

# INVESTIGATING THE DEPENDENCE ON MACHINE SIZE OF THE ENERGY CONFINEMENT IN TOKAMAKS USING DATA-DRIVEN METHODS

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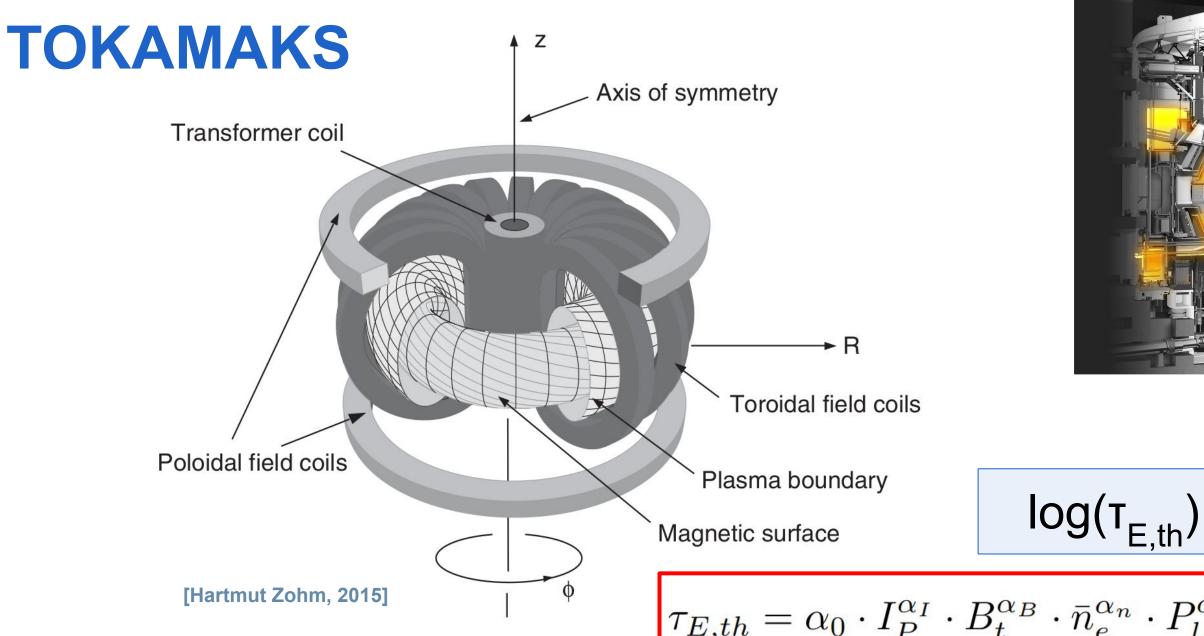
11/07/2023

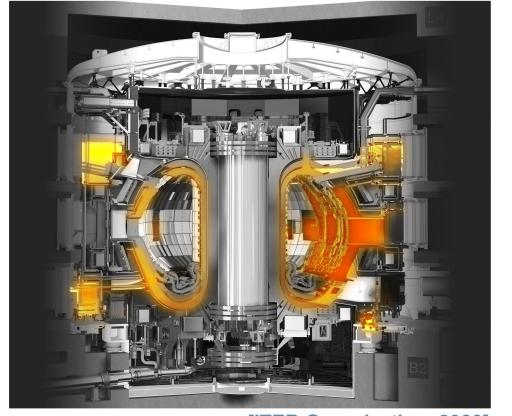






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[ITER Organization, 2023]

Plasma boundary
Magnetic surface
$$\log(\mathsf{T}_{\mathsf{E},\mathsf{th}}) \longrightarrow \mathsf{OLS}$$

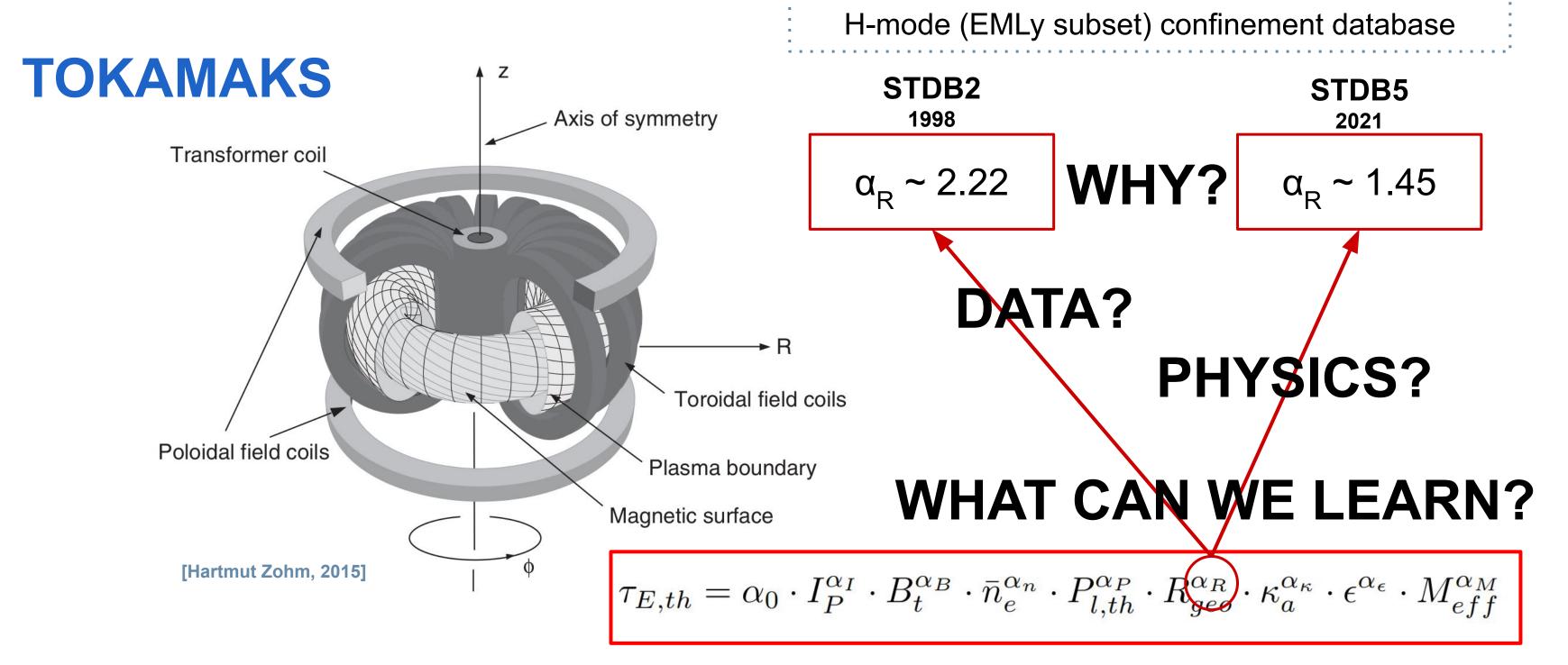
$$\tau_{E,th} = \alpha_0 \cdot I_P^{\alpha_I} \cdot B_t^{\alpha_B} \cdot \bar{n}_e^{\alpha_n} \cdot P_{l,th}^{\alpha_P} \cdot R_{geo}^{\alpha_R} \cdot \kappa_a^{\alpha_\kappa} \cdot \epsilon^{\alpha_\epsilon} \cdot M_{eff}^{\alpha_M}$$







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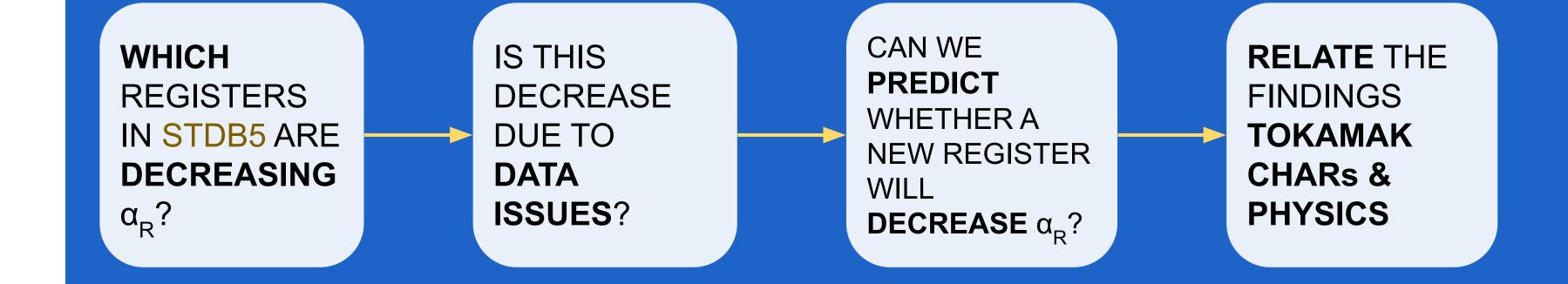








# WORKFLOW







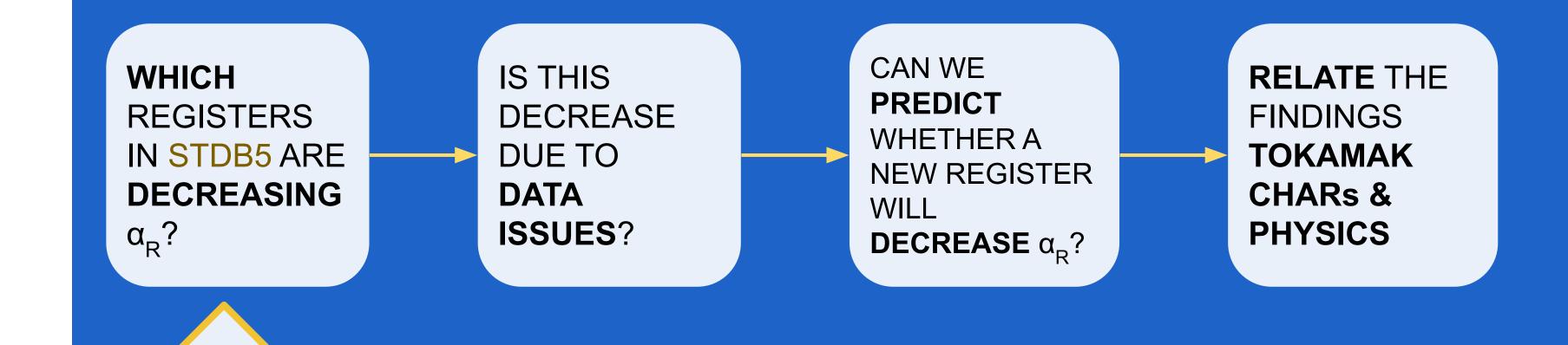




**GHENT** 

UNIVERSITY

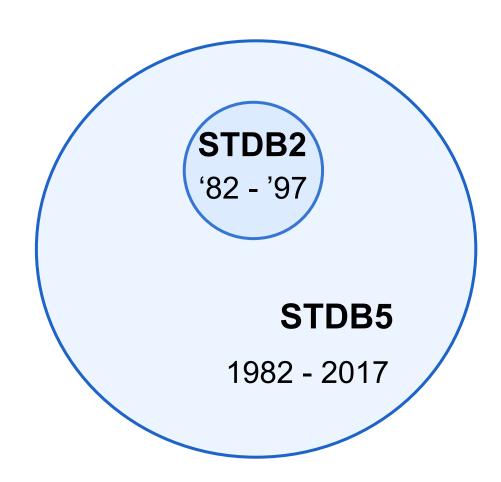
# WORKFLOW











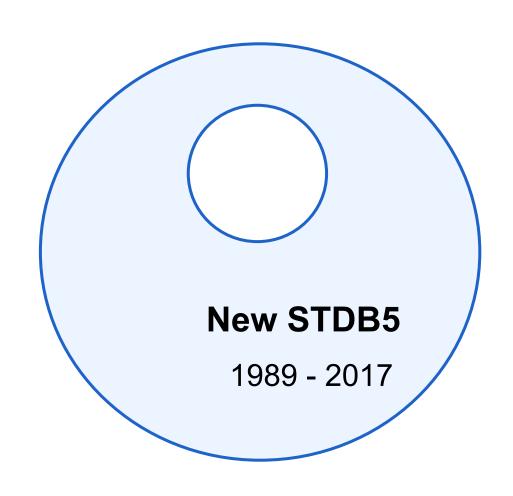
Dataset size = 6252



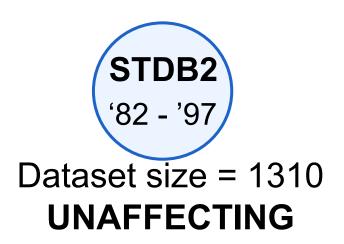










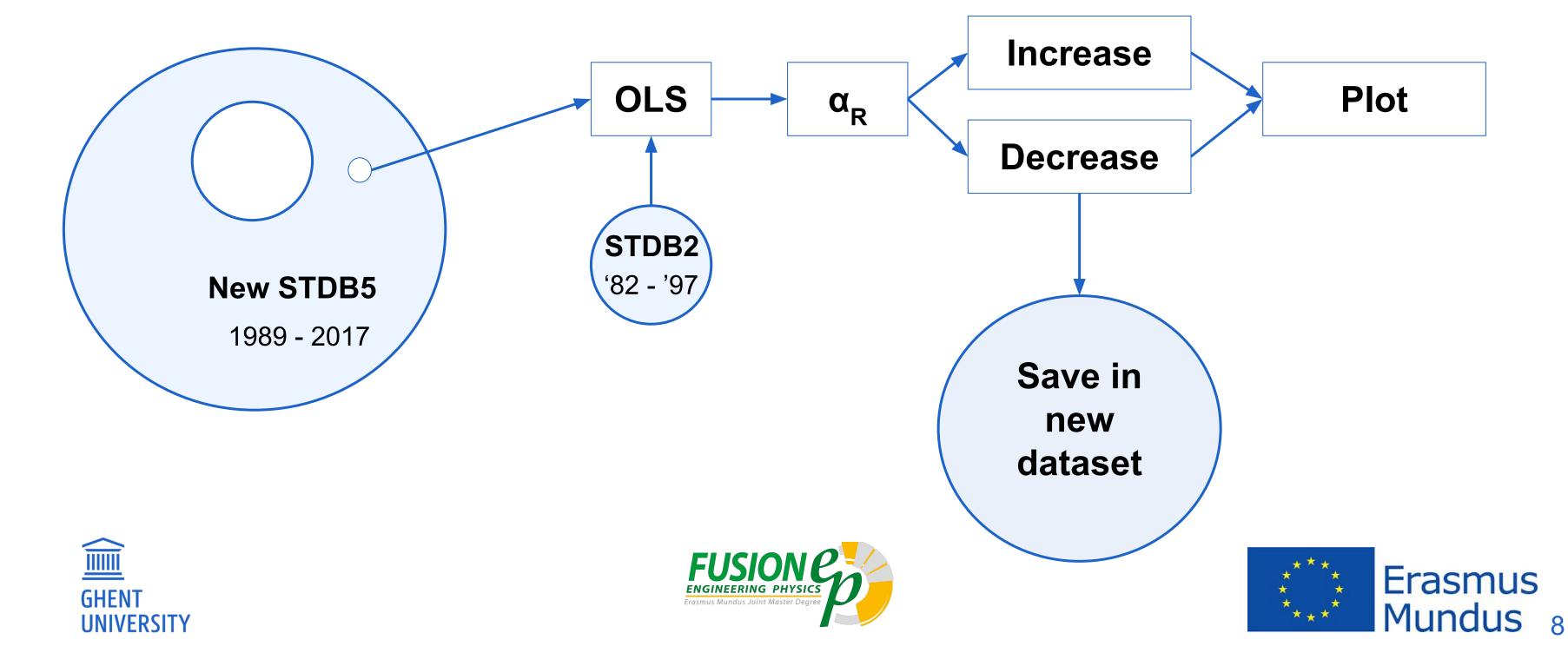




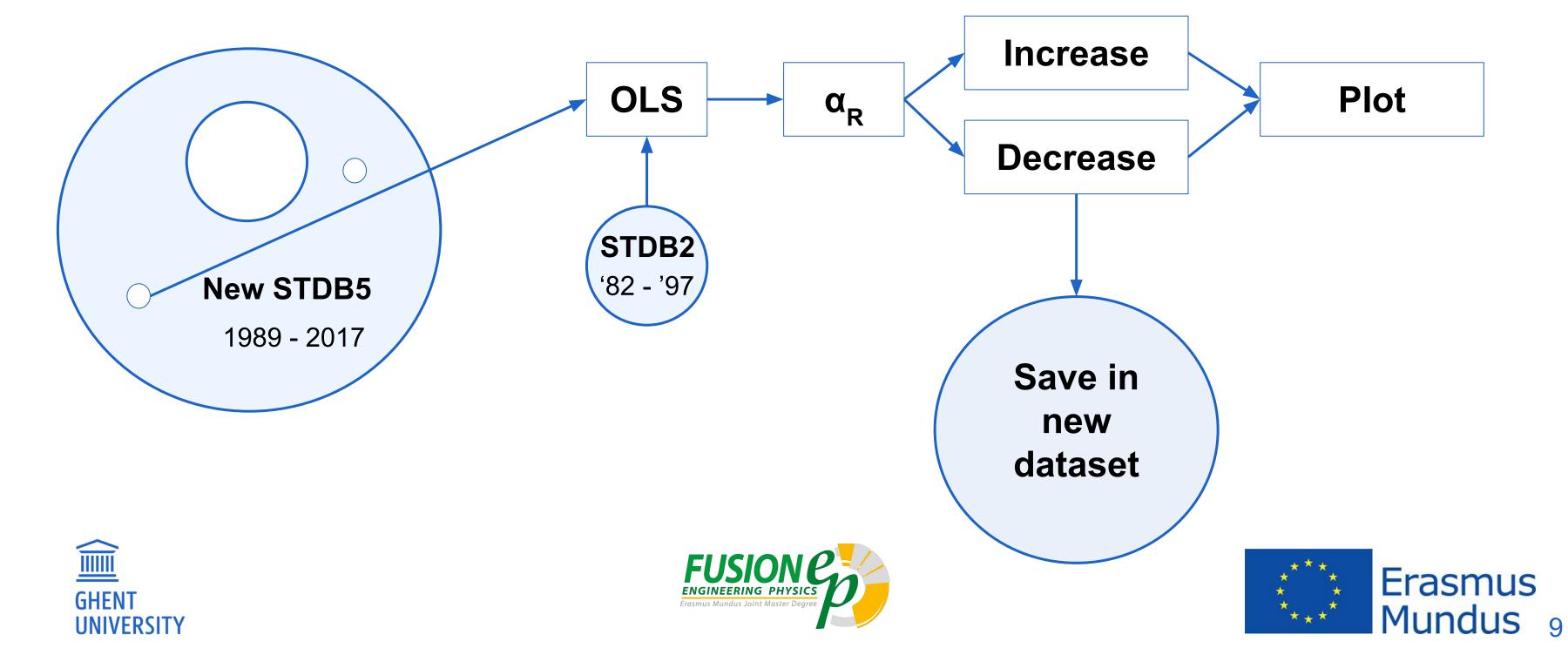




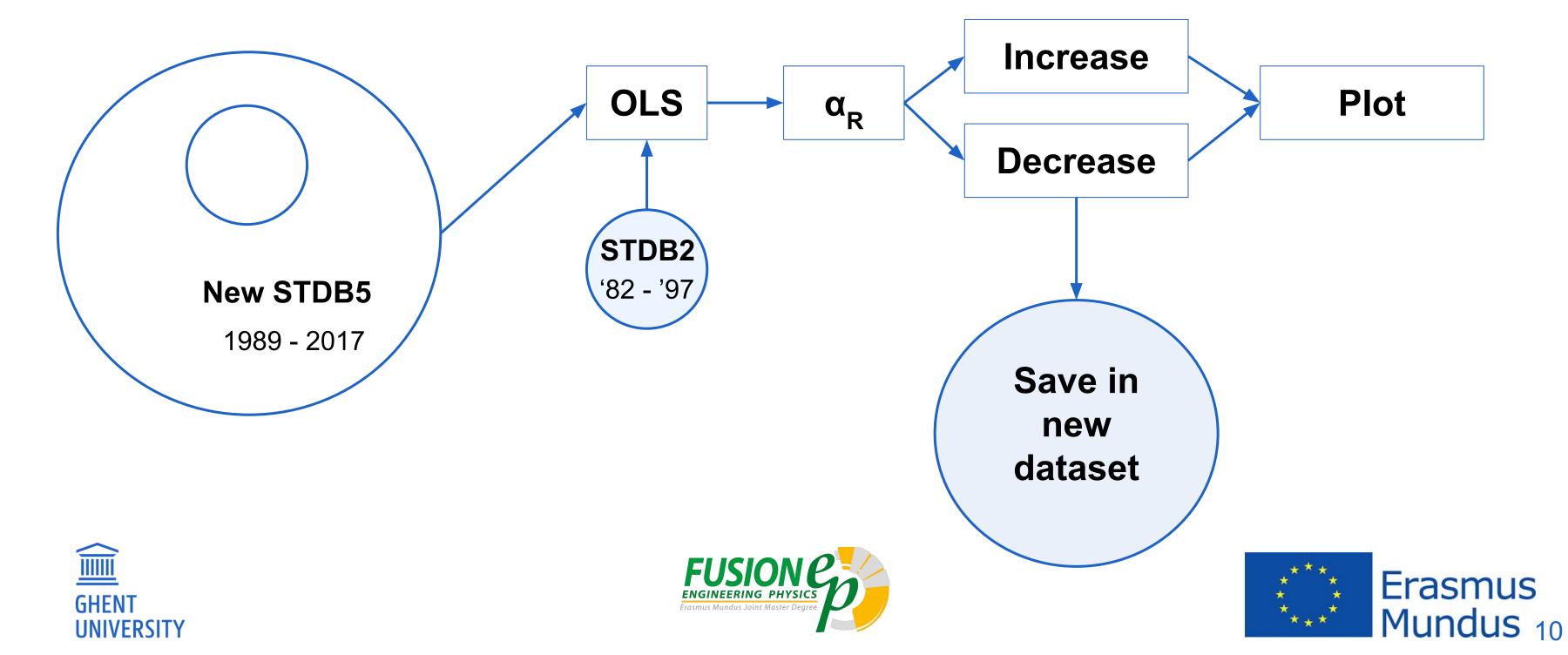






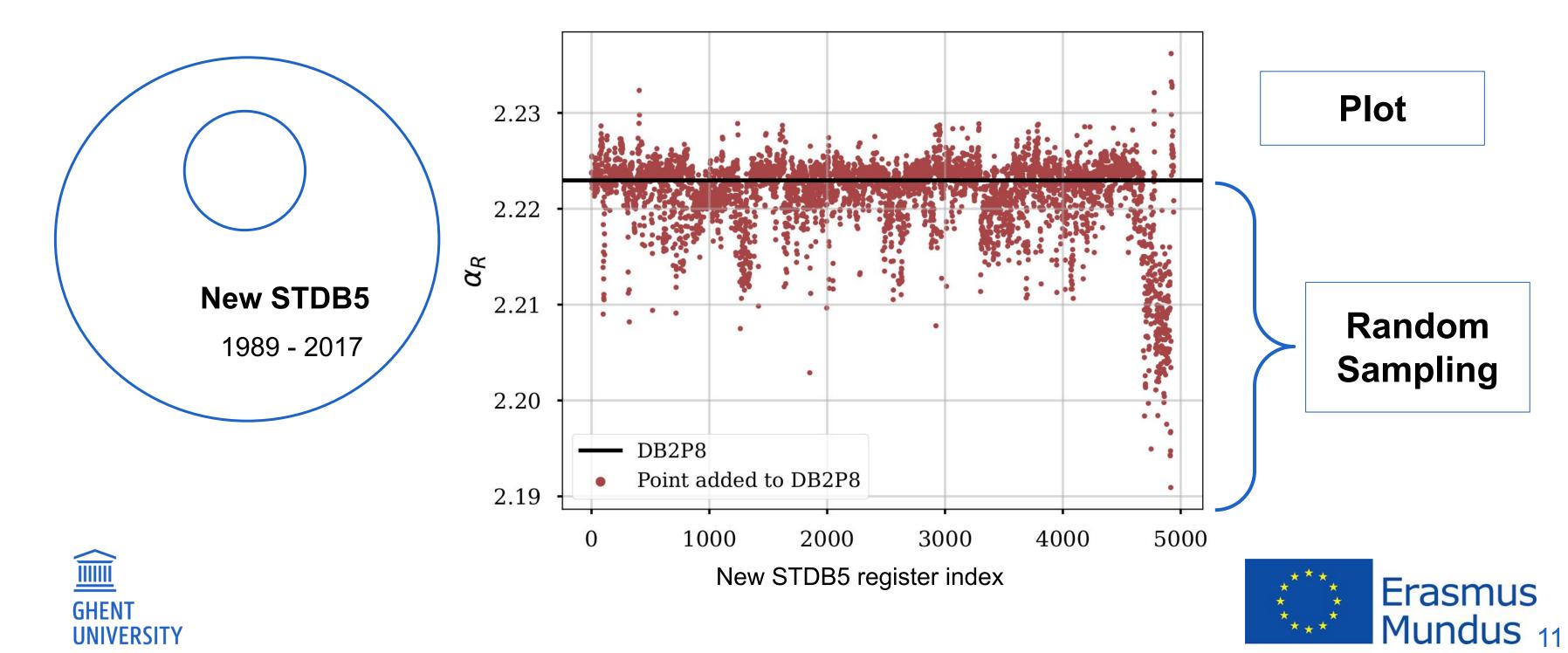








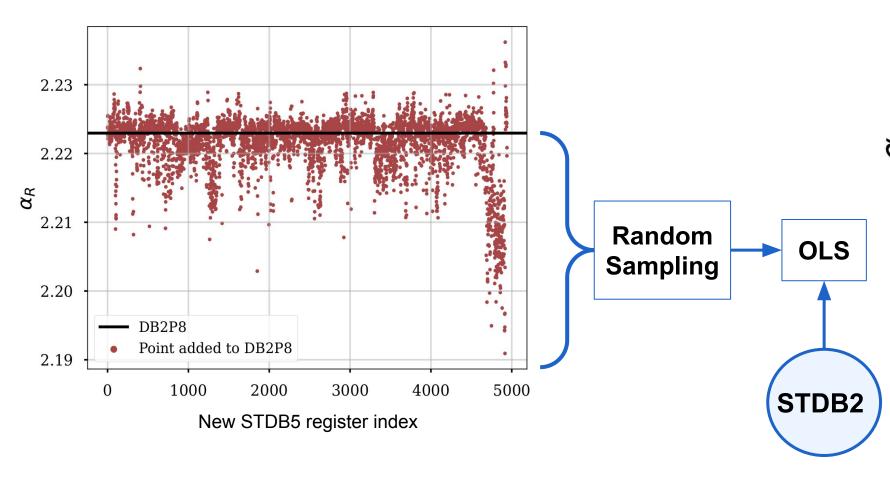
## **DECREASING REGISTERS**

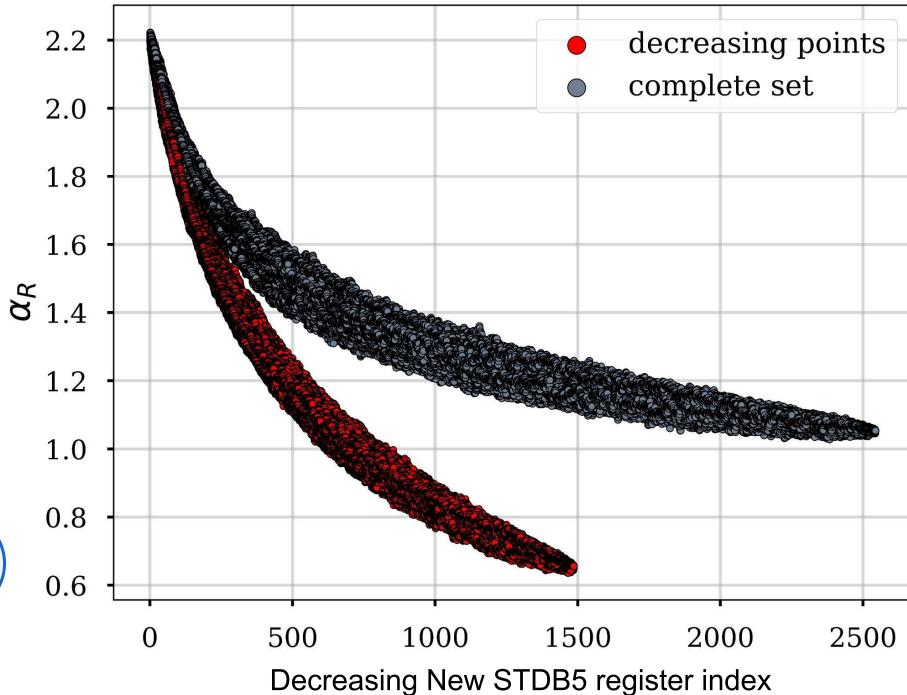






## **DECREASING REGISTERS**





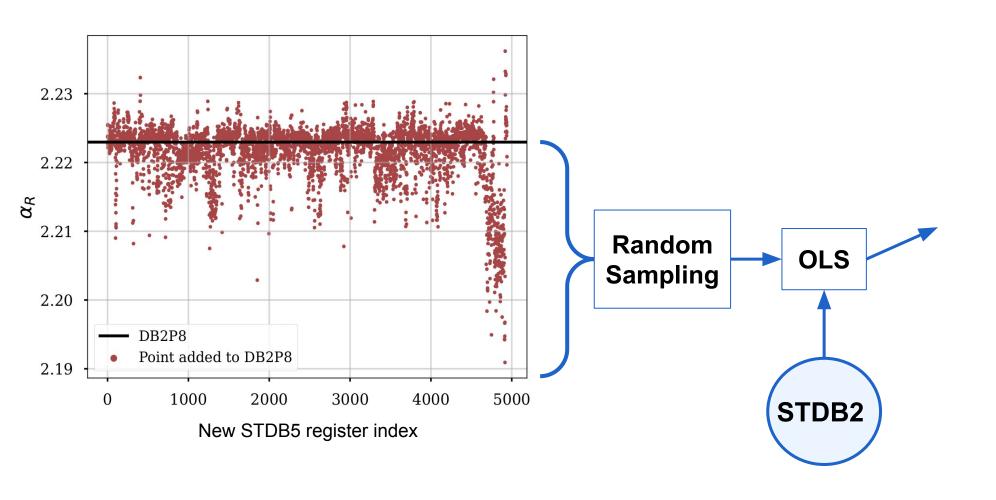








## **DECREASING REGISTERS**



"Small dataset"

"Big dataset"

Smallest subset for  $\alpha_R < 1$ 

 $\alpha_{R} \sim 0.9998$ 

Subset size = 618

9.88% decreased  $\alpha_R$  90.12% did not

Smallest  $\alpha_R$ 

 $\alpha_{R} \sim 0.6379$ 

Subset size = 1459

23.34% decreased  $\alpha_R$  76.66% did not

STDB5









# DECREASING AND UNAFFECTING REGISTERS

"Small dataset"

"Small dataset"

"Big dataset"

"Big dataset"

Smallest subset for

 $\alpha_R < 1$ 

 $\alpha_{\rm R} \sim 0.99$ 

dataset size = 618

**Complementary** for

 $\alpha_R < 1$ 

 $\alpha_{R} \sim 1.71$ 

dataset size = 5634

Complementary for smallest  $\alpha_R$ 

 $\alpha_{R} \sim 2.16$ 

dataset size = 4793

Smallest  $\alpha_R$ 

 $\alpha_{\rm R} \sim 0.64$ 

dataset size = 1459

+

+

Together they make **STDB5** 

dataset size = 6252





# WORKFLOW

**WHICH** REGISTERS IN STDB5 ARE **DECREASING**  $\alpha_R$ ?

Depends on how you

IS THIS DECREASE **DUE TO DATA ISSUES?** 

**CAN WE PREDICT** WHETHER A **NEW REGISTER** WILL **DECREASE**  $\alpha_{R}$ ?

**RELATE** THE **FINDINGS TOKAMAK** CHARs & **PHYSICS** 

classify / label them.







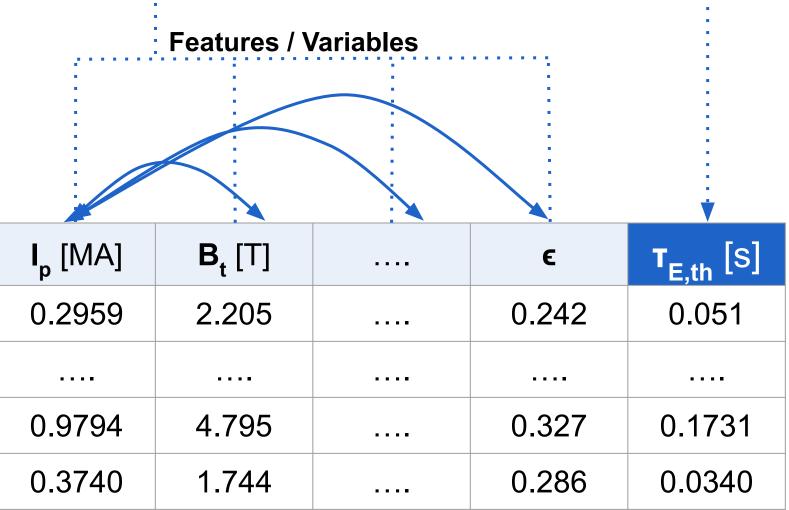
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## **MULTICOLLINEARITY?**

## **CONSEQUENCES**

- Increase in the standard errors of each feature
- Numerical instability
- Unreliable models





Used as predictors for the **target** 





[W.-M Lee, 2021]
[D.Besley, E. Kuh, and R. Welsch, 2004]
[R. M. O'Brien, 2007]





## VARIANCE INFLATION FACTOR

$$y = X\beta$$

$$\hat{\boldsymbol{\beta}} = \left(\mathbf{X}^T \mathbf{X}\right)^{-1} \mathbf{X}^T \mathbf{y}$$

$$VIF_k = \frac{1}{1 - R_k^2}$$

$R^2 - 1 -$	$\sum_{i=1}^{n} (y_i -$	$-\hat{y}_i)^2$
$\Lambda$ – 1 –	$\overline{\sum_{i=1}^{n}(y_i-y_i)}$	$\overline{(-\bar{y})^2}$

I <sub>p</sub> [MA]	<b>B</b> <sub>t</sub> [T]	 €	T <sub>E,th</sub> [s]
0.2959	2.205	 0.242	0.051
	I I ····	 	
0.9794	4.795	 0.327	0/1731
0.3740	1.744	 0.286	0.0340





[D.Besley, E. Kuh, and R. Welsch, 2004] [R. M. O'Brien, 2007]



## VARIANCE INFLATION FACTOR

$$y = X\beta$$

$$\hat{\boldsymbol{\beta}} = \left(\mathbf{X}^T \mathbf{X}\right)^{-1} \mathbf{X}^T \mathbf{y}$$

$$VIF_k = \frac{1}{1 - R_k^2}$$

$R^2 - 1 -$	$\sum_{i=1}^{n}$	$(y_i - \hat{y}_i)^2$
N - 1 -	$\sum_{i=1}^{n}$	$\overline{(y_i-\bar{y})^2}$

I <sub>p</sub> [MA]	B <sub>t</sub> [T]		E	T <sub>E,th</sub> [s]
0.2959	2.205		0.242	0.051
0.9794	4.795		0.327	0/1731
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		I		





[D.Besley, E. Kuh, and R. Welsch, 2004] [R. M. O'Brien, 2007]



## VARIANCE INFLATION FACTOR

$$y = X\beta$$

$$\hat{\boldsymbol{\beta}} = \left(\mathbf{X}^T \mathbf{X}\right)^{-1} \mathbf{X}^T \mathbf{y}$$

$$VIF_k = \frac{1}{1 - R_k^2}$$

$$Var[\beta_k] = \frac{\left(1 - R^2\right) \cdot \sum \left(y_i - \bar{y}\right)^2}{n - M - 1} \cdot \frac{1}{\sum x_k^2} \cdot VIF_k$$

$R^2 - 1 -$	$\sum_{i=1}^{n}$	$(y_i - \hat{y}_i)^2$
N - 1 -	$\sum_{i=1}^{n}$	$\overline{(y_i - \bar{y})^2}$

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0.0340	0.286	 1.744	0.3740





[D.Besley, E. Kuh, and R. Welsch, 2004] [R. M. O'Brien, 2007]



## **VIF: RESULTS**

$$y = X\beta$$

$$\hat{\boldsymbol{\beta}} = \left(\mathbf{X}^T \mathbf{X}\right)^{-1} \mathbf{X}^T \mathbf{y}$$

$$VIF_k = \frac{1}{1 - R_k^2}$$

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$I_p$ -	25	17	16	21	19	18	- 25
$B_t$ -	3.3	5.2	3.8	5.4	3.3	5.3	- 20
$ar{n}_e$ -	4.6	2.7	2.8	3	3.2	2.8	
$P_{l,th}$ -	5.8	4.1	5.4	4.4	5.8	4.2	- 15
$R_{geo}$ -	11	13	10	14	9.4	14	- 10
$\kappa_a$ -	5.6	2.7	3.6	3.2	4.4	2.8	
$\epsilon$ -	7.2	8.5	7.8	8.8	7.3	8.7	- 5
$M_{eff}$ -	1.7	1.3	1.5	1.4	1.5		- 0
	SIDBL	TOB'S	Big	Big	Small	Small	- 0
Ć	5,	1.3 STDBS Decreasing	Mattectin	g. Big	Small Small	) -	







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## **VIF: RESULTS**

$$\mathbf{y} = \mathbf{X}\boldsymbol{\beta} + \boldsymbol{\varepsilon}$$

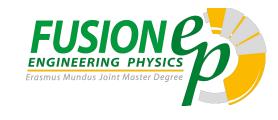
$$\mathbf{y} = \mathbf{X}\boldsymbol{\beta} + \boldsymbol{\varepsilon} \qquad \hat{\boldsymbol{\beta}} = (\mathbf{X}^T\mathbf{X})^{-1}\mathbf{X}^T\mathbf{y}$$

$$VIF_k = \frac{1}{1 - R_k^2}$$

$$Var[\beta_k] = \frac{\left(1 - R^2\right) \cdot \sum \left(y_i - \bar{y}\right)^2}{n - M - 1} \cdot \frac{1}{\sum x_k^2} \cdot VIF_k$$

$B_t$ -	1.7	2.6	2.1	2.4	1.7	2.6	- 25
$ar{n}_e$ -	3.9	2.4	2.5	2.6	2.9	2.5	- 20
$P_{l,th}$ -	5.6	4.1	5.3	4.3	5.7	4.2	- 15
$R_{geo}$ -	4.7	4.1	4.1	4.2	4	4.2	13
$\kappa_a$ -	2.8	1.9	2.3	2	2.5	1.9	- 10
$\epsilon$ -	2	2.6	2.4	2.3	2.1	2.5	- 5
$M_{eff}$ -					1.5	1.4	
	TOBL	STDBS	D. Bild	g Big	Small Small	Small	- 0
C	5,	Secreasin	Mattectin	creasing.	affecting	<b>)</b>	









# WORKFLOW

WHICH REGISTERS IN STDB5 ARE DECREASING  $\alpha_R$ ?

Depends on how you classify / label them.

IS THIS
DECREASE
DUE TO
DATA
ISSUES?

Yes, but not entirely.

CAN WE PREDICT WHETHER A NEW REGISTER WILL DECREASE  $\alpha_R$ ?

RELATE THE FINDINGS
TOKAMAK
CHARs & PHYSICS









# FEATURES OF INTEREST: Entropy

$$E = -\sum_{i,j=1}^{N} \left[ S_{ij} \log(S_{ij}) + \left( 1 - S_{ij} \right) \log(1 - S_{ij}) \right]$$

$$E = -\sum_{\substack{i,j=1\\i\neq j}}^{N} \left[ S_{ij} \log(S_{ij}) + \left( 1 - S_{ij} \right) \log(1 - S_{ij}) \right]$$

$$S_{ij} = \exp\left( -\gamma \cdot D_{ij} \right), \quad D_{ij} = \left[ \sum_{k=1}^{M} \left( \frac{x_{ik} - x_{jk}}{max(F_k) - min(F_k)} \right)^2 \right]^{1/2}$$

$$S_{ij} = \frac{1}{M} \sum_{k=1}^{M} \delta_{ij}(x^k); \quad \text{with } \delta_{ij}(x^k) = \begin{cases} 1, \text{ if } x_i^k = x_j^k \\ 0, \text{ if } x_i^k \neq x_j^k \end{cases}$$

columns 35 | 12

#### **IDEA**

Keep features that increased the entropy when removed.





[M. Dash and H. Liu, 2020]





# FEATURES OF INTEREST: Entropy + Low MCL

columns

35 | 12

$$E = -\sum_{i,j=1}^{N} \left[ S_{ij} \log(S_{ij}) + \left( 1 - S_{ij} \right) \log(1 - S_{ij}) \right]$$

$$E = -\sum_{\substack{i,j=1\\i\neq j}}^{N} \left[ S_{ij} \log(S_{ij}) + \left( 1 - S_{ij} \right) \log(1 - S_{ij}) \right]$$

$$S_{ij} = \exp\left( -\gamma \cdot D_{ij} \right), \quad D_{ij} = \left[ \sum_{k=1}^{M} \left( \frac{x_{ik} - x_{jk}}{max(F_k) - min(F_k)} \right)^2 \right]^{1/2}$$

$$S_{ij} = \frac{1}{M} \sum_{k=1}^{M} \delta_{ij}(x^k); \quad \text{with } \delta_{ij}(x^k) = \begin{cases} 1, \text{ if } x_i^k = x_j^k \\ 0, \text{ if } x_i^k \neq x_j^k \end{cases}$$

## **MULTICOLLINEARITY**

Removing features with high VIF and causing high condition indices.









# FEATURES OF INTEREST: Entropy + Low MCL

FEATURE	DESCRIPTION	FEATURE	DESCRIPTION
BETA	Plasma pressure normalised to <b>B</b> <sub>t</sub>	WFICFORM	Total fast ion energy due to ICRH estimated from approximate formula
RHOSTAR	Normalised Ion Gyroradius	WFFORM	Total fast ion energy due to NBI
NUSTAR	Normalised Ion collision frequency	ZEFFNEO	Line average plasma effective charge, from Bremsstrahlung
Q95	Plasma safety factor at the 95% poloidal flux surface	DWDIA	Time rate of change of the total plasma stored energy
PFLOSS	NBI power that is lost from the plasma through charge exchange and unconfined orbits		

Description of all columns in STDB5 found in [Princeton, 2021, http://arks.princeton.edu/ark:/88435/dsp01m900nx49h]



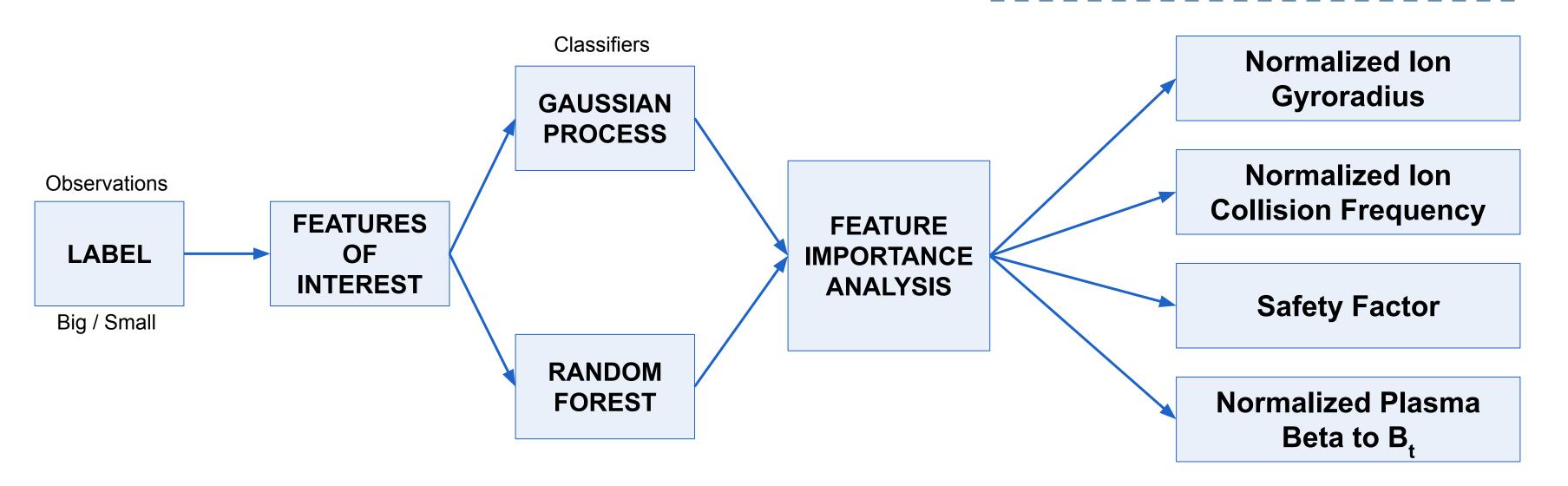






## FEATURES FOR PREDICTION

Best Performance with "big dataset"











## PREDICTIONS IN NEW DEVICES

9	$\overline{I_P}$	$B_t$	$\bar{n}_e$	$P_{l,th}$	$R_{geo}$	$\kappa_a$	$\epsilon$	$M_{eff}$	$ ho_*$	$\beta_t$	$ u_*$	995	$ au_{E,th}$
ITER	15	5.3	1.03	87	6.2	1.8	0.32	2.5	0.002	2.24	0.014	3	3.5
SPARC	8.7	12.2	3.1	25	1.85	1.97	0.31	2.5	0.003	1.20	0.03	3.2	0.77

## UNAFFECTED

$$\tau_{E,2.16} = 0.06 \cdot I_p^{0.78} \cdot B_t^{0.24} \cdot \bar{n}_e^{0.41} \cdot P_{l,th}^{-0.75} \cdot R_{geo}^{2.16} \cdot \kappa_a^{0.5} \cdot \epsilon^{0.79} \cdot M_{eff}^{0.22}$$









# WORKFLOW

WHICH REGISTERS IN STDB5 ARE DECREASING  $\alpha_R$ ?

Depends on how you classify / label them.

IS THIS
DECREASE
DUE TO
DATA
ISSUES?

Yes, but not entirely.

CAN WE PREDICT WHETHER A NEW REGISTER WILL DECREASE  $\alpha_R$ ?

Yes, with only four variables.

RELATE THE FINDINGS
TOKAMAK
CHARs & PHYSICS



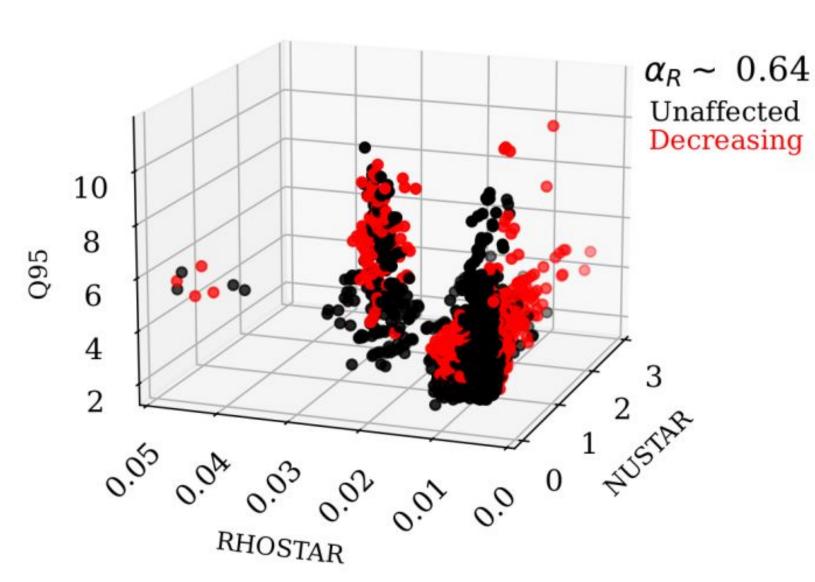


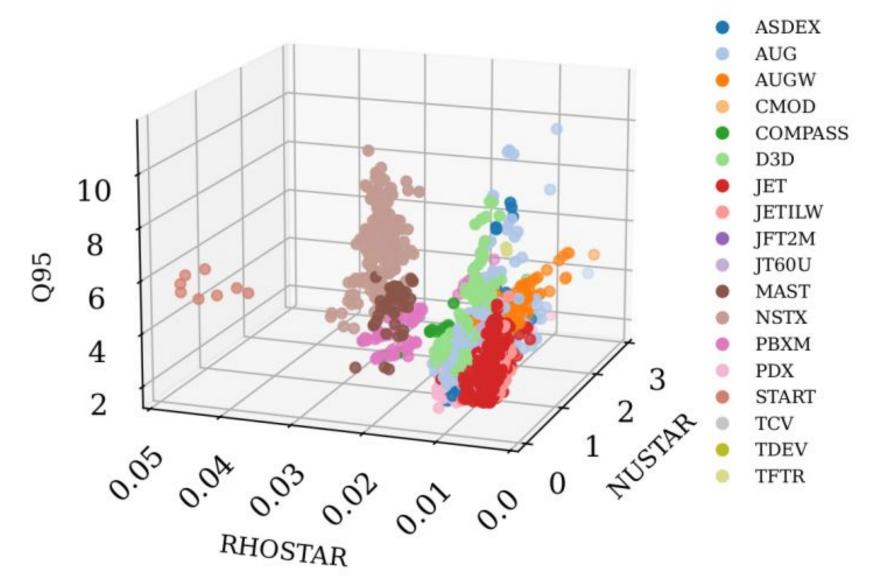




## **Characteristic Clusters: main features**

## **New STDB5**







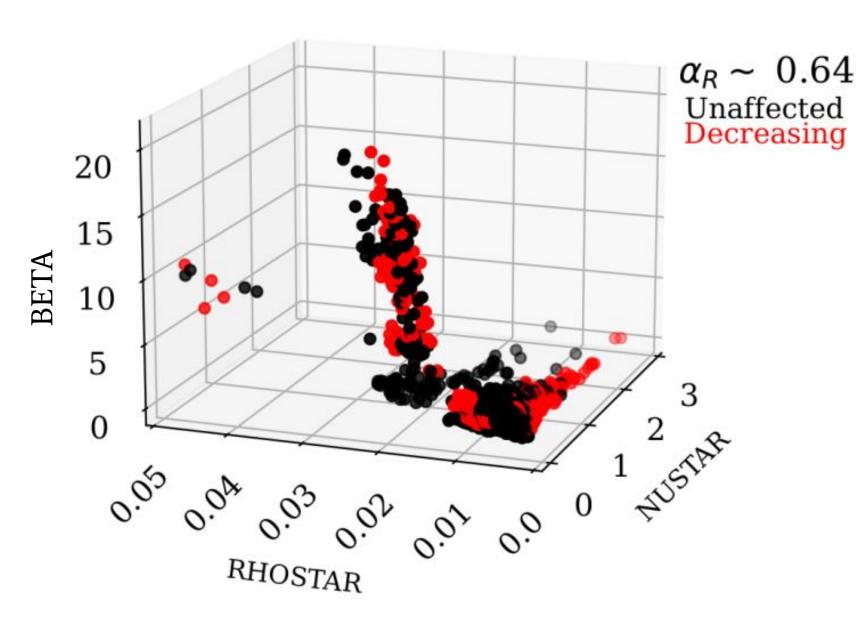


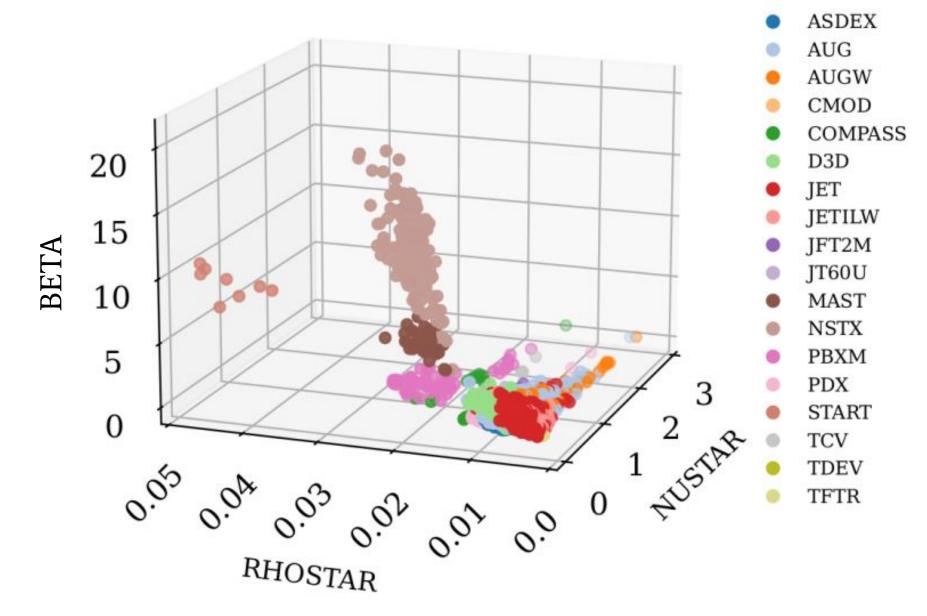




## **Characteristic Clusters: main features**

## **New STDB5**













# Dimensionless Scaling and the Normalized Ion Gyroradius

$$\omega_i \cdot \tau_{E,the} \propto \rho_*^{-(2+\chi_{\rho}')} \cdot F(\nu_*, \beta_t, \{p_i\})$$
 with  $0 \le \chi_{\rho}' \le 1$   $\chi_{\rho}' = 0.9 \pm 0.3$ 

H-mode

characteristic turbulence scale length  $~\ell pprox 
ho_*^{\chi_{
ho}'} \cdot a^{1-\chi_{
ho}'}$ 

				D1g	b1g_as		small_ds		
		STDB2	STDB5	Decr.	Unaff.	Decr.	Unaff.		
	$\chi_{\rho}$	-3.09	-1.80	-1.31	-2.63	-1.5	-2.08		
	$\chi'_{\rho}$	1.09	0.2	0.69	0.63	-0.5	0.08		
ITER	ℓ [m]	0.0012	7.99	227.43	0.026	64.99	1.15		
SPARC	ℓ [m]	0.0017	1.67	22.33	0.019	8.46	0.37		

[G. Verdoolaege et al., 2021] [I. P. E. G. et al., 1999]



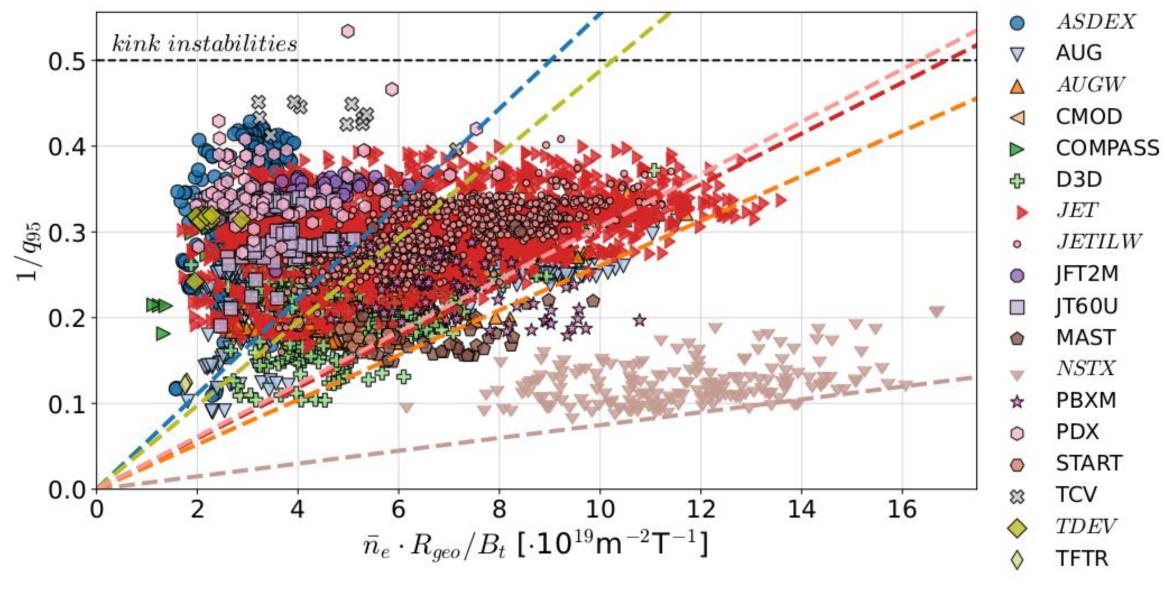




## **DENSITY LIMIT**

## STDB5

$$n_{GW} = \frac{10 \cdot I_p}{\pi a^2}$$



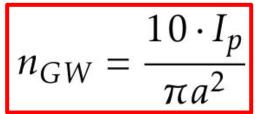




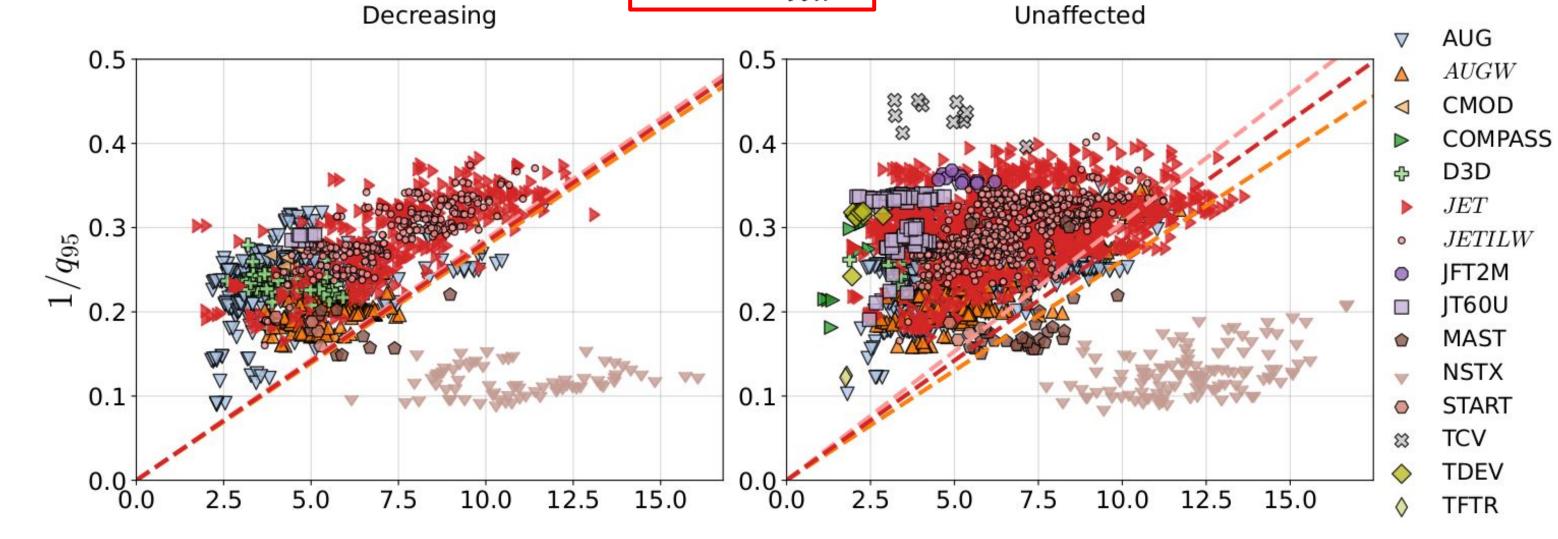




## **DENSITY LIMIT**



## **New STDB5**













# Conclusions and Further Work







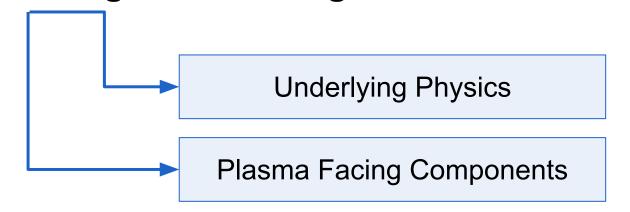


## **CONCLUSIONS**

## **FURTHER WORK**

- Multicollinearity is not the sole factor influencing α<sub>R</sub>.
- It is possible to predict influencing observations with
  - $\rho_* \beta_t \nu_* q_{95}$
- Characteristic clusters for spherical tokamaks.
- It is possible to tell if a dataset lacks of representative physics.
- Most of the observations that surpassed the GW limit are classified as unaffecting.
- ITER and SPARC are expected to follow a scaling law similar to the 1998 scaling (great news!).

- Obtain more data on spherical tokamaks.
- Discern between spherical and non spherical.
- Take time-series data of relevant shots and subject to other ML algorithms; e.g. surrogate modelling.











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- **f** Universiteit Gent
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- @ugent
- in Ghent University

# Thank you! Questions?









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