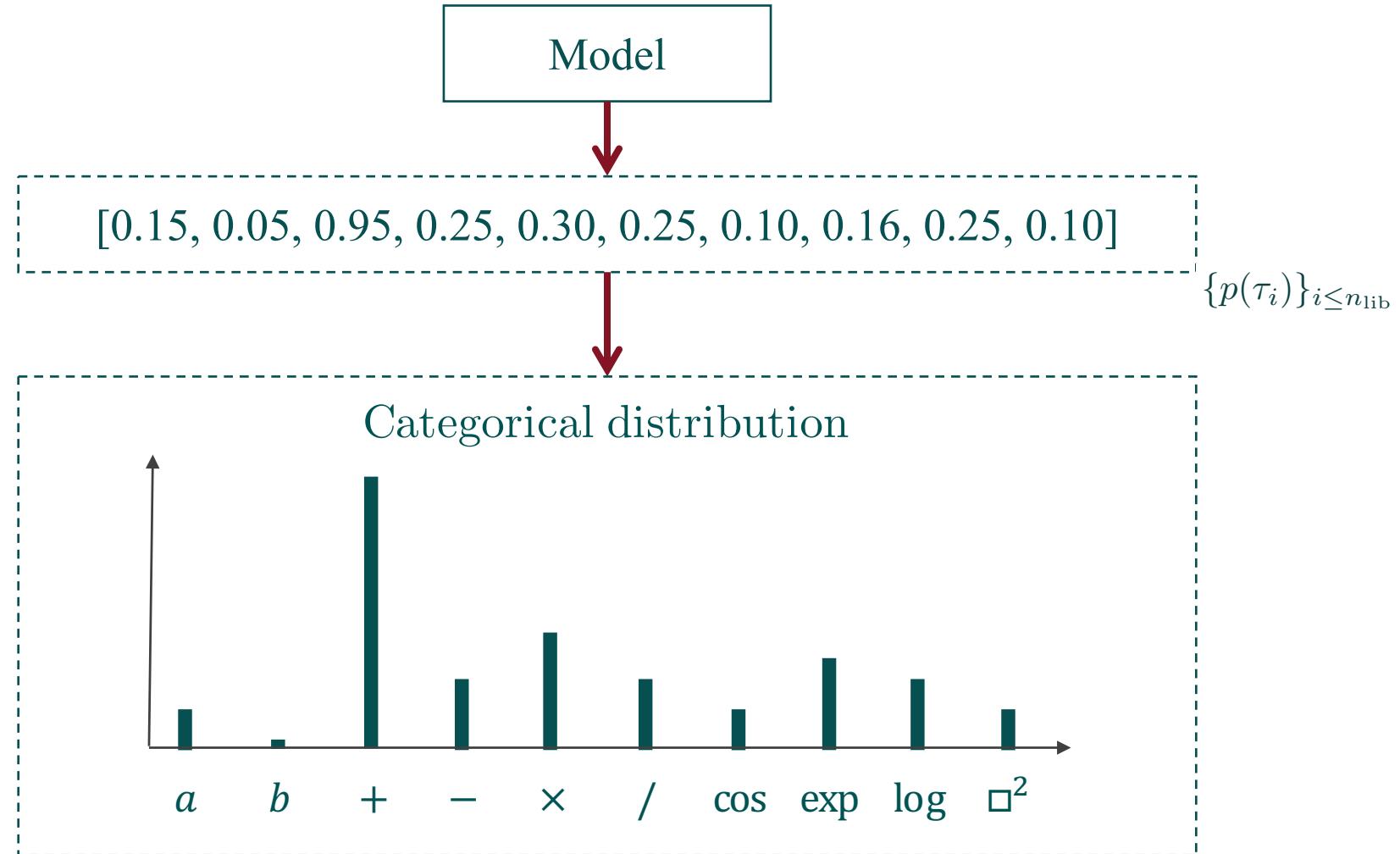
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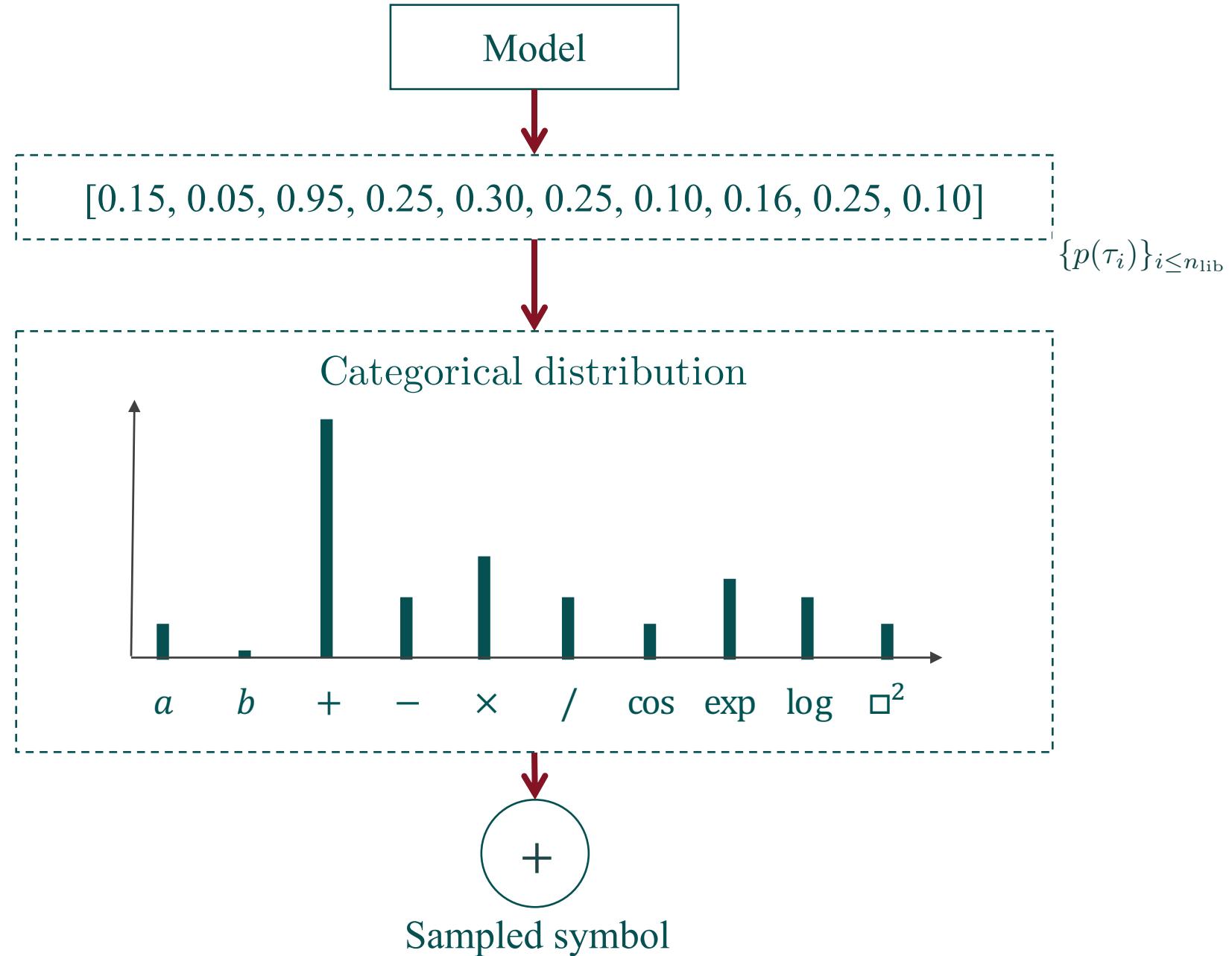
# Embedding (1) : how to go from numbers to symbols ?



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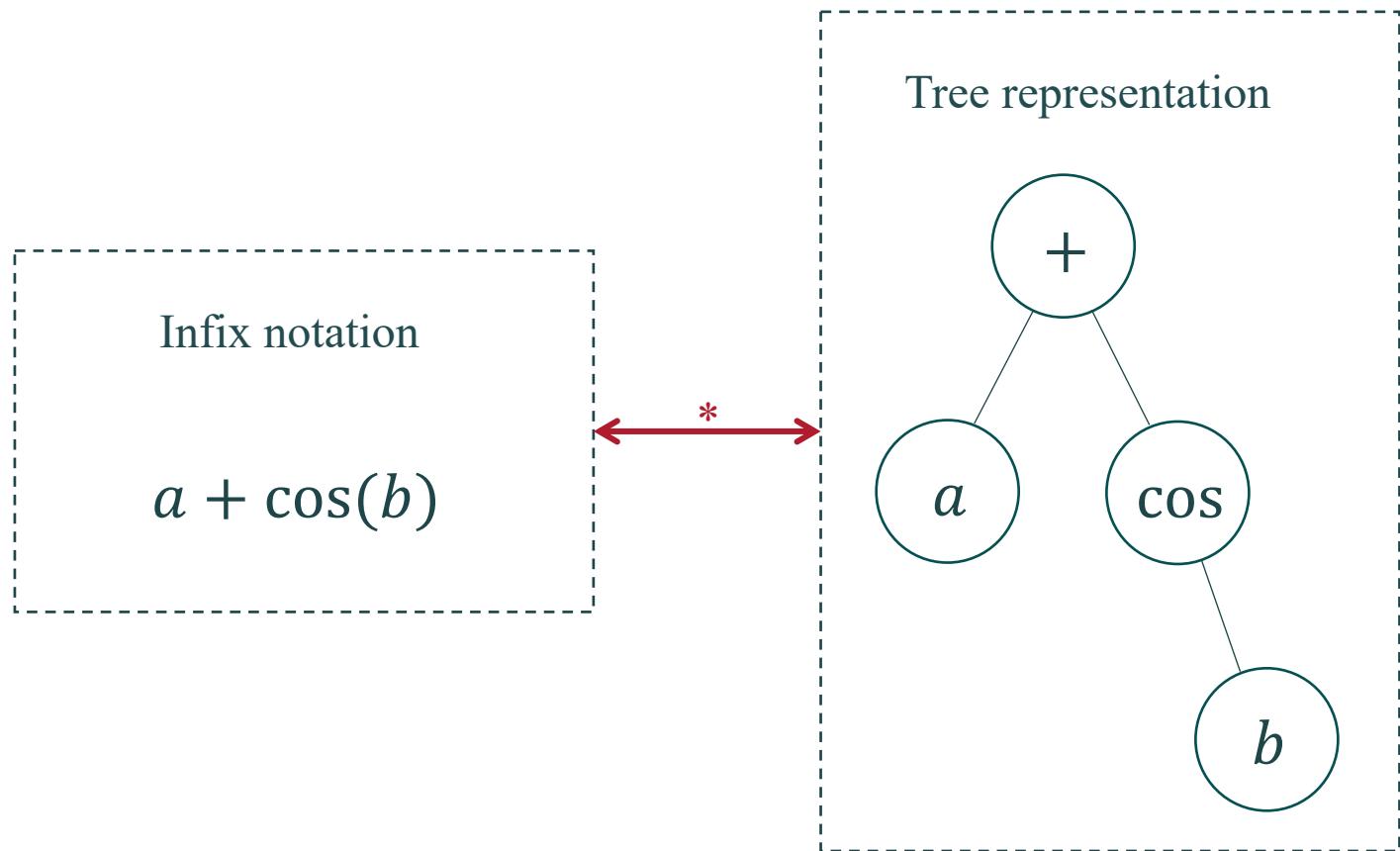
## Embedding (2) : how to go from vector of symbols to expressions ?

Φ-SO : Embedding

Infix notation

$$a + \cos(b)$$

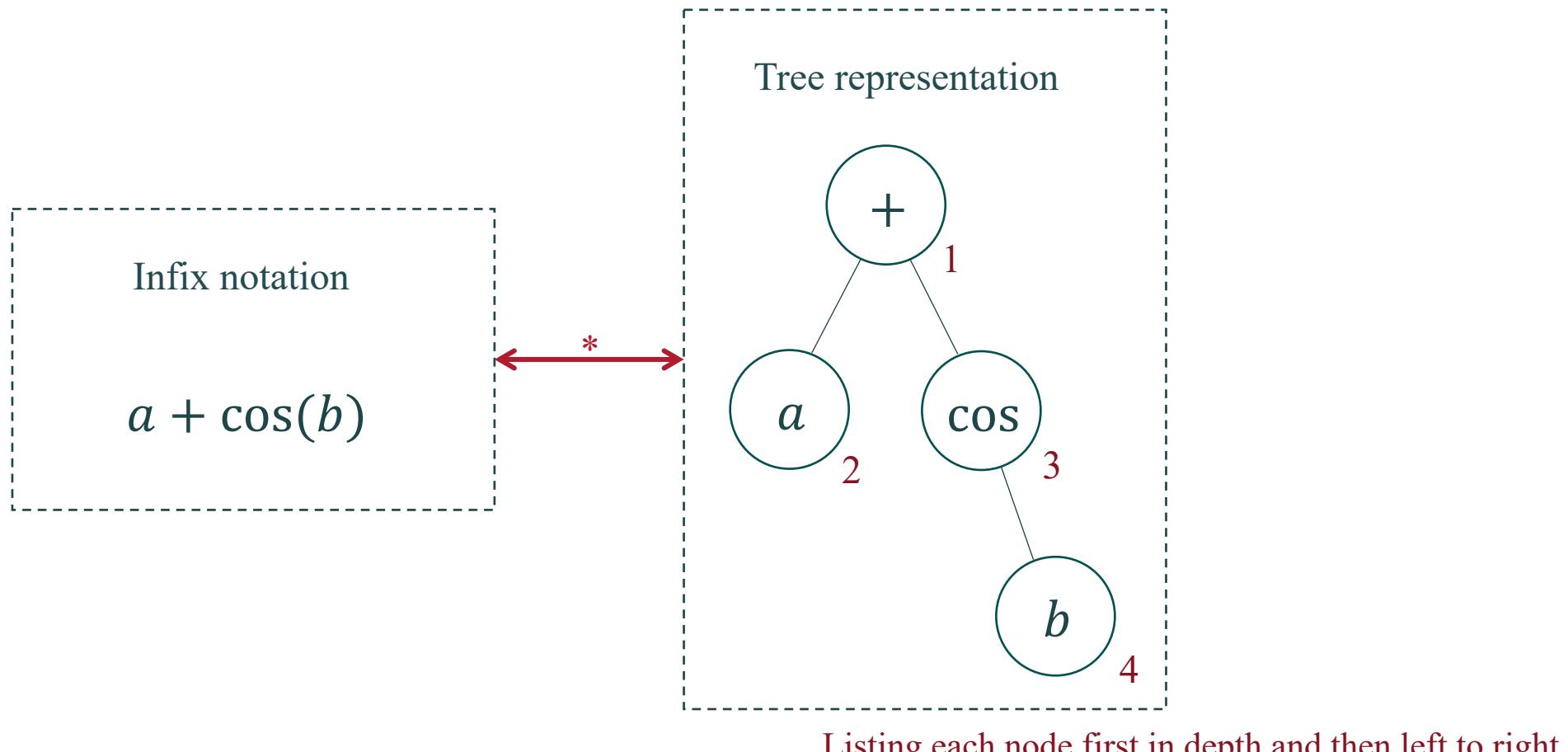
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\* 1:1 equivalence

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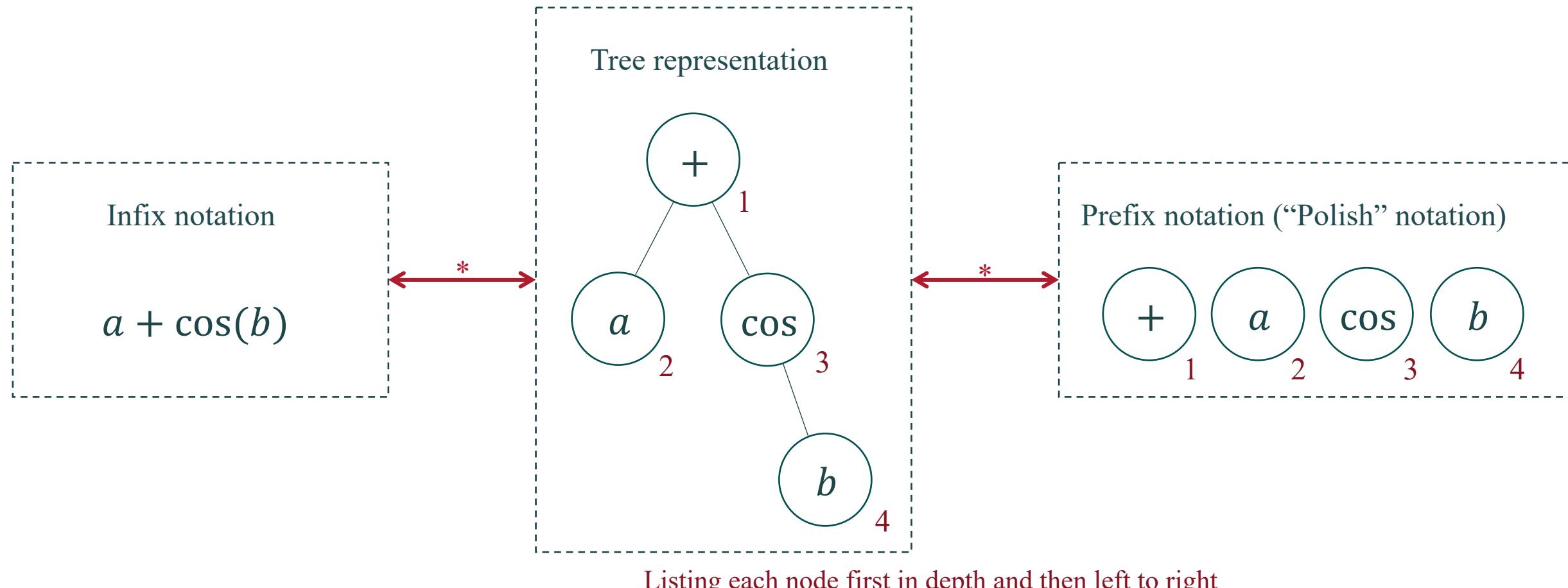
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## Embedding (2) : how to go from vector of symbols to expressions ?

$\Phi$ -SO : Embedding



\* 1:1 equivalence

# Reinforcement learning & risk seeking policy

Φ-SO : learning



## Categorical distributions

$$v = v_0 + x/t$$

$$v = v_0$$

$$v = v_0 - x/t$$

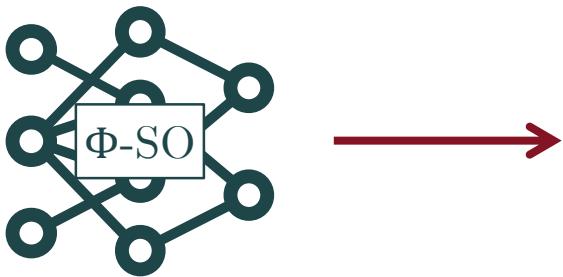
$$v = x/t$$

⋮

$$v = \frac{v_0^2}{x/t}$$

# Reinforcement learning & risk seeking policy

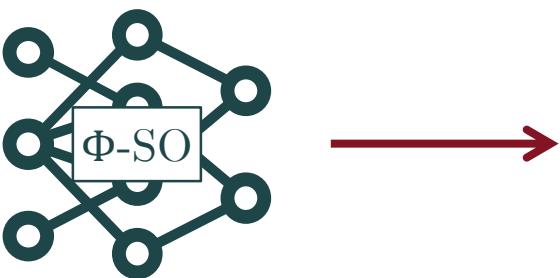
Φ-SO : learning



Categorical distributions	
Candidates vs data → reward	
$v = v_0 + x/t$	→ r = 0.99
$v = v_0$	→ r = 0.93
$v = v_0 - x/t$	→ r = 0.73
$v = x/t$	→ r = 0.28
⋮	
$v = \frac{v_0^2}{x/t}$	→ r = 0.64

# Reinforcement learning & risk seeking policy

Φ-SO : learning



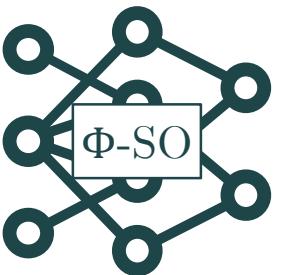
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Reinforce on the 5% best candidates

Not punished/rewarded on 95% of candidates  
→ “risk seeking”

# Reinforcement learning & risk seeking policy

→ Black box reward function



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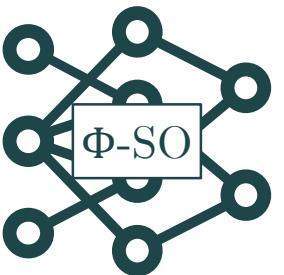
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Reward → no auto-differentiation  
(unlike most machine learning methods used in physics)

→ We can apply any physical constraints we want even if it is not differentiable

# Reinforcement learning & risk seeking policy

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## Categorical distributions

Candidates vs data → reward

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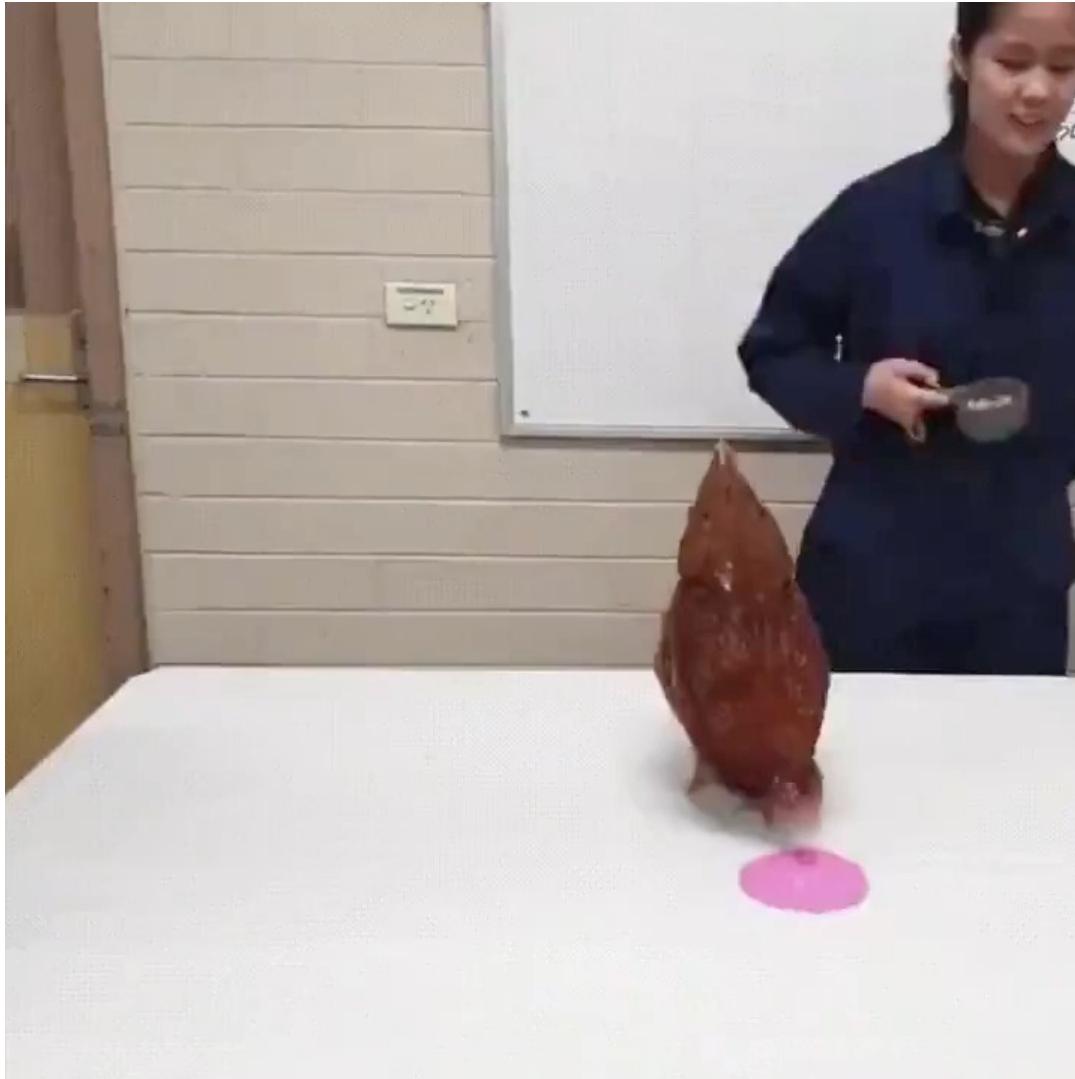
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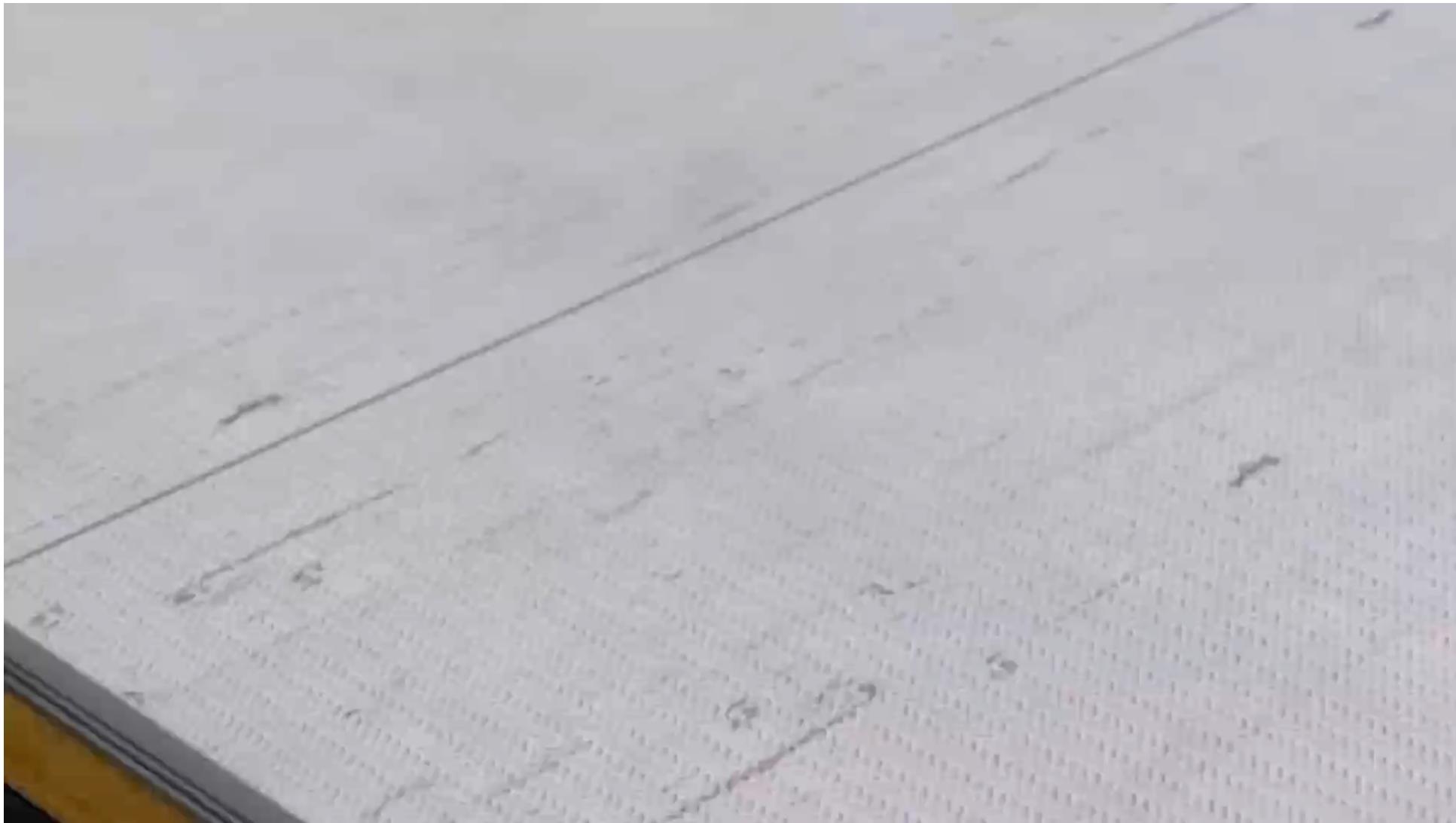
- Complexity (Occam's razor)
- Symmetries
- Constraints on derivatives/primitives
- Symbolic computing using Sympy/Mathematica
- Fitness in ODEs, limits values
- Behavior in N-body simulation
- ...

# Reinforcement learning in a nutshell (1)



<https://youtu.be/spfpBrBjntg>

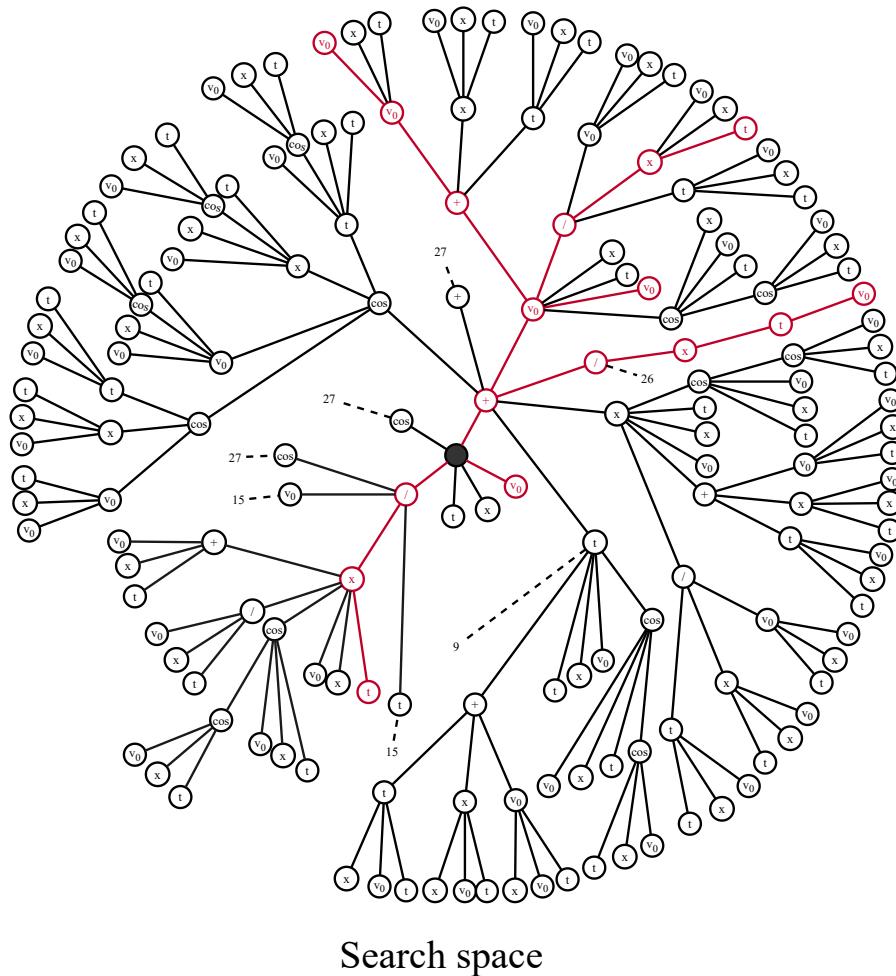
# Reinforcement learning in a nutshell (2)



<https://youtu.be/igZ6IPQimjQ>

# Search space reduction using physical units constraints

268 expressions

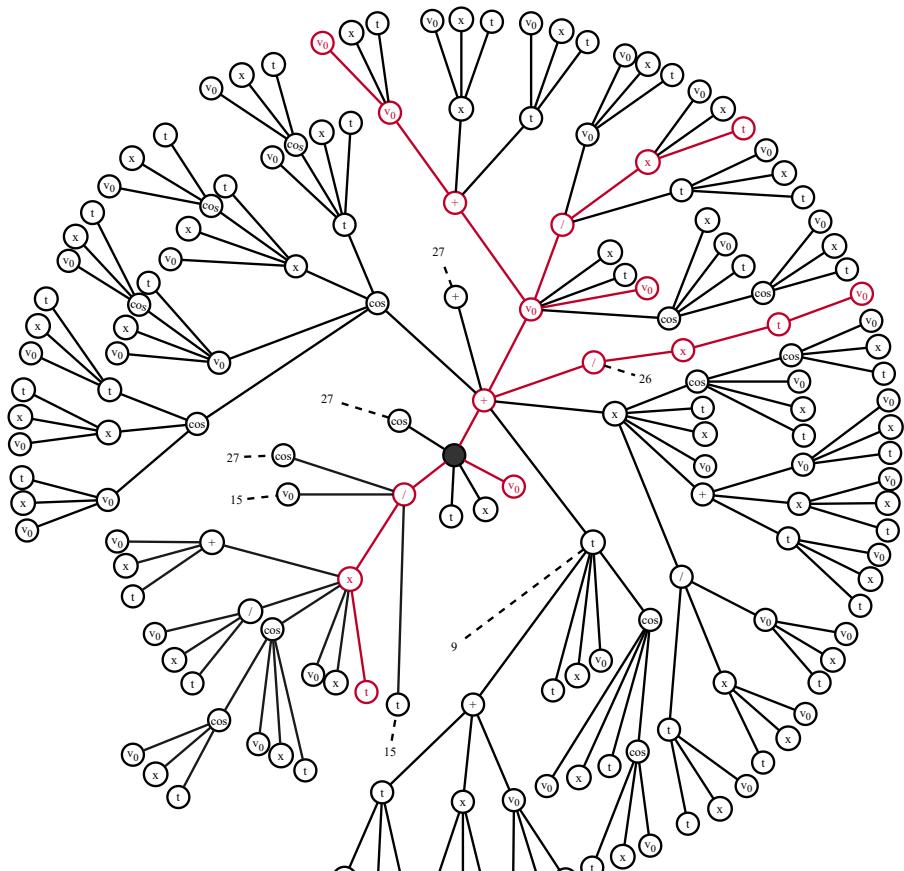


Prefix notation paths for expressing a velocity  $v$  using a library of symbols  $\{+, /, \cos, v_0, x, t\}$  where  $v_0$  is a velocity,  $x$  is a length, and  $t$  is a time (length < 5 for readability).

[Tenachi et al 2023a,b]

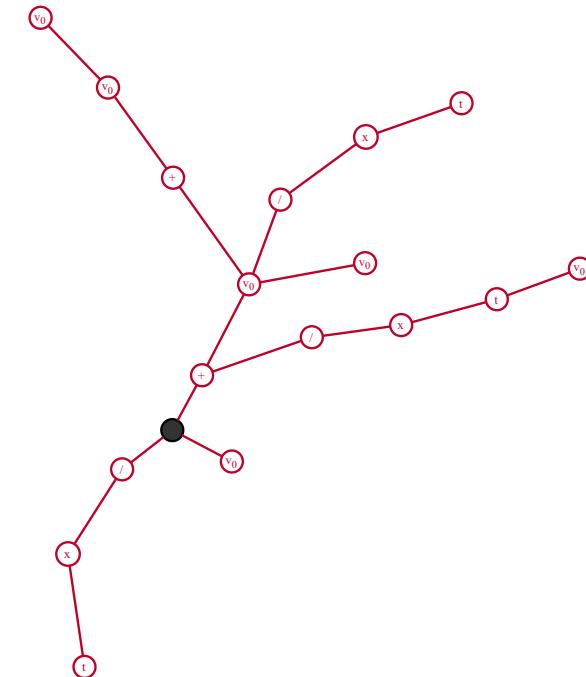
# Search space reduction using physical units constraints

268 expressions



Search space

6 expressions



Search space with our *in situ* physical units prior

Prefix notation paths for expressing a velocity  $v$  using a library of symbols  $\{+, /, \cos, v_0, x, t\}$  where  $v_0$  is a velocity,  $x$  is a length, and  $t$  is a time (length < 5 for readability).

[Tenachi et al 2023a,b]

# Feynman benchmark

120 equations from the Feynman Lectures on Physics (+ other physics textbooks)

Feynman eq.	Equation
II.2.42	$P = \frac{\kappa(T_2 - T_1)A}{d}$
III.3.24	$F_E = \frac{1}{4\pi r^2}$
III.4.23	$V_e = \frac{1}{4\pi r^2}$
III.6.11	$V_e = \frac{1}{4\pi r^2} \frac{P_d}{c}$
III.6.15a	$E_f = \frac{3}{4\pi r^2} \frac{P_d}{c}$
III.6.15b	$E_f = \frac{3}{4\pi r^2} \frac{P_d}{c}$
III.8.7	$E = \frac{3}{5} \frac{q^2}{4\pi r^2}$
III.8.31	$E_{den} = \frac{q^2}{2}$
III.10.9	$E_f = \frac{q E_{den}}{V}$
III.11.3	$x = \frac{q}{m(\omega_0^2 - \omega^2)}$
III.11.17	$n = n_0(1 + \frac{q^2}{m(\omega_0^2 - \omega^2)})$
III.11.20	$P_e = \frac{n_0 q^2 P_d}{3 R_k T}$
III.11.27	$P_e = \frac{1 - n_0 / \omega}{1 - \omega / \omega_0}$
III.11.28	$\theta = 1 + \frac{1 - \omega / \omega_0}{1 - \omega / \omega_0}$
III.13.17	$B = \frac{1}{4\pi c r^2} \frac{p_m \cdot r}{\rho}$
III.13.23	$\rho = \frac{\sqrt{1-u^2/c^2}}{\sqrt{1-v^2/c^2}}$
III.13.34	$j = \frac{\sqrt{1-v^2/c^2}}{\sqrt{1-u^2/c^2}}$
III.15.4	$E = -\mu_M B$
III.15.5	$E = -p_d E_f$
III.21.32	$V_e = \frac{1}{4\pi r^2} \frac{1}{V}$
III.24.17	$k = \sqrt{\frac{u^2}{c^2} - \frac{v^2}{c^2}}$
III.27.16	$F_E = ecE_f^2$
III.27.18	$E_{den} = eE_f^2$
III.34.2a	$I = \frac{qv}{2\pi r}$
III.34.2	$\mu_M = \frac{q}{2\pi r}$
III.34.11	$\omega = \frac{q_A B}{2m}$
III.34.29a	$\mu_M = \frac{q h}{4\pi m}$
III.34.29b	$E = \frac{q \mu_M}{h} B$
III.35.18	$n = \frac{\exp(-\mu_m E)}{1 - \exp(-\mu_m E)}$
III.35.21	$M = n_\rho \mu_M$
III.36.38	$f = \frac{\mu_M B}{k_1 T} + \frac{h}{k_1 T}$
III.37.1	$E = \mu_M (1 + \frac{h}{k_1 T})$
III.38.3	$F = \frac{Y_A g}{c}$
III.38.14	$\mu_S = \frac{2(Y_A g)}{c^2}$
III.43.2	$n = \frac{e^{\frac{q}{c} E}}{e^{\frac{q}{c} E} - 1}$
III.43.33	$E = \frac{h\omega}{e^{\frac{q}{c} E} - 1}$
III.7.38	$\omega = \frac{2\mu_M B}{h} - \frac{h}{k_1 T}$
III.8.54	$p_T = \sin(\frac{E_T}{h})$
III.9.52	$p_T = \frac{p_d E_f}{h} t$
III.10.19	$E = \mu_M \sqrt{B}$
III.12.43	$L = nh$
III.13.18	$v = \frac{2Ed^2 k}{h}$
III.14.14	$I = I_0(e^{-\frac{h\nu}{k_b T}})$
III.15.12	$E = 2U(1 - \frac{V}{n k_b T})$
III.15.14	$m = \frac{c^2}{2E d^2}$
III.15.27	$k = \frac{nd}{\gamma - 1}$
III.17.37	$f = \beta(1 + \alpha)$
III.19.51	$E = \frac{mc^2}{2(4\pi c)^2}$
III.20	$j = \frac{-\rho_{co} q A_c}{m c}$
III.6.20a	$f = e^{-\theta^2/2} / \sqrt{2\pi}$
III.6.20	$f = e^{-\theta^2/2} \sqrt{2\pi \sigma^2}$
III.6.20b	$f = e^{-\frac{(\theta - \theta_0)^2}{2\sigma^2}} / \sqrt{2\pi \sigma^2}$
III.8.14	$d = \sqrt{(x_2 - x_1)^2 + (y_2 - y_1)^2}$
III.9.18	$F = \frac{G m_1 m_2}{r^2}$
III.10.7	$m = \sqrt{1 - \frac{v^2}{c^2}}$
III.11.19	$A = x_1 y_1 + x_2 y_2 + x_3 y_3$
III.12.1	$F = \mu M v$
III.12.2	$F = \frac{q v q_2}{4\pi r^2}$
III.12.4	$F = q_2 E_f$
III.12.11	$F = q(E_f + B v \sin \theta)$
III.13.4	$K = \frac{1}{2} m(u^2 + v^2 + w^2)$
III.13.12	$U = G m_1 m_2 (\frac{1}{r_2} - \frac{1}{r_1})$
III.14.3	$U = mgz$
III.14.4	$U = \frac{k_e p_m z^2}{2}$
III.15.3x	$x_1 = \frac{\sqrt{1-u^2/c^2}}{\sqrt{1-v^2/c^2}}$
III.15.3t	$t_1 = \frac{\sqrt{1-u^2/c^2}}{\sqrt{1-v^2/c^2}}$
III.15.10	$p = \frac{\sqrt{1-u^2/c^2}}{\sqrt{1-v^2/c^2}}$
III.16.6	Source
III.18.4	Equation
III.18.12	Rutherford Scattering
III.18.16	Friedman Equation
III.24.6	$E = \frac{1}{2} m(\omega^2 r^2)$
III.25.13	Compton Scattering
III.26.2	Radiated gravitational wave power
III.27.6	Relativistic aberration
III.29.4	N-slit diffraction
III.29.16	Goldstein 3.16
III.30.3	Goldstein 3.55
III.30.5	$\theta = \arcsin(\frac{v}{c})$ (ellipse)
III.32.5	$P = \frac{q^2 c^2}{8\pi r^2}$
III.32.17	Goldstein 3.99
III.34.8	Goldstein 8.56
III.34.10	Goldstein 12.80
III.34.14	Jackson 2.11
III.34.27	Jackson 3.45
III.37.4	Jackson 4.60
III.38.12	Jackson 11.38 (Doppler)
III.39.10	Weinberg 15.2.1
III.39.11	Weinberg 15.2.2
III.39.22	Weinberg 15.2.2
III.40.1	Schwarz 13.132 (Klein-Nishina)
III.41.16	$L_{rad} = \frac{\pi^2 c^2 k_B T}{2} \frac{h\nu}{c}$
III.43.16	$v = \frac{\mu_{drift} q_e}{2} (e \frac{k_B T}{c})$
III.43.31	$D = \mu_e k_B T$
III.43.43	$\kappa = \frac{1}{\gamma - 1} \frac{k_B}{A}$
III.44.4	$E = n k_B T \ln(\frac{V_2}{V_1})$
III.47.23	$c = \sqrt{\frac{2pc}{\rho}}$
III.48.20	$E = \frac{mc^2}{\sqrt{1-v^2/c^2}}$
III.50.26	$x = x_1 [\cos(\omega t) + \alpha \cos(\omega t)^2]$

## Against 17 other algorithms

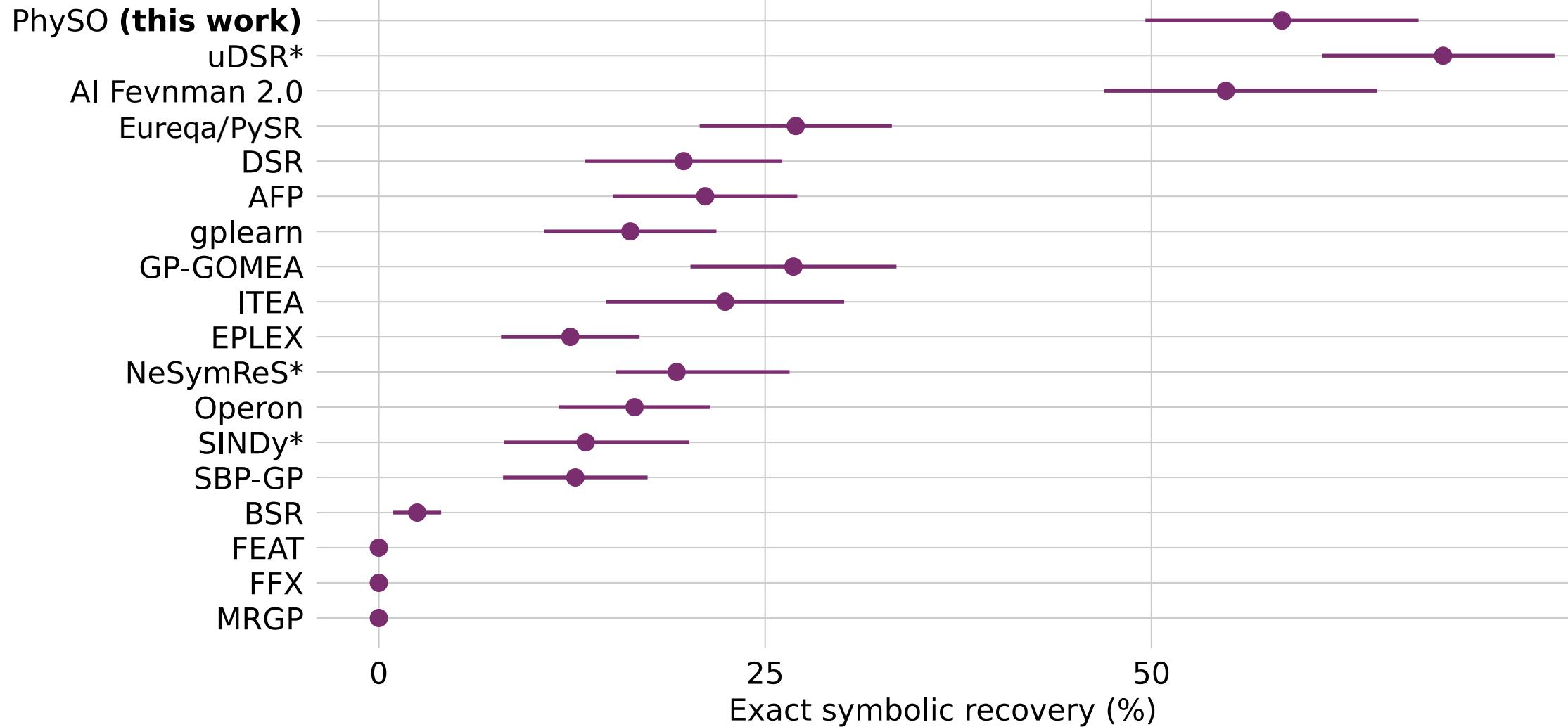
Method	Technique(s)	Description	Reference
PhySO	RL, DA	Physical Symbolic Optimization	This work
uDSR	RL, GP, Simp., Sup.	A Unified Framework for Deep Symbolic Regression	Landajuela et al. (2022)
AIFeynman 2.0	Simp., DA	Symbolic regression exploiting graph modularity	Udrescu et al. (2020)
AFP_FE	GP	AFP with co-evolved fitness estimates, Eureqa-esque	Schmidt & Lipson (2009)
DSR	RL	Deep Symbolic Regression	Petersen et al. (2021a)
AFP	GP	Age-fitness Pareto Optimization	Schmidt & Lipson (2011)
gplearn	GP	Koza-style symbolic regression in Python	Stephens (2015)
GP-GOMEA	GP	GP-Optimal Mixing Evolutionary Algorithm	Virgolin et al. (2021)
ITEA	GP	Interaction-Transformation EA	de Franca & Aldeia (2021)
EPLEX	GP	$\epsilon$ -lexicase selection	La Cava et al. (2019)
NeSymReS	Sup.	Neural Symbolic Regression that Scales	Biggio et al. (2021)
Operon	GP	SR with Non-linear least squares	Kommenda et al. (2020)
SINDy	NeuroSym	Sparse identification of non-linear dynamics	Brunton et al. (2016)
SBP-GP	GP	Semantic Back-propagation Genetic Programming	Virgolin et al. (2019)
BSR	MCMC	Bayesian Symbolic Regression	Jin et al. (2019)
FEAT	GP	Feature Engineering Automation Tool	Cava et al. (2019)
FFX	Rand.	Fast function extraction	McConaghy (2011)
MRGP	GP	Multiple Regression Genetic Programming	Arnaldo et al. (2014)

**Table 3.** Summary of baseline symbolic regression methods along with the the underlying techniques they rely on: reinforcement learning (RL), genetic programming (GP), problem simplification schemes (Simp.), end-to-end supervised learning (Sup.), dimensional analysis (DA), neuro-symbolic / auto-differentiation based sparse fitting techniques (NeuroSym), Markov chain Monte Carlo (MCMC) and random search (Rand.).

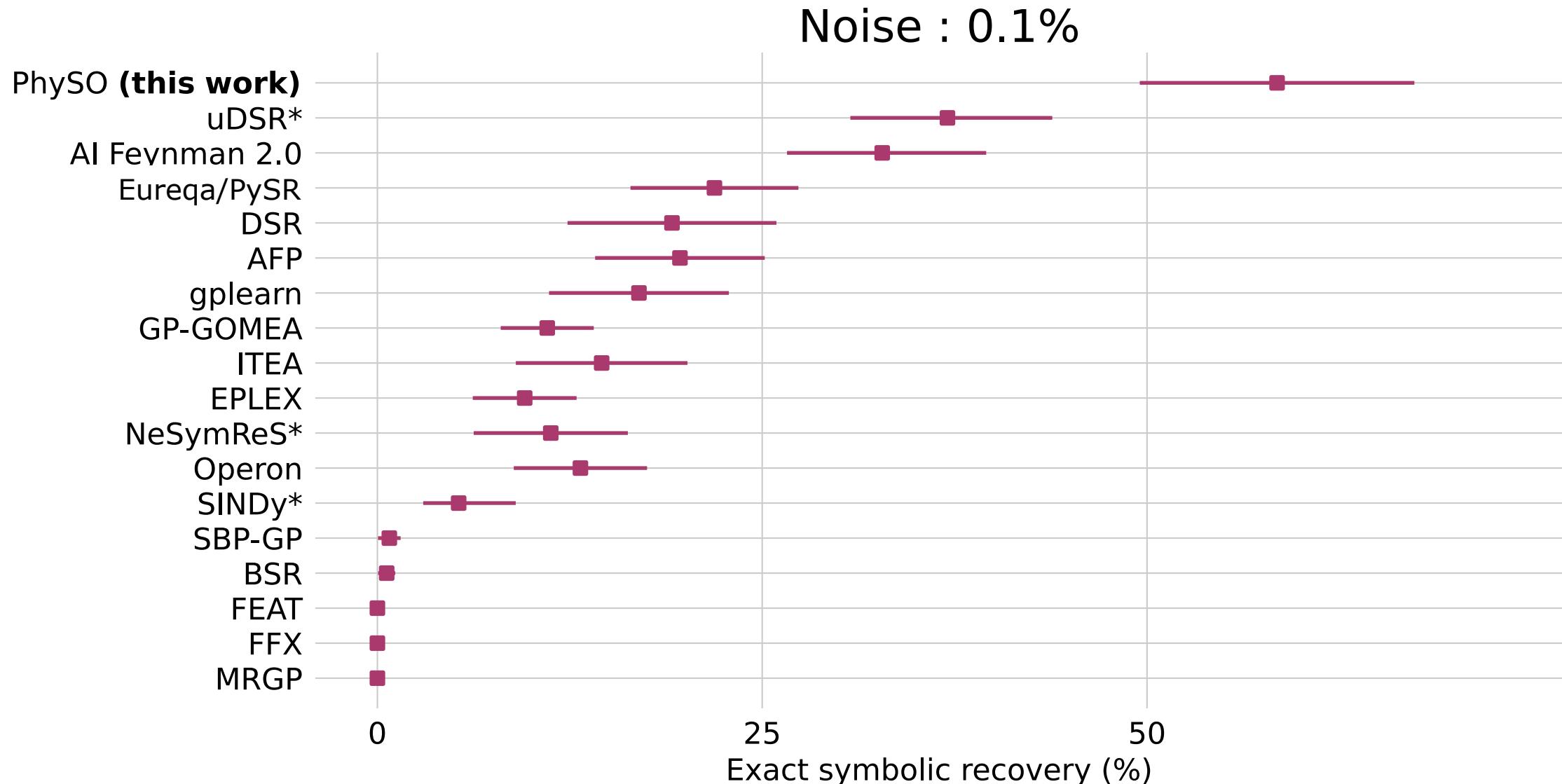
Introduced in [\[Udrescu & Tegmark 2020\]](#)  
Formalized benchmark in SRBench [\[LaCava et al 2021\]](#)

## Feynman benchmark

Noise : 0%



## Feynman benchmark



# Feynman benchmark

Noise : 1%

**PhySO (this work)**

uDSR\*

AI Feynman 2.0

Eureqa/PySR

DSR

AFP

gplearn

GP-GOMEA

ITEA

EPLEX

NeSymReS\*

Operon

SINDy\*

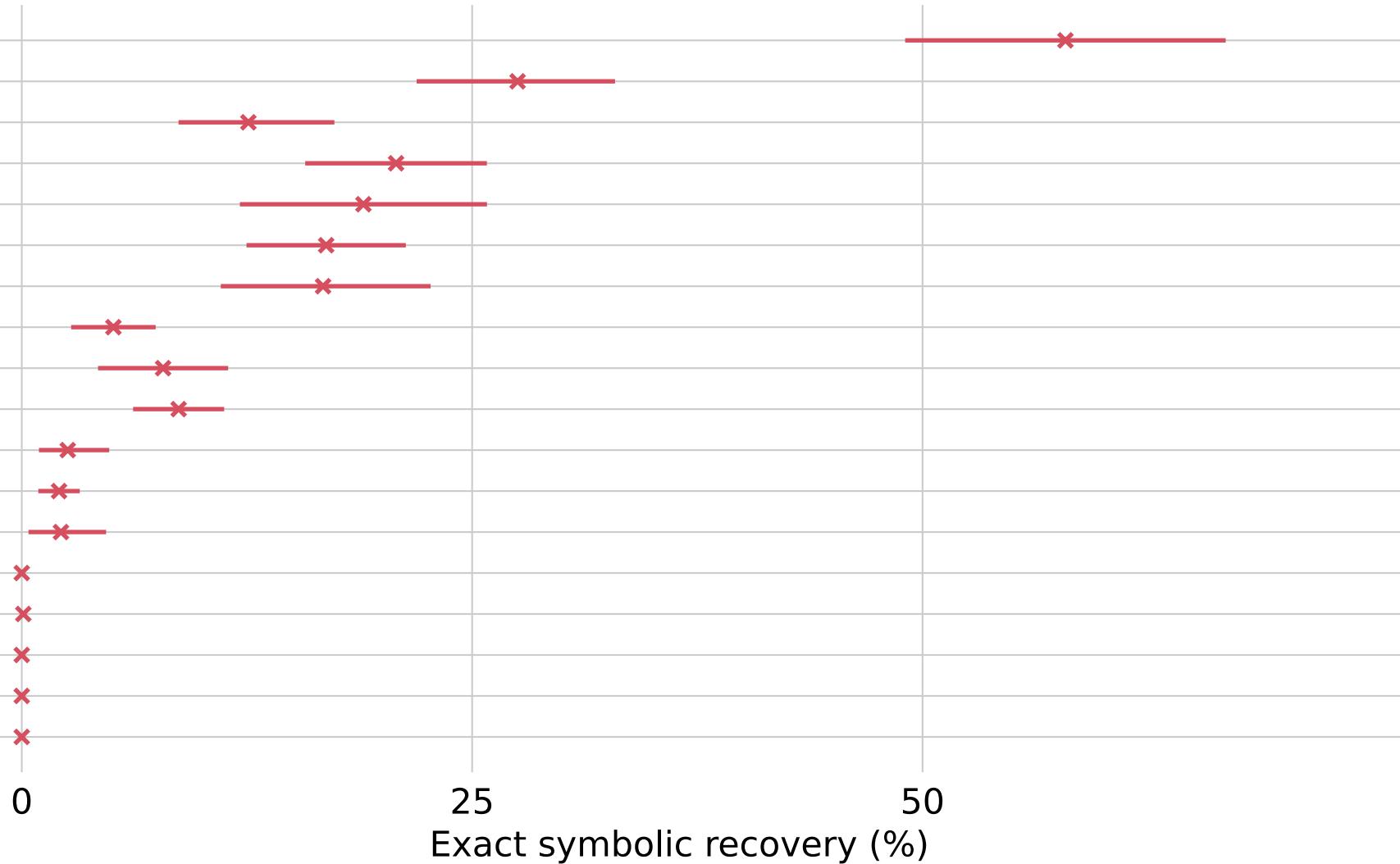
SBP-GP

BSR

FEAT

FFX

MRGP



## Feynman benchmark

Noise : 10%

**PhySO (this work)**

uDSR\*

AI Feynman 2.0

Eureqa/PySR

DSR

AFP

gplearn

GP-GOMEA

ITEA

EPLEX

NeSymReS\*

Operon

SINDy\*

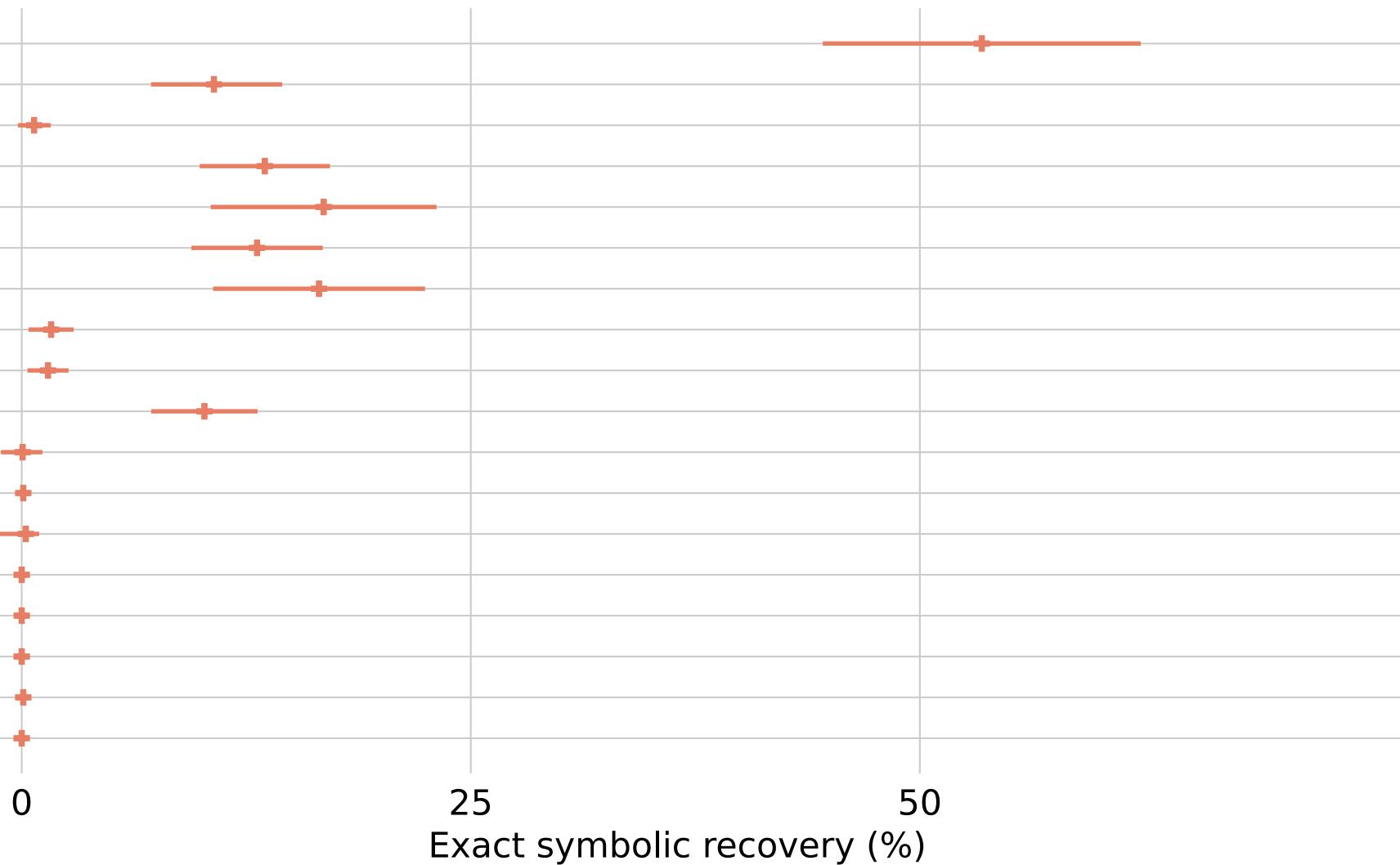
SBP-GP

BSR

FEAT

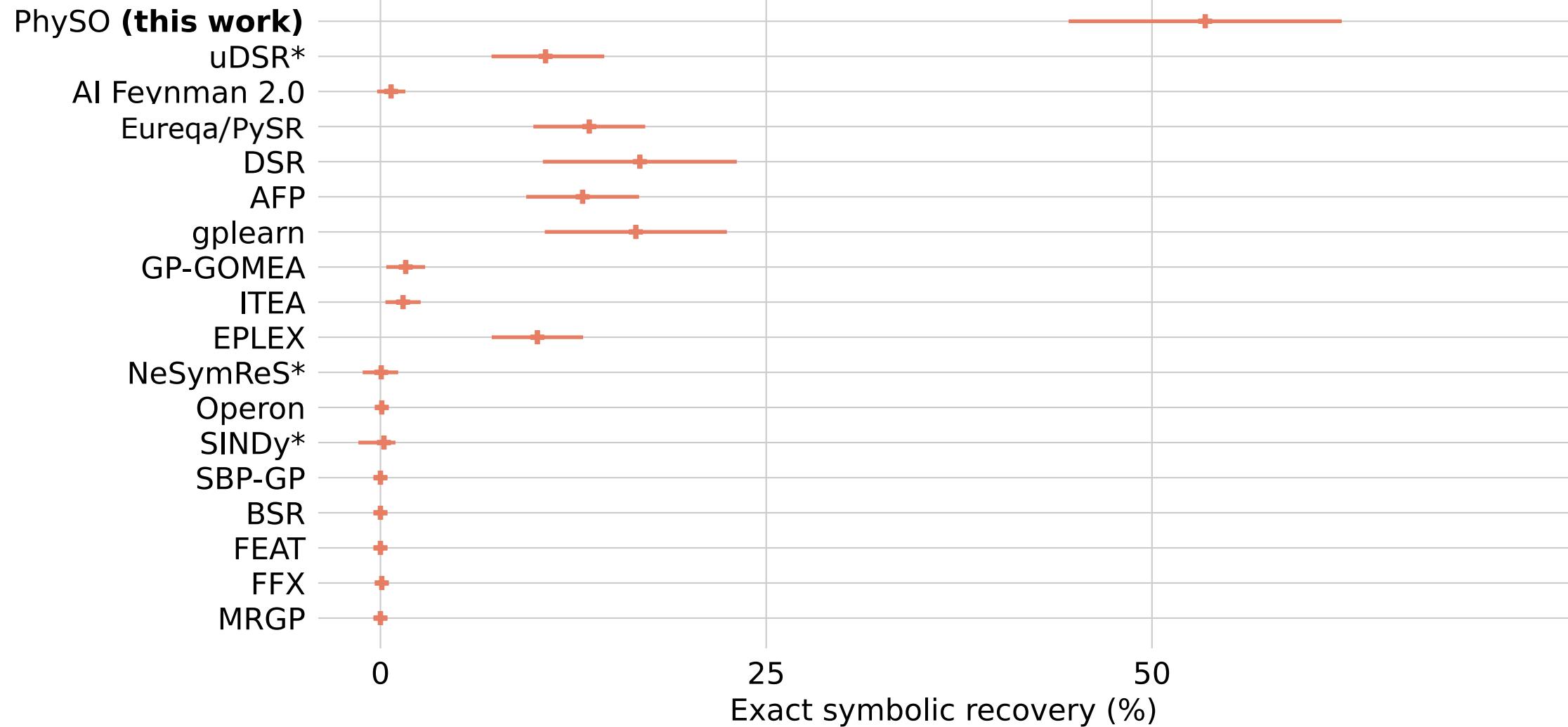
FFX

MRGP



# Feynman benchmark

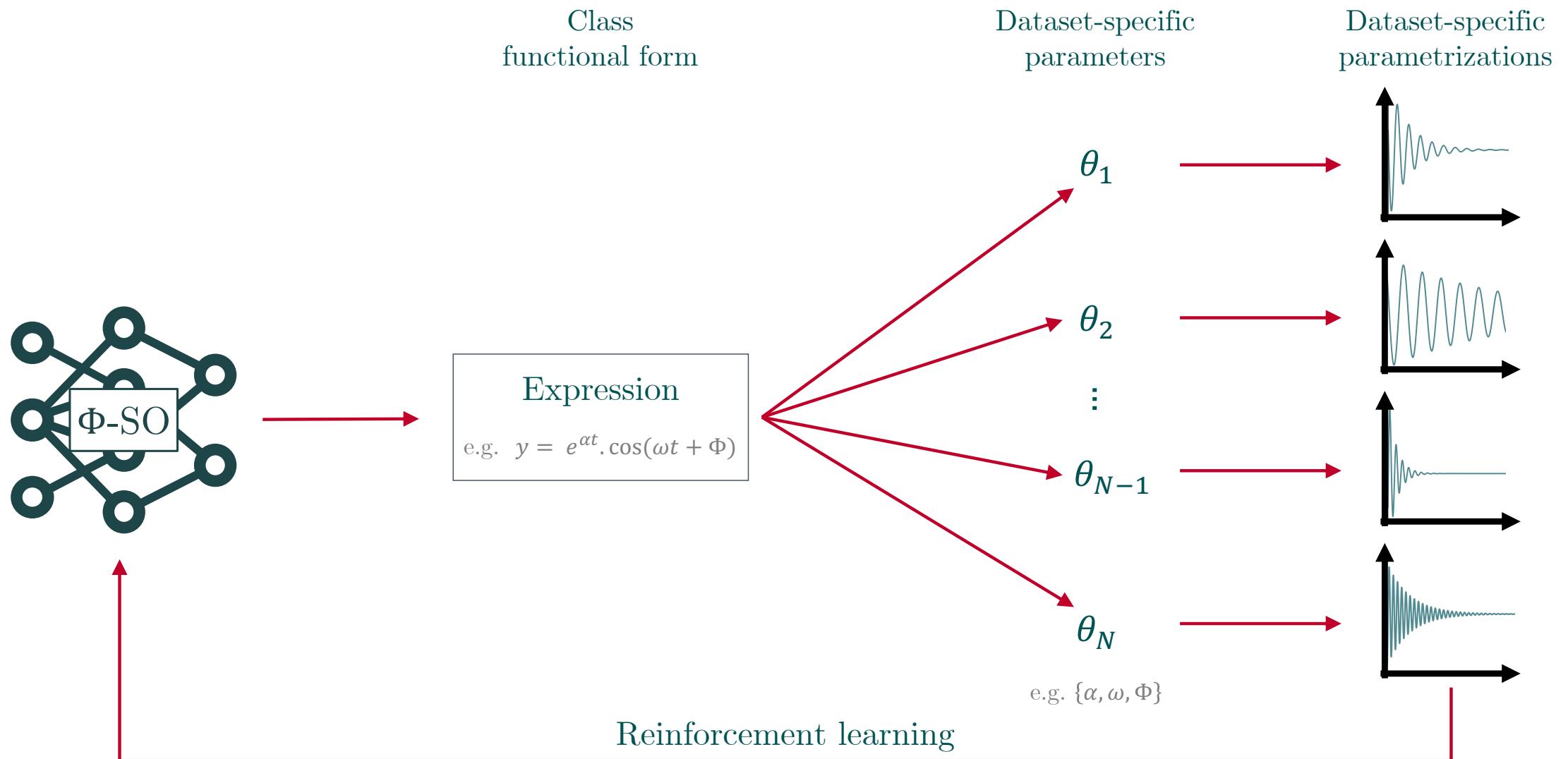
Noise : 10%



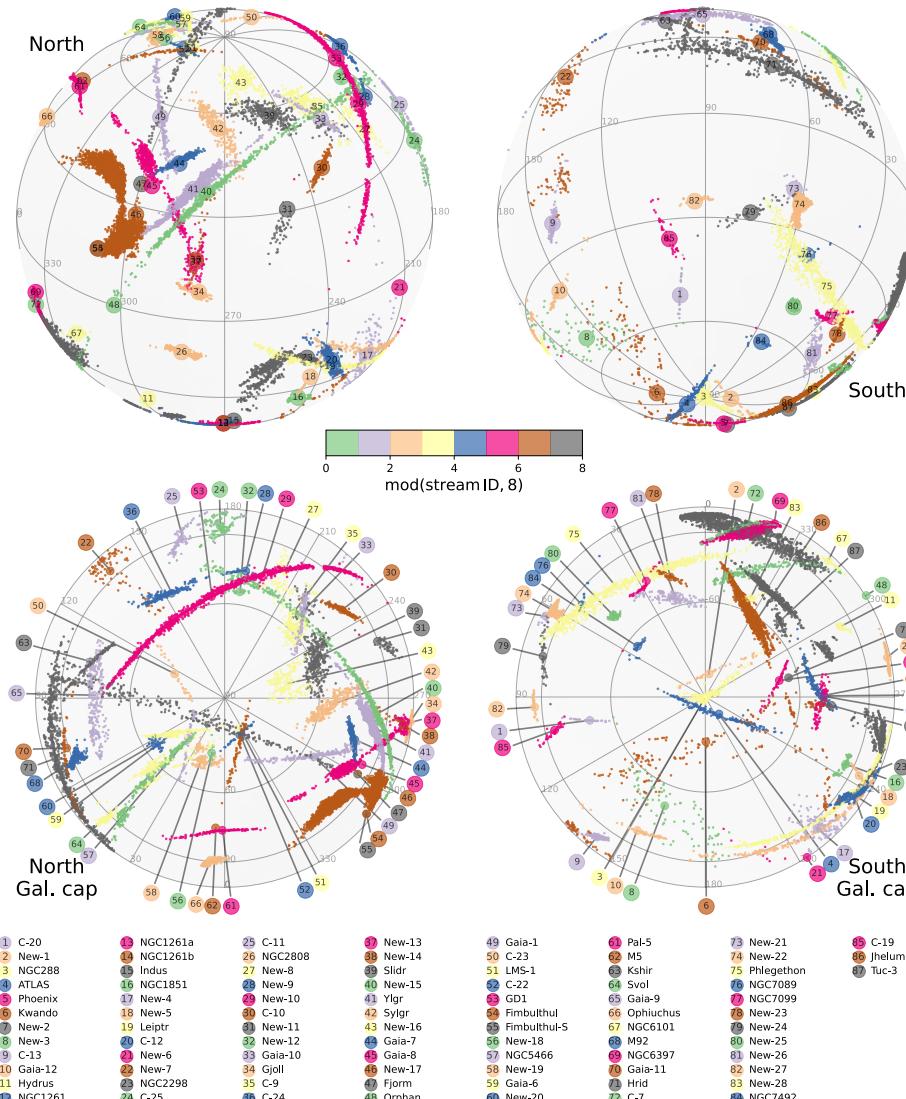
→ Φ-SO is the state-of-the-art in noisy scenarios

→ Neural net is better than no neural net (e.g. PySR, Operon etc.)

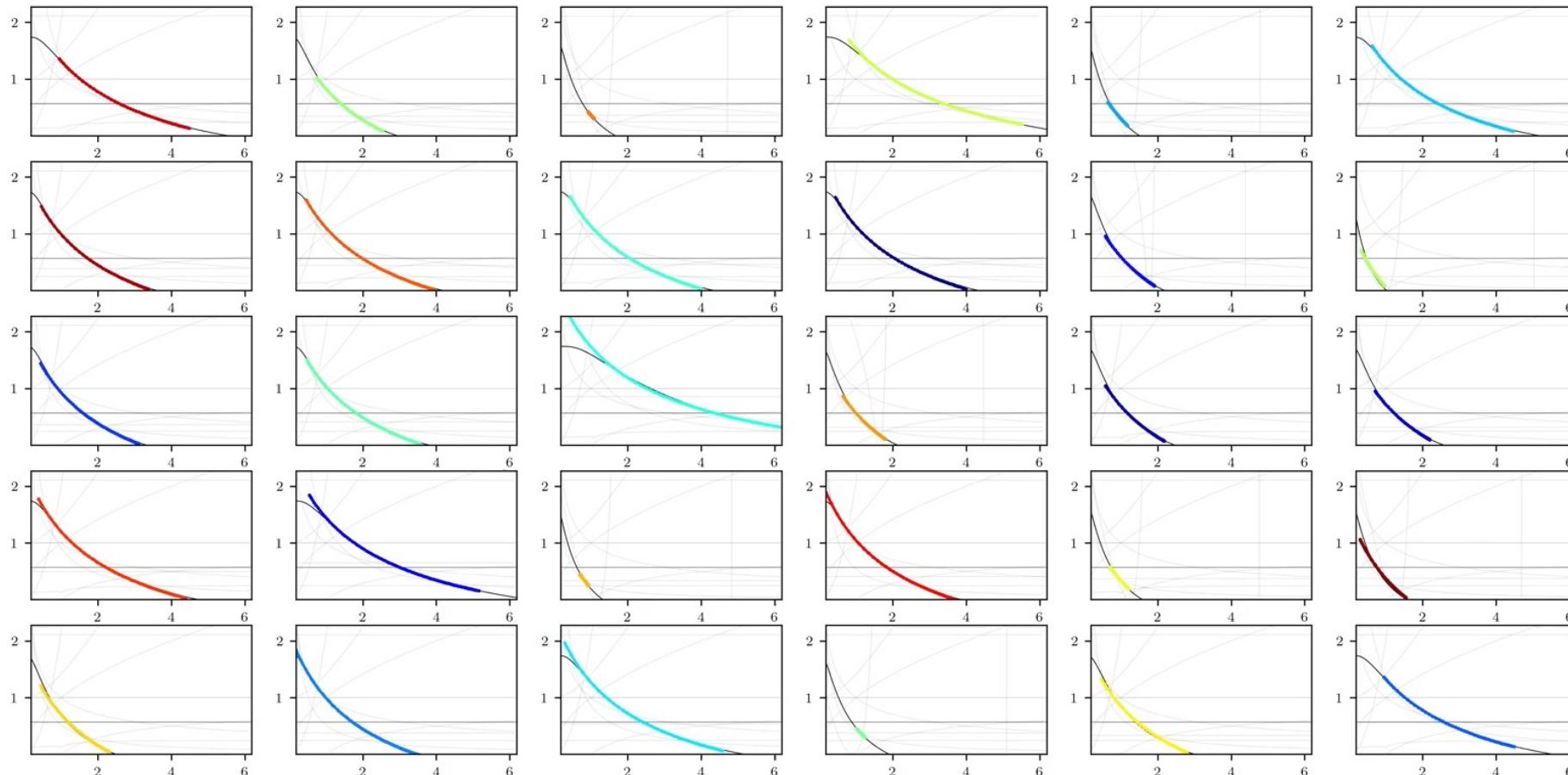
# Class symbolic regression



# Class SR : Milky Way stellar streams



# Class SR : Milky Way stellar streams



Best fit:

$$E(r) = \frac{-A c e^{\frac{R}{r}} - E_t c e^{\frac{R}{r}} + E_t}{c e^{\frac{R}{r}} - 1}$$

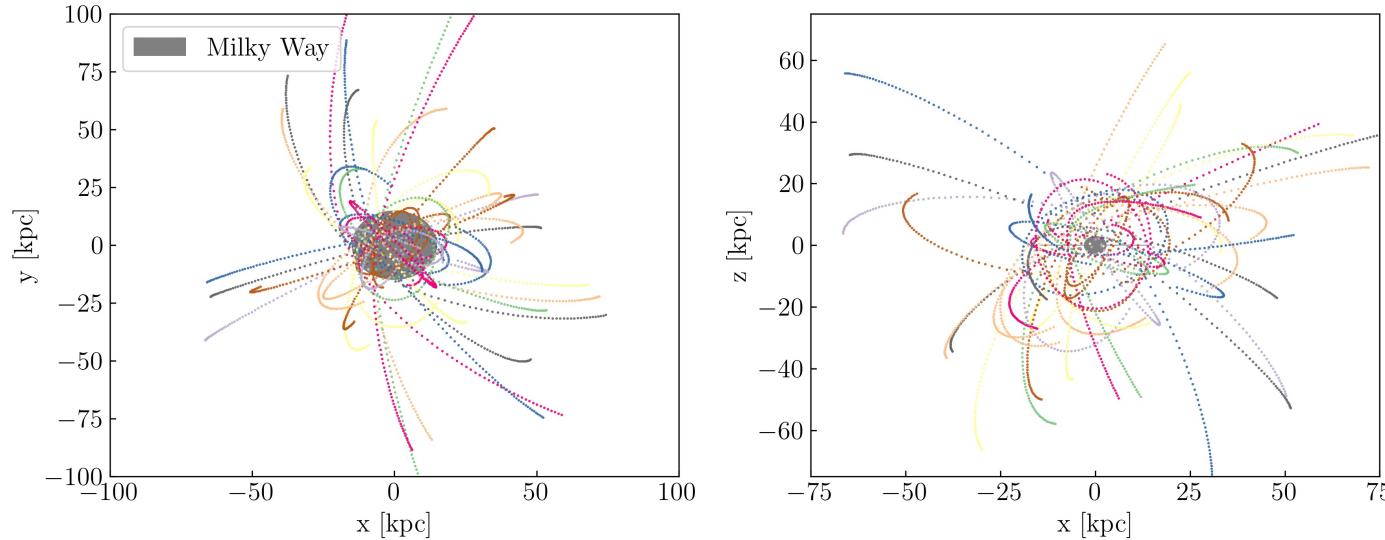
<https://youtu.be/iZKuCTG0YA4>

Trying:

$$E(r) = \frac{-A c e^{\frac{R}{r}} - E_t c e^{\frac{R}{r}} + E_t}{c e^{\frac{R}{r}} - 1}$$

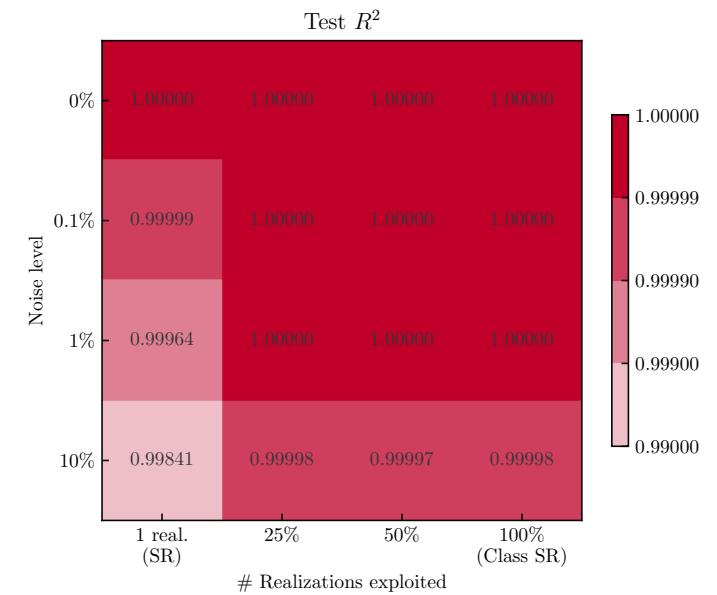
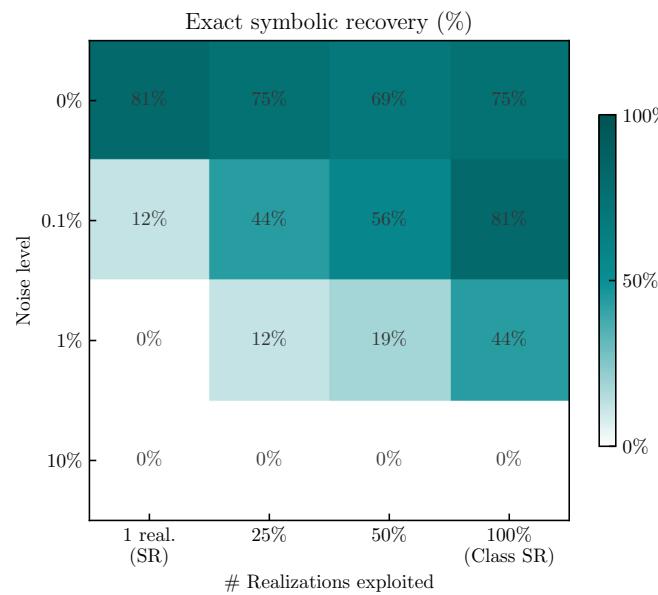
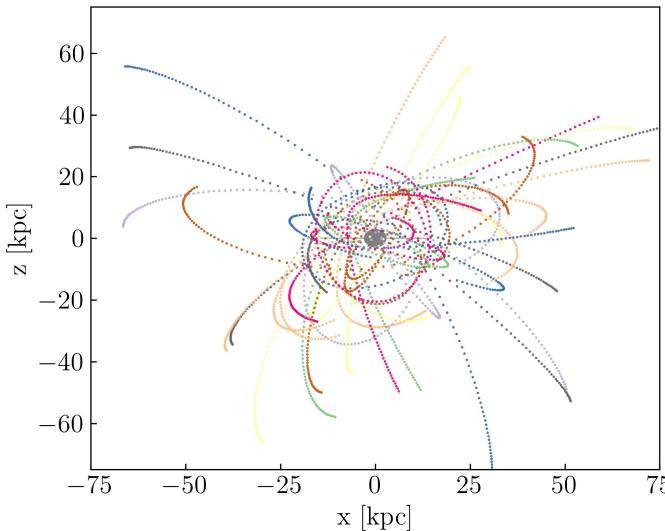
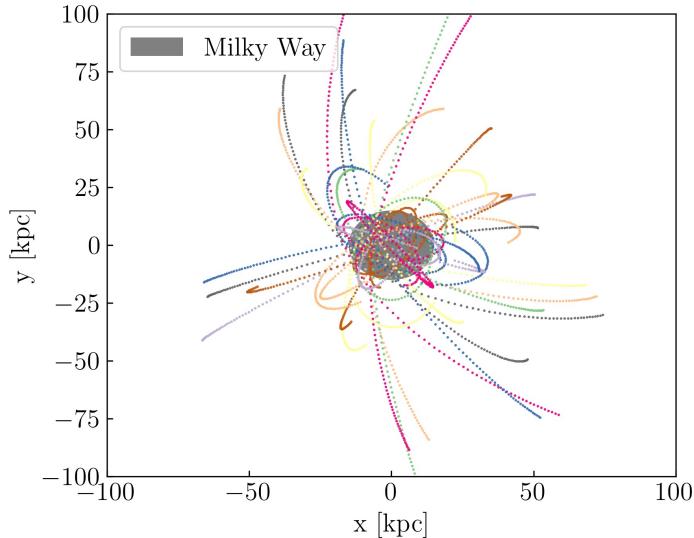
[Tenachi et al 2024] (Class SR)

## Class SR : Milky Way stellar streams



[Tenachi et al 2024] (Class SR)

# Class SR : Milky Way stellar streams



[Tenachi et al 2024] (Class SR)

# PhySO development plan

PhySO extension #1 – Dimensional analysis ✓

PhySO extension #2 – Class SR ✓

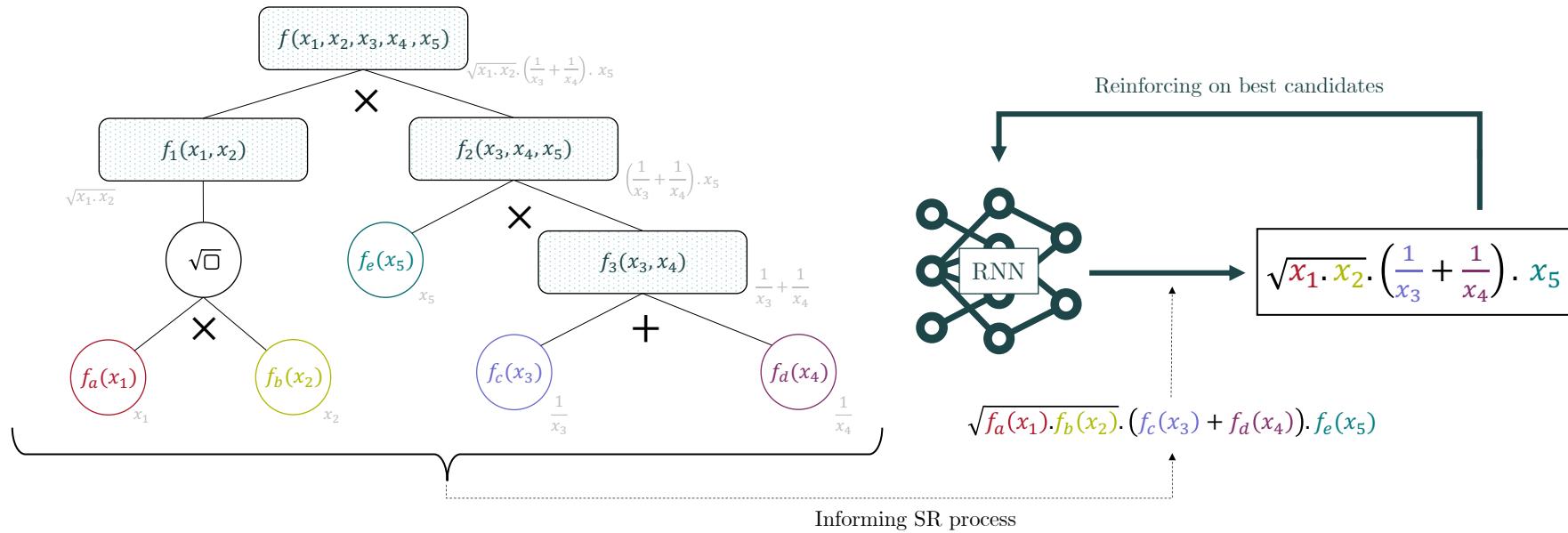
PhySO extension #3 – Structure analysis F

PhySO extension #4 – Differential equations F

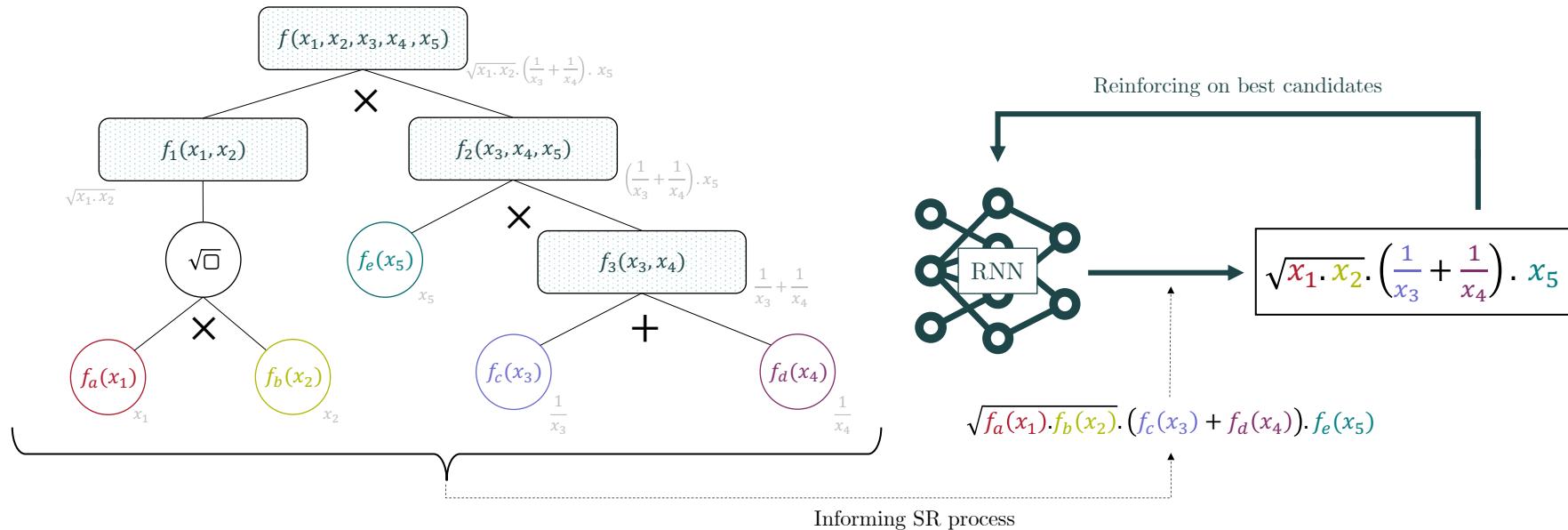
PhySO extension #5 – Hindsight relabeling RL F

PhySO extension #6 – Theory Φ-SO (planned)

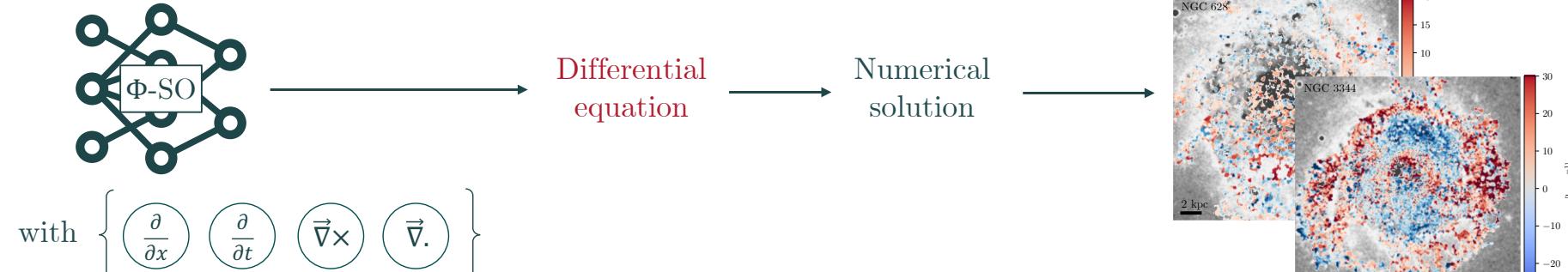
# Unveiling graph structures by detecting separabilities in data



# Unveiling graph structures by detecting separabilities in data

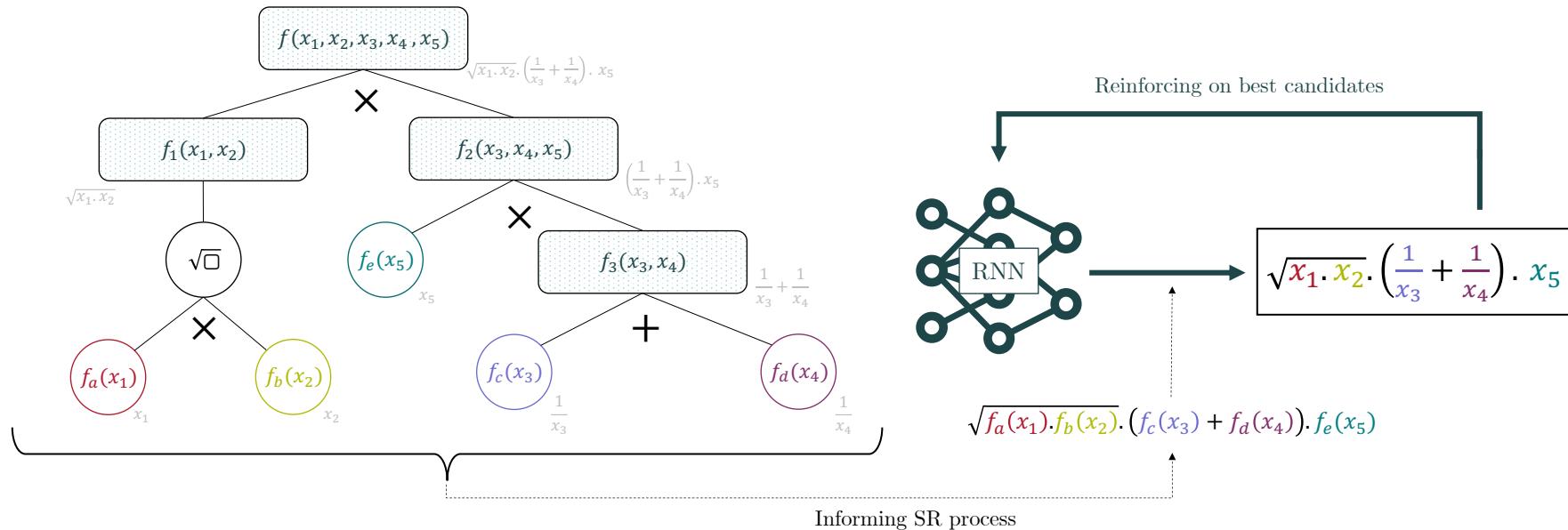


## Uncovering differential equations from data

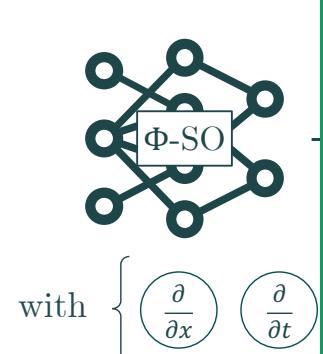


Adapted from [Urrejola-Mora et al 2022]

# Unveiling graph structures by detecting separabilities in data



## Uncovering differentiable graph structures

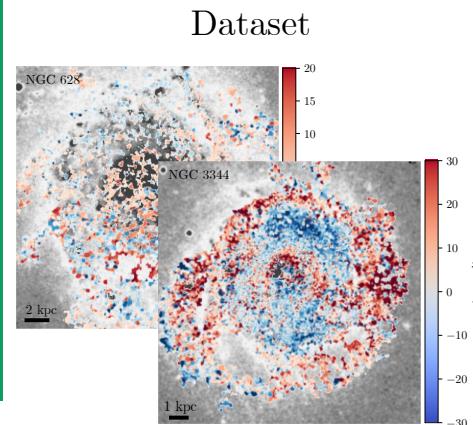


S15b: The changing macrocosm of astroinformatics : big data, catalogues, machine learning, challenges, and more!

Thu. 26th Jun.

15:00 - 16:30

Boole 2



Adapted from [Urrejola-Mora et al 2022]

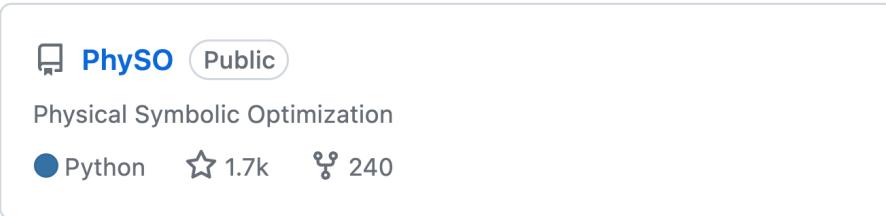
# Physical Symbolic Optimization

Φ-SO

An open source ...



Github repository: [WassimTenachi/PhySO](#)



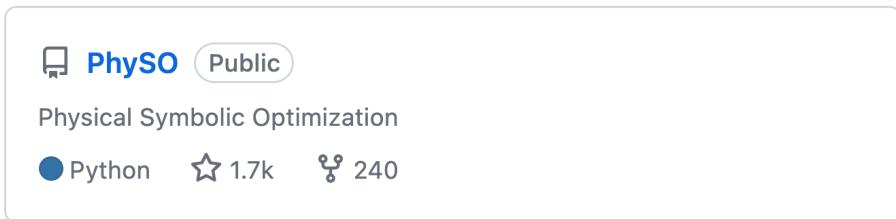
Physical Symbolic Optimization.

# Physical Symbolic Optimization

An open source ...

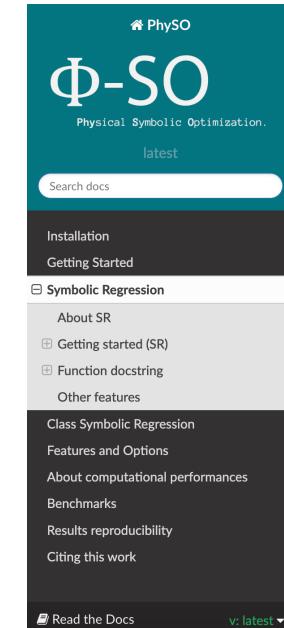


Github repository: [WassimTenachi/PhySO](https://github.com/WassimTenachi/PhySO)



... and documented package

Documentation: [physo.readthedocs.io](https://physo.readthedocs.io)



## Getting started (SR)

In this tutorial, we show how to use `physo` to perform Symbolic Regression (SR). The reference notebook for this tutorial can be found here: [sr\\_quick\\_start.ipynb](#).

### Setup

Importing the necessary libraries:

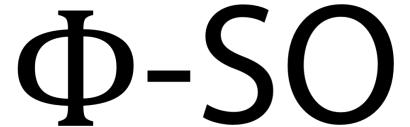
```
# External packages
import numpy as np
import matplotlib.pyplot as plt
import torch
```

Importing `physo`:

```
# Internal code import
import physo
import physo.learn.monitoring as monitoring
```

It is recommended to fix the seed for reproducibility:

```
# Seed
seed = 0
np.random.seed(seed)
torch.manual_seed(seed)
```



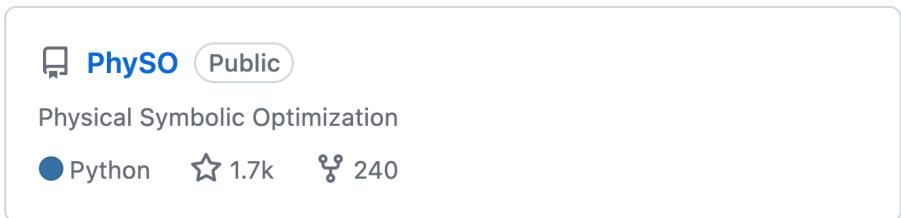
Physical Symbolic Optimization.

# Physical Symbolic Optimization

An open source ...



Github repository: [WassimTenachi/PhySO](https://github.com/WassimTenachi/PhySO)

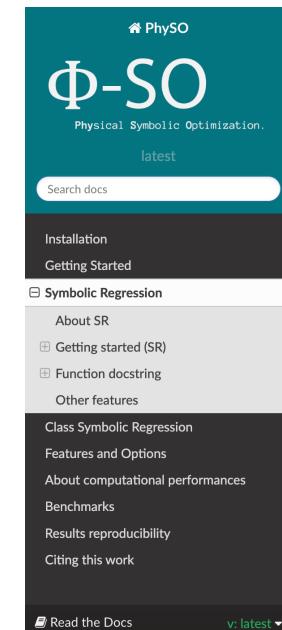


Creating a physical free-form symbolic analytical function  $f: \mathbb{R}^n \rightarrow \mathbb{R}$  fitting  $y = f(\mathbf{X})$  given  $(\mathbf{X}, y)$  data:

```
expression, logs = physo.SR(X, y,
    X_units = [ [1, 0, 0] , [1, -1, 0] ],
    y_units = [2, -2, 1],
    fixed_consts      = [ 1.      ],
    fixed_consts_units = [ [0,0,0] ],
    free_consts_units = [ [0, 0, 1] , [1, -2, 0] ],
)
```

... and documented package

Documentation: [physo.readthedocs.io](https://physo.readthedocs.io)



## Getting started (SR)

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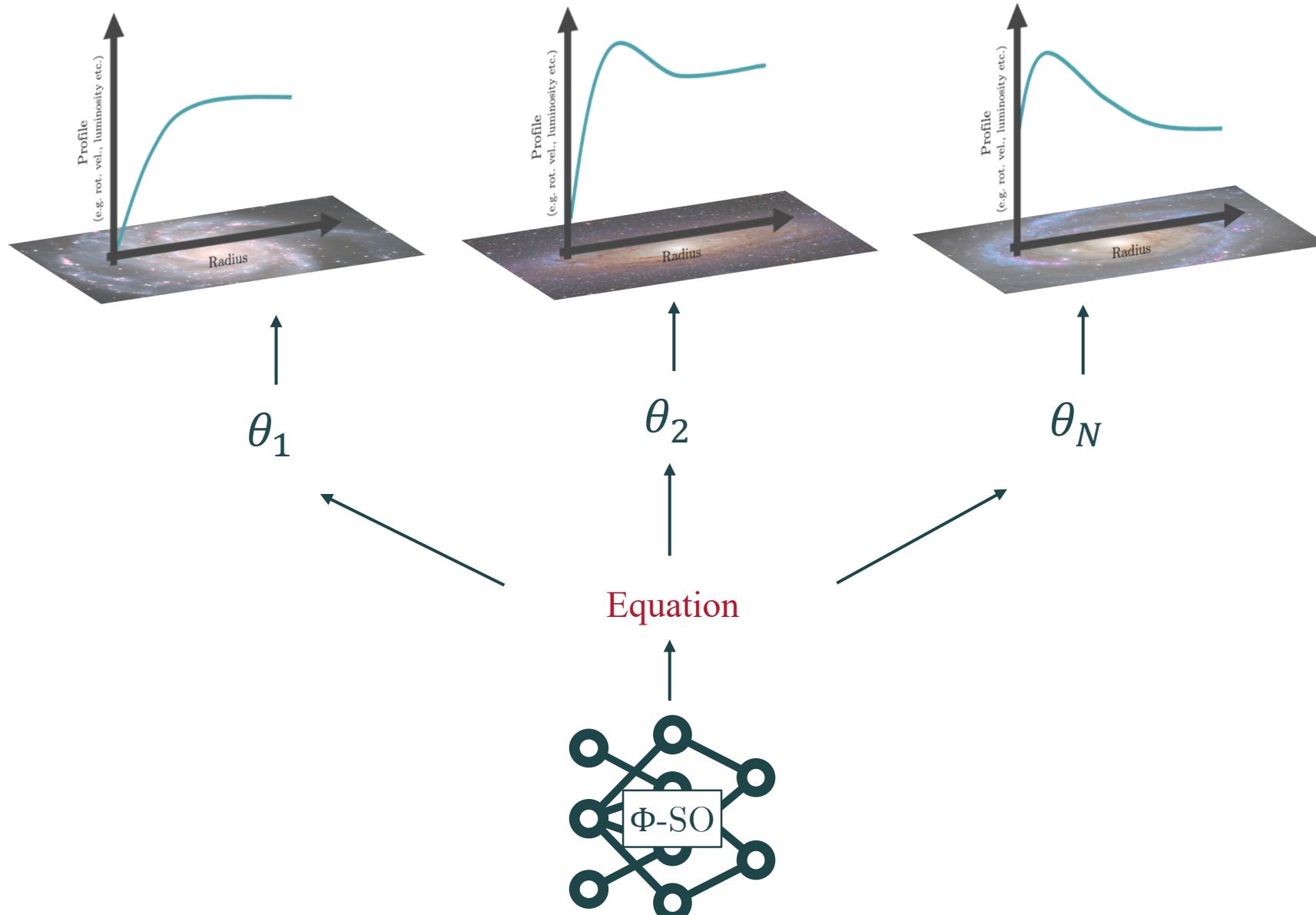
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```

It is recommended to fix the seed for reproducibility:

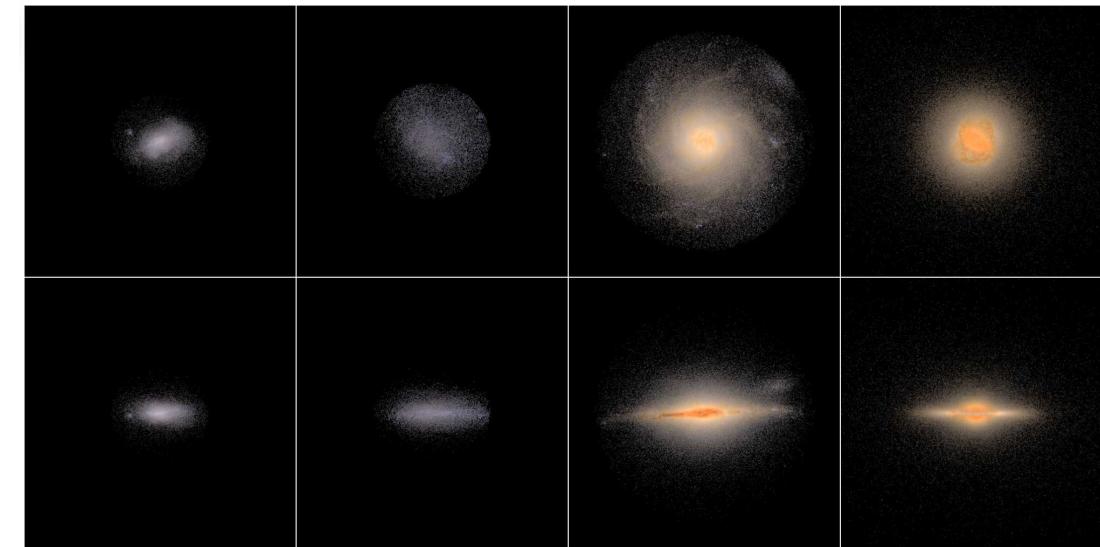
```
# Seed
seed = 0
np.random.seed(seed)
torch.manual_seed(seed)
```

# Class SR for galactic profiles



# Learning a dark matter profile

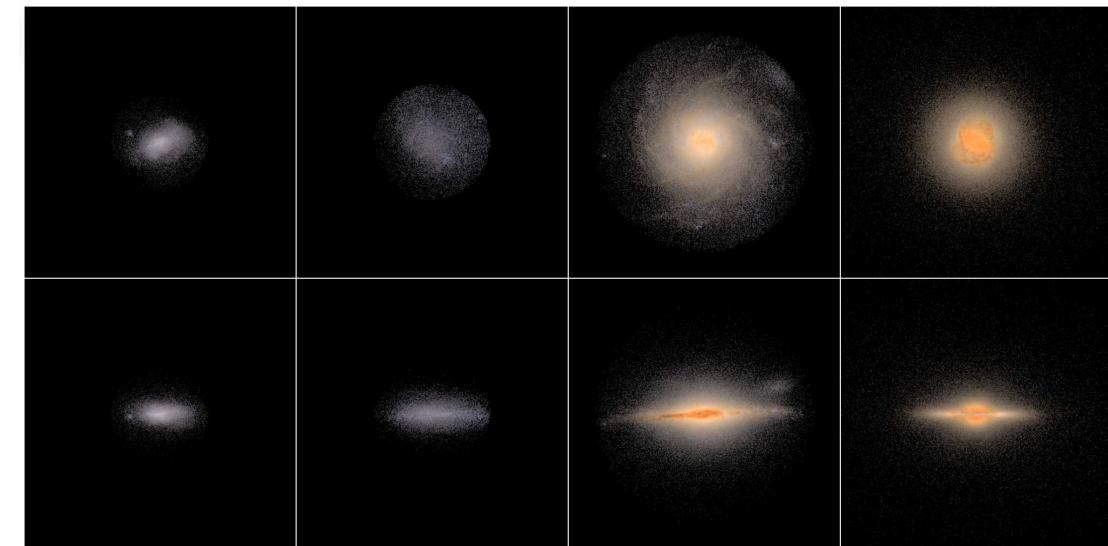
Optimal DM profile



Numerical Investigation of a Hundred Astrophysical Objects (NIHAO) [[Butsky 2015](#)]

# Learning a dark matter profile

Optimal DM profile



Numerical Investigation of a Hundred Astrophysical Objects (NIHAO) [Butsky 2015]

NFW profile

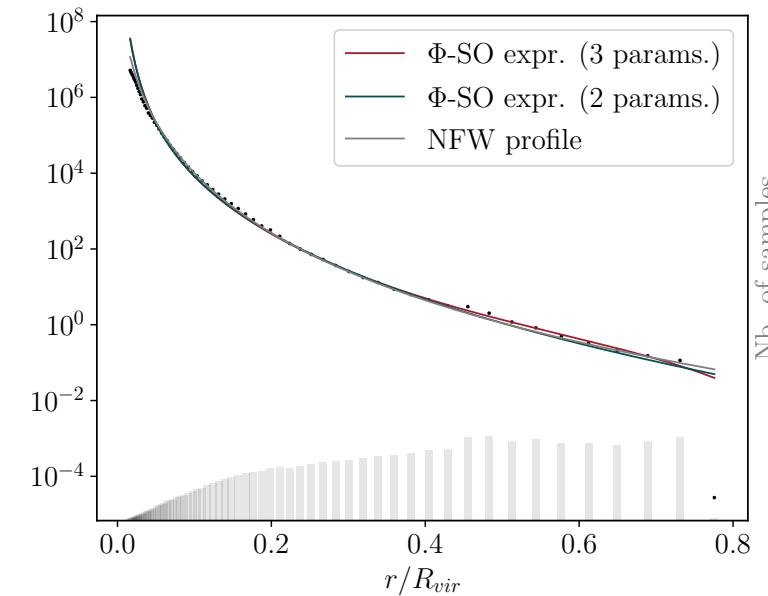
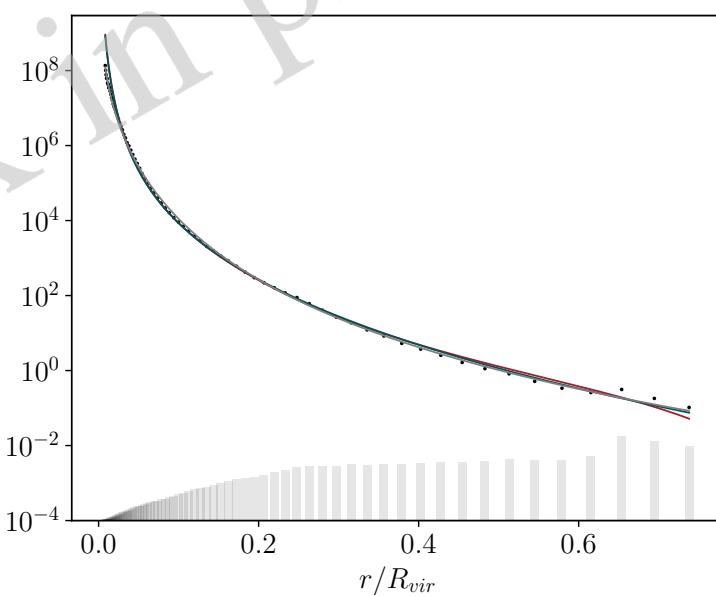
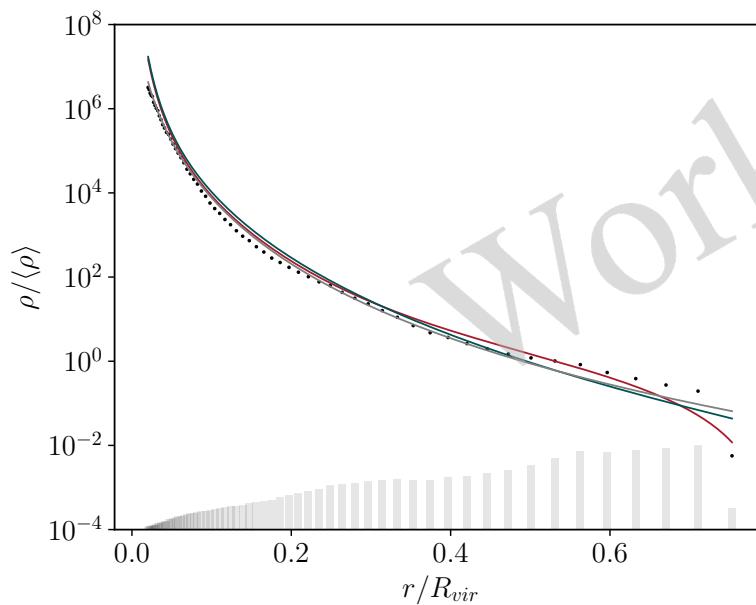
$$\frac{\rho_0 r_s^3}{r (r + r_s)^2}$$

$\Phi$ -SO expr. (2 params.)

$$\frac{\rho_0 r_s^2 \left( c_0 + e^{\frac{c_0 r_s}{r}} \right)}{c_0^2 r^2}$$

$\Phi$ -SO expr. (3 params.)

$$\rho_0 - \frac{\rho_0 \left( c_0 r_{s0} - r_{s1} e^{\frac{c_0 r_{s1}}{r}} \right)^2}{r^2}$$



Nb. of samples

# Derived quantities

Density profile       $\rho(r)$

# Derived quantities

Density profile

$$\rho(r)$$

Assumptions:

- Spherical symmetry
- Isotropic velocity distribution
- Steady-state (time-independent)

# Derived quantities

Density profile

$$\rho(r)$$

Enclosed mass

$$M(r) = 4\pi \int_0^r \rho(r') r'^2 dr'$$



Assumptions:

- Spherical symmetry
- Isotropic velocity distribution
- Steady-state (time-independent)

# Derived quantities

Density profile

$$\rho(r)$$

Enclosed mass

$$M(r) = 4\pi \int_0^r \rho(r') r'^2 dr'$$

Circular velocity

$$V_c(r) = \sqrt{\frac{GM(r)}{r}}$$

Assumptions:

- Spherical symmetry
- Isotropic velocity distribution
- Steady-state (time-independent)



# Derived quantities

Density profile

$$\rho(r)$$

Enclosed mass

$$M(r) = 4\pi \int_0^r \rho(r') r'^2 dr'$$

Circular velocity

$$V_c(r) = \sqrt{\frac{GM(r)}{r}}$$

Potential

$$\Phi(r) = -G \int_r^\infty \frac{M(r')}{r'^2} dr'$$

Assumptions:

- Spherical symmetry
- Isotropic velocity distribution
- Steady-state (time-independent)

$$\nabla^2 \Phi = 4\pi G \rho$$

(Poisson)

# Derived quantities

Density profile

$$\rho(r)$$

Enclosed mass

$$M(r) = 4\pi \int_0^r \rho(r') r'^2 dr'$$

Circular velocity

$$V_c(r) = \sqrt{\frac{GM(r)}{r}}$$

Potential

$$\Phi(r) = -G \int_r^\infty \frac{M(r')}{r'^2} dr'$$

Radial velocity dispersion

$$\sigma_r^2(r) = \frac{1}{\rho(r)} \int_r^\infty \rho(r') \frac{GM(r')}{r'^2} dr' \quad \leftarrow$$

Assumptions:

- Spherical symmetry
- Isotropic velocity distribution
- Steady-state (time-independent)

$$\frac{d(\rho\sigma_r^2)}{dr} + \frac{2\beta}{r}\rho\sigma_r^2 = -\rho \frac{GM(r)}{r^2} \quad (\text{Jeans})$$

# Derived quantities of DM profiles

		$M(r)$	$V_c(r)$	$\Phi(r)$	$\sigma^2(r)$
Cusps	NFW $\rho(r) = \frac{\rho_0}{\left(\frac{r}{r_s}\right)\left(1 + \frac{r}{r_s}\right)^2}$	✓	✓	✓	✗
	SuperNFW $\rho(r) = \frac{\rho_0}{\left(\frac{r}{r_s}\right)\left(1 + \frac{r}{r_s}\right)^{5/2}}$	✓	✓	✓	✓
Cores	pISO $\rho(r) = \frac{\rho_0}{1 + \left(\frac{r}{r_s}\right)^2}$	✓	✓	✓	✓
	Burkert $\rho(r) = \frac{\rho_0}{\left(1 + \frac{r}{r_s}\right)\left(1 + \left(\frac{r}{r_s}\right)^2\right)}$	✓	✓	✓	✗
Variable slope	Lucky13 $\rho(r) = \frac{\rho_0}{\left(1 + \frac{r}{r_s}\right)^3}$	✓	✓	✓	✗
	Einasto $\rho(r) = \rho_0 \exp\left(-\frac{2}{\alpha} \left[\left(\frac{r}{r_s}\right)^\alpha - 1\right]\right)$	✗	✗	✗	✗
	gNFW $\rho(r) = \frac{\rho_0}{\left(\frac{r}{r_s}\right)^\alpha \left(1 + \left(\frac{r}{r_s}\right)\right)^{3-\alpha}}$	✗	✗	✗	✗
	✗ Non-elementary functions				

# Learning a profile with analytic constraints

$$\rho(r)$$

$$-c_0\rho_0 - \rho_0 \log(-c_0^2) + \frac{\rho_0}{c_0}$$

$$M(r)$$

$$r^3 \left( -\frac{4\pi c_0 \rho_0}{3} - \frac{8\pi \rho_0 \log(c_0)}{3} - \frac{4i\pi^2 \rho_0}{3} + \frac{4\pi \rho_0}{3c_0} \right)$$

$$V_c(r)$$

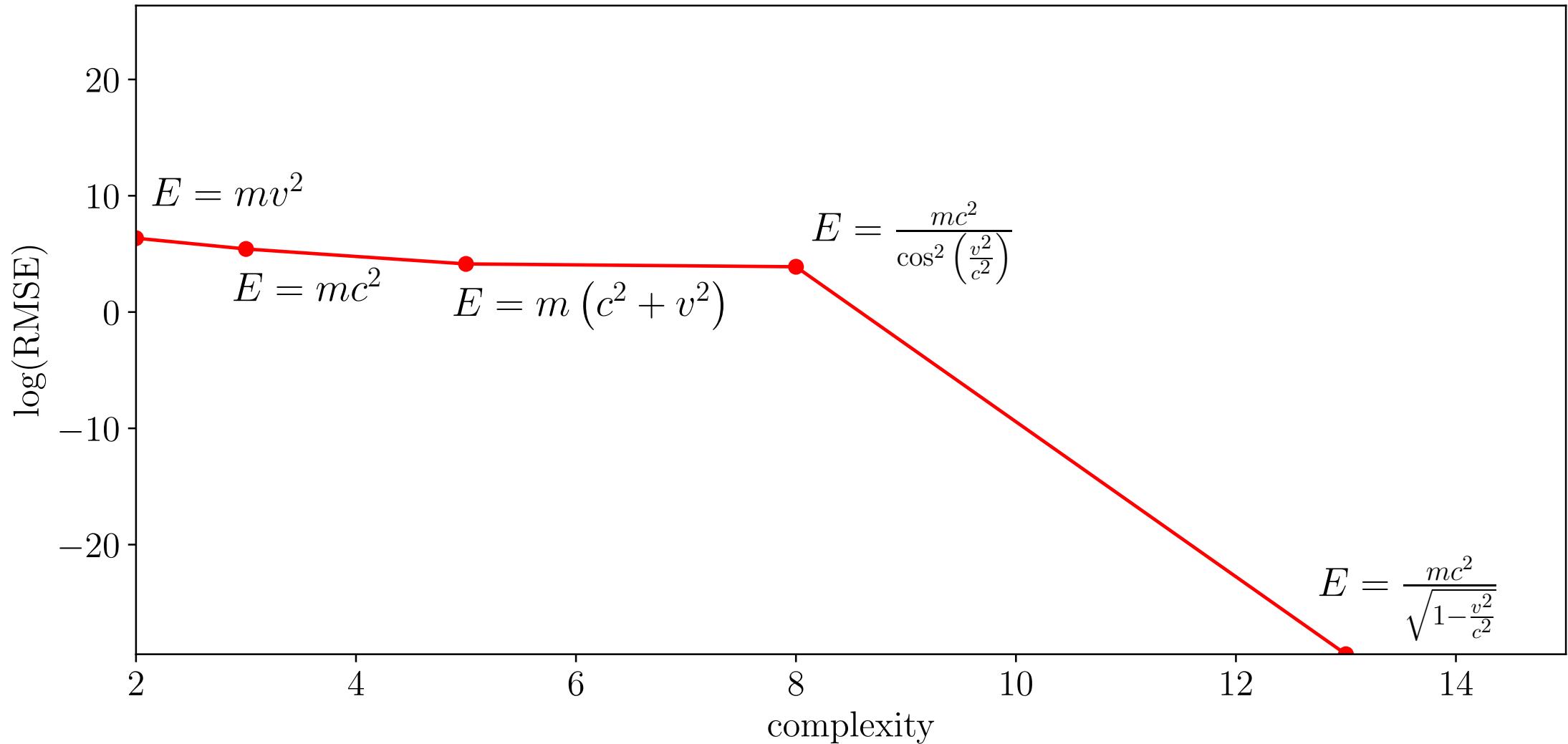
$$\frac{2\sqrt{3}\sqrt{\pi}\sqrt{G}r\sqrt{\rho_0}\sqrt{-c_0(c_0+2\log(c_0)+i\pi)+1}}{3\sqrt{c_0}}$$

$$\Phi(r)$$

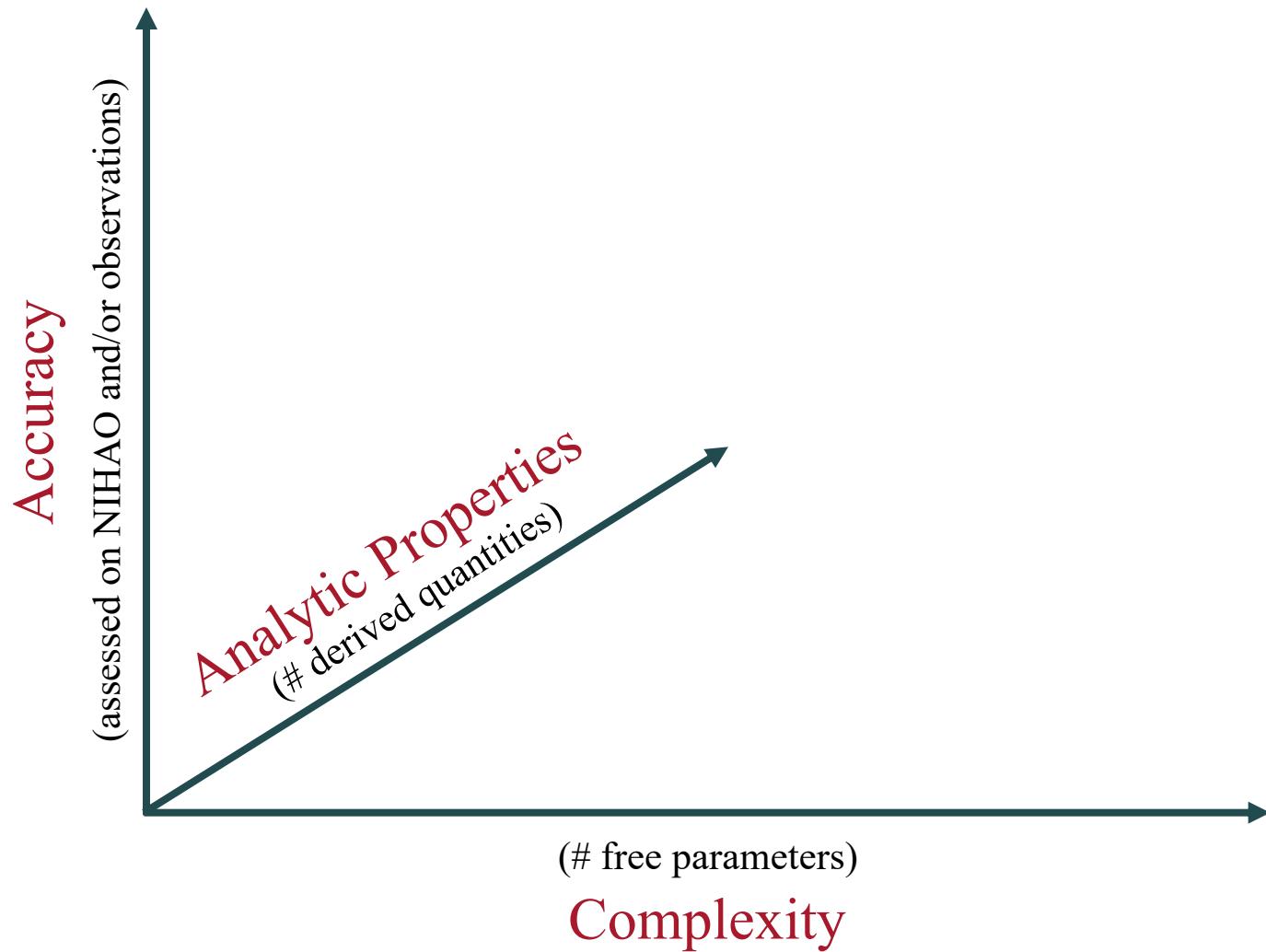
$$r^2 \left( -\frac{2\pi G c_0 \rho_0}{3} - \frac{4\pi G \rho_0 \log(c_0)}{3} - \frac{2i\pi^2 G \rho_0}{3} + \frac{2\pi G \rho_0}{3c_0} \right)$$

<https://youtu.be/Js3VbbmEaIo>

# Pareto fronts: accuracy vs complexity optima

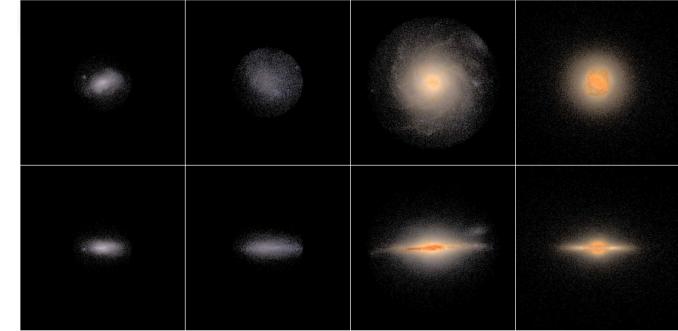
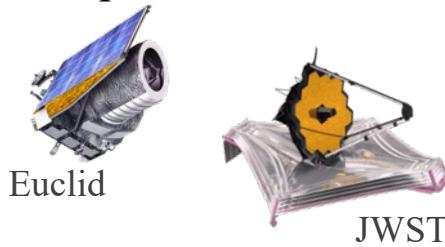


# Toward the ultimate DM profile(s)



# Perspectives

- Exploring anisotropic distributions
- Impact of the hydrodynamics on the functional form
- Including observational constraints via velocity-related quantities



Numerical Investigation of a Hundred Astrophysical Objects (NIHAO) [[Butsky 2015](#)]

- Other analytic constraints  
(eg. converging enclosed mass – unlike NFW)

$$\exists M_{\text{tot}} \in \mathbb{R}^+ : \lim_{r \rightarrow \infty} M(r) = M_{\text{tot}}$$

Thank you for your  
attention !

