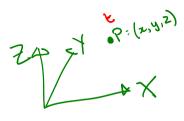


```
Speed of light* in ms<sup>-1</sup> is c  
Wavelength in m is \lambda  Gran \lambda = 532 × 10<sup>-4</sup> × Wave period in s is T = \lambda/c  
Wave frequency in Hz is \nu = 1/T  
Angular freq. in rad s<sup>-1</sup> is \omega = 2\pi/T  
Wave number in rad m<sup>-1</sup> is k = 2\pi/\lambda
```

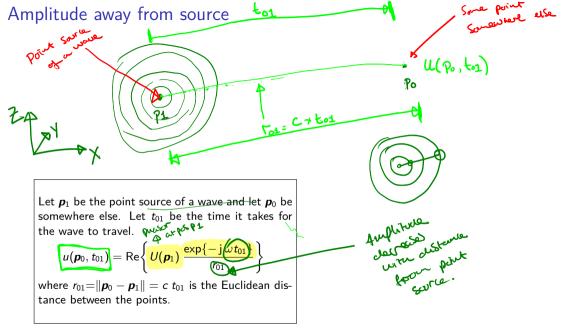


A phasor encodes max, amplitude $\widehat{A}[p]$ and phase $\widehat{\phi}(p)$ at position p, (2,3,2) = $\underbrace{U(p)} = \widehat{\phi}(p) \exp\{j \phi(p)\}$.

The scalar value of an EM wave vector component at time t can be found as,

$$u(\mathbf{p}, t) = (\text{Re}) U(\mathbf{p}) \exp\{-i\omega t\}$$

$$= (A(\mathbf{p}) \cos(\omega t - \phi(\mathbf{p})).$$



$$\omega \ t_{01} = \frac{2\pi}{T} \ t_{01} = \frac{2\pi}{\lambda/c} \ t_{01}$$

$$= \frac{2\pi \ c}{\lambda} \ t_{01} = \boxed{k \ r_{01}}$$
Since there is no explicit time term, this can be used to express the phasor at p_0 .
$$\omega \omega \nabla U(p_0) = U(p_1) \exp\{-j \ k \ r_{01}\}$$

Speed of light* in ms^{-1} is c

Wavelength in m is λ

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Let p_1 be the point source of a wave and let p_0 be somewhere else. Let t_{01} be the time it takes for the wave to travel.

$$u(\boldsymbol{p}_0,t_{01}) = \mathsf{Re} \Bigg\{ U(\boldsymbol{p}_1) \, rac{\mathsf{exp} \{ -\, \mathsf{j} \, \omega \, t_{01} \}}{r_{01}} \Bigg\}$$

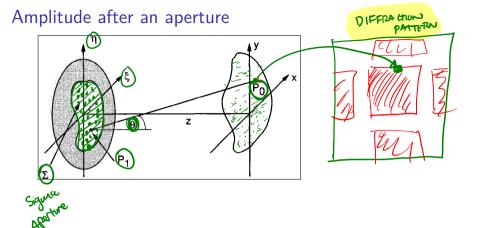
where $r_{01} = ||p_0 - p_1|| = c t_{01}$ is the Euclidean distance between the points.

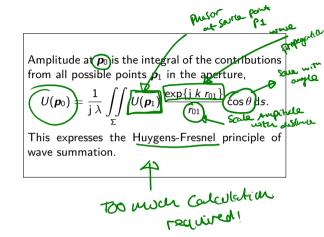
$$\omega t_{01} = \frac{2\pi}{T} t_{01} = \frac{2\pi}{\lambda/c} t_{01}$$

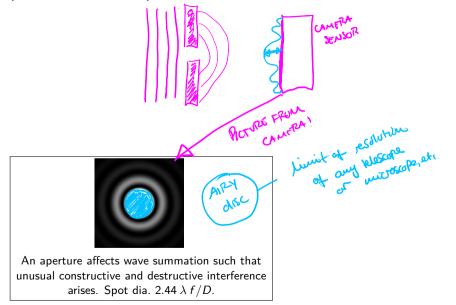
$$= \frac{2\pi c}{\lambda} t_{01} = k r_{01}$$

Since there is no explicit time term, this can be used to express the phasor at p_0 ,

$$U(\mathbf{p}_0) = U(\mathbf{p}_1) \frac{\exp\{-j \ k \ r_{01}\}}{r_{01}}$$



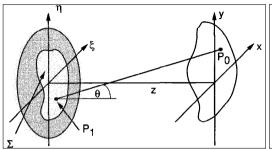




This is termed *diffraction* and it happens to all physical waves:

- ► light
- sound
- vibration (e.g. of water)
- gravitational waves

Diffuse reflection from a rough surface can also be understood as diffraction.



Amplitude at p_0 is the integral of the contributions from all possible points p_1 in the aperture,

$$U(\boldsymbol{p}_0) = \frac{1}{j \lambda} \iint_{\Sigma} U(\boldsymbol{p}_1) \frac{\exp\{j \ k \ r_{01}\}}{r_{01}} \cos \theta \, ds.$$

This expresses the Huygens-Fresnel principle of wave summation.

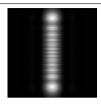


An aperture affects wave summation such that unusual constructive and destructive interference arises. Spot dia. $2.44 \lambda f/D$.

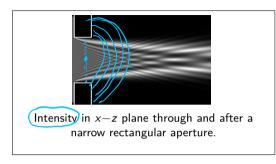
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Intensity in x-y plane after a narrow rectangular aperture.



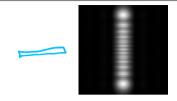
Significant computation is required for *numerical* solutions that simulate diffraction effects through summation of wave amplitudes.

Many different techniques can be used, e.g. *finite element methods*, to find wave amplitudes at discrete volumes in space at successive steps in time.

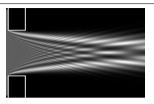
Note that intensity at distance r_{01} is distributed over a sphere whose surface area is $4\pi r_{01}^2$. So intensity scales $\propto 1/r_{01}^2$.

Since amplitude is the square root of intensity, it scales $\propto \sqrt{1/r_{01}^2} = 1/r_{01}$.

In two dimensional wave propagation (e.g. on water) the amplitude scales $\propto 1/\sqrt{r_{01}}.$



Intensity in x-y plane after a narrow rectangular aperture.



Intensity in x-z plane through and after a narrow rectangular aperture.

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Since $\cos \theta = z/r_{01}$, wave summation can be rewritten in more explicit rectangular coordinates as,

$$U(x, y, z) = \frac{z}{j \lambda} \iint_{\Sigma} U(\xi, \eta) \frac{\exp\{j k r_{01}\}}{r_{01}^2} d\xi d\eta$$

with distance calculated as,

$$r_{01} = \sqrt{z^2 + (x - \xi)^2 + (y - \eta)^2}$$

To facilitate analytical solutions, an approximation for distance $\it r_{01}$ uses a binomial expansion to replace the square root,

$$\sqrt{1+b} = (1+b)^{rac{1}{2}}$$

$$= 1 + rac{1}{2}b - rac{1}{8}b^2 + \dots$$
when $|b| < 1$.

$$\begin{split} r_{01} &= \sqrt{z^2} \sqrt{1 + \left(\frac{x-\xi}{z}\right)^2 + \left(\frac{y-\eta}{z}\right)^2} \\ &\approx z \Big[1 + \frac{1}{2} \Big(\frac{x-\xi}{z}\Big)^2 + \frac{1}{2} \Big(\frac{y-\eta}{z}\Big)^2\Big] \\ \text{using only the first two terms of the expansion} \\ \text{with } b &= \Big(\frac{x-\xi}{z}\Big)^2 + \Big(\frac{y-\eta}{z}\Big)^2. \\ \text{(cf. parabolic approx. of spherical wavefront.)} \end{split}$$

The same approximation for r_{01} doesn't have to be used for all occurances.

Using the first term only, the denominator $r_{01}^2 \approx z^2$. This can be factored out of the integral into the scaling term,

$$\frac{z}{j \lambda z^2} = \frac{1}{j \lambda z}.$$

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$$\mathbf{A}_{01} = \sqrt{z^2} \sqrt{1 + \left(\frac{x - \xi}{z}\right)^2 + \left(\frac{y - \eta}{z}\right)^2}$$

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using only the first two terms of the expansion with $b = \left(\frac{x-\xi}{2}\right)^2 + \left(\frac{y-\eta}{2}\right)^2$.

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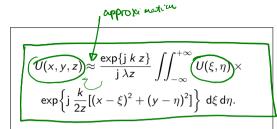
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$$\exp\{j\ k \text{ on}\} \approx \exp\left\{j\ k \overline{z\left[1+\frac{1}{2}\left(\frac{x-\xi}{z}\right)^2+\frac{1}{2}\left(\frac{y-\eta}{z}\right)^2\right]}\right\}$$

$$= \exp\{j\ k\ z\}\ \exp\left\{j\ k\ z \left[\frac{1}{2}\left(\frac{x-\xi}{z}\right)^2+\frac{1}{2}\left(\frac{y-\eta}{z}\right)^2\right]\right\}$$

$$= \exp\{j\ k\ z\}\ \exp\left\{j\ \frac{k}{2z}[(x-\xi)^2+(y-\eta)^2]\right\}$$



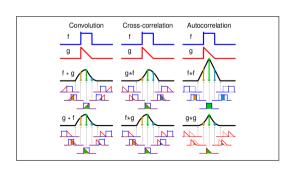
Accurate only for the "near field" close to the aperture because of distance approximation.

To facilitate analysis, it can be written as a convolution of the aperture with a function h.

$$U(x, y, z) \approx \iint_{-\infty}^{+\infty} U(\xi, \eta) \times h(x - \xi, y - \eta) \, d\xi \, d\eta.$$

Convolution kernel h(v, w) =

$$\frac{\exp\{j\;k\;z\}}{j\;\lambda z}\exp\Big\{j\;\frac{k}{2z}\big(v^2+w^2\big)\Big\}.$$



$$\exp\{j \ k \ r_{01}\} \approx \exp\{j \ k \ z \left[1 + \frac{1}{2} \left(\frac{x - \xi}{z}\right)^{2} + \frac{1}{2} \left(\frac{y - \eta}{z}\right)^{2}\right]$$

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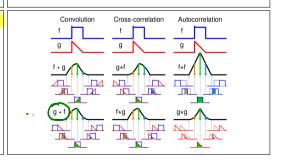
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$$(x - \xi)^2 = x^2 - 2x\xi + \xi^2,$$

 $(y - \eta)^2 = y^2 - 2y\eta + \eta^2.$

Hence further factorization outside the integral is possible since only those terms in ξ and η need to remain inside.

$$\exp\Bigl\{\mathrm{j}\,\frac{k}{2z}[(x-\xi)^2+(y-\eta)^2]\Bigr\}=$$

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This integral can be recognised as the (scaled) Fourier transform of the (scaled) aperture evaluated at spatial frequencies,

$$f_X = \frac{x}{\lambda z}$$
 $f_Y = \frac{y}{\lambda z}$.

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Note that $k = 2\pi/\lambda$,

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$$U(x, y, z) \approx \frac{\exp\{j \ k \ z\}}{j \ \lambda z} \exp\{j \ \frac{k}{2z}(x^2 + y^2)\} \times \int_{-\infty}^{+\infty} U(\xi, \eta) \exp\{j \ \frac{k}{2z}(\xi^2 + \eta^2)\} \times \exp\{-j \ \frac{2\pi}{\lambda z}(x\xi + y\eta)\} \ d\xi \ d\eta.$$

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When z in $\frac{k}{2z}(\xi^2+\eta^2)$ is very big,* this expression is close to 0 so its exponent is close to 1. So it is not essential to use it a scaling factor,

$$U(x, y, z) = \frac{\exp\{j kz\} \exp\{j \frac{k}{2z}(x^2 + y^2)\}}{j \lambda z}$$
$$\iint_{-\infty}^{+\infty} U(\xi, \eta) \exp\{-j \frac{2\pi}{\lambda z}(x\xi + y\eta)\} d\xi d\eta$$

Hence wave summation can be expressed as the (scaled) Fourier transform of the (unscaled) aperture evaluated at frequencies,

$$f_X = \frac{x}{\lambda z}$$
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Accurate only for the "far field" distant from the aperture because of distance assumption.

For intensity $I(x, y, z) = |U(x, y, z)|^2$, the numerator and denominator of the scaling term simplify as follows.

$$|\exp\{j kz\}|^2 = \exp\{+j kz\} \times \exp\{-j kz\} = \exp\{+j kz - j kz\} = \exp\{0\} = 1.$$

$$|j \lambda z|^2 = (+j \lambda z)(-j \lambda z)$$
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For a nice alternative derivation of the material in this lecture, see https://www.youtube.com/watch?v=JKxDa5D3GnQ.

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