

Deep Learning - Theory and Practice

*IE 643
Lecture 1*

July 28 & 31, 2025.

1 Introduction

2 Perceptron

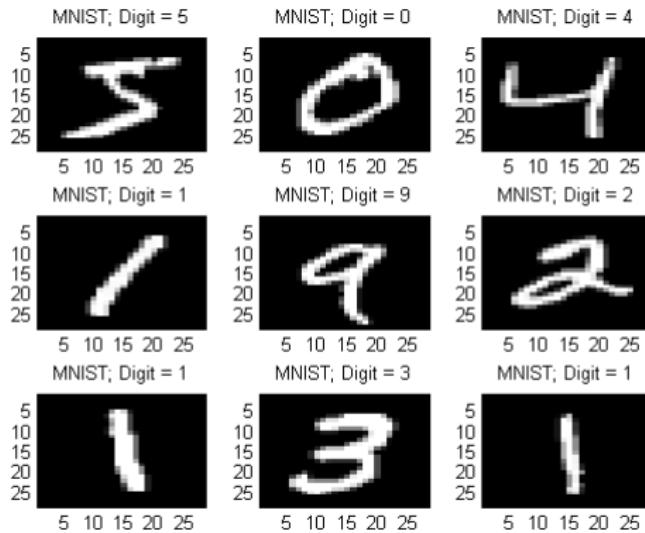
3 Nature of Machine Learning Tasks

- Supervised Machine Learning

Deep Learning - Motivating Applications

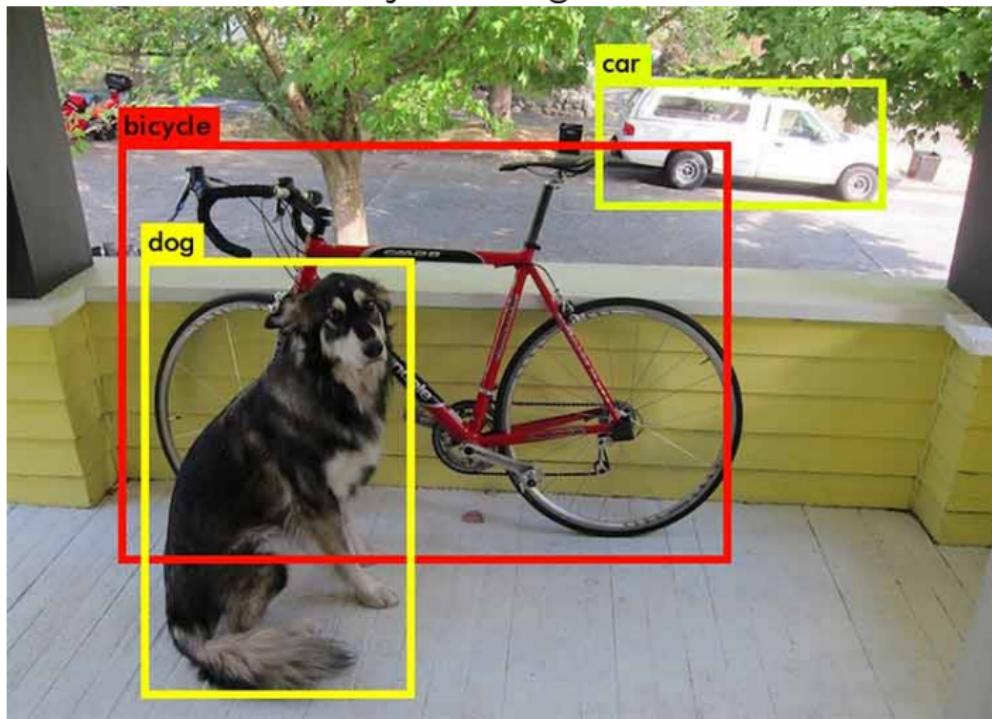
Deep Learning Applications

Handwritten Character Recognition



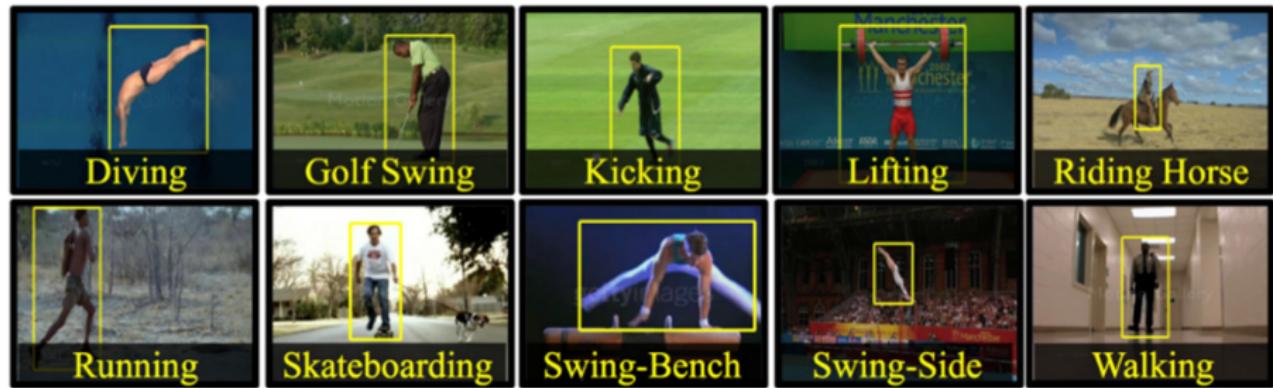
Deep Learning Applications

Object Recognition



Deep Learning Applications

Action Recognition



Deep Learning Applications

Machine Translation

English ▾



Thai ▾

I am a student.

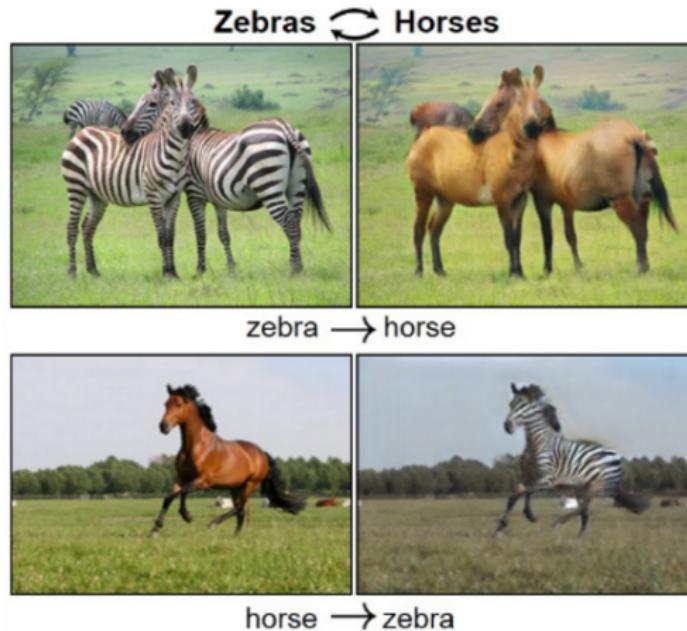


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Deep Learning Applications

Image Generation

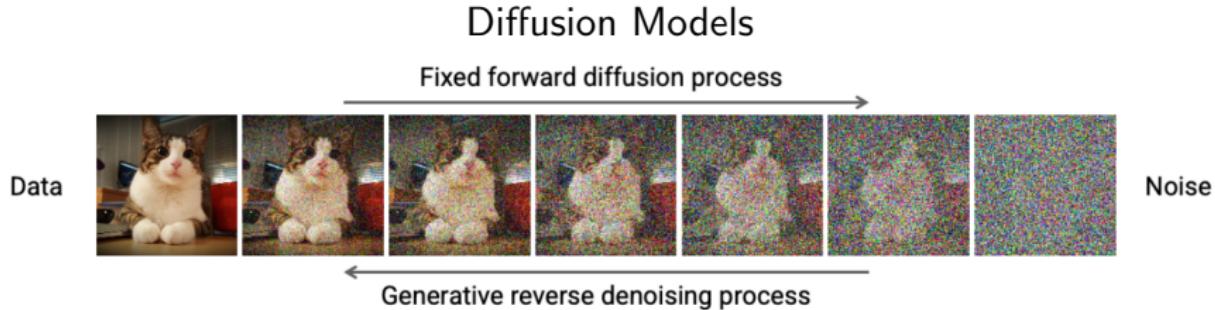


Deep Learning Applications

Other Applications

- Real-time object detection
- Real-time segmentation
- Chat bots
- Real-time generative tasks

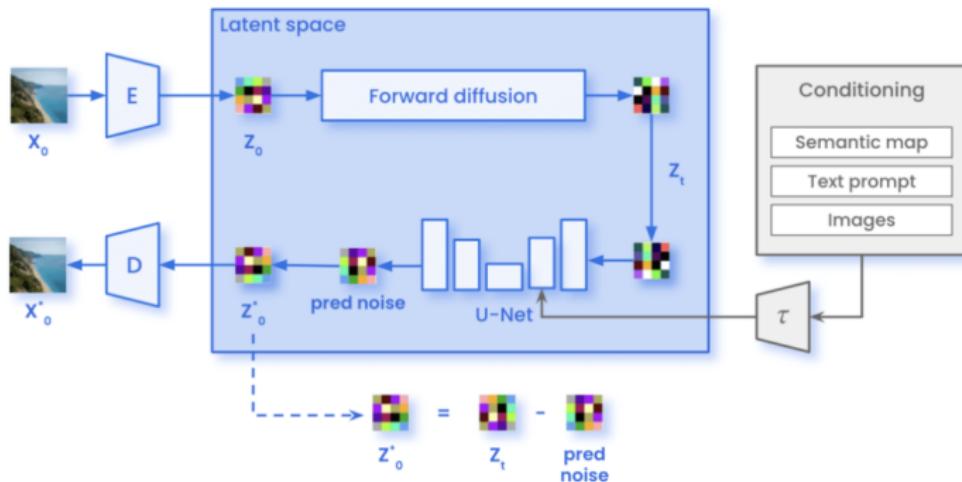
Generating images from text



source: <https://cvpr2022-tutorial-diffusion-models.github.io/>

Generating images from text

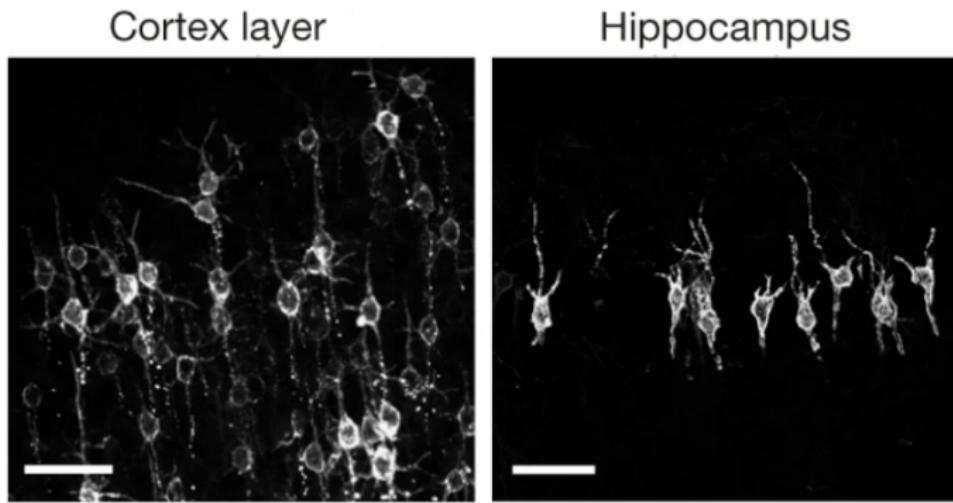
Diffusion Models



source: <https://blog.marvik.ai/>

Perceptron

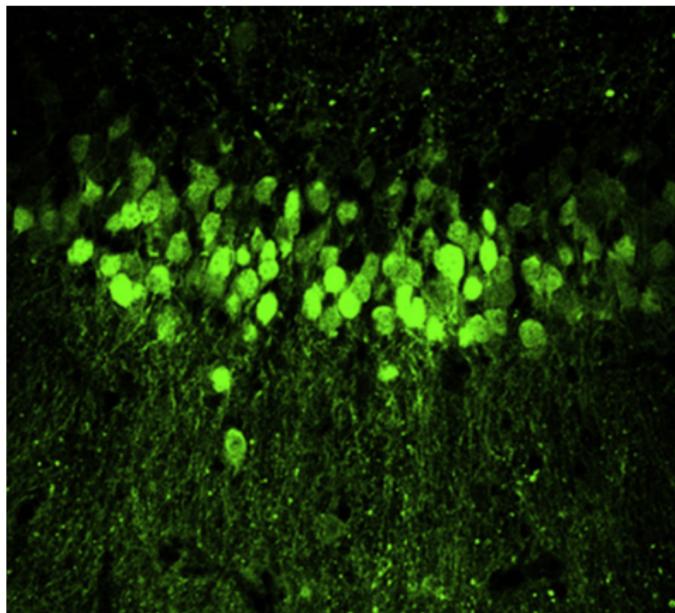
Biological Motivation



Population imaging of neural activity in awake behaving mice. K. D. Piatkevich et al. Nature, 574, pp. 413-117, 2019.

Biological Motivation

Neuronal activity in thalamus region of mouse¹



¹Ref: Shana Li et al. In vivo labeling and quantitative imaging of neuronal populations using MRI, *NeuroImage*, Vol. 281, 2023.

Perceptron

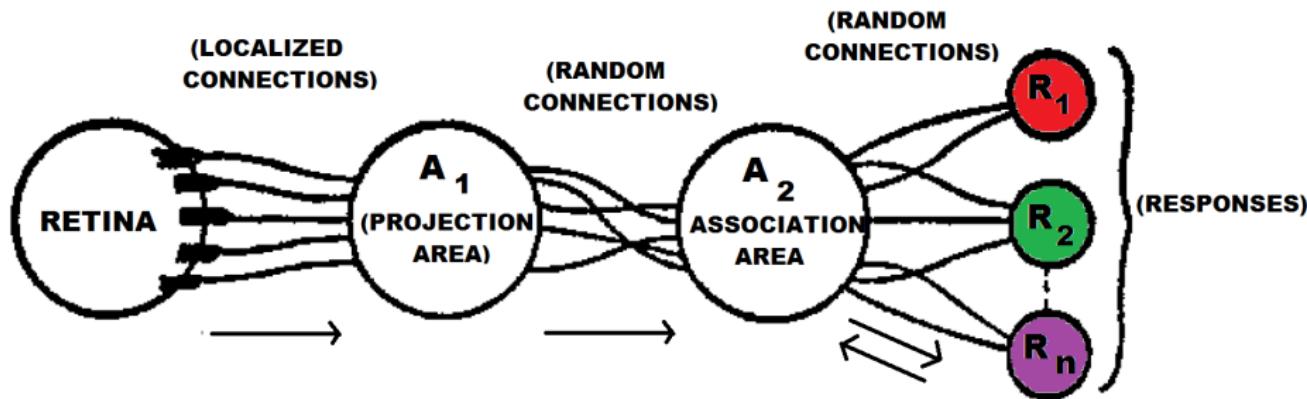
Psychological Review
Vol. 65, No. 6, 1958

THE PERCEPTRON: A PROBABILISTIC MODEL FOR INFORMATION STORAGE AND ORGANIZATION IN THE BRAIN

F. ROSENBLATT

Cornell Aeronautical Laboratory

Perceptron

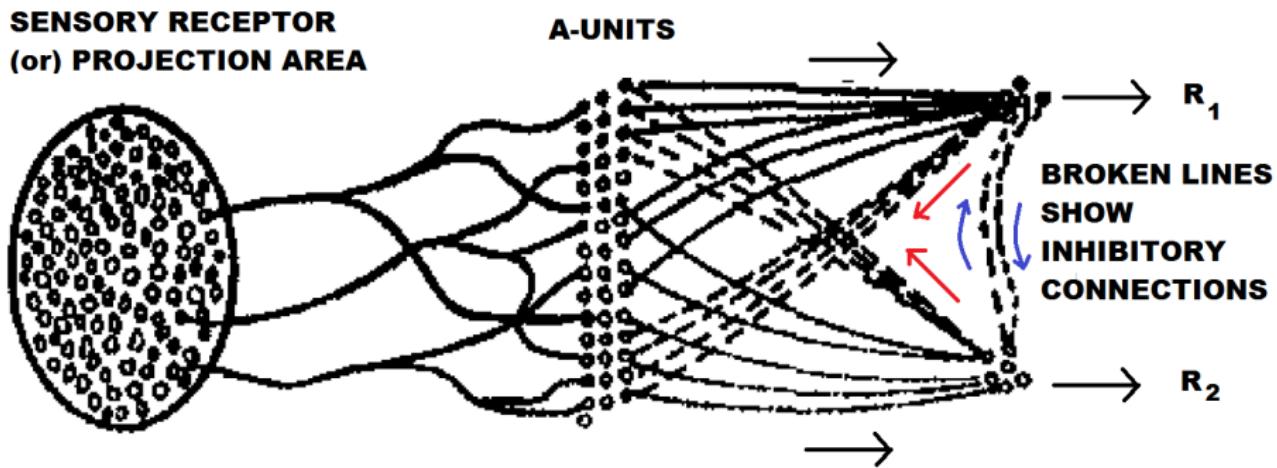


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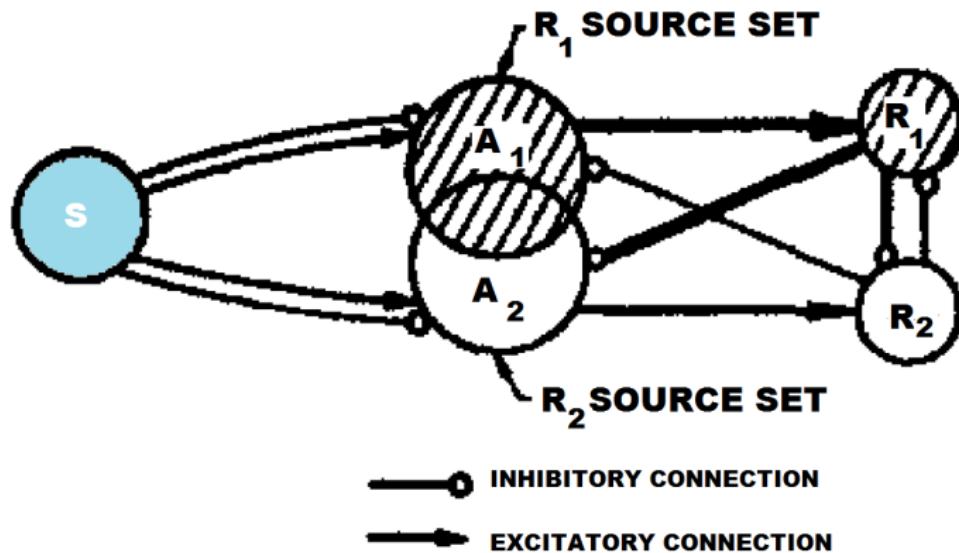
Key Assumptions

- Stimuli which are **similar** will tend to form pathways to same sets of response cells.
- Stimuli which are **dissimilar** will tend to form pathways to different sets of response cells.
- Application of positive or negative reinforcements may facilitate or hinder the formation of connections.
- **Similarity** of stimuli is a dynamically evolving attribute.

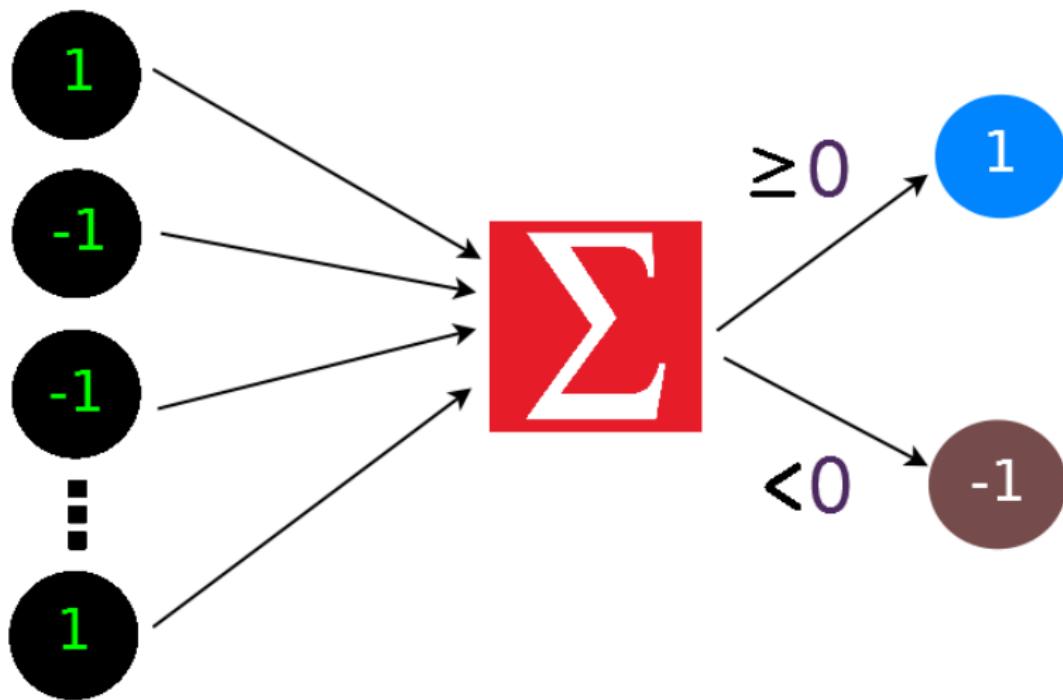
Perceptron



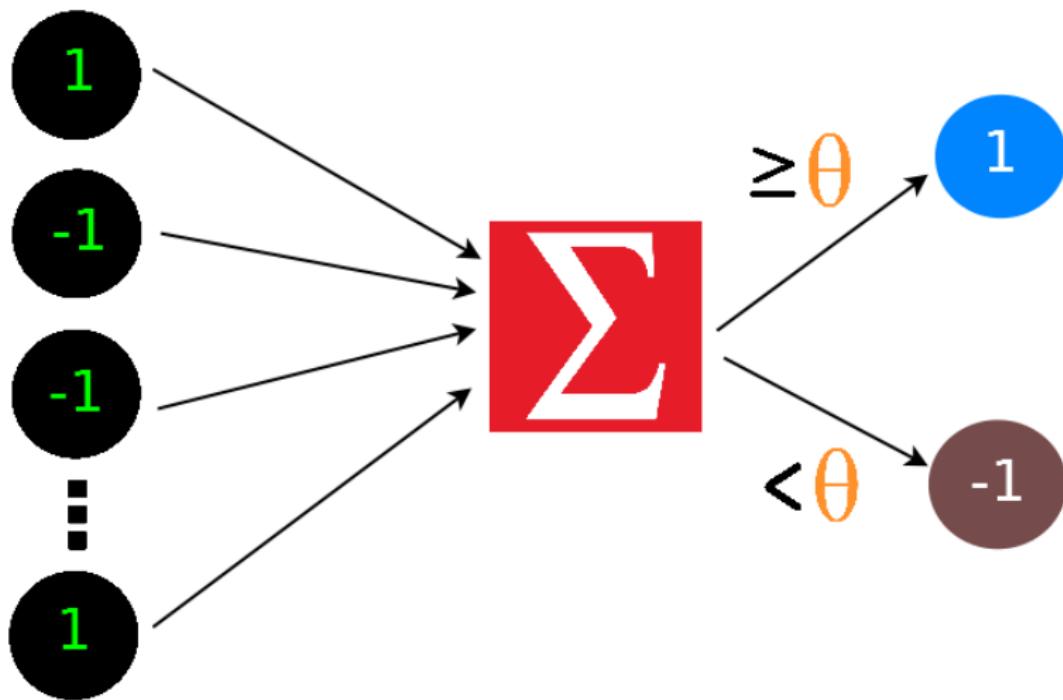
Perceptron



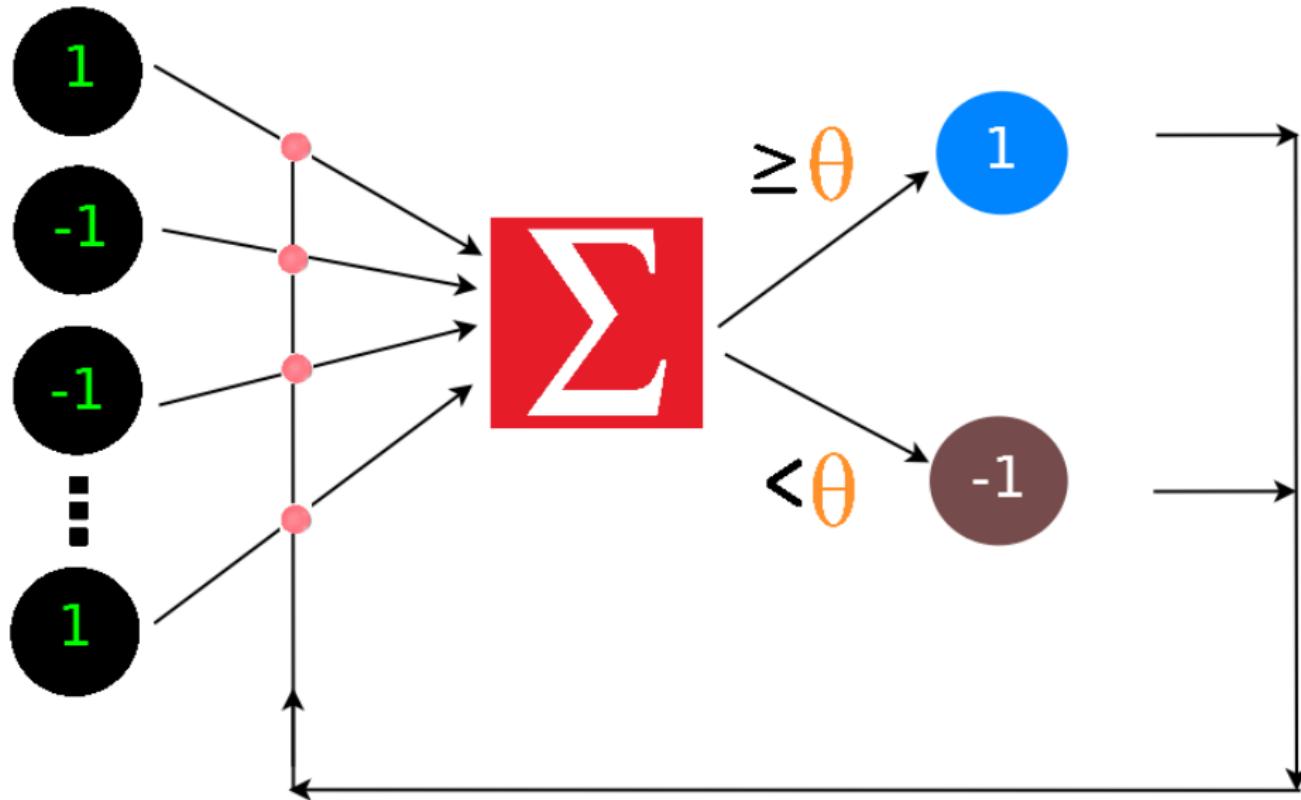
Perceptron



Perceptron



Perceptron



Nature of Machine Learning Tasks

Nature of Machine Learning Tasks

- Supervised Learning
 - ▶ Inputs and corresponding outputs are known during learning
 - ▶ e.g. Classification (Binary, Multi-class, Multi-label)
- Unsupervised Learning
 - ▶ Input objects are not generally labeled
 - ▶ e.g. Clustering, Principal-component Analysis
- Semi-supervised Learning
 - ▶ learning from a few labeled data
- Transfer Learning
 - ▶ transferring a learned model from task T_1 to T_2
 - ▶ e.g. transfer from image captioning to video captioning

Supervised Machine Learning

Binary Classification

Recall: e-mail Spam Classification

Genuine

Spam

Can't see the pictures? Select "Always deliver images" or view the message in your browser.

sent this message to
Your registered name is listed below to show this message
as originated from you.
[Learn more](#)

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Near IT Park, commonfloor

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Binary Classification

- **Input:** e-mail messages
- **Output:** Spam/Not spam

Binary Classification

- **Input:** e-mail messages \Rightarrow some feature space
- **Output:** Spam/Not spam $\Rightarrow \{+1, -1\}$
- Generally many input/output pairs are given for learning the machine learning model.

Binary Classification

Feature Extraction

W Dark matter - Wikipedia, the free encyclopedia en.wikipedia.org/w/index.php?title=Dark_matter&oldid=983000000

Galaxy clusters and gravitational lensing [edit]

Galaxy clusters are especially important for dark matter studies since their masses can be estimated in three independent ways:

- From the scatter in radial velocities of the galaxies within them (as in Zwicky's early observations, with much larger modern samples).
- From X-rays emitted by very hot gas within the clusters. The temperature and density of the gas can be estimated from the energy and flux of the X-rays, hence the gas pressure; assuming pressure and gravity balance, this enables the mass profile of the cluster to be derived. Many of the experiments of the Chandra X-ray Observatory use this technique to independently determine the mass of clusters. These observations generally indicate a ratio of baryonic to total mass approximately 12-15 percent, in reasonable agreement with the Planck spacecraft cosmic average of 15.5 - 16 percent.^[37]
- From their gravitational lensing effects on background objects, usually more distant galaxies. This is observed as "strong lensing" (multiple images) near the cluster core, and weak lensing (shape distortions) in the outer parts. Several large Hubble projects have used this method to measure cluster masses.

Generally these three methods are in reasonable agreement, that clusters contain much more matter than the visible galaxies and gas.

A gravitational lens is formed when the light from a more distant source (such as a quasar) is "bent" around a massive object (such as a cluster of galaxies) between the source object and the observer. The process is known as gravitational lensing.

The galaxy cluster Abell 2029 is composed of thousands of galaxies enveloped in a cloud of hot gas, and an amount of dark matter equivalent to more than 10^{14} Suns. At the center of this cluster is an enormous, elliptically shaped galaxy that is thought to have been formed from the mergers of many smaller galaxies.^[38] The measured orbital velocities of galaxies within galactic clusters have been found to be consistent with dark matter observations.

Another important tool for future dark matter observations is gravitational lensing. Lensing relies on the effects of general relativity to predict masses without relying on dynamics, and so is a completely independent means of measuring the dark matter. Strong lensing, the observed distortion of background galaxies into arcs when the light passes through a gravitational lens, has been observed around a few distant clusters including Abell 1689 (pictured right).^[39] By measuring the distortion geometry, the mass of the cluster causing the phenomena can be obtained. In the dozens of cases where this has been done, the mass-to-light ratios obtained correspond to the dynamic measurements of clusters.^[40]

Weak gravitational lensing looks at minute distortions of galaxies observed in vast galaxy surveys due to foreground objects through statistical analyses. By examining the apparent shear of the adjacent background galaxies, astrophysicists can characterize the mean distribution of dark matter by statistical means and have found mass-to-light ratios that correspond to dark matter predictions by other large-scale structure measurements.^[41] The correspondence of the two gravitational lens techniques to other dark matter measurements has convinced almost all astrophysicists that dark matter actually exists as a major component of the universe's composition.

The most direct observational evidence to date for dark matter is in a system known as the Bullet Cluster. In most regions of the universe, dark matter and visible material are found together,^[42] as expected because of their mutual gravitational attraction. In the Bullet Cluster, a collision between two galaxy clusters appears to have caused a separation of dark matter and baryonic matter. X-ray observations show that much of the baryonic matter (in the form of 10^7 - 10^8 Kelvin^[43] gas, or plasma) in the system is concentrated in the center of the system. Electromagnetic interactions between passing gas particles caused them to slow down and settle near the point of impact. However, weak gravitational lensing observations of the same system show that much of the mass resides outside of the central region of baryonic gas. Because dark matter does not interact by electromagnetic forces, it would not have been slowed in the same way as the X-ray visible gas, so the dark matter components of the two clusters passed through each other without slowing down substantially. This accounts for the separation. Unlike the galactic rotation curves, this evidence for dark matter is independent of the details of Newtonian gravity, so it is claimed to be direct evidence of the existence of dark matter.^[43]

Another galaxy cluster, known as the Train Wreck Cluster/Abell 520, appears to have an unusually massive and dark core containing few of the cluster's galaxies, which presents problems for standard dark matter models.^[44]

This may be explained by the dark core actually being a long, low-density dark matter filament (containing few galaxies) along the line of sight, projected onto the cluster core.^[45]

The observed behavior of dark matter in clusters constrains whether and how much dark matter scatters off other dark matter particles, quantified as its

Paraphasis

Capitalized First Letter

Numbers

Quote

Superscripts

Bigram

Trigram

Hyphenated

Name

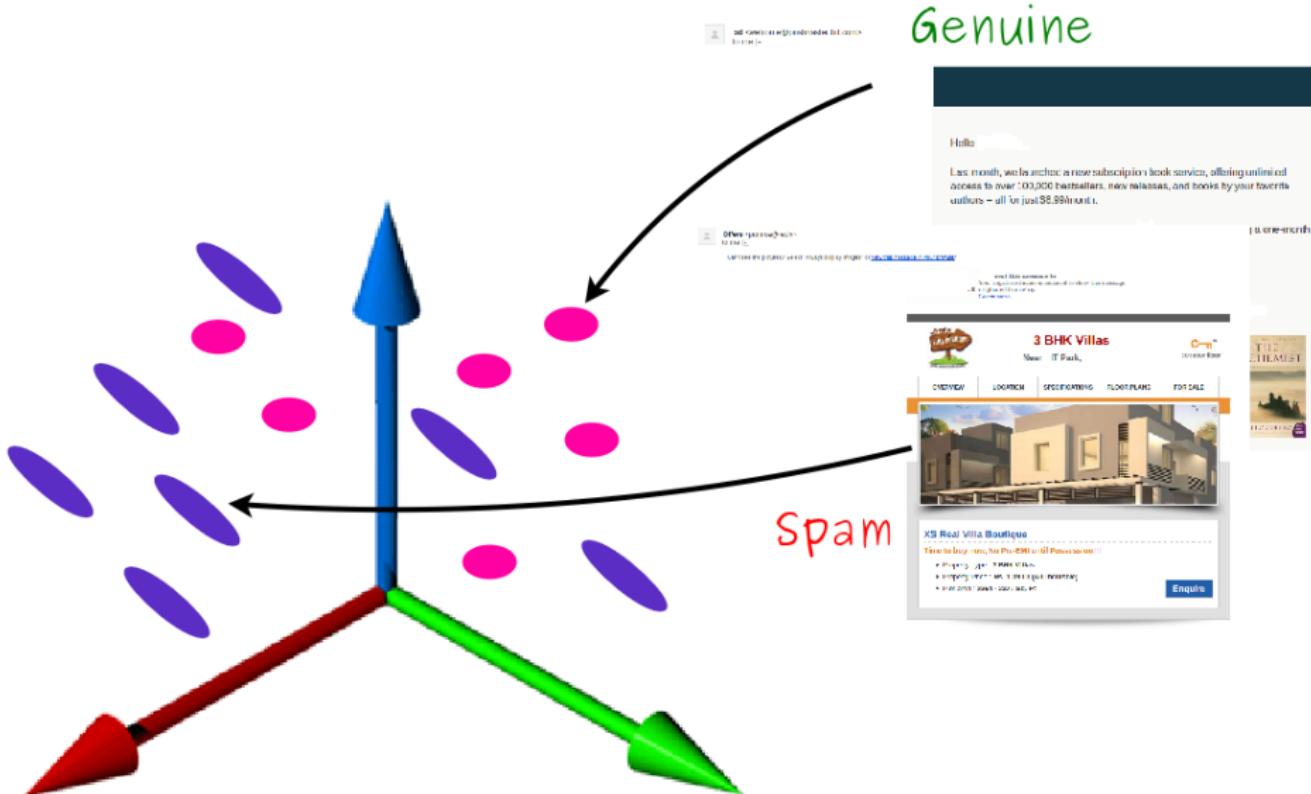







Binary Classification

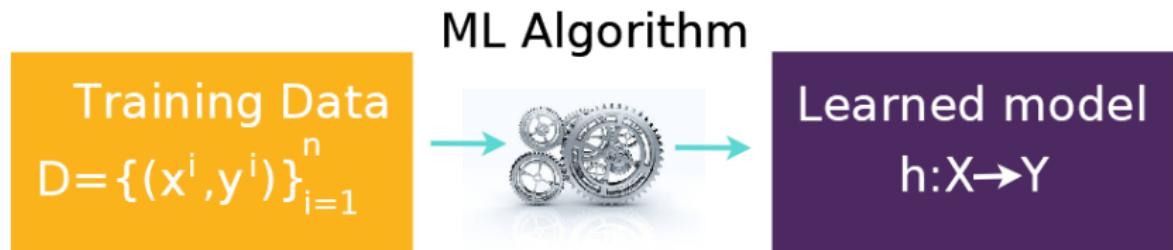
Feature Extraction



Binary Classification

- **Input:** e-mail messages \implies some feature space $\subseteq \mathbb{R}^d$.
 - ▶ $x \in \mathcal{X} \subseteq \mathbb{R}^d$
- **Output:** Spam/Not spam $\implies \{+1, -1\}$
 - ▶ $y \in \mathcal{Y} = \{+1, -1\}$
- Generally n input/output pairs $\{(x^i, y^i)\}_{i=1}^n \in (\mathcal{X} \times \mathcal{Y})^n$ are given for learning the machine learning model.
- $D = \{(x^i, y^i)\}_{i=1}^n$ called the training data.

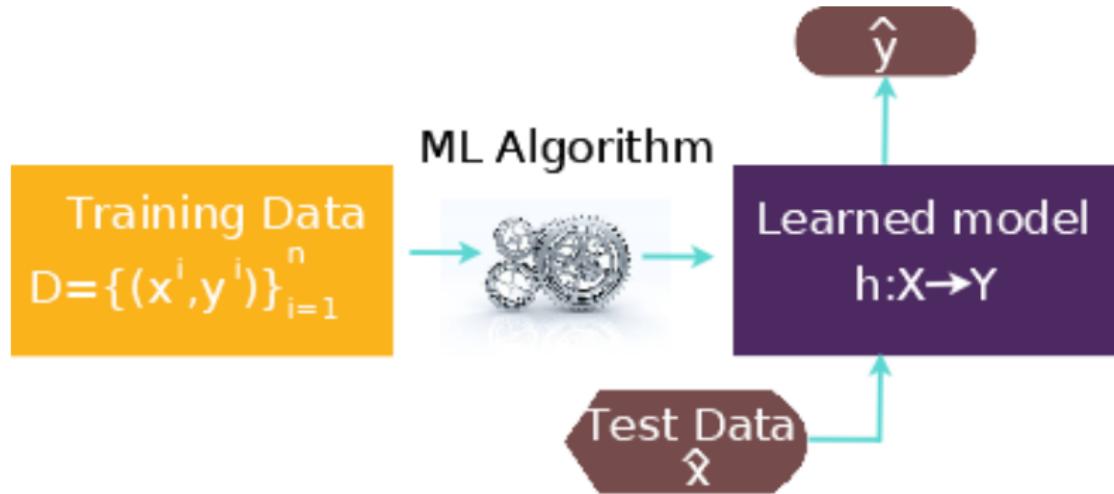
General Nature of a Supervised Machine Learning Task



Training

- **Input:** Training data $D = \{(x^i, y^i)\}_{i=1}^n$
- **Aim:** Learn a model $h : \mathcal{X} \rightarrow \mathcal{Y}$

General Nature of a Supervised Machine Learning Task



Training

- **Input:** Training data $D = \{(x^i, y^i)\}_{i=1}^n$
- **Aim:** Learn a model $h : \mathcal{X} \rightarrow \mathcal{Y}$

Testing

- Given \hat{x} , predict $\hat{y} = h(\hat{x})$

A sample data set for email classification

spam	to_multiple	from	cc	sent_email	image	attach	dollar	winner	inherit	viagra	password	num_char
0	0	1	0	0	0	0	0	no	0	0	0	11.37
0	0	1	0	0	0	0	0	no	0	0	0	10.504
0	0	1	0	0	0	0	0	no	0	0	0	13.256
0	0	1	0	0	0	0	0	no	0	0	2	1.231
1	0	1	0	0	0	0	0	no	0	0	0	5.108
1	0	1	4	0	0	2	0	no	0	0	0	0.631
1	0	1	0	0	0	0	1	yes	0	0	0	0.963
1	0	1	0	0	0	0	2	no	0	0	0	2.182

Representing image as vectors

