**G0.0+0.0** Sgr A East

**RA:**  $17^{\text{h}}45^{\text{m}}44^{\text{s}}$  **1-GHz flux/Jy:** 100? **Size/arcmin:**  $3.5 \times 2.5$  **Dec:**  $-29^{\circ}00'$  **Spectral index:** 0.8? **Type:** S

Radio: Non-thermal shell, in complex region, interacting with molecular material to the west.

**X-ray:** Diffuse emission, centrally peaked. **Point sources:** Compact X-ray/radio source.

#### References:

Ekers *et al.* 1983, A&A, 122, 143. VLA at 1.4 GHz and 5 GHz (both  $5'' \times 8''$ ),  $S_{1.4~\rm GHz} = 77~\rm Jy$ ,  $S_{5.0~\rm GHz} = 31~\rm Jy$ . Pedlar *et al.* 1989, ApJ, 342, 769. VLA at 332 MHz (12"),

1.4 GHz  $(1''.3 \times 2''.5)$  and 5 GHz  $(1''.3 \times 2''.5)$ .

Mezger et al. 1989, A&A, 209, 337. Nearby molecular material. Ho et al. 1991, Nature, 350, 309. VLA of NH<sub>3</sub> emission from surroundings.

Anantharamaiah *et al.* 1991, MNRAS, 249, 262. VLA at 330 MHz (17"×33").

Serabyn *et al.* 1992, ApJ, 395, 166. Nearby molecular material. Gray 1994, MNRAS, 270, 835. MOST at 843 MHz (43"×88"). Yusef-Zadeh & Mehringer 1995, ApJ, 452, L37. VLA of nearby H<sub>2</sub>O masers.

Yusef-Zadeh et al. 1999, ApJ, 512, 230. OH maser observations. Coil & Ho 2000, ApJ, 533, 245. NH<sub>3</sub> observations of surroundings.

LaRosa et al. 2000, AJ, 119, 207. VLA at 333 MHz (24"×43"). see also: LaRosa et al. 2000, AJ, 119, 3145. Erratum.

Fatuzzo et al. 2001, ApJ, 549, 293. Electron–positron lines from the vicinity.

Yusef-Zadeh *et al.* 2001, ApJ, 560, 749. Observations of nearby molecular hydrogen.

Maeda et al. 2002, ApJ, 570, 671. Chandra observations.Sakano et al. 2003, AN, 324 (No S1), 197. XMM-Newton observations.

Roy & Rao 2004, MNRAS, 349, L25. GMRT at 620 MHz (6.6×11.4).

Sakano et al. 2004, MNRAS, 350, 129. XMM-Newton observa-

Yusef-Zadeh *et al.* 2004, ApJS, 155, 421. VLA at 1.4 GHz (1"×2".2).

Park et al. 2005, ApJ, 631, 964. Chandra observations.
Aharonian et al. 2006, ApJ, 636, 777. HESS detection.
Koyama et al. 2007, PASJ, 59, S237. Suzaku observations.
Lee et al. 2008, ApJ, 674, 247. Molecular H<sub>2</sub> observations of surroundings.

Sjouwerman & Pihlström 2008, ApJ, 681, 1287. VLA at 1.7 GHz of OH masers.

Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers.

Tsuboi *et al.* 2012, PASJ, 64, 111. CO observations of SW. Minh *et al.* 2013, ApJ, 773, 31. NH<sub>3</sub> observations of region. Zhao *et al.* 2013, ApJ, 777, 146. VLA at 4.8 GHz (0".5 $\times$ 0".7), 5.5 GHz (0".6 $\times$ 1".6), and 8.3 GHz (1"0 $\times$ 2".0).

Nynka *et al.* 2013, ApJ, 778, L31. NuSTAR of compact X-ray source.

Pihlström et al. 2014, AJ, 147, 73. VLA observations of methanol masers

Ponti *et al.* 2015, MNRAS, 453, 172. XMM-Newton observations. Tsuboi *et al.* 2015, PASJ, 67, 109. CO observations of region. Lau *et al.* 2015, Science, 348, 413. IR observations.

Ajello et al. 2016, ApJ, 819, 44. Fermi observations.

Yusef-Zadeh *et al.* 2016, ApJ, 819, 60. VLA at 1.5 GHz 0".5×1".4) and other frequencies.

McEwen et al. 2016, ApJ, 832, 129. CH<sub>3</sub>OH maser observations in region.

Qiao et al. 2018, ApJS, 239, 15. OH maser observations.

G0.3+0.0

**RA:**  $17^{\text{h}}46^{\text{m}}15^{\text{s}}$  **1-GHz flux/Jy:** 22 **Size/arcmin:**  $15\times 8$  **Dec:**  $-28^{\circ}38'$  **Spectral index:** 0.6 **Type:** S

Has been called G0.33+0.04, G0.30+0.04 and G0.4+0.1.

Radio: Bilateral shell, near Galactic Centre.

X-ray: Diffuse emission.

### References:

Kassim & Frail 1996, MNRAS, 283, L51. VLA at 333 MHz  $(23''\times42'')$ , plus review of flux densities and other observations. LaRosa *et al.* 2000, AJ, 119, 207. VLA at 333 MHz (30''). *see also*: LaRosa *et al.* 2000, AJ, 119, 3145. Erratum.

Yusef-Zadeh *et al.* 2004, ApJS, 155, 421. VLA at 1.4 GHz (30"). Ponti *et al.* 2015, MNRAS, 453, 172. XMM-Newton observations. Ajello *et al.* 2016, ApJ, 819, 44. Fermi observations.

## G0.9 + 0.1

RA: 17<sup>h</sup>47<sup>m</sup>21<sup>s</sup>
1-GHz flux/Jy: 18?
Size/arcmin: 8
Dec: -28°09'
Spectral index: varies
Type: C

**Radio:** Flat spectrum core within steep spectrum shell. **X-ray:** Central core, with non-thermal spectrum.

Point sources: Central pulsar.

#### References:

Helfand & Becker 1987, ApJ, 314, 203. VLA at 1.4 and 5 GHz, and Einstein observations.

and Ellistein observations.

Gray 1994, MNRAS, 270, 835. MOST at 843 MHz (43"×91").

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

Mereghetti et al. 1998, A&A, 331, L77. X-ray detection.

LaRosa et al. 2000, AJ, 119, 207. VLA at 333 MHz (24"×43"). see also: LaRosa et al. 2000, AJ, 119, 3145. Erratum.

Sidoli et al. 2000, A&A, 361, 719. BeppoSAX observations.

Gaensler et al. 2001, ApJ, 556, L107. Chandra observations.

Porquet et al. 2003, A&A, 401, 197. XMM-Newton observations.

Yusef-Zadeh et al. 2004, ApJS, 155, 421. VLA at 1.4 GHz (8"6×11"3).

Aharonian *et al.* 2005, A&A, 432, L25. HESS detection. Dubner *et al.* 2008, A&A, 487, 1033. ATCA and VLA at 1.4 GHz  $(1.2 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 5 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 6 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 6 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 7 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 6 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 7 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 8 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ , 9 GHz  $(1.6 \times 2.5 : S = 8.3 \pm 0.7 \text{ Jy})$ 

Camilo et al. 2009, ApJ, 700, L34. Pulsar detection.Holler et al. 2012, A&A, 539, A24. Chandra and XMM-Newton observations.

Ponti *et al.* 2015, MNRAS, 453, 172. XMM-Newton observations. Archer *et al.* 2016, ApJ, 821, 129. γ-ray observations. Abdalla *et al.* 2018, A&A, 612, A1. HESS observations.

## G1.0 - 0.1

**RA:** 17<sup>h</sup>48<sup>m</sup>30<sup>s</sup> **1-GHz flux/Jy:** 15 **Size/arcmin:** 8 **Dec:** -28°09′ **Spectral index:** 0.6? **Type:** S

Has been called G1.05-0.1 and G1.05-0.15.

**Radio:** Incomplete shell, to the S of Sgr D.

X-ray: Possibly detected.

### References:

Downes *et al.* 1979, A&AS, 35, 1. Review of flux densities. Anantharamaiah *et al.* 1991, MNRAS, 249, 262. VLA at 330 MHz  $(64'' \times 100'' : S = 12.3 \text{ Jy})$ .

Liszt 1992, ApJS, 82, 495. VLA at 1.6 GHz (13"×23"). Gray 1994, MNRAS, 270, 835. MOST at 843 MHz (43"×91"). Mehringer *et al.* 1998, ApJ, 493, 274. VLA at 1.6 GHz (15"×28") and 5 GHz, including masers observations.

Yusef-Zadeh et al. 1999, ApJ, 527, 172. VLA of nearby OH masers

LaRosa *et al.* 2000, AJ, 119, 207. VLA at 333 MHz (24"×43"). *see also*: LaRosa *et al.* 2000, AJ, 119, 3145. Erratum.

Sidoli *et al.* 2001, A&A, 372, 651. BeppoSAX possible detection. Yusef-Zadeh *et al.* 2004, ApJS, 155, 421. VLA at 1.4 GHz (30"). Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH

Nobukawa et al. 2009, AdSpR, 43, 1045. Suzaku observations. Marquez-Lopez & Phillips 2010, MNRAS, 407, 94. Mid-IR observations.

Pihlström et al. 2014, AJ, 147, 73. VLA search for methanol masers.

Ponti et al. 2015, MNRAS, 453, 172. XMM-Newton observations. Froebrich et al. 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.

# G1.4 - 0.1

RA: 17<sup>h</sup>49<sup>m</sup>39<sup>s</sup> 1-GHz flux/Jy: 2? Size/arcmin: 10 Dec: -27°46′ Spectral index: ? Type: S

**Radio:** Shell, brightest in E. **X-ray:** Diffuse emission.

## References:

Gray 1994, MNRAS, 270, 847. MOST at 843 MHz  $(43'' \times 92'' : S = 2 \text{ Jy})$ .

Yusef-Zadeh et al. 1999, ApJ, 527, 172. VLA of nearby OH masers.

Bhatnagar 2002, MNRAS, 332, 1. GMRT at 327 MHz ( $2'.4 \times 2'.7$ :  $S=4.2\pm0.5$ ).

Yusef-Zadeh *et al.* 2004, ApJS, 155, 421. VLA at 1.4 GHz (8"2×12"2).

Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers.

Pihlström *et al.* 2014, AJ, 147, 73. VLA observations of methanol masers.

Ponti *et al.* 2015, MNRAS, 453, 172. XMM-Newton observations. Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. McEwen *et al.* 2016, ApJ, 826, 189. NH<sub>3</sub> and CH<sub>3</sub>OH observations.

## G1.9 + 0.3

**RA:** 17<sup>h</sup>48<sup>m</sup>45<sup>s</sup> **1-GHz flux/Jy:** 0.6 **Size/arcmin:** 1.5 **Dec:** -27°10′ **Spectral index:** 0.6 **Type:** S

Radio: Shell, brighter to the N, brightening; shows secular increase.

**X-ray:** Shell, with bright limbs to E and W. **Distance:** H<sub>I</sub> absorption gives < 10 kpc.

#### References:

Green & Gull 1984, Nature, 312, 527. VLA at 5 GHz (2"×4"4). Gray 1994, MNRAS, 270, 835. MOST at 843 MHz (43"×94"). Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Yusef-Zadeh *et al.* 2004, ApJS, 155, 421. VLA at 1.4 GHz (8"3×12").

Nord *et al.* 2004, AJ, 128, 1646. VLA at 330 MHz (7"×12"). Green 2004, BASI, 32, 335. VLA at 1.5 GHz (7".2×9".4). Reynolds *et al.* 2008, ApJ, 680, L41. Chandra observations. Green *et al.* 2008, MNRAS, 387, L54. VLA at 4.86 GHz (4"×10"), for expansion studies.

Murphy et al. 2008, MNRAS, 389, L23. MOST at 843 MHz for flux increase.

Gómez & Rodríguez 2009, RMxAA, 45, 91. VLA at 1.5 GHz (5".1×10".6).

Reynolds *et al.* 2009, ApJ, 695, L149. Chandra spectroscopy. Borkowski *et al.* 2010, ApJ, 724, L161. Chandra observations. Carlton *et al.* 2011, ApJ, 737, L22. Chandra expansion studies. Borkowski *et al.* 2013, ApJ, 771, L9. Chandra observations. Abramowski *et al.* 2014, MNRAS, 441, 790. HESS observations. Borkowski *et al.* 2014, ApJ, 790, L18. Chandra expansion studies. Roy & Pal 2014, IAUS, 296, 197. GMRT HI observations. De Horta *et al.* 2014, SerAJ, 189, 41. ATCA at 1.4 (5".4×10".4), 2.4 (2".9×6".1) and 5 GHz (1".2×2".8). Gök & Ergin 2015, AdSpR, 56, 1793. Suzaku and Fermi observa-

Zoglauer *et al.* 2015, ApJ, 798, 98. NuSTAR observations. Borkowski *et al.* 2017, ApJ, 837, L7. Chandra expansion studies.

## G3.7 - 0.2

**RA:**  $17^h55^m26^s$  **1-GHz flux/Jy:** 2.3 **Size/arcmin:**  $14\times11$  **Dec:**  $-25^{\circ}50'$  **Spectral index:** 0.65 **Type:** S

Has been called G003.8-00.3.

Radio: Double arc.

### References:

Gray 1994, MNRAS, 270, 847. MOST at 843 MHz  $(43'' \times 99'' : S = 2.4 \text{ Jy})$ .

Gaensler 1998, ApJ, 493, 781. VLA at 1.4 GHz ( $9'' \times 15'' : S = 1.7 \pm 0.1$  Jy).

Yusef-Zadeh *et al.* 2004, ApJS, 155, 421. VLA at 1.4 GHz (8".4×11".4).

# G3.8 + 0.3

**RA:** 17<sup>h</sup>52<sup>m</sup>55<sup>s</sup> **1-GHz flux/Jy:** 3? **Size/arcmin:** 18 **Dec:** -25°28′ **Spectral index:** 0.6 **Type:** S?

Radio: Incomplete shell.

### References:

Gray 1994, MNRAS, 270, 847. MOST at 843 MHz  $(43'' \times 100'' : S = 3.5 \text{ Jy})$ .

Bhatnagar 2002, MNRAS, 332, 1. GMRT at 327 MHz ( $17'' \times 27''$ :  $S = 6.0 \pm 0.4$ ).

# G4.2 - 3.5

**RA:** 18<sup>h</sup>08<sup>m</sup>55<sup>s</sup> **1-GHz flux/Jy:** 3.2? **Dec:** -27°03′ **Size/arcmin:** 28 **Type:** S

**Radio:** Elongated shell. **Optical:** Detected.

### References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4'.3).

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations.

**G4.5+6.8** Kepler, SN1604, 3C358

**RA:** 17<sup>h</sup>30<sup>m</sup>42<sup>s</sup> **1-GHz flux/Jy:** 19 **Size/arcmin:** 3 **Dec:** -21°29′ **Spectral index:** 0.64 **Type:** S

This is the remnant of Kepler's SN of AD1604.

**Radio:** Incomplete shell, brighter to the N.

**Optical:** Faint filaments. **X-ray:** Shell, brighter to the N.

**Distance:** Optical expansion and proper motion indicates about 2.9 kpc, H<sub>I</sub> observations suggest 3.4 to 6.4 kpc.

#### References:

halo.

van den Bergh & Kamper 1977, ApJ, 218, 617. Optical proper motions.

Leibowitz & Danziger 1983, MNRAS, 204, 273. Optical spectra. White & Long 1983, ApJ, 264, 196. Einstein observations. Matsui *et al.* 1984, ApJ, 287, 295. VLA at 1.4 (2".5×3".2) and 5 GHz (3".2×4".8) and Einstein image (5").

Dickel *et al.* 1988, ApJ, 330, 254. VLA at 1.4 (1"2×2"3) and 5 GHz (0".6×1".0) at two epochs.

Smith et al. 1989, ApJ, 347, 925. EXOSAT observations.
Hatsukade et al. 1990, PASJ, 42, 279. X-ray spectrum.
Blair et al. 1991, ApJ, 366, 484. Optical imaging and spectroscopy.

Bandiera & van den Bergh 1991, ApJ, 374, 186. Optical changes. van den Bergh 1991, PASP, 103, 194. Optical imaging. Predehl & Schmitt 1995, A&A, 293, 889. ROSAT of dust scattered

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations. Hughes 1999, ApJ, 527, 298. ROSAT and Einstein image comparison for expansion studies.

Reynoso & Goss 1999, AJ, 118, 926. VLA at 1.4 GHz  $(13'' \times 23'')$  for HI studies.

Kinugasa & Tsunemi 1999, PASJ, 51, 239. ASCA observations. Gerardy & Fesen 2001, AJ, 121, 2781. IR spectroscopy and imaging

DeLaney *et al.* 2002, ApJ, 580, 914. VLA at 1.3 to 1.5 GHz and 5 GHz (7".2) for spectral index studies.

Morgan et al. 2003, ApJ, 597, L33. Sub-mm dust observations. Sollerman et al. 2003, A&A, 407, 249. Optical spectroscopy.

Cassam-Chenaï et al. 2004, A&A, 414, 545. XMM-Newton observations.

Bamba et al. 2005, ApJ, 621, 793. Chandra observations of rim. Riesgo & López 2005, RMxAA, 41, 57. Optical observations of filament (previously classified as PN, H 2-12).

Blair et al. 2007, ApJ, 662, 998. Spitzer observations. Reynolds et al. 2007, ApJ, 668, L135. Chandra observations. Sankrit et al. 2008, AJ, 135, 538. HST observations. Aharonian et al. 2008, A&A, 488, 219. HESS upper limit. Enomoto et al. 2008, ApJ, 683, 383, Y-ray upper limit.

Katsuda *et al.* 2008, ApJ, 689, 225. Chandra proper motion studies.

Vink et al. 2008, ApJ, 689, 231. Chandra proper motion studies. Gomez et al. 2012, MNRAS, 420, 3557. Herschel IR dust observations.

Williams et al. 2012, ApJ, 755, 3. Spitzer spectroscopy.
Burkey et al. 2013, ApJ, 764, 63. Chandra observations.
Yang et al. 2013, ApJ, 766, 44. Suzaku spectroscopy.
Park et al. 2013, ApJ, 767, L10. Suzaku observations.
Katsuda et al. 2015, ApJ, 808, 49. XMM-Newton, Chandra and Suzaku observations.

Sankrit *et al.* 2016, ApJ, 817, 36. HST for proper motion studies. Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Sato & Hughes 2017, ApJ, 845, 167. Chandra expansion studies. Kasuga *et al.* 2018, PASJ, 70, 88. Chandra observations.

# G4.8+6.2

**RA:** 17<sup>h</sup>33<sup>m</sup>25<sup>s</sup> **1-GHz flux/Jy:** 3 **Size/arcmin:** 18 **Dec:** -21°34′ **Spectral index:** 0.6 **Type:** S

Has been called G4.5+6.2.

Radio: Faint shell.

### References:

Duncan et al. 1995, MNRAS, 277, 36. Parkes 64-m at 2.4-GHz (10'4).

Bhatnagar 2000, MNRAS, 317, 453. GMRT at 327 MHz (1' $3\times2'2:S=5.5\pm1.2$  Jy), and NVSS at 1.4 GHz.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

## G5.2 - 2.6

**RA:** 18<sup>h</sup>07<sup>m</sup>30<sup>s</sup> **1-GHz flux/Jy:** 2.6? **Size/arcmin:** 18 **Dec:** -25°45′ **Spectral index:** 0.6? **Type:** S

Radio: Poorly resolved shell.

#### References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4/3).

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

G5.4–1.2 Milne 56

**RA:** 18<sup>h</sup>02<sup>m</sup>10<sup>s</sup> **1-GHz flux/Jy:** 35? **Size/arcmin:** 35 **Dec:** -24°54′ **Spectral index:** 0.2? **Type:** C?

Part been called G5.3–1.0. Has been suggested that this is not a SNR.

Radio: Incomplete shell, including wide 'v' of emission to east with small flat-spectrum source at apex.

Optical: Detected.

**X-ray:** Pulsar detected, with faint extension.

**Point sources:** Pulsar nearby, in flat spectrum source.

**Distance:** HI absorption suggests > 4.3 kpc.

#### References:

Clark *et al.* 1975, AuJPA, 37, 75. Molonglo at 408 MHz (3': S = 38 Jy).

Becker & Kundu 1975, AJ, 80, 679. NRAO 140-ft at 10.6 GHz

Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8/4) and 5 GHz (4/4).

Angerhofer *et al.* 1977, A&A, 55, 11. NRAO 140-ft at 5 GHz (6'8:  $S = 21.9 \pm 2.4$  Jy).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'.6).

Zealey et al. 1979, A&AS, 38, 39. Optical detection.

Becker & Helfand 1985, Nature, 313, 115. VLA at 1.4 and 5 GHz.

Helfand & Becker 1985, Nature, 313, 118. Suggesting it is not a SNR.

Manchester *et al.* 1985, MNRAS, 212, 975. Pulsar detection. Caswell *et al.* 1987, MNRAS, 225, 329. MOST at 843 MHz (42"×110").

Frail & Kulkarni 1991, Nature, 352, 785. Pulsar and remnant association.

Manchester et al. 1991, MNRAS, 253, 7P. Pulsar and remnant association.

Milne *et al.* 1992, MNRAS, 255, 707. Parkes 64-m at 4.75 (4'.5: $S = 30.8 \pm 2.1$  Jy) and 8.4 GHz (3': $S = 24 \pm 3$  Jy), including polarisation

Kassim 1992, AJ, 103, 943. VLA at 327 MHz ( $3.0 \times 4.9 : S = 38 \text{ Jy}$ ).

Frail *et al.* 1994, AJ, 107, 1120. VLA at 327 MHz (68"×73"), plus H<sub>I</sub> absorption.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search.

Gaensler & Frail 2000, Nature, 406, 158. Pulsar observations, including proper motion.

Kaspi *et al.* 2001, ApJ, 562, L163. X-ray detection of pulsar, and upper limit for remnant.

Reich 2002, in NSPS, p1. Effelsberg 100-m at 10.6 GHz, including polarisation.

Blazek et al. 2006, ApJ, 652, 1523. Proper motion study of pulsar.Zeiger et al. 2008, ApJ, 674, 271. Proper motion study of pulsar.Liszt et al. 2009, A&A, 508, 1331. CO and IR observations of region.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser detection. Pihlström et al. 2014, AJ, 147, 73. VLA search for methanol masers.

# G5.5+0.3

**RA:**  $17^{\rm h}57^{\rm m}04^{\rm s}$  **1-GHz flux/Jy:** 5.5 **Size/arcmin:**  $15\times12$  **Dec:**  $-24^{\circ}00'$  **Spectral index:** 0.7 **Type:** S

Has been called G5.55+0.32.

Radio: Shell.
Optical: Detected.

### References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 14.3 \pm 0.3$  Jy), plus other observations.

Liszt et al. 2009, A&A, 508, 1331. CO and IR observations of region.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Stupar & Parker 2011, MNRAS, 414, 2282.  $\rm H\alpha$  observations. Froebrich *et al.* 2015, MNRAS, 454, 2586.  $\rm H_2$  IR observations.

# G5.9 + 3.1

**RA:** 17<sup>h</sup>47<sup>m</sup>20<sup>s</sup> **1-GHz flux/Jy:** 3.3? **Size/arcmin:** 20 **Dec:** -22°16′ **Spectral index:** 0.4? **Type:** S

Radio: Asymmetric shell.

#### References:

Reich *et al.* 1988, IAUCo, 101, 293. Effelsberg 100-m at 2.7 GHz (4/3).

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

# G6.1+0.5

Has been called G6.10+0.53.

Radio: Partial shell. Optical: Detected.

### References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 13.4 \pm 0.2$  Jy), plus other observations. Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations.

Froebrich et al. 2015, MNRAS, 454, 2586. H2 IR observations.

G6.1+1.2

RA:  $17^h54^m55^s$  1-GHz flux/Jy: 4.0? Size/arcmin:  $30\times26$  Dec:  $-23^\circ05'$  Spectral index: 0.3? Type: F

Has been called G6.1+1.15. **Radio:** Faint, diffuse emission.

## References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Junkes *et al.* 1988, LNP, 316, 134. Effelsberg 100-m at 2.7 GHz (4'.3), including polarisation.

Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4/3).

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

G6.4-0.1

RA: 18<sup>h</sup>00<sup>m</sup>30<sup>s</sup> 1-GHz flux/Jy: 310 Size/arcmin: 48 Dec: -23°26′ Spectral index: varies Type: C

Has been called G6.6–0.2.

**Radio:** Several non-thermal sources in a ring, with flat spectrum core.

**Optical:** Diffuse emission.

**X-ray:** Diffuse emission from most of the remnant.

**Point sources:** Young pulsar near edge of remnant, but not thought to be related.

**Distance:** HI observations suggest 1.9 kpc.

#### References:

Kundu & Velusamy 1972, A&A, 20, 237. NRAO 140-ft at 10 GHz (3').

van den Bergh *et al.* 1973, ApJS, 26, 19. Optical observations. Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8.4) and 5 GHz (4.4).

Dopita *et al.* 1977, ApJ, 214, 179. Some optical line ratios. Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'6)

Lozinskaya 1981, SvAL, 7, 17. Mean optical velocity.

Andrews *et al.* 1983, ApJ, 266, 684. VLA at 1.4 GHz (8") 4.9 GHz (3") and 15 GHz (1") of central region only, plus Einstein image of central region.

Bohigas *et al.* 1983, RMxAA, 8, 155. Optical spectra. Andrews *et al.* 1985, AJ, 90, 310. VLA of central component. Long *et al.* 1991, ApJ, 373, 567. Einstein and optical observations. Kassim 1992, AJ, 103, 943. VLA at 327 MHz (3′.1×4′.4 : *S* = 660 Jy).

Kaspi *et al.* 1993, ApJ, 409, L57. Pulsar association. Frail *et al.* 1993, Nature, 365, 136. VLA at 327 MHz (smoothed to 65"), plus pulsar association.

Frail et al. 1994, ApJ, 424, L111. VLA of associated OH masers. Frail et al. 1996, AJ, 111, 1651. OH maser emission.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Claussen *et al.* 1997, ApJ, 489, 143. VLA of OH masers.

Claussen *et al.* 1999, ApJ, 522, 349. High resolution observations of OH masers.

Yusef-Zadeh et al. 2000, ApJ, 540, 842. VLA at 327 MHz (2.6 $\times$ 5.5) and 1.48 GHz (40" $\times$ 65").

Dubner *et al.* 2000, AJ, 120, 1933. VLA at 328 MHz ( $52'' \times 97''$ :  $S = 425 \pm 40$  Jy) and 1415 MHz ( $48'' \times 88''$ :  $S = 246 \pm 20$  Jy), and comparison with other observations.

Reach & Rho 2000, ApJ, 544, 843. ISO observations of interactions with surroundings.

see also: Reach & Rho 2001, ApJ, 558, 943. Erratum. Douvion et al. 2001, A&A, 373, 281. ISO observations. Roberts et al. 2001, ApJS, 133, 451. ASCA observations. Velázquez et al. 2002, AJ, 124, 2145. Parkes 64-m at 1.4 GHz (15<sup>t</sup>) for HI.

Rho & Borkowski 2002, ApJ, 575, 201. ROSAT and ASCA observations.

Claussen et al. 2002, ApJ, 580, 909. Observations of nearby source.

Yusef-Zadeh *et al.* 2003, ApJ, 583, 267. OH observations. Caswell 2004, MNRAS, 349, 99. ATCA at 1.7 GHz of associated OH masers. Mavromatakis et al. 2004, A&A, 426, 567. Optical observations.Reach et al. 2005, ApJ, 618, 297. Molecular lines and near IR observations.

Hoffman *et al.* 2005, ApJ, 620, 257. OH maser observations. Kawasaki *et al.* 2005, ApJ, 631, 935. ASCA observations. Neufeld *et al.* 2007, ApJ, 664, 890. Spitzer observations. Aharonian *et al.* 2008, A&A, 481, 401. HESS observations. Casandjian & Grenier 2008, A&A, 489, 849. γ-ray observations. Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH

Giuliani *et al.* 2010, A&A, 516, L11. γ-ray observations. Marquez-Lopez & Phillips 2010, MNRAS, 407, 94. Mid-IR observations

vations.
Abdo *et al.* 2010, ApJ, 718, 348. Fermi observations.
Yuan & Neufeld 2011, ApJ, 726, 76. Spitzer observations.

Sawada et al. 2012, PASJ, 64, 81. Suzaku observations.

Nichols *et al.* 2012, MNRAS, 419, 251. CO observations of selected regions.

Gusdorf et al. 2012, A&A, 542, L19. CO observations of regions in NE.

Vaupreé et al. 2014, A&A, 568, A50. CO, HCO<sup>+</sup> and DCO<sup>+</sup> molecular line observations.

Pihlström *et al.* 2014, AJ, 147, 73. VLA observations of methanol masers.

Neufeld *et al.* 2014, ApJ, 781, 102. Herschel and Spitzer IR spectroscopy.

Hanabata *et al.* 2014, ApJ, 786, 145. Fermi observations of region. Zhou *et al.* 2014, ApJ, 791, 87. XMM-Newton observations. Gusdorf *et al.* 2014, IAUS, 296, 178. CO observations. Nakamura *et al.* 2014, PASJ, 66, 62. XMM-Newton observations of

NE.

Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. Maxted *et al.* 2016, MNRAS, 462, 532. NH<sub>3</sub> observations of region.

Acero *et al.* 2016, ApJS, 224, 8. Fermi observations.
Pannuti *et al.* 2017, ApJ, 839, 59. Optical and X-ray observations.
Abdalla *et al.* 2018, A&A, 612, A1. HESS observations.
Okon *et al.* 2018, PASJ, 70, 35. Suzaku observations.
Nobukawa *et al.* 2018, ApJ, 854, 87. Suzaku observations.

Cui et al. 2018, ApJ, 860, 69. Fermi observations.

Abdalla *et al.* 2018, A&A, 612, A1. HESS observations.

Ruiz-Lapuente et al. 2018, ApJ, 862, 124. HST search for progenitor companion.

# G6.4+4.0

**RA:** 17<sup>h</sup>45<sup>m</sup>10<sup>s</sup> 1-GHz flux/Jy: 1.3? Size/arcmin: 31 **Dec:**  $-21^{\circ}22'$ **Spectral index:** 0.4? Type: S

Radio: Faint asymmetric shell.

#### References:

Reich et al. 1988, IAUCo, 101, 293. Effelsberg 100-m at 2.7 GHz (4'.3).

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

# G6.5 - 0.4

**RA:**  $18^{h}02^{m}11^{s}$ **1-GHz flux/Jy: 27** Size/arcmin: 18 **Dec:** -23°34′ Type: S **Spectral index:** 0.6

Has been called G6.51–0.48, and part has been called G6.67–0.42.

**Radio:** Shell, overlapping G6.4–0.1.

**Optical:** Detected.

### References:

Yusef-Zadeh et al. 2000, ApJ, 540, 842. VLA at 330 MHz  $(2.6 \times 5.5)$  and 1.4 GHz  $(0.7 \times 1.1)$ .

Brogan et al. 2006, ApJ, 639, L25. VLA at 330 MHz (42'': S = $60.8\pm0.4$  Jy), plus other observations.

Casandjian & Grenier 2008, A&A, 489, 849.  $\gamma$ -ray observations. Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Stupar & Parker 2011, MNRAS, 414, 2282. Ha observations. Ajello et al. 2016, ApJ, 819, 44. Fermi observations.

# G7.0 - 0.1

**RA:** 18<sup>h</sup>01<sup>m</sup>50<sup>s</sup> 1-GHz flux/Jy: 2.5? Size/arcmin: 15 **Dec:**  $-22^{\circ}54'$ **Spectral index:** 0.5? Type: S

Has been called G7.06–0.12.

Radio: Double rim, brightest in W, confused by bright HII region M20 in SE.

### References:

Yusef-Zadeh et al. 2000, ApJ, 540, 842. VLA at 327 MHz  $(2.6 \times 5.5)$  and 1.48 GHz  $(40'' \times 65'')$ .

Dubner et al. 2000, AJ, 120, 1933. VLA at 328 MHz (52"×97") and 1415 MHz (48"×88").

# G7.2+0.2

**RA:** 18<sup>h</sup>01<sup>m</sup>07<sup>s</sup> 1-GHz flux/Jy: 2.8 Size/arcmin: 12 **Dec:** -22°38′ Spectral index: 0.6 Type: S

Has been called G7.20+0.20.

Radio: Partial shell.

### References:

Brogan et al. 2006, ApJ, 639, L25. VLA at 330 MHz (42": S =  $5.2\pm0.2$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

**G7.7–3.7** 

**RA:** 18<sup>h</sup>17<sup>m</sup>25<sup>s</sup> **1-GHz flux/Jy:** 11 **Size/arcmin:** 22 **Dec:** -24°04' **Spectral index:** 0.32 **Type:** S

Has been associated with the SN of AD386.

**Radio:** Shell, with high polarisation.

X-ray: Arc in S.

#### References:

Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8'.4) and 5 GHz (4'.4). Milne *et al.* 1986, MNRAS, 223, 487. MOST at 843 MHz  $(44'' \times 108'')$  and Parkes 64-m at 8.4 GHz (3':  $S = 4.6 \pm 0.5$  Jy), with polarisation, plus review of flux densities.

Dubner et al. 1996, AJ, 111, 1304. VLA at 1.4 GHz  $(35'' \times 70'' : S = 9.9 \pm 0.1 \text{ Jy})$ , including polarisation.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Zhou *et al.* 2018, ApJ, 865, L6. XMM-Newton observations.

## G8.3 - 0.0

**RA:**  $18^h04^m34^s$  **1-GHz flux/Jy:** 1.2 **Size/arcmin:**  $5\times4$  **Dec:**  $-21^\circ49'$  **Spectral index:** 0.6 **Type:** S

Has been called G8.31-0.09.

Radio: Shell.
References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 2.3 \pm 0.1$  Jy), plus other observations.

Higashi et al. 2008, ApJ, 683, 957. γ-ray detection.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Kilpatrick et al. 2016, ApJ, 816, 1. CO observations, including broad lines.

## G8.7 - 5.0

RA: 18<sup>h</sup>24<sup>m</sup>10<sup>s</sup> 1-GHz flux/Jy: 4.4 Size/arcmin: 26 Dec: -23°48′ Spectral index: 0.3 Type: S

Radio: Asymmetric shell.

## References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4/3). Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. de Wilt *et al.* 2017, MNRAS, 468, 2093. Molecular line observations of region.

G8.7-0.1 (W30)

**RA:** 18<sup>h</sup>05<sup>m</sup>30<sup>s</sup> **1-GHz flux/Jy:** 80 **Size/arcmin:** 45 **Dec:** -21°26' **Spectral index:** 0.5 **Type:** S?

Has been called G8.6–0.1.

Radio: Clumpy non-thermal shell, with low-frequency turnover.

**X-ray:** Northern edge detected.

**Point sources:** Pulsar inside western edge.

### References

Odegard 1986, AJ, 92, 1372. TPT at 57.5 MHz ( $7'.2 \times 9'.7 : S = 190 \pm 50$  Jy).

Kassim & Weiler 1990, Nature, 343, 146. VLA at 327 MHz (3!0×3!7).

Kassim & Weiler 1990, ApJ, 360, 184. VLA at 327 MHz  $(2.8 \times 4.1 : S = 129 \pm 11 \text{ Jy})$ , and part at 1.4 GHz  $(0.9 \times 1.8)$ , plus review of flux densities.

Frail *et al.* 1994, AJ, 107, 1120. VLA at 327 MHz (37"×55").
 Finley & Ögelman 1994, ApJ, 434, L25. ROSAT observations, including pulsar.

Aharonian et al. 2005, Science, 307, 1938. HESS detection. Aharonian et al. 2006, ApJ, 636, 777. HESS observations. Landi et al. 2006, ApJ, 651, 190. X-ray observations. Briskin et al. 2006, ApJ, 652, 554. Pulsar proper motion. Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser detection. Castro & Slane 2010, ApJ, 717, 372. Fermi observations. Ajello et al. 2012, ApJ, 744, 80. Fermi observations. Pihlström et al. 2014, AJ, 147, 73. VLA search for methanol masers.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

G8.9 + 0.4

RA: 18<sup>h</sup>03<sup>m</sup>58<sup>s</sup>
1-GHz flux/Jy: 9
Dec: -21°03'
Spectral index: 0.6
Type: S

Has been called G8.90+0.40.

Radio: Shell.
References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 18.2 \pm 0.5$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

G9.7 - 0.0

**RA:**  $18^{\text{h}}07^{\text{m}}22^{\text{s}}$  **1-GHz flux/Jy:** 3.7 **Size/arcmin:**  $15 \times 11$  **Dec:**  $-20^{\circ}35'$  **Spectral index:** 0.6 **Type:** S

Has been called G9.7-0.1 and G9.70-0.06.

Radio: Shell.
References:

Frail et al. 1994, AJ, 107, 1120. VLA at 327 MHz.

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 6.5\pm0.2$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser detection.

Pihlström et al. 2014, AJ, 147, 73. VLA search for methanol

maser

Yeung et al. 2016, ApJ, 827, 41. Fermi observations.

G9.8 + 0.6

RA: 18<sup>h</sup>05<sup>m</sup>08<sup>s</sup> 1-GHz flux/Jy: 3.9 Size/arcmin: 12 Dec: -20°14′ Spectral index: 0.5 Type: S

Radio: Asymmetric shell.

References:

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'6)

Caswell 1983, MNRAS, 204, 833. Molonglo at 408 MHz (3' :  $S = 5.8 \pm 0.6$  Jy).

Dubner *et al.* 1993, AJ, 105, 2251. VLA at 1.47 GHz  $(41'' \times 63'' : S = 3.5 \pm 0.4 \text{ Jy})$ .

Frail et al. 1994, AJ, 107, 1120. VLA at 327 MHz.

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

G9.9 - 0.8

RA: 18<sup>h</sup>10<sup>m</sup>41<sup>s</sup> 1-GHz flux/Jy: 6.7 Size/arcmin: 12 Dec: -20°43′ Spectral index: 0.4 Type: S

Has been called G9.95-0.81.

Radio: Shell.
Optical: Detected.

References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 11.0 \pm 0.3$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Stupar & Parker 2011, MNRAS, 414, 2282.  $H\alpha$  observations.

Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. Kilpatrick *et al.* 2016, ApJ, 816, 1. CO observations, including broad lines.

G10.5-0.0

**RA:** 18<sup>h</sup>09<sup>m</sup>08<sup>s</sup> 1-GHz flux/Jy: 0.9 Size/arcmin: 6 **Dec:** -19°47′ Spectral index: 0.6 Type: S

Has been called G10.59-0.04.

Radio: Partial shell. **X-ray:** Possibly detected.

References:

Sugizaki et al. 2001, ApJS, 134, 77. ASCA survey observations. Brogan et al. 2006, ApJ, 639, L25. VLA at 330 MHz (42'': S =

 $1.4\pm0.1$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

G11.0 - 0.0

**RA:**  $18^{h}10^{m}04^{s}$ 1-GHz flux/Jy: 1.3 Size/arcmin:  $11 \times 9$ **Dec:** -19°25′ Spectral index: 0.6 Type: S

Has been called G11.0+0.0 and G11.03-0.05.

Radio: Partial shell. **X-ray:** Diffuse emission.

**Distance:** Optical absorption suggests 2.4 kpc.

References:

Bamba et al. 2003, ApJ, 589, 253. ASCA observations. Brogan et al. 2004, AJ, 127, 355. VLA at 330 MHz (25"), 1.5 GHz (25"), and 74 MHz.

Brogan et al. 2006, ApJ, 639, L25. VLA at 330 MHz (42'' : S =

 $3.1\pm0.2$  Jy), plus other observations.

Castelletti et al. 2016, A&A, 587, A71. VLA at 1.4 GHz (4".4×8".3) and CO observations of region. Araya et al. 2018, ApJ, 859, 69. Fermi observations.

Shan et al. 2018, ApJS, 238, 35. Optical absorption for distance.

G11.1-1.0

**RA:** 18<sup>h</sup>14<sup>m</sup>03<sup>s</sup> 1-GHz flux/Jy: 5.8 Size/arcmin:  $18 \times 12$ **Dec:** -19°46′ Spectral index: 0.5 Type: S

Has been called G11.2-1.1 and G11.17-1.04.

Radio: Shell. **Optical:** Detected.

References:

Brogan et al. 2006, ApJ, 639, L25. VLA at 330 MHz (42": S =  $11.0\pm0.3$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Stupar & Parker 2011, MNRAS, 414, 2282. Ha observations.

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S= 3.40±0.25 Jy) including polarisation and review of flux densities.

G11.1-0.7

**RA:** 18<sup>h</sup>12<sup>m</sup>46<sup>s</sup> 1-GHz flux/Jy: 1.0 Size/arcmin:  $11 \times 7$ **Dec:**  $-19^{\circ}38'$ **Spectral index:** 0.7 Type: S

Has been called G11.15-0.71.

Radio: Partial shell.

References:

Brogan et al. 2004, AJ, 127, 355. VLA at 330 MHz (25"), 1.5 GHz (25"), and 74 MHz.

Brogan et al. 2006, ApJ, 639, L25. VLA at 330 MHz (42'': S = $2.3\pm0.1$  Jy), plus other observations.

# G11.1+0.1

**RA:**  $18^{\text{h}}09^{\text{m}}47^{\text{s}}$  **1-GHz flux/Jy:** 2.3 **Size/arcmin:**  $12\times10$  **Dec:**  $-19^{\circ}12'$  **Spectral index:** 0.4 **Type:** S

Has been called G11.18+0.11.

Radio: Shell.

**References:** Brogan *et al.* 2004, AJ, 127, 355. VLA at 330 MHz (25"),

1.5 GHz (25"), and 74 MHz.

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 3.5 \pm 0.2$  Jy), plus other observations.

Castelletti *et al.* 2016, A&A, 587, A71. VLA at 1.4 GHz (4".4×8".3) and CO observations of region.

## G11.2 - 0.3

**RA:** 18<sup>h</sup>11<sup>m</sup>27<sup>s</sup> **1-GHz flux/Jy:** 22 **Size/arcmin:** 4 **Dec:** -19°25′ **Spectral index:** 0.5 **Type:** C

Probably associated with the SN of AD386.

**Radio:** Symmetrical clumpy shell, with flatter spectrum core.

X-ray: Shell, with hard spectrum centrally brightened region around pulsar.

Point sources: Central pulsar.

**Distance:** Hi absorption indicates 4.4 kpc.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Radhakrishnan *et al.* 1972, ApJS, 24, 49. Hi absorption. Becker & Kundu 1975, AJ, 80, 679. NRAO 140-ft at 10.6 GHz

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'.6).

Downes 1984, MNRAS, 210, 845. VLA at 1465 MHz (20"×25") and Einstein observations, with review of flux densities.

Becker *et al.* 1985, ApJ, 296, 461. VLA at 1.4 and 5 GHz, plus HI absorption, Einstein observations.

Morsi & Reich 1987, A&AS, 71, 189. Effelsberg 100-m at 32 GHz  $(26''.5:S=4.04\pm0.24 \text{ Jy})$ .

Green *et al.* 1988, MNRAS, 231, 735. VLA at 1.4 and 5 GHz. Kassim 1992, AJ, 103, 943. VLA at 327 MHz  $(3'.2 \times 4'.1 : S = 39 \text{ Jy})$ .

Reynolds et al. 1994, MNRAS, 271, L1. ROSAT image and spectra.

Vasisht *et al.* 1996, ApJ, 456, L59. ASCA observations. Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Torii *et al.* 1997, ApJ, 489, L145. ASCA detection of pulsar. Torii *et al.* 1999, ApJ, 523, L69. X-ray timing observations of pulsar

Kothes & Reich 2001, A&A, 372, 627. Effelsberg 100-m at 4.25, 10.45, 14.9 and 32 GHz (2'.5, 1'.1, 0'.86 and 0'.45 :  $S = 9.6 \pm 0.5, 6.3 \pm 0.4, 5.7 \pm 0.4$  and  $3.8 \pm 0.4$ ).

Kaspi et al. 2001, ApJ, 560, 371. Chandra observations.

Tam *et al.* 2002, ApJ, 572, 202. VLA at 1.4/1.5 GHz (1".8×2".6:S = 16.6±0.9 Jy) and 5 GHz (1".5×2".1:S = 8.4±0.9 Jy) for spectral studies.

Reich 2002, in NSPS, p1. Effelsberg 100-m at 14.7 GHz. Roberts *et al.* 2003, ApJ, 588, 992. Chandra observations.

Tam & Roberts 2003, ApJ, 598, L27. Multi-epoch VLA observations at 1.4/1.5 GHz and 5 GHz, for expansion studies.

Brogan *et al.* 2004, AJ, 127, 355. VLA at 330 MHz (25"), 1.5 GHz (25"), and 74 MHz.

Bock & Gaensler 2005, ApJ, 626, 343. BIMA at 88.6 GHz (18"). Reach *et al.* 2006, AJ, 131, 1479. Spitzer observations.

Kaplan & Moon 2006, ApJ, 644, 1056. IR upper limit for pulsar. Koo *et al.* 2007, ApJ, 657, 308. IR observations.

Dean et al. 2008, MNRAS, 384, L29. INTEGRAL observations of pulsar and nebula.

Moon et al. 2009, ApJ, 703, L81. IR spectroscopy.

Mizuno et al. 2010, AJ, 139, 1542. Spitzer observations.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9 $.5:S=9.0\pm0.5$  Jy) including polarisation and review of flux densities.

Andersen et al. 2011, ApJ, 742, 7. Spitzer observations.

Lee et al. 2013, ApJ, 770, 143. IR observations.

Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. Kilpatrick *et al.* 2016, ApJ, 816, 1. CO observations, including broad lines.

Borkowski *et al.* 2016, ApJ, 819, 160. Chandra observations. Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

# G11.4-0.1

**RA:** 18<sup>h</sup>10<sup>m</sup>47<sup>s</sup> **1-GHz flux/Jy:** 6 **Size/arcmin:** 8 **Dec:** -19°05′ **Spectral index:** 0.5 **Type:** S?

**Radio:** Incomplete shell, possibly with central core.

#### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 9.4 Jy) contaminated by sidelobes of a nearby source, and Parkes 64-m at 5 GHz (4': S = 2.8 Jy).

Angerhofer *et al.* 1977, A&A, 55, 11. NRAO 140-ft at 5 GHz (6'8:  $S = 2.0 \pm 0.4$  Jy).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'6)

Kassim 1992, AJ, 103, 943. VLA at 327 MHz ( $3.2 \times 4.1 : S = 18 \text{ Jy}$ ).

Dubner *et al.* 1993, AJ, 105, 2251. VLA at 1.47 GHz ( $44'' \times 63''$ :  $S = 5.1 \pm 0.6$  Jy).

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Brogan *et al.* 2004, AJ, 127, 355. VLA at 330 MHz (25"), 1.5 GHz (25"), and 74 MHz.

Rangelov *et al.* 2014, ApJ, 796, 34. X-ray upper limit. Castelletti *et al.* 2016, A&A, 587, A71. VLA at 1.4 GHz (4".4×8".3) and CO observations of region.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

## G11.8 - 0.2

RA: 18<sup>h</sup>12<sup>m</sup>25<sup>s</sup> 1-GHz flux/Jy: 0.7 Size/arcmin: 4 Dec: -18°44′ Spectral index: 0.3 Type: S

Has been called G11.89-0.21.

Radio: Shell.

**X-ray:** Possibly detected.

#### References:

Sugizaki *et al.* 2001, ApJS, 134, 77. ASCA survey observations. Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 0.9 \pm 0.1$  Jy), plus other observations. Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

# G12.0-0.1

RA: 18<sup>h</sup>12<sup>m</sup>11<sup>s</sup> 1-GHz flux/Jy: 3.5 Size/arcmin: 7? Dec: -18°37′ Spectral index: 0.7 Type: ?

**Radio:** Incomplete shell, defined in E only. **X-ray:** Detected, including possible PWN.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 6.6 Jy) and Parkes 64-m at 5 GHz (4': S = 1.1 Jy).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2/6).

Kassim 1992, AJ, 103, 943. VLA at 327 MHz (3'.2×4'.1).

Dubner *et al.* 1993, AJ, 105, 2251. VLA at 1.47 GHz  $(41'' \times 61'' : S = 0.7 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.
Sugizaki et al. 2001, ApJS, 134, 77. ASCA survey observations.
Yamauchi et al. 2014, PASJ, 66, 20. Suzaku observations of possible PWN.

## G12.2+0.3

**RA:**  $18^{\rm h}11^{\rm m}17^{\rm s}$  **1-GHz flux/Jy:** 0.8 **Size/arcmin:**  $6\times5$  **Dec:**  $-18^{\circ}10'$  **Spectral index:** 0.7 **Type:** S

Has been called G12.26+0.30.

Radio: Partial shell.

## References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 1.5 \pm 0.1$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Kilpatrick *et al.* 2016, ApJ, 816, 1. CO observations, including broad lines.

## G12.5+0.2

**RA:**  $18^{\rm h}12^{\rm m}14^{\rm s}$  **1-GHz flux/Jy:** 0.6 **Size/arcmin:**  $6\times5$  **Dec:**  $-17^{\circ}55'$  **Spectral index:** 0.4 **Type:** C?

Has been called G12.58+0.22. **Radio:** Diffuse, central brightened.

References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 0.01 \text{ L}_{\odot}$ ) when the enterpretions of the state of t

 $0.8\pm0.1$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

## G12.7 - 0.0

**RA:** 18<sup>h</sup>13<sup>m</sup>19<sup>s</sup> **1-GHz flux/Jy:** 0.8 **Size/arcmin:** 6 **Dec:** -17°54′ **Spectral index:** 0.8 **Type:** S

Has been called G12.72-0.00.

Radio: Shell.

References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 2.0 \pm 0.1$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

G12.8 - 0.0

**RA:** 18<sup>h</sup>13<sup>m</sup>37<sup>s</sup> **1-GHz flux/Jy:** 0.8 **Dec:** -17°49′ **Spectral index:** 0.5 **Size/arcmin:** 3 **Type:** C?

Has been called G12.82-0.02 and G12.83-0.02.

Radio: Shell. X-ray: Diffuse.

Point sources: Central X-ray pulsar.

References:

Brogan *et al.* 2005, ApJ, 629, L105. VLA at 330 MHz  $(19'' \times 32'')$ , plus other observations.

Ubertini et al. 2005, ApJ, 629, L109. INTEGRAL and other observations.

Aharonian et al. 2006, ApJ, 636, 777. HESS detection. Albert et al. 2006, ApJ, 637, L41.  $\gamma$ -ray observations.

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 1.2\pm0.1$  Jy), plus other observations.

Landi et al. 2006, ApJ, 651, 190. X-ray observations.

Funk et al. 2007, A&A, 470, 249. XMM-Newton observations, CO observations of surroundings.

Helfand *et al.* 2007, ApJ, 665, 1297. Chandra observations. Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Gotthelf & Halpern 2009, ApJ, 700, L158. Pulsar detection. Halpern *et al.* 2012, ApJ, 753, L14. Pulsar observations. Araya *et al.* 2018, ApJ, 859, 69. Fermi observations. Dzib *et al.* 2018, ApJ, 866, 100. VLA of central source.

# G13.3-1.3

**RA:**  $18^h 19^m 20^s$  **1-GHz flux/Jy:** ? **Size/arcmin:**  $70 \times 40$  **Dec:**  $-18^{\circ}00'$  **Spectral index:** ? **Type:** S?

Radio: Amorphous emission.Optical: Filaments in S.X-ray: Elongated emission.

**Distance:** Absorption indicates 2–4 kpc.

References:

Seward *et al.* 1995, ApJ, 449, 681. ROSAT detection, optical studies and observations of CO.

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.
Koralesky et al. 1998, AJ, 116, 1323. VLA search for OH emission.

# G13.5+0.2

**RA:**  $18^{\rm h}14^{\rm m}14^{\rm s}$  **1-GHz flux/Jy:** 3.5? **Size/arcmin:**  $5\times4$  **Dec:**  $-17^{\circ}12'$  **Spectral index:** 1.0? **Size/arcmin:**  $5\times4$ 

Has been called G13.46+0.16.

Radio: Elongated, incomplete shell.

#### References:

Helfand *et al.* 1989, ApJ, 341, 151. VLA at 5 GHz ( $S = 0.65 \pm 0.05$  Jy) and 1.4 GHz ( $15'': S = 2.67 \pm 0.5$  Jy). Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.

## G14.1 - 0.1

**RA:**  $18^{\rm h}16^{\rm m}40^{\rm s}$  **1-GHz flux/Jy:** 0.5 **Size/arcmin:**  $6\times5$  **Dec:**  $-16^{\circ}41'$  **Spectral index:** 0.6 **Type:** S

Has been called G14.18-0.12.

Radio: Shell.

References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz (42'': S =

0.9±0.1 Jy), plus other observations. Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

# G14.3+0.1

**RA:**  $18^{\rm h}15^{\rm m}58^{\rm s}$  **1-GHz flux/Jy:** 0.6 **Size/arcmin:**  $5\times4$  **Dec:**  $-16^{\circ}27'$  **Spectral index:** 0.4 **Type:** S

Has been called G14.30+0.14. Has been suggested this is not an SNR.

Radio: Partial shell.

### References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 1.2\pm0.1$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Pinheiro Gonçalves *et al.* 2011, AJ, 142, 47. IR observations.

## G15.1 - 1.6

**RA:**  $18^{\text{h}}24^{\text{m}}00^{\text{s}}$  **1-GHz flux/Jy:** 5.5? **Size/arcmin:**  $30 \times 24$  **Dec:**  $-16^{\circ}34'$  **Spectral index:** 0.0? **Type:** S?

Radio: Elongated, incomplete shell.

Optical: Diffuse shell.

### References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz

Boumis et al. 2008, A&A, 481, 705. Optical detection.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 4.8±0.3 Jy) including polarisation and review of flux densities.

# G15.4+0.1

**RA:**  $18^h 18^m 02^s$  **1-GHz flux/Jy:** 5.6 **Size/arcmin:**  $15 \times 14$  **Dec:**  $-15^{\circ} 27'$  **Spectral index:** 0.62 **Type:** C?

Has been called G15.42+0.18.

Radio: Shell.

**X-ray:** Centrally brightened.

**Distance:** Hi observations suggest 4.8 kpc.

#### References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 10.9 \pm 0.3$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 2.1±0.2 Jy) including polarisation and review of flux densities.

Castelletti *et al.* 2013, A&A, 557, L15. GMRT at 1.4 GHz (15") plus SGPS H<sub>I</sub> and CO observations of region.

Abramowski *et al.* 2014, A&A, 562, A40. HESS and XMM-Newton observations.

Supan *et al.* 2015, A&A, 576, A81. GMRT at 624 MHz (10"). Abdalla *et al.* 2018, A&A, 612, A1. HESS observations.

# G15.9 + 0.2

**RA:**  $18^h 18^m 52^s$  **1-GHz flux/Jy:** 5.0 **Size/arcmin:**  $7 \times 5$  **Dec:**  $-15^{\circ}02'$  **Spectral index:** 0.63 **Type:** S?

Radio: Incomplete shell, with bright concentration to the E.

**X-ray:** Shell, brighter to S and E. **Point sources:** Central X-ray source.

#### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 7.7 Jy) and Parkes 64-m at 5 GHz (4': S = 1.9 Jy).

Angerhofer *et al.* 1977, A&A, 55, 11. NRAO 140-ft at 5 GHz (6.8:  $S=1.1\pm0.2$  Jy).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'6)

Caswell *et al.* 1982, MNRAS, 200, 1143. FIRST at 1415 MHz (44"×58").

Dubner *et al.* 1996, AJ, 111, 1304. VLA at 330 MHz ( $61'' \times 77''$ :  $S = 11.2 \pm 1.0$  Jy), 1.4 GHz ( $14'' \times 23''$ :  $S = 3.9 \pm 0.1$  Jy) and 4.9 GHz ( $13'' \times 16''$ ).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations. Reynolds et al. 2006, ApJ, 652, L45. Chandra observations.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 2.0±0.3 Jy) including polarisation and review of flux densities. Klochov *et al.* 2016, A&A, 592, L12. Chandra observations.

Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Maggi & Acero 2017, A&A, 597, A65. XMM-Newton observations.

Sasaki et al. 2018, MNRAS, 479, 3033. Chandra and IR observations, plus  $H\alpha$  non-detection.

# G16.0 - 0.5

**RA:**  $18^h21^m56^s$  **1-GHz flux/Jy:** 2.7 **Size/arcmin:**  $15\times10$  **Dec:**  $-15^{\circ}14'$  **Spectral index:** 0.6 **Type:** S

Has been called G16.05–0.57.

Radio: Shell.

## References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 4.9 \pm 0.2$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Beaumont *et al.* 2011, ApJ, 741, 14. CO of region.

Froebrich et al. 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.

# G16.2-2.7

**RA:** 18<sup>h</sup>29<sup>m</sup>40<sup>s</sup> **1-GHz flux/Jy:** 2.5 **Dec:** -16°08' **Size/arcmin:** 17 **Type:** S

Radio: Double rim.

#### References:

Trushkin 1999, A&A, 352, L103. Review of radio observations. Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 1.28±0.10 Jy) including polarisation and review of flux densities.

# G16.4 - 0.5

RA: 18<sup>h</sup>22<sup>m</sup>38<sup>s</sup> 1-GHz flux/Jy: 4.6 Size/arcmin: 13 Dec: -14°55′ Spectral index: 0.3? Type: S

Has been called G16.41–0.55.

Radio: Partial shell.

### References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 10.0 \pm 0.3$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9'.5: S = 3.0±0.3 Jy) including polarisation and review of flux densities.

## G16.7 + 0.1

RA: 18<sup>h</sup>20<sup>m</sup>56<sup>s</sup> 1-GHz flux/Jy: 3.0 Size/arcmin: 4 Dec: -14°20′ Spectral index: 0.6 Type: C

Has been called G16.73+0.08.

**Radio:** Asymmetric shell with flat-spectrum core.

**X-ray:** Non-thermal core.

### References:

Helfand *et al.* 1989, ApJ, 341, 151. VLA at 5 GHz (5:S=0.95 Jy) and 1.4 GHz (15'':S=2.43 Jy) and Ooty at 327 MHz (S=5.13 Jy).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations, including masers.

Reynoso & Mangum 2000, ApJ, 545, 874. CO observations of surroundings.

Sugizaki *et al.* 2001, ApJS, 134, 77. ASCA survey observations. Helfand *et al.* 2003, ApJ, 592, 941. XMM-Newton observations. Bock & Gaensler 2005, ApJ, 626, 343. BIMA at 88.6 GHz (19"×25").

Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers.

Bhatnagar et al. 2011, ApJ, 739, L20. VLA at 6 GHz.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9!5 : S = 1.23 $\pm$ 0.11 Jy) including polarisation and review of flux densities.

Pihlström *et al.* 2014, AJ, 147, 73. VLA search for methanol masers.

Kilpatrick *et al.* 2016, ApJ, 816, 1. CO observations, including broad lines.

Chang et al. 2018, MNRAS, 474, 2607. Chandra observations.

# G17.0 - 0.0

RA: 18<sup>h</sup>21<sup>m</sup>57<sup>s</sup> 1-GHz flux/Jy: 0.5 Size/arcmin: 5 Dec: -14°08′ Spectral index: 0.5 Type: S

Has been called G17.02-0.04.

Radio: Shell.

### References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 0.7 \pm 0.1$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

## G17.4 - 2.3

**RA:** 18<sup>h</sup>30<sup>m</sup>55<sup>s</sup> **1-GHz flux/Jy:** 5 **Size/arcmin:** 24? **Dec:** -14°52′ **Spectral index:** 0.5? **Type:** S

Radio: Incomplete, poorly defined shell.

Optical: Filaments to SE, and diffuse emission.

#### References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4'.3).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

Boumis *et al.* 2002, A&A, 385, 1042. Optical observations. Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 2.3±0.2 Jy) including polarisation and review of flux densities.

## G17.4 - 0.1

RA: 18<sup>h</sup>23<sup>m</sup>08<sup>s</sup> 1-GHz flux/Jy: 0.4 Size/arcmin: 6 Dec: -13°46′ Spectral index: 0.7 Type: S

Has been called G17.48-0.12.

Radio: Partial shell.

#### References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 0.9 \pm 0.1$  Jy), plus other observations.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

# G17.8 - 2.6

RA: 18<sup>h</sup>32<sup>m</sup>50<sup>s</sup> 1-GHz flux/Jy: 5 Size/arcmin: 24 Dec: -14°39′ Spectral index: 0.5 Type: S

Radio: Well defined shell.

### References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4/3).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9!5:  $S = 2.23 \pm 0.13$  Jy) including polarisation and review of flux densities.

## G18.1-0.1

RA: 18<sup>h</sup>24<sup>m</sup>34<sup>s</sup> 1-GHz flux/Jy: 4.6 Size/arcmin: 8 Dec: -13°11′ Spectral index: 0.5 Type: S

Has been called G18.1–0.2 and G18.16–0.16.

Radio: Shell.

**X-ray:** Possibly detected.

**Distance:** Hi absorption suggests 6.4 kpc.

## References:

Odegard 1986, AJ, 92, 1372. TPT at 57.5 MHz (8'). Kassim *et al.* 1989, ApJ, 338, 152. VLA at 330 MHz (2'.8×3'.2)

and 1.5 GHz  $(0.9 \times 1.5)$ .

Sugizaki *et al.* 2001, ApJS, 134, 77. ASCA survey observations. Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 7.6 \pm 0.1 \text{ Jy}$ ), plus other observations

 $7.6\pm0.1$  Jy), plus other observations.

Hewitt & Yusef-Zadeh 2009, ApJ, 694, L16. OH maser search.

Paron et al. 2013, MNRAS, 433, 1619. CO, optical and other observations of region.

Leahy *et al.* 2014, MNRAS, 438, 1813. VGPS HI observations. Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. Voison *et al.* 2016, MNRAS, 458, 2813. Molecular line observations of region.

Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

# G18.6 - 0.2

RA: 18<sup>h</sup>25<sup>m</sup>55<sup>s</sup> 1-GHz flux/Jy: 1.4 Size/arcmin: 6 Dec: -12°50′ Spectral index: 0.4 Type: S

Has been called G18.62-0.28.

Radio: Partial shell.

**Distance:** HI absorption suggests 4.4 kpc.

#### References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 1.9 \pm 0.1$  Jy), plus other observations.

Kilpatrick *et al.* 2016, ApJ, 816, 1. CO observations, including broad lines.

Voison et al. 2016, MNRAS, 458, 2813. Molecular line observations of region.

Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

G18.8+0.3 Kes 67

**RA:**  $18^h23^m58^s$  **1-GHz flux/Jy:** 33 **Size/arcmin:**  $17\times11$  **Dec:**  $-12^\circ23'$  **Spectral index:** 0.46 **Type:** S

Has been called G18.9+0.3.

Radio: Incomplete shell, in complex region near the HII region W39.

Distance: Association with molecular cloud suggests 12 kpc, and H<sub>I</sub> absorption suggests 13.8 kpc.

### References:

Willis 1973, A&A, 26, 237. NRAO 300-ft at 2.7 GHz (5' :  $S = 17\pm7$  Jy).

Clark *et al.* 1975, AuJPA, 37, 75. Molonglo at 408 MHz (3': S = 38 Jy).

Caswell *et al.* 1975, A&A, 45, 239. Parkes HI absorption. Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8/4) and 5 GHz (4/4).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'.6).

Milne *et al.* 1989, PASA, 8, 187. Parkes 64-m at 8.4 GHz ( $3!0:S = 12.9 \pm 1.0$  Jy), including polarisation.

Kassim 1992, AJ, 103, 943. VLA at 327 MHz ( $2'.9 \times 3'.5 : S = 55$  Jy).

Dubner *et al.* 1996, AJ, 111, 1304. VLA at 1.4 GHz  $(55'' \times 75'' : S = 29.9 \pm 0.3 \text{ Jy})$ .

Dubner *et al.* 1999, AJ, 118, 930. Parkes 64-m at 1.6 GHz (15") for HI, VLA at 1.6 GHz (12"×17") for OH, plus CO observations.

Dubner et al. 2004, A&A, 426, 201. CO observations of environment.

Tian et al. 2007, A&A, 474, 541. VGPS at 1.4 GHz (1') including HI, plus CO observations of region.

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5:  $S = 15.3 \pm 0.9$  Jy) including polarisation and review of flux densities. Vasquez et al. 2012, A&A, 545, A89. CO observations of region.

Paron et al. 2012, A&A, 547, A60. CO and other molecular observations of region.

Paron et al. 2015, A&A, 580, A51. CO observations in S.

Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS Hi absorption observations.

## G18.9 - 1.1

**RA:** 18<sup>h</sup>29<sup>m</sup>50<sup>s</sup> **1-GHz flux/Jy:** 37 **Size/arcmin:** 33 **Dec:** -12°58′ **Spectral index:** 0.39 **Type:** C?

Has been called G18.95–1.1 and G18.94–1.04.

**Radio:** Non-thermal, diffuse partially limb-brightened, with central ridge.

Optical: Detected. X-ray: Partial shell.

**Point sources:** Compact X-ray source, with diffuse nebula.

**Distance:** Optical absorption suggests 1.8 kpc.

#### References:

Fürst *et al.* 1985, Nature, 314, 720. Effelsberg 100-m at 4.75 GHz (2'.4: S=23.8 Jy), plus other flux densities.

Odegard 1986, AJ, 92, 1372. TPT at 57.5 MHz ( $7'.2 \times 8'$ :  $S = 82\pm15$  Jy), plus review of flux densities.

Barnes & Turtle 1988, IAUCo, 101, 347. Molonglo at 408 MHz  $(2.9 \times 3.1: S = 58 \pm 9 \text{ Jy})$  and Parkes 64-m at 5 GHz  $(4.1 \times 4.4: S = 23 \pm 6 \text{ Jy})$ .

Patnaik *et al.* 1988, Nature, 332, 136. Ooty at 327 MHz  $(0.6 \times 1.6)$ .

Fürst *et al.* 1989, A&A, 209, 361. Effelsberg 100-m at 4.75 GHz (2'.45: S=23.8 Jy) and VLA at 1.5 and 4.9 GHz ( $14''\times19''$ ), and Effelsberg 100-m at 1.4 GHz (9') for Hi.

Aschenbach et al. 1991, A&A, 246, L32. ROSAT observations.

Fürst et al. 1997, A&A, 319, 655. ROSAT observations, and Effelsberg 100-m at 10.6 GHz (1!1).

Reich 2002, in NSPS, p1. Effelsberg 100-m at 10.6 GHz, including polarisation.

Harrus et al. 2004, ApJ, 603, 152. ROSAT and ASCA observations.

Tüllmann et al. 2010, ApJ, 720, 848. Chandra detection of compact source.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 19.6±1.0 Jy) including polarisation and review of flux densities. Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations. Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. Acero *et al.* 2016, ApJS, 224, 8. Fermi observations.

Shan et al. 2018, ApJS, 238, 35. Optical absorption for distance.

# G19.1 + 0.2

RA: 18<sup>h</sup>24<sup>m</sup>56<sup>s</sup> 1-GHz flux/Jy: 10 Size/arcmin: 27 Dec: -12°07′ Spectral index: 0.5 Type: S

Has been called G19.15+0.27.

Radio: Partial shell.

### References:

Brogan *et al.* 2006, ApJ, 639, L25. VLA at 330 MHz ( $42'': S = 17.4 \pm 0.4$  Jy), plus other observations.

## G20.0-0.2

RA: 18<sup>h</sup>28<sup>m</sup>07<sup>s</sup> 1-GHz flux/Jy: 10 Size/arcmin: 10 Dec: -11°35′ Spectral index: 0.1 Type: F

Radio: Faint, filled-centre, polarised.

**X-ray:** Centrally brightened.

**Point sources:** OH source 20.1–0.1 is nearby. **Distance:** HI absorption suggests 11.2 kpc.

### References:

Becker & Helfand 1985, ApJ, 297, L25. VLA at 1.4 and 5 GHz (12").

Odegard 1986, AJ, 92, 1372. TPT at 57.5 MHz ( $7'.2 \times 8'$ :  $S = 8.5 \pm 2$  Jy), plus review of flux densities.

Junkes *et al.* 1988, LNP, 316, 134. Effelsberg 100-m at 2.7 GHz (4/3), including polarisation.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5:  $S = 9.2 \pm 0.5$  Jy) including polarisation and review of flux densities.

Petriella *et al.* 2013, A&A, 554, A73. Chandra observations, plus CO and H<sub>I</sub> observations of region.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS Hi absorption observations.

## G21.0 - 0.4

**RA:**  $18^{\text{h}}31^{\text{m}}12^{\text{s}}$  **1-GHz flux/Jy:** 1.1 **Size/arcmin:**  $9 \times 7$  **Dec:**  $-10^{\circ}47'$  **Spectral index:** 0.6 **Type:** S

Has been called G21.04-0.47.

Radio: Shell.
References:

Brogan et al. 2006, ApJ, 639, L25. VLA at 330 MHz (42'': S =

 $2.3\pm0.2$  Jy), plus other observations.

# G21.5-0.9

RA: 18h33m33s 1-GHz flux/Jy: 7 Size/arcmin: 5 Dec: -10°35′ Spectral index: varies Type: C

Early observations relate to the central core only. **Radio:** Filled-centre, with high frequency turnover.

X-ray: Central core, with extended, faint halo.

Point sources: Central pulsar.

**Distance:** HI absorption suggests 4.4 kpc.

#### References:

Wilson & Weiler 1976, A&A, 53, 89. WSRT at 5 GHz ( $6'' \times 35''$ ). Becker & Kundu 1976, ApJ, 204, 427. NRAO interferometer at 2.7 GHz ( $5'' \times 20''$ ) and 8 GHz ( $2'' \times 7''$ ), plus review of flux densities.

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'6)

Becker & Szymkowiak 1981, ApJ, 248, L23. VLA at 5 GHz (8"), and Einstein observations.

Davelaar et al. 1986, ApJ, 300, L59. EXOSAT X-ray spectrum, and VI.A HI observations

Morsi & Reich 1987, A&AS, 69, 533. Effelsberg 100-m at 32 GHz  $(26\%5:S=5.64\pm0.29 \text{ Jy})$ .

Fürst *et al.* 1988, PASJ, 40, 347. NRO array at 22.3 GHz (4".4×7".3).

Junkes et al. 1988, LNP, 316, 134. Effelsberg 100-m at 2.7 GHz (4'.3), including polarisation.

Salter *et al.* 1989, A&A, 225, 167. Observations at 90.7 (29".5:  $S = 3.8 \pm 0.4$  Jy) and 141.9 GHz ( $S = 2.5 \pm 1.2$  Jy).

Salter *et al.* 1989, ApJ, 338, 171. NRAO 12-m at 84.2 GHz ( $S = 3.94 \pm 0.70$  Jy), plus review of flux densities.

Asaoka & Koyama 1990, PASJ, 42, 625. Ginga X-ray spectrum. Kassim 1992, AJ, 103, 943. VLA at 327 MHz (3'.3×3'.4:S=9 Jy). Wallace *et al.* 1994, A&A, 286, 565. HI of surroundings.

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Slane *et al.* 2000, ApJ, 533, L29. Chandra observations identifying X-ray halo.

Warwick et al. 2001, A&A, 365, L248. XMM-Newton observations of X-ray halo.

Bock et al. 2001, ApJ, 561, L203. BIMA at 94 GHz (4".6×8".6).

Safi-Harb et al. 2001, ApJ, 561, 308. Chandra and other X-ray observations

Reich 2002, in NSPS, p1. Effelsberg 100-m at 22 GHz (8'') and 32 GHz, including polarisation.

La Palombara & Mereghetti 2002, A&A, 383, 916. XMM-Newton upper limit on pulsations.

Bocchino et al. 2005, A&A, 442, 539. XMM-Newton and Chandra observations.

Gupta et al. 2005, CSci, 89, 853. Pulsar discovery.

Camilo et al. 2006, ApJ, 637, 456. Pulsar discovery.

Bietenholz & Bartel 2008, MNRAS, 386, 1411. VLA at 4.75 GHz (0".53×0".82) for expansion studies.

Tian & Leahy 2008, MNRAS, 391, L54. VGPS at 1.4 GHz (1') including Hi.

Matheson & Safi-Harb 2010, ApJ, 724, 572. Chandra observations. Bietenholz *et al.* 2011, MNRAS, 412, 1221. VLA at 1.4 GHz (14"×18").

Bhatnagar et al. 2011, ApJ, 739, L20. VLA at 6 GHz.

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9 $.5:S=6.5\pm0.4$  Jy) including polarisation and review of flux densities.

Zajczyk et al. 2012, A&A, 542, A12. IR observations, including polarisation.

Nynka et al. 2014, ApJ, 789, 72. NuSTAR observations.

Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 70, 100 and 143 GHz.

Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations

Aharonian et al. 2018, PASJ, 70, 38. Hitomi observations.

### G21.6 - 0.8

**RA:** 18<sup>h</sup>33<sup>m</sup>40<sup>s</sup> **1-GHz flux/Jy:** 1.4 **Size/arcmin:** 13 **Dec:** -10°25′ **Spectral index:** 0.5? **Type:** S

Has been called G21.64–0.84. **Radio:** Faint, irregular shell.

## References:

Bietenholz *et al.* 2011, MNRAS, 412, 1221. VLA at 327 MHz (85'': S=2.8 Jy) and 1.4 GHz  $(14'' \times 18'')$ .

Froebrich et al. 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.

G21.8-0.6 Kes 69

**RA:** 18<sup>h</sup>32<sup>m</sup>45<sup>s</sup> **1-GHz flux/Jy:** 65 **Size/arcmin:** 20 **Dec:** -10°08' **Spectral index:** 0.56 **Type:** S

Radio: Incomplete shell.

X-ray: Detected.

**Distance:** Association with CO indicates 5.2 kpc, and H<sub>I</sub> absorption suggests 5.6 kpc.

#### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Wilson 1972, A&A, 19, 354. H<sub>2</sub>CO absorption.

Kundu *et al.* 1974, AJ, 79, 132. NRAO 140-ft at 5 GHz (6') and 10 GHz (3').

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft et 2.7 GHz (5':S=42.3±4.6 Jy).

Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8'4) and 5 GHz (4'4).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2.6).

Seward 1990, ApJS, 73, 781. Einstein observations.

Kassim 1992, AJ, 103, 943. VLA at 327 MHz  $(3.3 \times 3.5 : S = 132 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations, including masers.

Yusef-Zadeh et al. 2003, ApJ, 585, 319. X-ray observations. Reach et al. 2006, AJ, 131, 1479. Spitzer observations.

Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers.

Wood *et al.* 2008, AJ, 135, 2358. VLA at 4.8 GHz, including polarisation.

Tian & Leahy 2008, MNRAS, 391, L54. VGPS at 1.4 GHz (1 $^{\prime}$ ) including Hi.

Zhou *et al.* 2009, ApJ, 691, 516. CO and HCO<sup>+</sup> of region. Hewitt *et al.* 2009, ApJ, 694, 1266. Spitzer spectroscopy.

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 24.0±1.3 Jy) including polarisation and review of flux densities.

Andersen *et al.* 2011, ApJ, 742, 7. Spitzer observations. Pihlström *et al.* 2014, AJ, 147, 73. VLA search for methanol masers.

Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS H<sub>I</sub> absorption observations.

Sezer et al. 2018, MNRAS, 481, 1416. Suzaku observations.

# G22.7-0.2

**RA:** 18<sup>h</sup>33<sup>m</sup>15<sup>s</sup> **1-GHz flux/Jy:** 33 **Size/arcmin:** 26 **Dec:** -09°13′ **Spectral index:** 0.6 **Type:** S?

**Radio:** Non-thermal ring in complex region, overlapping G23.3–0.3.

**X-ray:** Possible detection.

**Point sources:** Variable radio source near centre, and  $\gamma$ -ray source near edge.

**Distance:** Association with CO indicates 4.4 kpc, and H<sub>I</sub> absorption suggests 4.7 kpc.

## References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'.6).

Kassim 1992, AJ, 103, 943. VLA at 327 MHz ( $3'.3 \times 3'.4 : S = 82 \text{ Jy}$ ).

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Reach *et al.* 2006, AJ, 131, 1479. Spitzer observations.

Becker et al. 2010, AJ, 140, 157. Variable radio source detection.

Andersen *et al.* 2011, ApJ, 742, 7. Spitzer observations. Su *et al.* 2014, ApJ, 796, 122. CO observations.

Su *et al.* 2015, ApJ, 811, 134. CO observations of region.

Abramowski *et al.* 2015, MNRAS, 446, 1163. HESS observations. Mori *et al.* 2017, ApJ, 848, 80. NuSTAR, Chandra and XMM-Newton observations.

Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS H<sub>I</sub> absorption observations.

G23.3-0.3

**RA:** 18<sup>h</sup>34<sup>m</sup>45<sup>s</sup> **1-GHz flux/Jy:** 70 **Size/arcmin:** 27 **Dec:** -08°48′ **Spectral index:** 0.5 **Type:** S

**Radio:** Distorted ring, in complex region, overlapping G22.7–0.2.

**X-ray:** Possible extended emission, with compact sources.

Point sources: Pulsar association suggested.

**Distance:** HI and CO observations suggest 4.2 or 4.8 kpc.

#### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2/6)

Kassim 1992, AJ, 103, 943. VLA at 327 MHz  $(3'.3 \times 3'.4 : S = 138 \text{ Jy})$ .

Gaensler & Johnston 1995, MNRAS, 275, L73. Possible pulsar association.

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Aharonian *et al.* 2005, Science, 307, 1938. HESS detection. Reach *et al.* 2006, AJ, 131, 1479. Spitzer possible detection. Aharonian *et al.* 2006, ApJ, 636, 777. HESS observations. Albert *et al.* 2006, ApJ, 643, L53. γ-ray observations. Landi *et al.* 2006, ApJ, 651, 190. X-ray observations.

Tian *et al.* 2007, ApJ, 657, L25. VGPS at 1.4 GHz (1') including HI, plus XMM-Newton observations.

Leahy & Tian 2008, AJ, 135, 167. VGPS at 1.4 GHz (1') including HI, plus CO observations.

Mukherjee et al. 2009, ApJ, 691, 1707. XMM-Newton and SWIFT observations.

Frail *et al.* 2013, ApJ, 773, L19. OH observations. Castro *et al.* 2013, ApJ, 774, 36. Fermi observations.

Abramowski *et al.* 2015, A&A, 574, A27. HESS observations. Su *et al.* 2015, ApJ, 811, 134. CO observations of region.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.
Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

G24.7 - 0.6

**RA:** 18<sup>h</sup>38<sup>m</sup>43<sup>s</sup> **1-GHz flux/Jy:** 8 **Size/arcmin:** 15? **Dec:** -07°32′ **Spectral index:** 0.5 **Type:** S?

**Radio:** Incomplete shell, defined in SW.

**Distance:** HI absorption and CO suggests 3.8 kpc.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 12.3 Jy) and Parkes 64-m at 5 GHz (4': S = 3.6 Jy).

Angerhofer *et al.* 1977, A&A, 55, 11. NRAO 140-ft at 5 GHz (6'.8:  $S = 2.2 \pm 0.5$  Jy).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2.6).

Dubner *et al.* 1993, AJ, 105, 2251. VLA at 1.47 GHz  $(44'' \times 56'' : S = 1.9 \text{ Jy})$ .

Frail *et al.* 1996, AJ, 111, 1651. OH emission near remnant. Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search.

Ranasinghe & Leahy 2018, MNRAS, 477, 2243. VGPS HI absorption, plus CO observations.

G24.7 + 0.6

**RA:** 18<sup>h</sup>34<sup>m</sup>10<sup>s</sup> **1-GHz flux/Jy:** 20? **Size/arcmin:** 30×15 **Dec:** -07°05′ **Spectral index:** 0.2? **Type:** C?

Radio: Filled-centre, with faint shell, and a compact HII region to the S.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2/6)

Reich *et al.* 1984, A&A, 133, L4. Effelsberg 100-m at 2.7 GHz  $(4.3:S=19\pm3 \text{ Jy})$  and 4.75 GHz  $(2.4:S=17\pm4 \text{ Jy})$  and NRO 45-m at 10.2 GHz  $(2.7:S=15\pm3 \text{ Jy})$ .

Becker & Helfand 1987, ApJ, 316, 660. VLA at 1.4 GHz (12"), and X-ray upper limit.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. Acero *et al.* 2016, ApJS, 224, 8. Fermi observations.

## G25.1-2.3

**RA:**  $18^{\rm h}45^{\rm m}10^{\rm s}$  **1-GHz flux/Jy:** 8 **Size/arcmin:**  $80\times30$ ? **Dec:**  $-08^{\circ}00'$  **Spectral index:** 0.5? **Type:** S

**Radio:** Incomplete shell, extent not well defined.

#### References:

Gao *et al.* 2011, A&A, 532, A144. Urumqi 25-m at 5 GHz (9.5:  $S = 3.7 \pm 0.4$  Jy), plus other observations.

**G27.4+0.0** 4C-04.71

**RA:** 18<sup>h</sup>41<sup>m</sup>19<sup>s</sup> **1-GHz flux/Jy:** 6 **Size/arcmin:** 4 **Dec:** -04°56′ **Spectral index:** 0.68 **Type:** S

Early references refer to G27.3–0.1 (Kes 73), a supposed larger remnant.

Radio: Incomplete shell.

X-ray: Diffuse emission, with central low period pulsar.

**Point sources:** Central AXP.

**Distance:** HI absorption suggests 7.5 to 9.8 kpc or 5.8 kpc, and association with CO suggests 9 kpc.

#### References:

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz (5').

Clark *et al.* 1975, AuJPA, 37, 75. Molonglo at 408 MHz (3': S = 4.4 Jy).

Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8.4) and 5 GHz (4.4).

Angerhofer *et al.* 1977, A&A, 55, 11. NRAO 140-ft at 5 GHz (6.8:  $S = 2.0 \pm 0.5$  Jy).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2.6).

Caswell *et al.* 1982, MNRAS, 200, 1143. FIRST at 1415 MHz  $(45'' \times 60'' : S = 3.5 \text{ Jy})$ . Revise  $S_{408 \text{ MHz}} = 10.4 \text{ Jy}$ , and  $S_{5 \text{ GHz}} = 1.9 \pm 0.2 \text{ Jy}$ .

Kriss et al. 1985, ApJ, 288, 703. Einstein observations, plus VLA at 1.4 and 5 GHz (12").

Sanbonmatsu & Helfand 1992, AJ, 104, 2189. VLA at 1.4 GHz for HI absorption.

Helfand *et al.* 1994, ApJ, 434, 627. ROSAT observations, particularly of central source.

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations, including nearby masers.

Vasisht & Gotthelf 1997, ApJ, 486, L129. ASCA detection of pulsar.

Gotthelf & Vasisht 1997, ApJ, 486, L133. ASCA observations.

Gotthelf *et al.* 1999, ApJ, 522, L49. X-ray timing observations of pulsar.

Sugizaki et al. 2001, ApJS, 134, 77. ASCA survey observations. Mereghetti et al. 2001, MNRAS, 321, 143. Search for optical/IR counterpart to pulsar.

Kuiper et al. 2004, ApJ, 613, 1173. X-ray observations of pulsar. Tian & Leahy 2008, ApJ, 677, 292. VGPS at 1.4 GHz (1') including HI.

Mizuno *et al.* 2010, AJ, 139, 1542. Spitzer observations.

An *et al.* 2013, ApJ, 779, 163. NuSTAR and other observations.

Kumar *et al.* 2014, ApJ, 781, 41. Chandra and XMM-Newton observations

Froebrich et al. 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. Kilpatrick et al. 2016, ApJ, 816, 1. CO observations, including broad lines.

Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Yeung et al. 2017, ApJ, 837, 69. Fermi observations.

Borkowski & Reynolds 2017, ApJ, 846, 13. Chandra expansion studies.

Liu et al. 2017, ApJ, 851, 37. Fermi and CO observations.Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

## G27.8+0.6

**RA:**  $18^{\text{h}}39^{\text{m}}50^{\text{s}}$  **1-GHz flux/Jy:** 30 **Size/arcmin:**  $50 \times 30$  **Dec:**  $-04^{\circ}24'$  **Spectral index:** varies **Type:** F

**Radio:** Filled-centre, with spectral turnover.

**X-ray:** Possible pulsar wind nebula.

### References:

Reich *et al.* 1984, A&A, 133, L4. Effelsberg 100-m at 2.7 GHz  $(4.3:S=23\pm2 \text{ Jy})$ , and 4.75 GHz  $(2.4:S=18\pm2 \text{ Jy})$  and NRO 45-m at 10.2 GHz (smoothed to  $4.3:S=8.5\pm2 \text{ Jy}$ ).

Frail et al. 1996, AJ, 111, 1651. OH emission near remnant. Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Casandjian & Grenier 2008, A&A, 489, 849. γ-ray observations. Misanovic *et al.* 2010, ApJ, 725, 931. XMM-Newton pulsar/wind nebula search.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9 $.5:S=21.0\pm1.1$  Jy) including polarisation and review of flux densities. Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.

# G28.6 - 0.1

**RA:**  $18^{\text{h}}43^{\text{m}}55^{\text{s}}$  **1-GHz flux/Jy:** 3? **Size/arcmin:**  $13 \times 9$  **Dec:**  $-03^{\circ}53'$  **Spectral index:** ? **Type:** S

**Radio:** Poorly defined regions of non-thermal emission. **X-ray:** Diffuse shell, with thermal and non-thermal emission.

Point sources: Pulsar in NE.

**Distance:** HI absorption and CO suggests 9.6 kpc.

#### References:

Helfand *et al.* 1989, ApJ, 341, 151. VLA at 1.4 GHz (15") and 5 GHz (15").

Bamba et al. 2001, PASJ, 53, L21. ASCA observations. Ueno et al. 2003, ApJ, 588, 338. Chandra observations.

Ebisawa et al. 2005, ApJ, 635, 214. Chandra observations.Zyuzin et al. 2018, MNRAS, 476, 2177. Pulsar observations.Ranasinghe & Leahy 2018, MNRAS, 477, 2243. VGPS HI absorption, plus CO observations.

## G28.8 + 1.5

**RA:** 18<sup>h</sup>39<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** ? **Size/arcmin:** 100? **Dec:** -02°55′ **Spectral index:** 0.4? **Type:** S?

Radio: Part of rim detected.

X-ray: Diffuse, Centrally brightened.

#### References:

Schwentker 1994, A&A, 286, L47. ROSAT observations. Song *et al.* 2000, PASJ, 52, 181. ASCA observations. Casandjian & Grenier 2008, A&A, 489, 849. γ-ray observations.

Misanovic et al. 2010, ApJ, 725, 931. XMM-Newton pulsar/wind nebula search.

Froebrich et al. 2015, MNRAS, 454, 2586. H2 IR observations.

# G29.6+0.1

**RA:** 18<sup>h</sup>44<sup>m</sup>52<sup>s</sup> **1-GHz flux/Jy:** 1.5? **Size/arcmin:** 5 **Dec:** -02°57' **Spectral index:** 0.5? **Type:** S

Radio: Diffuse shell.

Point sources: AXP associated.

## References:

Gaensler et al. 1999, ApJ, 526, L37. VLA at 5 GHz (13") and 8 GHz (8")

Vasisht *et al.* 2000, ApJ, 542, L49. X-ray observations of AXP. Kilpatrick *et al.* 2016, ApJ, 816, 1. CO observations, including broad lines.

G29.7-0.3 Kes 75

**RA:** 18<sup>h</sup>46<sup>m</sup>25<sup>s</sup> **1-GHz flux/Jy:** 10 **Size/arcmin:** 3 **Dec:** -02°59′ **Spectral index:** 0.63 **Type:** C

Has erroneously been called G29.6+0.1.

**Radio:** Shell with flatter spectrum emission from centre.

**X-ray:** Thermal shell and non-thermal core, and central pulsar.

**Point sources:** X-ray pulsar.

**Distance:** Association with CO implies 11 kpc, and H<sub>I</sub> absorption suggests 5.6 kpc.

#### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Milne & Dickel 1974, AuJPh, 27, 549. Parkes 64-m at 2.7 GHz  $(8.4:S=5\pm20\% \text{ Jy})$ .

Becker & Kundu 1975, AJ, 80, 679. NRAO 140-ft at 10.6 GHz (3').

Becker & Kundu 1976, ApJ, 204, 427. NRAO interferometer at 2.7 GHz ( $7'' \times 20''$ ) and 8 GHz ( $8'' \times 25''$ ), plus review of flux densities.

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'.6).

Becker et al. 1983, ApJ, 268, L93. VLA at 1.4 GHz (3"), plus Einstein observations.

Becker & Helfand 1984, ApJ, 283, 154. VLA at 5 GHz (2".6), plus

Morsi & Reich 1987, A&AS, 71, 189. Effelsberg 100-m at 32 GHz  $(26\%5: S=1.02\pm0.07 \text{ Jy})$ .

Salter et al. 1989, ApJ, 338, 171. NRAO 12-m at 84.2 GHz of core, plus review of flux densities.

Kassim 1992, AJ, 103, 943. VLA at 327 MHz  $(3'.4 \times 3'.7 : S = 27.4 \text{ Jy})$ .

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search.
Blanton & Helfand 1996, ApJ, 470, 961. ASCA observations.
Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.
Gotthelf et al. 2000, ApJ, 542, L37. X-ray pulsar detection.
Sugizaki et al. 2001, ApJS, 134, 77. ASCA survey observations.

Mereghetti et al. 2002, ApJ, 574, 873. BeppoSAX observations of pulsar.

Helfand et al. 2003, ApJ, 582, 783. Chandra observations.

Bock & Gaensler 2005, ApJ, 626, 343. BIMA at 88.6 GHz  $(10'' \times 13'')$ .

Morton et al. 2007, ApJ, 667, 219. Spitzer and Chandra observations.

McBride et al. 2008, A&A, 477, 249. INTEGRAL and Chandra observations.

Leahy & Tian 2008, A&A, 480, L25. VGPS at 1.4 GHz (1 $^{\prime}$ ) including Hı.

Kumar & Safi-Harb 2008, ApJ, 678, L43. Chandra observations. Ng *et al.* 2008, ApJ, 686, 508. Chandra observations.

Gavriil et al. 2008, Science, 319, 1802. X-ray observations of pul-

Su et al. 2009, ApJ, 694, 376. CO observations of region, plus Chandra observations.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9!5 : S =  $3.6\pm0.6$  Jy) including polarisation and review of flux densities.

Kilpatrick *et al.* 2016, ApJ, 816, 1. CO observations, including broad lines.

Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

Reynolds *et al.* 2018, ApJ, 856, 133. Chandra observations for expansion and brightness change studies.

## G30.7 - 2.0

**RA:** 18<sup>h</sup>54<sup>m</sup>25<sup>s</sup> **1-GHz flux/Jy:** 0.5? **Size/arcmin:** 16 **Dec:** -02°54′ **Spectral index:** 0.7? **Type:** ?

Radio: Poorly defined.

### References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4/3).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

G30.7 + 1.0

**RA:**  $18^{\rm h}44^{\rm m}00^{\rm s}$  **1-GHz flux/Jy:** 6 **Size/arcmin:**  $24 \times 18$  **Dec:**  $-01^{\circ}32'$  **Spectral index:** 0.4 **Type:** S?

**Radio:** Non-thermal, highly polarised part shell? **Point sources:** Compact source near centre.

## References:

Reich *et al.* 1986, A&A, 155, 185. Effelsberg 100-m at 4.75 GHz  $(2.4:S=3.4\pm0.4 \text{ Jy})$ , plus other flux densities.

Kassim 1992, AJ, 103, 943. VLA at 327 MHz ( $3.6 \times 3.9 : S = 8.6 \text{ Jy}$ ).

Gorham *et al.* 1996, ApJ, 458, 257. Pulsar search. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 2.93±0.19 Jy) including polarisation and review of flux densities.

## G31.5 - 0.6

RA: 18<sup>h</sup>51<sup>m</sup>10<sup>s</sup> 1-GHz flux/Jy: 2? Size/arcmin: 18? Dec: -01°31′ Spectral index: ? Type: S?

Has been called G31.55–0.65.

Radio: Distorted shell? near HII region. Optical: Diffuse, incomplete shell.

#### References:

Fürst *et al.* 1987, A&AS, 69, 403. Effelsberg 100-m at 4.75 GHz (2.4), plus other flux densities.

Mavromatakis et al. 2001, A&A, 370, 265. Optical observations.

G31.9+0.0 3C391

**RA:**  $18^{\rm h}49^{\rm m}25^{\rm s}$  **1-GHz flux/Jy:** 25 **Size/arcmin:**  $7\times5$  **Dec:**  $-00^{\circ}55'$  **Spectral index:** varies **Type:** S

**Radio:** Shell, brightest in NW, with low frequency turnover.

**X-ray:** Diffuse with central core.

**Distance:** HI absorption and CO association suggests 7.1 kpc.

#### References:

Radhakrishnan *et al.* 1972, ApJS, 24, 49. Hi absorption. Becker & Kundu 1975, AJ, 80, 679. NRAO 140-ft at 10.6 GHz (3').

Green *et al.* 1975, A&A, 44, 187. Effelsberg 100-m at 15 GHz (58'': S > 1.6).

Becker & Kundu 1976, ApJ, 204, 427. NRAO interferometer at 2.7 GHz (20"×23") and 8 GHz (9"×24"), plus review of flux densities

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz

Goss *et al.* 1979, A&A, 78, 75. FIRST at 1.4 GHz ( $54'' \times 66''$ :  $S = 20 \pm 2$  Jy) and Effelsberg 100-m at 10.7 GHz (77'':  $S = 7.5 \pm 0.8$  Jy).

Wang & Seward 1984, ApJ, 279, 705. Einstein observations. Kassim 1992, AJ, 103, 943. VLA at 327 MHz (3'.4×3'.7: S = 44.8 Jy).

Reynolds & Moffett 1993, AJ, 105, 2226. VLA at 1.4 GHz (6"), including possible associated CO.

Moffett & Reynolds 1994, ApJ, 425, 668. VLA at 330 MHz (smoothed to  $30'': S = 38.5 \pm 0.5$  Jy) 1.46 GHz (6".7) and 4.85 GHz (6".2×6".4), including spectral index and polarisation studies.

Frail et al. 1996, AJ, 111, 1651. OH emission near remnant, including masers.

Gorham et al. 1996, ApJ, 458, 257. Pulsar search.

Reach & Rho 1996, A&A, 315, L277. ISO spectroscopy.

Rho & Petre 1996, ApJ, 467, 698. ROSAT observations.

Wilner *et al.* 1998, AJ, 115, 247. CO observations of surroundings. Reach & Rho 1998, ApJ, 507, L93. ISO observations.

Reach & Rho 1999, ApJ, 511, 836. CO, HCO+ and CS observations of surroundings.

Reach & Rho 2000, ApJ, 544, 843. ISO observations of interactions with surroundings.

see also: Reach & Rho 2001, ApJ, 558, 943. Erratum.

Sugizaki *et al.* 2001, ApJS, 134, 77. ASCA survey observations. Chen & Slane 2001, ApJ, 563, 202. ASCA observations. Reach *et al.* 2002, ApJ, 564, 302. Observations of shocked molecular species.

Chen *et al.* 2004, ApJ, 616, 885. Chandra observations. Brogan *et al.* 2005, AJ, 130, 148. VLA at 74 MHz, 330 MHz and 1.5 GHz (70").

Kawasaki *et al.* 2005, ApJ, 631, 935. ASCA observations. Reach *et al.* 2006, AJ, 131, 1479. Spitzer observations. Neufeld *et al.* 2007, ApJ, 664, 890. Spitzer observations. Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers

Castro & Slane 2010, ApJ, 717, 372. Fermi observations. Yuan & Neufeld 2011, ApJ, 726, 76. Spitzer observations. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5:  $S = 8.9 \pm 0.6$  Jy) including polarisation and review of flux densities. Pihlström *et al.* 2014, AJ, 147, 73. VLA search for methanol

Neufeld *et al.* 2014, ApJ, 781, 102. Spitzer and Herschel IR spectroscopy.

Ergin et al. 2014, ApJ, 790, 65. Fermi and Suzaku observations. Gusdorf et al. 2014, IAUS, 296, 178. CO observations.

Su *et al.* 2014, IAUS, 296, 372. VGPS for H<sub>I</sub> absorption. Lee *et al.* 2014, MNRAS, 443, 2650. [Fe<sub>II</sub>] IR survey observations.

Sato et al. 2014, PASJ, 66, 124. Suzaku observations.
Froebrich et al. 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.
Kilpatrick et al. 2016, ApJ, 816, 1. CO observations, including broad lines.

Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Ranasinghe & Leahy 2017, ApJ, 843, 119. HI and CO observa-

**G32.0–4.9** 3C396.1

**RA:** 19<sup>h</sup>06<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** 22? **Size/arcmin:** 60? **Dec:** -03°00′ **Spectral index:** 0.5? **Type:** S?

Radio: Possible large shell?

#### References:

Milne & Hill 1969, AuJPh, 22, 211. Parkes 64-m at 635 MHz  $(31': S = 25\pm30\% \text{ Jy})$ , 1410 MHz  $(15': S = 19\pm15\% \text{ Jy})$  and 2650 MHz  $(8'.4: S = 8.6\pm30\% \text{ Jy})$ . Fluxes if size is 60', plus review of flux densities.

Caswell 1970, AuJPh, 23, 105. Revision of low frequency flux densities.

Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo at 111 MHz ( $S = 105\pm30$  Jy).

G32.1 - 0.9

RA: 18<sup>h</sup>53<sup>m</sup>10<sup>s</sup> 1-GHz flux/Jy: ? Size/arcmin: 40? Dec: -01°08′ Spectral index: ? Type: C?

Radio: Possible faint shell, not well defined.

**X-ray:** Diffuse, with clumps.

#### References:

Folgheraiter *et al.* 1997, MNRAS, 292, 365. ROSAT and ASCA observations.

Froebrich et al. 2015, MNRAS, 454, 2586. H2 IR observations.

G32.4+0.1

RA: 18<sup>h</sup>50<sup>m</sup>05<sup>s</sup> 1-GHz flux/Jy: 0.25? Size/arcmin: 6 Dec: -00°25′ Spectral index: ? Type: S

Has been called G32.45+0.1.

Radio: Shell. X-ray: Shell.

**Distance:** X-ray absorption suggests 17 kpc.

References:

Yamaguchi et al. 2004, PASJ, 56, 1059. XMM-Newton and other observations.

Ueno et al. 2005, in XRRC, E4.18. XMM-Newton observations.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.Kilpatrick et al. 2016, ApJ, 816, 1. CO observations, including broad lines.

**G32.8–0.1** Kes 78

**RA:**  $18^h51^m25^s$  **1-GHz flux/Jy:** 11? **Size/arcmin:**  $22\times15$  **Dec:**  $-00^\circ08'$  **Spectral index:** 0.2? **Type:** S?

Part has been called G33.1–0.1.

Radio: Incomplete, elongated shell.

Optical: Detected.

**X-ray:** Patchy, elongated shell.

**Distance:** Association with CO and H<sub>I</sub> absorption indicate 4.8 kpc.

#### References:

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft et 2.7 GHz (5': $S=7.2\pm0.5$  Jy).

Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo at 430 MHz ( $S = 19.0 \pm 15.5$  Jy).

Becker & Kundu 1975, AJ, 80, 679. NRAO 140-ft at 10.6 GHz (3').

Caswell *et al.* 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3':S=12.8 Jy) and Parkes 64-m at 5 GHz (4':S=7.7 Jy).

Altenhoff et al. 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'6)

Kassim 1992, AJ, 103, 943. VLA at 327 MHz ( $3.6 \times 3.8 : S = 31.3 \text{ Jy}$ ).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

Koralesky et al. 1998, AJ, 116, 1323. VLA detection of compact OH emission.

Stupar & Parker 2011, MNRAS, 414, 2282. H $\alpha$  observations. Zhou & Chen 2011, ApJ, 743, 4. XMM-Newton observations, plus CO of region.

Pihlström *et al.* 2014, AJ, 147, 73. VLA search for methanol masers.

G33.2-0.6

RA: 18<sup>h</sup>53<sup>m</sup>50<sup>s</sup> 1-GHz flux/Jy: 3.5 Size/arcmin: 18 Dec: -00°02′ Spectral index: varies Type: S

Radio: Incomplete shell.

Optical: Filaments and diffuse emission.

## References:

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'.6).

Reich 1982, A&A, 106, 314. Effelsberg 100-m at 2.7 GHz (4.4:  $S = 2.6 \pm 0.3$  Jy) and 4.75 GHz (2.5:  $S = 1.75 \pm 0.2$  Jy).

Dubner *et al.* 1996, AJ, 111, 1304. VLA at 1.4 GHz  $(52'' \times 68'' : S = 2.7 \pm 0.3 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations. Boumis et al. 2009, A&A, 499, 789. Optical observations. Froebrich et al. 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.

**G33.6+0.1** Kes 79, 4C00.70, HC13

**RA:** 18<sup>h</sup>52<sup>m</sup>48<sup>s</sup> **1-GHz flux/Jy:** 20 **Size/arcmin:** 10 **Dec:** +00°41′ **Spectral index:** 0.51 **Type:** S

Has been called G33.7+0.0.

**Radio:** Shell, with bright central region, in complex region.

**X-ray:** Multiple shells and filaments. **Point sources:** Central X-ray pulsar.

**Distance:** HI absorption suggests 7.8 kpc or 3.5 kpc, and CO observations suggest 5.5 kpc.

### References:

Caswell *et al.* 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3':S=35.5 Jy) and Parkes 64-m at 5 GHz (4':S=7.8 Jy).

Caswell et al. 1975, A&A, 45, 239. Hi absorption.

Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo  $S_{430~\mathrm{MHz}} = 69 \pm 33~\mathrm{Jy}.$ 

Becker & Kundu 1975, AJ, 80, 679. NRAO 140-ft at 10.6 GHz (3':S=6.8±1.5 Jy).

Angerhofer *et al.* 1977, A&A, 55, 11. NRAO 140-ft at 5 GHz (6.8:  $S = 11.4 \pm 1.1$  Jy).

Caswell *et al.* 1981, MNRAS, 195, 89. FIRST at 1415 MHz (1'), plus observations of the nearby point source.

van Gorkom *et al.* 1982, MNRAS, 198, 757. WSRT HI absorption to nearby point source, possibly extragalactic.

Seaquist & Gilmore 1982, AJ, 87, 378. VLA observations of nearby source, plus Einstein observations.

Green 1989, MNRAS, 238, 737. OH absorption.

Frail & Clifton 1989, ApJ, 336, 854. VLA at 1.4 GHz ( $1'\times2'.9$ ), including H<sub>I</sub> absorption.

Velusamy  $\widetilde{et}$  al. 1991, AJ, 102, 676. VLA at 327 MHz (1'), 1.5 (7"×14") and 5 GHz (7"), including spectral comparison.

Green & Dewdney 1992, MNRAS, 254, 686. Observations of adjacent molecular material.

Kassim 1992, AJ, 103, 943. VLA at 327 MHz ( $3'.6 \times 3'.8 : S = 34.8 \text{ Jy}$ ).

Seward & Velusamy 1995, ApJ, 439, 715. ROSAT observations. Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

Koralesky et al. 1998, AJ, 116, 1323. VLA search for OH emission.

Sugizaki *et al.* 2001, ApJS, 134, 77. ASCA survey observations. Tsunemi & Enoguchi 2002, PASJ, 54, 735. ASCA observations. Seward *et al.* 2003, ApJ, 584, 414. Chandra observations. Stanimirović 2003, ApJ, 592, 953. Arecibo OH absorption. Sun *et al.* 2004, ApJ, 605, 742. Chandra observations. Gotthelf *et al.* 2005, ApJ, 627, 390. XMM-Newton pulsar detec-

Reach *et al.* 2006, AJ, 131, 1479. Spitzer possible detection. Halpern *et al.* 2007, ApJ, 665, 1304. X-ray pulsar timing observa-

Giacani *et al.* 2009, A&A, 507, 841. VLA at 74 MHz  $(36'' \times 39'' : S = 76 \pm 10 \text{ Jy})$ , 324 MHz  $(13'' : S = 39 \pm 8 \text{ Jy})$  and 1.5 GHz  $(17'' \times 19'' : S = 11.5 \pm 1.5 \text{ Jy})$ , plus review of flux densities and XMM-Newton observations.

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz ( $9.5:S=9.4\pm0.5$  Jy) including polarisation and review of flux densities. Auchettl et al. 2014, ApJ, 783, 32. Fermi detection.

Kilpatrick *et al.* 2016, ApJ, 816, 1. CO observations, including broad lines.

Zhou et al. 2016, ApJ, 831, 192. CO and other observations.Sato et al. 2016, PASJ, 68, S8. Suzaku observations.Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations

Kuriki et al. 2018, ApJ, 864, 161. CO observations.

**G34.7–0.4** W44, 3C392

**RA:**  $18^h 56^m 00^s$  **1-GHz flux/Jy:** 240 **Size/arcmin:**  $35 \times 27$  **Dec:**  $+01^{\circ}22'$  **Spectral index:** 0.37 **Type:** C

Has been called G34.6–0.5.

**Radio:** Distorted shell, brighter to the E, with pulsar and associated nebula.

Optical: Diffuse emission.

X-ray: Centrally concentrated, thermal spectrum, plus pulsar wind nebula.

**Point sources:** Pulsar within the boundary of the remnant.

**Distance:** HI absorption suggests 3.0 kpc, and optical absorption suggests 2.1 kpc.

#### References:

Caswell *et al.* 1975, A&A, 45, 239. HI absorption. Clark *et al.* 1975, AuJPA, 37, 75. Molonglo at 408 MHz (3': *S* = 299 Jy).

Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo  $S_{430 \text{ MHz}} = 540 \pm 187 \text{ Jy}$ .

Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8.4) and 5 GHz (4.4).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2/6).

Smith *et al.* 1985, MNRAS, 217, 99. Einstein observations. Wolszczan *et al.* 1991, ApJ, 372, L99. Pulsar detection. Kassim 1992, AJ, 103, 943. VLA at 327 MHz (3'.6×3'.8: *S* = 469 Jy).

Jones et al. 1993, MNRAS, 265, 631. VLA at 1.4 GHz (15"), plus X-ray spectra.

Rho *et al.* 1994, ApJ, 430, 757. Optical and ROSAT observations. Koo & Heiles 1995, ApJ, 442, 679. HI of surrounding shell. Esposito *et al.* 1996, ApJ, 461, 820. Possible associated  $\gamma$ -ray emis-

Harrus *et al.* 1996, ApJ, 464, L161. ASCA observations. Frail *et al.* 1996, ApJ, 464, L165. VLA at 1.5 and 8.4 GHz (7".8×8".9) of pulsar nebula.

Frail *et al.* 1996, AJ, 111, 1651. OH maser emission. Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Reach & Rho 1996, A&A, 315, L277. ISO spectroscopy. Harrus *et al.* 1997, ApJ, 488, 781. X-ray observations. Claussen *et al.* 1997, ApJ, 489, 143. VLA of associated OH masers.

Giacani et al. 1997, AJ, 113, 1379. VLA at 1.4 GHz (15"), plus optical images.

Seta *et al.* 1998, ApJ, 505, 286. CO observations of surroundings. Cox *et al.* 1999, ApJ, 524, 179. Revision of distance.

Reach & Rho 2000, ApJ, 544, 843. ISO observations of interactions with surroundings.

see also: Reach & Rho 2001, ApJ, 558, 943. Erratum. Roberts et al. 2001, ApJS, 133, 451. ASCA observations. Petre et al. 2002, ApJ, 579, 404. Chandra observations of pulsar and wind nebula.

Mavromatakis et al. 2003, A&A, 405, 591. Optical observations.Shelton et al. 2004, ApJ, 611, 906. Chandra observations of part.Reach et al. 2005, ApJ, 618, 297. Molecular line and near-IR observations.

Hoffman et al. 2005, ApJ, 627, 803. Observations of OH masers.

Kawasaki *et al.* 2005, ApJ, 631, 935. ASCA observations. Reach *et al.* 2006, AJ, 131, 1479. Spitzer observations. Neufeld *et al.* 2007, ApJ, 664, 890. Spitzer observations. Castelletti *et al.* 2007, A&A, 471, 537. VLA at 74 MHz ( $36'' \times 39''$ :  $S = 634 \pm 70$  Jy) and 324 MHz (13'':  $S = 411 \pm 50$  Jy). Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers.

Abdo *et al.* 2010, Science, 327, 1103. Fermi observations. Yuan & Neufeld 2011, ApJ, 726, 76. Spitzer observations. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9/5: S=118±6 Jy) including polarisation and review of flux densities. Giuliani *et al.* 2011, ApJ, 742, L30. γ-ray observations. Uchiyama *et al.* 2012, ApJ, 749, L35. Fermi observations. Uchida *et al.* 2012, PASJ, 64, 141. Suzaku observations. Ackermann *et al.* 2013, Science, 339, 807. Fermi observations. Yoshiike *et al.* 2013, ApJ, 768, 179. CO and HI observations of region.

Sashida et al. 2013, ApJ, 774, 10. HCO+ and CO observations of region.

Park et al. 2013, ApJ, 777, 14. Arecibo HI observations of region. Pihlström et al. 2014, AJ, 147, 73. VLA search for methanol masers.

Anderl et al. 2014, A&A, 569, A81. CO observations of regions in NF

Su *et al.* 2014, IAUS, 296, 372. VGPS for H<sub>I</sub> absorption. Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 30, 44 and 70 GHz.

Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

McEwen et al. 2016, ApJ, 826, 189. NH<sub>3</sub> and CH<sub>3</sub>OH observations.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Génova-Santos et al. 2017, MNRAS, 464, 4107. Radio observations at 10 to 20 GHz.

Egron *et al.* 2017, MNRAS, 470, 1329. SRT at 1.5 GHz (11 $^\prime$  : S = 214 $\pm$ 6 Jy) and 7 GHz (2 $^\prime$ 7 : S = 94 $\pm$ 4 Jy).

Yamada *et al.* 2017, ApJ, 834, L3. CO and HCO<sup>+</sup> observations. Shan *et al.* 2018, ApJS, 238, 35. Optical absorption for distance. Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

G35.6 - 0.4

**RA:**  $18^h57^m55^s$  **1-GHz flux/Jy:** 9 **Size/arcmin:**  $15\times11$  **Dec:**  $+02^{\circ}13'$  **Spectral index:** 0.5 **Type:** S?

Re-identified as SNR in 2009.

**Radio:** Diffuse, with some limb brightening. **Distance:** HI absorption suggests 3.8 kpc.

References:

Green 2009, MNRAS, 399, 177. Identification in the radio as a

Paron & Giacani 2010, A&A, 509, L4. CO and IR observations of region.

Zhu et al. 2013, ApJ, 775, 95. HI and other observations.

Paredes *et al.* 2014, A&A, 561, A56. GMRT at 610 MHz (4"8×12"2).

Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

G36.6 - 0.7

Radio: polarised arc, possibly part of a larger shell?

References:

Fürst *et al.* 1987, A&AS, 69, 403. Effelsberg 100-m at 4.75 GHz (2'.4), plus other flux densities.

Kassim 1992, AJ, 103, 943. VLA at 327 MHz ( $3'.4 \times 3'.7 : S = 6.7 \text{ Jy}$ ).

Gorham et al. 1996, ApJ, 458, 257. Pulsar search.

Reach *et al.* 2006, AJ, 131, 1479. Spitzer possible detection. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5:  $S = 0.39 \pm 0.04$  Jy) including polarisation and review of flux densities

G36.6 + 2.6

**RA:** 18<sup>h</sup>48<sup>m</sup>49<sup>s</sup> **1-GHz flux/Jy:** 0.7? **Size/arcmin:** 17×13? **Dec:** +04°26′ **Spectral index:** 0.5? **Type:** S

Radio: Poorly resolved shell.

References

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4'.3).

G38.7 - 1.3

**RA:**  $19^h06^m40^s$  **1-GHz flux/Jy:** ? **Size/arcmin:**  $32\times19$ ? **Dec:**  $+04^\circ28'$  **Spectral index:** ? **Type:** S

G38.7–1.4 refers to the E portion.

Radio: Incomplete shell.

Optical: Arc of filaments, brighter to E.

**X-ray:** Detected in E.

References:

Schaudel et al. 2002, ASPC, 271, 391. ROSAT of E, and radio survey observations.

Sabin et al. 2013, MNRAS, 431, 279. Hα and radio survey observations.

Huang et al. 2014, ApJ, 785, 118. XMM-Newton and Chandra observations.

Froebrich et al. 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.

**G39.2–0.3** 3C396, HC24, NRAO 593

**RA:**  $19^h04^m08^s$  **1-GHz flux/Jy:** 18 **Size/arcmin:**  $8\times6$  **Dec:**  $+05^{\circ}28'$  **Spectral index:** 0.34 **Type:** C

**Radio:** Shell, brighter to W, with faint 'tail' to E. **X-ray:** Diffuse, brighter to W, with central core.

**Point sources:** Central X-ray source. **Distance:** HI absorption suggests 8.5 kpc.

#### References

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3'). Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo  $S_{430~\rm MHz}=54\pm38~\rm Jy.$ 

Becker & Kundu 1975, AJ, 80, 679. NRAO 140-ft at 10.6 GHz  $(3': S=4.1\pm1.0 \text{ Jy})$ .

Caswell et al. 1975, A&A, 45, 239. HI absorption.

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2/6).

Caswell *et al.* 1982, MNRAS, 200, 1143. FIRST at 1.4 GHz (48"×65").

Becker & Helfand 1987, AJ, 94, 1629. VLA at 1.4 GHz (12'': S = 14 Jy) and 5 GHz, plus Einstein observations.

Patnaik *et al.* 1990, A&A, 232, 467. VLA at 1.5 GHz (25") and 1.4 GHz (7".5×7"8) and 5 GHz (25") including polarisation, plus Ooty at 327 MHz (31"×100"), including review of flux densities.

Kassim 1992, AJ, 103, 943. VLA at 327 MHz  $(3.5 \times 3.6 : S = 42.5 \text{ Jy})$ .

Anderson & Rudnick 1993, ApJ, 408, 514. VLA at 1.45 and 4.89 GHz for spectral index studies.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Gorham *et al.* 1996, ApJ, 458, 257. Pulsar search. Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Koralesky *et al.* 1998, AJ, 116, 1323. VLA search for OH emission.

Harrus & Slane 1999, ApJ, 516, 811. ASCA observations. Aharonian *et al.* 2001, A&A, 375, 1008. HESS limit. Olbert *et al.* 2003, ApJ, 592, L45. Chandra observations. Reach *et al.* 2006, AJ, 131, 1479. Spitzer observations. Scaife *et al.* 2007, MNRAS, 377, L69. 33 GHz observations. Lee *et al.* 2009, ApJ, 691, 1042. IR observations. Hewitt *et al.* 2009, ApJ, 694, 1266. Spitzer spectroscopy. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9!5: *S* = 8.8±0.5 Jy) including polarisation and review of flux densities.

Su et al. 2011, ApJ, 727, 43. Chandra and CO observations of region.

Andersen *et al.* 2011, ApJ, 742, 7. Spitzer observations.
Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.
Kilpatrick *et al.* 2016, ApJ, 816, 1. CO observations, including broad lines.

Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Cruciani *et al.* 2016, MNRAS, 459, 4224. Parkes 64 m at 8.4, 13.5, 18.6 and 21.5 GHz, plus review of flux densities. Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

**G39.7–2.0** W50, SS433

**RA:**  $19^{\rm h}12^{\rm m}20^{\rm s}$  **1-GHz flux/Jy:** 85? **Size/arcmin:**  $120\times60$  **Dec:**  $+04^{\circ}55'$  **Spectral index:** 0.7? **Type:** ?

Eastern part has been called G40.0–3.1. Is this a SNR?

Radio: Elongated shell, containing SS433, adjacent to the H<sub>II</sub> region S74.

**Optical:** Faint filaments at the edge of the radio emission.

X-ray: Emission from SS433 and two lobes.

**Point sources:** SS433 is the compact source in the centre of the W50.

**Distance:** HI absorption and CO observations indicate 4.9 kpc.

### References:

van den Bergh 1980, ApJ, 236, L23. Optical in Hα and [SII]. Zealey *et al.* 1980, MNRAS, 192, 731. Optical spectra. van Gorkom *et al.* 1982, MNRAS, 198, 757. WSRT HI absorption to nearby point source (not SS433).

Seaquist & Gilmore 1982, AJ, 87, 378. VLA observations of nearby source.

Watson *et al.* 1983, ApJ, 273, 688. X-ray observations. Downes *et al.* 1986, MNRAS, 218, 393. Effelsberg 100-m at 4.75 GHz ( $2'.4:S=34\pm4$  Jy), plus previous 1.7 and 2.7 GHz data

Romney et al. 1987, ApJ, 321, 822. VLBI of SS433, including distance.

Elston & Baum 1987, AJ, 94, 1633. Mosaic with VLA at  $1.4~\mathrm{GHz}$  (30'') of fine structure only.

Kawai et al. 1989, PASJ, 41, 491. X-ray observations of SS433.
Band 1989, ApJ, 336, 937. Einstein and EXOSAT observations.
Yamauchi et al. 1994, PASJ, 46, L109. X-ray spectral observations.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Brinkmann *et al.* 1996, A&A, 312, 306. ROSAT observations.

Kotani *et al.* 1996, PASJ, 48, 619. X-ray line emission from SS433.

Safi-Harb & Ögelman 1997, ApJ, 483, 868. ROSAT observations. Dubner *et al.* 1998, AJ, 116, 1842. VLA at 328 MHz ( $60'' \times 64''$ :  $S = 160 \pm 20$  Jy), and 1.4 GHz ( $54'' \times 56''$ ), plus NRAO 140-ft at 1.4 GHz (21') for HI observations.

Safi-Harb & Petre 1999, ApJ, 512, 784. X-ray observations. Aharonian *et al.* 2001, A&A, 375, 1008. HESS limit. Brinkmann *et al.* 2007, A&A, 463, 611. XMM-Newton observations of E lobe.

Boumis *et al.* 2007, MNRAS, 381, 308. Optical observations. Lockman *et al.* 2007, MNRAS, 381, 881. Hi observations. Gao *et al.* 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9.5:

 $S=37\pm4$  Jy), including polarisation and review of flux densities. Farnes *et al.* 2017, MNRAS, 467, 4777. ATCA at 1.4 to 3.1 GHz, including polarisation and H $\alpha$  observations.

Broderick *et al.* 2018, MNRAS, 475, 5360. LOFAR at 115 to 189 MHz, including 140 MHz (55"×78").

Su et al. 2018, ApJ, 863, 103. CO and HI observations.

## G40.5 - 0.5

**RA:** 19h07m10s **1-GHz flux/Jy:** 11 **Size/arcmin:** 22 **Dec:** +06°31′ **Spectral index:** 0.4 **Type:** S

**Radio:** Shell, brightest to the NE. **Point sources:** Central pulsar.

#### References:

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2!6).

Downes *et al.* 1980, A&A, 92, 47. Effelsberg 100-m at 1.7 GHz  $(7.6: S = 9.3 \pm 1.3 \text{ Jy})$ , and 2.7 GHz  $(4.4: S = 7.2 \pm 0.5 \text{ Jy})$ , plus review of flux densities.

Aharonian *et al.* 2001, A&A, 375, 1008. HESS limit. Yang *et al.* 2006, ChJAA, 6, 210. CO observations of surroundings Abdo *et al.* 2007, ApJ, 664, L91. γ-ray observations. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5:  $S = 6.4 \pm 0.3$  Jy) including polarisation and review of flux densities. Aliu *et al.* 2014, ApJ, 787, 166. γ-ray observations of region. Lyne *et al.* 2017, ApJ, 834, 137. Pulsar detection.

G41.1-0.3

**RA:**  $19^h07^m34^s$  **1-GHz flux/Jy:** 25 **Size/arcmin:**  $4.5\times2.5$  **Dec:**  $+07^\circ08'$  **Spectral index:** 0.50 **Type:** S

Radio: 3C397 is two sources: the E is the SNR, the W is a HII region.

**X-ray:** Brighter to the E and W, with central component.

**Distance:** HI absorption suggest 8.5 kpc.

#### References:

Kundu *et al.* 1974, AJ, 79, 132. NRAO 140-ft at 10 GHz (3') and 5 GHz (6').

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz  $(5': S=21.3\pm1.2 \text{ Jy for both components})$ .

Caswell et al. 1975, A&A, 45, 239. Hi absorption.

Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo  $S_{430~\mathrm{MHz}} = 82 \pm 51~\mathrm{Jy}$ , also Algonquin 46-m at 10.6 GHz (3':  $S = 12 \pm 2~\mathrm{Jy}$ ), and Haystack 36-m at 15.5 GHz (2'.3:  $S = 8.5 \pm 3.0~\mathrm{Jy}$ ).

Caswell *et al.* 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3': S = 29.8 Jy) and Parkes 64-m at 5 GHz (4': S = 8.7 Jy).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2/6).

Caswell *et al.* 1982, MNRAS, 200, 1143. FIRST at 1.4 GHz (52"×58").

Becker *et al.* 1985, ApJ, 296, 461. VLA at 1.4 and 5 GHz (8"), plus Einstein observations.

Morsi & Reich 1987, A&AS, 71, 189. Effelsberg 100-m at 32 GHz (smoothed to 30'':  $S=4.10\pm0.19$  Jy).

Kassim 1992, AJ, 103, 943. VLA at 327 MHz ( $3'.5 \times 3'.6 : S = 46.3 \text{ Jy}$ ).

Anderson & Rudnick 1993, ApJ, 408, 514. VLA at 1.45 and 4.89 GHz, for spectral index studies.

Gorham et al. 1996, ApJ, 458, 257. Pulsar search.

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

Chen et al. 1999, ApJ, 520, 737. ASCA and ROSAT observations. Dyer & Reynolds 1999, ApJ, 526, 365. VLA at 1.5 GHz (6".6×6".9) and 4.8 GHz (5".6×6".4), including polarisation and comparison with ROSAT image.

Safi-Harb *et al.* 2000, ApJ, 545, 922. ROSAT, ASCA and other X-ray observations.

Sugizaki *et al.* 2001, ApJS, 134, 77. ASCA survey observations. Aharonian *et al.* 2001, A&A, 375, 1008. HESS limit.

Safi-Harb *et al.* 2005, ApJ, 618, 321. Chandra observations. Reach *et al.* 2006, AJ, 131, 1479. Spitzer observations.

Jiang *et al.* 2010, ApJ, 712, 1147. CO observations of region. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S =

 $18.5\pm1.1$  Jy) including polarisation and review of flux densities. Yang *et al.* 2013, ApJ, 766, 44. Suzaku spectroscopy.

Yamaguchi et al. 2015, ApJ, 801, L31. Suzaku observations. Clark et al. 2015, ApJ, 809, L2. Fermi observations.

Kilpatrick et al. 2016, ApJ, 816, 1. CO observations, including broad lines.

Leahy & Ranasinghe 2016, ApJ, 817, 74. VGPS for HI absorption. Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS Hi absorption observations.

## G41.5+0.4

**RA:** 19<sup>h</sup>05<sup>m</sup>50<sup>s</sup> **1-GHz flux/Jy:** 1? **Size/arcmin:** 10 **Dec:** +07°46′ **Spectral index:** ? **Type:** S?

Radio: Partial clumpy shell, brighter to NE.

### References

Kaplan *et al.* 2002, ApJ, 566, 378. VLA at 332 MHz (20'':  $S = 1.8 \pm 0.4$  Jy).

Alves et al. 2012, MNRAS, 422, 2429. Radio observations.

# G42.0-0.1

**RA**: 19<sup>h</sup>08<sup>m</sup>10<sup>s</sup> 1-GHz flux/Jy: 0.5? Size/arcmin: 8 **Dec:**  $+08^{\circ}00'$ **Spectral index:** ? Type: S?

Radio: Irregular shell.

#### References:

Kaplan et al. 2002, ApJ, 566, 378. VLA at 332 MHz (20": S =  $1.8\pm0.4 \text{ Jy}$ ).

Alves et al. 2012, MNRAS, 422, 2429. Radio observations.

## G42.8+0.6

RA: 19h07m20s **1-GHz flux/Jy: 3**? Size/arcmin: 24 **Dec:** +09°05′ **Spectral index:** 0.5? Type: S

Has been called G42.8+0.65.

Radio: Faint shell.

**Point sources:** Near soft gamma repeater, and young pulsar.

#### References:

Fürst et al. 1987, A&AS, 69, 403. Effelsberg 100-m at 4.75 GHz  $(2.4:S=1.5\pm0.2 \text{ Jy})$ , plus other flux densities.

Vasisht et al. 1994, ApJ, 431, L35. VLA at 327 MHz (3'.2×3'.4). Hurley et al. 1996, ApJ, 463, L13. Observations of soft gamma

repeater field.

Lorimer & Xilouris 2000, ApJ, 545, 385. Pulsar detection. Aharonian et al. 2001, A&A, 375, 1008. HESS limit. Kaplan et al. 2002, ApJ, 566, 378. VLA at 333 MHz (50"), and other observations of the region.

G43.3-0.2W49B

**RA:** 19<sup>h</sup>11<sup>m</sup>08<sup>s</sup> Size/arcmin:  $4 \times 3$ 1-GHz flux/Jy: 38 **Dec:** +09°06′ **Spectral index:** 0.46 Type: S

Radio: Shell, brightest to the SE and W, near the HII region W49A.

**X-ray:** Centrally brightened, elongated E–W. **Point sources:** Compact X-ray source. **Distance:** Hi absorption suggests 11.3 kpc.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 2.7 GHz (4').

Downes & Wilson 1974, A&A, 34, 133. Effelsberg 100-m at 10.7 GHz (1'.3).

Green et al. 1975, A&A, 44, 187. Effelsberg 100-m at 15.0 GHz  $(58'': S=9.0\pm0.7 \text{ Jy}).$ 

Lockhart & Goss 1978, A&A, 67, 355. Hi absorption.

Altenhoff et al. 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2.6).

Pye et al. 1984, MNRAS, 207, 649. Einstein observations, and VLA at 1.4 and 5 GHz (both 12").

Smith et al. 1985, ApJ, 296, 469. EXOSAT spectrum.

Morsi & Reich 1987, A&AS, 71, 189. Effelsberg 100-m at 32 GHz (smoothed to 30'':  $S=6.90\pm0.38$  Jy).

Moffett & Reynolds 1994, ApJ, 437, 705. VLA at 330 MHz  $(6''.7 \times 7''.7 : S = 64.4 \text{ Jy}), 1.48 \text{ GHz} (4''.8 \times 5''.2 : S = 31.8 \text{ Jy}) \text{ and}$ 4.85 GHz (4".0×4".1), including polarisation.

Fujimoto et al. 1995, PASJ, 47, L31. ASCA observations. Gorham et al. 1996, ApJ, 458, 257. Pulsar search.

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations. Hwang et al. 2000, ApJ, 532, 970. ROSAT image and ASCA spec-

Sugizaki et al. 2001, ApJS, 134, 77. ASCA survey observations. Aharonian et al. 2001, A&A, 375, 1008. HESS limit. Brogan & Troland 2001, ApJ, 550, 799. VLA at 1.4 GHz

 $(24'' \times 27'')$  and 5'') for Hi Zeeman splitting.

Lacey et al. 2001, ApJ, 559, 954. VLA at 74 MHz  $(23'' \times 26'' : S =$ 55.6 Jy) and 326 MHz (6".2 $\times$ 6".6:S=56.0 Jy).

Kaplan et al. 2002, ApJ, 566, 378. VLA at 333 MHz (50"), and other observations of the region.

Kawasaki et al. 2005, ApJ, 631, 935. ASCA observations. Reach et al. 2006, AJ, 131, 1479. Spitzer observations.

Miceli et al. 2006, A&A, 453, 567. XMM-Newton observations. Keohane et al. 2007, ApJ, 654, 938. IR and Chandra observations. Ozawa et al. 2009, ApJ, 706, L71. Suzaku observations.

Abdo et al. 2010, ApJ, 722, 1303. Fermi observations.

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S= 19.1±1.0 Jy) including polarisation and review of flux densities.

Rodes-Roca 2013, A&A, 555, A115. IR of compact X-ray source. Yang et al. 2013, ApJ, 766, 44. Suzaku spectroscopy.

Lopez et al. 2013, ApJ, 777, 145. Chandra spectroscopy. Zhu *et al.* 2014, ApJ, 793, 95. Spitzer and other observations. Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Kilpatrick et al. 2016, ApJ, 816, 1. CO observations, including broad lines.

Koo et al. 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Abdalla et al. 2018, A&A, 612, A1. HESS observations. Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

Abdalla et al. 2018, A&A, 612, A5. HESS observations. Zhou & Vink 2018, A&A, 615, A150. Chandra observations. Tanaka et al. 2018, ApJ, 866, L26. NuSTAR observations. Yamaguchi et al. 2018, ApJ, 868, L35. NuSTAR observations.

# G43.9 + 1.6

**RA:** 19h05m50s **1-GHz flux/Jy:** 9.0 **Size/arcmin:** 60? **Dec:** +10°30′ **Spectral index:** 0.5 **Type:** S?

**Radio:** Large, poorly defined faint shell. **Point sources:** Soft gamma repeater nearby.

#### References

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4'.3).

Vasisht *et al.* 1994, ApJ, 431, L35. VLA at 327 MHz (3'.2×3'.4). Hurley *et al.* 1996, ApJ, 463, L13. Observations of soft gamma repeater field.

Kaplan *et al.* 2002, ApJ, 566, 378. VLA at 333 MHz (50"), and other observations of the region.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9!5: S =  $4.55\pm0.24$  Jy) including polarisation and review of flux densities.

## G45.7 - 0.4

RA: 19<sup>h</sup>16<sup>m</sup>25<sup>s</sup> 1-GHz flux/Jy: 4.2? Size/arcmin: 22 Dec: +11°09′ Spectral index: 0.4? Type: S

**Radio:** Shell, brightest to the SE, poorly defined to NW.

#### References:

Fürst *et al.* 1987, A&AS, 69, 403. Effelsberg 100-m at 4.75 GHz (2'.4: S=2.6±0.3 Jy), plus other flux densities.

Gorham *et al.* 1996, ApJ, 458, 257. Pulsar search.

Reach *et al.* 2006, AJ, 131, 1479. Spitzer possible detection.

Acero *et al.* 2016, ApJS, 224, 8. Fermi observations.

G46.8-0.3 (HC30)

**RA:** 19<sup>h</sup>18<sup>m</sup>10<sup>s</sup> **1-GHz flux/Jy:** 17 **Size/arcmin:** 15 **Dec:** +12°09′ **Spectral index:** 0.54 **Type:** S

Has been called G46.6-0.2.

**Radio:** Shell, two bright arcs to NNW and SSE. **Distance:** HI absorption suggests 5.7 to 11.4 kpc.

### References:

Willis 1973, A&A, 26, 237. NRAO 300-ft at 2.7 GHz (5' :  $S = 9.8 \pm 0.9$  Jy), and 37-m at 1.7 GHz ( $S = 14.5 \pm 5.5$  Jy).

Caswell *et al.* 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3':S = 20.3 Jy) and Parkes 64-m at 5 GHz (4':S = 7.1 Jy).

Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo  $S_{430~\mathrm{MHz}} = 46 \pm 21~\mathrm{Jy}.$ 

Angerhofer *et al.* 1977, A&A, 55, 11. NRAO 140-ft at 5 GHz (6'.8:  $S = 5.9 \pm 0.6$  Jy).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2!6).

Dubner *et al.* 1996, AJ, 111, 1304. VLA at 1.4 GHz  $(53'' \times 58'' : S = 13.3 \pm 0.1 \text{ Jy})$ .

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz ( $9.5:S=7.02\pm0.18$  Jy) including polarisation and review of flux densities.

Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS Hi absorption observations.

G49.2-0.7 (W51)

**RA:** 19<sup>h</sup>23<sup>m</sup>50<sup>s</sup> **1-GHz flux/Jy:** 160? **Size/arcmin:** 30 **Dec:** +14°06′ **Spectral index:** 0.3? **Type:** S?

Has erroneously been called G49.1–0.1.

**Radio:** In complex region, parameters uncertain. **Optical:** Some diffuse emission possibly associated.

X-ray: Elongated east—west.

**Distance:** Association with CO gives 6 kpc, optical absorption suggests 5.7 kpc, and H<sub>I</sub> absorption suggests 5.4 kpc.

#### References:

Shaver & Goss 1970, AuJPA, 14, 133. Parkes 64-m at 5 GHz (4'). Sato 1973, PASJ, 25, 135. HI absorption.

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz (5'), S=51.5±3.2 Jy, for the non-thermal component, but probably confused

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'.6).

Seward 1990, ApJS, 73, 781. Einstein observations.

Copetti & Schmidt 1991, MNRAS, 250, 127. 151 MHz observations

Kassim 1992, AJ, 103, 943. VLA at 327 MHz (3.1×3.5).

Subrahmanyan & Goss 1995, MNRAS, 275, 755. VLA at 330 MHz (1'.1).

Koo et al. 1995, ApJ, 447, 211. ROSAT observations.

Frail et al. 1996, AJ, 111, 1651. OH emission near remnant.

Gorham et al. 1996, ApJ, 458, 257. Pulsar search.

Green et al. 1997, AJ, 114, 2058. OH masers.

Koo & Moon 1997, ApJ, 475, 194. Arecibo (3') and VLA  $(40'' \times 42'')$  at 1.4 GHz for Hi.

Koo & Moon 1997, ApJ, 485, 263. NRAO 12-m CO and HCO<sup>+</sup> observations.

Brogan et al. 2000, ApJ, 537, 875. VLA at 1.7 GHz for OH Zeeman splitting.

Mavromatakis *et al.* 2001, A&A, 370, 265. Optical observations. Koo *et al.* 2002, AJ, 123, 1629. ASCA observations.

Koo et al. 2005, ApJ, 633, 946. Chandra observations.

Kang & Koo 2007, ApJS, 173, 85. SGPS of high velocity HI.
Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers.

Abdo et al. 2009, ApJ, 706, L1. Fermi observations.

Koo et al. 2010, AJ, 140, 262. Hi Zeeman splitting observations.
Ceccarelli et al. 2011, ApJ, 740, L4. Molecular line observations of region

Aleksić et al. 2012, A&A, 541, A13. γ-ray observations.

Hanabata *et al.* 2013, PASJ, 65, 42. Suzaku observations. Tian & Leahy 2013, ApJ, 769, L17. HI observations of region.

Brogan et al. 2013, ApJ, 771, 91. VLA at 74 MHz (84"×92") and 320 MHz (33"×35"), plus OH, molecular line and other observations

Park *et al.* 2013, ApJ, 777, 14. Arecibo of Ht in region. Sasaki *et al.* 2014, A&A, 563, A9. XMM-Newton observations. Pihlström *et al.* 2014, AJ, 147, 73. VLA search for methanol masers.

Dumes *et al.* 2014, ApJ, 786, L24. Molecular line observations. Jogler & Funk 2016, ApJ, 816, 100. Fermi observations. McEwen *et al.* 2016, ApJ, 826, 189. NH<sub>3</sub> and CH<sub>3</sub>OH observa-

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Shan et al. 2018, ApJS, 238, 35. Optical absorption for distance.Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

G53.4+0.0

**RA:** 19<sup>h</sup>29<sup>m</sup>57<sup>s</sup> **1-GHz flux/Jy:** 1.5 **Size/arcmin:** 10? **Dec:** +18°10′ **Spectral index:** 0.6? **Type:** S

Has been called G53.41+0.03. **Radio:** Asymmetric shell.

X-ray: Detected.

### References:

Anderson *et al.* 2017, A&A, 605, A58. VLA at 1 to 2 GHz. Driessen *et al.* 2018, ApJ, 860, 133. LOFAR at 140 MHz, plus other radio observations, and XMM-Newton observations.

**G53.6–2.2** 3C400.2, NRAO 611

**RA:** 19<sup>h</sup>38<sup>m</sup>50<sup>s</sup> **1-GHz flux/Jy:** 8 **Size/arcmin:** 33×28 **Dec:** +17°14′ **Spectral index:** 0.50 **Type:** S

Has been called G53.7–2.2.

**Radio:** Ring of emission, with extension to NW.

**Optical:** Filaments and diffuse emission. **X-ray:** Centrally brightened, offset to NW. **Distance:** Association with H<sub>I</sub> gives 2.8 kpc.

#### References:

Willis 1973, A&A, 26, 237. NRAO 300-ft at 2.7 GHz (5' :  $S = 5.3 \pm 0.6$  Jy).

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz  $(5': S=4.8\pm0.3 \text{ Jy})$ .

Clark *et al.* 1975, AuJPA, 37, 75. Molonglo at 408 MHz (3': S = 11.7 Jy).

Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo  $S_{430~\mathrm{MHz}} = 20 \pm 10~\mathrm{Jy}, \ S_{318~\mathrm{MHz}} = 20 \pm 3.6~\mathrm{Jy}.$ 

Goss *et al.* 1975, A&A, 43, 459. WSRT at 610 MHz ( $1' \times 3' : S = 13.2 \pm 1.6 \text{ Jy}$ ).

Sabbadin & d'Odorico 1976, A&A, 49, 119. Optical spectra. van den Bergh 1978, ApJS, 38, 119. Optical observations. Rosado 1983, RMxAA, 8, 59. Optical spectra.

Blair & Long 1988, PASP, 100, 461. Optical imaging and spectroscopy.

see also: Blair & Long 1988, PASP, 100, 651. Erratum. Long et al. 1991, ApJ, 373, 567. Einstein and optical observations. Winkler *et al.* 1993, ApJ, 405, 608. Optical imaging. Dubner *et al.* 1994, AJ, 108, 207. VLA at 327 MHz (59") and 1.49 GHz (52"), plus X-rays.

Saken *et al.* 1995, ApJ, 443, 231. ROSAT observations. Gorham *et al.* 1996, ApJ, 458, 257. Pulsar search.

Giacani *et al.* 1998, A&AS, 133, 61. DRAO at 1.4 GHz for HI studies, including distance.

Yoshita *et al.* 2001, PASJ, 53, 93. ASCA observations, and spectral comparison with ROSAT.

Ambrocio-Cruz et al. 2006, RMxAA, 42, 241. Optical imaging and spectroscopy.

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5:  $S = 4.00 \pm 0.22$  Jy) including polarisation and review of flux densities.

Broersen & Vink 2015, MNRAS, 446, 3885. Chandra observa-

Ergin et al. 2017, ApJ, 842, 22. Suzaku and Fermi observations.

# G54.1+0.3

Radio: Filled-centre core, with possible faint diffuse emission.

**X-ray:** Centrally concentrated, with more extended diffuse emission.

Point sources: Central pulsar.

**Distance:** HI absorption suggests 4.9, association with CO suggests 8.2 kpc, and optical absorption suggests 6.3 kpc.

### References:

Green 1985, MNRAS, 216, 691. Radio at 2.7 GHz ( $7'' \times 20''$ ). Reich *et al.* 1985, A&A, 151, L10. Effelsberg 100-m at 4.75 GHz ( $2'.4:S=0.37\pm0.04$  Jy)

Velusamy & Becker 1988, AJ, 95, 1162. VLA at 1.4 ( $14'': S = 0.48 \pm 0.03$  Jy), 1.6 ( $14'': S = 0.42 \pm 0.03$  Jy) and 5 GHz ( $5'': S = 0.33 \pm 0.02$  Jy), Ooty at 327 MHz ( $S = 0.50 \pm 0.08$  Jy), plus review of flux densities.

Seward 1989, AJ, 97, 481. Einstein observations.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search.

Gorham et al. 1996, ApJ, 458, 257. Pulsar search.

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Lu *et al.* 2001, A&A, 370, 570. ROSAT and ASCA observations. Lu *et al.* 2002, ApJ, 568, L49. Chandra observations.

Camilo et al. 2002, ApJ, 574, L71. Pulsar detection.

Kaplan & Moon 2006, ApJ, 644, 1056. IR upper limit for pulsar. Leahy *et al.* 2008, AJ, 136, 1477. VGPS at 1.4 GHz (1') including Hr.

Koo et al. 2008, ApJ, 673, L147. Akari observations of surroundings.

Hurley-Walker et al. 2009, MNRAS, 396, 365. Radio observations at 14 to 18 GHz.

Bocchino et al. 2010, A&A, 520, A71. XMM-Newton and Suzaku observations.

Lang *et al.* 2010, ApJ, 709, 1125. VLA at 1.4 GHz (6".6×6".8), 4.7 GHz (3".2×3".3), and 8.2 GHz (3".0×3".2) and Spitzer observations.

Acciari *et al.* 2010, ApJ, 719, L69. γ-ray observations. Lee *et al.* 2012, JKAS, 45, 117. CO observations of region. Krivonos *et al.* 2017, MNRAS, 470, 512. INTEGRAL observa-

Temim et al. 2017, ApJ, 836, 129. Spitzer, Herschel and Akari observations.

Shan et al. 2018, ApJS, 238, 35. Optical absorption for distance.Ranasinghe & Leahy 2018, AJ, 155, 204. VGPS HI absorption observations.

Rho et al. 2018, MNRAS, 479, 5101. Spitzer, Herschel and other observations.

Driessen *et al.* 2018, ApJ, 860, 133. LOFAR observations at 144 MHz, plus other observations.

G54.4-0.3 (HC40)

**RA:** 19h33m20s **1-GHz flux/Jy:** 28 **Size/arcmin:** 40 **Dec:** +18°56′ **Spectral index:** 0.5 **Type:** S

Has been called G54.5–0.3.

**Radio:** Shell, in complex region.

**Optical:** Faint filaments.

Point sources: Pulsar outside NW rim.

**Distance:** HI and CO observations suggest 6.6 kpc.

References:

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz  $(5':S=34.4\pm5.0 \text{ Jy})$ .

Altenhoff et al. 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz

(2'.6). Caswell 1985, AJ, 90, 1224. DRAO at 1.4 GHz (1'.3×2'.6: S =

18±4 Jy). Velusamy *et al.* 1986, JApA 7, 105, WSRT at 609 MHz

Velusamy *et al.* 1986, JApA, 7, 105. WSRT at 609 MHz  $(50'' \times 191'')$  smoothed to  $100'' \times 200''$ ).

Junkes *et al.* 1992, A&AS, 96, 1. Surrounding CO. Junkes *et al.* 1992, A&A, 261, 289. Nearby IRAS sources.

Boumis et al. 2005, A&A, 443, 175. Optical observations.
Reach et al. 2006, AJ, 131, 1479. Spitzer observations.
Kang & Koo 2007, ApJS, 173, 85. SGPS of high velocity HI.
Andersen et al. 2011, ApJ, 742, 7. Spitzer observations.
Park et al. 2013, ApJ, 777, 14. Arecibo of HI in region.
Froebrich et al. 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.
Karpova et al. 2017, MNRAS, 466, 1757. X-ray observations of pulsar.

Ranasinghe & Leahy 2017, ApJ, 843, 119. HI and CO observations.

G55.0+0.3

**RA:**  $19^{\text{h}}32^{\text{m}}00^{\text{s}}$  **1-GHz flux/Jy:** 0.5? **Size/arcmin:**  $20 \times 15$ ? **Oec:**  $+19^{\circ}50'$  **Spectral index:** 0.5? **Type:** S

Has been called G55.2+0.5. **Radio:** Faint, partial shell.

**Point sources:** Old pulsar nearby.

**Distance:** Association with H<sub>I</sub> features implies 14 kpc.

References:

Taylor *et al.* 1992, AJ, 103, 931. WSRT at 327 MHz ( $1.0 \times 2.5$ ), and northern sky survey at 4.9 GHz.

Matthews *et al.* 1998, ApJ, 493, 312. WSRT at 327 MHz  $(1.0 \times 2.9 : S = 0.98 \pm 0.15 \text{ Jy})$ , DRAO at 1.4 GHz  $(1.0 \times 2.9 : S = 0.25 \pm 0.12 \text{ Jy})$ , plus HI observations.

Reach et al. 2006, AJ, 131, 1479. Spitzer possible detection.

G55.7 + 3.4

**RA:** 19<sup>h</sup>21<sup>m</sup>20<sup>s</sup> **1-GHz flux/Jy:** 1? **Size/arcmin:** 23 **Dec:** +21°44′ **Spectral index:** 0.3? **Type:** S

Radio: Incomplete shell.

**Point sources:** Old pulsar within the boundary of the remnant.

References:

Goss *et al.* 1977, A&A, 61, 93. WSRT observations at 610 MHz  $(57'' \times 156'' : S = 1.9 \pm 0.2 \text{ Jy})$  and 1415 MHz  $(27'' \times 72'' : S = 1.0 \pm 0.1 \text{ Jy})$ .

Bhatnagar *et al.* 2011, ApJ, 739, L20. VLA at 1.3 to 1.9 GHz (30").

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9 $^{\prime}$ 5:  $S = 0.52 \pm 0.03$  Jy) including polarisation and review of flux densities.

G57.2+0.8 (4C21.53)

**RA:** 19<sup>h</sup>34<sup>m</sup>59<sup>s</sup> **1-GHz flux/Jy:** 1.8 **Size/arcmin:** 12? **Dec:** +21°57′ **Spectral index:** 0.35 **Type:** S?

**Radio:** Extended non-thermal arc. **Point sources:** Central magnetar.

**Distance:** HI observations suggest 12.5 kpc.

#### References:

Sieber & Seiradakis 1984, A&A, 130, 257. Effelsberg 100-m at 1.4 GHz ( $8.8:1.34\pm0.1$ ), 2.7 GHz ( $4.3:0.86\pm0.1$ ), plus other surveys of the area

Caswell et al. 1985, AJ, 90, 488. DRAO at 1.4 GHz ( $1' \times 3'$ ). Hurley-Walker et al. 2009, MNRAS, 396, 365. Radio observations

at 14 to 18 GHz. Sun *et al.* 2011. A&A, 536, A83, U

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9!5 : S = 0.74 $\pm$ 0.04 Jy) including polarisation and review of flux densities

Surnis *et al.* 2016, ApJ, 826, 184. GMRT at 610 MHz and VLA at 1.4 GHz.

Israel *et al.* 2016, MNRAS, 457, 3448. Chandra, XMM-Newton and Swift observations of magnetar.

Kothes *et al.* 2018, ApJ, 852, 54. DRAO at 408 MHz (2'.8×8'.5) and 1.4 GHz (0'.82×2'.5), including H<sub>I</sub> and polarisation, plus other radio observations.

# G59.5+0.1

RA: 19<sup>h</sup>42<sup>m</sup>33<sup>s</sup> 1-GHz flux/Jy: 3? Size/arcmin: 15 Dec: +23°35′ Spectral index: ? Type: S

Has been called G59.6+0.1. **Radio:** Incomplete shell. **Optical:** Diffuse shell.

#### References:

Taylor *et al.* 1992, AJ, 103, 931. WSRT at 327 MHz ( $1.0 \times 2.5 : S = 5.1 \pm 0.2$  Jy), and northern sky survey at 4.9 GHz.

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Gök *et al.* 2008, Ap&SS, 318, 207. Optical observations.

Hurley-Walker et al. 2009, MNRAS, 396, 365. Radio observations at 14 to 18 GHz.

Xu & Wang 2012, A&A, 543, A24. CO observations of SE.

# G63.7 + 1.1

**RA:** 19<sup>h</sup>47<sup>m</sup>52<sup>s</sup> **1-GHz flux/Jy:** 1.8 **Size/arcmin:** 8 **Dec:** +27°45′ **Spectral index:** 0.24 **Type:** F

Radio: Centrally brightened, with core.

**X-ray:** Diffuse emission.

### References:

Taylor *et al.* 1992, AJ, 103, 931. WSRT at 327 MHz ( $1.0 \times 2.2$ ), and northern sky survey at 4.9 GHz.

Wallace *et al.* 1997, AJ, 114, 2068. WSRT at 1.4 GHz ( $14'' \times 26''$ : S = 1.63 Jy), DRAO at 1.4 GHz (smoothed to 2'), plus review of flux densities and other observations.

Hurley-Walker et al. 2009, MNRAS, 396, 365. Radio observations at 14 to 18 GHz.

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9!5: S =  $1.12\pm0.06$  Jy) including polarisation and review of flux densities.

Matheson et al. 2016, ApJ, 825, 134. XMM-Newton and Chandra observations.

# G64.5 + 0.9

**RA:** 19<sup>h</sup>50<sup>m</sup>25<sup>s</sup> **1-GHz flux/Jy:** 0.15? **Size/arcmin:** 8 **Dec:** +28°16′ **Spectral index:** 0.5 **Type:** S?

**Radio:** Shell with central source. **Optical:** Filaments in N and W.

### References:

Hurley-Walker et al. 2009, MNRAS, 398, 249. Radio identifica-

Neustadt et al. 2017, MNRAS, 469, 516. Optical observations.

# G65.1+0.6

**RA:**  $19^h54^m40^s$  **1-GHz flux/Jy:** 5.5 **Size/arcmin:**  $90\times50$  **Dec:**  $+28^{\circ}35'$  **Spectral index:** 0.61 **Type:** S

Radio: Large, faint shell.

Point sources: Old pulsar nearby.

**Distance:** Possible association with H<sub>I</sub> suggests 9 kpc.

#### References:

Landecker *et al.* 1990, A&A, 232, 207. DRAO at 408 MHz  $(3.5 \times 7.0 : S = 9.5 \pm 0.1 \text{ Jy})$ , and 1.4 GHz  $(1.0 \times 2.0 : S = 5.4 \pm 1.0 \text{ Jy})$ .

Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search. Tian & Leahy 2006, A&A, 455, 1053. CGPS at 408 MHz (2'.8×5'.9: S=8.6±0.8 Jy) and 1.4 GHz (0'.8×1'.7: S=4.9±0.5 Jy) including HI.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3':  $S = 9.1 \pm 1.0$  Jy) and 1420 MHz ( $\sim$  1':  $S = 3.9 \pm 0.5$  Jy), including review of flux densities.

Aleksić *et al.* 2010, ApJ, 725, 1629. γ-ray observations. Gao *et al.* 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9'.5:  $S = 3.2 \pm 0.3$  Jy), including polarisation and review of flux densities.

Froebrich et al. 2015, MNRAS, 454, 2586. H2 IR observations.

# G65.3 + 5.7

**RA:** 19<sup>h</sup>33<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** 42 **Size/arcmin:** 310×240 **Dec:** +31°10′ **Spectral index:** 0.6 **Type:** S?

Has been called G65.2+5.7.

Radio: Large, faint ring, near S91 and S94.

**Optical:** Filamentary ring.

**X-ray:** Diffuse, centrally brightened.

**Distance:** Optical proper motions and velocities indicates 0.8 kpc.

# References:

Gull *et al.* 1977, ApJ, 215, L69. Optical plates. Reich *et al.* 1979, A&A, 72, 270. Effelsberg 100-m observations at 1.42 GHz (smoothed to  $11': S = 42.4 \pm 1.6$  Jy), estimate  $S_{408 \text{ MHz}} = 91 \pm 5$  Jy from previous sky survey. Lozinskaya 1981, SvAL, 7, 17. Mean optical velocity. Rosado 1981, ApJ, 250, 222. Optical interferometry. Fesen *et al.* 1983, ApJS, 51, 337. Deep [OIII] imagery. Fesen *et al.* 1985, ApJ, 292, 29. Optical spectra. Seward 1990, ApJS, 73, 781. Einstein observations. Gorham *et al.* 1996, ApJ, 458, 257. Pulsar search.

Mavromatakis et al. 2002, A&A, 388, 355. Optical observations.
Boumis et al. 2004, A&A, 424, 583. Optical observations.
Shelton et al. 2004, ApJ, 615, 275. ROSAT observations.
Kaplan et al. 2006, ApJS, 163, 344. X-ray upper limit on compact sources.

Xiao *et al.* 2009, A&A, 503, 827. Effelsberg 100-m at 2.7 GHz  $(4.4:S=22\pm3 \text{ Jy})$ , and Urumqi 21-m at 4.8 GHz  $(9.5:S=16.8\pm1.8 \text{ Jy})$  including polarisation and review of flux densities. Gosachinskii 2010, AstL, 36, 260. HI observations. Kim *et al.* 2010, ApJ, 722, 388. Far UV observations.

G65.7 + 1.2DA 495

**RA:** 19<sup>h</sup>52<sup>m</sup>10<sup>s</sup> 1-GHz flux/Jy: 5.1 Size/arcmin: 22 **Dec:** +29°26′ Spectral index: varies Type: F

Has mistakenly been called G55.7+1.2.

Radio: Centrally brightened with thick shell?

**X-ray:** Centrally brightened.

**Point sources:** Compact X-ray source near centre. **Distance:** Hi polarisation observations suggest 1.5 kpc.

### References:

Willis 1973, A&A, 26, 237. NRAO 300-ft at 2.7 GHz (5' : S = $2.8\pm0.4$  Jy), and 37-m at 1.7 GHz ( $S=4.4\pm0.5$  Jy), plus review of flux densities.

see also: Willis 1973, A&A, 27, 483. Erratum.

Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo  $S_{430 \text{ MHz}}$  = 8.7 $\pm$ 4.9 Jy,  $S_{318~\text{MHz}} = 9.7 \pm 2.2$  Jy.

Landecker & Caswell 1983, AJ, 88, 1810. DRAO at 1.4 GHz

 $(0.9 \times 1.5 : S = 4.4 \pm 0.2 \text{ Jy}).$ 

Velusamy et al. 1989, JApA, 10, 161. Ooty at 327 MHz  $(36'' \times 64'')$ , WSRT at 610 MHz (62'') and VLA at 1.4 GHz (36'')and 12"), including IRAS imaging.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Gorham et al. 1996, ApJ, 458, 257. Pulsar search. Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Kothes et al. 2004, ApJ, 607, 855. HI polarisation absorption.

Arzoumanian et al. 2004, ApJ, 610, L101. ROSAT and ASCA observations of compact source.

Kothes et al. 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3' : S =$  $6.5\pm0.6 \text{ Jy}$ ) and 1420 MHz ( $\sim 1': S = 4.0\pm0.2 \text{ Jy}$ ), including polarisation and review of flux densities.

Arzoumanian et al. 2008, ApJ, 687, 505. Chandra observations. Kothes et al. 2008, ApJ, 687, 516. CGPS at 408 MHz (2'.9×6'.0:  $S=6.5\pm0.5$  Jy) and 1.4 GHz (0'.82×1'.75:  $S=4.0\pm0.2$  Jy), Effelsberg 100-m at 4.85 GHz (2'.45 :  $S = 1.6 \pm 0.1$  Jy) and 10.55 GHz (smoothed to  $2.45: S = 1.1 \pm 0.1$  Jy), plus review of flux densities. Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S= 1.95±0.10 Jy) including polarisation and review of flux densi-

Karpova et al. 2015, MNRAS, 453, 2241. Chandra and XMM-Newton observations.

# G66.0 - 0.0

**RA:** 19<sup>h</sup>57<sup>m</sup>50<sup>s</sup> 1-GHz flux/Jy: ? Size/arcmin:  $31 \times 25$ ? **Dec:** +29°03′ **Spectral index:** ? Type: S

**Radio:** Some emission in N. **Optical:** Incomplete shell.

**Distance:** Optical absorption suggests 2.3 kpc.

References:

Sabin et al. 2013, MNRAS, 431, 279. Ha and radio survey observations.

Shan et al. 2018, ApJS, 238, 35. Optical absorption for distance.

G67.6+0.9

**RA:** 19<sup>h</sup>57<sup>m</sup>45<sup>s</sup> 1-GHz flux/Jy: ? Size/arcmin:  $50 \times 45$ ? **Dec:** +30°53′ **Spectral index:** ? Type: S

**Radio:** Arc in S.

**Optical:** Filamentary shell.

**Distance:** Optical absorption suggests 2.0 kpc.

References:

Sabin et al. 2013, MNRAS, 431, 279. Ha and radio survey obser-

vations.

Shan et al. 2018, ApJS, 238, 35. Optical absorption for distance.

G67.7 + 1.8

**RA:**  $19^h54^m32^s$  **1-GHz flux/Jy:** 1.0 **Size/arcmin:**  $15\times12$  **Dec:**  $+31^{\circ}29'$  **Spectral index:** 0.61 **Type:** S

Radio: Double arc shell. Optical: Filaments in N. X-ray: Detected.

**Point sources:** Compact X-ray source.

**Distance:** Optical absorption suggests 1.5–5.7 kpc.

References:

Taylor *et al.* 1992, AJ, 103, 931. WSRT at 327 MHz ( $1.0 \times 1.9 \pm 0.1$  Jy), and northern sky survey at 4.9 GHz ( $S = 0.42 \pm 0.05$  Jy).

Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search. Mavromatakis *et al.* 2001, A&A, 370, 265. Optical observations. Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3': S =  $1.1\pm0.1$  Jy) and 1420 MHz ( $\sim$  1': S =  $0.68\pm0.04$  Jy), including polarisation and review of flux densities.

Gök *et al.* 2008, Ap&SS, 318, 207. Optical observations. Hurley-Walker *et al.* 2009, MNRAS, 396, 365. Radio observations at 14 to 18 GHz.

Hui & Becker 2009, A&A, 494, 1005. Chandra observations. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5:  $S = 0.30 \pm 0.03$  Jy) including polarisation and review of flux densities.

Shan et al. 2018, ApJS, 238, 35. Optical absorption for distance.

G67.8 + 0.5

**RA:**  $20^h00^m00^s$  **1-GHz flux/Jy:** ? **Size/arcmin:**  $7\times 5$  **Dec:**  $+30^\circ 51'$  **Spectral index:** ? **Type:** ?

Radio: Poorly resolved arc.

**Optical:** Diffuse shell, brighter to W.

References:

Sabin *et al.* 2013, MNRAS, 431, 279. Hα and radio survey observations.

G68.6 - 1.2

**RA:** 20<sup>h</sup>08<sup>m</sup>40<sup>s</sup> **1-GHz flux/Jy:** 1.1 **Size/arcmin:** 23 **Dec:** +30°37′ **Spectral index:** 0.2 **Type:** ?

Radio: Faint, poorly defined source.

References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Junkes *et al.* 1988, LNP, 316, 134. Effelsberg 100-m at 2.7 GHz (4'.3), including polarisation.

Reich et al. 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4'.3).

Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 1420 MHz ( $\sim 1'$ :  $S = 0.57 \pm 0.08$  Jy), including review of flux densities

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5:  $S = 0.80 \pm 0.04$  Jy), including polarisation and review of flux densition

G69.0+2.7

RA: 19<sup>h</sup>53<sup>m</sup>20<sup>s</sup>
1-GHz flux/Jy: 120?
Size/arcmin: 80?
Spectral index: varies
Type: ?

An association with a SN in AD1408 has been suggested. Has been called G68.8+2.8.

**Radio:** Compact core, flat spectrum plateau, and steeper spectrum extensions, with spectral break?

**Optical:** Expanding nebulosity near centre, with filaments to the SW and far NE.

**X-ray:** Diffuse emission with compact source. **Point sources:** Pulsar at western edge of core.

**Distance:** HI observations suggest 1.5 kpc, and optical absorption suggests 4.6 kpc.

#### References:

Angerhofer *et al.* 1981, A&A, 94, 313. WSRT at 610 MHz  $(56'' \times 103'')$  1.4 GHz  $(24'' \times 44'')$  and 5 GHz  $(7'' \times 13'')$ , plus optical.

Becker *et al.* 1982, ApJ, 255, 557. X-ray observations. Sofue *et al.* 1983, PASJ, 35, 437. NRO 45-m at 10.2 GHz (2'7). Velusamy & Kundu 1983, JApA, 4, 253. VLA of compact sources. Blair *et al.* 1984, ApJ, 282, 161. Optical images and spectra. Wang & Seward 1984, ApJ, 285, 607. Einstein observations. Strom *et al.* 1984, A&A, 139, 43. Radio observations of flat spec-

trum component, VLA 5 GHz (1"7) and 1.4 GHz. Mantovani *et al.* 1985, A&A, 145, 50. Bologna at 408 MHz  $(2.6 \times 4.9 : S = 67.5 \pm 10.5 \text{ Jy})$ , Effelsberg 100-m at 1.41 (9':  $S = 62 \pm 9 \text{ Jy})$ , 1.72 (7.6:  $S = 66 \pm 5 \text{ Jy})$ , 2.7 (4.5:  $S = 52 \pm 4 \text{ Jy}$ ) and

4.75 GHz (2.4: S=44±3.3 Jy), plus review of flux densities. Kulkarni et al. 1988, Nature, 331, 50. Pulsar detection. Angelini et al. 1988, ApJ, 330, L43. EXOSAT spectra. Fesen et al. 1988, Nature, 334, 229. IRAS of surrounding shell. Junkes et al. 1988, LNP, 316, 134. Effelsberg 100-m at 2.7 GHz (4.3), including polarisation.

Whitehead *et al.* 1989, MNRAS, 237, 1109. Optical of core. Salter *et al.* 1989, ApJ, 338, 171. NRAO 12-m at 84.2 GHz of core, plus review of flux densities.

Hester & Kulkarni 1989, ApJ, 340, 362. Optical imaging and spectroscopy.

Koo *et al.* 1990, ApJ, 364, 178. Large, expanding HI shell. Greidanus & Strom 1990, A&A, 240, 376. Optical kinematics of core

Koo *et al.* 1993, ApJ, 417, 196. VLA at 1.4 GHz of surrounding Hi (1'), plus IRAS.

Safi-Harb *et al.* 1995, ApJ, 439, 722. ROSAT observations. Srinivasan 1997, ApJ, 489, 170.  $\gamma$ -ray observations of pulsar. Mavromatakis *et al.* 2001, A&A, 371, 300. Optical observations. Butler *et al.* 2002, A&A, 395, 845. HST detection of pulsar. Migliazzo *et al.* 2002, ApJ, 567, L141. Pulsar proper motion study. Castelletti *et al.* 2003, AJ, 126, 2114. GMRT at 240 and 618 MHz  $(17''\times26'')$  and  $(17''\times26'')$  and  $(17''\times26'')$  and VLA at 324 and 1380 MHz  $(63''\times73'')$  and  $(17''\times26'')$ .

Moon et al. 2004, ApJ, 610, L33. Chandra and HST observations of core

Li et al. 2005, ApJ, 628, 931. Chandra observations of pulsar and surroundings.

Golden et al. 2005, ApJ, 635, L153. High resolution radio observations of pulsar and surroundings.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3':  $S = 72\pm7$  Jy) and 1420 MHz ( $\sim$  1':  $S = 56\pm5$  Jy), including polarisation and review of flux densities.

Kang & Koo 2007, ApJS, 173, 85. SGPS of high velocity HI.
Albert et al. 2007, ApJ, 669, 1143. γ-ray observations.
Zeiger et al. 2008, ApJ, 674, 271. Proper motion of pulsar.
Gao et al. 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9/5: S=36±4 Jy), including polarisation and review of flux densities.
Leahy & Ranasinghe 2012, MNRAS, 423, 718. CGPS at 1.4 GHz, including HI, plus ROSAT observations.

Park et al. 2013, ApJ, 777, 14. Arecibo of HI in region. Arnaud et al. 2016, A&A, 586, A134. Planck flux densities at 30 and 44 GHz.

Shan et al. 2018, ApJS, 238, 35. Optical absorption for distance.

# G69.7 + 1.0

**RA:**  $20^{\rm h}02^{\rm m}40^{\rm s}$  **1-GHz flux/Jy:** 2.0 **Size/arcmin:**  $16\times14$  **Dec:**  $+32^{\circ}43'$  **Spectral index:** 0.7 **Type:** S

Radio: Poorly resolved source.

X-ray: Detected.

### References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Junkes *et al.* 1988, LNP, 316, 134. Effelsberg 100-m at 2.7 GHz (4'.3), including polarisation.

Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4/3).

Lorimer et al. 1998, A&A, 331, 1002. Pulsar search. Yoshita et al. 2000, PASJ, 52, 867. ROSAT and ASCA observations. Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3': S = 3.2 \pm 0.4$  Jy) and 1420 MHz ( $\sim 1': S = 1.5 \pm 0.1$  Jy), including review of flux densities.

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5:  $S = 0.78 \pm 0.07$  Jy) including polarisation and review of flux densities.

# G70.0 - 21.5

**RA:**  $21^h24^m00^s$  **1-GHz flux/Jy:** ? **Size/arcmin:**  $330\times240$  **Dec:**  $+19^{\circ}23'$  **Spectral index:** ? **Type:** S

Radio: Not detected.

Optical: Large, faint shell of filaments.

**X-ray:** Partially detected.

### References:

Boumis et al. 2002, A&A, 396, 225. Optical and ROSAT observa-

tions.

Fesen et al. 2015, ApJ, 812, 37. Hα and other optical/UV line and

ROSAT observations.

G73.9 + 0.9

**RA:** 20<sup>h</sup>14<sup>m</sup>15<sup>s</sup> **1-GHz flux/Jy:** 9 **Size/arcmin:** 27 **Dec:** +36°12′ **Spectral index:** 0.23 **Type:** S?

Radio: Diffuse, centrally brightened to SW.

Optical: Faint shell.

### References:

Reich *et al.* 1986, A&A, 155, 185. Effelsberg 100-m at 4.75 GHz  $(2.4:S=6.7\pm0.5\ \mathrm{Jy})$ , plus other flux densities.

Chastenay & Pineault 1988, IAUCo, 101, 297. DRAO at 408 MHz (3'.5×5'.9) and 1.4 GHz (1'.0×1'.7).

Pineault & Chastenay 1990, MNRAS, 246, 169. DRAO at 408 MHz ( $3'.4 \times 5'.8 : S = 12.7 \pm 1.2 \text{ Jy}$ ) and 1.4 GHz ( $1'.0 \times 1'.7 : S = 7.4 \pm 1.0 \text{ Jy}$ ).

Gorham et al. 1996, ApJ, 458, 257. Pulsar search.

Pineault *et al.* 1996, AJ, 112, 201. DRAO at 1.4 GHz (smoothed to 2') for HI.

Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Mavromatakis 2003, A&A, 398, 153. Optical observations. Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3': S =  $10.0\pm1.7$  Jy) and 1420 MHz ( $\sim$  1': S =  $7.6\pm0.6$  Jy), including polarisation and review of flux densities.

Sitnik 2010, ARep, 54, 317. Hα and CO observations of region. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9'.5: *S* = 6.2±0.3 Jy) including polarisation and review of flux densities. Jeong *et al.* 2012, Ap&SS, 342, 389. CO observations of region. Acero *et al.* 2016, ApJS, 224, 8. Fermi observations. Zdziarski *et al.* 2016, MNRAS, 455, 1451. Fermi observations.

G74.0–8.5 Cygnus Loop

**RA:**  $20^h51^m00^s$  **1-GHz flux/Jy:** 210 **Size/arcmin:**  $230\times160$  **Dec:**  $+30^\circ40'$  **Spectral index:** varies **Type:** S

Has been suggested that this is two overlapping remnants.

**Radio:** Shell, brightest to the NE, with fainter breakout region to S, with spectral variations.

**Optical:** Large filamentary loop, brightest to the NE, not well defined to the S or W.

**X-ray:** Shell in soft X-rays.

**Point sources:** Several compact radio sources within the boundary of the remnant, including CL4, plus X-ray sources in S

**Distance:** Stellar interactions gives 0.74 kpc.

#### References:

Green 1990, AJ, 100, 1927. DRAO at 408 MHz ( $3'.3 \times 6'.7$ ) for spectral index study, plus X-ray and optical.

Graham et al. 1991, AJ, 101, 175. Shocked molecular H outside rim in NE.

Shull & Clarke 1991, PASP, 103, 811. Optical spectroscopy of nearside filaments.

Fesen et al. 1992, AJ, 104, 719. Ha imagery.

Long et al. 1992, ApJ, 400, 214. Optical of Balmer dominated filament.

Arendt et al. 1992, ApJ, 400, 562. IRAS observations.

Hester *et al.* 1994, ApJ, 420, 721. Hα, [O<sub>III</sub>] and other optical observations of Balmer dominated filaments in NE.

Gorham et al. 1996, ApJ, 458, 257. Pulsar search.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search.

Leahy et al. 1997, AJ, 114, 2081. DRAO at 1.4 GHz ( $1' \times 2'$ ), including polarisation.

Leahy & Roger 1998, ApJ, 505, 784. DRAO at 1.4 GHz ( $1.0 \times 1.9$ ) and 408 MHz ( $3.4 \times 6.9$ ), for spectral index studies in comparison with other radio observations.

Levenson *et al.* 1998, ApJS, 118, 541. Optical images. Roger *et al.* 1999, A&AS, 137, 7. 22 MHz flux density (*S* = 1400±400 Jy).

Bohigas et al. 1999, ApJ, 518, 324. Optical spectroscopy of surroundings.

Levenson et al. 1999, ApJ, 526, 874. ROSAT images.

Blair *et al.* 1999, AJ, 118, 942. HST observations, for distance. Danforth *et al.* 2000, AJ, 119, 2319. UV, optical and X-ray comparison of selected regions.

Ghavamian et al. 2001, ApJ, 547, 995. Optical spectroscopy.

Miyata et al. 2001, ApJ, 550, 1023. ASCA observations of compact
X-ray sources.

Danforth *et al.* 2001, AJ, 122, 938. Far-UV spectroscopy, Hα and other optical observations of NE region.

Levenson & Graham 2001, ApJ, 559, 948. HST of SE region. Uyanıker *et al.* 2002, A&A, 389, L61. Effelsberg 100-m at 2.7 GHz (4'3) including polarisation, and comparison with ROSAT data.

Leahy 2002, AJ, 123, 2689. DRAO at 1.4 GHz  $(2'\times4')$  for HI. Blair *et al.* 2002, ApJS, 140, 367. UV spectroscopy. Levenson *et al.* 2002, ApJ, 576, 798. Chandra observations of W

Levenson et al. 2002, ApJ, 576, 798. Chandra observations of W edge.

Reich *et al.* 2003, A&A, 408, 961. Effelsberg 100-m at 863 MHz  $(14.5: S = 184 \pm 18 \text{ Jy})$ .

Leahy 2004, MNRAS, 351, 385. Chandra observations of SW. Uyanıker *et al.* 2004, A&A, 426, 909. Effelsberg 100-m at 2.7 GHz (4'.3), with comparison with other data for spectral index studies. Blair *et al.* 2005, AJ, 129, 2268. HST of outer filaments.

Leahy 2005, AJ, 130, 165. DRAO at 1.4 GHz of SE.

Levenson & Graham 2005, ApJ, 622, 366. Chandra observations of knot in SE.

Sun *et al.* 2006, A&A, 447, 937. Urumqi 25-m at 4.8 GHz (9'.5:  $S=90\pm9$  Jy), with comparisons with other data for spectral index studies

Kaplan et al. 2006, ApJS, 163, 344. X-ray upper limit on compact sources.

Seon *et al.* 2006, ApJ, 644, L175. Far UV observations. Sankrit *et al.* 2007, AJ, 133, 1383. UV observations of part.

Tsunemi et al. 2007, ApJ, 671, 1717. XMM-Newton observations of NE to SW.

Nemes et al. 2008, ApJ, 675, 1293. XMM-Newton observations of NE.

Katsuda *et al.* 2008, ApJ, 680, 1198. Chandra observations of NE. Uchida *et al.* 2008, ApJ, 688, 1102. XMM-Newton observations. Uchida *et al.* 2009, PASJ, 61, 503. Suzaku observations of N.

Kimura *et al.* 2009, PASJ, 61, S137. Suzaku observations from NE to SW.

Tsunemi *et al.* 2009, PASJ, 61, S147. Suzaku observations of SE. Kosugi *et al.* 2010, PASJ, 62, 1035. Suzaku observations of SE. Sankrit *et al.* 2010, ApJ, 712, 1092. Spitzer observations.

McEntaffer & Brantseg 2011, ApJ, 730, 99. Chandra observations of E.

Katagiri et al. 2011, ApJ, 741, 44. Fermi observations.

Katsuda et al. 2012, ApJ, 754, L7. Suzaku and XMM-Newton observations.

Leahy & Hassan 2013, ApJ, 764, 55. Suzaku and XMM-Newton observations of SW.

Oakley et al. 2013, ApJ, 766, 51. X-ray spectroscopy.

Kim et al. 2014, ApJ, 784, 12. Far UV observations.

Sankrit et al. 2014, ApJ, 787, 3. Spitzer spectroscopy in SE.

Medina et al. 2014, ApJ, 791, 30. Optical spectroscopy in NE. Raymond et al. 2015, ApJ, 805, 152. HST observations in NE.

see also: Raymond et al. 2015, ApJ, 814, 165. Erratum.

Roberts & Wang 2015, MNRAS, 449, 1340. Suzaku observations. Arnaud *et al.* 2016, A&A, 586, A134. Planck flux density 30 GHz.

Katsuda *et al.* 2016, ApJ, 819, L32. Hα spectroscopy of NE. Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Boubert et al. 2017, A&A, 606, A14. Gaia search for runaway progenitor companion.

Fesen *et al.* 2018, MNRAS, 481, 1786. Optical observations for distance.

G74.9+1.2

RA:  $20^h 16^m 02^s$ 1-GHz flux/Jy: 9Size/arcmin:  $8 \times 6$ Dec:  $+37^{\circ}12'$ Spectral index: variesType: F

**Radio:** Filled-centre, with high polarisation and high frequency turnover.

X-ray: Centrally brightened.

**Point sources:** Compact X-ray source in SE. **Distance:** Optical extinction gives 6.1 kpc.

#### References:

Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo  $S_{430~\rm MHz}=12.2\pm9.7~\rm Jy,~S_{318~\rm MHz}=17.7\pm5.0~\rm Jy.$  Weiler & Shaver 1978, A&A, 70, 389. WSRT at 610 MHz

Weiler & Shaver 1978, A&A, 70, 389. WSRT at 610 MHz  $(57'' \times 94'' : S = 9.1 \pm 1.2 \text{ Jy})$ , 1.4  $(24'' \times 40'' : S = 8.7 \pm 1.2 \text{ Jy})$  and 5 GHz  $(24'' \times 40'' : S = 5.6 \pm 1.3 \text{ Jy})$ .

Geldzahler *et al.* 1980, A&A, 84, 237. Effelsberg 100-m at 2.7 GHz ( $4.4:S=7.6\pm0.5$  Jy).

Wilson 1980, ApJ, 241, L19. Einstein observations.

van Gorkom *et al.* 1982, MNRAS, 198, 757. WSRT HI absorption of nearby compact source.

Seaquist & Gilmore 1982, AJ, 87, 378. VLA observations of nearby source.

Morsi & Reich 1987, A&AS, 69, 533. Effelsberg 100-m at 32 GHz (smoothed to 40'':  $S=1.47\pm0.19$  Jy).

Green & Gull 1989, MNRAS, 237, 555. VLA at 1.4 GHz (1'.2×1'.4) including Hı.

Salter et al. 1989, ApJ, 338, 171. NRAO 12-m at 84.2 GHz, plus review of flux densities.

Pineault & Chastenay 1990, MNRAS, 246, 169. DRAO at 408 MHz ( $3'.4 \times 5'.8 : S = 11.6 \pm 0.4$  Jy) and 1.4 GHz ( $1'.0 \times 1'.7 : S = 7.2 \pm 0.3$  Jy).

Asaoka & Koyama 1990, PASJ, 42, 625. Ginga X-ray spectrum. Wendker *et al.* 1991, A&A, 241, 551. DRAO at 408 MHz  $(3.5 \times 5.2 : S = 13.3 \pm 0.8 \text{ Jy})$  and Effelsberg 100-m at 4.8 GHz  $(S = 7.5 \pm 0.7 \text{ Jy})$ .

Wallace et al. 1994, A&A, 286, 565. HI of surroundings.

Cho et al. 1994, AJ, 108, 634. CO of adjacent molecular clouds.

Gorham et al. 1996, ApJ, 458, 257. Pulsar search.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Wallace *et al.* 1997, A&A, 317, 212. DRAO at 408 MHz (3'.4×5'.5)

Wallace et al. 1997, A&A, 317, 212. DRAO at 408 MHz (3: $4 \times 5$ :5) and 1.4 GHz (1: $0 \times 1$ :6) including HI (smoothed to 2').

Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Kothes *et al.* 2003, ApJ, 588, 852. CGPS at 1.4 GHz ( $1' \times 1'$ .6) including H<sub>I</sub>, plus CO observations.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3': S = 11.9 $\pm$ 0.9 Jy) and 1420 MHz ( $\sim$  1': S = 7.1 $\pm$ 1.1 Jy), including polarisation and review of flux densities.

Hurley-Walker *et al.* 2009, MNRAS, 396, 365. Radio observations at 14 to 18 GHz.

Sitnik 2010, ARep, 54, 317. H $\alpha$  and CO observations of region. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5:  $S = 6.4 \pm 0.4$  Jy) including polarisation and review of flux densities.

Matheson *et al.* 2013, ApJ, 774, 33. Chandra observations.

Bassani *et al.* 2014, A&A, 561, A108. X-ray and γ-ray observations.

Aliu *et al.* 2014, ApJ, 788, 78. γ-ray detection. Saha 2016, MNRAS, 460, 3563. Fermi observations. Acero *et al.* 2016, ApJS, 224, 8. Fermi observations. Liu *et al.* 2018, ApJ, 859, 173. CO observations. Abeysekara *et al.* 2018, ApJ, 861, 134. γ-ray observations.

# G76.9 + 1.0

RA: 20<sup>h</sup>22<sup>m</sup>20<sup>s</sup> 1-GHz flux/Jy: 2? Size/arcmin: 9 Dec: +38°43′ Spectral index: ? Type: C

Radio: Bipolar shell.

Point sources: Central pulsar.

### References:

Taylor *et al.* 1992, AJ, 103, 931. WSRT at 327 MHz ( $1.0 \times 1.6$ ), and northern sky survey at 4.9 GHz.

Landecker *et al.* 1993, A&A, 276, 522. VLA at 1.49 GHz (14"), 4.86 GHz (13"×16") and 8.55 GHz (11"×12"), including polarisation and review of flux densities.

Landecker *et al.* 1997, A&AS, 123, 199. Miyun at 232-MHz (3!8×5!4).

Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3': S =  $2.3\pm0.2$  Jy) and 1420 MHz ( $\sim$  1': S =  $1.35\pm0.07$  Jy), including polarisation and review of flux densities.

Hurley-Walker et al. 2009, MNRAS, 396, 365. Radio observations at 14 to 18 GHz.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9 $^{\prime}$ 5: S = 0.79 $\pm$ 0.07 Jy) including polarisation and review of flux densities.

Marthi *et al.* 2011, MNRAS, 416, 2560. GMRT at 618 MHz  $(51'' \times 54'')$ , 1160 MHz  $(2''.2 \times 3''.4)$ , and Chandra observations of central source.

Arzoumanian *et al.* 2011, ApJ, 739, 39. Pulsar detection. Jeong *et al.* 2012, Ap&SS, 342, 389. CO observations of region.

G78.2+2.1 DR4, γ Cygni SNR

Has been called G78.1+1.8.

**Radio:** In complex region (early catalogues refer to other proposed remnants in this region).

Optical: Faint filaments, spectra indicate a SNR superposed on a HII region.

**X-ray:** Weak emission from the SE of the remnant.

**Point sources:** X-ray pulsar at edge of remnant, with nebula.

**Distance:** Associations with other objects suggests 1.7 to 2.6 kpc.

#### References:

Higgs *et al.* 1977, AJ, 82, 718. DRAO at 1.4 GHz ( $2' \times 3' : S = 270\pm40$  Jy) plus some 10 GHz (4') survey data, reveals true extent of remnant.

d'Odorico & Sabbadin 1977, A&AS, 28, 439. Optical spectra. van den Bergh 1978, ApJS, 38, 119. Optical observations. Landecker *et al.* 1980, A&AS, 39, 133. DRAO HI observations (2'×3'1).

Higgs et al. 1983, AJ, 88, 97. CO of surroundings.

Bohigas et al. 1983, RMxAA, 8, 155. Optical spectra, find thermal only.

Braun & Strom 1986, A&AS, 63, 345. WSRT HI observations. Fukui & Tatematsu 1988, IAUCo, 101, 261. CO observations of the vicinity (2'7).

Green 1989, MNRAS, 238, 737. OH observations.

Pineault & Chastenay 1990, MNRAS, 246, 169. DRAO at 408 MHz ( $3'.4 \times 5'.8 : S = 480 \pm 60$  Jy) and 1.4 GHz ( $1'.0 \times 1'.7 : S = 270 \pm 40$  Jy).

Wendker *et al.* 1991, A&A, 241, 551. DRAO at 408 MHz  $(3.5 \times 5.2 : S = 540 \pm 40 \text{ Jy})$  and Effelsberg 100-m at 4.8 GHz  $(S = 150 \pm 15 \text{ Jy})$ .

Esposito *et al.* 1996, ApJ, 461, 820. Associated γ-ray emission. Brazier *et al.* 1996, MNRAS, 281, 1033. γ-ray and X-ray point source.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search.

Zhang et al. 1997, A&A, 324, 641. Multi-frequency radio comparison.

Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search. Roberts *et al.* 2001, ApJS, 133, 451. ASCA observations.

Uchiyama *et al.* 2002, ApJ, 571, 866. ASCA observations. Mavromatakis 2003, A&A, 408, 237. Optical observations.

Bykov 2004, A&A, 427, L21. Hard X-ray observations.

Becker et al. 2004, ApJ, 615, 897. Chandra and other observations

of compact sources. Weisskopf *et al.* 2006, ApJ, 652, 387. Chandra and other observa-

tions of compact sources. Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3'$ :  $S = 500 \pm 35$  Jy) and 1420 MHz ( $\sim 1'$ :  $S = 226 \pm 19$  Jy), including

review of flux densities. Kang & Koo 2007, ApJS, 173, 85. SGPS of high velocity HI. Casandjian & Grenier 2008, A&A, 489, 849.  $\gamma$ -ray observations.

Casandjian & Grenier 2008, A&A, 489, 849. 7-ray observations. Ladouceur & Pineault 2008, A&A, 490, 197. CGPS at 408 MHz (2'.9×4'.5) and 1.4 GHz (0'.8×1'.5).

Gao *et al.* 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9'.5:  $S = 170 \pm 18$  Jy), including polarisation and review of flux densities.

Leahy et al. 2013, MNRAS, 436, 968. ROSAT and Chandra observations, and CGPS for HI.

Aliu et al. 2013, ApJ, 770, 93. γ-ray observations.

Lin et al. 2013, ApJ, 770, L9. Pulsar detection.

Hui et al. 2015, ApJ, 799, 76. XMM-Newton and Chandra observations of pulsar, including proper motion.

Fraija & Araya 2016, ApJ, 826, 31. Fermi observations. Acero *et al.* 2016, ApJS, 224, 8. Fermi observations.

Abeysekara *et al.* 2018, ApJ, 861, 134. γ-ray observations.

G82.2+5.3 W63

**RA:**  $20^{\rm h}19^{\rm m}00^{\rm s}$  **1-GHz flux/Jy:** 120? **Size/arcmin:**  $95\times65$  **Dec:**  $+45^{\circ}30'$  **Spectral index:** 0.5? **Size/arcmin:**  $95\times65$ 

Has been called G82.5+5.3.

**Radio:** Shell in the Cygnus X complex.

Optical: In complex region, but spectra indicate SNR filaments.

X-rav: Detected.

**Distance:** Optical absorption suggests 3.2 kpc.

# References:

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz (5':S=59.0±3.5 Jy).

Sabbadin 1976, A&A, 51, 159. Optical spectra.

Angerhofer et al. 1977, A&A, 55, 11. NRAO 140-ft at 5 GHz (6/8). Incomplete mapping.

Rosado & González 1981, ŘMxAA, 5, 93. Optical spectra. Seward 1990, ApJS, 73, 781. Einstein observations.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search.

Reich *et al.* 2003, A&A, 408, 961. Effelsberg 100-m at 863 MHz  $(14'5: S = 82.5 \pm 5.5 \text{ Jy})$ .

Uyanıker et al. 2003, ApJ, 585, 785. CGPS at 1.4 GHz (1') including polarisation, of part.

Mavromatakis *et al.* 2004, A&A, 415, 1051. ROSAT, ASCA and optical observations.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3': S = 144 $\pm$ 12 Jy) and 1420 MHz ( $\sim$  1': S = 93 $\pm$ 5 Jy), including review of flux densities.

Gao *et al.* 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9'.5:  $S=49\pm5$  Jy), including polarisation and review of flux densities. Shan *et al.* 2018, ApJS, 238, 35. Optical absorption for distance.

# G83.0 - 0.3

**RA:**  $20^{\text{h}}46^{\text{m}}55^{\text{s}}$  **1-GHz flux/Jy:** 1 **Size/arcmin:**  $9\times7$  **Dec:**  $+42^{\circ}52'$  **Spectral index:** 0.4 **Type:** S

Radio: Incomplete shell.

#### References:

Taylor et al. 1992, AJ, 103, 931. WSRT at 327 MHz ( $1.0 \times 1.5$ ), and northern sky survey at 4.9 GHz.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3':  $S = 1.2 \pm 0.3$  Jy) and 1420 MHz ( $\sim$  1':  $S = 0.8 \pm 0.1$  Jy, including polarisation and review of flux densities.

G84.2 - 0.8

**RA:**  $20^h53^m20^s$  **1-GHz flux/Jy:** 11 **Size/arcmin:**  $20\times16$  **Dec:**  $+43^\circ27'$  **Spectral index:** 0.5 **Type:** S

**Radio:** Elongated shell, with a filament aligned with the major axis.

X-ray: Detected.

**Distance:** Hi absorption suggests 6 kpc.

### References:

Matthews *et al.* 1977, A&A, 55, 1. WSRT at 610 MHz ( $56'' \times 81''$ :  $S = 12.4 \pm 1.5$  Jy) and Effelsberg 100-m at 2.7 GHz ( $4'.4: S = 6.8 \pm 1.3$  Jy).

Matthews & Shaver 1980, A&A, 87, 255. WSRT at 1415 MHz  $(23'' \times 32'')$ , and Effelsberg 100-m at 2.7 GHz (4.4 :  $S = 5.6 \pm 0.5$  Jy).

Feldt & Green 1993, A&A, 274, 421. DRAO at 1.4 GHz (1'×1'.5), including HI, plus CO observations.

Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Uyanıker et al. 2003, ApJ, 585, 785. CGPS at 1.4 GHz (1') including polarisation.

Kaplan et al. 2004, ApJS, 153, 269. Chandra limits for any compact source.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3': S = 14.5 \pm 0.5$  Jy) and 1420 MHz ( $\sim 1': S = 7.2 \pm 0.8$  Jy), including review of flux densities.

Leahy & Green 2012, ApJ, 760, 25. CGPS, including HI, plus Chandra observations.

Jeong et al. 2012, Ap&SS, 342, 389. CO observations of region.

G85.4 + 0.7

RA: 20<sup>h</sup>50<sup>m</sup>40<sup>s</sup> 1-GHz flux/Jy: ? Size/arcmin: 24? Dec: +45°22′ Spectral index: 0.2 Type: S

**Radio:** Faint, incomplete shell, within larger thermal shell.

X-ray: Centrally brightened.

**Distance:** Hi observations suggest 3.5 kpc, and optical absorption suggests 4.4 kpc.

### References:

Kothes *et al.* 2001, A&A, 376, 641. CGPS at 408 MHz ( $2'.8\times4'.4:S<0.45$  Jy) and 1.4 GHz ( $0'.8\times1'.1$ ), plus H<sub>I</sub> and X-ray data.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3': S = 2.9 \pm 0.5$  Jy) and 1420 MHz ( $\sim 1': S = 2.3 \pm 0.2$  Jy), including review of flux densities.

Jackson et al. 2008, ApJ, 674, 936. XMM-Newton and Hi observa-

Jeong et al. 2012, Ap&SS, 342, 389. CO observations of region. Shan et al. 2018, ApJS, 238, 35. Optical absorption for distance.

# G85.9 - 0.6

**RA:** 20<sup>h</sup>58<sup>m</sup>40<sup>s</sup> **1-GHz flux/Jy:** ? **Size/arcmin:** 24 **Dec:** +44°53′ **Spectral index:** 0.2 **Type:** S

Radio: Faint, incomplete shell.

**Optical:** Diffuse shell. **X-ray:** Centrally brightened.

**Distance:** Hi observations suggest 4.8 kpc.

#### References:

Kothes *et al.* 2001, A&A, 376, 641. CGPS at 408 MHz ( $2'.8 \times 4'.4$ : S < 0.9 Jy) and 1.4 GHz ( $0'.8 \times 1'.1$ ), plus HI, X-ray and optical data

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3': S = 3.0 \pm 1.3$  Jy) and 1420 MHz ( $\sim 1': S = 2.2 \pm 0.8$  Jy), including review of flux densities.

Jackson et al. 2008, ApJ, 674, 936. XMM-Newton and Hi observations.

Gök et al. 2009, Ap&SS, 324, 17. Optical observations. Jeong et al. 2012, Ap&SS, 342, 389. CO observations of region.

G89.0+4.7

**RA:** 20<sup>h</sup>45<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** 220 **Size/arcmin:** 120×90 **Dec:** +50°35′ **Spectral index:** 0.38 **Type:** S

Radio: Distorted shell (4C50.52, an extragalactic double, is within the boundary of the remnant).

**Optical:** Filaments and patches. **X-ray:** Centrally brightened.

**Distance:** Various associations suggest 0.8 kpc, and optical absorption suggests 1.9 kpc.

### References:

Hirabayashi & Takahashi 1972, PASJ, 24, 231. 30-m dish at 4.2 GHz (11' :  $S = 160 \pm 40$  Jy).

Willis 1973, A&A, 26, 237. NRAO 300-ft at 2.7 GHz (5' :  $S = 148\pm16$  Jy), plus optical filaments.

Hill 1974, MNRAS, 169, 59. Half-Mile Telescope at 1.4 GHz  $(3'\times3'.9)$ .

Haslam et al. 1975, A&A, 39, 453. Effelsberg 100-m at 2.7 GHz (4/4)

Fukui & Tatematsu 1988, IAUCo, 101, 261. CO observations of the vicinity (2'.7).

Tatematsu *et al.* 1990, A&A, 237, 189. DRAO at 408 MHz (3'.5×4'.5) and 1.4 GHz (1'.0×1'.3), including HI, plus CO observations of adjacent molecular cloud.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search.

Leahy & Aschenbach 1996, A&A, 315, 260. ROSAT observations. Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search.

Koo et al. 2001, ApJ, 552, 175. NRAO 12-m and other CO observations (27'' and 45'') of eastern part.

Reich *et al.* 2003, A&A, 408, 961. Effelsberg 100-m at 863 MHz  $(14/5: S = 228 \pm 5 \text{ Jy})$ .

Uyanıker et al. 2003, ApJ, 585, 785. CGPS at 1.4 GHz (1') including polarisation.

Byun *et al.* 2006, ApJ, 637, 283. CO observations of surroundings. Lazendic & Slane 2006, ApJ, 647, 350. X-ray observations.

Leahy 2006, ApJ, 647, 1125. CGPS at 408 MHz ( $2'.8 \times 3'.7$ ) and 1.4 GHz ( $0'.8 \times 1'.1$ ).

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3':  $S = 259 \pm 19$  Jy) and 1420 MHz ( $\sim$  1':  $S = 183 \pm 9$  Jy), including polarisation and review of flux densities.

Kang & Koo 2007, ApJS, 173, 85. SGPS of high velocity HI. Mavromatakis *et al.* 2007, A&A, 461, 991. Optical observations. Shinn *et al.* 2009, ApJ, 693, 1883. IR observations.

Pannuti *et al.* 2010, AJ, 140, 1787. ASCA and observations. Shinn *et al.* 2010, AdSpR, 45, 445. IR observations in S.

Sninn et al. 2010, Adspk, 45, 445. In observations in S. Gao et al. 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9'.5:  $S = 107 \pm 11$  Jy), including polarisation and review of flux densities

Reichardt *et al.* 2012, A&A, 546, A21. Fermi detection.

Shinn et al. 2012, ApJ, 759, 34. Akari observations of H<sub>2</sub>.

Pivato *et al.* 2013, ApJ, 779, 179. Fermi observations. Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 30

Arnaud et al. 2016, A&A, 586, A134. Planck flux densities at 30 and 44 GHz.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Boubert et al. 2017, A&A, 606, A14. Gaia search for runaway progenitor companion.

Shan *et al.* 2018, ApJS, 238, 35. Optical absorption for distance. Suzuki *et al.* 2018, PASJ, 70, 75. Suzuku observations.

**G93.3+6.9** DA 530, 4C(T)55.38.1

**RA:**  $20^h52^m25^s$  **1-GHz flux/Jy:** 9 **Size/arcmin:**  $27\times20$  **Dec:**  $+55^{\circ}21'$  **Spectral index:** 0.45 **Type:** C?

Has been called G93.2+6.7.

**Radio:** Shell, with two bright limbs, highly polarised.

**X-ray:** Compact central source.

**Distance:** Hi observations suggest 2.2 kpc.

#### References:

Roger & Costain 1976, A&A, 51, 151. DRAO at 1.42 GHz  $(2' \times 2'.4 : S = 6.9 \text{ Jy})$ .

Haslam *et al.* 1980, A&A, 92, 57. Effelsberg 100-m at 1.72 GHz  $(7.6:S=6.47\pm0.52 \text{ Jy})$  and 2.7 GHz  $(4.4:S=5.64\pm0.64 \text{ Jy})$ , plus review of flux densities.

Lalitha et al. 1984, A&A, 131, 196. Effelsberg 100-m at 4.75 GHz (smoothed to 3': S=4.01 $\pm$ 0.57 Jy).

Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search. Landecker *et al.* 1999, ApJ, 527, 866. DRAO at 408 MHz (3'.5×4'.3) and 1.4 GHz (1'.0×1'.2), including polarisation and H<sub>I</sub>. Foster & Routledge 2003, ApJ, 598, 1005. HI for distance. Kaplan *et al.* 2004, ApJS, 153, 269. Chandra limits for any compact source.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3': S = 10.5 \pm 0.7$  Jy) and 1420 MHz, including review of flux densities. Jiang *et al.* 2007, ApJ, 670, 1142. Chandra observations. Bocchino *et al.* 2008, AdSpR, 41, 407. XMM-Newton observa-

Jeong et al. 2012, Ap&SS, 342, 389. CO observations of region.

**G93.7–0.2** CTB 104A, DA 551

Has been called G93.6-0.2 and G93.7-0.3.

Radio: Distorted, faint shell.

**Distance:** Association with H<sub>I</sub> features suggests 1.5 kpc.

### References:

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz (5':S=18.4±1.0 Jy).

Mantovani *et al.* 1982, A&A, 105, 176. Effelsberg 100-m at 1.7 GHz ( $7.6:S=53.5\pm5.0$  Jy), plus review of flux densities. Landecker *et al.* 1985, AJ, 90, 1082. DRAO at 1.4 GHz (smoothed to  $2':S=58\pm6$  Jy).

Mantovani *et al.* 1991, A&A, 247, 545. Effelsberg 100-m at 4.75 GHz (smoothed to  $3': S = 33.5 \pm 4.0$  Jy), including polarisation, plus review of flux densities.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search.

Koralesky et al. 1998, AJ, 116, 1323. VLA search for OH emission

Uyanıker *et al.* 2002, ApJ, 565, 1022. CGPS 1.4 GHz (49"×54"), including HI, and 408 MHz (2'.8×3'.7).

Uyanıker et al. 2003, ApJ, 585, 785. CGPS at 1.4 GHz (1') including polarisation.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3':  $S = 67 \pm 6$  Jy) and 1420 MHz ( $\sim$  1':  $S = 35 \pm 4$  Jy), including polarisation and review of flux densities.

Gao et al. 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9.5:  $S=25.0\pm2.5$  Jy), including polarisation and review of flux densities

**G94.0+1.0** 3C434.1

RA:  $21^h24^m50^s$  1-GHz flux/Jy: 13 Size/arcmin:  $30\times25$  Dec:  $+51^{\circ}53'$  Spectral index: 0.45 Type: S

Radio: Incomplete shell, containing HI shell.

**Distance:** Association with stellar wind bubble implies 5.2 kpc.

#### References:

Willis 1973, A&A, 26, 237. NRAO 300-ft at 2.7 GHz (5' :  $S = 6.1 \pm 0.8$  Jy), and 37-m at 1.7 GHz ( $S = 11 \pm 3$  Jy).

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz (5':S=5.8±0.4 Jy). Also NRAO 140-ft at 5 GHz (6').

Mantovani *et al.* 1982, A&A, 105, 176. Effelsberg 100-m at 1.7 GHz ( $7.6:S=12.0\pm1.3$  Jy), plus review of flux densities. Goss *et al.* 1984, A&A, 138, 469. WSRT at 610 MHz (smoothed to

Goss *et al.* 1984, A&A, 138, 469. WSRT at 610 MHz (smoothed to  $100'': S = 16 \pm 1.7$  Jy) and Effelsberg 100-m at 4.75 GHz (2!4:  $S = 7.2 \pm 0.5$  Jy).

Landecker et al. 1985, AJ, 90, 1082. DRAO at 1.4 GHz (smoothed to  $2': S=16\pm3$  Jy).

Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Uyanıker *et al.* 2003, ApJ, 585, 785. CGPS at 1.4 GHz (1') including polarisation.

Foster *et al.* 2004, A&A, 417, 79. DRAO at 1.4 GHz, including Hi.

Foster 2005, A&A, 441, 1043. CGPS at 408 MHz ( $2.8 \times 3.6$ ) and 1.4 GHz ( $0.8 \times 1.0$ ) for spectral index studies, plus other observations.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3':  $S = 20 \pm 2$  Jy) and 1420 MHz ( $\sim$  1':  $S = 11.3 \pm 1.0$  Jy), including review of flux densities.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 6.2±0.4 Jy) including polarisation and review of flux densities. Jeong *et al.* 2012, Ap&SS, 342, 389. CO observations of region. Jeong *et al.* 2013, ApJ, 770, 105. CO observations of region.

G96.0 + 2.0

**RA:** 21<sup>h</sup>30<sup>m</sup>30<sup>s</sup> **1-GHz flux/Jy:** 0.35 **Dec:** +53°59′ **Size/arcmin:** 26 **Type:** S

**Radio:** Faint, arc in S, poorly defined in N. **Distance:** Association for H<sub>I</sub> indicates 4 kpc.

### References:

Kothes *et al.* 2005, A&A, 444, 871. CGPS at 408 MHz (2'.8×3'.5) and 1.4 GHz (50"×61") including HI.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3': S = 0.42 \pm 0.06$  Jy) and 1420 MHz ( $\sim 1': S = 0.24 \pm 0.02$  Jy), including review of flux densities.

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9 $^{\prime}$ 5:  $S = 0.14 \pm 0.02$  Jy) including polarisation and review of flux densities

G106.3 + 2.7

**RA:**  $22^h27^m30^s$  **1-GHz flux/Jy:** 6 **Size/arcmin:**  $60\times24$  **Dec:**  $+60^\circ50'$  **Spectral index:** 0.6 **Type:** C?

Incorporates the pulsar wind nebula G106.6+2.9 (the 'Boomerang').

Radio: Faint extended source, which brighter 'head' to NE.

**X-ray:** Pulsar and wind nebula.

**Point sources:** Pulsar.

### References:

Pineault & Joncas 2000, AJ, 120, 3218. DRAO at 408 MHz  $(3.5 \times 3.9 : S = 10.5 \pm 0.3 \text{ Jy})$  and 1.4 GHz  $(1.0 \times 1.2 : S = 4.9 \pm 0.6 \text{ Jy})$ , plus HI.

Halpern et al. 2001, ApJ, 547, 323. X-ray and radio observations of the 'head'.

Halpern et al. 2001, ApJ, 552, L125. Pulsar detection.

Kothes et al. 2001, ApJ, 560, 236. CGPS at 1.4 GHz, including HI, plus CO and other observations.

Ng & Romani 2004, ApJ, 601, 479. Chandra detection of pulsar wind nebula.

Kothes et al. 2004, ApJ, 607, 855. Hi polarisation absorption.

Kothes *et al.* 2006, ApJ, 638, 225. Effelsberg 100-m at 4.85 (2<sup>!</sup>4), 8.35 (1<sup>!</sup>4), 10.5 (1<sup>!</sup>2) and 32 GHz (0<sup>!</sup>45) of pulsar wind nebula, including polarisation.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3': S =  $8.6\pm1.0$  Jy) and 1420 MHz ( $\sim$  1': S =  $4.8\pm0.5$  Jy), including polarisation and review of flux densities.

Abdo et al. 2007, ApJ, 664, L91. γ-ray observations.

Casandjian & Grenier 2008, A&A, 489, 849. γ-ray observations.

Acciari *et al.* 2009, ApJ, 703, L6. γ-ray observations. Gao *et al.* 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9.5:

Gao *et al.* 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9.5:  $S = 2.0 \pm 0.3$  Jy), including polarisation and review of flux densities.

# G108.2 - 0.6

RA:  $22^h53^m40^s$  1-GHz flux/Jy: 8 Size/arcmin:  $70\times54$  Dec:  $+58^\circ50'$  Spectral index: 0.5 Type: S

Radio: Faint shell.

**Distance:** Possible associated H<sub>I</sub> structures suggest 3.2 kpc.

#### References:

Tian et al. 2007, A&A, 465, 907. DRAO at 408 MHz ( $2'.8 \times 3'.3:S = 11.5 \pm 1.2$  Jy) and 1.4 GHz ( $1'.0 \times 1'.2:S = 6.6 \pm 0.7$  Jy) including H<sub>1</sub>

G109.1-1.0 CTB 109

**RA:** 23<sup>h</sup>01<sup>m</sup>35<sup>s</sup> **1-GHz flux/Jy:** 20 **Size/arcmin:** 28 **Dec:** +58°53′ **Spectral index:** 0.45 **Type:** S

**Radio:** Semicircular shell, with the Molecular cloud S152 is to the immediate W.

**Optical:** Faint optical filaments.

**X-ray:** Semicircular shell, with pulsar at W edge. **Point sources:** Long period X-ray pulsar (magnetar). **Distance:** Various observations imply 3.2 kpc.

### References:

Hughes et al. 1981, ApJ, 246, L127. WSRT at 610 MHz (1':  $S = 40\pm5$  Jy) shows bad CLEAN artefacts.

Blair & Kirshner 1981, Nature, 291, 132. Optical spectra. Downes 1983, MNRAS, 203, 695. Effelsberg 100-m at 2.7 GHz

Downes 1983, MNRAS, 203, 695. Effelsberg 100-m at 2.7 GHz (4'.4: S=13.0±1.5 Jy). Sofue *et al.* 1983, PASJ, 35, 447. NRO 45-m at 10.2 GHz (2'.7).

Solue et al. 1983, 1783, 33, 347. INO 43-III at 10.2 GHz (2.7). Hughes et al. 1984, ApJ, 283, 147. WSRT at 610 MHz (20":  $S = 26 \pm 3$  Jy) 1.4 GHz (10"), DRAO at 1.4 GHz (1 $^{\prime}$ ×1/2:  $S = 16.8 \pm 2$  Jy) and Algonquin 46-m at 6.5 GHz (4 $^{\prime}$ 5:  $S = 6.7 \pm 1$  Jy). Braun & Strom 1986, A&AS, 63, 345. WSRT Hi observations.

Tatematsu *et al.* 1987, PASJ, 39, 755. NRO 45-m at 10 GHz (2'7), plus polarisation.

Tatematsu et al. 1987, A&A, 184, 279. CO observations of the surroundings (2'.7).

Hanson *et al.* 1988, A&A, 195, 114. EXOSAT of pulsar. *see also:* Hanson *et al.* 1988, A&A, 207, 204. Erratum. Morini *et al.* 1988, ApJ, 333, 777. EXOSAT observations. Koyama *et al.* 1989, PASJ, 41, 461. X-ray observations of pulsar. Coe *et al.* 1989, MNRAS, 238, 649. IRAS observations of surroundings

Green 1989, MNRAS, 238, 737. OH observations.

Tatematsu *et al.* 1990, ApJ, 351, 157. CO of surroundings, plus X-ray observations.

Davies & Coe 1991, MNRAS, 249, 313. Optical and IR observations near pulsar.

Fesen & Hurford 1995, AJ, 110, 747. Optical observations. Hurford & Fesen 1995, MNRAS, 277, 549. ROSAT imaging. Rho & Petre 1997, ApJ, 484, 828. ROSAT observations.

Parmar *et al.* 1998, A&A, 330, 175. X-ray observations. Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search.

Koralesky et al. 1998, AJ, 116, 1323. VLA search for OH emission

Patel *et al.* 2001, ApJ, 563, L45. Chandra observations of pulsar. Kothes *et al.* 2002, ApJ, 576, 169. CGPS at 1.4 GHz ( $59'' \times 68''$ ), including HI, plus CO observations.

Sasaki et al. 2004, ApJ, 617, 322. XMM-Newton observations.
 Sasaki et al. 2006, ApJ, 642, L149. CO observations of surroundings, plus Chandra observations.

Kothes et al. 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3':  $S = 26 \pm 3$  Jy) and 1420 MHz ( $\sim$  1':  $S = 17.4 \pm 1.2$  Jy), including polarisation and review of flux densities.

Tian *et al.* 2010, MNRAS, 404, L1. CGPS for H<sub>I</sub> absorption. Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 9.8±0.5 Jy) including polarisation and review of flux densities. Kothes & Foster 2012, ApJ, 746, L4. H<sub>I</sub> and CO observations of region.

Castro et al. 2012, ApJ, 756, 88. Fermi observations. Sasaki et al. 2013, A&A, 552, A45. Chandra observations of NE. Tendulakar et al. 2013, ApJ, 772, 31. Pulsar proper motion study. Vogel et al. 2014, ApJ, 789, 75. NuSTAR observations of pulsar. Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Nakano et al. 2017, PASJ, 69, 40. Suzaku observations.

Sánchez-Cruces et al. 2018, MNRAS, 473, 1705. Optical observations.

G111.7-2.1 Cassiopeia A, 3C461

**RA:** 23<sup>h</sup>23<sup>m</sup>26<sup>s</sup> **1-GHz flux/Jy:** 2300 **Size/arcmin:** 5 **Dec:** +58°48′ **Spectral index:** 0.77 **Type:** S

Presumably the remnant of a late 17th century SN.

**Radio:** Bright shell with compact knots and extended plateau of emission; shows secular decline. **Optical:** Fast knots and quasi-stationary flocculli, with many filaments at large radii, and NE 'jet'.

**X-ray:** Incomplete shell, with hard spectral component.

**Point sources:** Central compact X-ray source. **Distance:** Optical expansion gives 3.3 kpc.

#### References:

Anderson & Rudnick 1995, ApJ, 441, 307. VLA at 1.4 GHz and 4.8 GHz, for proper motion studies.

Kassim *et al.* 1995, ApJ, 455, L59. VLA at 74 (25") and 332 MHz (6"), including spectral index comparisons.

Anderson & Rudnick 1996, ApJ, 456, 234. VLA for spectral index comparisons.

Schwarz *et al.* 1997, A&AS, 123, 43. WSRT at 1.4 GHz (30'') for HI absorption.

O'Sullivan & Green 1999, MNRAS, 303, 575. Flux density changes at 13.5, 15.5 and 16.5 GHz.

Agüeros & Green 1999, MNRAS, 305, 957. 151 MHz observations for bulk expansion studies.

Gotthelf et al. 2001, ApJ, 552, L39. Chandra observations, showing outer shock.

Hwang *et al.* 2001, ApJ, 560, L175. Chandra observations of Doppler shifted lines.

Fesen et al. 2001, AJ, 122, 2644. HST observations.

Reynoso & Goss 2002, ApJ, 575, 871. VLA at 5 GHz  $(6''.0 \times 6''.4)$  for H<sub>2</sub>CO absorption studies.

Krause et al. 2005, Science, 308, 1604. Spitzer light echoes from surroundings.

Ennis et al. 2006, ApJ, 652, 376. Spitzer observations.

Kang & Koo 2007, ApJS, 173, 85. SGPS of high velocity HI.

Rho et al. 2008, ApJ, 673, 271. Spitzer observations.

Rest et al. 2008, ApJ, 681, L81. SN light echo.

Krause et al. 2008, Science, 320, 1195. SN light echo spectrum. Helmboldt & Kassim 2009, AJ, 138, 838. Low radio frequency temporal variations.

Patnaude & Fesen 2009, ApJ, 697, 535. X-ray proper motions. Barlow *et al.* 2010, A&A, 518, L138. Herschel observations.

Sibthorpe et al. 2010, ApJ, 719, 1553. Akari and sub-mm observations of region.

DeLaney et al. 2010, ApJ, 725, 2038. Spitzer and Chandra observations for 3-D structure.

Patnaude et al. 2011, ApJ, 729, L28. Chandra observations of fad-

Rest et al. 2011, ApJ, 732, 3. SN light echo observations. Fesen et al. 2011, ApJ, 736, 109. HST variability studies.

Besel & Krause 2012, A&A, 541, L3. IR light echoes.

Rho et al. 2012, ApJ, 747, L6. Akari CO observations.

Vogt et al. 2012, ApJ, 750, 155. Spitzer light echoes.

Asgekar et al. 2013, A&A, 551, L11. LOFAR of carbon recombination lines.

Yang et al. 2013, ApJ, 766, 44. Suzaku spectroscopy.

Rutherford et al. 2013, ApJ, 769, 64. Chandra spectroscopy.

Milisavljevic & Fesen 2013, ApJ, 772, 134. Optical spectroscopy for 3-D structure.

Koo et al. 2013, Science, 342, 1346. IR spectroscopy.

DeLaney *et al.* 2014, ApJ, 785, 7. VLA plus Pie Town at 74 MHz (9") and comparison with higher frequencies.

Arendt et al. 2014, ApJ, 786, 55. Spitzer and Herschel IR observations.

Patnaude & Fesen 2014, ApJ, 789, 138. Multi-epoch optical and X-ray observations.

Lee et al. 2014, ApJ, 789, 7. Spitzer and CO observations.

Vinyaikin 2014, ARep, 58, 626. Time evolution of radio emission.

Alarie et al. 2014, MNRAS, 441, 2996. Optical imag-

ing/spectroscopy, including distance from expansion.

Grefenstette et al. 2014, Nature, 506, 339. NuSTAR <sup>44</sup>Ti observations.

Grefenstette et al. 2015, ApJ, 802, 15. NuSTAR observations.Lee et al. 2015, ApJ, 808, 98. WISE, Spitzer and other IR observations.

Milisavljevic & Fesen 2015, Science, 347, 526. Near-IR observations.

Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 7 frequencies between 30 and 353 GHz.

Kilpatrick *et al.* 2016, ApJ, 816, 1. CO observations, including broad lines.

Fesen & Milisavljevic 2016, ApJ, 818, 17. HST [SIII] and [SII] observations.

Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Wang & Li 2016, ApJ, 825, 102. INTEGRAL observations.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Oonk *et al.* 2017, MNRAS, 465, 1066. LOFAR observations, including H and C lines, 33 to 78 MHz, including 69 MHz (9".8×11".2).

De Looze *et al.* 2017, MNRAS, 465, 3309. Spitzer and Herschel observations.

Trotter *et al.* 2017, MNRAS, 469, 1299. Time evolution of radio emission.

Grefenstette et al. 2017, ApJ, 834, 19. NuSTAR observations. Sato et al. 2017, ApJ, 836, 225. Multi-epoch Chandra observa-

Lee et al. 2017, ApJ, 837, 118. IR observations.

Arias *et al.* 2018, A&A, 612, A110. LOFAR at 30 to 77 MHz  $(7'' \times 17'')$ .

Salas *et al.* 2018, MNRAS, 475, 2496. LOFAR C recombination lines as 43, 54, 148 and 340 MHz (70").

Sato et al. 2018, ApJ, 853, 46. Chandra and NuSTAR observations.

Zhou et al. 2018, ApJ, 865, 6. CO observations.

Raymond et al. 2018, ApJ, 866, 128. IR observations.

Koo et al. 2018, ApJ, 866, 139. IR observations.

# G113.0+0.2

**RA:**  $23^{\text{h}}26^{\text{m}}50^{\text{s}}$  **1-GHz flux/Jy:** 4 **Size/arcmin:**  $40 \times 17$ ? **Dec:**  $+61^{\circ}26'$  **Spectral index:** 0.5? **Type:** ?

**Radio:** Elongated, extent not well defined. **Point sources:** Contains old pulsar.

**Distance:** Association for H<sub>I</sub> indicates 3.1 kpc.

### References:

Kothes et al. 2005, A&A, 444, 871. CGPS at 408 MHz ( $2'.8\times3'.1$ ) and 1.4 GHz ( $49''\times55''$ ) including HI.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3') and 1420 MHz ( $\sim$  1').

Sun et al. 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9!5: S =  $1.9\pm0.5$  Jy) including polarisation and review of flux densities.

# G114.3+0.3

**RA:**  $23^h37^m00^s$  **1-GHz flux/Jy:** 5.5 **Size/arcmin:**  $90\times55$  **Dec:**  $+61^\circ55'$  **Spectral index:** 0.5 **Type:** S

Radio: Shell, with HII region S165 within the boundary of the remnant.

**Optical:** Faint emission in centre and to S. **Point sources:** Pulsar near centre of remnant.

**Distance:** Association with H<sub>I</sub> and other features implies 0.7 kpc.

#### References:

Reich & Braunsfurth 1981, A&A, 99, 17. Effelsberg 100-m at 2.7 GHz (4'4: S = 3.6 Jy) and S<sub>1.4 GHz</sub> = 4.4 Jy from 1.4 GHz survey data, plus HI from Maryland–Green Bank survey. Kulkarni *et al.* 1993, Nature, 362, 135. Pulsar association. Fürst *et al.* 1993, A&A, 276, 470. Pulsar association. Becker *et al.* 1996, A&A, 306, 464. ROSAT of pulsar. Fesen *et al.* 1997, AJ, 113, 767. Optical observations. Reich 2002, in NSPS, p1. Effelsberg 100-m at 2.7 GHz. Mavromatakis *et al.* 2002, A&A, 383, 1011. Optical observations.

Yar-Uyanıker *et al.* 2004, ApJ, 616, 247. CGPS at 1.4 GHz  $(49'' \times 55'')$ , including Hi  $(1'.0 \times 1'.1)$ .

Tian & Leahy 2006, ChJAA, 6, 543. CGPS at 408 MHz ( $3'.4\times3'.9: S=12.0\pm6.0$  Jy) and 1.4 GHz ( $1'.0\times1'.1:S=9.8\pm0.8$  Jy).

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 1420 MHz ( $\sim 1'$ :  $S = 5.4 \pm 0.8$  Jy), including review of flux densities.

Gao et al. 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9'.5:  $S = 6.9 \pm 0.7$  Jy), including polarisation and review of flux densities.

# G116.5 + 1.1

**RA:**  $23^h53^m40^s$  **1-GHz flux/Jy:** 10 **Size/arcmin:**  $80\times60$  **Dec:**  $+63^{\circ}15'$  **Spectral index:** 0.5 **Type:** S

**Radio:** Distinct shell, with high polarisation.

**Optical:** Detected.

**Distance:** Association with H<sub>I</sub> features implies 1.6 kpc.

# References:

Reich & Braunsfurth 1981, A&A, 99, 17. Effelsberg 100-m at 2.7 GHz ( $4/4:S=4.7\pm0.4$  Jy) and  $S_{1.4~GHz}=8.0\pm0.8$  Jy from 1.4 GHz survey data, plus HI from Maryland–Green Bank survey.

Fesen *et al.* 1997, AJ, 113, 767. Optical observations. Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search. Yar-Uyanıker *et al.* 2004, ApJ, 616, 247. CGPS at 1.4 GHz (49"×55"), including H<sub>I</sub> (1.0×1.1).

Mavromatakis et al. 2005, A&A, 435, 141. Optical observations.

Tian & Leahy 2006, ChJAA, 6, 543. CGPS at 408 MHz ( $3'.4\times3'.8:$   $S=15.0\pm1.5$  Jy) and 1.4 GHz ( $1'.0\times1'.1:S=10.6\pm0.$  Jy).

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3': S = 12.5 $\pm$ 1.6 Jy) and 1420 MHz ( $\sim$  1': S = 10.3 $\pm$ 0.70 Jy), including polarisation and review of flux densities.

Gao et al. 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9'.5:  $S = 5.7 \pm 0.6$  Jy), including polarisation and review of flux densities.

G116.9+0.2

Has been called G117.3+0.1 and G116.9+0.1.

Radio: Incomplete shell.

**Optical:** Filaments on sky survey.

**X-ray:** Centrally brightened, with NE 'breakout'.

**Point sources:** Pulsar outside rim to E.

**Distance:** Association with H<sub>I</sub> features implies 1.6 kpc.

#### References:

Willis 1973, A&A, 26, 237. NRAO 300-ft at 2.7 GHz (5' :  $S = 3.9 \pm 1.0$  Jy), and 37-m at 1.7 GHz ( $S = 5.5 \pm 2.0$  Jy), plus review of flux densities.

van den Bergh *et al.* 1973, ApJS, 26, 19. Optical observations. Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz  $(5': S=4.2\pm0.2 \text{ Jy})$ .

Angerhofer *et al.* 1977, A&A, 55, 11. NRAO 140-ft at 5 GHz (6.8:  $S = 3.0 \pm 0.3$  Jy).

Dickel & Willis 1980, A&A, 85, 55. WSRT at 610 MHz (56"×64") and 1.4 GHz (22"×25").

Reich & Braunsfurth 1981, A&A, 99, 17. Effelsberg 100-m at 2.7 GHz ( $4/4:S=4.8\pm0.4$  Jy) and  $S_{1.4~GHz}=7.8\pm0.8$  Jy from 1.4 GHz survey data, plus HI from Maryland–Green Bank survey.

Lozinskaya 1981, SvAL, 7, 17. Mean optical velocity. Landecker *et al.* 1982, AJ, 87, 1379. DRAO at 1.42 GHz and HI (2'×2'.3:S=8.3±0.5 Jy), plus review of flux densities. *see also*: Landecker *et al.* 1983, AJ, 88, 877. Erratum. Fesen *et al.* 1985, ApJ, 292, 29. Optical spectra. Hailey & Craig 1994, ApJ, 434, 635. Optical spectroscopy. Hailey & Craig 1995, ApJ, 455, L151. ROSAT of nearby pulsar. Fesen *et al.* 1997, AJ, 113, 767. Optical observations.

Craig *et al.* 1997, ApJ, 488, 307. ROSAT observations. Reich 2002, in NSPS, p1. Effelsberg 100-m at 10.6 GHz (69"), including polarisation.

Yar-Uyanıker *et al.* 2004, ApJ, 616, 247. CGPS at 1.4 GHz  $(49'' \times 55'')$ , including Hi  $(1'.0 \times 1'.1)$ .

Lazendic & Slane 2006, ApJ, 647, 350. X-ray observations. Tian & Leahy 2006, ChJAA, 6, 543. CGPS at 408 MHz ( $3'.4\times3'.8:$   $S=15.0\pm1.5$  Jy) and 1.4 GHz ( $1'.0\times1'.1:S=8.1\pm0.4$  Jy).

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3': S = 10.5 $\pm$ 0.8 Jy) and 1420 MHz ( $\sim$  1': S = 7.0 $\pm$ 0.8 Jy), including review of flux densities.

Pannuti et al. 2010, AJ, 140, 1787. ASCA and Chandra observations.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9!5 : S =  $3.6\pm0.4$  Jy) including polarisation and review of flux densities.

Clark *et al.* 2017, ApJ, 834, 106. Fermi pulsar detection. Wu *et al.* 2018, ApJ, 854, 99. Radio detection of pulsar.

Zyuzin et al. 2018, MNRAS, 476, 2177. X-ray observations of pulsar.

Katsuragawa et al. 2018, PASJ, 70, 110. Suzaku observations.

**G119.5+10.2** CTA 1

**RA:** 00<sup>h</sup>06<sup>m</sup>40<sup>s</sup> **1-GHz flux/Jy:** 36 **Size/arcmin:** 90? **Dec:** +72°45′ **Spectral index:** 0.6 **Type:** S

Has been called G119.5+10.3.

Radio: Incomplete shell, with 'breakout' to NW.

**Optical:** Faint diffuse nebulosities. **X-ray:** Centrally brightened. **Point sources:** Central pulsar.

**Distance:** Associated H<sub>I</sub> shell indicates 1.4 kpc.

#### References:

Sieber *et al.* 1979, A&A, 74, 361. Effelsberg 100-m at 2.7 GHz  $(4.4:S=23.6\pm2.8 \text{ Jy})$ .

Sieber *et al.* 1981, A&A, 103, 393. Effelsberg 100-m at 2.7 GHz and 1.7 GHz ( $7'.6: S = 31.6 \pm 2.5$  Jy), and 151 MHz ( $4': S = 62.6 \pm 6$  Jy).

Fesen *et al.* 1981, ApJ, 247, 148. Optical, including spectra. Fesen *et al.* 1983, ApJS, 51, 337. Deep [OIII] imagery. Seward 1990, ApJS, 73, 781. Einstein observations.

Simonetti 1992, ApJ, 386, 170. VLA observations of background sources for rotation measure studies.

Pineault *et al.* 1993, AJ, 105, 1060. DRAO at 1.4 GHz (1 $^\prime$  :  $S = 34 \pm 4$  Jy), plus HI and IRAS.

Seward *et al.* 1995, ApJ, 453, 284. ROSAT observations. Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Slane *et al.* 1997, ApJ, 485, 221. ASCA and ROSAT observations.

Pineault *et al.* 1997, A&A, 324, 1152. DRAO at 408 MHz (3'.5) and 1.4 GHz (1'.0). Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search.

Brazier *et al.* 1998, A&A, 331, 1002. Fulsal search.

Brazier *et al.* 1998, MNRAS, 295, 819. Studies of central, compact γ-ray source.

Mavromatakis et al. 2000, A&A, 353, 371. Optical imagery.

Roberts et al. 2001, ApJS, 133, 451. ASCA observations.Reich 2002, in NSPS, p1. Effelsberg 100-m at 1.4 GHz (9'.4).Slane et al. 2004, ApJ, 601, 1045. XMM-Newton observations of central source.

Halpern et al. 2004, ApJ, 612, 398. Chandra observations of central nebula, plus optical and radio limits for compact source.
Casandjian & Grenier 2008, A&A, 489, 849. γ-ray observations.
Abdo et al. 2008, Science, 322, 1218. Fermi detection of pulsar.
Lin et al. 2010, ApJ, 725, L1. XMM-Newton observations of pulsar.

Caraveo et al. 2010, ApJ, 725, L6. XMM-Newton observations of pulsar.

Sun *et al.* 2011, A&A, 535, A64. Urumqi 25-m at 4.8 GHz (9.5:  $S = 11.6 \pm 1.2$  Jy) and Effelsberg 100-m at 2.6 GHz (4.4:  $S = 20.3 \pm 2.0$  Jy) including polarisation.

Lin *et al.* 2012, MNRAS, 426, 2283. Suzaku observations. Mignani *et al.* 2013, MNRAS, 430, 1354. Optical limits for pulsar. Aliu *et al.* 2013, ApJ, 764, 38. γ-ray observations. Li *et al.* 2016, ApJ, 831, 19. Fermi observations of pulsar. Ackermann *et al.* 2018, ApJS, 237, 32. Fermi observations.

G120.1+1.4 Tycho, 3C10, SN1572

RA:  $00^h25^m18^s$ 1-GHz flux/Jy: 50Size/arcmin: 8Dec:  $+64^\circ09'$ Spectral index: 0.58Type: S

This is the remnant of the Tycho's SN of AD1572.

**Radio:** Shell, brightest to the NE.

Optical: Faint filaments/knots to the NNW, NE and E.

**X-ray:** Shell, brighter to the NE.

**Point sources:** Faint radio source near centre of the remnant, thought to be extragalactic.

**Distance:** HI observations suggest 2.3–3 kpc, optical proper motion and shock velocity gives 2.4 kpc.

### References:

Duin & Strom 1975, A&A, 39, 33. WSRT at 610 MHz  $(57''\times64'')$  and 5 GHz  $(7''\times8'').$ 

Klein et al. 1979, A&A, 76, 120. Effelsberg 100-m at 10.7 GHz  $(1'.2:S=13.1\pm0.8 \text{ Jy})$ , plus review of flux densities.

Strom et al. 1982, MNRAS, 200, 473. WSRT at 1415 MHz (27"×31") from 1971 and 1979, for expansion.

Dickel et al. 1982, ApJ, 257, 145. Comparison of radio, X-ray and optical observations.

Reid et al. 1982, ApJ, 261, 485. Einstein observations.

Seward et al. 1983, ApJ, 266, 287. Einstein observations.

Tan & Gull 1985, MNRAS, 216, 949. Radio at 2.7 GHz (4") from 1980 and 1983, and One-Mile Telescope at 1.4 GHz (23") from 1965 and 1980, for expansion.

Albinson et al. 1986, MÑRAS, 219, 427. Hi observations.

Green & Gull 1987, MNRAS, 224, 1055. VLA HI absorption observations towards central radio source.

Wood *et al.* 1992, AJ, 103, 1338. VLA at 5 GHz (1"5) polarisation studies

Vancura et al. 1995, ApJ, 441, 680. X-ray spectra and ROSAT im-

age.
Predehl & Schmitt 1995, A&A, 293, 889. ROSAT of dust scattered halo

Schwarz et al. 1995, A&A, 299, 193. WSRT and Effelsberg 100-m at 1.4 GHz (50"), and VLA at 1.4 GHz (13"), for neutral hydrogen studies.

Hwang & Gotthelf 1997, ApJ, 475, 665. ASCA observations. Reynoso *et al.* 1997, ApJ, 491, 816. VLA at 1.4 GHz (1".4×1".5) from 1984 and 1994 for expansion studies.

Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Reynoso et al. 1999, AJ, 117, 1827. VLA at 1.4 GHz (13"), for HI studies.

Katz-Stone *et al.* 2000, ApJ, 529, 453. VLA at 333 MHz and 1.4 GHz (7".7×9".5), for spectral index studies.

Ghavamian *et al.* 2000, ApJ, 535, 266. Optical observations of shock precursors.

see also: Ghavamian et al. 2017, ApJ, 843, 77. Erratum.

Hughes 2000, ApJ, 545, L53. ROSAT X-ray expansion.

Decourchelle et al. 2001, A&A, 365, L218. XMM-Newton observations

Ghavamian et al. 2001, ApJ, 547, 995. Optical spectroscopy. Douvion et al. 2001, A&A, 373, 281. ISO observations.

Lee et al. 2004, ApJ, 605, L113. Observations of molecular clouds in vicinity.

Warren et al. 2005, ApJ, 634, 376. Chandra observations.

Kothes et al. 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3':  $S=86\pm5$  Jy) and 1420 MHz ( $\sim$  1':  $S=40.5\pm1.5$  Jy), including polarisation and review of flux densities.

Cassam-Chenaï et al. 2007, ApJ, 665, 315. Chandra observations. Rest et al. 2008, ApJ, 681, L81. SN light echo.

Krause et al. 2008, Nature, 456, 617. SN light echo spectrum. Hurley-Walker et al. 2009, MNRAS, 396, 365. Radio observations at 14 to 18 GHz.

Raymond et al. 2010, ApJ, 712, 901. Ha spectroscopy.

Lee et al. 2010, ApJ, 715, L146. Hα observations.

Hayato et al. 2010, ApJ, 725, 894. Suzaku observations of expansion.

Ishihara et al. 2010, A&A, 521, L61. Akari observations.

Tian & Leahy 2011, ApJ, 729, L15. HI and CO observations.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 20.0±2.0 Jy) including polarisation and review of flux densities. Giordano *et al.* 2012, ApJ, 744, L2. Fermi observations.

Gomez et al. 2012, MNRAS, 420, 3557. Herschel IR dust observations.

Yang et al. 2013, ApJ, 766, 44. Suzaku spectroscopy.

Williams *et al.* 2013, ApJ, 770, 129. Spitzer observations.

Yamaguchi et al. 2014, ApJ, 780, 136. Suzaku observations.

Wang & Li 2014, ApJ, 789, 123. INTEGRAL observations.

Troja et al. 2014, ApJ, 797, L6. Swift observations.

Miceli et al. 2015, ApJ, 805, 120. XMM-Newton observations.

Lu et al. 2015, ApJ, 805, 142. Chandra observations.

Katsuda et al. 2015, ApJ, 808, 49. Suzaku observations.

Tran *et al.* 2015, ApJ, 812, 101. Chandra observations. Lopez *et al.* 2015, ApJ, 814, 132. NuSTAR observations.

Arnaud et al. 2016, A&A, 586, A134. Planck flux densities at 5 frequencies between 30 and 143 GHz.

Koo et al. 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Williams *et al.* 2016, ApJ, 823, L32. Multi-epoch X-ray and radio observations for expansion.

Zhou et al. 2016, ApJ, 826, 34. CO observations.

Chen et al. 2017, A&A, 604, A113. CO observations.

Yamaguchi et al. 2017, ApJ, 834, 124. Suzaku observations of E

Archambault *et al.* 2017, ApJ, 836, 23. γ-ray observations.

Sato & Hughes 2017, ApJ, 840, 112. Chandra and Suzaku observations.

Williams et al. 2017, ApJ, 842, 28. Multi-epoch Chandra observations for expansion studies.

Knežević et al. 2017, ApJ, 846, 167. Hα observations of NE.

Kerzendorf et al. 2018, MNRAS, 479, 5696. HST search for progenitor companion.

Vinyaikin et al. 2018, ARep, 62, 130. Time evolution of radio emission.

# G126.2+1.6

**RA**: 01<sup>h</sup>22<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: 6 Size/arcmin: 70 **Dec:**  $+64^{\circ}15'$ Spectral index: 0.5 Type: S?

Radio: Poorly defined shell. Optical: Filaments, mostly in W.

Reich et al. 1979, A&A, 78, L13. Effelsberg 100-m at 1.4 GHz  $(9': S=6.8\pm0.7 \text{ Jy})$  and 2.7 GHz  $(4.4: S=3.9\pm0.4 \text{ Jy})$ .

Blair et al. 1980, ApJ, 242, 592. Optical detection and spectra.

Rosado 1982, RMxAA, 5, 127. Optical spectra.

Fesen et al. 1983, ApJS, 51, 337. Deep [OIII] imagery. Fürst et al. 1984, A&A, 133, 11. Effelsberg 100-m at 2.7 GHz

(4.4) and 4.8 GHz (2.6).

Joncas et al. 1989, A&A, 219, 303. DRAO at 408 MHz (3'.5×3'.9:  $S = 12\pm2.5$  Jy) and part at 1.4 GHz (1.0×1.1), plus review of flux densities.

Xilouris et al. 1993, A&A, 270, 393. Optical imaging. Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Reich et al. 2003, A&A, 408, 961. Effelsberg 100-m at 863 MHz  $(14.5:S=6.1\pm1.6 \text{ Jy}).$ 

Boumis *et al.* 2005, A&A, 443, 175. Optical observations. Tian & Leahy 2006, A&A, 447, 205. CGPS at 408 MHz (3'.4×3'.8:  $S = 9.7 \pm 3.9$  Jy) and 1.4 GHz (1'.0×1'.1:  $S = 6.7 \pm 2.1$  Jy), plus other observations for spectral index studies.

Kothes et al. 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3' : S =$  $5.7\pm0.7 \text{ Jy}$ ) and 1420 MHz ( $\sim 1': S = 6.4\pm1.1 \text{ Jy}$ ), including review of flux densities.

Sun et al. 2007, A&A, 463, 993. Urumqi 25-m at 5 GHz (9.5: S =  $2.6\pm0.6$  Jy), including polarisation.

see also: Sun et al. 2007, A&A, 469, 1003. Erratum.

R5 G127.1+0.5

**RA**: 01<sup>h</sup>28<sup>m</sup>20<sup>s</sup> 1-GHz flux/Jv: 12 Size/arcmin: 45 **Dec:** +63°10′ Spectral index: 0.45 Type: S

Has been called G127.3+0.7.

**Radio:** Distinct shell, with bright central source.

Optical: Detected.

**Point sources:** Flat radio spectrum (extragalactic) source at centre of remnant.

**Distance:** 1.2–1.3 kpc if associated with NGC 559.

### References:

Caswell 1977, MNRAS, 181, 789. Half-Mile Telescope at 1.42 GHz  $(3.5 \times 3.9)$ , plus other observations of central source. Pauls 1977, A&A, 59, L13. Effelsberg 100-m at 1.4 GHz (9': S =

Salter et al. 1978, A&A, 66, 77. Effelsberg 100-m at 2.7 GHz (4.4), plus 5 and 8.7 GHz of the central source.

Reich et al. 1979, A&A, 78, L13. Effelsberg 100-m at 1.4 GHz  $(9': S=10.8\pm 1.3 \text{ Jy}).$ 

Pauls et al. 1982, A&A, 112, 120. WSRT at 610 MHz (56"×62") and HI absorption to the point source with the VLA.

Geldzahler & Shaffer 1982, ApJ, 260, L69. Observations of central source.

Fürst et al. 1984, A&A, 133, 11. Effelsberg 100-m at 2.7 GHz (4.4) and 4.8 GHz (2.6).

Goss & van Gorkom 1984, JApA, 5, 425. WSRT HI absorption of central source.

Joncas et al. 1989, A&A, 219, 303. DRAO at 408 MHz (3:5×3:9:  $S = 17.9 \pm 2.0 \text{ Jy}$ ) and 1.4 GHz (1'.0×1'.13:  $S = 10.1 \pm 0.8 \text{ Jy}$ ), plus review of flux densities.

Xilouris et al. 1993, A&A, 270, 393. Optical imaging.

Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Reich et al. 2003, A&A, 408, 961. Effelsberg 100-m at 863 MHz  $(14.5: S = 14.6 \pm 0.8 \text{ Jy}).$ 

Kaplan et al. 2004, ApJS, 153, 269. Chandra limits for any compact sources.

Leahy & Tian 2006, A&A, 451, 251. CGPS at 408 MHz (3'.4×3'.8:  $S = 17.1 \pm /1.7$  Jy) and 1.4 GHz (1'0×1'.2:  $S = 10.0 \pm /0.8$  Jy).

Kothes et al. 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3' : S =$  $15.9\pm1.0 \text{ Jy}$ ) and 1420 MHz ( $\sim 1': S = 9.7\pm0.6 \text{ Jy}$ ), including polarisation and review of flux densities.

Sun et al. 2007, A&A, 463, 993. Urumqi 25-m at 5 GHz (9'.5: S =  $6.3\pm0.7$  Jy), including polarisation.

see also: Sun et al. 2007, A&A, 469, 1003. Erratum. Zhou et al. 2014, ApJ, 791, 109. CO observations.

G130.7+3.1 3C58, SN1181

**RA:** 02<sup>h</sup>05<sup>m</sup>41<sup>s</sup> **1-GHz flux/Jy:** 33 **Size/arcmin:** 9×5 **Dec:** +64°49′ **Spectral index:** 0.07 **Type:** F

This is the remnant of the SN of AD1181.

**Radio:** Filled-centre, highly polarised, with high frequency turnover.

**Optical:** Faint filaments.

**X-ray:** Centrally brightened, with faint jet.

Point sources: Central pulsar.

**Distance:** HI absorption indicates 2 kpc.

#### References:

Green *et al.* 1975, A&A, 44, 187. Effelsberg 100-m at 15.0 GHz  $(58'': S=26.7\pm0.5 \text{ Jy})$ .

Wilson & Weiler 1976, A&A, 49, 357. WSRT at 610 MHz  $(58'' \times 64'')$ , 1.4 GHz  $(24'' \times 27'')$  and 5 GHz  $(7'' \times 8'')$ . van den Bergh 1978, ApJ, 220, L9. Optical observations.

Becker et al. 1982, ApJ, 255, 557. X-ray observations.

Green & Gull 1982, Nature, 299, 606. HI absorption distance. Fesen 1983, ApJ, 270, L53. Optical spectra.

Reynolds & Aller 1985, AJ, 90, 2312. VLA at 1.4 GHz (2'), for limits of shell.

Davelaar et al. 1986, ApJ, 300, L59. EXOSAT spectrum.

Green 1986, MNRAS, 218, 533. 151 MHz observations (1.2×1.3:  $S = 36 \pm 4$  Jy), plus 2.7 GHz (4"), plus Einstein observations for limit on shell.

Green 1987, MNRAS, 225, 11P. Flux density increase at 408 MHz

Morsi & Reich 1987, A&AS, 69, 533. Effelsberg 100-m at 32 GHz (26".5: S=24.2±1.4 Jy).

Reynolds & Aller 1988, ApJ, 327, 845. VLA at 1.4 (2".4) and 4.9 GHz (2".5).

Salter *et al.* 1989, ApJ, 338, 171. NRAO 12-m at 84.2 GHz (90":  $S=15.0\pm2.0$  Jy), plus review of flux densities.

S=15.0±2.0 Jy), plus review of flux densities. Asaoka & Koyama 1990, PASJ, 42, 625. Ginga X-ray spectrum. Green & Scheuer 1992, MNRAS, 258, 833. IRAS upper limits.

Roberts *et al.* 1993, A&A, 274, 427. HI absorption.

Wallace et al. 1994, A&A, 286, 565. HI of surroundings.

Helfand et al. 1995, ApJ, 453, 741. ROSAT observations.

Torii *et al.* 2000, PASJ, 52, 875. ASCA observations. Bietenholz *et al.* 2001, ApJ, 560, 772. VLA at 74 MHz (26'': S = 33.6 Jy) and 327 MHz (8''2: S = 33.9 Jy), for spectral index and

expansion studies. Bocchino et al. 2001, A&A, 369, 1078. XMM-Newton observa-

Murray et al. 2002, ApJ, 568, 226. Chandra pulsar detection.

Camilo et al. 2002, ApJ, 571, L41. Pulsar detection in radio. Reich 2002, in NSPS, p1. Effelsberg 100-m at 32 GHz (26") for polarised intensity.

Slane *et al.* 2002, ApJ, 571, L45. Chandra observations. Slane *et al.* 2004, ApJ, 616, 403. Deep Chandra imaging.

Bietenholz 2006, ApJ, 645, 1180. VLA at 1.4 GHz (1".36) for expansion studies.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3': S = 32.2 \pm 2.0$  Jy) and 1420 MHz ( $\sim 1': S = 31.9 \pm 1.0$  Jy), including polarisation and review of flux densities.

Gotthelf *et al.* 2007, ApJ, 654, 267. XMM-Newton observations. Slane *et al.* 2008, ApJ, 676, L33. Spitzer and other IR observations.

Fesen *et al.* 2008, ApJS, 174, 379. Optical observations for proper motion studies.

Shearer & Neustroev 2008, MNRAS, 390, 235. Optical observations of pulsar nebula.

Shibanov et al. 2008, A&A, 486, 273. Optical observations of pulsar nebula.

Abdo *et al.* 2009, ApJ, 699, L102. Fermi observations of pulsar. Livingstone *et al.* 2009, ApJ, 706, 1163. Pulsar observations.

Hurley-Walker *et al.* 2009, MNRAS, 396, 365. Radio observations at 14 to 18 GHz.

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9'.5: S = 31.7±3.0 Jy) including polarisation and review of flux densities. Bietenholz *et al.* 2013, MNRAS, 431, 2590. Proper motion study of pulsar.

Kothes et al. 2013, A&A, 560, A18. CGPS HI observations for dis-

Aleksić et al. 2014, A&A, 567, L8. γ-ray detection.

Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 7 frequencies between 30 and 353 GHz.

G132.7+1.3

**RA:** 02<sup>h</sup>17<sup>m</sup>40<sup>s</sup> **1-GHz flux/Jy:** 45 **Size/arcmin:** 80 **Dec:** +62°45′ **Spectral index:** 0.6 **Type:** S

Has been called G132.4+2.2.

**Radio:** Faint shell, adjacent to W3/4/5 complex.

**Optical:** Complete, filamentary shell, shock excited spectra.

X-ray: Partial shell.

**Point sources:** Pulsar nearby.

**Distance:** Interaction with surroundings gives 2 kpc.

#### References:

van den Bergh *et al.* 1973, ApJS, 26, 19. Optical observations. Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz  $(5': S=33.8\pm7.0 \text{ Jy})$ .

d'Odorico & Sabbadin 1977, A&AS, 28, 439. Optical spectra. Read 1981, MNRAS, 194, 863. Radio at 151 MHz (4.4) and 1.4 GHz (2') showing HI shell.

Lozinskaya 1981, SvAL, 7, 17. Mean optical velocity. Fesen & Gull 1983, PASP, 95, 196. Optical image. Leahy *et al.* 1985, ApJ, 294, 183. Einstein observations. Landecker *et al.* 1987, AJ, 94, 111. DRAO at 408 MHz (3'.5×4'):

 $S=75\pm15$  Jy), plus review of flux densities. Routledge *et al.* 1991, A&A, 247, 529. DRAO at 1.4 GHz (1'.0×1'.1) for HI, plus CO observations.

Fesen *et al.* 1995, AJ, 110, 2876. Optical imaging and spectroscopy, DRAO at 408 MHz (3/5×4′) and 1.4 GHz (1/0×1/1).

Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search. Koralesky *et al.* 1998, AJ, 116, 1323. VLA detection of

Koralesky *et al.* 1998, AJ, 116, 1323. VLA detection of compact OH emission.

Reich *et al.* 2003, A&A, 408, 961. Effelsberg 100-m at 863 MHz  $(14'5:S=51.5\pm3.5 \text{ Jy})$ .

Tian & Leahy 2005, A&A, 436, 187. CGPS at 408 MHz (3'.4×3'.8:) and 1.4 GHz (1'.0×1'.1), for spectral index studies. see also: Tian & Leahy 2006, A&A, 451, 991. Erratum.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3':  $S = 61 \pm 9$  Jy) and 1420 MHz ( $\sim$  1':  $S = 29.4 \pm 2.7$  Jy), including review of flux densities.

Lazendic & Slane 2006, ApJ, 647, 350. X-ray observations. Green 2007, BASI, 35, 77. Review of radio spectrum. Shi *et al.* 2008, A&A, 487, 601. Urumqi 25-m at 4.8 GHz (9/.5), plus other survey observations for spectral studies.

Casandjian & Grenier 2008, A&A, 489, 849. γ-ray observations. Kilpatrick et al. 2016, ApJ, 816, 1. CO observations, including broad lines.

Acero *et al.* 2016, ApJS, 224, 8. Fermi observations. Katagiri *et al.* 2016, ApJ, 818, 114. Fermi observations. Zhou *et al.* 2016, ApJ, 833, 4. CO observations of region.

# G150.3 + 4.5

**RA:** 04<sup>h</sup>27<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** ? **Size/arcmin:** 180×150 **Dec:** +55°28′ **Spectral index:** ? **Type:** S

Radio: Faint radio shell.

### References:

Gao & Han 2014, A&A, 567, A59. Radio surveys at 1.4, 2.7 and 5 GHz (9'.4,4'.3,9'.5).

Ackermann et al. 2017, ApJ, 843, 139. Fermi observations. Ackermann et al. 2018, ApJS, 237, 32. Fermi observations.

# G152.4 - 2.1

**RA:**  $04^h07^m50^s$  **1-GHz flux/Jy:** 3.5? **Size/arcmin:**  $100\times95$  **Dec:**  $+49^\circ11'$  **Spectral index:** 0.7? **Type:** S

Radio: Bilateral shell.

### References

Foster *et al.* 2013, A&A, 549, A107. Effelsberg 100-m at 2.7 GHz, including polarisation, plus various radio survey observations.

# G156.2 + 5.7

**RA:** 04<sup>h</sup>58<sup>m</sup>40<sup>s</sup> **1-GHz flux/Jy:** 5 **Dec:** +51°50′ **Spectral index:** 0.5 **Size/arcmin:** 110 **Type:** S

Radio: Faint shell, brighter in E and W.

Optical: Filamentary ring and smaller patchy ring.

X-ray: Faint shell.

**Distance:** Optical/X-ray observations imply >1.7 kpc.

#### References:

Pfeffermann *et al.* 1991, A&A, 246, L28. ROSAT detection. Reich *et al.* 1992, A&A, 256, 214. Effelsberg 100-m at 1.4 (9' :  $S = 4.2 \pm 1.0$  Jy) and 2.7 GHz (4'.3 :  $S = 3.0 \pm 1.0$  Jy), plus HI and IRAS

Yamauchi et al. 1993, PASJ, 45, 795. Hard X-ray observations. Lorimer et al. 1998, A&A, 331, 1002. Pulsar search.

Yamauchi et al. 1999, PASJ, 51, 13. ASCA observations of some regions.

Reich 2002, in NSPS, p1. Effelsberg 100-m at 2.7 GHz, including polarisation.

Pannuti & Allen 2004, AdSpR, 33, 434. ASCA and RXTE observations. Kaplan et al. 2006, ApJS, 163, 344. X-ray upper limit on compact sources.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3'$ :  $S = 8.1 \pm 1.3$  Jy), including review of flux densities.

Xu *et al.* 2007, A&A, 470, 969. Urumqi 25-m at 4.8 GHz (9.5:  $S = 2.5 \pm 0.5$  Jy), including polarisation.

Gerardy & Fesen 2007, MNRAS, 376, 929. Optical observations. Katsuda *et al.* 2009, PASJ, 61, S155. Suzaku observations. Uchida *et al.* 2012, PASJ, 64, 61. Suzaku observations. Katsuda *et al.* 2016, ApJ, 826, 108. Multi-epoch Hα observations

Katsuda *et al.* 2016, ApJ, 826, 108. Multi-epoch Hα observations for expansion.

# G159.6 + 7.3

**RA:**  $05^{\rm h}20^{\rm m}00^{\rm s}$  **1-GHz flux/Jy:** ? **Size/arcmin:**  $240\times180$ ? **Dec:**  $+50^{\circ}00'$  **Spectral index:** ? **Type:** S

Radio: Not detected.
Optical: Large, faint shell.
X-ray: Possible emission.

### References:

Fesen & Milisavljevic 2010, AJ, 140, 1163. H $\alpha$  and ROSAT observations

G160.9+2.6

 RA:  $05^h01^m00^s$  1-GHz flux/Jy: 110
 Size/arcmin:  $140 \times 120$  

 Dec:  $+46^{\circ}40'$  Spectral index: 0.64 Type: S

Has been called G160.5+2.8 and G160.4+2.8.

Radio: Large, filamentary shell.Optical: Incomplete shell.X-ray: Centrally brightened.

**Point sources:** Pulsar within boundary of the remnant, plus several nearby compact radio sources.

**Distance:** Various observations suggests less than 4 kpc.

#### References:

d'Odorico & Sabbadin 1977, A&AS, 28, 439. Optical spectra. Damashek *et al.* 1978, ApJ, 225, L31. Pulsar.

Lozinskaya 1981, SvAL, 7, 17. Mean optical velocity.

Dwarakanath *et al.* 1982, JApA, 3, 207. Radio observations at 34.5 MHz  $(26' \times 40' : S = 750 \pm 150 \text{ Jy})$ , plus review of flux densities.

van Gorkom *et al.* 1982, MNRAS, 198, 757. WSRT HI absorption to nearby point source.

Seaquist & Gilmore 1982, AJ, 87, 378. VLA observations of nearby source.

Leahy 1987, ApJ, 322, 917. Einstein observations.

Leahy & Roger 1991, AJ, 101, 1033. DRAO at 408 MHz (3'.5×4'.8) and 1.4 GHz (1'.0×1'.4), including H<sub>I</sub> and discussion of distance.

Yamauchi & Koyama 1993, PASJ, 45, 545. Hard X-ray observations.

Leahy & Aschenbach 1995, A&A, 293, 853. ROSAT observations. Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search.

Leahy *et al.* 1998, A&A, 339, 601. 232 MHz (3'.8×5'.2), 151 MHz (4'.2×5'.8) and Effelsberg 100-m at 4.7 GHz (2'.5) for spectral index studies

Roger *et al.* 1999, A&AS, 137, 7. 22 MHz flux density ( $S = 1130\pm340$  Jy).

Reich *et al.* 2003, A&A, 408, 961. Effelsberg 100-m at 863 MHz  $(14.5:S=91\pm3 \text{ Jy})$ .

Fürst & Reich 2004, in MIM, p141. Effelsberg 100-m at 1.4 and 2.7 GHz (9'3), including polarisation.

Kaplan *et al.* 2006, ApJS, 163, 344. X-ray upper limit on compact sources.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim$  3': S =  $102\pm7$  Jy) and 1420 MHz ( $\sim$  1': S = 54.0 $\pm$ 2.9 Jy), including polarisation and review of flux densities.

Leahy & Tian 2007, A&A, 461, 1013. CGPS at 408 MHz  $(2.8 \times 3.9 : S = 117.8 \pm 5.3 \text{ Jy})$  and 1.4 GHz  $(58'' \times 80'' : S = 65.9 \pm 3.4 \text{ Jy})$ .

Gao *et al.* 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9'.5:  $S=34\pm3$  Jy), including polarisation and review of flux densities. Gosachinskii 2013, AstL, 39, 179. Hi observations of region. Araya *et al.* 2014, MNRAS, 444, 860. Fermi detection.

**G166.0+4.3** VRO 42.05.01

**RA:**  $05^h26^m30^s$  **1-GHz flux/Jy:** 7 **Size/arcmin:**  $55\times35$  **Dec:**  $+42^\circ56'$  **Spectral index:** 0.37 **Type:** S

Radio: Two arcs of strikingly different radii.

**Optical:** Nearly complete ring. **X-ray:** Predominantly in SW. **Distance:** H<sub>I</sub> indicates 4.5 kpc.

### References:

van den Bergh *et al.* 1973, ApJS, 26, 19. Optical observations. Willis 1973, A&A, 26, 237. NRAO 300-ft at 2.7 GHz ( $5': S = 5.2\pm1.0 \text{ Jy}$ ).

Lozinskaya 1979, AuJPh, 32, 113. Hα interferometry. Landecker *et al.* 1982, ApJ, 261, L41. DRAO at 1.4 GHz (1′.0×1′.4), plus review of flux densities.

Fesen *et al.* 1983, ApJS, 51, 337. Deep [OIII] imagery. Pineault *et al.* 1985, A&A, 151, 52. VLA at 1.4 GHz (16"×20") of part of remnant, and optical observations.

Fesen *et al.* 1985, ApJ, 292, 29. Optical spectra.

Braun & Strom 1986, A&AS, 63, 345. WSRT HI Observations. Pineault *et al.* 1987, ApJ, 315, 580. DRAO and VLA combined at 1.4 GHz (20").

Landecker et al. 1989, MNRAS, 237, 277. DRAO at 1.4 GHz  $(1'.0 \times 1'.4)$ , including HI.

Burrows & Guo 1994, ApJ, 421, L19. ROSAT images and spectra.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Guo & Burrows 1997, ApJ, 480, L51. ASCA observations. Lorimer *et al.* 1998, A&A, 331, 1002. Pulsar search. Leahy & Tian 2005, A&A, 440, 929. CGPS at 408 MHz (3'.4×5'.0)

and 1.4 GHz (1.0×1.4), for spectral index studies. see also: Tian & Leahy 2006, A&A, 451, 991. Erratum.

Kothes *et al.* 2006, A&A, 457, 1081. CGPS at 408 MHz ( $\sim 3'$ :  $S = 8.1 \pm 0.9$  Jy) and 1420 MHz ( $\sim 1'$ :  $S = 5.1 \pm 0.4$  Jy), including polarisation and review of flux densities.

Bocchino et al. 2009, A&A, 498, 139. XMM-Newton observations.

Gao et al. 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9'.5:  $S = 3.3 \pm 0.3$  Jy), including polarisation and review of flux densities

Araya 2013, MNRAS, 434, 2202. Fermi observations. Matsumura *et al.* 2017, PASJ, 69, 30. Suzaku observations.

# G178.2-4.2

**RA:**  $05^h25^m05^s$  **1-GHz flux/Jy:** 2 **Size/arcmin:**  $72\times62$  **Dec:**  $+28^{\circ}11'$  **Spectral index:** 0.5 **Type:** S

Radio: Faint shell, brighter in NE.

#### References:

Gao *et al.* 2011, A&A, 532, A144. Urumqi 25-m at 5 GHz (9.5:  $S=1.0\pm0.1$  Jy), plus other observations.

# G179.0 + 2.6

RA: 05<sup>h</sup>53<sup>m</sup>40<sup>s</sup> 1-GHz flux/Jy: 7 Size/arcmin: 70 Dec: +31°05′ Spectral index: 0.4 Type: S?

**Radio:** Thick shell, with background extragalactic sources near centre.

**Optical:** Nearly complete shell. **Point sources:** Pulsar near centre.

#### References:

Fürst & Reich 1986, A&A, 154, 303. Effelsberg 100-m at 1.4 (9'.4), 2.7 (4'.3) and 4.75 GHz (2'.4).

Fürst et al. 1989, A&A, 223, 66. Observations of central, extragalactic source.

Lorimer et al. 1998, A&A, 331, 1002. Pulsar search. Reich 2002, in NSPS, p1. Effelsberg 100-m at 2.7 GHz.

Gao *et al.* 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9'.5:  $S = 3.2 \pm 0.3$  Jy), including polarisation and review of flux densities.

Jeong *et al.* 2012, Ap&SS, 342, 389. CO observations of region. Pletsch *et al.* 2013, ApJ, 779, L11. Pulsar detection. Acero *et al.* 2016, ApJS, 224, 8. Fermi observations. How *et al.* 2018, MNRAS, 478, 1987. Optical observations.

G180.0-1.7

RA: 05<sup>h</sup>39<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: 65 Size/arcmin: 180 Dec: +27°50′ Spectral index: varies Type: S

Radio: Large faint shell, with spectral break.

Optical: Wispy ring. X-ray: Possible detection.

Point sources: Pulsar within boundary, with faint wind nebula.

**Distance:** Various observations suggest about 1.2 kpc.

### References:

van den Bergh *et al.* 1973, ApJS, 26, 19. Optical observations. Sofue *et al.* 1980, PASJ, 32, 1. Effelsberg 100-m at 5 GHz (2'.6) of parts.

Kundu *et al.* 1980, A&A, 92, 225. Effelsberg 100-m at 2.7 GHz (5!5:S=34.9±4 Jy) and 1.6 GHz (10':S=60.2±6 Jy).

Angerhofer & Kundu 1981, AJ, 86, 1003. Arecibo at 430 MHz  $(9': S=97\pm20 \text{ Jy})$ .

Fürst *et al.* 1982, A&A, 115, 428. Observations of compact radio sources near the remnant.

Fesen et al. 1985, ApJ, 292, 29. Optical spectra.

Fürst & Reich 1986, A&A, 163, 185. Effelsberg 100-m at 1.4, 2.7 and 4.7/5.0 GHz (9'.4.4'.3 and 2'.4/2'.6).

Sauvageot et al. 1990, A&A, 227, 183. EXOSAT possible detection.

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Anderson *et al.* 1996, ApJ, 468, L55. Pulsar detection.

Reich 2002, in NSPS, p1. Effelsberg 100-m at 2.7 GHz. Reich *et al.* 2003, A&A, 408, 961. Effelsberg 100-m at 863 MHz

 $(14.5: S = 77 \pm 10 \text{ Jy}).$ 

Romani & Ng 2003, ApJ, 585, L41. Chandra of pulsar.

Kramer et al. 2003, ApJ, 593, L31. Pulsar observations.Sallmen & Welsh 2004, A&A, 426, 555. Optical absorption towards background stars.

Drew et al. 2005, MNRAS, 362, 753. Ha imaging.

Ng et al. 2007, ApJ, 654, 487. Pulsar and wind nebula observations

Xiao et al. 2008, A&A, 482, 783. Urumqi 25-m at 4.8 GHz  $(9.5: S = 15.4\pm3.0 \text{ Jy})$  and Effelsberg 100-m at 2.6 GHz  $(4.4: S = 34.6\pm4.0 \text{ Jy})$ .

Jeong et al. 2012, Ap&SS, 342, 389. CO observations of region. Katsuta et al. 2012, ApJ, 752, 135. Fermi observations.

Dinçel et al. 2015, MNRAS, 448, 3196. Identification of OB runaway star near centre.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Boubert et al. 2017, A&A, 606, A14. Gaia search for runaway progenitor companion.

Chen et al. 2017, MNRAS, 472, 3924. Optical and IR observations, including distance.

Ren et al. 2018, RAA, 18, 111. Optical spectroscopy.

# G181.1+9.5

RA: 06<sup>h</sup>26<sup>m</sup>40<sup>s</sup> 1-GHz flux/Jy: ? Size/arcmin: 74 Dec: +32°30′ Spectral index: 0.45? Type: S

Radio: Faint shell.X-ray: Detected.

**Distance:** HI observations suggest 0.5–2.5 kpc.

References:

Kothes *et al.* 2017, A&A, 597, A116. DRAO at 1.4 GHz (50"×90") including H<sub>I</sub> and Effelsberg 100-m at 4850 MHz (2'.45) including polarisation, plus ROSAT survey observations.

# G182.4+4.3

**RA:** 06<sup>h</sup>08<sup>m</sup>10<sup>s</sup> **1-GHz flux/Jy:** 0.5 **Dec:** +29°00′ **Spectral index:** 0.4 **Size/arcmin:** 50 **Type:** S

Radio: Incomplete shell.

Optical: Brighter in S and NW.

X-ray: Diffuse emission.

### References

Kothes *et al.* 1998, A&A, 331, 661. Effelsberg 100-m at 1.4, 2.7, 4.9 and 10.5 GHz (9'.4: $S = 0.36 \pm 0.08$  Jy, 4'.4: $S = 0.25 \pm 0.04$  Jy, 2'.5:  $S = 0.20 \pm 0.02$  Jy and 1'.2:  $S = 0.15 \pm 0.03$  Jy), plus X-ray upper limit.

Reich 2002, in NSPS, p1. Effelsberg 100-m at 2.7 GHz and 4.9 GHz (3').

Sun *et al.* 2011, A&A, 536, A83. Urumqi 25-m at 5 GHz (9.5: S = 0.26±0.5 Jy) including polarisation and review of flux densities. Sezer *et al.* 2012, MNRAS, 427, 1168. Optical and XMM-Newton observations.

Jeong et al. 2012, Ap&SS, 342, 389. CO observations of region.

# G184.6-5.8 Crab Nebula, 3C144, SN1054

**RA:**  $05^h34^m31^s$  **1-GHz flux/Jy:** 900 **Size/arcmin:**  $7\times5$  **Dec:**  $+22^\circ01'$  **Spectral index:** 0.30 **Type:** F

This is the remnant of the SN of AD1054.

**Radio:** Filled-centre, central pulsar, with faint 'jet' (or tube) extending from the N edge; shows secular decline.

**Optical:** Strongly polarised filaments, diffuse synchrotron emission, with 'jet' faintly visible.

**X-ray:** Central 'torus' around the pulsar. **Point sources:** Pulsar powering the remnant.

**Distance:** Proper motions and radial velocities give 2 kpc.

#### References:

Velusamy 1984, Nature, 308, 251. VLA at 1.4 GHz, radio detection of 'jet'.

Velusamy 1985, MNRAS, 212, 359. VLA at 1.4 GHz (15").
Aller & Reynolds 1985, ApJ, 293, L73. Radio flux density decrease.

Fesen & Gull 1986, ApJ, 306, 259. Deep [OIII] imagery of 'jet'. Marcelin *et al.* 1990, A&A, 228, 471. Optical of 'jet'. Bietenholz & Kronberg 1990, ApJ, 357, L13. VLA at 1.4 GHz

Hester *et al.* 1990, ApJ, 357, 539. Optical and IR images. Hickson & van den Bergh 1990, ApJ, 365, 224. Optical polarisation

tion.
Bietenholz & Kronberg 1992, ApJ, 393, 206. VLA at 1.5 and 5 GHz (1".8) and 1.5 and 14 GHz (6".5) for spectral studies.

Fesen & Staker 1993, MNRAS, 263, 69. [OIII] imaging of 'jet', and proper motion studies.

Kassim *et al.* 1993, AJ, 106, 2218. VLA at 74 MHz (20"). Véron-Cetty & Woltjer 1993, A&A, 270, 370. Continuum and [OIII] photometry.

Wallace et al. 1994, A&A, 286, 565. HI of surroundings.
Frail et al. 1995, ApJ, 454, L129. VLA at 333 MHz (20") for limits on shell

Bietenholz *et al.* 1997, ApJ, 490, 291. Comparison of VLA observations at 74 MHz, 327 MHz, 1.5 GHz and 5 GHz for spectral index studies.

Blair *et al.* 1997, ApJS, 109, 473. HST imaging. Fesen *et al.* 1997, AJ, 113, 354. Limits on  $H\alpha$  halo.

Nugent 1998, PASP, 110, 831. Optical expansion.

Sankrit et al. 1998, ApJ, 504, 344. HST images.

Wallace et al. 1999, ApJS, 124, 181. DRAO at 1.4 GHz ( $1.0 \times 2.8$ ), plus Effelsberg 100-m, for HI studies.

Greiveldinger & Aschenbach 1999, ApJ, 510, 305. X-ray variability of torus.

Weisskopf *et al.* 2000, ApJ, 536, L81. Chandra observations.

Sollerman *et al.* 2000, ApJ, 537, 861. HST observations. Aharonian *et al.* 2000, ApJ, 539, 317. HESS observations.

Willingale et al. 2001, A&A, 365, L212. XMM-Newton observations.

Douvion et al. 2001, A&A, 373, 281. ISO observations.

Bietenholz *et al.* 2001, ApJ, 560, 254. Multi-epoch VLA observations, showing variations near the pulsar.

Bandiera et al. 2002, A&A, 386, 1044. 1.3 mm observations. Hester et al. 2002, ApJ, 577, L49. HST and Chandra multi-epoch observations.

Atkins et al. 2003, ApJ, 595, 803. γ-ray observations.

Green et al. 2004, MNRAS, 355, 1315. Sub-mm and ISO observations.

Mori *et al.* 2004, ApJ, 609, 186. Chandra observations. Čadež *et al.* 2004, ApJ, 609, 797. Optical observations.

Bietenholz *et al.* 2004, ApJ, 615, 794. VLA at 5 GHz (1".4) and HST multi-epoch observation for proper motion studies.

Melatos et al. 2005, ApJ, 633, 931. Multi-epoch near-IR observations central region.

Seward *et al.* 2006, ApJ, 636, 873. Chandra observations of scattering halo.

Temim *et al.* 2006, AJ, 132, 1610. Spitzer observations. see *also*: Temim *et al.* 2009, AJ, 137, 5155. Erratum.

Aharonian et al. 2006, A&A, 457, 899. HESS observations

Seward et al. 2006, ApJ, 652, 1277. Chandra observations. MacAlpine et al. 2007, AJ, 133, 81. Optical spectroscopy.

Kaplan et al. 2008, ApJ, 677, 1201. HST proper motion of pulsar.
 Rudie et al. 2008, MNRAS, 384, 1200. [OIII] observations of 'jet', for proper motion.

Hurley-Walker et al. 2009, MNRAS, 396, 365. Radio observations at 14 to 18 GHz.

Tziamtzis *et al.* 2009, A&A, 497, 167. Limits on Hα halo.

Carlebois et al. 2010, AJ, 139, 2083. Optical imaging spectroscopy.

Aumount *et al.* 2010, A&A, 514, A70. IRAM 30-m at 150 GHz  $(16\rlap.{''}7:S=244\pm24$  Jy) and GBT at 90 GHz  $(9\rlap.{''}3)$ .

Satterfield et al. 2012, AJ, 144, 27. Optical spectroscopy.

Temim *et al.* 2012, ApJ, 753, 72. Spitzer observations of dust. Gomez *et al.* 2012, ApJ, 760, 96. Herschel IR and sub-mm observations of dust.

Loh *et al.* 2012, MNRAS, 421, 789. IR observations of  $\rm H_2$ . Lundqvist & Tziamtzis 2012, MNRAS, 423, 1571. Optical limits on outer shell.

Loll *et al.* 2013, ApJ, 765, 152. HST optical line and continuum observations.

Barlow et al. 2013, Science, 342, 1343. Herschel far-IR spectroscopy.

Black & Fesen 2015, MNRAS, 447, 2540. [OIII] observations of 'jet'.

Owen & Barlow 2015, ApJ, 801, 141. Herschel and Spitzer observations.

Bietenholz & Nugent 2015, MNRAS, 454, 2416. Multi-epoch VLA and optical observations for expansion study.

Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 9 frequencies between 30 and 857 GHz.

Trotter et al. 2017, MNRAS, 469, 1299. Time evolution of radio

Dubner *et al.* 2017, ApJ, 840, 82. VLA, HST and Chandra observations, plus ALMA of central region, and UV observations.

Ritacco *et al.* 2018, A&A, 616, A35. Observations 150 GHz (18"), including polarisation.

G189.1+3.0 IC443, 3C157

**RA:** 06<sup>h</sup>17<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** 165 **Dec:** +22°34′ **Spectral index:** 0.36 **Size/arcmin:** 45 **Type:** C

**Radio:** Limb-brightened to NE, with faint extension to the E.

**Optical:** Brightest to the NE, with faint filaments outside the NE boundary.

X-ray: Shell, brightest to the NE, plus compact source with nebula.

**Point sources:** X-ray source and nebula in S. **Distance:** Optical observations imply 1.9 kpc.

#### References:

Fesen 1984, ApJ, 281, 658. Optical of filament to far NE. Erickson & Mahoney 1985, ApJ, 290, 596. TPT at 4 frequencies between 31 and 74 MHz (11'×13' to 4'.7×5'.4), plus review of flux densities.

Braun & Strom 1986, A&A, 164, 193. WSRT at 327 MHz  $(72'' \times 185'')$  and 1.4 GHz  $(17'' \times 43'')$ , plus HI and IRAS.

Green 1986, MNRAS, 221, 473. 151 MHz observations  $(1'2\times3'.1)$  and Half-Mile Telescope at 1.4 GHz  $(2'.1\times5'.4)$ .

Mufson *et al.* 1986, AJ, 92, 1349. Radio, IR, optical, UV and X-ray comparison, including VLA at 1.6 GHz (3"3×3"8 and 40"). Petre *et al.* 1988, ApJ, 335, 215. Einstein and other X-ray observations.

Dickel *et al.* 1989, AJ, 98, 1363. VLA at 1.4 GHz (1".1×1".2) of NE.

Wood *et al.* 1991, AJ, 102, 224. VLA at 5 GHz (3".6×3".8) of northeast, including polarisation.

Asaoka & Aschenbach 1994, A&A, 284, 573. X-ray, including possible overlapping remnant.

Tauber et al. 1994, ApJ, 421, 570. Observations of shocked molecular species.

Claussen et al. 1999, ApJ, 522, 349. High resolution observations of OH masers.

Rho et al. 2001, ApJ, 547, 885. ISO and 2Mass IR spectroscopy and imaging.

Olbert et al. 2001, ApJ, 554, L205. Chandra of compact X-ray source and surrounding nebula.

Reich *et al.* 2003, A&A, 408, 961. Effelsberg 100-m at 863 MHz  $(14'5: S=160\pm 5 \text{ Jy})$ .

Welsh & Sallmen 2003, A&A, 408, 545. Optical absorption studies

Leahy 2004, AJ, 127, 2277. DRAO at 408 MHz (3'.3×8'.6) and 1.4 GHz (1'.0×2'.6), for spectral index studies. *see also*: Leahy 2004, AJ, 128, 1478. Addendum.

Bykov *et al.* 2005, ApJ, 624, L41. Chandra observations of compact source.

Gaensler et al. 2006, ApJ, 648, 1037. Chandra of X-ray source and nebula.

Hewitt *et al.* 2006, ApJ, 652, 1288. GBT at 1.6 and 1.7 GHz (7.2) for OH, and VLA at 330 MHz (64"×74").

Rosado et al. 2007, AJ, 133, 89. Observations of shocked  $H_2$ . Albert et al. 2007, ApJ, 664, L87.  $\gamma$ -ray observations.

Albert et al. 2007, ApJ, 664, L87. γ-ray observations. Troja et al. 2008, A&A, 485, 777. XMM-Newton observations.

Casandjian & Grenier 2008, A&A, 489, 849. γ-ray observations. Bykov *et al.* 2008, ApJ, 676, 1050. XMM-Newton, Chandra and Spitzer observations.

Hewitt et al. 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers.

Lee et al. 2008, AJ, 135, 796. VLA at 1.4 GHz  $(39'' \times 42'')$  including HI.

Bocchino et al. 2008, AdSpR, 41, 396. INTEGRAL observations.

Bocchino et al. 2009, A&A, 498, 139. XMM-Newton observations.

Acciari *et al.* 2009, ApJ, 698, L133. γ-ray observations. Yamaguchi *et al.* 2009, ApJ, 705, L6. Suzaku observations. Koo *et al.* 2010, AJ, 140, 262. Ht Zeeman splitting observations. Gao *et al.* 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9.5:

 $S=85\pm9$  Jy), including polarisation and review of flux densities. Castelletti *et al.* 2011, A&A, 534, A21. VLA at 74 MHz (35":  $S=470\pm51$  Jy) and 330 MHz (17":  $S=248\pm15$  Jy), plus review of flux densities.

Yuan & Neufeld 2011, ApJ, 726, 76. Spitzer observations.
Xu et al. 2011, ApJ, 727, 81. CO observations of region.
Shinn et al. 2011, ApJ, 732, 124. Akari observations of H<sub>2</sub>.
Lee et al. 2012, ApJ, 749, 34. CO and HCO+ of region.
Taylor et al. 2012, ApJ, 750, L15. Optical absorption of background stars.

Yuan et al. 2012, ApJ, 753, 126. Spitzer spectroscopy. Ackermann et al. 2013, Science, 339, 807. Fermi observations. Hezareth et al. 2013, A&A, 558, A45. CO observations, include

Hezareth et al. 2013, A&A, 558, A45. CO observations, including polarisation, of region.

Kokusho *et al.* 2013, ApJ, 768, L8. IR observations of [FeII], plus Akari and Spitzer observations.

Pihlström *et al.* 2014, AJ, 147, 73. VLA search for methanol masers.

Ohnishi *et al.* 2014, ApJ, 784, 74. Suzaku observations. Yamaguchi *et al.* 2014, ApJ, 785, L27. Suzaku observations.

Yamaguchi et al. 2014, ApJ, 785, L27. Suzaku observations. Su et al. 2014, ApJ, 788, 122. CO observations.

Gusdorf et al. 2014, IAUS, 296, 178. CO observations. Mitra et al. 2014, IAUS, 296, 376. Combined GMRT and CLFST

observations at 150 MHz ( $19'' \times 24''$ ). Swartz *et al.* 2015, ApJ, 808, 84. Chandra observations of compact source and nebula.

Kokusho *et al.* 2015, P&SS, 116, 92. IR [FeII] observations. Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 9 frequencies between 30 and 857 GHz.

Kilpatrick et al. 2016, ApJ, 816, 1. CO observations, including broad lines.

Koo et al. 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

McEwen et al. 2016, ApJ, 826, 189. NH<sub>3</sub> and CH<sub>3</sub>OH observations.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Egron *et al.* 2017, MNRAS, 470, 1329. SRT at 1.5 GHz (11':  $S = 134 \pm 4$  Jy) and 7 GHz (2'.7:  $S = 67 \pm 3$  Jy).

Ambrocio-Cruz et al. 2017, MNRAS, 472, 51. Optical observations of NW, for distance.

Madsen et al. 2017, ApJ, 841, 56. NuSTAR observations.
Greco et al. 2018, A&A, 615, A157. XMM-Newton observations.
Zhang et al. 2018, ApJ, 859, 141. NuSTAR, XMM-Newton and Chandra observations.

# G190.9 - 2.2

**RA**: 06<sup>h</sup>01<sup>m</sup>55<sup>s</sup> 1-GHz flux/Jy: 1.3? Size/arcmin:  $70 \times 60$ **Dec:**  $+18^{\circ}24'$ **Spectral index:** 0.7? Type: S

Radio: Incomplete shell.

#### References:

Foster et al. 2013, A&A, 549, A107. Various radio survey observations.

G205.5+0.5Monoceros Nebula

**RA:** 06<sup>h</sup>39<sup>m</sup>00<sup>s</sup> Size/arcmin: 220 1-GHz flux/Jy: 140 **Dec:**  $+06^{\circ}30'$ Spectral index: 0.4 Type: S

Radio: In complex region, parts may be HII regions.

Optical: Large ring, near Rosette nebula.

**X-ray:** Possibly detected.

**Distance:** Association with molecular cloud suggests 1.6 to 2.0 kpc.

#### References:

Milne & Dickel 1974, AuJPh, 27, 549. Parkes 64-m at 2.7 GHz (9').

Velusamy & Kundu 1974, A&A, 32, 375. NRAO 300-ft at 2.7 GHz (5'), part only.

Dickel & DeNoyer 1975, AJ, 80, 437. Arecibo at 111 MHz (1':S=  $462\pm180 \text{ Jy}$ ) and  $S_{610 \text{ MHz}} = 245 \text{ Jy}$ .

Davies et al. 1978, A&AS, 31, 271. Deep optical plates.

Lozinskaya 1981, SvAL, 7, 17. Mean optical velocity.

Graham et al. 1982, A&A, 109, 145. Effelsberg 100-m at 2.7 GHz  $(4.4:S=97.6\pm12.5 \text{ Jy})$ , plus review of flux densities.

Fesen et al. 1985, ApJ, 292, 29. Optical spectra. Leahy et al. 1986, MNRAS, 220, 501. Einstein observations. Odegard 1986, ApJ, 301, 813. TPT at 20.6, 25.6 and 30.9 MHz (24', 19' and 16').

Esposito et al. 1996, ApJ, 461, 820. Possible associated γ-ray emis-

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search.

Oliver et al. 1996, A&A, 315, 578. CO observations of some of surroundings.

Jaffe et al. 1997, ApJ, 484, L129. γ-ray detection.

Aharonian et al. 2004, A&A, 417, 973. HESS limit.

Kaplan et al. 2006, ApJS, 163, 344. X-ray upper limit on compact

Casandjian & Grenier 2008, A&A, 489, 849. γ-ray observations. Xiao & Zhu 2012, A&A, 545, A86. Review of radio, H<sub>I</sub> and Hα observations.

Dirks & Meyer 2016, ApJ, 819, 45. Time variation of optical line absorption.

Katagiri et al. 2016, ApJ, 831, 106. Fermi observations.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Boubert et al. 2017, A&A, 606, A14. Gaia search for runaway progenitor companion.

Su et al. 2017, ApJ, 836, 211. CO observations. Zhao et al. 2018, ApJ, 855, 12. IR observations.

G206.9 + 2.3PKS 0646+06

**RA:** 06<sup>h</sup>48<sup>m</sup>40<sup>s</sup> 1-GHz flux/Jy: 6 Size/arcmin:  $60 \times 40$ **Dec:** +06°26′ Spectral index: 0.5 Type: S?

**Radio:** Diffuse source near the Monoceros Nebula.

Optical: Filaments detected. **X-ray:** Possibly detected.

### References:

Davies & Meaburn 1978, A&A, 69, 443. Optical observations. Nousek et al. 1981, ApJ, 248, 152. HEAO-1 X-ray limit. Graham et al. 1982, A&A, 109, 145. Effelsberg 100-m at 2.7 GHz  $(4.4:S=4.1\pm0.6 \text{ Jy})$ , plus review of flux densities.

Rosado 1982, RMxAA, 5, 127. Optical observations. Fesen et al. 1985, ApJ, 292, 29. Optical spectra. Leahy 1986, A&A, 156, 191. Einstein observations.

Odegard 1986, ApJ, 301, 813. TPT at 20.6, 25.6 and 30.9 MHz (24', 19' and 16').

Biggs & Lyne 1996, MNRAS, 282, 691. Pulsar search. Gao et al. 2011, A&A, 529, A159. Urumqi 25-m at 5 GHz (9'.5:  $S = 2.9 \pm 0.3$  Jy), including polarisation and review of flux densi-

Ambrocio-Cruz et al. 2014, RMxAA, 50, 323. [SII] spectroscopy. Su et al. 2017, ApJ, 836, 211. CO observations.

# G213.0-0.6

**RA:** 06<sup>h</sup>50<sup>m</sup>50<sup>s</sup> **1-GHz flux/Jy:** 21 Size/arcmin:  $160 \times 140$ ? **Dec:**  $-00^{\circ}30'$ Spectral index: 0.4 Type: S

Has also been called G213.3-0.4.

Radio: Large, faint shell. **Optical:** Filamentary shell.

**Point sources:** Central X-ray source.

**Distance:** Association with molecular cloud suggests about 1 kpc.

# References:

Reich *et al.* 2003, A&A, 408, 961. Effelsberg 100-m at 863 MHz (14'.5: *S* = 22.0±3.7 Jy).

Stupar & Parker 2012, MNRAS, 419, 1413. Ha, radio survey and

other observations.
Su *et al.* 2017, ApJ, 836, 211. CO observations.

G260.4 - 3.4Puppis A, MSH 08–44

**RA:**  $08^{\text{h}}22^{\text{m}}10^{\text{s}}$ 1-GHz flux/Jy: 130 Size/arcmin:  $60 \times 50$ **Dec:**  $-43^{\circ}00'$ Spectral index: 0.5

This remnant overlaps the Vela SNR (G263.9–3.3).

**Radio:** Angular shell, brightest to the E, poorly defined to the W.

Optical: Nebulosity and wisps. **X-ray:** Brightest to the E.

**Point sources:** Central possible pulsating X-ray source.

**Distance:** Association with H<sub>I</sub> implies 1.3 kpc.

#### References:

Green 1971, AuJPh, 24, 773. Molonglo at 408 MHz (3': S = 198±20 Jy).

Goudis & Meaburn 1978, A&A, 62, 283, Hα+[NII] optical image. Petre et al. 1982, ApJ, 258, 22. Einstein observations.

Milne et al. 1983, MNRAS, 204, 237. FIRST at 1415 MHz (50"), large scale emission missing.

Teske & Petre 1987, ApJ, 318, 370. Coronal optical line emission. Dubner & Arnal 1988, A&AS, 75, 363. HI and CO observations of surroundings.

Arendt et al. 1990, ApJ, 350, 266. MOST at 843 MHz (44"×65"), with large scale emission added, plus IR, optical and X-ray ob-

Dubner et al. 1991, AJ, 101, 1466. VLA at 327 MHz (43"×73") and 1.5 GHz (43"×77").

Arendt et al. 1991, ApJ, 368, 474. IR observations.

Milne et al. 1993, MNRAS, 261, 366. Parkes 64-m at 4.75 (4.5: S= 59 $\pm$ 5 Jy) and 8.4 GHz (3': $S=38\pm4$  Jy), plus polarisation.

Berthiaume et al. 1994, ApJ, 425, 132. X-ray spectroscopy. Sutherland & Dopita 1995, ApJ, 439, 365. Spectrophotometry. Reynoso et al. 1995, AJ, 110, 318. VLA at 1.4 GHz (90") including neutral hydrogen.

Blair et al. 1995, ApJ, 454, L35. Far UV spectroscopy. Petre et al. 1996, ApJ, 465, L43. ROSAT of central source. Bock et al. 1998, AJ, 116, 1886. MOST at 843 MHz (43"×60"). Pavlov et al. 1999, ApJ, 511, L45. Possible pulsation detection from central X-ray source.

Zavlin et al. 1999, ApJ, 525, 959. X-ray observations of central source.

Bocchino et al. 2000, A&A, 359, 316. Optical studies of selected filaments in N.

Woermann et al. 2000, MNRAS, 317, 421. OH observations. Gaensler et al. 2000, ApJ, 537, L35. Radio limit for nebula around possible pulsar.

Reynoso et al. 2003, MNRAS, 345, 671. ATCA at 1.4 GHz (90") for HI near central X-ray source.

Hwang et al. 2005, ApJ, 635, 355. Chandra observations of E edge.

Hui & Becker 2006, A&A, 454, 543. XMM-Newton and Chandra

Hui & Becker 2006, A&A, 457, L33. Chandra proper motion study of central source.

Castelletti et al. 2006, A&A, 459, 535. VLA at 1.4 GHz  $(16'' \times 34'' : S = 114 \pm 8 \text{ Jy})$  and 327 MHz  $(45'' \times 90'' : S =$ 263±20 Jy).

Winkler & Petre 2007, ApJ, 670, 635. Chandra proper motion study of central source.

Type: S

Paron et al. 2008, A&A, 480, 439. CO observations of E. Hwang et al. 2008, ApJ, 676, 378. Suzaku observations. Katsuda et al. 2008, ApJ, 678, 297. XMM-Newton observations. Mignani et al. 2009, A&A, 500, 1211. Optical limits for compact X-ray source.

Katsuda et al. 2010, ApJ, 714, 1725. Chandra and XMM-Newton observations.

Arendt et al. 2010, ApJ, 725, 585. Spitzer observations.

Becker et al. 2012, ApJ, 755, 141. Chandra proper motion study of central X-ray source.

Katsuda et al. 2012, ApJ, 756, 49. XMM-Newton spectroscopy. Hewitt et al. 2012, ApJ, 759, 89. Fermi and WMAP 23 to 90 GHz observations.

Dubner et al. 2013, A&A, 555, A9. XMM-Newton and Chandra observations.

Katsuda et al. 2013, ApJ, 768, 182. XMM-Newton observations. Abramowski et al. 2015, A&A, 575, A81. HESS limit.

Reynoso & Walsh 2015, MNRAS, 451, 3044. ATCA at 1.4 GHz  $(51'' \times 82'')$  and 1.7 GHz for spectral index study.

Arnaud et al. 2016, A&A, 586, A134. Planck flux densities at 4 frequencies between 30 and 100 GHz.

Luna et al. 2016, A&A, 590, A70. XMM-Newton and Chandra observations.

Koo et al. 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Reynoso et al. 2017, MNRAS, 464, 3029. ATCA at 1.4 GHz including HI.

García et al. 2017, A&A, 604, L5. XMM-Newton observations of

Slane et al. 2018, ApJ, 865, 86. XMM-Newton and Chandra observations of Vela X.

Reynoso et al. 2018, MNRAS, 477, 2087. ATCA at 1.3 to 2.6 GHz, including polarisation.

# G261.9 + 5.5

RA: 09h04m20s **1-GHz flux/Jy:** 10? Size/arcmin:  $40 \times 30$ **Dec:**  $-38^{\circ}42'$ **Spectral index:** 0.4? Type: S

**Radio:** Faint shell with little limb brightening.

### References:

Hill 1967, AuJPh, 20, 297. Parkes 64-m at 2650 MHz (7.5: S =7 Jy) also  $S_{1410 \text{ MHz}} = 8$  Jy,  $S_{81.5 \text{ MHz}} = 25$  Jy.

Colomb & Dubner 1980, A&A, 82, 244. Argentine 30-m dish at 1.4 GHz, for H<sub>I</sub> possibly associated with remnant.

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz  $(44'' \times 71'')$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

G263.9-3.3 Vela (XYZ)

RA: 08<sup>h</sup>34<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: 1750 Size/arcmin: 255 Dec: -45°50′ Spectral index: varies Type: C

This refers to the whole Vela XYZ complex, of which X has at times been classified as a separate (filled-centre) remnant. This remnant is overlapped by G260.4–3.4 and G266.2–1.2.

**Radio:** Large shell, with flatter spectrum component (Vela X), and pulsar nebula.

**Optical:** Filaments.

**X-ray:** Patchy shell, with extensions, central nebula and pulsar.

**Point sources:** Pulsar within Vela X, with one-sided 'jet'.

**Distance:** Vela pulsar parallax gives 0.3 kpc, optical spectra and H<sub>I</sub> studies suggest 0.25 kpc.

#### References:

Milne 1968, AuJPh, 21, 201. Parkes 64-m at various frequencies, including 408 MHz ( $48':S=2300\pm300$  Jy), 635 MHz ( $31':S=2360\pm300$  Jy), 1410 MHz ( $14':S=1640\pm300$  Jy) and 2650 MHz ( $7'.5:S=1400\pm250$  Jy), plus discussion of the distance.

Milne 1980, A&A, 81, 293. Maps of Vela X with Parkes 64-m at 1.66, 2.7 and 5 GHz (12', 8'.4 and 4'.4).

Weiler & Panagia 1980, A&A, 90, 269. Clarification of notation of this region and review previous observations.

Bignami & Caraveo 1988, ApJ, 325, L5. Pulsar proper motion from optical observations.

Dwarakanath 1991, JApA, 12, 199. Gauribidanur 'T' array at 34.5 MHz ( $26' \times 84' : S = 1800$  Jy for Vela X and S = 3900 Jy for Vela YZ), plus review of flux densities.

Bietenholz et al. 1991, ApJ, 376, L41. VLA at 5 GHz in vicinity of pulsar.

Willmore et al. 1992, MNRAS, 254, 139. Hard X-ray observations.

Dubner *et al.* 1992, A&AS, 96, 505. Argentine 30-m at 1.4 GHz (30') of surrounding Hi.

Strom *et al.* 1995, Nature, 373, 590. Radio of X-ray extensions. Markwardt & Ögelman 1995, Nature, 375, 40. X-ray jet from pulsar

Jenkins & Wallerstein 1995, ApJ, 440, 227. Optical absorption of associated neutral carbon cloud.

Milne 1995, MNRAS, 277, 1435. Parkes 64-m at 8.4 GHz (3'), including polarisation, of Vela X.

Danks & Sembach 1995, AJ, 109, 2627. Optical spectroscopy of background stars.

Blair et al. 1995, AJ, 110, 312. UV spectroscopy.

Duncan et al. 1996, MNRAS, 280, 252. Parkes 64-m at 2.4 GHz

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 62'')$  of part.

Green et al. 1997, AJ, 114, 2058. Parkes 64-m and ATCA OH observations.

Frail et al. 1997, ApJ, 475, 224. VLA at 330 MHz ( $1' \times 1'$ 8) and comparison with ROSAT observations of Vela X.

Markwardt & Ögelman 1997, ApJ, 480, L13. ASCA observations of pulsar 'iet'.

Raymond et al. 1997, ApJ, 482, 881. UV spectroscopy.

Yoshikoshi *et al.* 1997, ApJ, 487, L65. γ-ray observations of pulsar.

Jenkins et al. 1998, ApJ, 492, L147. UV absorption to background star.

Dubner *et al.* 1998, AJ, 116, 813. Parkes 64-m at 1.4 GHz (15') for HI studies.

Bock *et al.* 1998, AJ, 116, 1886. MOST at 843 MHz  $(43'' \times 60'')$ . Cha *et al.* 1999, ApJ, 515, L25. Optical spectra, for distance.

Lu & Aschenbach 2000, A&A, 362, 1083. ROSAT spatially resolved spectroscopy.

Chadwick *et al.* 2000, ApJ, 537, 414. Limit on high energy  $\gamma$ -rays from pulsar.

Sankrit *et al.* 2001, ApJ, 549, 416. Far-UV observations of selected region.

Pavlov et al. 2001, ApJ, 554, L189. Chandra two epoch observations of pulsar nebula.

Helfand et al. 2001, ApJ, 556, 380. Chandra observations of pulsar and its nebula.

Miyata *et al.* 2001, ApJ, 559, L45. Chandra of X-ray extension. Alvarez *et al.* 2001, A&A, 372, 636. Radio spectral index studies. *see also:* Alvarez *et al.* 2001, A&A, 379, 323. Erratum.

Caraveo et al. 2001, ApJ, 561, 930. HST parallax observations of nulsar

Moriguchi et al. 2001, PASJ, 53, 1025. CO observations.

Dodson *et al.* 2003, MNRAS, 343, 116. ATCA at 1.4, 2.4, 5.2 and 8.5 GHz (6".2×8".1, 26"×36", 10".5×12".1 and 10".6×11".2) of pulsar nebula.

Sankrit *et al.* 2003, ApJ, 589, 242. Optical nebulosity to NE of G266.2–1.2.

Nichols & Slavin 2004, ApJ, 610, 285. UV absorption toward background sources.

Hales et al. 2004, ApJ, 613, 977. Vela X at 31 GHz (4'.1).

Mongano *et al.* 2005, A&A, 436, 917. XMM-Newton and other X-ray observations of pulsar nebula.

Miceli *et al.* 2005, A&A, 442, 513. XMM-Newton observations of N rim.

Katsuda & Tsunemi 2005, PASJ, 57, 621. XMM-Newton observations of E.

Aharonian et al. 2006, A&A, 448, L43. HESS observations.

McConnell et al. 2006, AJ, 131, 648. ATCA at 4.9 GHz (12') including polarisation.

Katsuda & Tsunemi 2006, ApJ, 642, 917. XMM-Newton observations of NE.

Nishikida et al. 2006, ApJ, 644, L171. Far UV observations.

Miceli et al. 2008, ApJ, 676, 1064. XMM-Newton observations. LaMassa et al. 2008, ApJ, 689, L121. XMM-Newton observations of Vela X.

Grondin *et al.* 2013, ApJ, 774, 110. Fermi observations of Vela X. Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 30, 44 and 70 GHz.

Rao et al. 2016, MNRAS, 455, 2529. Time variation of optical line

Abdalla et al. 2018, A&A, 612, A1. HESS observations of Vela X.

**G266.2–1.2** RX J0852.0–4622

**RA:** 08<sup>h</sup>52<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** 50? **Size/arcmin:** 120 **Dec:** -46°20′ **Spectral index:** 0.3? **Type:** S

This remnant overlaps the Vela SNR (G263.9–3.3).

**Radio:** Incomplete shell, confused by the Vela SNR.

**Optical:** Nebulosity offset to NE.

X-ray: Non-thermal shell, confused by the Vela SNR, with central source, and possible associated pulsar.

**Point sources:** Central X-ray source, with optical nebula, and possible associated pulsar.

**Distance:** X-ray data suggest an upper limit of 1 kpc.

#### References:

Aschenbach 1998, Nature, 396, 141. ROSAT identification. Iyudin *et al.* 1998, Nature, 396, 142. γ-ray observations. Combi *et al.* 1999, ApJ, 519, L177. Radio observations. Duncan & Green 2000, A&A, 364, 732. Parkes 64-m at 1.4 GHz

(14/0), and comparison with other observations. Redman *et al.* 2000, ApJ, 543, L153. Optical of nearly nebulosity. Tsunemi *et al.* 2000, PASJ, 52, 887. ASCA spectral observations.

Slane et al. 2001, ApJ, 548, 814. ASCA observations.
Mereghetti et al. 2001, ApJ, 548, L213. BeppoSAX observations of central sources.

Pavlov et al. 2001, ApJ, 559, L131. Chandra of central X-ray source.

Moriguchi et al. 2001, PASJ, 53, 1025. CO observations.Pellizzoni et al. 2002, A&A, 393, L65. Optical observations of central source.

Redman et al. 2002, MNRAS, 336, 1093. Optical nebulosity to NE.

Kargaltsev et al. 2002, ApJ, 580, 1060. Chandra observations of central source.

Sankrit et al. 2003, ApJ, 589, 242. Optical nebulosity to NE. Redman & Meaburn 2005, MNRAS, 356, 969. Possible pulsar association. Iyudin *et al.* 2005, A&A, 429, 225. XMM-Newton observations. Aharonian *et al.* 2005, A&A, 437, L7. HESS observations. Katagiri *et al.* 2005, ApJ, 619, L163.  $\gamma$ -ray observations. Bamba *et al.* 2005, ApJ, 632, 294. Chandra of NW rim. Reynoso *et al.* 2006, A&A, 449, 243. ATCA at 1.38 GHz  $(32''\times37'')$ .

Enomoto *et al.* 2006, ApJ, 652, 1268. γ-ray observations. Mignani *et al.* 2007, A&A, 473, 883. Deep optical observations of compact X-ray source.

Aharonian et al. 2007, ApJ, 661, 236. HESS observations.
Katsuda et al. 2008, ApJ, 678, L35. XMM-Newton proper motion study.

Pannuti et al. 2010, ApJ, 721, 1492. Chandra observations of NW. Allen et al. 2015, ApJ, 798, 82. Two epoch Chandra observations for expansion.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.
Takeda et al. 2016, PASJ, 68, S10. Suzaku observations.
Fukui et al. 2017, ApJ, 850, 71. CO and HI observations.
Maxted et al. 2018, ApJ, 866, 76. ATCA at 1.4 to 2.9 GHz, plus other observations.

Abdalla et al. 2018, A&A, 612, A7. HESS observations.

# G272.2 - 3.2

**RA:** 09<sup>h</sup>06<sup>m</sup>50<sup>s</sup> **1-GHz flux/Jy:** 0.4 **Size/arcmin:** 15? **Dec:** -52°07′ **Spectral index:** 0.6 **Type:** S?

**Radio:** Diffuse shell. **Optical:** Detected.

X-ray: Centrally brightened.

### References:

Greiner et al. 1994, A&A, 286, L35. ROSAT observations, plus optical observations.

Duncan *et al.* 1997, MNRAS, 289, 97. Parkes 64-m at 1.4 GHz  $(18': S = 0.38 \pm 0.09 \text{ Jy})$  2.4 GHz  $(10'.6: S = 0.25 \pm 0.04 \text{ Jy})$  and 4.8 GHz  $(5'.7: S = 0.17 \pm 0.02 \text{ Jy})$ , MOST at 843 MHz  $(45'' \times 70'': S = 0.45 \pm 0.10 \text{ Jy})$ , and ATCA at 2.4 GHz  $(37'' \times 52'')$ , plus ROSAT observations.

Harrus et al. 2001, ApJ, 552, 614. ASCA and ROSAT observations, plus review of earlier observations.

Kamitsukasa et al. 2016, PASJ, 68, S7. Suzaku observations.

## G279.0+1.1

**RA:** 09<sup>h</sup>57<sup>m</sup>40<sup>s</sup> 1-GHz flux/Jy: 30? Size/arcmin: 95 **Dec:** -53°15′ **Spectral index:** 0.6? Type: S

Radio: Faint, incomplete shell.

Optical: Detected.

Point sources: Pulsar nearby.

#### References:

Woermann & Jonas 1988, MNRAS, 234, 971. Hartesbeesthoek 26m at 1.6 (30': $S = 25.2 \pm 4$  Jy) and 2.3 GHz (20': $S = 20.7 \pm 3$  Jy).

Duncan et al. 1995, MNRAS, 277, 319. Parkes 64-m at 1.4 (18':  $S = 28 \pm 3$  Jy) and 2.4 GHz (11':  $S = 20 \pm 2$  Jy), including polarisa-

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 53'')$  of part.

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations. Stupar & Parker 2011, MNRAS, 414, 2282. Ha observations.

G284.3-1.8 MSH 10-53

**RA:**  $10^{\rm h}18^{\rm m}15^{\rm s}$ Size/arcmin: 24? **1-GHz flux/Jy:** 11? **Dec:**  $-59^{\circ}00'$ **Spectral index:** 0.3? Type: S

Has been called G284.2–1.8.

**Radio:** Incomplete, poorly defined shell. **Optical:** Faint filament near edge.

**X-ray:** Diffuse emission, brighter in N and W.

**Point sources:** Central X- $/\gamma$ -ray source, not thought to be related.

#### References:

Ruiz & May 1986, ApJ, 309, 667. CO and optical observations. Milne et al. 1989, PASA, 8, 187. MOST at 843 MHz (43"×50") and Parkes 64-m at 8.4 GHz (3': $S=5.4\pm0.8$  Jy) including polarisation, plus earlier flux densities.

Green et al. 1997, AJ, 114, 2058. Parkes 64-m and ATCA OH observations.

Camilo et al. 2001, ApJ, 557, L51. Observations of nearby pulsar.

Camilo et al. 2004, ApJ, 616, 1118. Chandra observations of pulsar and nebula.

Abramowski et al. 2012, A&A, 541, A5. HESS observations. Williams et al. 2015, ApJ, 808, L19. Chandra and XMM-Newton observations.

Marcote et al. 2018, A&A, 619, A76. Observations of central source.

## G286.5-1.2

**RA:** 10<sup>h</sup>35<sup>m</sup>40<sup>s</sup> 1-GHz flux/Jy: 1.4? Size/arcmin:  $26 \times 6$ **Dec:** -59°42′ **Spectral index:** ? Type: S?

**Radio:** Double, elongated arc.

**Optical:** Detected.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 50'' : S = 1.6 \text{ Jy}).$ 

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations. Stupar & Parker 2011, MNRAS, 414, 2282.  $H\alpha$  observations.

## G289.7 - 0.3

**RA:** 11<sup>h</sup>01<sup>m</sup>15<sup>s</sup> Size/arcmin: 18×14 1-GHz flux/Jy: 6.2 **Dec:** -60°18′ **Spectral index:** 0.2? Type: S

**Radio:** Incomplete shell.

**Point sources:** Compact radio source near centre.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 50'' : S = 6.4 \pm 0.5 \text{ Jy})$ , plus Parkes 64-m at 4.5 GHz (S =  $7.5\pm2.5$  Jy) and 8.55 GHz ( $S=3.6\pm0.9$  Jy).

**G290.1–0.8** MSH 11–6*1*A

**RA:**  $11^h03^m05^s$  **1-GHz flux/Jy:** 42 **Size/arcmin:**  $19 \times 14$  **Dec:**  $-60^{\circ}56'$  **Spectral index:** 0.4 **Type:** S

Radio: Elongated, clumpy shell. Optical: Filaments detected. X-ray: Centrally brightened.

**Point sources:** Pulsar nearby, with PWN and extended 'jet' in X-rays.

**Distance:** HI absorption indicates  $7\pm1$  kpc.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Milne & Dickel 1975, AuJPh, 28, 209. Parkes 64-m at 5 GHz (4/4: S=20.2 Jy).

Elliott & Malin 1979, MNRAS, 186, 45P. Optical image and spectra.

Kirshner & Winkler 1979, ApJ, 227, 853. Optical observations. Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz  $(44'' \times 50'')$ .

Milne *et al.* 1989, PASA, 8, 187. MOST at 843 MHz  $(43'' \times 49'' : S = 45 \pm 11 \text{ Jy})$ , and Parkes 64-m at 8.4 GHz  $(3' : S = 19.5 \pm 1.0 \text{ Jy})$ , including polarisation.

Seward 1990, ApJS, 73, 781. Einstein observations.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×49": S=43 Jy).

Rosado *et al.* 1996, A&A, 315, 243. Optical kinematics. Kaspi *et al.* 1997, ApJ, 485, 820. Pulsar detection.

Gotthelf & Kaspi 1998, ApJ, 497, L29. ASCA observations of pulsar.

Slane et al. 2002, ApJ, 564, 284. ASCA observations.

Filipović *et al.* 2005, SerAJ, 170, 47. ATCA at 1.4 GHz (21"), plus other observations, including CO of surroundings.

Reynoso *et al.* 2006, MNRAS, 369, 416. ATCA at 1.4 GHz (22".5×25".0) including HI.

Pavan *et al.* 2014, A&A, 562, A122. Chandra and ATCA observations at 2 GHz (4".1×5".2) of PWN.

Halpern et al. 2014, ApJ, 795, L27. XMM-Newton observations of PWN.

Auchettl et al. 2015, ApJ, 810, 43. Fermi and Suzaku observations.

Kamitsukasa et al. 2015, PASJ, 67, 16. Suzaku observations.Pavan et al. 2016, A&A, 591, A91. Chandra observations, including PWN and 'jet'.

**G291.0–0.1** (MSH 11–62)

**RA:** 11<sup>h</sup>11<sup>m</sup>54<sup>s</sup> **1-GHz flux/Jy:** 16 **Size/arcmin:** 15×13 **Dec:** -60°38′ **Spectral index:** 0.29 **Type:** C

Radio: Centrally brightened core, with surrounding arcs.

Optical: Detected.

**X-ray:** Centrally brightened.

**Point sources:** Central compact X-ray source.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Roger *et al.* 1986, MNRAS, 219, 815. MOST at 843 MHz  $(43'' \times 50'' : S = 17.2 \pm 1.0 \text{ Jy})$ , and Parkes 64-m at 5 and 8.4 GHz  $(4.6 \text{ and } 3' : S = 10.4 \pm 0.4 \text{ Jy})$  and  $9.1 \pm 0.2$ , with polarisation.

Wilson 1986, ApJ, 302, 718. Einstein observations. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×49": S=12.7 Jy).

Harrus et al. 1998, ApJ, 499, 273. ASCA observations.
Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations.
Slane et al. 2012, ApJ, 749, 131. Chandra, XMM-Newton and Fermi observations.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

G292.0+1.8 MSH 11-54

**RA:**  $11^h24^m36^s$  **1-GHz flux/Jy:** 15 **Size/arcmin:**  $12\times8$  **Dec:**  $-59^{\circ}16'$  **Spectral index:** 0.4 **Type:** C

Radio: Centrally brightened source surrounded by a plateau of faint emission.

Optical: Oxygen rich.

X-ray: Ring of emission, with diffuse central nebula and pulsar.

**Point sources:** Central pulsar.

**Distance:** HI absorption implies 6.0 kpc.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Lockhart *et al.* 1977, MNRAS, 179, 147. Fleurs at 1415 MHz (50": S=13.0 Jy).

Goss et al. 1979, MNRAS, 188, 357. Optical spectra. Clark et al. 1980, MNRAS, 193, 129. X-ray spectrum. Tuohy et al. 1982, ApJ, 260, L65. Einstein observations. Dopita & Tuohy 1984, ApJ, 282, 135. Optical spectra. Braun et al. 1986, A&A, 162, 259. MOST at 843 MHz (0/8) and IRAS.

Hughes & Singh 1994, ApJ, 422, 126. EXOSAT spectrum.Sutherland & Dopita 1995, ApJ, 439, 365. Spectrophotometry.Hughes *et al.* 2001, ApJ, 559, L153. Chandra observations, including central nebula.

Park et al. 2002, ApJ, 564, L39. Chandra observations.
Camilo et al. 2002, ApJ, 567, L71. Pulsar detection.
Gonzalez & Safi-Harb 2003, ApJ, 583, L91. Chandra observations.
Hughes et al. 2003, ApJ, 591, L139. Chandra observations of pulsar.

Gaensler & Wallace 2003, ApJ, 594, 326. ATCA at 1.4, 2.3 and 5.2 GHz ( $8''.0 \times 9''.6, 6''.2 \times 7''.2$  and  $4''.8 \times 5''.5$ :  $S = 11.9 \pm 0.1, 11.4 \pm 0.1$  and  $8.8 \pm 0.1$ ), plus HI observations. Park *et al.* 2004, ApJ, 602, L33. Chandra observations.

Vink et al. 2004, NuPhS, 132, 62. XMM-Newton observations. Ghavamian et al. 2005, ApJ, 635, 365. Optical imaging spectroscopy.

Winkler & Long 2006, AJ, 132, 360. Optical observations.
Park et al. 2007, ApJ, 670, L121. Chandra observations.
Zharikov et al. 2008, A&A, 492, 805. Possible optical counterpart to pulsar and nebula.

Winkler et al. 2009, ApJ, 692, 1489. Optical proper motion studies.

Ghavamian et al. 2009, ApJ, 696, 1307. Spitzer spectroscopy. Lee et al. 2009, ApJ, 706, 441. IR observations. Lee et al. 2010, ApJ, 711, 861. Chandra observations. Ghavamian et al. 2012, ApJ, 750, 39. Spitzer observations. Zharikov et al. 2013, A&A, 554, A120. IR of pulsar wind nebula. Yamaguchi et al. 2014, ApJ, 785, L27. Suzaku observations. Kamitsukasa et al. 2014, PASJ, 66, 64. Suzaku observations. Bhalerao et al. 2015, ApJ, 800, 65. Chandra observations. Koo et al. 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Ghavamian & Williams 2016, ApJ, 831, 188. Spitzer observations. Acero *et al.* 2016, ApJS, 224, 8. Fermi observations.

## G292.2-0.5

**RA:**  $11^h 19^m 20^s$  **1-GHz flux/Jy:** 7 **Size/arcmin:**  $20 \times 15$  **Dec:**  $-61^{\circ}28'$  **Spectral index:** 0.5 **Type:** S

Radio: Shell.

**X-ray:** Shell, brighter to W, with central nebula.

**Point sources:** Central, young pulsar. **Distance:** HI absorption indicates 8.4 kpc.

## References:

Camilo *et al.* 2000, ApJ, 541, 367. Pulsar detection. Crawford *et al.* 2001, ApJ, 554, 152. ATCA at 1.4 GHz  $(25'' \times 29'' : S = 5.6 \pm 0.3 \text{ Jy})$  and 2.5 GHz  $(20'' \times 21'')$ .

Pivovaroff et al. 2001, ApJ, 554, 161. ROSAT and ASCA observations.

Gonzalez & Safi-Harb 2003, ApJ, 591, L143. Chandra observations of pulsar

Caswell *et al.* 2004, MNRAS, 352, 1405. ATCA at 5 GHz (1': S = 2.8 Jy), including polarisation, and 1.4 GHz for H<sub>I</sub> absorption.

Gonzalez & Safi-Harb 2005, ApJ, 619, 856. Chandra observations. Gonzalez *et al.* 2005, ApJ, 630, 489. XMM-Newton observations of pulsar.

Sañ-Harb et al. 2008, ApJ, 684, 532. Chandra observations of pulsar and nebula.

Kumar *et al.* 2012, ApJ, 754, 96. Chandra and XMM-Newton observations.

Ng et al. 2012, ApJ, 761, 65. XMM-Newton observations. Acero et al. 2016, ApJS, 224, 8. Fermi observations.

## G293.8+0.6

**RA:** 11<sup>h</sup>35<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** 5? **Size/arcmin:** 20 **Dec:** -60°54′ **Spectral index:** 0.6? **Type:** C

Radio: Central source, with faint extended plateau.

#### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 9.0 Jy) and Parkes 64-m at 5 GHz (4': S = 2.1 Jy).

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz (44"×51").

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×49": S=2.6 Jy).

G294.1 - 0.0

**RA:**  $11^h 36^m 10^s$  **1-GHz flux/Jy:** >2? **Size/arcmin:** 40 **Dec:**  $-61^\circ 38'$  **Spectral index:** ? **Type:** S

Radio: Faint shell.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 49'' : S > 2)$ .

G296.1 - 0.5

**RA:**  $11^h51^m10^s$  **1-GHz flux/Jy:** 8? **Size/arcmin:**  $37\times25$  **Dec:**  $-62^{\circ}34'$  **Spectral index:** 0.6? **Type:** S

Incorporates the previously catalogued remnant G296.1–0.7. Has been called G296.05–0.50.

Radio: Irregular shell, with nearby HII regions.

Optical: Detected.

**X-ray:** Irregular, incomplete shell.

### References

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 6.9 Jy) and Parkes 64-m at 5 GHz (4': S > 0.74). Longmore *et al.* 1977, MNRAS, 181, 541. Optical spectra. van den Bergh 1978, ApJS, 38, 119. Optical observations. Markert *et al.* 1981, ApJ, 248, L17. Einstein observations. Caswell & Barnes 1983, ApJ, 271, L55. Molonglo at 408 MHz (3': S = 12.4 Jy).

Bignami et al. 1986, ApJ, 302, 606. EXOSAT and Einstein observations

Hwang & Markert 1994, ApJ, 431, 819. ROSAT observations. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 48'' : S > 2.4)$ .

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Castro *et al.* 2011, ApJ, 734, 86. XMM-Newton observations. Gök & Sezer 2012, MNRAS, 419, 1603. Suzaku observations.

**G296.5+10.0** PKS 1209–51/52

**RA:**  $12^h09^m40^s$  **1-GHz flux/Jy:** 48 **Size/arcmin:**  $90\times65$  **Dec:**  $-52^{\circ}25'$  **Spectral index:** 0.5 **Type:** S

Has been called G296.5+9.7, and erroneously G295.5+9.7.

Radio: Shell with two bright limbs.

**Optical:** Detected.

**X-ray:** Incomplete shell, with central pulsar.

Point sources: Central pulsar.

#### References:

Irvine & Irvine 1974, ApJ, 192, L111. Optical observations. Danziger & Dennefeld 1976, PASP, 88, 44. Optical spectra. Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8/4) and 5 GHz (4/4).

Tuohy *et al.* 1979, ApJ, 230, L27. X-ray detection with HEAO-1 A2 experiment.

Ruiz 1983, AJ, 88, 1210. Optical spectra.

Dubner *et al.* 1986, AJ, 91, 343. Argentine 30-m dish at 1.4 GHz (34'), plus HI.

Kellett *et al.* 1987, MNRAS, 225, 199. EXOSAT of the west of the remnant, including the compact source.

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz  $(44'' \times 56'')$ .

Matsui et al. 1988, ApJ, 329, 838. Einstein observations, including compact source.

Roger *et al.* 1988, ApJ, 332, 940. MOST at 843 MHz (44"×56"). Bignami *et al.* 1992, ApJ, 389, L67. Optical in vicinity of X-ray source.

Milne & Haynes 1994, MNRAS, 270, 106. Parkes 64-m at 2.4 GHz (8'.3:  $S = 33\pm3$  Jy), 4.8 GHz (4'.5:  $S = 23.3\pm3$  Jy) and 8.4 GHz (3'.0: 18.8 $\pm3$ ), including polarisation and review of flux densities

Mereghetti et al. 1996, ApJ, 464, 842. Radio, optical and X-ray observations of central source.

Green et al. 1997, AJ, 114, 2058. Parkes 64-m and ATCA OH observations.

Vasisht et al. 1997, ApJ, 476, L43. ASCA observations of neutron star.

Zavlin *et al.* 1998, A&A, 331, 821. ROSAT and ASCA observations of neutron star.

Giacani *et al.* 2000, AJ, 119, 281. ATCA at 1.4 GHz (2'.7×4'.0) for HI studies.

Zavlin et al. 2000, ApJ, 540, L25. Chandra observations of central pulsar.

Gotthelf & Halpern 2007, ApJ, 664, L35. X-ray timing observations of pulsar.

Harvey-Smith *et al.* 2010, ApJ, 712, 1157. ATCA at 1.4 GHz (1'.8×3'.3), including polarisation.

Araya 2013, MNRAS, 434, 2202. Fermi observations.

Halpern & Gotthelf 2015, ApJ, 812, 61. Two epoch Chandra observations for pulsar proper motion.

Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 4 frequencies between 30 and 100 GHz.

Acero *et al.* 2016, ApJS, 224, 8. Fermi observations. Ackermann *et al.* 2018, ApJS, 237, 32. Fermi observations.

G296.7 - 0.9

**RA:**  $11^h55^m30^s$  **1-GHz flux/Jy:** 3 **Size/arcmin:**  $15\times8$  **Dec:**  $-63^\circ08'$  **Spectral index:** 0.5 **Type:** S

**Radio:** Bilateral shell. **X-ray:** Brighter to SE.

### References:

Schaudel *et al.* 2002, ASPC, 271, 391. ROSAT observations. Robbins *et al.* 2012, MNRAS, 419, 2623. ATCA at 1.4 GHz  $(33'' \times 45'' : S = 2.5 \pm 0.2 \text{ Jy})$ , plus MOST at 843 MHz  $(43'' \times 49'')$ , plus other observations.

Prinz & Becker 2013, A&A, 550, A33. XMM-Newton observations.

Green *et al.* 2014, PASA, 31, 42. Radio observations at 843 MHz  $(45'' \times 50'')$ .

G296.8 - 0.31156–62

Size/arcmin: 20×14 **RA:** 11<sup>h</sup>58<sup>m</sup>30<sup>s</sup> 1-GHz flux/Jy: 9 **Dec:**  $-62^{\circ}35'$ Spectral index: 0.6 Type: S

Radio: Shell, brighter to the NW.

X-ray: Detected.

**Distance:** Hi absorption gives 9.6 kpc.

References:

Caswell et al. 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3':S =15.0 Jy) and Parkes 64-m at 5 GHz (4':S=3.2 Jy). Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8'.4) and 5 GHz (4'.4). Hwang & Markert 1994, ApJ, 431, 819. ROSAT observations. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz

 $(43'' \times 48'' : S = 9.2 \text{ Jy}).$ 

Gaensler et al. 1998, MNRAS, 296, 813. ATCA at 1.3 GHz  $(22'' \times 24'' : S = 7.0 \pm 0.3 \text{ Jy})$ , including polarisation and H<sub>I</sub> observations, plus review of flux densities.

Sánchez-Ayaso et al. 2012, Ap&SS, 337, 573. XMM-Newton and IR observations.

G298.5 - 0.3

**RA:** 12<sup>h</sup>12<sup>m</sup>40<sup>s</sup> 1-GHz flux/Jy: 5? Size/arcmin: 5? **Dec:**  $-62^{\circ}52'$ **Spectral index:** 0.4? **Type:** ?

**Radio:** Not well resolved, may be part of a larger ring?

References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Hwang & Markert 1994, ApJ, 431, 819. ROSAT upper limit.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 48'' : S = 1.8 \text{ Jy}).$ 

Reach et al. 2006, AJ, 131, 1479. Spitzer possible detection.

G298.6 - 0.0

**RA:** 12<sup>h</sup>13<sup>m</sup>41<sup>s</sup> 1-GHz flux/Jy: 5? Size/arcmin:  $12 \times 9$ **Dec:** -62°37′ Spectral index: 0.3 Type: S

Has been called G298.6-0.1.

**Radio:** Incomplete shell, in complex region.

X-ray: Centrally brightened.

References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz  $(44'' \times 50'')$ 

Hwang & Markert 1994, ApJ, 431, 819. ROSAT upper limit.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 48'' : S = 7.4 \text{ Jy}).$ 

Reach et al. 2006, AJ, 131, 1479. Spitzer possible detection. Acero et al. 2016, ApJS, 224, 8. Fermi observations. Bamba et al. 2016, PASJ, 68, S5. Suzaku observations.

G299.2-2.9

**RA:** 12<sup>h</sup>15<sup>m</sup>13<sup>s</sup> 1-GHz flux/Jy: 0.5? Size/arcmin: 18×11 **Dec:**  $-65^{\circ}30'$ **Spectral index:** ? Type: S

Radio: Faint source. Optical: Filaments in W.

**X-ray:** Centrally brightened with shell at higher energies.

References:

Busser et al. 1996, A&A, 310, L1. ROSAT detection, plus optical

Slane et al. 1996, ApJ, 465, 840. Einstein, IRAS and radio observations.

Bai & Wang 2000, ApJ, 539, 760. ASCA observations. Park et al. 2007, ApJ, 665, 1173. Chandra observations. Post et al. 2014, ApJ, 792, L20. Chandra observations.

## G299.6 - 0.5

RA: 12<sup>h</sup>21<sup>m</sup>45<sup>s</sup> 1-GHz flux/Jy: 1.0? Size/arcmin: 13 Dec: -63°09′ Spectral index: ? Type: S

Radio: Faint shell, brightest to E.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 48'' : S = 1.1 \text{ Jy})$ .

## G301.4 - 1.0

**RA:**  $12^{h}37^{m}55^{s}$  **1-GHz** flux/Jy: 2.1? **Size/arcmin:**  $37 \times 23$  **Dec:**  $-63^{\circ}49'$  **Spectral index:** ? **Type:** S

**Radio:** Faint, incomplete shell, with possible extension to southwest.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 48'' : S = 2.3 \text{ Jy})$ .

## G302.3+0.7

**RA:** 12<sup>h</sup>45<sup>m</sup>55<sup>s</sup> **1-GHz flux/Jy:** 5? **Size/arcmin:** 17 **Dec:** -62°08′ **Spectral index:** 0.4? **Type:** S

**Radio:** Distorted shell, in complex region, with possibly associated filament.

### References

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 7.5 Jy) and Parkes 64-m at 5 GHz (4': S = 3.0 Jy). Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz

 $(43'' \times 49'' : S = 3.2 \text{ Jy}).$ 

Frail et al. 1996, AJ, 111, 1651. OH emission near remnant.

**G304.6+0.1** Kes 17

**RA:** 13<sup>h</sup>05<sup>m</sup>59<sup>s</sup> **1-GHz flux/Jy:** 14 **Size/arcmin:** 8 **Dec:** -62°42′ **Spectral index:** 0.5 **Type:** S

Radio: Incomplete shell.

X-ray: Detected.

**Distance:** Possible limit of > 9.7 kpc from H<sub>I</sub> absorption.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Milne & Dickel 1975, AuJPh, 28, 209. Parkes 64-m at 5 GHz (4.4: S=6.9 Jy).

Caswell *et al.* 1975, A&A, 45, 239. Parkes HI absorption. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×48": S=18 Jy).

Frail *et al.* 1996, AJ, 111, 1651. OH emission near remnant. Reach *et al.* 2006, AJ, 131, 1479. Spitzer observations. Hewitt *et al.* 2009, ApJ, 694, 1266. Spitzer spectroscopy. Combi *et al.* 2010, A&A, 523, A76. XMM-Newton observations Lee *et al.* 2011, ApJ, 740, 31. Akari and Spitzer observations.

Wu et al. 2011, ApJ, 740, L12. Fermi observations. Andersen et al. 2011, ApJ, 742, 7. Spitzer observations. Gök & Sezer 2012, MNRAS, 423, 1215. Suzaku observations. Gelfand et al. 2013, ApJ, 777, 148. ATCA at 1.4 GHz ( $8'' \times 23''$ :  $S = 10.9 \pm 0.4$  Jy), plus other observations. Pannuti et al. 2014, AJ, 147, 55. ASCA and XMM-Newton observations. Koo et al. 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties). Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Washino et al. 2016, PASJ, 68, S4. Suzaku observations.

## G306.3 - 0.9

**RA:** 13<sup>h</sup>21<sup>m</sup>50<sup>s</sup> **1-GHz flux/Jy:** 0.16? **Size/arcmin:** 4 **Dec:** -63°34′ **Spectral index:** 0.5? **Type:** S?

Radio: Diffuse emission.X-ray: Partial shell.

#### References:

Reynolds et al. 2013, ApJ, 766, 112. Chandra observations, and ATCA at 5.5 GHz  $(23'' \times 26'')$ , and 9 GHz.

Combi et al. 2016, A&A, 592, A125. XMM-Newton and Chandra observations. Sezar et al. 2017, MNRAS, 466, 3434. Suzaku and Fermi observa-

## G308.1 - 0.7

RA: 13<sup>h</sup>37<sup>m</sup>37<sup>s</sup>
1-GHz flux/Jy: 1.2?
Spectral index: ?

Size/arcmin: 13
Type: S

Radio: Faint shell.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×48": S=1.3 Jy).

## G308.4-1.4

**RA:**  $13^{h}41^{m}30^{s}$  **1-GHz flux/Jy:** 0.4? **Size/arcmin:**  $12\times6$ ? **Oec:**  $-63^{\circ}44'$  **Spectral index:** ? **Type:** S?

W part has been called G308.3-1.4.

**Radio:** Complex structure, with multiple arcs. **X-ray:** Limb brightened partial shell in W.

### References:

Schaudel *et al.* 2002, ASPC, 271, 391. ROSAT observations. Prinz & Becker 2012, A&A, 544, A7. ATCA at 1.4 GHz  $(53'' \times 64'' : S = 0.33 \text{ Jy})$  and 2.5 GHz  $(29'' \times 35'' : S = 0.24 \text{ Jy})$ , plus Chandra and other observations.

Hui et al. 2012, ApJ, 750, 7. XMM-Newton and other observa-

De Horta *et al.* 2013, MNRAS, 428, 1980. ATCA at 1.4 GHz  $(54'' \times 65'')$  and 2.5 GHz  $(29'' \times 35'')$ , plus other observations. Green *et al.* 2014, PASA, 31, 42. Radio observations at 843 MHz  $(45'' \times 50'')$ .

## G308.8 - 0.1

**RA:**  $13^{\text{h}}42^{\text{m}}30^{\text{s}}$  **1-GHz flux/Jy:** 15? **Size/arcmin:**  $30 \times 20$ ? **Dec:**  $-62^{\circ}23'$  **Spectral index:** 0.4? **Type:** C?

Incorporates previous catalogued remnant G308.7+0.0.

**Radio:** Bright ridge in north, and arc to south. **Point sources:** Pulsar near centre of remnant.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 16.7 Jy) and Parkes 64-m at 5 GHz (4': S = 7.0 Jy). Milne & Dickel 1975, AuJPh, 28, 209. Parkes 64-m at 5 GHz

(4.4:S=6.5 Jy).

Caswell *et al.* 1981, MNRAS, 195, 89. FIRST at 1415 MHz (50"). Wilson 1986, ApJ, 302, 718. Einstein upper limit.

Caswell *et al.* 1992, ApJ, 399, L151. MOST at 843 MHz (43"×49").

Kaspi *et al.* 1992, ApJ, 399, L155. Pulsar observations. Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Reach *et al.* 2006, AJ, 131, 1479. Spitzer possible detection.

## G309.2 - 0.6

**RA:**  $13^{\text{h}}46^{\text{m}}31^{\text{s}}$  **1-GHz flux/Jy:** 7? **Size/arcmin:**  $15 \times 12$  **Dec:**  $-62^{\circ}54'$  **Spectral index:** 0.4? **Type:** S

Has been called G309.2-0.7.

**Radio:** Distorted shell.

**X-ray:** Extended emission, with unrelated central source.

#### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 10.0 Jy) and Parkes 64-m at 5 GHz (4': S = 3.9 Jy).

Caswell *et al.* 1981, MNRAS, 195, 89. FIRST at 1415 MHz (1.5). Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz (44"×50").

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 48'' : S = 6 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

Gaensler *et al.* 1998, MNRAS, 299, 812. ATCA at 1.3 GHz  $(23'' \times 24'' : S = 5.2 \pm 0.2 \text{ Jy})$ .

Rakowski et al. 2001, ApJ, 548, 258. ASCA and ROSAT observa-

Rakowski et al. 2006, ApJ, 649, L111. Observations of central

Safi-Harb *et al.* 2007, ApJ, 659, 407. Observations of central source.

## G309.8 + 0.0

**RA:**  $13^h50^m30^s$  **1-GHz flux/Jy:** 17 **Size/arcmin:**  $25\times19$  **Dec:**  $-62^\circ05'$  **Spectral index:** 0.5 **Type:** S

Radio: Distorted shell.

**Point sources:** Steep radio spectrum source near the centre of the remnant.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 26.4 Jy) and Parkes 64-m at 5 GHz (4': S = 7.4 Jy).

Caswell et al. 1980, MNRAS, 190, 881. FIRST at 1415 MHz (1').

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×49": S > 8.8).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

## G310.6-1.6

**RA:**  $14^h00^m45^s$  **1-GHz flux/Jy:** ? **Size/arcmin:** 2.5 **Dec:**  $-63^\circ26'$  **Spectral index:** ? **Type:** C?

Radio: Not detected.

**X-ray:** Bright central nebula, with faint shell. **Point sources:** X-ray pulsar, and extended emission.

References:

Tomsick *et al.* 2009, ApJ, 701, 811. Chandra observations. Renaud *et al.* 2010, ApJ, 716, 663. Chandra and other observations

**G310.6–0.3** Kes 20B

RA: 13<sup>h</sup>58<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: 5? Size/arcmin: 8 Dec: -62°09′ Spectral index: ? Type: S

Has been called G310.65–0.29.

Radio: Asymmetric shell.

## References:

Whiteoak *et al.* 1994, MNRAS, 269, 294. MOST at 843 MHz. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 49'' : S = 5.4 \text{ Jy})$ .

Reach et al. 2006, AJ, 131, 1479. Spitzer possible detection.

G310.8-0.4 Kes 20A

**RA:**  $14^h00^m00^s$  **1-GHz flux/Jy:** 6? **Size/arcmin:** 12 **Dec:**  $-62^\circ17'$  **Spectral index:** ? **Type:** S

Has been called G310.80–0.41. **Radio:** Arc in E, in complex region.

#### References:

Whiteoak *et al.* 1994, MNRAS, 269, 294. MOST at 843 MHz. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 49'' : S = 6.9 \text{ Jy})$ .

Reach et al. 2006, AJ, 131, 1479. Spitzer observations. Andersen et al. 2011, ApJ, 742, 7. Spitzer observations.

## G311.5-0.3

RA: 14<sup>h</sup>05<sup>m</sup>38<sup>s</sup> 1-GHz flux/Jy: 3? Size/arcmin: 5 Dec: -61°58′ Spectral index: 0.5 Type: S

Has been called G311.52–0.37. **Radio:** Shell, not well resolved.

X-ray: Detected.

#### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3 $^{\prime}$ ) and Parkes 64-m at 5 GHz (4 $^{\prime}$ ).

Caswell & Barnes 1985, MNRAS, 216, 753. Molonglo at 408 MHz (3').

Whiteoak et al. 1994, MNRAS, 269, 294. MOST at 843 MHz.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 49'' : S = 2.9 \text{ Jy})$ .

Reach *et al.* 2006, AJ, 131, 1479. Spitzer observations. Andersen *et al.* 2011, ApJ, 742, 7. Spitzer observations. Pannuti *et al.* 2014, AJ, 147, 55. ASCA observations.

## G312.4 - 0.4

RA: 14<sup>h</sup>13<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: 45 Size/arcmin: 38 Dec: -61°44′ Spectral index: 0.36 Type: S

Has been called G312.44–0.36. **Radio:** Irregular, incomplete shell. **X-ray:** Weak emission in W.

**Point sources:** Nearby  $\gamma$ -ray sources and pulsars.

**Distance:** HI absorption suggests > 6 kpc and possibly > 14 kpc.

## References:

Caswell & Barnes 1985, MNRAS, 216, 753. Molonglo at 408 MHz (3':S=56 Jy).

Whiteoak *et al.* 1994, MNRAS, 269, 294. MOST at 843 MHz. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 49'' : S > 19)$ , plus Parkes 64-m at 4.5 GHz  $(S = 30 \pm 2 \text{ Jy})$  and 8.55 GHz  $(S = 17 \pm 4 \text{ Jy})$ .

Frail *et al.* 1996, AJ, 111, 1651. OH emission near remnant. Roberts *et al.* 1999, ApJ, 515, 712. MOST at 843 MHz (43"×49").

Case & Bhattacharya 1999, ApJ, 521, 246. Nearby γ-ray sources. Doherty *et al.* 2003, MNRAS, 339, 1048. ATCA at 1.4 GHz (25") plus H<sub>I</sub> absorption, and Chandra observations.

## G312.5 - 3.0

**RA:**  $14^{\text{h}}21^{\text{m}}00^{\text{s}}$  **1-GHz flux/Jy:** 3.5? **Size/arcmin:**  $20 \times 18$  **Dec:**  $-64^{\circ}12'$  **Spectral index:** ? **Type:** S

**Radio:** Distorted shell.

### References:

Duncan et al. 1995, MNRAS, 277, 36. Parkes 64-m at 2.4-GHz

Kane & Vaughan 2003, MNRAS, 344, 625. ATCA at 1.4 GHz  $(116'' \times 129'')$  and 2.4 GHz  $(67'' \times 75'')$ .

## G315.1+2.7

**RA:**  $14^{\text{h}}24^{\text{m}}30^{\text{s}}$  **1-GHz flux/Jy:** ? **Size/arcmin:**  $190 \times 150$  **Dec:**  $-57^{\circ}50'$  **Spectral index:** ? **Type:** S

Radio: Poorly defined shell.

Optical: Filaments, brighter in NE.

#### References

Duncan et al. 1997, MNRAS, 287, 722. Parkes 64-m at 2.4 GHz

 $(11':S=19\pm3 \text{ Jy}).$ 

Combi et al. 1998, A&A, 333, 298. Radio survey observations.

Stupar et al. 2007, MNRAS, 374, 1441. Optical and various radio

Stupar & Parker 2011, MNRAS, 414, 2282. Ha observations.

G315.4-2.3 RCW 86, MSH 14-63

**RA:** 14<sup>h</sup>43<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** 49 **Size/arcmin:** 42 **Dec:** -62°30′ **Spectral index:** 0.6 **Type:** S

Possibly the remnant of the SN of AD185?

**Radio:** Shell, brightest to the SW.

**Optical:** Bright, radiative filaments, with some faint Balmer dominated filaments.

**X-ray:** Partial shell, with thermal and non-thermal emission.

**Point sources:** Several X-ray sources.

**Distance:** Optical observations imply 2.3 kpc.

### References:

van den Bergh *et al.* 1973, ApJS, 26, 19. Optical observations. Caswell *et al.* 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3': S = 86 Jy) and Parkes 64-m at 5 GHz (4': S=18.2 Jy).

Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8'4) and 5 GHz (4'4).

Winkler 1978, ApJ, 221, 220. X-ray detection.

Leibowitz & Danziger 1983, MNRAS, 204, 273. Optical spectra. Pisarski *et al.* 1984, ApJ, 277, 710. Einstein observations

Nugent et al. 1984, ApJ, 284, 612. X-ray spectrum.

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz  $(44'' \times 50'')$ .

Long & Blair 1990, ApJ, 358, L13. Balmer dominated optical filaments.

Greidanus & Strom 1990, A&A, 240, 385. IRAS observations. Kaastra *et al.* 1992, A&A, 264, 654. Ginga X-ray spectra.

Strom 1994, MNRAS, 268, L5. Historical association

Chin & Huang 1994, Nature, 371, 398. Questioning of historical association.

Schaefer 1995, AJ, 110, 1793. Questioning of historical association.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×48": S > 22).

Rosado et al. 1996, A&A, 315, 243. Optical kinematics.

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Smith *et al.* 1997, AJ, 114, 2664. Observations of Balmer dominated filaments.

Vink et al. 1997, A&A, 328, 628. ASCA spectroscopy.

Bocchino et al. 2000, A&A, 360, 671. BeppoSAX observations of N and SW.

Bamba et al. 2000, PASJ, 52, 1157. ASCA observations.

Dickel *et al.* 2001, ApJ, 546, 447. ATCA at 1.34 GHz (8"), including polarisation.

Ghavamian *et al.* 2001, ApJ, 547, 995. Optical spectroscopy. Borkowski *et al.* 2001, ApJ, 550, 334. ASCA observations. Rho *et al.* 2002, ApJ, 581, 1116. Chandra observations.

Gvaramadze & Vikhlinin 2003, A&A, 401, 625. Chandra point source search.

Sollerman *et al.* 2003, A&A, 407, 249. Optical spectroscopy. Kaplan *et al.* 2004, ApJS, 153, 269. Chandra limits for any compact source.

Bamba et al. 2005, ApJ, 621, 793. Chandra observations of rim. Vink et al. 2006, ApJ, 648, L33. Chandra and XMM-Newton observations.

Ueno et al. 2007, PASJ, 59, S171. Suzaku observations.
Aharonian et al. 2009, ApJ, 692, 1500. HESS observations.
Yamaguchi et al. 2011, PASJ, 63, S837. Suzaku observations.
Helder et al. 2011, ApJ, 737, 85. XMM-Newton and optical observations.

Williams *et al.* 2011, ApJ, 741, 96. Spitzer observations. Mignani *et al.* 2012, MNRAS, 425, 2309. X-ray and optical observations of compact X-ray sources.

Lemoine-Goumard et al. 2012, A&A, 545, A28. Fermi observations.

Helder et al. 2013, MNRAS, 435, 910. Optical proper motion studies.

Castro et al. 2013, ApJ, 779, 49. Chandra of NW.

Yuan et al. 2014, ApJ, 785, L22. Fermi observations.

Tsubone et al. 2014, in EFXU, p72. Suzaku observations.

Broersen et al. 2014, MNRAS, 441, 3040. XMM-Newton observations.

Arnaud et al. 2016, A&A, 586, A134. Planck flux densities at 30, 44 and 70 GHz.

Ajello et al. 2016, ApJ, 819, 98. Fermi observations, and ATCA HI of surroundings.

Yamaguchi *et al.* 2016, ApJ, 820, L3. Chandra observations of NE. Koo *et al.* 2016, ApJ, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Tsubone et al. 2017, ApJ, 835, 34. Suzaku observations.

Sano et al. 2017, JHEAp, 15, 1. CO and HI observations.

Abramowski et al. 2018, A&A, 612, A4. HESS observations.

## G315.4 - 0.3

**RA:**  $14^{\text{h}}35^{\text{m}}55^{\text{s}}$  **1-GHz flux/Jy:** 8 **Size/arcmin:**  $24 \times 13$  **Dec:**  $-60^{\circ}36'$  **Spectral index:** 0.4 **Type:** ?

Radio: Irregular non-thermal emission, with HII region superposed in E.

Optical: Detected.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 15.9 Jy) and Parkes 64-m at 5 GHz (4': S = 4.9 Jy).

Caswell *et al.* 1981, MNRAS, 195, 89. FIRST at 1415 MHz (50'': S = 6.25 Jy), re-assessment of earlier flux densities.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×49": S=3.1 Jy).

Frail *et al.* 1996, AJ, 111, 1651. OH emission near remnant. Reach *et al.* 2006, AJ, 131, 1479. Spitzer possible detection. Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations.

## G315.9 - 0.0

**RA:**  $14^{h}38^{m}25^{s}$  **1-GHz flux/Jy:** 0.8? **Size/arcmin:**  $25 \times 14$  **Dec:**  $-60^{\circ}11'$  **Spectral index:** ? **Type:** S

Has been called G315.8–0.0.

**Radio:** Faint, distorted shell, with elongated trail to pulsar.

Point sources: Pulsar at end of radio trail.

### References:

Kesteven *et al.* 1987, AuJPh, 40, 855. MOST at 843 MHz  $(44'' \times 50'')$ .

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 50'' : S = 0.9 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m and ATCA OH observations.

Camilo et al. 2009, ApJ, 703, L55. Pulsar detection.Ng et al. 2012, ApJ, 746, 105. ATCA at 1.4, 2.5, 5 and 9 GHz, including polarisation, of pulsar trail.

G316.3-0.0 (MSH 14-57)

**RA:**  $14^{\text{h}}41^{\text{m}}30^{\text{s}}$  **1-GHz flux/Jy:** 20? **Size/arcmin:**  $29 \times 14$  **Dec:**  $-60^{\circ}00'$  **Spectral index:** 0.4 **Type:** S

Radio: Distorted shell, with possible 'blowout'.

X-ray: Detected.

**Distance:** HI absorption data suggests > 7.2 kpc.

## References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Caswell *et al.* 1975, A&A, 45, 239. Parkes H<sub>I</sub> absorption. Milne & Dickel 1975, AuJPh, 28, 209. Parkes 64-m at 5 GHz (4:4:S=16.7 Jy).

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz  $(44'' \times 51'')$ .

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 50'' : S = 20 \text{ Jy}).$ 

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations. Sugizaki et al. 2001, ApJS, 134, 77. ASCA survey observations. Acero et al. 2016, ApJS, 224, 8. Fermi observations.

## G317.3-0.2

**RA:** 14<sup>h</sup>49<sup>m</sup>40<sup>s</sup> **1-GHz flux/Jy:** 4.7? **Size/arcmin:** 11 **Dec:** -59°46′ **Spectral index:** ? **Type:** S

**Radio:** Incomplete shell. **Optical:** Detected.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 50'' : S = 5.2 \text{ Jy})$ .

Stupar & Parker 2011, MNRAS, 414, 2282. Ha observations.

## G318.2+0.1

Radio: Faint shell, with central HII region.

**X-ray:** Sources within remnant.

#### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 50'' : S > 4.3)$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

Bocchino et al. 2001, A&A, 367, 629. BeppoSAX and ROSAT observations

Onako et al. 2016, ApJ, 829, 106. Akari observations.

G318.9+0.4

**RA:**  $14^h58^m30^s$  **1-GHz flux/Jy:** 4? **Size/arcmin:**  $30\times14$  **Dec:**  $-58^\circ29'$  **Spectral index:** 0.2? **Type:** C

May not be a SNR?

Radio: Complex arcs, with off-centre core.

### References:

Whiteoak 1990, Nature, 347, 157. MOST at 843 MHz ( $43'' \times 51''$ ). Whiteoak 1993, ApJ, 415, 701. MOST at 843 MHz ( $43'' \times 51''$ :  $S = 4.8 \pm 0.6$  Jy), Parkes 64-m at 4.5 GHz ( $4'.7:S = 3.7 \pm 0.2$  Jy) and 8.4 GHz ( $2'.8:S = 3.0 \pm 0.4$  Jy) including polarisation, and ATCA at 1.4 GHz and 4.8 GHz ( $11'' \times 13''$ ) of core.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 50'' : S = 4.8 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

**G320.4–1.2** MSH 15–52, RCW 89

**RA:** 15<sup>h</sup>14<sup>m</sup>30<sup>s</sup> **1-GHz flux/Jy:** 60? **Size/arcmin:** 35 **Dec:** -59°08′ **Spectral index:** 0.4 **Type:** C

Has been suggested as the remnant of the SN of AD185?

Radio: Ragged shell.

**Optical:** RCW 89 is the  $H\alpha$  emitting region to the NW. **X-ray:** Partial shell, central nebula and pulsar and 'jet'. **Point sources:** Radio and X-ray pulsar, with wind nebula.

**Distance:** HI absorption indicates 5.2 kpc.

#### References:

Caswell *et al.* 1975, A&A, 45, 239. Parkes HI absorption. Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8'.4) and 5 GHz (4'.4).

Dopita et al. 1977, ApJ, 214, 179. Optical spectra.

Caswell *et al.* 1981, MNRAS, 195, 89. FIRST at 1415 MHz (50").
Seward & Harnden 1982, ApJ, 256, L45. X-ray observations of pulsar.

Manchester et al. 1982, ApJ, 262, L31. Radio observations of pulsar.

Seward et al. 1983, ApJ, 267, 698. X-ray, Optical and IR. van den Bergh & Kamper 1984, ApJ, 280, L51. Optical expansion. Seward et al. 1984, ApJ, 281, 650. X-ray observations of pulsar and nebulosity.

Lortet et al. 1987, A&A, 180, 65. Optical observations.
Trussoni et al. 1990, A&A, 234, 403. EXOSAT observations.
Asaoka & Koyama 1990, PASJ, 42, 625. Ginga X-ray spectrum.
Arendt 1991, AJ, 101, 2160. IRAS observations, including compact source.

Milne *et al.* 1993, MNRAS, 264, 853. Parkes 64-m at 4.8 GHz  $(4.5: S = 37\pm7 \text{ Jy})$  and 8.4 GHz  $(3.0: S = 24\pm4 \text{ Jy})$ , including polarisation and review of flux densities.

Strom 1994, MNRAS, 268, L5. Historical association.
Chin & Huang 1994, Nature, 371, 398. Questioning of historical association.

Matz et al. 1994, ApJ, 434, 288. X-ray observations of pulsar. Schaefer 1995, AJ, 110, 1793. Questioning of historical association.

Du Plessis *et al.* 1995, ApJ, 453, 746. Hartesbeesthoek 26-m at 2.3, 5 and 8.5 GHz (*S*=42,35.6 and 14.5 Jy).

Greiveldinger et al. 1995, ApJ, 454, 855. ROSAT observations. Trussoni et al. 1996, A&A, 306, 581. ROSAT observations.

Tamura *et al.* 1996, PASJ, 48, L33. ASCA observations. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 50'' : S = 62 \text{ Jy})$ .

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Marsden *et al.* 1997, ApJ, 491, L39. X-ray spectroscopy. Brazier & Becker 1997, MNRAS, 284, 335. ROSAT observations. Gaensler *et al.* 1999, MNRAS, 305, 724. ATCA at 1.4 GHz (21"×24"), plus Hi observations, and 5.3 GHz (10"×15"). Sako *et al.* 2000, ApJ, 537, 422. Possible γ-ray detection of pulsar. Mineo *et al.* 2001, A&A, 380, 695. BeppoSAX observations. Gaensler *et al.* 2002, ApJ, 569, 878. Chandra observations of pulsar and nebula.

Dubner *et al.* 2002, AJ, 123, 337. ATCA at 1.4 GHz (2'.7×4'.0), plus Hi observations.

Aharonian et al. 2005, A&A, 435, L17. HESS detection.
Yatsu et al. 2005, ApJ, 631, 312. Chandra observations of pulsar and jet.

DeLaney et al. 2006, ApJ, 640, 929. Chandra and ROSAT multiepoch observations of pulsar wind nebula.

Kaplan & Moon 2006, ApJ, 644, 1056. Possible IR detection of pulsar.

Forot et al. 2006, ApJ, 651, L45. X-ray observations of pulsar and nebula.

Nakamori *et al.* 2008, ApJ, 677, 297. γ-ray observations. Yatsu *et al.* 2009, PASJ, 61, 129. Chandra observations of pulsar wind nebula.

Koo *et al.* 2011, ApJ, 732, 6. Akari and Spitzer observations. An *et al.* 2014, ApJ, 793, 90. NuSTAR observations of PWN. Abdalla *et al.* 2018, A&A, 612, A1. HESS observations.

## G320.6-1.6

**RA:**  $15^{\rm h}17^{\rm m}50^{\rm s}$  **1-GHz flux/Jy:** ? **Size/arcmin:**  $60\times30$  **Dec:**  $-59^{\circ}16'$  **Spectral index:** ? **Type:** S

Radio: Faint shell, overlapping G320.4–1.2 in W.

Optical: Detected.

### References:

Milne *et al.* 1993, MNRAS, 264, 853. Parkes 64-m at 4.8 GHz (4'.5) and 8.4 GHz (3'.0), including polarisation.

Duncan et al. 1995, MNRAS, 277, 36. Parkes 64-m at 2.4-GHz (10/4).

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 50'' : S > 9.3)$ .

Stupar & Parker 2011, MNRAS, 414, 2282. Ha observations.

## G321.9-1.1

RA:  $15^h23^m45^s$  1-GHz flux/Jy: >3.4? Size/arcmin: 28 Dec:  $-58^{\circ}13'$  Spectral index: ? Type: S

Radio: Faint shell.

#### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×51": S > 3.8).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

## G321.9-0.3

**RA:**  $15^{\rm h}20^{\rm m}40^{\rm s}$  **1-GHz flux/Jy:** 13 **Size/arcmin:**  $31\times23$  **Dec:**  $-57^{\circ}34'$  **Spectral index:** 0.3 **Type:** S

**Radio:** Shell brighter to the W, with Cir X-1 to N.

Point sources: Pulsar near E edge.

#### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 18.3 Jy) and Parkes 64-m at 5 GHz (4': S = 7.8 Jy).

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz  $(44'' \times 52'')$ .

Stewart *et al.* 1993, MNRAS, 261, 593. ATCA at 1.5 GHz (21"). Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×51": S > 8.3).

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Mignani *et al.* 2002, A&A, 386, 487. HST proper motion study of Cir X-1.

Tudose *et al.* 2006, MNRAS, 372, 417. MOST at 843 MHz. Pletsch *et al.* 2013, ApJ, 779, L11. Pulsar detection. Acero *et al.* 2016, ApJS, 224, 8. Fermi observations.

## G322.1+0.0

**RA:**  $15^{\text{h}}20^{\text{m}}49^{\text{s}}$  **1-GHz flux/Jy:** ? **Size/arcmin:**  $8\times4.5$ ? **Dec:**  $-57^{\circ}10'$  **Spectral index:** ? **Type:** S?

Radio: Circular shell, with extension to S.

**X-ray:** Diffuse emission.

**Point sources:** Cir X-1 HMXB at centre.

## References:

Heinz *et al.* 2013, ApJ, 779, 171. Chandra observations and ATCA at 1.1–3.1 GHz (4"0×4".9).

## G322.5 - 0.1

**RA:** 15<sup>h</sup>23<sup>m</sup>23<sup>s</sup> **1-GHz flux/Jy:** 1.5 **Dec:** -57°06' **Spectral index:** 0.4 **Size/arcmin:** 15 **Type:** C

**Radio:** Shell with central extended source. **Point sources:** PN Pe 2-8 within boundary.

### References

Whiteoak 1992, MNRAS, 256, 121. MOST at 843 MHz  $(43'' \times 51'' : S = 2.0 \pm 0.3 \text{ Jy})$ .

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 51'' : S = 1.7 \text{ Jy})$ , plus Parkes 64-m at 4.5 GHz (=  $0.89 \pm 0.13$ ).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

## G323.5+0.1

**RA:** 15<sup>h</sup>28<sup>m</sup>42<sup>s</sup> **1-GHz flux/Jy:** 3? **Size/arcmin:** 13 **Dec:** -56°21′ **Spectral index:** 0.4? **Type:** S

**Radio:** Distorted shell, confused with thermal emission. **Point sources:** Compact, probably thermal source near centre.

#### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 4.2 Jy) and Parkes 64-m at 5 GHz (4': S = 1.5 Jy).

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×52": S=4.2 Jy).

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Reach *et al.* 2006, AJ, 131, 1479. Spitzer possible detection.

## G323.7 - 1.0

**RA:**  $15^{h}34^{m}30^{s}$  **1-GHz flux/Jy:** ? **Size/arcmin:**  $51\times38$  **Dec:**  $-57^{\circ}12'$  **Spectral index:** ? **Type:** S

Radio: Faint shell.

**X-ray:** Faint diffuse emission.

**Distance:** CO and H<sub>I</sub> observations suggest 3.5 kpc.

References:

Green et al. 2014, PASA, 31, 42. MGPS observations at 843 MHz

 $(43'' \times 51'')$ .

Araya et al. 2017, ApJ, 843, 12. Fermi observations.

Abdalla *et al.* 2018, A&A, 612, A8. HESS observations. Maxted *et al.* 2018, MNRAS, 480, 134. CO and HI observations. Saji *et al.* 2018, PASJ, 70, 23. Suzaku observations.

G326.3-1.8 MSH 15-56

RA: 15<sup>h</sup>53<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: 145 Size/arcmin: 38 Dec: -56°10′ Spectral index: varies Type: C

Has been called G326.2-1.7.

Radio: Shell, with elongated, flat-spectrum core.

Optical: Emission around the shell.

**X-ray:** Shell, with central extended emission.

**Point sources:** Compact X-ray source.

### References:

Clark *et al.* 1975, AuJPA, 37, 75. Molonglo at 408 MHz (3': S = 180 Jy)

Caswell *et al.* 1975, A&A, 45, 239. Parkes Hi absorption. Milne *et al.* 1979, MNRAS, 188, 437. FIRST at 1415 MHz (0.8:

S > 95) and Parkes 64-m at 14.7 GHz (2'.2:  $S = 69 \pm 8$  Jy). van den Bergh 1979, ApJ, 227, 497. Optical observations.

Zealey *et al.* 1979, A&AS, 38, 39. Optical observations.

Dennefeld 1980, PASP, 92, 603. Optical observations

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz (44"×53").

Milne et al. 1989, PASA, 8, 187. MOST at 843 MHz ( $43'' \times 52''$ :  $S = 153 \pm 40$  Jy), and Parkes 64-m at 8.4 GHz (3':  $S = 68 \pm 5$  Jy), including polarisation.

Seward 1990, ApJS, 73, 781. Einstein observations. Kassim *et al.* 1993, ApJ, 419, 733, ROSAT image.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×52": S > 130).

Rosado et al. 1996, A&A, 315, 243. Optical kinematics.

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Dickel *et al.* 2000, ApJ, 543, 840. ATCA at 1.34 GHz ( $6''.4 \times 8''.5 : S > 60$  Jy), 4.80 GHz (3''.8 : S = 25 Jy for core only) and 8.64 GHz (3''.2 : S = 15 Jy for core only).

Temim *et al.* 2013, ApJ, 768, 61. XMM-Newton and Chandra observations.

Yatsu et al. 2013, ApJ, 773, 25. XMM-Newton and Chandra observations.

Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 5 frequencies between 30 and 143 GHz.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

Temim et al. 2017, ApJ, 851, 128. Chandra proper motion study of central source.

Devin et al. 2018, A&A, 617, A5. Fermi observations.

## G327.1-1.1

RA: 15<sup>h</sup>54<sup>m</sup>25<sup>s</sup> 1-GHz flux/Jy: 7? Size/arcmin: 18 Dec: -55°09′ Spectral index: ? Type: C

Radio: Shell, with off-centre core.

**X-ray:** Diffuse, with core.

#### References:

Milne & Dickel 1974, AuJPh, 27, 549. Parkes 64-m at 2.7 GHz  $(8.4:S=10\pm15\% \text{ Jy})$ .

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 10.6 Jy) and Parkes 64-m at 5 GHz (4': S = 4.3 Jy).

Lamb & Markert 1981, ApJ, 244, 94. Einstein observations. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×52": S=7.6 Jy).

Seward *et al.* 1996, ApJ, 471, 887. ROSAT observations. Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Sun *et al.* 1999, ApJ, 511, 274. ASCA and ROSAT observations. Bocchino & Bandiera 2003, A&A, 398, 195. BeppoSAX observations.

Temim et al. 2009, ApJ, 691, 895. Chandra and XMM-Newton observations.

Temim *et al.* 2015, ApJ, 808, 100. Chandra observations. Ma *et al.* 2016, ApJ, 820, 100. ATCA at 1.4  $(19'' \times 23'')$ , 2.3  $(11'' \times 16'')$ , 4.8  $(13'' \times 15'')$  and 8.6 GHz  $(8'' \times 10'')$  of core, including polarisation.

Abdalla et al. 2018, A&A, 612, A1. HESS observations.

## G327.2 - 0.1

**RA:**  $15^h50^m55^s$  **1-GHz flux/Jy:** 0.4 **Size/arcmin:** 5 **Dec:**  $-54^{\circ}18'$  **Spectral index:** ? **Type:** S

Has been called G327.24-0.13.

**Radio:** Shell, possibly with central emission. **Point sources:** Central pulsar (magnetar).

#### References:

Camilo *et al.* 2007, ApJ, 666, L93. Pulsar observations. Gelfand & Gaensler 2007, ApJ, 667, 1111. Chandra and XMM-Newton observations of magnetar, and 843 GHz  $(43'' \times 53'' : S = 0.5 \pm 0.1 \text{ Jy})$  and SGPS at 1.4 GHz  $(S = 0.3 \pm 0.1 \text{ Jy})$ .

G327.4+0.4 Kes 27

**RA:** 15<sup>h</sup>48<sup>m</sup>20<sup>s</sup> **1-GHz flux/Jy:** 30? **Size/arcmin:** 21 **Dec:** -53°49′ **Spectral index:** 0.6 **Type:** S

Has been called G327.3+0.4 and G327.3+0.5.

Radio: Incomplete, multi-arc shell, brightest to the SE.

**X-ray:** Diffuse, brighter in E.

**Distance:** HI absorption indicates 4.3 to 5.4 kpc.

### References:

Caswell *et al.* 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3':S = 58 Jy) and Parkes 64-m at 5 GHz (4':S = 12.4 Jy).

Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8.4) and 5 GHz (4.4).

Lamb & Markert 1981, ApJ, 244, 94. Einstein observations. Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz (44"×55").

Milne et al. 1989, PASA, 8, 187. MOST at 843 MHz  $(43'' \times 53'' : S = 32.2 \pm 6 \text{ Jy})$ , and Parkes 64-m at 8.4 GHz  $(3' : S = 9.4 \pm 0.8 \text{ Jy})$ , including polarisation.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×53": S=25 Jy).

Seward et al. 1996, ApJ, 471, 887. ROSAT observations.
Green et al. 1997, AJ, 114, 2058. Parkes 64-m and ATCA OH observations.

McClure-Griffiths *et al.* 2001, ApJ, 551, 394. ATCA and Parkes 64-m at 1.4 GHz  $(1'8\times2'0)$ , plus HI.

Enoguchi et al. 2002, PASJ, 54, 229. ASCA observations. Kawasaki et al. 2005, ApJ, 631, 935. ASCA observations. Reach et al. 2006, AJ, 131, 1479. Spitzer possible detection. Chen et al. 2008, ApJ, 676, 1040. Chandra observations. Minami et al. 2014, in EFXU, p48. Suzaku observations. Xing et al. 2015, ApJ, 805, 19. Fermi observations.

# G327.4+1.0

**RA:** 15<sup>h</sup>46<sup>m</sup>48<sup>s</sup> 1-GHz flux/Jy: 1.9? Size/arcmin: 14 **Dec:** -53°20′ **Spectral index:** ? Type: S

Radio: Asymmetric shell.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×54": S=2.1 Jy).

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations.

McClure-Griffiths *et al.* 2001, ApJ, 551, 394. ATCA and Parkes 64-m at 1.4 GHz (1'.8×2'.0), plus HI.

**G327.6+14.6** SN1006, PKS 1459–41

RA: 15<sup>h</sup>02<sup>m</sup>50<sup>s</sup> 1-GHz flux/Jy: 19 Size/arcmin: 30 Dec: -41°56′ Spectral index: 0.6 Type: S

This is the remnant of the SN of AD1006.

**Radio:** Shell, with two bright arcs.

**Optical:** Filaments to the NW, with broad  $H\alpha$  component.

**X-ray:** Thermal shell, with non-thermal limb-brightened arcs.

**Point sources:** The background Schweizer–Middleditch star is near the middle of the remnant.

**Distance:** Optical spectra and proper motion indicate 1.85 kpc.

#### References:

van den Bergh 1976, ApJ, 208, L17. Optical observations. Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8.4) and 5 GHz (4.4).

Schweizer & Middleditch 1980, ApJ, 241, 1039. Possible stellar remnant.

Caswell *et al.* 1983, MNRAS, 204, 921. FIRST at 1415 MHz (77").

Reynolds & Gilmore 1986, AJ, 92, 1138. VLA at 1.37 and 1.67 GHz  $(16''\times20'')$ .

Kirshner et al. 1987, ApJ, 315, L135. Broad Hα optical component

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz  $(44'' \times 66'')$ .

Fesen *et al.* 1988, ApJ, 327, 164. UV absorption spectra of the Schweizer–Middleditch star.

Roger *et al.* 1988, ApJ, 332, 940. MOST at 843 MHz ( $44'' \times 66''$ :  $S = 17.5 \pm 1.5$  Jy).

Long et al. 1988, ApJ, 333, 749. Optical proper motion for distance.

Wu et al. 1993, ApJ, 416, 247. UV spectra.

Reynolds & Gilmore 1993, AJ, 106, 272. VLA at 1.37 and 1.67 GHz (24"), including polarisation.

Moffett *et al.* 1993, AJ, 106, 1566. VLA at 1.37 and 1.67 GHz (10"×14"), for proper motion studies.

Ozaki *et al.* 1994, PASJ, 46, 367. X-ray observations. Raymond *et al.* 1995, ApJ, 454, L31. Far UV spectroscopy. Koyama *et al.* 1995, Nature, 378, 255. ASCA observations. Willingale *et al.* 1996, MNRAS, 278, 749. ROSAT observations.

Blair *et al.* 1996, ApJ, 468, 871. Optical absorption studies. Laming *et al.* 1996, ApJ, 472, 267. Modelling of optical spectra,

including distance. Wu *et al.* 1997, ApJ, 477, L53. Far UV absorption spectra of the Schweizer–Middleditch star.

Winkler & Long 1997, ApJ, 486, L137. UV absorption spectra of background quasar.

Winkler & Long 1997, ApJ, 491, 829. ROSAT and optical images. Vink *et al.* 2000, A&A, 354, 931. X-ray spectroscopy.

Burleigh *et al.* 2000, A&A, 356, 585. Optical spectroscopy of the Schweizer–Middleditch star.

Allen et al. 2001, ApJ, 558, 739. ASCA, ROSAT and other X-ray observations.

Dubner *et al.* 2002, A&A, 387, 1047. ATCA at 1.4 GHz  $(3.0 \times 4.7)$  for HI, plus CO observations.

Ghavamian et al. 2002, ApJ, 572, 888. Optical of filaments in NW. Sollerman et al. 2003, A&A, 407, 249. Optical spectroscopy. Winkler et al. 2003, ApJ, 585, 324. Optical proper motion studies. Vink et al. 2003, ApJ, 587, L31. XMM-Newton observations.

Bamba et al. 2003, ApJ, 589, 827. Chandra observations of NE. Winkler et al. 2005, ApJ, 624, 189. HST absorption towards background sources.

Kalemci et al. 2006, ApJ, 644, 274. INTEGRAL observations.
Acero et al. 2007, A&A, 475, 883. XMM-Newton observations.
Raymond et al. 2007, ApJ, 659, 1257. HST Hα observations.
Hamilton et al. 2007, MNRAS, 381, 771. UV spectroscopy of the Schweizer-Middleditch star.

Bamba *et al.* 2008, AdSpR, 41, 411. Suzaku observations. Cassam-Chenaï *et al.* 2008, ApJ, 680, 1180. ATCA and VLA at 1.5 GHz (6"×9") plus Chandra observations.

Yamaguchi et al. 2008, PASJ, 60, S141. Suzaku observations. Dyer et al. 2009, AJ, 137, 2956. GBT and VLA observations at 1.4 GHz.

Winkler et al. 2011, ApJ, 742, 80. Multi-epoch UV spectroscopy of the Schweizer-Middleditch star.

Miceli et al. 2013, A&A, 550, A80. XMM-Newton observations. Broersen et al. 2013, A&A, 552, A9. XMM-Newton observations of NW knot.

Katsuda et al. 2013, ApJ, 763, 85. Chandra proper motion study of NW.

Winkler et al. 2013, ApJ, 764, 156. Spitzer observations. Uchida et al. 2013, ApJ, 771, 56. Suzaku observations.

Reynoso *et al.* 2013, ÅJ, 145, 104. VLA and ATCA at 1.4 GHz (10"), including polarisation.

Nikolić et al. 2013, Science, 340, 45. Optical spectroscopy.Winkler et al. 2014, ApJ, 781, 65. Chandra observations and Hα spectroscopy.

Miceli *et al.* 2014, ApJ, 782, L33. XMM-Newton observations of SW.

Acero et al. 2015, A&A, 580, A74. Fermi limit.

Sparls et al. 2015, ApJ, 815, L9. Polarised optical line observations.

Li *et al.* 2015, MNRAS, 453, 3953. XMM-Newton observations. Arnaud *et al.* 2016, A&A, 586, A134. Planck flux densities at 4 frequencies between 30 and 100 GHz.

Xing et al. 2016, ApJ, 823, 44. Fermi observations.

Li et al. 2016, MNRAS, 462, 158. XMM-Newton observations. Raymond et al. 2017, ApJ, 851, 12. UV and optical spectra, for proper motions and distance.

Condon et al. 2017, ApJ, 851, 100. Fermi observations.

Kerzendorf et al. 2018, MNRAS, 479, 192. Optical search for progenitor companion.

Li et al. 2018, ApJ, 864, 85. NuSTAR observations of NE and SW limbs.

G328.4+0.2 (MSH 15–57)

**RA:** 15<sup>h</sup>55<sup>m</sup>30<sup>s</sup> **1-GHz flux/Jy:** 15 **Dec:** -53°17' **Spectral index:** 0.0 **Type:** F

Radio: Amorphous emission, with central bar.

**X-ray:** Detected at high energies.

**Distance:** H<sub>I</sub> absorption indicates > 17.4 kpc.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Caswell et al. 1980, MNRAS, 190, 881. FIRST at 1415 MHz

Wilson 1986, ApJ, 302, 718. Lack of detection with Einstein. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 54'' : S = 15 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

Gaensler *et al.* 2000, ApJ, 542, 380. ATCA at 1.4 GHz  $(16.5 \times 19.5 : S = 14.3 \pm 0.1 \text{ Jy})$  and 4.5 GHz  $(1.5 \times 2.0 : S = 12.5 \pm 0.2 \text{ Jy})$ .

Hughes *et al.* 2000, ApJ, 542, 386. ASCA observations. McClure-Griffiths *et al.* 2001, ApJ, 551, 394. ATCA and Parkes 64-m at 1.4 GHz (1/8×2/0), plus HI.

Johnston *et al.* 2004, MNRAS, 348, L19. ATCA at 19 GHz (6"1×7".7), including polarisation.

Gelfand *et al.* 2007, Apl, 663, 468. ATCA at 1.4 GHz (5".8 $\times$ 7".0:  $S=13.8\pm0.4$  Jy), plus XMM-Newton observations.

G329.7 + 0.4

**RA:**  $16^h01^m20^s$  **1-GHz flux/Jy:** >34? **Size/arcmin:**  $40\times33$  **Dec:**  $-52^\circ18'$  **Spectral index:** ? **Type:** S

Radio: Diffuse shell, in complex region.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×53": S > 38).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

McClure-Griffiths *et al.* 2001, ApJ, 551, 394. ATCA and Parkes 64-m at 1.4 GHz (1/8×2/0), plus HI.

Reach et al. 2006, AJ, 131, 1479. Spitzer possible detection.

G330.0+15.0 Lupus Loop

**RA:** 15<sup>h</sup>10<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** 350? **Size/arcmin:** 180? **Dec:** -40°00′ **Spectral index:** 0.5? **Type:** S

**Radio:** Low surface brightness loop with H<sub>I</sub> shell.

**X-ray:** Detected, with central source.

**Point sources:** Central, possibly pulsating, X-ray source.

### References:

Milne 1971, AuJPh, 24, 757. Parkes 64-m at 408 MHz (48'), 635 MHz (31') and 1410 MHz (15').

Milne & Dickel 1974, AuJPh, 27, 549. Parkes 64-m at 2.7 GHz  $(8.4:S=120\pm30\% \text{ Jy})$ .

Toor 1980, A&A, 85, 184. X-ray image and spectrum.

Colomb & Dubner 1982, A&A, 112, 141. Argentine 30-m dish at 1.42 GHz (30'), HI observations.

Leahy *et al.* 1991, ApJ, 374, 218. HEAO-1 X-ray spectra. Ozaki *et al.* 1994, PASJ, 46, 367. X-ray observations.

Kaplan *et al.* 2006, ApJS, 163, 344. X-ray upper limit on compact sources.

Shinn et al. 2006, ApJ, 644, L189. Far UV observations.

## G330.2 + 1.0

**RA:** 16<sup>h</sup>01<sup>m</sup>06<sup>s</sup> **1-GHz flux/Jy:** 5? **Size/arcmin:** 11 **Dec:** -51°34′ **Spectral index:** 0.3 **Type:** S?

Radio: Clumpy non-thermal emission, possibly a distorted shell.

X-ray: Shell.

**Point sources:** Central compact X-ray source. **Distance:** HI absorption indicates > 4.9 kpc.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 8.6 Jy) and Parkes 64-m at 5 GHz (4': S = 4.0 Jy).

Caswell *et al.* 1983, MNRAS, 204, 915. FIRST at 1415 MHz (47"×52"), and MOST at 843 MHz (43"×55").

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×55": S=4.7 Jy).

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. McClure-Griffiths *et al.* 2001, ApJ, 551, 394. ATCA and Parkes 64-m at 1.4 GHz (1/8×2/0), plus HI.

Torii et al. 2006, PASJ, 58, L11. ASCA detection.

Park et al. 2009, ApJ, 695, 431. XMM-Newton and Chandra observations.

Abramowski et al. 2014, MNRAS, 441, 790. HESS upper limit. Doroshenko et al. 2018, A&A, 619, A26. XMM-Newton observations of central source.

Williams et al. 2018, ApJ, 855, 118. XMM-Newton observations. Borkowski et al. 2018, ApJ, 868, L21. Chandra observations for expansion study.

## G332.0+0.2

**RA:** 16<sup>h</sup>13<sup>m</sup>17<sup>s</sup> **1-GHz flux/Jy:** 8? **Size/arcmin:** 12 **Dec:** -50°53′ **Spectral index:** 0.5 **Type:** S

**Radio:** Incomplete shell.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3 $^{\prime}$ ) and Parkes 64-m at 5 GHz (4 $^{\prime}$ ).

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz  $(44'' \times 57'')$ .

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 55'' : S = 8.9 \text{ Jy})$ .

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. McClure-Griffiths *et al.* 2001, ApJ, 551, 394. ATCA and Parkes 64-m at 1.4 GHz  $(1.8\times2.0)$ , plus HI.

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

G332.4-0.4 RCW 103

RA: 16<sup>h</sup>17<sup>m</sup>33<sup>s</sup> 1-GHz flux/Jy: 28 Size/arcmin: 10 Dec: -51°02′ Spectral index: 0.5 Type: S

Radio: Shell, brightest to the S.

**Optical:** Filaments correspond well to the radio shell, brightest in SE.

**X-ray:** Brightest to NW, with point source near centre.

**Point sources:** Central, variable X-ray source, and nearby pulsar.

**Distance:** HI absorption indicates 3.1 kpc.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Caswell et al. 1975, A&A, 45, 239. Parkes HI absorption.

Tuohy & Garmire 1980, ApJ, 239, L107. Einstein detection of X-ray point source.

Caswell *et al.* 1980, MNRAS, 190, 881. FIRST at 1415 MHz (50").

Lamb & Markert 1981, ApJ, 244, 94. Einstein observations. Ruiz 1983, AJ, 88, 1210. Optical spectra.

Tuohy et al. 1983, ApJ, 268, 778. X-ray observations.

Leibowitz & Danziger 1983, MNRAS, 204, 273. Optical spectra. Nugent *et al.* 1984, ApJ, 284, 612. X-ray spectrum (and Einstein image from Tuohy, private communication).

Meaburn & Allen 1986, MNRAS, 222, 593. Optical spectra. Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz (44"×57").

Oliva et al. 1990, A&A, 240, 453. IR spectroscopy

Dickel *et al.* 1996, AJ, 111, 340. ATCA at 1.36 (8") and 2.37 GHz (4".5"), including polarisation.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×55": S=34 Jy).

Frail *et al.* 1996, AJ, 111, 1651. OH emission near remnant. Gotthelf *et al.* 1997, ApJ, 487, L175. ASCA observations of compact X-ray source.

Carter et al. 1997, PASP, 109, 990. Optical expansion. Torii et al. 1998, ApJ, 494, L207. ASCA detection of nearby pul-

Kaspi et al. 1998, ApJ, 503, L161. Pulsar observations.

Gotthelf *et al.* 1999, ApJ, 514, L107. X-ray variability of central source.

Oliva et al. 1999, A&A, 343, 943. ISO spectroscopy.

Torii *et al.* 2000, ApJ, 534, L71. X-ray timing on pulsar, including glitch.

Reynoso *et al.* 2004, PASA, 21, 82. ATCA at 1.4 GHz (50"), including H<sub>I</sub> absorption to central source.

Russeil *et al.* 2005, A&A, 429, 497. Hα observations. Reach *et al.* 2006, AJ, 131, 1479. Spitzer observations.

Paron 2006, PASA, 23, 69. CO and HCO+ observations of surroundings.

De Luca 2006, Science, 313, 814. XMM-Newton observations of central source.

De Luca 2007, Ap&SS, 308, 231. XMM-Newton observation of periodicity of central source.

Matsumoto *et al.* 2007, PASJ, 59, S199. Suzaku observations. De Luca 2008, ApJ, 682, 1185. IR observations of central source.

Andersen et al. 2011, ApJ, 742, 7. Spitzer observations.

Xing et al. 2014, ApJ, 781, 64. Fermi observations. Frank et al. 2015, ApJ, 810, 113. Chandra observations.

Koo *et al.* 2016, Api, 821, 20. Spitzer and Herschel flux densities (and comparison with X-ray properties).

Rea et al. 2016, ApJ, 828, L13. Chandra, NuSTAR and Swift observations.

Borghese et al. 2018, MNRAS, 478, 741. NuSTAR and Swift observations of central source.

**G332.4+0.1** MSH 16–51, Kes 32

RA: 16<sup>h</sup>15<sup>m</sup>20<sup>s</sup> 1-GHz flux/Jy: 26 Size/arcmin: 15 Dec: -50°42′ Spectral index: 0.5 Type: S

Has been called G332.4+0.2.

Radio: Distorted shell, with thermal jet and plume adjacent.

Optical: Detected.

**X-ray:** Shell, brightest to NW. **Point sources:** Pulsar nearby.

### References

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Roger *et al.* 1985, Nature, 316, 44. MOST at 843 MHz  $(44'' \times 57'')$ .

Caraveo 1993, ApJ, 415, L111. Nearby pulsar.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×56": S=29 Jy).

Frail *et al.* 1996, AJ, 111, 1651. OH emission near remnant. Brinkmann *et al.* 1999, A&A, 346, 599. ROSAT image of remnant and nearby pulsar.

Vink *et al.* 2004, ApJ, 604, 693. Chandra observations. Reach *et al.* 2006, AJ, 131, 1479. Spitzer possible detection. Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations.

## G332.5-5.6

RA: 16<sup>h</sup>43<sup>m</sup>20<sup>s</sup> 1-GHz flux/Jy: 2? Size/arcmin: 35 Dec: -54°30′ Spectral index: 0.7? Type: S

Radio: Bipolar shell, with central emission also.

**Optical:** Patchy filaments. **X-ray:** Emission from centre.

**Distance:** Optical extinction suggests 3 kpc.

#### References

Reynoso & Green 2007, MNRAS, 375, 92. ATCA at 1.4 GHz  $(40'': S = 1.90 \pm 0.15 \text{ Jy})$  and 2.4 GHz  $(90'': S = 1.3 \pm 0.2 \text{ Jy})$  including polarisation, and 843 MHz  $(43'' \times 53'')$  plus ROSAT observations.

Stupar *et al.* 2007, MNRAS, 381, 377. Optical observations, plus ATCA at 1.4 GHz (58") and 2.4 GHz (95") including polarisation, and other observations.

Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations. Zhu *et al.* 2015, MNRAS, 452, 3470. Suzaku observations, and distance estimate.

Suárez *et al.* 2015, A&A, 583, A84. XMM-Newton observations of central region.

Ackermann et al. 2018, ApJS, 237, 32. Fermi observations.

## G335.2+0.1

RA: 16<sup>h</sup>27<sup>m</sup>45<sup>s</sup> 1-GHz flux/Jy: 16 Size/arcmin: 21 Dec: -48°47′ Spectral index: 0.5 Type: S

Radio: Well defined shell.

 $(43'' \times 57'' : S = 16 \text{ Jy}).$ 

**Point sources:** Old pulsar within remnant boundary.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': *S* = 27.1 Jy) and Parkes 64-m at 5 GHz (4': *S* = 8.6 Jy). Kaspi *et al.* 1996, AJ, 111, 2028. Pulsar observations. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Reach *et al.* 2006, AJ, 131, 1479. Spitzer possible detection. de Wilt *et al.* 2017, MNRAS, 468, 2093. Molecular line observations of region.

## G336.7 + 0.5

**RA:**  $16^{\text{h}}32^{\text{m}}11^{\text{s}}$  **1-GHz flux/Jy:** 6 **Size/arcmin:**  $14\times10$  **Dec:**  $-47^{\circ}19'$  **Spectral index:** 0.5 **Type:** S

Radio: Irregular shell. Optical: Detected.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3 $^{\prime}$ ) and Parkes 64-m at 5 GHz (4 $^{\prime}$ ).

Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz  $(44'' \times 60'')$ .

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×58": S=6.1 Jy).

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations.

G337.0-0.1 (CTB 33)

**RA:** 16<sup>h</sup>35<sup>m</sup>57<sup>s</sup> **1-GHz flux/Jy:** 1.5 **Dec:** -47°36′ **Size/arcmin:** 1.5 **Type:** S

This entry refers to a small (1.5) SNR, not the larger previously catalogued G337.0–0.1. Has mistakenly been called G337.7–0.1.

Radio: Shell, in a complex region.

**Point sources:** Associated with a soft gamma repeater. **Distance:** Association with CTB 33 gives 11 kpc.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 58'' : S = 21 \text{ Jy})$ .

Frail et al. 1996, AJ, 111, 1651. OH emission near remnant, including masers.

Sarma *et al.* 1997, ApJ, 483, 335. ATCA at 1.4 GHz (12") including HI, and 8.9 GHz (13"×15") for recombination lines, clarifying extent of the remnant.

Woods et al. 1999, ApJ, 519, L139. Soft gamma repeater observations.

Hurley et al. 2000, ApJ, 528, L21. ASCA observations of soft gamma repeater.

Brogan et al. 2000, ApJ, 537, 875. VLA at 1.7 GHz for OH Zeeman splitting.

Castro et al. 2013, ApJ, 774, 36. Fermi observations.

G337.2 - 0.7

**RA:** 16<sup>h</sup>39<sup>m</sup>28<sup>s</sup> **1-GHz flux/Jy:** 1.5 **Size/arcmin:** 6 **Dec:** -47°51′ **Spectral index:** 0.4 **Type:** S

**Radio:** Shell, brighter in S. **X-ray:** Extended emission.

**Distance:** Hi absorption suggests 2.0 to 9.3 kpc.

References

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 3.8 Jy) and Parkes 64-m at 5 GHz (4': S = 0.70 Jy).

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43" × 58": S = 2.0 Jy).

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Rakowski *et al.* 2001, ApJ, 548, 258. ASCA and ROSAT observations.

Rakowski *et al.* 2006, ApJ, 646, 982. ATCA observations at 1.3 GHz (15":  $S=1.55\pm0.05$  Jy) and 5 GHz (15":  $S=0.93\pm0.02$  Jy), plus HI and Chandra observations. Yamaguchi *et al.* 2014, ApJ, 785, L27. Suzaku observations.

Takata et al. 2016, PASJ, 68, S3. Suzaku observations.

G337.2+0.1

RA:  $16^{\text{h}}35^{\text{m}}55^{\text{s}}$  1-GHz flux/Jy: 1.5? Size/arcmin:  $3\times2$  Dec:  $-47^{\circ}20'$  Spectral index: ? Type: ?

**Radio:** Not well defined.

X-ray: Detected.

**Distance:** Association with H<sub>I</sub> hole gives 14 kpc.

References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz

 $(43'' \times 59'' : S = 1.6 \text{ Jy})$ 

Combi et al. 2005, A&A, 431, L9. ASCA and other observations.

Aharonian et al. 2006, ApJ, 636, 777. HESS observations of nearby source

Combi et al. 2006, ApJ, 653, L41. XMM-Newton observations.

G337.3+1.0 Kes 40

**RA:**  $16^{\text{h}}32^{\text{m}}39^{\text{s}}$  **1-GHz flux/Jy:** 16 **Size/arcmin:**  $15 \times 12$  **Dec:**  $-46^{\circ}36'$  **Spectral index:** 0.55 **Type:** S

Radio: Nearly complete shell.

#### References:

 $(44'' \times 61'')$ .

Caswell *et al.* 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3': S = 24.6 Jy) and Parkes 64-m at 5 GHz (4': S = 7.2 Jy). Dickel & Milne 1976, AuJPh, 29, 435. Comparison of earlier Parkes 64-m maps at 2.7 GHz (8'.4) and 5 GHz (4'.4). Kesteven & Caswell 1987, A&A, 183, 118. MOST at 843 MHz

Milne *et al.* 1989, PASA, 8, 187. MOST at 843 MHz  $(43'' \times 59'' : S = 14.8 \pm 3.0 \text{ Jy})$ , and Parkes 64-m at 8.4 GHz  $(3' : S = 5.1 \pm 0.6 \text{ Jy})$ , including polarisation.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×59": S=20 Jy).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

**G337.8–0.1** Kes 41

**RA:**  $16^{\text{h}}39^{\text{m}}01^{\text{s}}$  **1-GHz flux/Jy:** 15 **Size/arcmin:**  $9\times6$  **Dec:**  $-46^{\circ}59'$  **Spectral index:** 0.5 **Type:** S

Radio: Distorted shell.X-ray: Centrally brightened.

**Distance:** HI absorption suggests 11 kpc.

#### References

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Caswell *et al.* 1975, A&A, 45, 239. Parkes HI absorption. Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 59'' : S = 18 \text{ Jy})$ .

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Koralesky *et al.* 1998, AJ, 116, 1323. VLA detection of compact OH emission.

Caswell 2004, MNRAS, 349, 99. ATCA at 1.7 GHz, for associated OH masers.

Reach et al. 2006, AJ, 131, 1479. Spitzer possible detection.

Kothes & Dougherty 2007, A&A, 468, 993. SGPS at 1.4 GHz including Hi.

Combi et al. 2008, A&A, 488, L25. XMM-Newton observations.
Zhang et al. 2015, ApJ, 799, 103. XMM-Newton, Chandra and CO observations.

Liu et al. 2015, ApJ, 809, 102. Fermi observations.
Acero et al. 2016, ApJS, 224, 8. Fermi observations.
Qiao et al. 2016, ApJS, 227, 26. OH maser observations.
Supan et al. 2018, A&A, 619, A108. CO, HI and Spitzer observations.

Supan et al. 2018, A&A, 619, A109. Fermi observations and review of radio flux densities.

# G338.1+0.4

**RA:** 16<sup>h</sup>37<sup>m</sup>59<sup>s</sup> **1-GHz flux/Jy:** 4? **Size/arcmin:** 15? **Dec:** -46°24′ **Spectral index:** 0.4 **Type:** S

**Radio:** Arc in NE, merging with thermal emission in S.

Optical: Detected.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Zealey et al. 1979, A&AS, 38, 39. Optical detection.

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 59'' : S = 3.8 \text{ Jy}).$ 

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

### G338.3-0.0

**RA:**  $16^{\text{h}}41^{\text{m}}00^{\text{s}}$  **1-GHz flux/Jy:** 7? **Size/arcmin:** 8 **Dec:**  $-46^{\circ}34'$  **Spectral index:** ? **Type:** C?

**Radio:** Irregular shell, in complex region. **X-ray:** Central X-ray source and nebula.

Point sources: Central pulsar.

**Distance:** Hi observations suggest 8 to 13 kpc.

#### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×59": S=7.4 Jy).

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Sugizaki *et al.* 2001, ApJS, 134, 77. ASCA survey observations. Aharonian *et al.* 2005, Science, 307, 1938. HESS detection. Aharonian *et al.* 2006, ApJ, 636, 777. HESS observations. Landi *et al.* 2006, ApJ, 651, 190. X-ray observations.

Funk et al. 2007, ApJ, 662, 517. XMM-Newton observations. Lemiere et al. 2009, ApJ, 706, 1269. Chandra observations of central source and nebula.

Slane et al. 2010, ApJ, 720, 266. Fermi observations.

Castelletti *et al.* 2011, A&A, 536, A98. GMRT at 235 MHz (10"×26"), 610 MHz (5".0×12".6) and 1280 MHz (5".0×6".4), plus ATCA at 2.3 GHz (4".3×4".8).

Gotthelf *et al.* 2014, ApJ, 788, 155. NuSTAR observations. Lemoine-Goumard *et al.* 2014, ApJ, 794, L16. Fermi observations. Abramowski *et al.* 2014, MNRAS, 439, 2828. HESS observations. *see also:* Abramowski *et al.* 2014, MNRAS, 441, 3640. Erratum. Supan *et al.* 2016, A&A, 589, A51, SGPS HI and other observa-

Supan et al. 2016, A&A, 589, A51. SGPS HI and other observations.

Lau et al. 2017, MNRAS, 464, 3757. CO and other molecular line observations, plus HI observations.de Wilt et al. 2017, MNRAS, 468, 2093. Molecular line observa-

de Wilt et al. 2017, MNRAS, 468, 2093. Molecular line observations of region.

Xin et al. 2018, ApJ, 867, 55. Fermi observations.

## G338.5+0.1

**RA:**  $16^{\rm h}41^{\rm m}09^{\rm s}$  **1-GHz flux/Jy:** 12? **Size/arcmin:** 9 **Dec:**  $-46^{\circ}19'$  **Spectral index:** ? **Type:** ?

Radio: Circle of non-thermal emission in complex region, not well defined.

Distance: HI absorption suggests 11 kpc.

### References:

Shaver & Goss 1970, AuJPA, 14, 133. Molonglo at 408 MHz (3') and Parkes 64-m at 5 GHz (4').

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 59'' : S = 13 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.
 Kothes & Dougherty 2007, A&A, 468, 993. SGPS at 1.4 GHz including Hr.

Abramowski et al. 2014, ApJ, 794, L1. HESS observations. Lau et al. 2017, MNRAS, 464, 3757. CO and other molecular line observations, plus HI observations.

de Wilt *et al.* 2017, MNRAS, 468, 2093. Molecular line observations of region.

### G340.4+0.4

**RA:**  $16^{\rm h}46^{\rm m}31^{\rm s}$  **1-GHz flux/Jy:** 5 **Size/arcmin:**  $10\times7$  **Dec:**  $-44^{\circ}39'$  **Spectral index:** 0.4 **Type:** S

Radio: Distorted shell, elongated east-west.

Optical: Detected.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 8.2 Jy) and Parkes 64-m at 5 GHz (4': S = 2.9 Jy).

Caswell *et al.* 1983, MNRAS, 203, 595. FIRST at 1415 MHz (50").

Dubner *et al.* 1996, AJ, 111, 1304. VLA at 330 MHz  $(48'' \times 75'' : S = 9.8 \pm 0.9 \text{ Jy})$  and 1.4 GHz  $(9'' \times 27'' : S = 3.6 \pm 0.1 \text{ Jy})$ .

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 61'' : S = 5.9 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations. Stupar & Parker 2011, MNRAS, 414, 2282.  $H\alpha$  observations.

## G340.6+0.3

**RA:** 16<sup>h</sup>47<sup>m</sup>41<sup>s</sup> **1-GHz flux/Jy:** 5? **Size/arcmin:** 6 **Dec:** -44°34′ **Spectral index:** 0.4? **Type:** S

Radio: Incomplete shell.

**Optical:** Possible associated filaments. **Distance:** HI absorption suggests 15 kpc.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 7.0 Jy) and Parkes 64-m at 5 GHz (4': S = 2.8 Jy).

Zealey *et al.* 1979, A&AS, 38, 39. Optical observations.

Caswell *et al.* 1983, MNRAS, 203, 595. FIRST at 1415 MHz (50").

Dubner *et al.* 1996, AJ, 111, 1304. VLA at 330 MHz  $(48'' \times 75'' : S = 9.2 \pm 0.9 \text{ Jy})$  and 1.4 GHz  $(9'' \times 27'' : S = 5.8 \pm 0.1 \text{ Jy})$ . Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz

 $(43'' \times 61'' : S = 4.5 \text{ Jy}).$ 

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.
Koralesky et al. 1998, AJ, 116, 1323. VLA search for OH emission.

Reach et al. 2006, AJ, 131, 1479. Spitzer possible detection. Kothes & Dougherty 2007, A&A, 468, 993. SGPS at 1.4 GHz including HI.

## G341.2+0.9

**RA:**  $16^{\rm h}47^{\rm m}35^{\rm s}$  **1-GHz flux/Jy:** 1.5? **Size/arcmin:**  $22\times16$  **Dec:**  $-43^{\circ}47'$  **Spectral index:** 0.6? **Type:** C

**Radio:** Incomplete shell, with extension to SW. **Point sources:** Pulsar in W, with wind nebula.

#### References:

Frail *et al.* 1994, ApJ, 437, 781. VLA at 330 MHz  $(54'' \times 116'' : S = 3.0 \pm 0.1 \text{ Jy})$  and 1.4 GHz  $(21'' \times 25'' : S = 12.5 \pm 0.05 \text{ Jy})$ . Giacani *et al.* 2001, AJ, 121, 3133. VLA at 1.4 and 4.9 GHz (25'') of pulsar wind nebula.

## G341.9 - 0.3

RA: 16<sup>h</sup>55<sup>m</sup>01<sup>s</sup> 1-GHz flux/Jy: 2.5 Size/arcmin: 7 Dec: -44°01′ Spectral index: 0.5 Type: S

**Radio:** Incomplete shell, brightest to NE.

### References:

Caswell *et al.* 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3':S=7.4 Jy) and Parkes 64-m at 5 GHz (4':S=1.7 Jy).

Caswell *et al.* 1983, MNRAS, 203, 595. FIRST at 1415 MHz (50"), revision of previous flux densities.

Dubner *et al.* 1996, ÂJ, 111, 1304. VLA at 1.4 GHz  $(11'' \times 35'' : S = 2.2 \pm 0.1 \text{ Jy})$ .

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 62'' : S = 2.7 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.
Koralesky et al. 1998, AJ, 116, 1323. VLA search for OH emission.

## G342.0-0.2

**RA:**  $16^h54^m50^s$  **1-GHz flux/Jy:** 3.5? **Size/arcmin:**  $12\times9$  **Dec:**  $-43^\circ53'$  **Spectral index:** 0.4? **Type:** S

Radio: Distorted shell.

### References

Caswell *et al.* 1983, MNRAS, 203, 595. FIRST at 1415 MHz (50''), estimate  $S_{408 \text{ MHz}} = 5 \text{ Jy}$ ,  $S_{5 \text{ GHz}} = 2 \text{ Jy from previous maps}$ .

Dubner et al. 1996, AJ, 111, 1304. VLA at 1.4 GHz (11"×35").

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 62'' : S = 3.5 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.Koralesky et al. 1998, AJ, 116, 1323. VLA search for OH emission.

G342.1+0.9

**RA:** 16<sup>h</sup>50<sup>m</sup>43<sup>s</sup> 1-GHz flux/Jy: 0.5? Size/arcmin:  $10 \times 9$ **Dec:** -43°04′ **Spectral index:** ? Type: S

Radio: Incomplete shell.

References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 63'' : S = 0.6 \text{ Jy}).$ 

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

G343.0-6.0**RCW 114** 

RA: 17<sup>h</sup>25<sup>m</sup>00<sup>s</sup> Size/arcmin: 250 1-GHz flux/Jy: ? **Dec:**  $-46^{\circ}30'$ **Spectral index:** ? Type: S

Radio: Faint, poorly defined. **Optical:** Filamentary shell.

References:

Walker & Zealey 2001, MNRAS, 325, 287. Optical observations,

and review of earlier observations.

Welsh et al. 2003, A&A, 403, 605. Optical spectroscopy. Casandjian & Grenier 2008, A&A, 489, 849. γ-ray observations.

Kim et al. 2010, ApJ, 709, 823. UV observations.

G343.1-2.3

**RA:** 17<sup>h</sup>08<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jv:** 8? Size/arcmin: 32? **Dec:** -44°16′ **Spectral index:** 0.5? Type: C?

**Radio:** Incomplete shell? **X-ray:** Pulsar wind nebula.

Point sources: Pulsar near edge, with wind nebula.

References:

McAdam et al. 1993, Nature, 361, 516. MOST at 843 MHz (smoothed to 2').

Frail et al. 1994, ApJ, 437, 781. VLA at 330 MHz  $(56'' \times 111'' : S =$ 10.6 Jy) and 1.4 GHz (22"×27") near pulsar.

Becker et al. 1995, A&A, 298, 528. ROSAT of pulsar, and limit for remnant.

Giacani et al. 2001, AJ, 121, 3133. VLA at 1.4, 4.9 and 8.5 GHz (25") of pulsar wind nebula.

Gotthelf et al. 2002, ApJ, 567, L125. Pulsar x-ray detection. Dodson & Golap 2002, MNRAS, 334, L1, ATCA at 1.4 GHz  $(47'' \times 70'')$  including polarisation, and Chandra observations of pulsar wind nebula.

Aharonian et al. 2005, A&A, 432, L9. HESS limit. Abramowski et al. 2011, A&A, 528, A143. HESS observations.

G343.1 - 0.7

**RA:** 17<sup>h</sup>00<sup>m</sup>25<sup>s</sup> 1-GHz flux/Jy: 7.8 Size/arcmin: 27×21 **Dec:** -43°14′ Spectral index: 0.55 Type: S

Radio: Shell, with smaller thermal shell adjacent.

References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 63'' : S8.5 \pm 0.6)$ , plus Parkes 64-m at 4.5 GHz (S =  $3.9\pm0.6 \text{ Jy}$ ) and 8.55 GHz ( $S=2.4\pm0.5 \text{ Jy}$ )

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations. Koralesky et al. 1998, AJ, 116, 1323. VLA search for OH emission.

## G344.7 - 0.1

**RA:** 17<sup>h</sup>03<sup>m</sup>51<sup>s</sup> **1-GHz flux/Jy:** 2.5? **Dec:** -41°42′ **Spectral index:** 0.3? **Size/arcmin:** 8 **Type:** C?

**Radio:** Asymmetric shell, with possible core.

X-ray: Detected.

**Distance:** Hi absorption and association with features suggests 6.3 kpc.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 4.7 Jy) and Parkes 64-m at 5 GHz (4': S = 1.3 Jy).

Dubner *et al.* 1993, AJ, 105, 2251. VLA at 1.47 GHz ( $30'' \times 43''$ :  $S = 1.7 \pm 0.1$  Jy).

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×65": S=2.5 Jy).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations. Sugizaki et al. 2001, ApJS, 134, 77. ASCA survey observations.

Yamauchi et al. 2005, PASJ, 57, 459. ASCA observations. Reach et al. 2006, AJ, 131, 1479. Spitzer observations. Giacani et al. 2011, A&A, 531, A138. VLA and ATCA at 1.4 GHz  $(5''6\times8'')$ , ATCA at 5 GHz  $(10''\times13'')$  and 8.4 GHz  $(5''\times10'')$ , plus XMM-Newton observations and HI from SGPS. Andersen et al. 2011, ApJ, 742, 7. Spitzer observations. Yamaguchi et al. 2012, ApJ, 749, 137. Suzaku observations. Yang et al. 2013, ApJ, 766, 44. Suzaku spectroscopy.

## G345.7 - 0.2

**RA:**  $17^h07^m20^s$  **1-GHz flux/Jy:** 0.6? **Size/arcmin:** 6 **Dec:**  $-40^{\circ}53'$  **Spectral index:** ? **Type:** S

**Radio:** Poorly defined diffuse shell. **Point sources:** Old pulsar nearby.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×66": S=0.7 Jy).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

### G346.6-0.2

**RA:**  $17^h10^m19^s$  **1-GHz flux/Jy:** 8? **Size/arcmin:** 8 **Dec:**  $-40^\circ11'$  **Spectral index:** 0.5? **Type:** S

Radio: Irregular shell.

**X-ray:** Centrally brightened, clumpy.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 14.9 Jy) and Parkes 64-m at 5 GHz (4': S = 4.3 Jy).

Dubner *et al.* 1993, AJ, 105, 2251. VLA at 1.47 GHz ( $31'' \times 43''$ :  $S = 8.1 \pm 0.9$  Jy).

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 67'' : S = 8.7 \text{ Jy})$ .

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Koralesky *et al.* 1998, AJ, 116, 1323. VLA detection of compact OH emission.

Reach et al. 2006, AJ, 131, 1479. Spitzer observations.
Hewitt et al. 2009, ApJ, 694, 1266. Spitzer spectroscopy.
Sezer et al. 2011, MNRAS, 415, 301. Suzaku observations.
Andersen et al. 2011, ApJ, 742, 7. Spitzer observations.
Yamauchi et al. 2013, PASJ, 65, 6. Suzaku observations.
Pihlström et al. 2014, AJ, 147, 73. VLA search for methanol masers.

Pannuti *et al.* 2014, AJ, 147, 55. ASCA observations. Auchettl *et al.* 2017, ApJ, 847, 121. XMM-Newton observations. **G347.3–0.5** RX J1713.7–3946

**RA:**  $17^h 13^m 50^s$  **1-GHz flux/Jy:** 30? **Size/arcmin:**  $65 \times 55$  **Dec:**  $-39^{\circ}45'$  **Spectral index:** ? **Type:** S?

Radio: Faint emission.

**X-ray:** Non-thermal, limb-brightened to W, with central source.

**Point sources:** Central X-ray source.

**Distance:** Association with molecular clouds and X-ray observations imply 1.3 kpc.

#### References:

Koyama *et al.* 1997, PASJ, 49, L7. ASCA of NW. Slane *et al.* 1999, ApJ, 525, 357. ASCA and other observations. Muraishi *et al.* 2000, A&A, 354, L57. γ-ray detection.

Butt *et al.* 2001, ApJ, 562, L167. Associated  $\gamma$ -ray emission. Ellison *et al.* 2001, ApJ, 563, 191. ATCA at 1.4 GHz ( $36'' \times 46''$ ), and ASCA observations.

Enomoto *et al.* 2002, Nature, 416, 823. γ-ray observations.
Uchiyama *et al.* 2002, PASJ, 54, L73. ASCA observations.
Uchiyama *et al.* 2003, A&A, 400, 567. Chandra spectroscopy.
Pannuti *et al.* 2003, ApJ, 593, 377. ROSAT, ASCA and RXTE observations.

Lazendic et al. 2003, ApJ, 593, L27. Chandra, XMM-Newton and other X-ray observations of central source.

Fukui et al. 2003, PASJ, 55, L61. CO observations of surroundings.

Cassam-Chenaï *et al.* 2004, A&A, 427, 199. XMM-Newton and other observations.

Aharonian *et al.* 2004, Nature, 432, 75. HESS detection. Lazendic *et al.* 2004, ApJ, 602, 271. Chandra observations of parts, and ATCA at 1.4 GHz (36"×46").

Hiraga et al. 2005, A&A, 431, 953. XMM-Newton observations, Moriguchi et al. 2005, ApJ, 631, 947. CO observations of surroundings.

Aharonian et al. 2007, A&A, 464, 235. HESS observations. see also: Aharonian et al. 2011, A&A, 531, C1. Erratum.

Uchiyama et al. 2007, Nature, 449, 576. Chandra multi-epoch observations for study of small-scale variability.

Tanaka et al. 2008, ApJ, 685, 988. Suzaku observations.Takahashi et al. 2008, PASJ, 60, S131. Suzaku observations of SW

Mignani et al. 2008, A&A, 484, 457. Optical and IR observations of central source.

Acero *et al.* 2009, A&A, 505, 157. XMM-Newton observations, plus other radio, IR and HESS observations.

Sano et al. 2010, ApJ, 724, 59. CO observations of region.

Maxted et al. 2012, MNRAS, 422, 2230. Molecular line observations

see also: Maxted et al. 2013, MNRAS, 430, 2511. Erratum. Sano et al. 2013, ApJ, 778, 59. CO, SGPS HI and Suzaku observations

Federici *et al.* 2015, A&A, 577, A12. Fermi detection. Sano *et al.* 2015, ApJ, 799, 175. Suzaku observations.

Katsuda *et al.* 2015, ApJ, 814, 29. XMM-Newton observations. Acero *et al.* 2016, ApJS, 224, 8. Fermi observations.

Tsuji & Uchiyama 2016, PASJ, 68, 108. Chandra observations of NW.

Acero et al. 2017, A&A, 597, A106. XMM-Newton observations for proper motion studies.

Okuno et al. 2018, PASJ, 70, 77. Chandra observations. Abdalla et al. 2018, A&A, 612, A6. HESS observations.

## G348.5 - 0.0

**RA:** 17<sup>h</sup>15<sup>m</sup>26<sup>s</sup> **1-GHz flux/Jy:** 10? **Size/arcmin:** 10? **Dec:** -38°28′ **Spectral index:** 0.4? **Type:** S?

**Radio:** Arc, overlapping G348.5+0.1.

### References:

Kassim *et al.* 1991, ApJ, 374, 212. VLA at 333 MHz  $(46'' \times 53'')$ , 1.4 GHz  $(18'' \times 33'')$  and part at 5 GHz  $(2''.5 \times 3''.9)$ .

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 69'' : S = 10.2 \text{ Jy})$ .

Frail et al. 1996, AJ, 111, 1651. OH emission near remnant, including masers.

Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Koralesky *et al.* 1998, AJ, 116, 1323. VLA search for OH emission

Reach *et al.* 2006, AJ, 131, 1479. Spitzer observations. Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers.

Hewitt et al. 2009, ApJ, 694, 1266. Spitzer spectroscopy.
Andersen et al. 2011, ApJ, 742, 7. Spitzer observations.
Pihlström et al. 2014, AJ, 147, 73. VLA search for methanol masers.

Pannuti et al. 2014, AJ, 147, 55. XMM-Newton upper limit.

G348.5+0.1 CTB 37A

RA: 17<sup>h</sup>14<sup>m</sup>06<sup>s</sup> 1-GHz flux/Jy: 72 Size/arcmin: 15 Dec: -38°32′ Spectral index: 0.3 Type: S

**Radio:** Shell, poorly define to S and W, overlapping G348.5–0.0 in E.

X-ray: Brighter to W.

**Distance:** Hi absorption indicates 8.0 kpc.

### References:

Clark *et al.* 1975, AuJPA, 37, 75. Molonglo at 408 MHz (3': S = 97 Jy)

Milne & Dickel 1975, AuJPh, 28, 209. Parkes 64-m at 5 GHz (4/4:S=43 Jy).

Caswell et al. 1975, A&A, 45, 239. Parkes HI absorption.

Milne *et al.* 1979, MNRAS, 188, 437. FIRST at 1415 MHz (0'.8: S > 50) and Parkes 64-m at 14.7 GHz (2'.2:  $S = 18 \pm 5$  Jy).

Downes 1984, MNRAS, 210, 845. VLA at 1465 MHz ( $20'' \times 45''$ ). Kassim *et al.* 1991, ApJ, 374, 212. VLA at 333 MHz ( $46'' \times 53''$ ), 1.4 GHz ( $18'' \times 33''$ ) and part at 5 GHz ( $2''.5 \times 3''$ 9).

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×69": S=71 Jy).

Brogan et al. 2000, ApJ, 537, 875. VLA at 1.7 GHz for OH Zeeman splitting.

Reynoso & Mangum 2000, ApJ, 545, 874. CO observations.
Reach et al. 2006, AJ, 131, 1479. Spitzer observations.
Aharonian et al. 2008, A&A, 490, 685. HESS observations.
Castro & Slane 2010, ApJ, 717, 372. Fermi observations.
Andersen et al. 2011, ApJ, 742, 7. Spitzer observations.
Sezer et al. 2011, MNRAS, 417, 1387. Suzaku observations.
Tian & Leahy 2012, MNRAS, 421, 2593. SGPS Hi absorption observations.

Maxted *et al.* 2013, MNRAS, 434, 2188. Molecular line observations of region.

Pannuti et al. 2014, AJ, 147, 55. XMM-Newton and Chandra observations.

Yamauchi et al. 2014, PASJ, 66, 2. Suzaku observations. Acero et al. 2016, ApJS, 224, 8. Fermi observations.

G348.7+0.3

**RA:** 17<sup>h</sup>13<sup>m</sup>55<sup>s</sup> **1-GHz flux/Jy:** 26 **Size/arcmin:** 17? **Dec:** -38°11′ **Spectral index:** 0.3 **Type:** S

Radio: Incomplete shell with faint eastern extensions.

**X-ray:** Diffuse emission. **Point sources:** X-ray pulsar.

**Distance:** Hi absorption suggests 13 kpc.

### References:

Clark *et al.* 1975, AuJPA, 37, 75. Molonglo at 408 MHz (3' : S = 34 Jy).

Milne & Dickel 1975, AuJPh, 28, 209. Parkes 64-m at 5 GHz (4'.4: S=32 Jy).

Caswell *et al.* 1975, A&A, 45, 239. Parkes Hi absorption. Milne *et al.* 1979, MNRAS, 188, 437. FIRST at 1415 MHz (0.8:

S > 20) and Parkes 64-m at 14.7 GHz (2'.2: $S = 8 \pm 3$  Jy). Downes 1984, MNRAS, 210, 845. VLA at 1465 MHz (20"×45").

Kassim *et al.* 1991, ApJ, 374, 212. VLA at 333 MHz (46"×53"). Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×69": S=33 Jy).

Frail et al. 1996, AJ, 111, 1651. OH emission near remnant.

Aharonian et al. 2006, ApJ, 636, 777. HESS detection.

Aharonian et al. 2008, A&A, 486, 829. HESS and Chandra observations.

Nakamura et al. 2009, PASJ, 61, S197. Suzaku and Chandra observations.

Halpern & Gotthelf 2010, ApJ, 710, 941. Chandra detection of pul-

Halpern & Gotthelf 2010, ApJ, 725, 1384. Chandra observations of pulsar.

Tian & Leahy 2012, MNRAS, 421, 2593. SGPS Hi observations. Xin *et al.* 2016, ApJ, 817, 64. Fermi observations.

G349.2 - 0.1

**RA:**  $17^h17^m15^s$  **1-GHz flux/Jy:** 1.4? **Size/arcmin:**  $9\times6$  **Dec:**  $-38^{\circ}04'$  **Spectral index:** ? **Type:** S

Radio: Elongated shell, adjacent to bright HII region.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 70'' : S = 1.6 \text{ Jy})$ .

Frail et al. 1996, AJ, 111, 1651. OH emission near remnant.

## G349.7 + 0.2

**RA**: 17<sup>h</sup>17<sup>m</sup>59<sup>s</sup> 1-GHz flux/Jy: 20 Size/arcmin:  $2.5 \times 2$ **Dec:**  $-37^{\circ}26'$ Spectral index: 0.5 Type: S

Radio: Incomplete clumpy shell, with enhancement to the S.

**X-ray:** Irregular shell, brighter to S and E. **Distance:** HI absorption indicates 11.5 kpc.

### References:

Becker & Kundu 1975, AJ, 80, 679. NRAO 140-ft at 10.6 GHz

Caswell et al. 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3': S = 31.0 Jy) and Parkes 64-m at 5 GHz (4':S = 9.1 Jy), no maps. Caswell et al. 1975, A&A, 45, 239. Parkes HI absorption.

Shaver et al. 1985, Nature, 313, 113. VLA at 1.4 GHz  $(3''.4 \times 14''.5)$ .

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 71'' : S = 22 \text{ Jy}).$ 

Frail et al. 1996, AJ, 111, 1651. OH emission near remnant, including masers

Brogan et al. 2000, ApJ, 537, 875. VLA at 1.7 GHz for OH Zeeman splitting.

Reynoso & Mangum 2001, AJ, 121, 347. CO observations of the vicinity.

Sugizaki et al. 2001, ApJS, 134, 77. ASCA survey observations. Slane et al. 2002, ApJ, 580, 904. ASCA observations.

Dubner et al. 2004, A&A, 426, 201. CO observations of surround-

Lazendic et al. 2005, ApJ, 618, 733. Chandra observations.

Reach et al. 2006, AJ, 131, 1479. Spitzer observations. Hewitt et al. 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH

Hewitt et al. 2009, ApJ, 694, 1266. Spitzer spectroscopy. Castro & Slane 2010, ApJ, 717, 372. Fermi observations.

Lazendic et al. 2010, MNRAS, 409, 371. OH, CO and other molecular line observations of region.

Andersen et al. 2011, ApJ, 742, 7. Spitzer observations. Pihlström et al. 2014, AJ, 147, 73. VLA search for methanol masers

Tian & Leahy 2014, ApJ, 783, L2. SGPS HI observations. Yasumi et al. 2014, PASJ, 66, 68. Suzaku observations. Abramowski et al. 2015, A&A, 574, A100. HESS observations. see also: Abramowski et al. 2015, A&A, 580, C1. Corrigendum.

Ergin et al. 2015, ApJ, 804, 124. Fermi and Suzaku observations. Rho et al. 2015, ApJ, 812, 44. Herschel observations. Koo et al. 2016, ApJ, 821, 20. Spitzer and Herschel flux densities

(and comparison with X-ray properties).

Acero et al. 2016, ApJS, 224, 8. Fermi observations. Abdalla et al. 2018, A&A, 612, A1. HESS observations.

## G350.0-2.0

**RA:** 17<sup>h</sup>27<sup>m</sup>50<sup>s</sup> Size/arcmin: 45 1-GHz flux/Jy: 26 **Dec:**  $-38^{\circ}32'$ Spectral index: 0.4 Type: S

Incorporates the previously catalogued G350.0-1.8 in the NW.

Radio: Shell, brightest in NW.

Optical: Detected. **X-ray:** Brighter to NW.

### References:

Clark et al. 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 49.5 Jy) and Parkes 64-m at 5 GHz (4': S = 13.6 Jy). Milne & Dickel 1975, AuJPh, 28, 209. Parkes 64-m at 5 GHz

(4'.4).

Gaensler 1998, ApJ, 493, 781. VLA and Parkes 64-m at 1.4 GHz  $(18'' \times 21'' : S = 22.3 \pm 0.3 \text{ Jy})$ , clarifying extent of remnant. Stupar & Parker 2011, MNRAS, 414, 2282.  $H\alpha$  observations. Karpova et al. 2016, MNRAS, 462, 3845. XMM-Newton observations.

## G350.1 - 0.3

**RA:** 17<sup>h</sup>21<sup>m</sup>05<sup>s</sup> **1-GHz flux/Jy:** 6? Size/arcmin: 4? **Dec:**  $-37^{\circ}27'$ **Spectral index:** 0.8? Type: ?

Radio: Several clumps of emission.

**X-ray:** Diffuse emission, with compact source.

**Point sources:** X-ray source.

**Distance:** HI absorption indicates 4.5 to 10.7 kpc, possible interaction with molecular cloud indicates 4.5 kpc.

Clark et al. 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 10.7) and Parkes 64-m at 5 GHz (4':S=1.7).

Salter *et al.* 1986, A&A, 162, 217. VLA at 1.5 ( $4'' \times 10''.4$ ) and 4.8 GHz ( $15'' \times 35''$ ) and 15 GHz ( $2'' \times 5''$ ).

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×71")...

Sugizaki et al. 2001, ApJS, 134, 77. ASCA survey observations. Gaensler et al. 2008, ApJ, 680, L37. VLA at 4.8 GHz (5".6×11".4) plus XMM-Newton observations.

Lovchinsky et al. 2011, ApJ, 731, 70. Chandra and Spitzer observa-

Yasumi et al. 2014, PASJ, 66, 68. Suzaku observations.

## G351.0-5.4

RA: 17<sup>h</sup>46<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: ? Size/arcmin: 30 Dec: -39°25′ Spectral index: ? Type: S

Radio: Faint shell, brighter to E and W.

#### References:

de Gasperin *et al.* 2014, A&A, 568, A107. GMRT at 325 MHz  $(89'' \times 141'')$ , plus other observations.

## G351.2+0.1

**RA:** 17<sup>h</sup>22<sup>m</sup>27<sup>s</sup> **1-GHz flux/Jy:** 5? **Dec:** -36°11′ **Spectral index:** 0.4 **Size/arcmin:** 7 **Type:** C?

Has been called G351.3+0.2.

**Radio:** Distorted shell, with possible flat-spectrum core.

#### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 8.1 Jy) and Parkes 64-m at 5 GHz (4': S = 3.1 Jy).

Becker & Helfand 1988, AJ, 95, 883. VLA at 5 GHz (15"), and at 15 GHz of core.

Dubner *et al.* 1993, AJ, 105, 2251. VLA at 1.47 GHz  $(32'' \times 36'')$ :  $S=4.8\pm0.2$  Jy).

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×73": S=5.5 Jy).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

## G351.7 + 0.8

**RA:** 17<sup>h</sup>21<sup>m</sup>00<sup>s</sup> **1-GHz flux/Jy:** 10 **Size/arcmin:** 18×14 **Dec:** -35°27′ **Spectral index:** 0.5? **Type:** S

Radio: Elongated shell, adjacent to bright HII region.

Point sources: Pulsar nearby.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×74": S=11 Jy).

Tian et al. 2007, MNRAS, 378, 1283. SGPS at 1.4 GHz ( $100'': S = 8.4 \pm 0.7$  Jy) including HI.

## G351.9-0.9

**RA:**  $17^h28^m52^s$  **1-GHz flux/Jy:** 1.8? **Size/arcmin:**  $12\times9$  **Dec:**  $-36^{\circ}16'$  **Spectral index:** ? **Type:** S

Radio: Asymmetric shell.

### References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 73'' : S = 2.0 \text{ Jy})$ .

## G352.7 - 0.1

**RA:**  $17^h27^m40^s$  **1-GHz flux/Jy:** 4 **Size/arcmin:**  $8\times6$  **Dec:**  $-35^\circ07'$  **Spectral index:** 0.6 **Type:** S

Radio: Distorted shell.X-ray: Detected.

**Distance:** HI absorption indicates 6.8 to 8.4 kpc.

### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 9.6 Jy) and Parkes 64-m at 5 GHz (4': S = 2.3 Jy).

Caswell *et al.* 1983, MNRAS, 203, 595. FIRST at 1415 MHz (1/1).

Dubner *et al.* 1993, AJ, 105, 2251. VLA at 1.47 GHz ( $34'': S = 3.4 \pm 0.4$  Jy).

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz (43"×75": S=4.4 Jy).

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.

Kinugasa et al. 1998, PASJ, 50, 249. ASCA observations.Koralesky et al. 1998, AJ, 116, 1323. VLA search for OH emission.

Sugizaki *et al.* 2001, ApJS, 134, 77. ASCA survey observations. Giacani *et al.* 2009, A&A, 507, 841. VLA at 4.8 GHz (9"×12"), plus HI and XMM-Newton observations.

Pannuti et al. 2014, ApJ, 782, 102. XMM-Newton and Chandra observations.

Sezer & Gök 2014, ApJ, 790, 81. Suzaku observations.

## G353.6 - 0.7

**RA:**  $17^h32^m00^s$  **1-GHz flux/Jy:** 2.5? **Size/arcmin:** 30 **Dec:**  $-34^\circ44'$  **Spectral index:** ? **Type:** S

Has erroneously been called G353.6–0.37.

Radio: Shell, brighter to S.

**X-ray:** Patchy shell, brighter to E. **Point sources:** Central X-ray source.

**Distance:** Various observations suggest 3.2 kpc.

### References:

Tian et al. 2008, ApJ, 679, L85. SGPS at 1.4 GHz (100''), plus 843 MHz (43'') and X-ray observations.

Halpern & Gotthelf 2010, ApJ, 710, 941. XMM-Newton observations.

Tian et al. 2010, ApJ, 712, 790. XMM-Newton, Suzaku and CO observations.

Abramowski *et al.* 2011, A&A, 531, A81. HESS observations. Bamba *et al.* 2012, ApJ, 756, 149. Suzaku observations. *see also:* Bamba *et al.* 2012, ApJ, 761, 80. Erratum.

Klochkov *et al.* 2013, A&A, 556, A41. XMM-Newton and other X-ray observations.

Fukuda et al. 2014, ApJ, 788, 94. HESS, SGPS HI and CO observations.

Nayana *et al.* 2017, MNRAS, 467, 155. GMRT at 325 MHz (97"×135") and 610 MHz (105"×150"), plus other observations

de Wilt *et al.* 2017, MNRAS, 468, 2093. Molecular line observations of region.

Doroshenko *et al.* 2017, A&A, 608, A23. XMM-Newton observations.

Maxted et al. 2018, MNRAS, 474, 662. CO and HI observations.

Condon *et al.* 2017, ApJ, 851, 100. Fermi observations. Guo *et al.* 2018, ApJ, 853, 2. Fermi observations. Abdalla *et al.* 2018, A&A, 612, A1. HESS observations.

### G353.9 - 2.0

**RA:** 17<sup>h</sup>38<sup>m</sup>55<sup>s</sup> **1-GHz flux/Jy:** 1? **Size/arcmin:** 13 **Dec:** -35°11′ **Spectral index:** 0.5? **Type:** S

Radio: Shell, with central double source.

## References:

Green 2001, MNRAS, 326, 283. VLA at 327 MHz  $(2'.7 \times 3'.0)$  and 1.4 GHz  $(36'' \times 42'')$ , plus 8.4 GHz  $(6''.1 \times 8''.4)$  of central source only.

G354.1+0.1

RA: 17<sup>h</sup>30<sup>m</sup>28<sup>s</sup>
1-GHz flux/Jy: ?
Size/arcmin: 15×3?
Dec: -33°46′
Spectral index: varies
Type: C?

Is this a SNR?

**Radio:** Elongated N–S. **Point sources:** Pulsar at S tip.

References:

Frail *et al.* 1994, ApJ, 437, 781. VLA at 330 MHz (47"×99") and 1.4 GHz (8"8×21"). Ajello *et al.* 2016, ApJ, 819, 44. Fermi observations.

G354.8 - 0.8

RA: 17<sup>h</sup>36<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: 2.8? Size/arcmin: 19 Dec: -33°42′ Spectral index: ? Type: S

Radio: Distorted shell.

References:

Whiteoak & Green 1996, A&AS, 118, 329. MOST at 843 MHz  $(43'' \times 78'' : S = 3.1 \text{ Jy})$ .

Green et al. 1997, AJ, 114, 2058. Parkes 64-m OH observations.
Koralesky et al. 1998, AJ, 116, 1323. VLA search for OH emission.

G355.4+0.7

RA: 17<sup>h</sup>31<sup>m</sup>20<sup>s</sup> 1-GHz flux/Jy: 5? Size/arcmin: 25 Dec: -32°26′ Spectral index: ? Type: S

Radio: Faint, incomplete shell.

References:

Gray 1994, MNRAS, 270, 835. MOST at 843 MHz ( $43'' \times 77''$ ). Roy & Bhatnagar 2006, JPhCS, 54, 152. GMRT at 330 MHz (1':  $S = 8.9 \pm 1.3$  Jy).

Ajello et al. 2016, ApJ, 819, 44. Fermi observations. Acero et al. 2016, ApJS, 224, 8. Fermi observations.

G355.6 - 0.0

RA: 17<sup>h</sup>35<sup>m</sup>16<sup>s</sup> 1-GHz flux/Jy: 3? Size/arcmin: 8×6 Dec: -32°38′ Spectral index: ? Type: S

Radio: Well defined shell.X-ray: Centrally brightened.

References:

Gray 1994, MNRAS, 270, 847. MOST at 843 MHz  $(43'' \times 80'' : S = 2.6 \text{ Jy})$ 

Sugizaki *et al.* 2001, ApJS, 134, 77. ASCA survey observations. Roy & Bhatnagar 2006, JPhCS, 54, 152. GMRT at 330 MHz (1':  $S=3.3\pm0.5$  Jy).

Marquez-Lopez & Phillips 2010, MNRAS, 407, 94. Mid-IR observations.

Minami et al. 2013, PASJ, 65, 99. Suzaku observations.

## G355.9 - 2.5

RA: 17<sup>h</sup>45<sup>m</sup>53<sup>s</sup>
1-GHz flux/Jy: 8
Size/arcmin: 13
Dec: -33°43′
Spectral index: 0.5
Type: S

Radio: Distorted shell, brightest to SE.

#### References:

Clark *et al.* 1975, AuJPA, 37, 1. Molonglo at 408 MHz (3': S = 12.3 Jy) and Parkes 64-m at 5 GHz (4': S = 3.4 Jy).

Dubner *et al.* 1993, AJ, 105, 2251. VLA at 1.47 GHz ( $32'' \times 34''$ :  $S = 5.0 \pm 0.3$  Jy).

Gray 1994, MNRAS, 270, 835. MOST at 843 MHz (43"×77"). Green *et al.* 1997, AJ, 114, 2058. Parkes 64-m OH observations. Marquez-Lopez & Phillips 2010, MNRAS, 407, 94. Mid-IR observations.

## G356.2 + 4.5

RA: 17<sup>h</sup>19<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: 4 Size/arcmin: 25 Dec: -29°40′ Spectral index: 0.7 Type: S

Has been called G356.2+4.4.

Radio: Faint shell.

### References:

Duncan et al. 1995, MNRAS, 277, 36. Parkes 64-m at 2.4-GHz (10'.4).

Bhatnagar 2000, MNRAS, 317, 453. GMRT at 327 MHz (1'.5×3':  $S=8.1\pm1.7$  Jy), and NVSS at 1.4 GHz.

## G356.3-1.5

**RA:**  $17^{\text{h}}42^{\text{m}}35^{\text{s}}$  **1-GHz flux/Jy:** 3? **Size/arcmin:**  $20 \times 15$  **Dec:**  $-32^{\circ}52'$  **Spectral index:** ? **Type:** S

Radio: Double arc.

### References:

Gray 1994, MNRAS, 270, 847. MOST at 843 MHz  $(43'' \times 79'' : S = 2.8 \text{ Jy})$ .

Bhatnagar 2002, MNRAS, 332, 1. GMRT at 327 MHz (0'.8×1'.7:  $S = 5.7 \pm 0.2$ ).

## G356.3-0.3

**RA:**  $17^h37^m56^s$  **1-GHz flux/Jy:** 3? **Size/arcmin:**  $11\times7$  **Dec:**  $-32^{\circ}16'$  **Spectral index:** ? **Type:** S

Has been suggested this part of a larger SNR.

Radio: Diffuse emission.

### References:

Gray 1994, MNRAS, 270, 847. MOST at 843 MHz  $(43'' \times 81'' : S = 2.6 \text{ Jy})$ 

Roy & Pramesh Rao 2002, MNRAS, 329, 775. GMRT at 330 MHz (2'.7×4'.8).

Acero et al. 2016, ApJS, 224, 8. Fermi observations.

G357.7-0.1 MSH 17-39

**RA:**  $17^{\text{h}}40^{\text{m}}29^{\text{s}}$  **1-GHz flux/Jy:** 37 **Size/arcmin:**  $8\times3$ ? **Dec:**  $-30^{\circ}58'$  **Spectral index:** 0.4 **Type:** ?

Has been suggested that this is not a SNR.

**Radio:** Multiple arcs and filaments, brighter to NW 'head'.

X-ray: Detected from NW 'head', and SW 'tail'.

**Distance:** Hi absorption suggests beyond Galactic Centre.

#### References:

Caswell *et al.* 1975, AuJPA, 37, 39. Molonglo at 408 MHz (3':S = 54.2 Jy) and Parkes 64-m at 5 GHz (4':S = 18.5 Jy).

Milne & Dickel 1975, AuJPh, 28, 209. Parkes 64-m at 5 GHz (4'.4: S = 14.6 Jy).

Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2.6).

Caswell *et al.* 1980, MNRAS, 190, 881. FIRST at 1415 MHz (50").

Weiler & Panagia 1980, A&A, 90, 269. Effelsberg 100-m at 9 GHz (1'.5) (private communication from Baker).

Shaver *et al.* 1985, Nature, 313, 113. VLA at 1.4 GHz  $(3''8\times10''9)$  and 5 GHz  $(12''\times26'')$ .

Becker & Helfand 1985, Nature, 313, 115. VLA at 1.4 GHz and 5 GHz.

Helfand & Becker 1985, Nature, 313, 118. Suggest it is not a SNR

Shaver et al. 1985, A&A, 147, L23. Observations of peripheral compact source

Caswell *et al.* 1989, PASA, 8, 184. MOST at 843 MHz (43"×83").

Gray 1994, MNRAS, 270, 835. MOST at 843 MHz (43"×84"). Stewart *et al.* 1994, ApJ, 432, L39. ATCA at 4.79 and 5.84 GHz (12"×22") and Effelsberg 100-m at 10.6 GHz (1'), including polarisation.

Frail et al. 1996, AJ, 111, 1651. OH emission near remnant, including masers.

Yusef-Zadeh et al. 1999, ApJ, 527, 172. VLA of nearby OH masers.

LaRosa *et al.* 2000, AJ, 119, 207. VLA at 333 MHz (24"×43"). *see also:* LaRosa *et al.* 2000, AJ, 119, 3145. Erratum.

Brogan *et al.* 2000, ApJ, 537, 875. VLA at 1.7 GHz for OH Zeeman splitting.

Brogan & Goss 2003, AJ, 125, 272. VLA at 1.4 GHz (11".4×13".6), including HI, and 8.3 GHz (6".8×14".3) recombination line observation of HII region.

Gaensler et al. 2003, ApJ, 594, L35. Chandra detection.

Lazendic *et al.* 2003, AN, 324 (No S1), 157. Molecular line observations.

Burton *et al.* 2004, MNRAS, 348, 638. IR and radio observations of H<sub>II</sub> region.

Lazendic *et al.* 2004, MNRAS, 354, 393. IR and molecular line observations.

Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers.

Phillips et al. 2009, MNRAS, 397, 1215. Observations of interactions with surroundings.

Castro et al. 2013, ApJ, 774, 36. Fermi observations.

Pihlström *et al.* 2014, AJ, 147, 73. VLA search for methanol masers.

Acero *et al.* 2016, ApJS, 224, 8. Fermi observations. Qiao *et al.* 2018, ApJS, 239, 15. OH maser observations.

## G357.7 + 0.3

RA: 17<sup>h</sup>38<sup>m</sup>35<sup>s</sup> 1-GHz flux/Jy: 10 Size/arcmin: 24 Dec: -30°44′ Spectral index: 0.4? Type: S

Radio: Non-thermal shell in complex region.

### References

Reich & Fürst 1984, A&AS, 57, 165. Effelsberg 100-m at 2.7 GHz  $(4'.3:S=7\pm1.5 \text{ Jy})$ ,  $S_{5 \text{ GHz}}=5.5\pm1.5 \text{ Jy}$  from surveys.

Gray 1994, MNRAS, 270, 835. MOST at 843 MHz (43"×84"). Yusef-Zadeh *et al.* 1999, ApJ, 527, 172. VLA of nearby OH masers.

Hewitt et al. 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers. Phillips & Marquez-Lugo 2010, MNRAS, 409, 701. Spitzer observations of region.

Pihlström *et al.* 2014, AJ, 147, 73. VLA search for methanol masers.

Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. Rho *et al.* 2017, ApJ, 834, 12. Molecular line observations.

## G358.0 + 3.8

RA: 17<sup>h</sup>26<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: 1.5? Size/arcmin: 38 Dec: -28°36′ Spectral index: ? Type: S

Radio: Faint shell.

### References:

Duncan *et al.* 1995, MNRAS, 277, 36. Parkes 64-m at 2.4-GHz (10/4).

Bhatnagar 2000, MNRAS, 317, 453. GMRT at 327 MHz (1'.3×2'.2:  $S=2.5\pm1.3$  Jy), and NVSS at 1.4 GHz.

## G358.1+1.0

RA: 17<sup>h</sup>37<sup>m</sup>00<sup>s</sup> 1-GHz flux/Jy: 2? Size/arcmin: 20 Dec: -29°59′ Spectral index: ? Type: S

Was erroneously called G358.1+0.1.

Radio: Faint shell.

### References:

Gray 1994, MNRAS, 270, 847. MOST at 843 MHz ( $43'' \times 77''$ ). Roy & Bhatnagar 2006, JPhCS, 54, 152. GMRT at 330 MHz (1':  $S=6.0\pm2.5$  Jy).

G358.5 - 0.9

**RA:**  $17^{\text{h}}46^{\text{m}}10^{\text{s}}$  **1-GHz flux/Jy:** 4? **Size/arcmin:** 17 **Dec:**  $-30^{\circ}40'$  **Spectral index:** ? **Type:** S

Radio: Shell, brighter to NE.

### References:

Gray 1994, MNRAS, 270, 835. MOST at 843 MHz ( $43'' \times 77''$ ). Roy & Bhatnagar 2006, JPhCS, 54, 152. GMRT at 330 MHz (1':  $S = 8.0 \pm 2.5$  Jy).

G359.0 - 0.9

RA: 17<sup>h</sup>46<sup>m</sup>50<sup>s</sup> 1-GHz flux/Jy: 23 Size/arcmin: 23 Dec: -30°16′ Spectral index: 0.5 Type: S

Radio: Incomplete shell.

Optical: Detected.

X-ray: Partial shell.

### References:

Reich *et al.* 1988, IAUCo, 101, 293. Summary of parameters. Reich *et al.* 1990, A&AS, 85, 633. Effelsberg 100-m at 2.7 GHz (4'.3).

Gray 1994, MNRAS, 270, 835. MOST at 843 MHz (43"×86"). LaRosa *et al.* 2000, AJ, 119, 207. VLA at 333 MHz (24"×43"). *see also*: LaRosa *et al.* 2000, AJ, 119, 3145. Erratum. Bamba *et al.* 2000, PASJ, 52, 259. ASCA observations. Yusef-Zadeh *et al.* 2004, ApJS, 155, 421. VLA at 1.4 GHz (8".4×12".8) of part.

Bamba *et al.* 2009, ApJ, 691, 1854. Suzaku observations. Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations. Ponti *et al.* 2015, MNRAS, 453, 172. XMM-Newton observations. Froebrich *et al.* 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations. de Wilt *et al.* 2017, MNRAS, 468, 2093. Molecular line observations of region.

## G359.1 - 0.5

**RA:** 17<sup>h</sup>45<sup>m</sup>30<sup>s</sup> **1-GHz flux/Jy:** 14 **Size/arcmin:** 24 **Dec:** -29°57′ **Spectral index:** 0.4? **Type:** S

Has been called G359.10–0.5.

**Radio:** Non-thermal shell in complex region, crossed by the 'snake'.

Optical: Detected.

**X-ray:** Centrally brightened.

**Point sources:** Several compact radio sources near centre, OH masers around edge.

#### References:

Downes *et al.* 1979, A&AS, 35, 1. From observations by Altenhoff *et al.* 1979, A&AS, 35, 23. Effelsberg 100-m at 4.9 GHz (2'.6: *S* = 13 Jy).

Reich & Fürst 1984, A&AS, 57, 165. Effelsberg 100-m at 2.7 GHz  $(4.3:S=10\pm1.5 \text{ Jy})$  and 4.8 GHz  $(2.4:S=8.1\pm0.5 \text{ Jy})$ .

Uchida *et al.* 1992, ApJ, 398, 128. VLA at 1.5 GHz (10"×11"), and observations of nearby molecular material.

Uchida et al. 1992, AJ, 104, 1533. VLA at 1.4 GHz.

Gray 1994, MNRAS, 270, 835. MOST at 843 MHz (43"×85"). Yusef-Zadeh *et al.* 1995, Science, 270, 1801. VLA at 1.4 GHz (31"×33"), and 1.7 GHz for OH survey.

LaRosa *et al.* 2000, AJ, 119, 207. VLA at 333 MHz (24"×43"). *see also:* LaRosa *et al.* 2000, AJ, 119, 3145. Erratum.

Bamba *et al.* 2000, PASJ, 52, 259. ASCA observations.Lazendic *et al.* 2002, MNRAS, 331, 537. Observations of shocked molecular gas where the 'snake' crosses the remnant.

Yusef-Zadeh *et al.* 2004, ApJS, 155, 421. VLA at 1.4 GHz (8".4×12".8).

Aharonian *et al.* 2008, A&A, 483, 509. XMM-Newton and HESS observations.

Hewitt *et al.* 2008, ApJ, 683, 189. GBT at 1.6 and 1.7 GHz for OH masers.

Bamba et al. 2009, ApJ, 691, 1854. Suzaku observations.
Stupar & Parker 2011, MNRAS, 414, 2282. Hα observations.
Ohnishi et al. 2011, PASJ, 63, 527. Suzaku observations.
Pihlström et al. 2014, AJ, 147, 73. VLA search for methanol masers.

Ponti et al. 2015, MNRAS, 453, 172. XMM-Newton observations.
Froebrich et al. 2015, MNRAS, 454, 2586. H<sub>2</sub> IR observations.
Hui et al. 2016, MNRAS, 457, 4262. Fermi observations of region.

de Wilt et al. 2017, MNRAS, 468, 2093. Molecular line observations of region.

Qiao et al. 2018, ApJS, 239, 15. OH maser observations.

# G359.1+0.9

**RA:** 17<sup>h</sup>39<sup>m</sup>36<sup>s</sup> **1-GHz flux/Jy:** 2? **Size/arcmin:** 12×11 **Dec:** -29°11′ **Spectral index:** ? **Type:** S

Radio: Shell, brightest in E.

### References:

Gray 1994, MNRAS, 270, 847. MOST at 843 MHz  $(43'' \times 88'' : S = 4.3 \text{ Jy})$ .

Roy & Bhatnagar 2006, JPhCS, 54, 152. GMRT at 330 MHz (1':  $S=4.3\pm1.0$  Jy).

Law et al. 2008, ApJS, 177, 515. VLA at 1.4 GHz  $(10''.9 \times 15''.9 : S = 1.3 \pm 0.5 \text{ Jy})$ .