

卫星摄影测量(Satellite Photogrammetry)

地形无关的控制方案 建立RPC模型

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旋转矩阵 (Euler Angle)

rotation matrix to transform **column 3-vectors** from one cartesian coordinate system to another. Final system is formed by rotating original system about its own axis by angle ϕ (counterclockwise as viewed from the +axis direction):

$$\begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos\omega & -\sin\omega \\ 0 & \sin\omega & \cos\omega \end{bmatrix} \begin{bmatrix} X_\omega \\ Y_\omega \\ Z_\omega \end{bmatrix} \quad \begin{bmatrix} X_\omega \\ Y_\omega \\ Z_\omega \end{bmatrix} = R_1(\omega) \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} \quad R_1(\omega) = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos\omega & \sin\omega \\ 0 & -\sin\omega & \cos\omega \end{bmatrix}$$

$$\begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} \cos\phi & 0 & \sin\phi \\ 0 & 1 & 0 \\ -\sin\phi & 0 & \cos\phi \end{bmatrix} \begin{bmatrix} X_\phi \\ Y_\phi \\ Z_\phi \end{bmatrix} \quad \begin{bmatrix} X_\phi \\ Y_\phi \\ Z_\phi \end{bmatrix} = R_2(\phi) \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} \quad R_2(\phi) = \begin{bmatrix} \cos\phi & 0 & -\sin\phi \\ 0 & 1 & 0 \\ \sin\phi & 0 & \cos\phi \end{bmatrix}$$

$$\begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = \begin{bmatrix} \cos\kappa & -\sin\kappa & 0 \\ \sin\kappa & \cos\kappa & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} X_\kappa \\ Y_\kappa \\ Z_\kappa \end{bmatrix} \quad \begin{bmatrix} X_\kappa \\ Y_\kappa \\ Z_\kappa \end{bmatrix} = R_3(\kappa) \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} \quad R_3(\kappa) = \begin{bmatrix} \cos\kappa & \sin\kappa & 0 \\ -\sin\kappa & \cos\kappa & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

四元数 (Quaternion) 描述姿态

空间任意一个旋转由一个 **旋转轴** 和一个 **转角** 进行描述

旋转轴：矢量表示 $\overrightarrow{OA} = [X_A, Y_A, Z_A]$

旋 转：旋转角 θ

分别定义四个元素：

$$x = c \bullet X_A \quad y = c \bullet Y_A$$

$$z = c \bullet Z_A \quad w = s$$

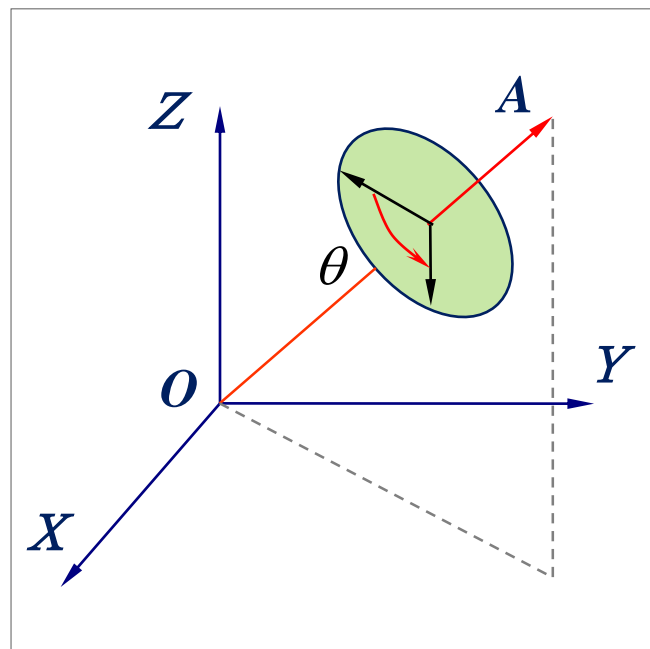
$$c = \sin(\theta/2) \quad s = \cos(\theta/2)$$

$$Q = [x, \quad y, \quad z, \quad w]$$

$$Q = w + xi + yj + zk$$

$$Q = [q_1 \quad q_2 \quad q_3 \quad q_4]$$

称为为四元数(Quaternion)



$$N(Q) = x^2 + y^2 + z^2 + w^2 = 1$$

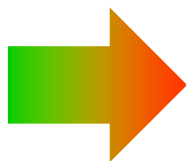
单位四元数

旋转矩阵 (Quaternion)

$$\mathbf{R} = \begin{bmatrix} 1 - 2(y^2 + z^2) & 2(xy - wz) & 2(xz + wy) \\ 2(xy + wz) & 1 - 2(x^2 + z^2) & 2(yz - wx) \\ 2(xz - wy) & 2(yz + wx) & 1 - 2(x^2 + y^2) \end{bmatrix}$$

$$\mathbf{R}_E = \mathbf{R}_\varphi \mathbf{R}_\omega \mathbf{R}_\kappa = \begin{bmatrix} \mathbf{a}_1 & \mathbf{a}_2 & \mathbf{a}_3 \\ \mathbf{b}_1 & \mathbf{b}_2 & \mathbf{b}_3 \\ \mathbf{c}_1 & \mathbf{c}_2 & \mathbf{c}_3 \end{bmatrix}$$

$$\mathbf{R}_E = \mathbf{R}$$



$$\varphi = -\arctg\left(\frac{\mathbf{a}_3}{\mathbf{c}_3}\right)$$

$$\omega = -\arcsin(\mathbf{b}_3)$$

$$\kappa = -\arctg\left(\frac{\mathbf{b}_1}{\mathbf{b}_2}\right)$$

四元数插值

三维单位矢量定义了一个球状上点。单位四元数则定义了一个四维超球面上的一个点。沿着超球状上两点之间的大弧进行插值就可得到平滑轨迹

$$t_0 \quad q_0 = [x_0 \quad y_0 \quad z_0 \quad w_0]$$

$$t \quad q_t = [x_t \quad y_t \quad z_t \quad w_t]$$

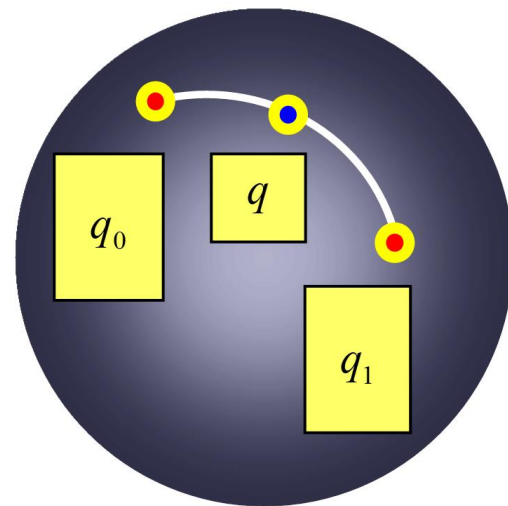
$$t_1 \quad q_1 = [x_1 \quad y_1 \quad z_1 \quad w_1]$$

$$q_t = \eta_0(t)q_0 + \eta_1(t)q_1$$

$$\eta_0 = \frac{\sin(\theta(t_1 - t)/(t_1 - t_0))}{\sin(\theta)}$$

$$\eta_1 = \frac{\sin(\theta(t - t_0)/(t_1 - t_0))}{\sin(\theta)}$$

$$\cos(\theta) = q_0 q_1$$



四元数球面线性内插

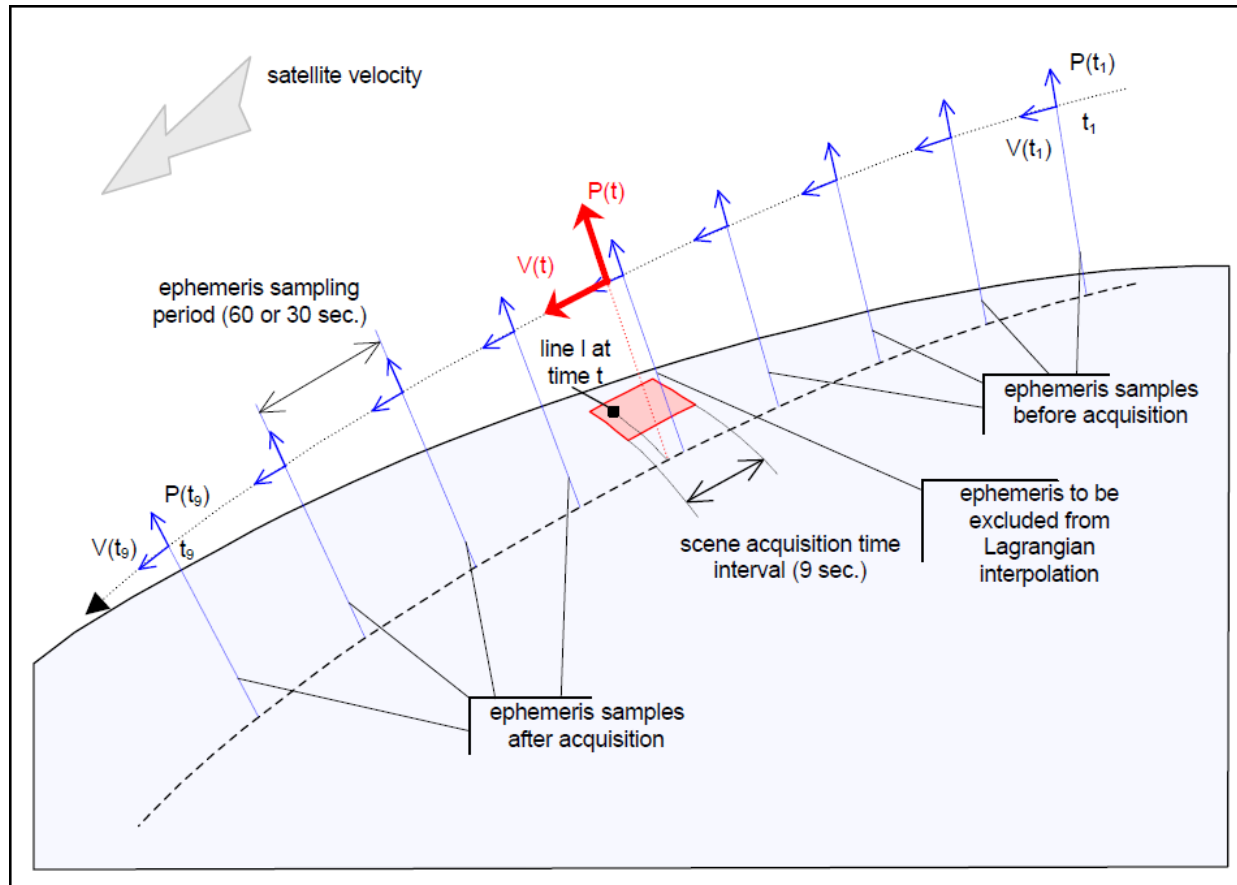
$|q_0 q_1| \rightarrow 1 \Rightarrow$ 退化为线性内插

$q_0 q_1 < 0 \Rightarrow \theta > \frac{\pi}{2}$, q_0 或 q_1 取反
使得插值的角距离最小化

姿态数据

```
attData_00 =  
{  
    timeCode = 97499270.0736999959 ;//卫星累积秒，从2009-01-01-0:0:0开始计  
    dateTime = "2012 02 03 11:07:50.073700" ;           //格林尼治时间  
    roll = -0.00001245 ;           // roll/pitch/yaw 卫星本体到轨道坐标系姿态  
    pitch = -0.00000509 ;  
    yaw = 0.03356850 ;  
    roll_velocity = 0.00000000 ;  
    pitch_velocity = 0.00000000 ;  
    yaw_velocity = 0.00000000 ;  
    q1 = -0.76259744 ; // q1/q2/q3/q4 卫星本体到J2000坐标系姿态；  
    q2 = 0.58472162 ;  
    q3 = 0.25788319 ;  
    q4 = 0.23561892 ;  
}
```

星历内插 (GPS)



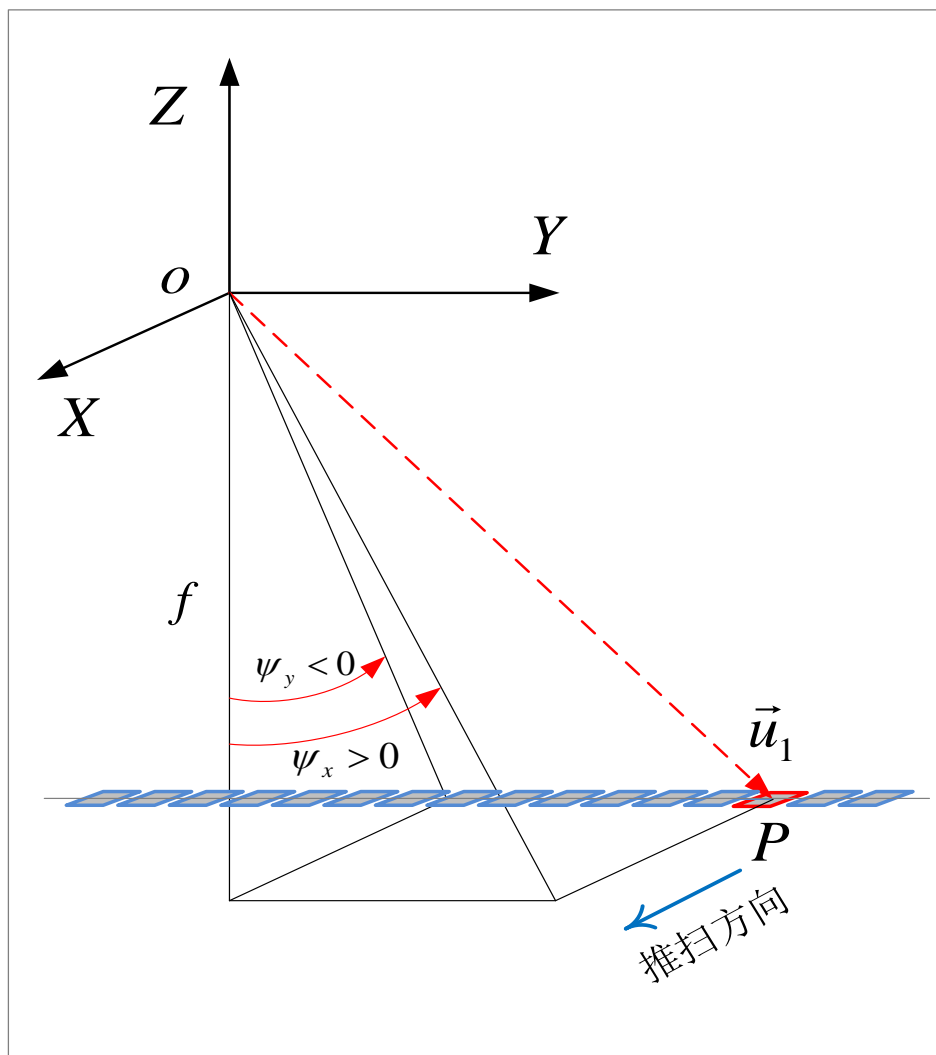
$$\vec{P}(t) = \sum_{j=1}^8 \frac{\vec{P}(t_j) \times \prod_{\substack{i=1 \\ i \neq j}}^8 (t - t_i)}{\prod_{\substack{i=1 \\ i \neq j}}^8 (t_j - t_i)}$$

$$\vec{V}(t) = \sum_{j=1}^8 \frac{\vec{V}(t_j) \times \prod_{\substack{i=1 \\ i \neq j}}^8 (t - t_i)}{\prod_{\substack{i=1 \\ i \neq j}}^8 (t_j - t_i)}$$

轨道数据

```
coordinateType = WGS84;//GPS测量卫星位置、速度矢量所在坐标系
dataType = GPS;
groupNumber = 542 ;      //测量记录总数
gpsData_00 =
{
    timeCode = 97499268.0000000000 ;//卫星累积秒，从2009-01-01-0:0:0开始计
    dateTime = “2012 02 03 11:07:48.000000” ; //格林尼治时间
    PX = -1912991.5000000000 ;    //卫星位置 (X,Y,Z) ， 单位米
    PY = 3092157.1000000001 ;
    PZ = 5828414.6000000006 ;
    VX = -1475.4500000000 ;      //卫星速度 (VX,VY,VZ) ， 单位米/秒
    VY = 6456.2100000000 ;
    VZ = -3905.8700000000 ;
}
```


相机坐标系\像元指向角



$$\frac{x - x_0 - \Delta x}{f} = \tan \psi_y$$

$$\frac{y - y_0 - \Delta y}{f} = \tan \psi_x$$

Look direction Angles

ψ_x / ψ_y

焦平面上像元的指向角
描述像元的位置

SPOT/IKONOS/ZY3等

相机校正文件

像元ID	垂轨指向角	沿轨指向角
0.000000	0.050668614594336	-0.000791106890569
1.000000	0.050664505955175	-0.000791109449488
2.000000	0.050660397314306	-0.000791112008175
3.000000	0.050656288671731	-0.000791114566631
4.000000	0.050652180027448	-0.000791117124855
5.000000	0.050648071381458	-0.000791119682847
6.000000	0.050643962733761	-0.000791122240609
7.000000	0.050639854084357	-0.000791124798139
8.000000	0.050635745433247	-0.000791127355437
9.000000	0.050631636780430	-0.000791129912504
10.000000	0.050627528125907	-0.000791132469341
11.000000	0.050623419469678	-0.000791135025946
.....		

$$\psi(p) = \psi(i) + (\psi(p_{i+1}) - \psi(p_i)) \times \frac{p - p_i}{p_{i+1} - p_i} \quad (p_i < p < p_{i+1})$$

偏置矩阵文件

starttime = 0.000000000000000000

pitch = -0.000511776876952

Vpitch = -0.0000000000125194

roll = 0.001828916699906

Vroll = 0.0000000000892101

yaw = 0.003770429577750

Vyaw = 0.000000000000000000

$$R = \begin{bmatrix} \cos\varphi & 0 & \sin\varphi \\ 0 & 1 & 0 \\ -\sin\varphi & 0 & \cos\varphi \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos\omega & -\sin\omega \\ 0 & \sin\omega & \cos\omega \end{bmatrix} \begin{bmatrix} \cos\kappa & -\sin\kappa & 0 \\ \sin\kappa & \cos\kappa & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

时间标签文件

RelLine	Time	deltaTime
0	131862405.0003719300000000000000	131862405.0003719300000000000000
1	131862405.0007438700000000000000	0.00037193298339843750
2	131862405.0011158000000000000000	0.00037193298339843750
3	131862405.0014877300000000000000	0.00037193298339843750
4	131862405.0018596600000000000000	0.00037193298339843750
5	131862405.0022316000000000000000	0.00037193298339843750
6	131862405.0026035300000000000000	0.00037193298339843750
7	131862405.0029754600000000000000	0.00037193298339843750
8	131862405.0033474000000000000000	0.00037193298339843750
9	131862405.0037193300000000000000	0.00037193298339843750
10	131862405.0040912600000000000000	0.00037193298339843750
11	131862405.0044632000000000000000	0.00037193298339843750
12	131862405.0048351300000000000000	0.00037193298339843750
.....		

相机严格模型

$$\begin{bmatrix} X \\ Y \\ Z \end{bmatrix}_{WGS84} = \begin{bmatrix} X_{GPS} \\ Y_{GPS} \\ Z_{GPS} \end{bmatrix} + R_{J2000}^{WGS84}(t) R_{star}^{J2000} (R_{star}^{body})^T \left[\begin{bmatrix} D_x \\ D_y \\ D_z \end{bmatrix} + \begin{bmatrix} d_x \\ d_y \\ d_z \end{bmatrix} + \lambda \cdot R_{camera}^{body} \begin{bmatrix} \tan(\psi_y) \\ \tan(\psi_x) \\ -1 \end{bmatrix} \cdot f \right]$$



$$\begin{bmatrix} X \\ Y \\ Z \end{bmatrix}_{WGS84} = \begin{bmatrix} X_{GPS} \\ Y_{GPS} \\ Z_{GPS} \end{bmatrix} + \lambda \cdot R_{J2000}^{WGS84}(t) \cdot R_{body}^{J2000} \cdot R_u \cdot R_{camera}^{body} \begin{bmatrix} \tan(\psi_y) \\ \tan(\psi_x) \\ -1 \end{bmatrix} \cdot f$$



$$R_u$$

http://hpiers.obspm.fr/eop-pc/index.php?index=matrice_php&lang=en



-
- EOP TIME SERIES ▾
- FTP products
- Reference C04 series
Each day since 1962
- Reference C01 series
Each 0.05 year from 1846
- EOP series & analysis
- EOP series & comparison
- Bulletins B, C, D
- Last days for EOP
- Rotation matrix/vector
- WEB Service
- EOP series: synoptic
- THEORY AND MODELLING ▾
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- LINKS ▾

Compute the Earth rotation matrix

updated on December 2007

Include combined EOP C04 (1962 - today + 6 months prediction) :

Celestial pole offsets ☒ UT1 - UTC ☒ Polar motion ☒

Real time
stop ☐ yes ☐

☐ Include diurnal and semidiurnal variations produced by ocean tides in polar motion and UT1 (IERS conventions 2000)

Submit request

Civil date : year month day
UTC : h min s

2020 4 17 15 h 36 min 22 s UTC

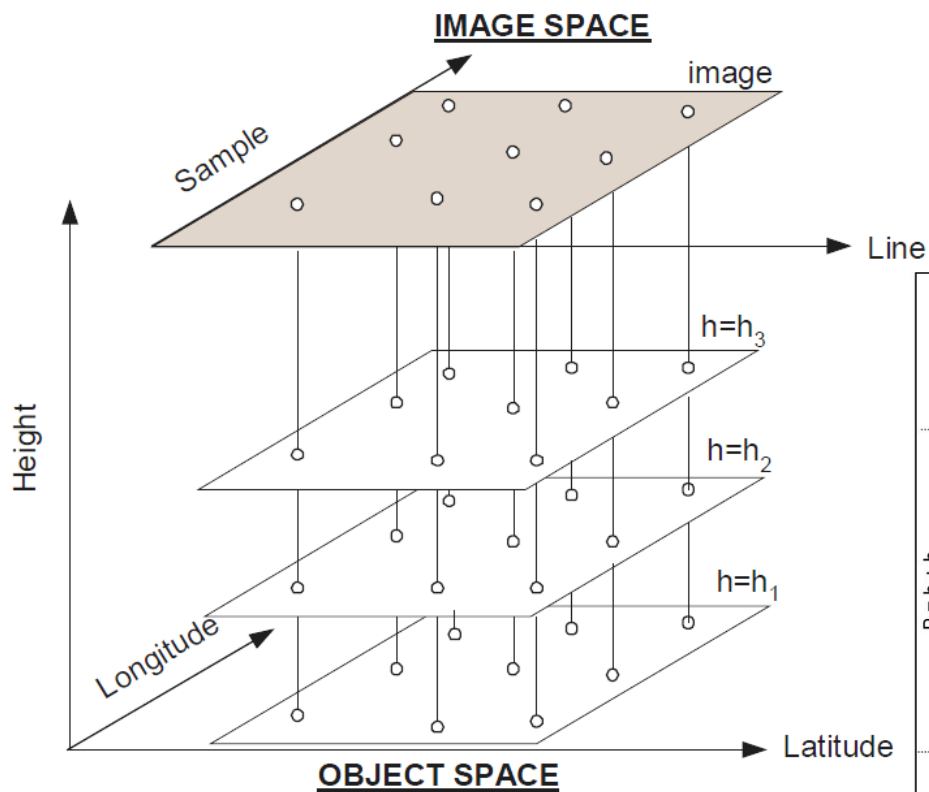
matrice:

0.172639552573	-0.984983163469	0.001937155147
0.984985012993	0.172639868400	-0.000004241506
-0.000330252398	0.001908801039	0.999998123704

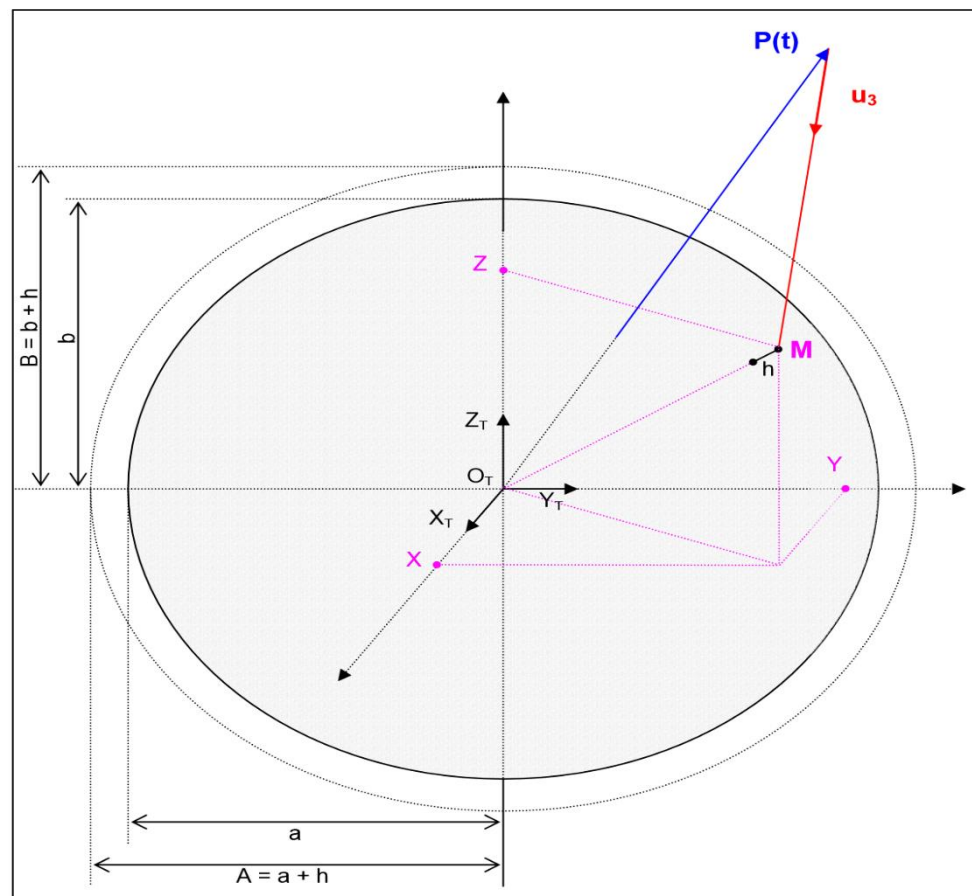
Transformation coordinate M from the international terrestrial reference system (ITRF) to the international celestial reference frame (ICRF) : **Celestial coordinates (X Y Z) = M x Terrestrial coordinates (x y z)**

- from 1962 to the current week the matrix can include the EOP of the IERS combined series C04 (time resolution for EOP fluctuations is about 6 days), as well as diurnal and semidiurnal variations in UT1 and polar motion produced by oceanic tides (IERS 2000 conventions). Accuracy is 0.0001", corresponding to $5 \cdot 10^{-10}$ rad.

Intersection of look direction with the Earth ellipsoid



Getting the look direction u_3 from the satellite at the position $P(t)$, we now compute the intersection with the ellipsoid located at an altitude h above the standard ITRF ellipsoid (see fig. below).



Intersection of look direction with the Earth ellipsoid

Let $M = (X, Y, Z)$ the geocentric coordinates to be found, point M is involved within the two following equations:

$$\overrightarrow{O_3M} = \vec{P}(t) + \mu \times \vec{u}_3 \Rightarrow \begin{cases} X = X_P + \mu \times (u_3)_X \\ Y = Y_P + \mu \times (u_3)_Y \\ Z = Z_P + \mu \times (u_3)_Z \end{cases} \quad Eq. 1$$

$$\frac{X^2 + Y^2}{A^2} + \frac{Z^2}{B^2} = 1 \quad \text{with} \quad \begin{cases} A = a + h \\ B = b + h \end{cases} \quad Eq. 2$$

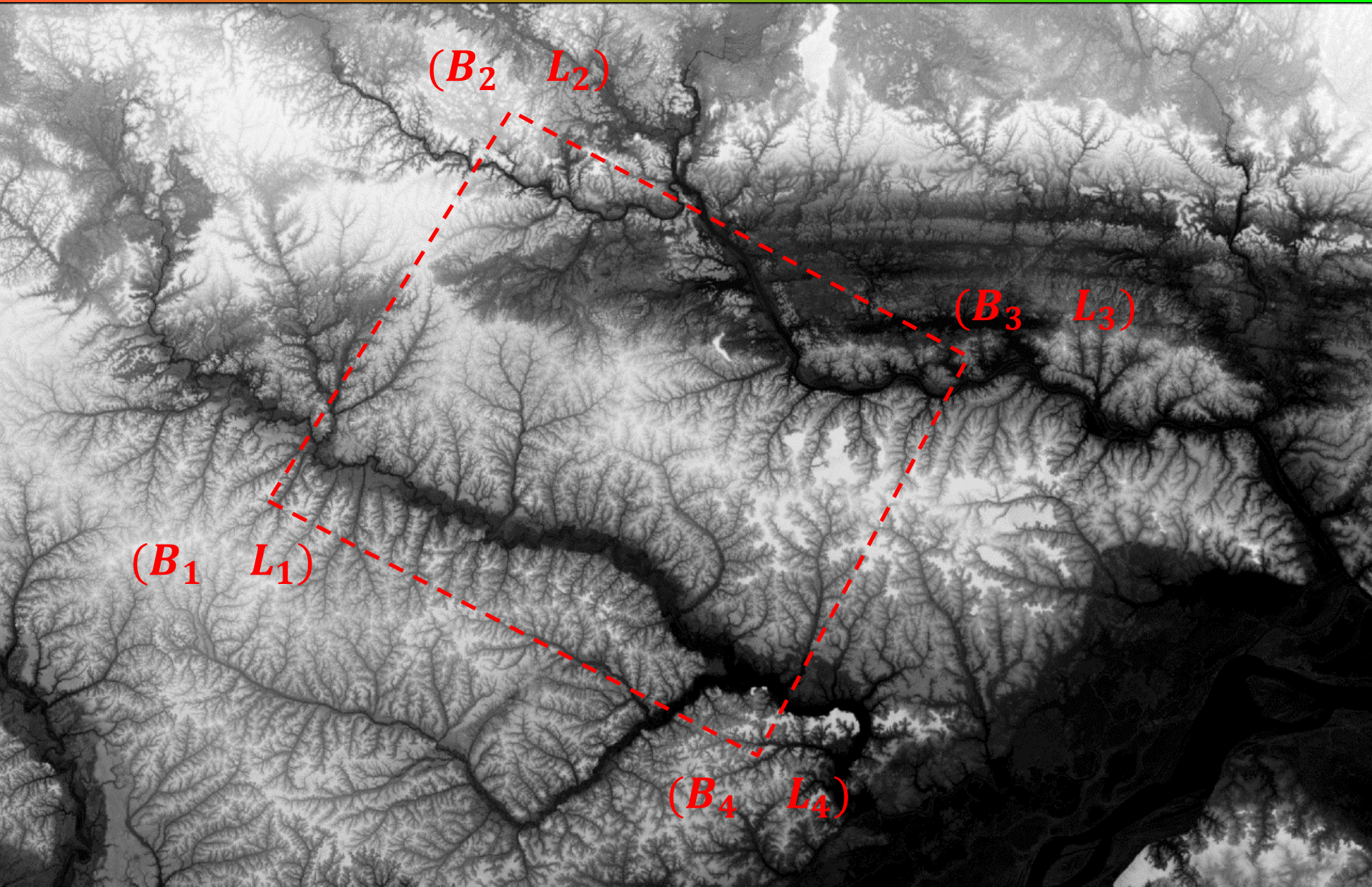
Leading to solve the 2nd degree equation:

$$\left[\frac{(u_3)_X^2}{A^2} + \frac{(u_3)_Y^2}{B^2} + \frac{(u_3)_Z^2}{B^2} \right] \times \mu^2 + 2 \times \left[\frac{X_P(u_3)_X}{A^2} + \frac{Y_P(u_3)_Y}{B^2} + \frac{Z_P(u_3)_Z}{B^2} \right] \times \mu + \left[\frac{X_P^2}{A^2} + \frac{Y_P^2}{B^2} + \frac{Z_P^2}{B^2} \right] = 1$$

This equation has necessarily two distinct solutions (μ_1, μ_2). The smallest one (μ_{\min}) shall be kept. Re-introducing this value within equation 1 gives the geocentric coordinates (X,Y,Z) of point M. The geodetic coordinates (λ, ϕ, h) may be computed.

视向量 \vec{u}_3 与椭球面求交算法

SRTM-DEM



物方到像方投影变换

- ◆ 实现最佳扫描线搜索
- ◆ 基本思想

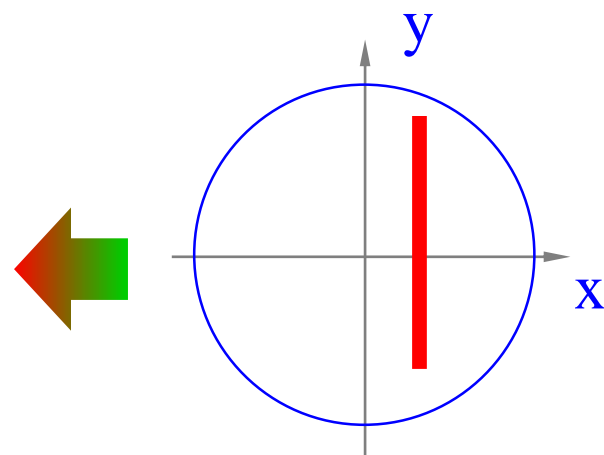
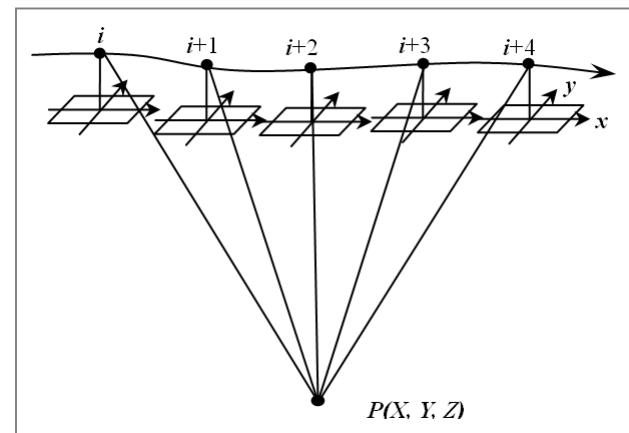
扫描线搜索是一个迭代过程；

选取扫描线对应的 6 EO 计算像点坐标；

比较 **计算的坐标值** 与焦平面上像元 **标定坐标**？

判断是否一致？

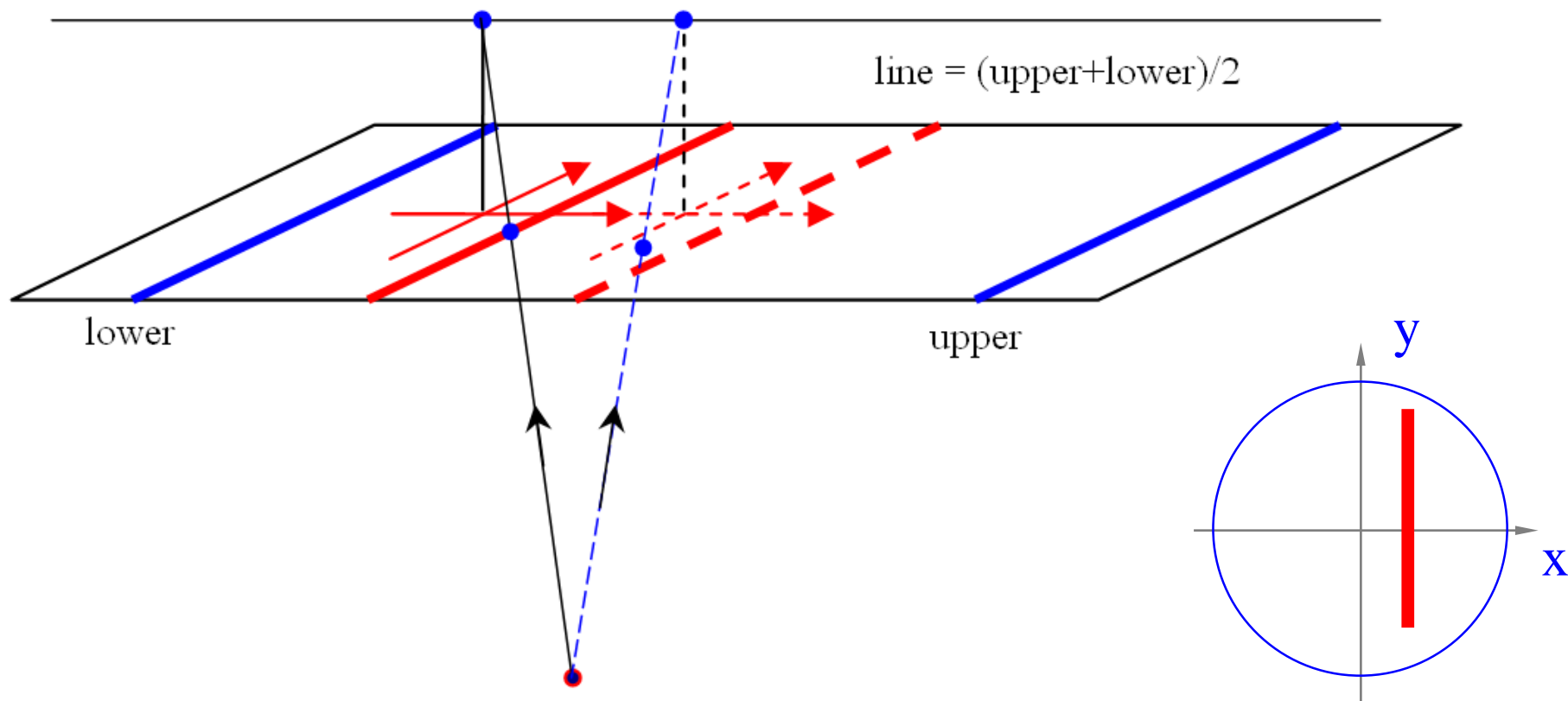
$$x = -f \frac{a_1(X - X_s) + b_1(Y - Y_s) + c_1(Z - Z_s)}{a_3(X - X_s) + b_3(Y - Y_s) + c_3(Z - Z_s)}$$
$$y = -f \frac{a_2(X - X_s) + b_2(Y - Y_s) + c_2(Z - Z_s)}{a_3(X - X_s) + b_3(Y - Y_s) + c_3(Z - Z_s)}$$



经典的像方搜索采用顺序搜索的策略

窗口二分法

迭代时，将搜索窗口二等分；分别将搜索窗口的中间、首或尾扫描行的6EO代入共线方程；基于像方搜索判据判断并取窗口的二分之一作为新的搜索窗口，快速、有效地缩小搜索范围，提高了算法效率。



1、最佳扫描行和搜索窗口的迭代初始值

设影像由 N 条扫描行构成。

$L=N/2$ ，初始搜索窗口取为 $[1,N]$ ， $L_S=1$ ， $L_E=N$ 。

2、计算像点坐标

根据 L_S 、 L 、 L_E ，计算像点坐标 x_S 、 x 、 x_E 。

3、计算 $x_S \cdot x$ 及 $x \cdot x_E$ ，并判断其符号

如果 $x_S \cdot x \leq 0$ ，则二分 L_S 及 L 之间的窗口；

如果 $x \cdot x_E \leq 0$ ，则二分 L 及 L_E 之间的窗口；

4、判断搜索窗口大小 $|L_S - L_E|$

如果 $|L_S - L_E| \leq \delta$ ，则进行下一步；否则转2步

5、搜索窗口 $[L_S, L_E]$ 内确定最佳扫描行

搜索窗口内，依次取扫描线计算像点坐标，与焦平面上 CCD 线阵距离最近的像点对应最佳扫描线。

RPC模型

$$\left\{ \begin{array}{l} r_n = \frac{r - r_0}{r_s} \\ c_n = \frac{c - c_0}{c_s} \\ X_n = \frac{X - X_0}{X_s} \\ Y_n = \frac{Y - Y_0}{Y_s} \\ Z_n = \frac{Z - Z_0}{Z_s} \end{array} \right\}$$

$$\begin{aligned} r_n &= \frac{p_1(X_n, Y_n, Z_n)}{p_2(X_n, Y_n, Z_n)} \\ c_n &= \frac{p_3(X_n, Y_n, Z_n)}{p_4(X_n, Y_n, Z_n)} \end{aligned}$$

$$\left\{ \begin{array}{l} X_0 = \frac{\sum X}{m} \\ Y_0 = \frac{\sum Y}{m} \\ Z_0 = \frac{\sum Z}{m} \end{array} \right\} \quad \left\{ \begin{array}{l} r_0 = \frac{\sum r}{m} \\ c_0 = \frac{\sum c}{m} \end{array} \right.$$

$$\left\{ \begin{array}{l} X_s = \max(|X_{\max} - X_0|, |X_{\min} - X_0|) \\ Y_s = \max(|Y_{\max} - Y_0|, |Y_{\min} - Y_0|) \\ Z_s = \max(|Z_{\max} - Z_0|, |Z_{\min} - Z_0|) \end{array} \right.$$

$$\left\{ \begin{array}{l} r_s = \max(|r_{\max} - r_0|, |r_{\min} - r_0|) \\ c_s = \max(|c_{\max} - c_0|, |c_{\min} - c_0|) \end{array} \right.$$

多项式形式

多项式形式 关于坐标的3次多项式 $(i+j+k) \leq 3$

$$\begin{aligned} p &= \sum_{i=0}^{m_1} \sum_{j=0}^{m_2} \sum_{k=0}^{m_3} a_{ijk} X^i Y^j Z^k \\ &= a_0 + a_1 Z + a_2 Y + a_3 X + \\ &\quad a_4 ZY + a_5 ZX + a_6 YX + a_7 Z^2 + a_8 Y^2 + a_9 X^2 + \\ &\quad a_{10} ZYX + a_{11} Z^2 Y + a_{12} Z^2 X + a_{13} Y^2 Z + a_{14} Y^2 X + \\ &\quad a_{15} ZX^2 + a_{16} YX^2 + a_{17} Z^3 + a_{18} Y^3 + a_{19} X^3 \end{aligned}$$

$$a_i (i = 0, 1, \dots, 19)$$

多项式系数

a_{ijk} are polynomial coefficients called
rational function coefficients (RFCS)

线性化

$$r_n = \frac{p_1(X_n, Y_n, Z_n)}{p_2(X_n, Y_n, Z_n)} \quad B = (1 \quad Z \quad Y \quad X \quad \dots \quad Y^3 \quad X^3) \cdot (1 \quad b_1 \quad \dots \quad b_{19})^T$$

$$c_n = \frac{p_3(X_n, Y_n, Z_n)}{p_4(X_n, Y_n, Z_n)} \quad J = (a_0 \quad a_1 \quad \dots \quad a_{19} \quad b_1 \quad b_2 \quad \dots \quad b_{19})^T$$

$$D = (1 \quad Z \quad Y \quad X \quad \dots \quad Y^3 \quad X^3) \cdot (1 \quad d_1 \quad \dots \quad d_{19})^T$$

$$K = (c_0 \quad c_1 \quad \dots \quad c_{19} \quad d_1 \quad d_2 \quad \dots \quad d_{19})^T$$

$$v_r = \left[\frac{1}{B} \quad \frac{Z}{B} \quad \frac{Y}{B} \quad \frac{X}{B} \quad \dots \quad \frac{Y^3}{B} \quad \frac{X^3}{B} \quad \frac{-rZ}{B} \quad \frac{-rY}{B} \quad \dots \quad \frac{-rY^3}{B} \quad \frac{-rX^3}{B} \right] \cdot J - \frac{r}{B}$$

$$v_c = \left[\frac{1}{D} \quad \frac{Z}{D} \quad \frac{Y}{D} \quad \frac{X}{D} \quad \dots \quad \frac{Y^3}{D} \quad \frac{X^3}{D} \quad \frac{-cZ}{D} \quad \frac{-cY}{D} \quad \dots \quad \frac{-cY^3}{D} \quad \frac{-cX^3}{D} \right] \cdot K - \frac{c}{D}$$

误差方程

假定有n个地面控制点及对应像点

$$\begin{bmatrix} v_{r1} \\ v_{r2} \\ \vdots \\ v_{rn} \end{bmatrix} = \begin{bmatrix} \frac{1}{B_1} & 0 & \cdots & 0 \\ 0 & \frac{1}{B_2} & 0 & \vdots \\ \vdots & 0 & \ddots & 0 \\ 0 & \cdots & 0 & \frac{1}{B_n} \end{bmatrix} \cdot \begin{bmatrix} 1 & Z_1 & \cdots & X_1^3 & -r_1 Z_1 & \cdots & -r_1 X_1^3 \\ 1 & Z_2 & \cdots & X_2^3 & -r_2 Z_2 & \cdots & -r_2 X_2^3 \\ \vdots & \vdots & \ddots & \vdots & \vdots & \ddots & \vdots \\ 1 & Z_n & \cdots & X_n^3 & -r_n Z_n & \cdots & -r_n X_n^3 \end{bmatrix} \cdot J$$

$$- \begin{bmatrix} \frac{1}{B_1} & 0 & \cdots & 0 \\ 0 & \frac{1}{B_2} & 0 & \vdots \\ \vdots & 0 & \ddots & 0 \\ 0 & \cdots & 0 & \frac{1}{B_n} \end{bmatrix} \cdot \begin{bmatrix} r_1 \\ r_2 \\ \vdots \\ r_n \end{bmatrix}$$

最小二乘解算

写成矩阵形式

$$V_r = W_r M J - W_r R$$

$$M = \begin{bmatrix} 1 & Z_1 & \cdots & X_1^3 & -r_1 Z_1 & \cdots & -r_1 X_1^3 \\ 1 & Z_2 & \cdots & X_2^3 & -r_2 Z_2 & \cdots & -r_2 X_2^3 \\ \vdots & \vdots & \ddots & \vdots & \vdots & \ddots & \vdots \\ 1 & Z_n & \cdots & X_n^3 & -r_n Z_n & \cdots & -r_n X_n^3 \end{bmatrix} \quad W_r = \begin{bmatrix} \frac{1}{B_1} & 0 & \cdots & 0 \\ 0 & \frac{1}{B_2} & 0 & \vdots \\ \vdots & 0 & \ddots & 0 \\ 0 & \cdots & 0 & \frac{1}{B_n} \end{bmatrix} \quad R = \begin{bmatrix} r_1 \\ r_2 \\ \vdots \\ r_n \end{bmatrix}$$

法方程形式

$$M^T W_r^2 M \cdot J = M^T W_r^2 R$$