

Radio Galaxy Zoo: Data Release 1 of 82,071 radio sources

Radio Galaxy Zoo science team et al.

Accepted year month day. Received year month day; in original form year month day

ABSTRACT

Data Release 1 for the Radio Galaxy Zoo (RGZ) project.

Key words: methods: data analysis — radio continuum: galaxies — infrared: galaxies.

1 INTRODUCTION

A complete introduction and description of the project is provided in Banfield et al. (2015, hereafter B15).

2 RADIO GALAXY ZOO

2.1 FIRST + WISE

2.2 ATLAS + SWIRE

ATLAS images are 2×2 arcmin.

3 DATA REDUCTION

Images in RGZ are presented for classification by independent users. The users are treated as having equivalent levels of skill, with consensus accomplished by a simple majority vote.

Current statistics of the project.

4 CATALOG

There are three basic types of data products for Radio Galaxy Zoo: raw classifications, consensus catalogs, and static versions.

Raw classifications are the individual clicks that each user performs; they contain the raw pixel information corresponding to the selection of radio components and the IR counterpart, if available. These are unlikely to be used by most science team members, since they don’t have consensus or weighting, require linking to the subject, and are stored only in MongoDB format. Raw classifications are updated daily on the Zooniverse servers.

The consensus catalog is the aggregated classifications over all users, sorted for each subject. This is run through a Python pipeline, combining the 20 total votes (or 5, in the case of single-component radio sources) and finding the most common answer. Only retired subjects with the full number of classifications are analyzed. We then add physical parameters to each match by measuring the properties in the radio image and positionally cross-matching to the AllWISE

(Cutri 2013) and SDSS DR12 (Alam et al. 2015) catalogs. The consensus is updated whenever the latest raw classifications are re-run against the consensus algorithm (every couple weeks, usually). The data is stored in MongoDB format.

Static versions of the catalogs can be generated from the MongoDB versions. These are “flat” versions that are more like the data products typically used in astronomy; a data table in CSV or FITS format where each row corresponds to a unique source and each column is a measured parameter. It’s different from the consensus catalog in two ways: firstly, it’s not updated as often and so represents a “static” version of the total classifications. Secondly, there are parameters that will have different numbers of elements for each source — for example, the number of distinct radio sources or peaks in a given source. Since that can’t be included in a flat table, these data are not included — use the MongoDB version of the consensus if you want data on that.

Neither the consensus nor static catalogs have the ATLAS subjects incorporated yet; they only contain FIRST images.

The fundamental entry in the DR1 catalog is a radio source, which contains one or more radio components and a possible IR counterpart (Tables 2 and 4).

Column 1 contains the unique ID for the RGZ source. Columns 2 and 3 contain the J2000.0 coordinates for the infrared counterpart of the radio source. Column 4 gives the kernel width (in arcsec) of the aggregate clicks used to pinpoint the IR source, providing a measure of positional uncertainty for the host identification. Columns 2-4 are only populated if at least 50% of the users positively identified an infrared counterpart from the WISE data. Columns 5-6 give the total integrated flux and error (in mJy) for all radio components associated with this source. Columns 7-8 give the peak integrated flux density and error (in mJy/beam) for the brightest radio peak in the source. Columns 9-10 give the integrated luminosity and error for all radio components associated with the source. Column 11 gives the maximum angular extent (in arcsec) of the bounding boxes for all radio components, as measured corner-to-corner. Column 12 is the transverse physical size (in kpc) corresponding to the maximum angular extent. Column 13 is the total solid angle for the radio source, calculated by summing the individual

solid angles subtended by the outermost contours for each radio component. Column 14 gives the cross-sectional area (in kpc^2) corresponding to the total solid angle. Column 15 gives the total number of radio peaks in the source, defined as the sum of the number of individual components plus any additional local maxima within a single component.

All components relating to the radio luminosity, transverse size, or cross-sectional area are only calculated if a redshift has been detected for the radio source's optical counterpart, since all such values require a distance.

5 RESULTS

5.1 FIRST

5.2 ATLAS

6 SUMMARY

REFERENCES

- Alam S. et al., 2015, *Ap. J. Suppl.*, 219, 12
 Banfield J. K. et al., 2015, *MNRAS*, 453, 2326
 Cutri R. M., 2013, VizieR Online Data Catalog, 2328

Table 1. RGZ consensus classifications of FIRST radio morphologies

RGZ ID	FIRST ID	Zooniverse ID	N_{class}	C_l	N_{comp}	IR counterpart
1	FIRSTJ145834.5+140942	ARG0002qe4	18	0.833	1	Y
2	FIRSTJ130905.4+433849	ARG0000yc4	5	1.000	1	Y
3	FIRSTJ102805.7+542412	ARG0000dcs	4	0.800	1	Y

Note. — The full, machine-readable version of this table is available at on the journal website and at <http://data.galaxyzoo.org/radio>. A portion is shown here for guidance on form and content.

Table 2. Matched catalog for RGZ-FIRST consensus sources

RGZ ID	IR counterpart		e_{IR}	S_{ν}	$\sigma_{S_{\nu}}$	$S_{\nu,peak}$	$\sigma_{S_{\nu,peak}}$	Radio				A_{tot}	N_{peaks}
	RA	dec						L_{ν}	$\sigma_{L_{\nu}}$	θ_{max}	$D_{A,max}$		
	J2000	J2000	[arcsec]	[mJy]	[mJy]	[mJy beam $^{-1}$]	[mJy beam $^{-1}$]	[W/Hz]	[W/Hz]	[arcmin]	[kpc]	[kpc 2]	
1	23.38219	251.6794	err	10.37	0.20	7.21	0.02	2.06e+24	3.92e+22	0.28	65.80	1666.7	1

Note. — The full, machine-readable version of this table is available at on the journal website and at <http://data.galaxyzoo.org/radio>. A portion is shown here for guidance on form and content.

Table 3. RGZ consensus classifications of ATLAS radio morphologies

RGZ ID	ATLAS ID	Zooniverse ID	N_{class}	C_l	N_{comp}	IR counterpart
1	CI002	ARG0002qe4	18	0.833	1	Y
2	CI003	ARG0000yc4	5	1.000	1	Y
3	CI004	ARG0000dcs	4	0.800	1	Y

Note. — The full, machine-readable version of this table is available at on the journal website and at <http://data.galaxyzoo.org/radio>. A portion is shown here for guidance on form and content.

Table 4. Matched catalog for RGZ-ATLAS consensus sources

RGZ ID	IR counterpart		e_{IR}	S_ν	σ_{S_ν}	$S_{\nu,peak}$	$\sigma_{S_{\nu,peak}}$	Radio				A_{tot}	N_{peaks}
	RA	dec						L_ν	σ_{L_ν}	θ_{max}	$D_{A,max}$		
	J2000	J2000	[arcsec]	[mJy]	[mJy]	[mJy beam ⁻¹]	[mJy beam ⁻¹]	[W/Hz]	[W/Hz]	[arcmin]	[kpc]	[arcsec ²]	[kpc ²]
1	23.38219	251.6794	err	10.37	0.20	7.21	0.02	2.06e+24	3.92e+22	0.28	65.80	105.9	1666.7
													1

Note. — The full, machine-readable version of this table is available at on the journal website and at <http://data.galaxyzoo.org/radio>. A portion is shown here for guidance on form and content.