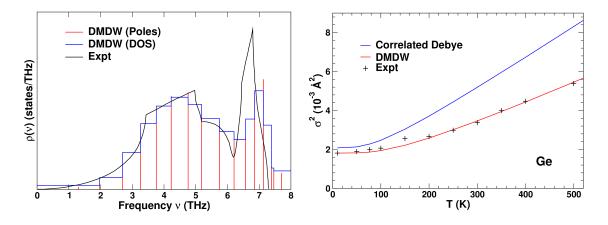
DMDW: A set of tools to calculate Debye-Waller factors and other vibrational properties using dynamical matrices

Introduction

DMDW is a set of tools developed to calculate vibrational properties such as EXAFS and crystallographic Debye-Waller (DW) factors, vibrational free energies, phonon densities of state and other related quantities from dynamical matrices (*i.e.* the matrix of force constants or Hessian matrix) using the Lanczos recursion algorithm.

DMDW provides an alternative to the correlated Einstein and Debye models. These semi-empirical models are unsatisfactory for several reasons:

- There are typically many more independent DW factors in the XAFS MS path expansion than can be fit reliably to the available data.
- They require separate fits to appropriate Debye or Einstein temperatures for each multiplescattering path.
- They typically ignore anisotropic contributions, and hence do not capture the detailed structure of the phonon spectra and associated DW factors.



Left: Dynamical matrix (in pole and continuous representations) and experimental phonon density of states. Right: Correlated Debye, dynamical matrix and experimental near neighbor DW factor in Ge.

A detailed description of the methodology can be found in:

"Theoretical x-ray absorption Debye-Waller factors", F. D. Vila, J. J. Rehr, H. H. Rossner and H. J. Krappe, Phys. Rev. B **76**, 014301 (2007), DOI: 10.1103/PhysRevB.76.014301.

"Recursion method for multiple-scattering XAFS Debye-Waller factors", A. V. Poiarkova and J. J. Rehr, J. Synchrotron Rad. **6**, 313 (1999) DOI: 10.1107/S0909049599001685.

How to use DMDW

Depending on the objectives and expertise of the user, DMDW can be used in different ways:

<u>FEFF integrated version:</u> FEFF users are referred to the DEBYE card section in the FEFF manual [LINK] for detailed instructions on how to use the fully integrated dmdw module. This module takes a dynamical matrix file and calculates DW factors for either EXAFS or XANES. In addition, for users with no experience in the calculation of dynamical matrices, JFEFF, the FEFF GUI offers a transparent way to interface with Cloud Computing versions of FEFF and ABINIT that automate the process.

<u>AI2PS version:</u> DMDW relies on dynamical matrices generated by other codes. Thus, to ease the learning curve required to obtain the dynamical matrices we have developed AI2PS, a program that partially automates the process. Users with little expertise should refer to the AI2PS project [LINK].

<u>Standalone version:</u> This page focuses on the capabilities of DMDW after a dynamical matrix is available. The User's Guide [LINK] explains the input, output and data file formats in detail, while the tutorials [LINK] included are meant to provide a quick start guide.

How to obtain DMDW

DMDW comes integrated in both the FEFF [LINK] and AI2PS [LINK] distributions. Please refer to their pages for instructions on how to install. The most current standalone version can be downloaded from here [LINK].

How to install

Please refer to the User Guide [LINK] for instructions on how to compile DMDW.

Tutorials

The tutorials input, output and dynamical matrix files are located in the "tutorials" directory of the distribution. These tutorials assume that the dynamical matrix files have been generated. The generation of these files is beyond the scope of these tutorials. The users should refer to the dynamical matrix generation guide. [LINK]

Case 1: Near-neighbor EXAFS DW factor of Cu between 1 and 300K

The input file (dmdw.inp) for this case is:

```
1 # Output level: 0 - Terse, 1 - Verbose
10 # Lanczos recursion order
20 1.0 300.0 # Temperature grid: Number of temps, T min, T max
0 # Type of output: 0 - s^2, 1 - VFE, 2 - SE, 3 - u^2
feff.dym # Dynamical matrix file name
1 # Number of paths descriptors
2 1 2 30.0 # Single scatt. descriptor from 1 to 2, ignore if R Path>30 au
```

The first line set the the verbosity of the output. Currently only two levels are available, with 0 reporting only the target quantity (σ^2 , u^2 , etc) and 1 producing other interesting quantities (see below). The second line sets the Lanczos recursion order. In this case 10 is more than sufficient to produce converged EXAFS DW factors. Users should always monitor convergence with respect to this parameter. Please refer to the User's Guide for a more detailed discussion of the lower and upper recursion order limits. The third line determines the target temperatures. Both single temperatures and ranges can be computed. The fourth line sets the type of DMDW output. (σ^2 and u^2 calculations are fully implemented, while vibrational free energies and self-energies are still under development.) The next line defines the name of the dynamical matrix file, in .dym format (please refer to the User's Guide for a detailed description of this format). The last two lines define the EXAFS paths for which the DW factors will be computed. In this case there is only one type of path, a single scattering path between the near-neighbor atoms 1 and 2. A path length cutoff is also included to ignore paths longer than 30.0 au.

The output file (dmdw.out) looks like this:

```
# Lanczos recursion order:
# Temperature: (See list Below)
# Dynamical matrix file: feff.dym
______
Path Indices: 1 2
pDOS Poles:
              Weight
   Freq. (THz)
      2.165
                0.00126
      2.888
                0.03260
      3.630
                0.11056
      4.340
                0.13929
      5.025
                0.13509
      5.704
                0.10975
      6.380
                0.10676
      6.805
                0.21940
      7.118
                0.06292
      7.332
                0.08235
pDOS Einstein freq (single pole) and associated temp:
Freq (THz) Temp (K)
  5.784
           277.60
pDOS n Moments and associated Einstein freqs and temps:
      Mom (THz^n) Freq (THz) Temp (K)
 n
                 5.07524
-2
      0.03882
                              243.57
                   5.27433
-1
       0.18960
                              253.12
       0.99998
 Ω
                   0.00000
                               0.00
 1
       5.63323
                   5.63323
                             270.34
                   5.78435
       33.45866
                             277.60
Path Length (Ang): 2.532
Temp (K) s^2 (1e-3 Ang^2)
         3.016
  1.00
            3.016
  16.74
            3.029
  32.47
  48.21
            3.091
            3.218
  63.95
  79.68
            3.404
            3.636
  95.42
           3.904
 111.16
```

. . .

The main portions of the output are the table of poles and weights determined by the Lanczos algorithm, the moments of the projected phonon density of states (pDOS) from both the single-pole and multi-pole approximations, their associated Einstein frequencies and temperatures and, finally, the table of DW factors at each target temperature. The poles and weights can be used to plot the pDOS (as in the figure above, which corresponds to the total, rather than projected pDOS, see below) while the moments can be used in simplified models and provide interesting properties of the systems, like its Einstein temperature.

Case 2: All single and some select double scattering EXAFS DW factors from one atom for Ge at 300K

As in the first case, the input file is very simple:

```
0  # Output level: 0 - Terse, 1 - Verbose
10  # Lanczos recursion order
1 300.0  # Temperature grid: Number of temps, T min, T max
0  # Type of output: 0 - s^2, 1 - VFE, 2 - SE, 3 - u^2
feff.dym  # Dynamical matrix file name
2  # Number of paths descriptors
2 1 0  30.0  # All single scatt. descriptor from 1, ignore if R_Path>30 au
3 1 4 0  30.0  # Double scatt. descriptor from 1, 4, ignore if R_Path>30 au
```

The main difference in this case is in the use of a terse output, just a single temperature and targeting single scattering parts from atoms 1, and double scattering paths that pass through atoms 1, 4 and any other atom in the system.

The output only lists the path indices, length and associated DW factors:

```
# Lanczos recursion order: 10
# Temperature: 300.00
# Dynamical matrix file: feff.dym

Path Indices: 1 2
Path Length (Ang), s^2 (1e-3 Ang^2): 2.417 3.448

Path Indices: 1 3
Path Length (Ang), s^2 (1e-3 Ang^2): 2.417 3.448
```

Case 3: Crystallographic u² for Ge at 300K

The main difference in this case is in the type of DMDW calculation, and that the descriptor now targets a single atom, rather than an EXAFS scattering path:

```
1  # Output level: 0 - Terse, 1 - Verbose
16  # Lanczos recursion order
1 300.0  # Temperature grid: Number of temps, T min, T max
3  # Type of output: 0 - s^2, 1 - VFE, 2 - SE, 3 - u^2
feff.dym  # Dynamical matrix file name
1  # Number of paths descriptors
1 1  0.0 # Target atom 1
```

The Lanczos order has been increased to provide better results since the u^2 calculations have slower convergence than the σ^2 . (Please refer to the User's Guide for a more detailed discussion of the lower and upper recursion order limits.)

The output file looks like this:

```
# Lanczos recursion order:
# Temperature: 300.00
# Dynamical matrix file: feff.dym
______
Atom Index: 1
----- Direction x ------
pDOS Poles:
   Freq. (THz)
               Weight
      1.787
              0.07747
      2.205
               0.10172
      2.627
              0.12653
      3.211
               0.02538
      3.931
              0.02408
              0.03918
      4.650
               0.02747
      5.325
      6.117
               0.03222
                0.08839
      6.575
      7.051
               0.02895
               0.03463
      7.455
               0.06450
      7.850
      8.100
               0.14295
      8.366
               0.11729
      8.458
               0.01356
               0.05567
      8.635
pDOS Einstein freq (single pole) and associated temp:
Freq (THz) Temp (K)
  6.241
         299.51
pDOS n Moments and associated Einstein freqs and temps:
 n Mom (THz^n) Freq (THz) Temp (K)
-2
      0.08025 3.53009
                             169.41
                            199.43
-1
       0.24064
                   4.15556
       1.00000
 0
                   0.00000
                              0.00
       5.68136
                            272.65
 1
                   5.68136
                  6.24108
                            299.51
      38.95106
u^2 (1e-3 Ang^2): 7.166
----- Direction y ------
pDOS Poles:
    Freq. (THz)
              Weight
      1.792 0.06711
               0.10530
      2.228
              0.13148
      2.629
      3.217
                0.02493
      3.926
                0.02336
      4.660
                0.04435
               0.02809
      5.318
               0.03694
      6.098
      6.557
               0.08771
      7.015
               0.02365
      7.436
               0.03234
             0.05568
0.14464
0.09064
      7.848
      8.104
      8.356
```

```
8.424 0.04703
8.641 0.05674
 pDOS Einstein freq (single pole) and associated temp:
 Freq (THz) Temp (K)
                 299.51
   6.241
 pDOS n Moments and associated Einstein freqs and temps:
       Mom (THz^n) Freq (THz) Temp (K)
          0.07803 3.57984 171.80

0.23833 4.19591 201.37

0.99999 0.00000 0.00

5.68823 5.68823 272.98

38.95074 6.24105 299.51
 -2
 -1
  0
  1
                                6.24105
u^2 (1e-3 Ang^2): 6.973
----- Direction z -----
 pDOS Poles:
          eq. (THz) Weight
1.786 0.05930
2.251 0.10504
2.642 0.13850
3.242 0.02409
3.937 0.02334
4.678 0.04897
5.318 0.02858
6.098 0.04633
6.567 0.08597
7.114 0.02174
7.511 0.02893
7.924 0.06717
8.143 0.13487
8.396 0.12810
8.620 0.04108
      Freq. (THz) Weight
           8.658
                            0.04108
 pDOS Einstein freq (single pole) and associated temp:
 Freq (THz) Temp (K)
   6.241
                  299.51
 pDOS n Moments and associated Einstein freqs and temps:
 n Mom (THz^n) Freq (THz) Temp (K)
          0.07612 3.62463 173.95

0.23623 4.23317 203.15

0.99999 0.00000 0.00

5.69498 5.69498 273.31

38.95122 6.24109 299.52
 -2
 -1
 0
u^2 (1e-3 Ang^2): 6.806
```

The output is very similar to the σ^2 case, but with three sections corresponding to displacement along the three Cartesian axes. The total pDOS can be obtained by collecting all poles and renormalizing the weights by 1/3. The total u^2 corresponds to the average of the three Cartesian components.

Contact information

For help please contact: fdv@uw.edu