

I : Effective Field Theory.

Why EFT?

For some integrals are hard to calculate directly.

Or itself contains divergence.

$$\text{e.g. } \text{---} \circlearrowleft \text{---} = \int d^4k \cdot \frac{1}{(k^2 - m^2)(k'^2 - m^2)}$$

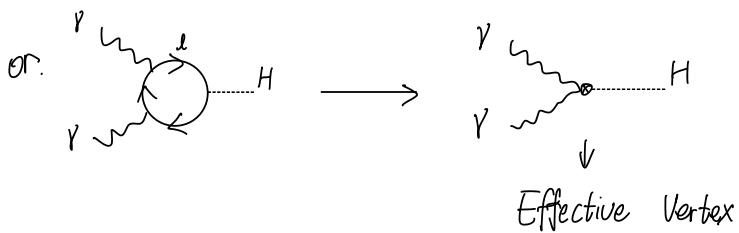
threshold.

$$= 4\pi \cdot \int \frac{k^2 dk}{(k^2 - m^2)^2}$$

\rightarrow divergent.

$$\text{or. } \text{---} \begin{matrix} p \\ k \\ l \end{matrix} \text{---} \propto \int d^4k \frac{1}{k^2 (k+l)^2 (k+p)^2} \rightarrow \text{hard to calculate.}$$

\rightarrow contain angular analysis.

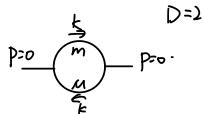


1. Regions Strategy.

Ex. 1. $I = \int_0^{+\infty} dk \frac{k}{(k^2+m^2)(k^2+M^2)}$ (Assuming $M \gg m$)

$$\begin{aligned} &= \frac{1}{2} \int_0^{+\infty} \frac{dk^2}{(k^2+m^2)(k^2+M^2)} \\ &= \frac{1}{2} \int_0^{+\infty} \frac{dk^2}{\left(k^2 + \frac{m^2+M^2}{2}\right)^2 - \left(\frac{M-m}{2}\right)^2} \\ &= \frac{1}{2} \cdot \frac{2}{M^2-m^2} \cdot \int_{\frac{m^2+M^2}{2}}^{+\infty} \frac{dx}{x^2-1} \\ &= \frac{1}{m^2-M^2} \cdot \left(\frac{1}{2} \ln 1 - \frac{1}{2} \ln \frac{2M^2}{2m^2} \right) \\ &= \frac{\ln M - \ln m}{M^2 - m^2} \end{aligned}$$

$\int \frac{dx}{x^2-1} = \frac{1}{2} \left[\frac{1}{x-1} - \frac{1}{x+1} \right] dx = \frac{1}{2} \ln \frac{x-1}{x+1} + C$



order of $N: 2$
order of $D: 4 \rightarrow$ convergent.

We can expand the result with $\frac{M}{m} \gg 1 \rightarrow$ Large Logs.

$$I = \frac{\ln \frac{M}{m}}{M^2} \frac{1}{1 - \frac{m^2}{M^2}} = \frac{\ln \frac{M}{m}}{M^2} \left(1 + \frac{m^2}{M^2} + \frac{m^4}{M^4} + \dots \right)$$

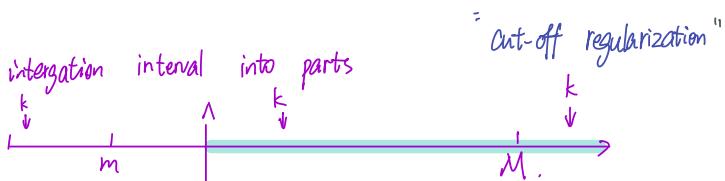
But, what if we just expand the integrand? (Assuming m is quite small)

$$\Rightarrow \frac{k}{(k^2+m^2)(k^2+M^2)} = \frac{k}{k^2(k^2+M^2)} \left(1 - \frac{m^2}{k^2} + \frac{m^4}{k^4} + \dots \right)$$

then integrate it from 0 to $+\infty$, every term divergent.

Why?

Divide the integration interval into parts



The expansion above only works for $k \gg m$ (the Miku Blue region).

We can expand the integrands with $\frac{k}{m} \ll 1$, it is valid (since $m \ll M$).

$$\Rightarrow I = \underbrace{\int_0^\Lambda dk \frac{k}{(k^2+m^2)(k^2+M^2)}}_{\substack{\downarrow \text{low-energy / infra-red} \\ (\text{with } m \sim k \ll M) \text{ I}_1}} + \underbrace{\int_\Lambda^{+\infty} dk \frac{k}{(k^2+m^2)(k^2+M^2)}}_{\substack{\uparrow \text{high-energy / ultra-violet} \\ (\text{with } m \ll k \ll M) \text{ I}_2}}$$

$$\Rightarrow I_1 = \int_0^\Lambda dk \frac{k}{(k^2+m^2)(k^2+M^2)} = \int_0^\Lambda dk \frac{k}{(k^2+m^2) \cdot M^2} \left(1 - \frac{k^2}{M^2} + \frac{k^4}{M^4} + \dots \right) \rightarrow \text{All terms are convergent}$$

$$I_2 = \int_\Lambda^{+\infty} dk \frac{k}{(k^2+m^2)(k^2+M^2)} = \int_\Lambda^{+\infty} dk \frac{k}{k^2(k^2+M^2)} \left(1 - \frac{m^2}{k^2} + \frac{m^4}{k^4} + \dots \right) \rightarrow \text{Also convergent}$$

Infra-red Cut-off ($m \ll \Lambda \ll M$)

So, we can calculate I_1, I_2 independently. To emphasize, the final answer couldn't contain " λ " terms.

$$I_1 = \frac{1}{2M^2} \ln(1 + \frac{\lambda^2}{m^2}) - \frac{1}{2M^4} \left(\lambda^2 - m^2 \ln(1 + \frac{\lambda^2}{m^2}) \right) + O(\frac{\lambda^4}{M^6}, \frac{\lambda^4 m^2}{M^6}, \frac{m^4}{M^6} \ln(1 + \frac{\lambda^2}{m^2}))$$

$$= \frac{M^2 + m^2}{2M^4} \ln(1 + \frac{\lambda^2}{m^2}) - \frac{\lambda^2}{2M^4}$$

$$I_2 = \frac{1}{2m^2} \ln(1 + \frac{M^2}{\lambda^2}) - \frac{m^2}{2\lambda^2 M^4} \left(M^2 - \lambda^2 \ln(1 + \frac{M^2}{\lambda^2}) \right) + O(\frac{m^4}{M^4 \lambda^2}, \frac{m^4}{M^6} \ln(1 + \frac{M^2}{\lambda^2}))$$

$$= \frac{M^2 + m^2}{2M^4} \ln(1 + \frac{M^2}{\lambda^2}) - \frac{m^2}{2\lambda^2 M^2} + O(\dots)$$

Second expand I_1, I_2 with $m \ll \lambda \ll M$; $\ln(1 + \frac{\lambda^2}{m^2}) = \ln \frac{\lambda^2}{m^2} + \ln(1 + \frac{M^2}{\lambda^2})$

$$I_1 = \frac{M^2 + m^2}{2M^4} \ln \frac{\lambda^2}{m^2} + \frac{M^2 + m^2}{2M^4} \cdot \frac{m^2}{\lambda^2} - \frac{\lambda^2}{2M^4}$$

$$I_2 = \frac{M^2 + m^2}{2M^4} \ln \frac{M^2}{\lambda^2} + \frac{M^2 + m^2}{2M^4} \cdot \frac{\lambda^2}{M^2} - \frac{m^2}{2\lambda^2 M^2}$$

$$\Rightarrow I_1 + I_2 = \frac{\ln \frac{M}{m}}{M^2} + \frac{m^2}{M^4} \ln \frac{M}{m} + \frac{m^2}{2M^4} \left(\frac{m^2}{\lambda^2} + \frac{\lambda^2}{M^2} \right)$$

high order (can be cancelled by calculating high order terms)
leading order ; next-leading order. \hookrightarrow do not contain " λ ".

$$\Rightarrow I = I_1 + I_2 \quad (\text{equals in terms})$$

Goods: "Physical", easy to calculate, explicit.

Bads: needs to deal with more integrals, break some symmetries (e.g. fringe symmetry)

"Dimensional regularization"

$$I = \int_0^{+\infty} dk \frac{k}{(k^2 + m^2)(k^2 + M^2)} \xrightarrow{\text{(t'Hooft scale)}} \mu^\varepsilon \int_0^{+\infty} d^{\varepsilon} k \frac{k}{(k^2 + m^2)(k^2 + M^2)}$$

mass dimension counter term dimension

using $d^N k = (N-1)! dk \cdot d\Omega_N$. ; where Ω_N is the N dimension angular variables / sphere surface,

here we ignore the angular contributions (Actually, the contribution of $\Omega_{4-\varepsilon}$ is nearly "1")

$$\left(\Omega_N = \frac{2\pi^{\frac{N}{2}}}{\Gamma(\frac{N}{2})} \Rightarrow \Omega_{4-\varepsilon} = \frac{2\pi^{\frac{\varepsilon}{2}}}{\Gamma(\frac{1-\varepsilon}{2})} = 2 \rightarrow \text{for } k > 0 \Rightarrow \int d\Omega_{4-\varepsilon} = \frac{1}{2} \times \Omega_{4-\varepsilon} = 1 \right)$$

$$\Rightarrow I = \mu^\varepsilon \int_0^{+\infty} dk \cdot k^{-\varepsilon} \frac{k}{(k^2 + m^2)(k^2 + M^2)} \quad (\text{This } \varepsilon \text{ is a hand-put and arbitrary parameter, just make integrals convergent.})$$

$$I_{(1)} = \mu^\varepsilon \int_0^{+\infty} dk k^{-\varepsilon} \frac{k}{(k^2 + m^2) M^2} \left(1 - \frac{k^2}{M^2} + \frac{k^4}{M^4} + \dots \right) \xrightarrow{\text{UV safe}}$$

\Rightarrow At least the first term is convergent.

$$I_{(11)} = \mu^\varepsilon \int_0^{+\infty} dk k^{-\varepsilon} \frac{k}{k^2 (k^2 + M^2)} \left(1 - \frac{m^2}{k^2} + \frac{m^4}{k^4} + \dots \right) \xrightarrow{\text{IR safe.}}$$

At first term:

$$I_{(1)} = \frac{1}{M^2 \varepsilon} - \frac{\ln \frac{M}{m}}{M^2} + O(\varepsilon) \quad \Rightarrow \quad I = I_{(1)} + I_{(11)} = \frac{1}{M^2} \ln \frac{M}{m} \quad \rightarrow \text{leading order in expansion of "2".}$$

$$I_{(11)} = -\frac{1}{M^2 \varepsilon} + \frac{\ln \frac{M}{m}}{M^2} + O(\varepsilon)$$

Actually, we do the integral without considering the convergence radius of MacLaurin series.
 That is the reason why the other terms are divergent. ↳ the "power" of regularization is not enough.

All also, we integrate the high-energy part over IR part and vice versa, but this part is trivial.

$$\begin{aligned} \text{Since, } I_{(1)} &= \int_0^\Lambda dk k^{-\varepsilon} \frac{k}{(k^2 + m^2)^{1/2}} \left(1 - \frac{k^2}{m^2} + \frac{k^4}{m^4} + \dots\right) \\ &= \left[\int_0^{t\Lambda} dk - \int_\Lambda^{t\Lambda} dk \right] k^{-\varepsilon} \frac{k}{(k^2 + m^2)^{1/2}} \left(1 - \frac{k^2}{m^2} + \frac{k^4}{m^4} + \dots\right) \\ &= I_{(1)} - R_{(1)} \xrightarrow{\text{Rest}} \end{aligned}$$

$$R_{(1)} = \int_\Lambda^{t\Lambda} dk k^{-\varepsilon} \frac{k}{k^2 + m^2} \left(1 - \frac{m^2}{k^2} - \frac{k^2}{m^2} + \dots\right) \quad (k \gg m)$$

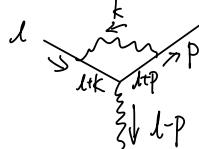
do the same, we get $R_{(1)} = \int_0^\Lambda dk k^{-\varepsilon} \frac{k}{k^2 + m^2} \left(1 - \frac{m^2}{k^2} - \frac{k^2}{m^2} + \dots\right)$
 $\Rightarrow R_{(1)} + R_{(1)} = \int_0^{t\Lambda} dk k^{-\varepsilon} \frac{k}{k^2 + m^2} \left(1 - \frac{m^2}{k^2} - \frac{k^2}{m^2} + \dots\right) \rightarrow \text{do not depend on the cut-off.}$

\hookrightarrow gives the scaleless integrals. $\Rightarrow R_{(1)} + R_{(1)} = 0$
 "Zero-bin" contribution

one can check by integrating out this integral: (1st term: $\int_0^\Lambda dk \frac{k^{1-\varepsilon}}{m^2} + \int_\Lambda^{t\Lambda} dk \frac{k^{1-\varepsilon}}{m^2} = -\frac{\Lambda^{1-\varepsilon}}{2\varepsilon m^2} + \frac{\Lambda^{1-\varepsilon}}{2\varepsilon m^2} = 0 \quad \checkmark$)

It seems neither easy nor explicit, but it's useful for the intrinsic divergent integrals.

2. Sudakov Problem.



The general strategy to obtain the expansion of a given Feynman diagram in a given kinematic limit is:

(i) Identify all regions of the integrand which leads to singularities (UV/IR, hard/soft ...)

(ii) Expand the integrand in each region, then integrate the expansion over the full phase space. (I_1, I_2, \dots)

(iii) Add all the results of different region's integrations to obtain the expansion of the original full integral. ($I = I_1 + I_2 + \dots$)

The Feynman diagrams above can write as (with the dimension of "d")

$$I = i \pi^{\frac{d}{2}} \mu^{4-d} \int dk^d \frac{1}{(k^2 + i\epsilon)[(k+q)^2 + i\epsilon][(k+p)^2 + i\epsilon]}$$

(it has IR divergence when $k \rightarrow 0$)

* Here we introduce the light-cone coordinate.

the four basis : $n_u = (1, 0, 0, 1)$; $\bar{n}_u = (1, 0, 0, -1)$, $n_u^x = (0, 1, 0, 0)$, $n_u^y = (0, 0, 1, 0)$.

usually combine n^x and n^y together as n_\perp

we have $n \cdot \bar{n} = \bar{n} \cdot n = 2$; $n \cdot n = \bar{n} \cdot \bar{n} = 0$; $n \cdot n_\perp = \bar{n} \cdot n_\perp = 0$

an arbitrary vector, let : $k_u = A n_u + B \bar{n}_u + C n_u^\perp$

$$k \cdot n = 2B \Rightarrow B = \frac{k \cdot n}{2}$$

$$k \cdot \bar{n} = 2A \Rightarrow A = \frac{k \cdot \bar{n}}{2}$$

$$k \cdot n^\perp = C \Rightarrow C = (k_x, k_y) = k_\perp \Rightarrow \text{Combine } k \pm n_u^\perp = k_u^\perp = (0, k_x, k_y, 0)$$

$$\Rightarrow k_u = k_n \frac{n_u}{2} + k_{\bar{n}} \frac{\bar{n}_u}{2} + k_\perp^\perp \equiv k_+ \frac{n_u}{2} + k_- \frac{\bar{n}_u}{2} + k_\perp^\perp \Rightarrow (k_+, k_-, k_\perp)$$

Light-cone coordinate is usually used in large momentum or near light speed problems.

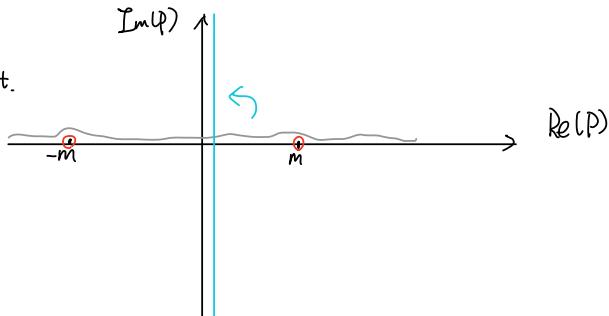
$$\text{In the light-cone coordinate } P \cdot g = \frac{1}{2}(P_+ g_+ + P_- g_-) + P_\perp^2 = \frac{1}{2}[(P \cdot \bar{n})(g \cdot n) + (g \cdot \bar{n})(P \cdot n)] + P_\perp^2$$

Introduce the following notation:

$$L^2 \equiv \lambda^2 - i\epsilon \quad ; \quad P^2 \equiv -p^2 - i\epsilon \quad ; \quad Q^2 \equiv -(l-p)^2 - i\epsilon \quad \rightarrow \text{the inject momentum}$$

Why 2 ?

$\text{Im}(P)$



Then we can introduce our expansion parameter λ

$$\lambda^2 \sim \frac{P^2}{Q^2} \sim \frac{L^2}{Q^2} \quad \& \quad P^2 \sim L^2 \sim \lambda^2 Q^2$$

choose the large component of momentum flow as the reference vector.

$$P^\mu \sim Q \frac{m}{2} ; \quad l^\mu \sim Q \frac{\bar{n}^\mu}{2}$$

$$\Rightarrow P^\mu \sim (\lambda^2, 1, \lambda) Q \quad ; \quad l^\mu \sim (1, \lambda^2, \lambda) Q \quad \rightarrow \text{Arbitrary, like } g \sim (\lambda^a, \lambda^b, \lambda^c) Q \\ \text{but it's necessary to make sure } g^2 \sim \lambda^n Q^n \\ \Leftrightarrow a+b=2c$$

in this Sudakov problem, only four regions below contributes.

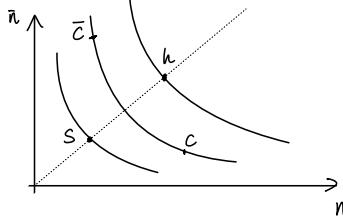
(i) Hard region (h) : $k_h^{\mu} \sim (1, 1, 1) Q$

(ii) Region Collinear to p (c) : $k_c^{\mu} \sim (\lambda^2, 1, \lambda) Q$

(iii) Region Collinear to ℓ (\bar{c}) : $k_{\bar{c}}^{\mu} \sim (1, \lambda^2, \lambda) Q$ (in two-body problem, p & ℓ are always back-to-back so it's also call anti-collinear)

(iv) Soft region (s) : $k_s^{\mu} \sim (\lambda, \lambda, \lambda) Q$

| Region | Theory scale |
|-----------|---|
| Hard | $\leftrightarrow QCD \leftrightarrow Q$ |
| Collinear | $\leftrightarrow SCET \leftrightarrow M$ |
| soft | $\leftrightarrow SET \leftrightarrow \frac{Q}{\lambda}$ |



why not soft region looks like $k_s^{\mu} \sim (\lambda, \lambda, \lambda) Q$ ↑ semi-hard

$$\text{consider } (k_s + k_0)^2 = (k_s + k_0)_L (k_s + k_0)_T + (k_s + k_0)_L^2 \\ = k_s^+ k_s^- + k_s^+ k_T^- + k_T^+ k_s^- + k_T^+ k_T^- \sim Q \lambda^2$$

$\uparrow \text{out} \quad \uparrow \text{out} \quad \uparrow \text{out} \quad \uparrow \text{out}$

since the observable of collinear momentum is $L^+ \sim \lambda Q^2$
there should be $(k_s + k_0)^2 \sim L^+ \sim Q \lambda^2$
so it breaks this observation order. If $k_s^{\mu} \sim (\lambda, \lambda, \lambda) Q$,
then $(k_s + k_0)^2 \sim k_s^+ k_s^- \sim Q \lambda^2$ reverses order.

$$SCET_I: P_s^+ \sim \lambda^4 \quad (\text{ultra-soft}) \\ SCET_{II}: P_s^+ \sim \lambda^2 \quad (\text{semi-hard})$$

We have $k_h^{\mu} \sim Q^2$; $k_c^{\mu} \sim k_{\bar{c}}^{\mu} \sim \lambda^2 Q^2$; $k_s^{\mu} \sim \lambda^4 Q^2$ ↑ ultra-soft
order hierarchy $\Rightarrow SCET_I$

Only those four contributes, other contributions vanished.

take the semi-hard region as the example. $(k_l^{\mu} = (1, 0, 0) \frac{Q}{\lambda}; P_c^{\mu} = (0, 1, 0) \frac{Q}{\lambda})$

$$I_{sh} = i \pi^{\frac{d}{2}} \mu^{4-d} \int d^d k \frac{1}{(k^2 + i0)(2k \cdot l + i0)(2k \cdot p - i0)} \\ = i \pi^{\frac{d}{2}} \mu^{4-d} \int d^d k dx dy dz \frac{2}{(k^2 + 2k(l + \lambda p) + \lambda^2 p^2)^{\frac{3-d}{2}}} \\ = i \pi^{\frac{d}{2}} \mu^{4-d} \frac{4\Gamma(\frac{3-d}{2})}{\Gamma(3)} \int dx dy dz \frac{2}{((x l + \lambda y p)^2 + \lambda^2 p^2)^{\frac{3-d}{2}}} \\ = i \pi^{\frac{d}{2}} \mu^{4-d} P(l + \lambda p)^{2-d} \int_0^\infty dx dy dz (x l + \lambda y p)^{-d} \\ = 0$$

— Feynman Parameter 2

$$\frac{1}{ABC} = \int_0^\infty dx \int_0^\infty dy \int_0^\infty dz \frac{2}{(A + Bx + Cz)^3}$$

(I) Hard Region:

$$(k + l)^2 = k^2 + 2l \cdot k_- + O(\lambda)$$

$$(k + p)^2 = k^2 + 2p \cdot k_+ + O(\lambda)$$

$$I_h = i \pi^{\frac{d}{2}} \mu^{4-d} \int d^d k \frac{1}{k^2 (k^2 + 2k \cdot l + i0) (k^2 + 2k \cdot p - i0)}$$

$$= i \pi^{\frac{d}{2}} \mu^{4-d} \int d^d k dx dy \frac{2}{[k^2 + 2k \cdot (x l + y p) + (x l + y p)^2 - xy l \cdot p]^{-d}}$$

$$= i \pi^{\frac{d}{2}} \mu^{4-d} \int_0^1 dy \cdot (1) \cdot x l \frac{2}{(x l + y p)^2 \Gamma(3-\frac{d}{2})} \left(\frac{1}{2xy l \cdot p} \right)^{3-\frac{d}{2}}$$

$$= \frac{m^d \Gamma(1+\varepsilon)}{(2l+p)^{1+\varepsilon}} \int_0^1 dy \frac{2}{(x l + y p)^{1-\varepsilon}}$$

$$= \frac{\Gamma(1+\varepsilon)}{2l+p} \cdot \left(\frac{m^2}{2l+p} \right)^\varepsilon \cdot \frac{1}{\varepsilon^2} = \frac{\Gamma(1+\varepsilon)}{Q^2} \left(\frac{1}{\varepsilon^2} + \frac{1}{\varepsilon} \ln \frac{1}{\varepsilon^2} + \frac{1}{2} \ln^2 \frac{1}{\varepsilon^2} + O(\varepsilon) \right)$$

— Feynman Parameter 1.

$$\frac{1}{ABC} = \int_0^1 dx \int_0^1 dy \int_0^1 dz \frac{2}{(xA + x_2 b + x_3 c)^3}$$

$$= \int_0^1 dy \frac{2}{((x-y)A + xB + yC)^3}$$

$$\begin{cases} l_+ \sim Q \\ p_- \sim Q \end{cases} \Rightarrow 2l \cdot p = 2x \frac{1}{2} Q^2 = Q^2$$

Or we can change our parameter to: $x_1 = y; \underline{x_2 = x-y}; x_3 = 1-x$
 \downarrow
 $\text{gives } y \propto x$

$$I_h = M^{4\epsilon} \int_0^1 dy \Gamma(1+\epsilon) \left(\frac{1}{2M^2(1-y)} \right)^{1-\epsilon}$$

$$= \frac{\Gamma(1+\epsilon)}{Q^2} \left(\frac{M^2}{Q^2} \right)^\epsilon \cdot \left[\int_0^1 dx \int_0^x dy \left[(1+x)y \right]^{-1-\epsilon} \right] = - \frac{\Gamma(-\epsilon) \Gamma(-\epsilon)}{\epsilon \Gamma(1-2\epsilon)} = \frac{\Gamma^2(-\epsilon)}{\Gamma(1-2\epsilon)}$$

$$= \frac{\Gamma(1+\epsilon)}{Q^2} \left(\frac{1}{\epsilon^2} + \frac{1}{\epsilon} \ln \frac{M^2}{Q^2} + \frac{1}{2} \ln^2 \frac{M^2}{Q^2} \right) \boxed{\frac{\pi^2}{6}} + O(\epsilon) \quad \text{alternative form}$$

$$= \frac{1}{Q^2} \left(\frac{1}{\epsilon^2} + \frac{1}{\epsilon} \ln \frac{M^2}{Q^2} + \frac{1}{2} \ln^2 \frac{M^2}{Q^2} - \frac{\pi^2}{12} + O(\epsilon) \right)$$

* So we can see that different parameter choice can derive different kind of constant term, but there scale-relying terms and dimensional divergence terms are all the same.

(II) Collinear Region. ($k \sim (x^2, 1, \lambda)Q$, or $k \parallel p$)

$$(k+\lambda)^2 = 2k \cdot k_t + O(\lambda^2) \quad \rightarrow \text{Only consider the leading order.}$$

$$(k+p)^2 = k^2 + 2k \cdot p + p^2 = O(\lambda^2)$$

$$I_c = i \pi^{\frac{d}{2}} \mu^{4-d} \int dk \frac{1}{k^2 (2k \cdot k_t) (k^2 + 2k \cdot p + p^2)} \quad \text{- Feynman Parameter 2}$$

$$= i \pi^{\frac{d}{2}} \mu^{4-d} \int_0^{+\infty} dy dy \int dk \frac{2}{(k^2 + 2k \cdot k_t + y^2 + 2y k \cdot p - y p^2)^3}$$

$$= i \pi^{\frac{d}{2}} \mu^{4-d} \int_0^{+\infty} dy dy \int dk \frac{2}{\{(1+y) \left[k^2 + 2k \cdot \left(\frac{k_t + y p}{1+y} \right) - \frac{y p^2}{1+y} \right]\}^3}$$

$$= i \pi^{\frac{d}{2}} \mu^{4-d} \frac{(2\pi)^d \chi^{\frac{d}{2}} \Gamma(3-\frac{d}{2}) \Gamma^2}{\Gamma(3)} \int_0^{+\infty} dy dy \frac{1}{(1+y)^3 \left(\frac{-y p^2}{(1+y)} + \frac{2xy k \cdot p + y p^2}{(1+y)^2} \right)^{3-\frac{d}{2}}}$$

$$= M^{2\epsilon} \Gamma(1+\epsilon) \int_0^{+\infty} dy dy \frac{1}{(1+y)^{1+2\epsilon} \left[-y p^2 + 2xy k \cdot p \right]^{\frac{d}{2}}}$$

$$= M^{2\epsilon} \Gamma(1+\epsilon) \int_0^{+\infty} dy dy \frac{1}{(1+y)^{1+2\epsilon} y^{\frac{d}{2}}} \left[\frac{1}{P^2 + \alpha Q^2} \right]^{\frac{d}{2}} \quad \text{let } r = \frac{Q^2}{P^2}$$

$$= \left(\frac{M^2}{P^2} \right)^\epsilon \frac{\Gamma(1+\epsilon)}{P^2} \int_0^{+\infty} dy \frac{1}{(1+y)^{1+2\epsilon} y^{\frac{d}{2}}} \int_0^{+\infty} dx \frac{1}{(1+rxy)^{1+\epsilon}}$$

$$= \frac{\Gamma(1+\epsilon)}{P^2} \left(\frac{M^2}{P^2} \right)^\epsilon \frac{\Gamma(-\epsilon) \Gamma(-\epsilon)}{\Gamma(1-2\epsilon)} \cdot \frac{1}{\epsilon \cdot r}$$

$$= \frac{\Gamma(1+\epsilon)}{Q^2} \left(\frac{M^2}{P^2} \right)^\epsilon \left[- \frac{\Gamma^2(-\epsilon)}{\Gamma(1-2\epsilon)} \right]$$

$$= \frac{\Gamma(1+\epsilon)}{Q^2} \left(-\frac{1}{\epsilon^2} - \frac{1}{\epsilon} \ln \frac{M^2}{P^2} - \frac{1}{2} \ln^2 \frac{M^2}{P^2} + \frac{\pi^2}{6} + O(\epsilon) \right) \quad \text{Collinear scale}$$

$$AF = \frac{1}{Q^2} \left(-\frac{1}{\epsilon^2} - \frac{1}{\epsilon} \ln \frac{M^2}{P^2} - \frac{1}{2} \ln^2 \frac{M^2}{P^2} + \frac{\pi^2}{12} \right)$$

Similarly:

$$I_c = \frac{\Gamma(1+\epsilon)}{Q^2} \left(-\frac{1}{\epsilon^2} - \frac{1}{\epsilon} \ln \frac{M^2}{L^2} - \frac{1}{2} \ln^2 \frac{M^2}{L^2} + \frac{\pi^2}{6} + O(\epsilon) \right)$$

$$AF = \frac{1}{Q^2} \left(-\frac{1}{\epsilon^2} - \frac{1}{\epsilon} \ln \frac{M^2}{L^2} - \frac{1}{2} \ln^2 \frac{M^2}{L^2} + \frac{\pi^2}{12} \right)$$

(III) Soft Region

$$(k+p)^2 = 2k \cdot p + p^2 + O(\lambda^3)$$

$$(k+\lambda)^2 = 2k \cdot \lambda + \lambda^2 + O(\lambda^3)$$

$$I_s = i\pi^{-\frac{d}{2}} \mu^{4-d} \int dk \frac{1}{k^2 (2k \cdot l + l^2) (2k \cdot p + p^2)}$$

$$= i\pi^{-\frac{d}{2}} \mu^{4-d} \int dk \int_0^{+\infty} dy \frac{1}{(k^2 + 2xk \cdot l + l^2 + 2y k \cdot p + y^2)^3}$$

$$= \mu^{2\varepsilon} P(1+\varepsilon) \int_0^{+\infty} dx dy \frac{1}{(-(\lambda x^2 + y p^2) + 2xy \lambda + p^2)^{1+\varepsilon}}$$

$$= \mu^{2\varepsilon} \frac{P(1+\varepsilon)}{L^2 p^2} \int_0^{+\infty} dx dy (x + y + \alpha xy)^{-1-\varepsilon}$$

$$= \mu^{2\varepsilon} \frac{P(1+\varepsilon)}{L^2 p^2} \int_0^{+\infty} dx \int_0^{+\infty} du x^{-\varepsilon} (1 + u + \alpha xu)^{-1-\varepsilon} \quad \left\{ \begin{array}{l} u = \frac{y}{x} \\ dy = x du \end{array} \right. \text{ Parameter.}$$

$$= \mu^{2\varepsilon} \frac{P(1+\varepsilon)}{L^2 p^2} \int_0^{+\infty} dx \int_0^{+\infty} du x^{-\varepsilon} (1 + u + \alpha xu)^{-1-\varepsilon} \quad \left\{ \begin{array}{l} v = (1 + \alpha x)u \\ dv = (1 + \alpha x)du \end{array} \right.$$

$$= \mu^{2\varepsilon} \frac{P(1+\varepsilon)}{L^2 p^2} \int_0^{+\infty} \frac{x^{-\varepsilon}}{1 + \alpha x} dx \int_0^{+\infty} dv (1 + v)^{-1-\varepsilon}$$

$$= \mu^{2\varepsilon} \frac{P(1+\varepsilon)}{L^2 p^2} \bar{\alpha}^{-1-\varepsilon} P(1-\varepsilon) P(\varepsilon) \cdot \frac{1}{\varepsilon}$$

$$= \frac{P(1+\varepsilon)}{Q^2} \left(\frac{\mu^2 Q^2}{L^2 p^2} \right)^2 [-P(\varepsilon) \cdot P(-\varepsilon)] \quad AF = \frac{1}{Q^2} \left(\frac{1}{\varepsilon^2} + \frac{1}{\varepsilon} \ln \frac{\mu^2 Q^2}{L^2 p^2} + \frac{1}{2} \ln^2 \frac{\mu^2 Q^2}{L^2 p^2} + \frac{\pi^2}{6} \right)$$

$$= \frac{P(1+\varepsilon)}{Q^2} \left(\frac{1}{\varepsilon^2} + \frac{1}{\varepsilon} \ln \frac{\mu^2 Q^2}{L^2 p^2} + \frac{1}{2} \ln^2 \frac{\mu^2 Q^2}{L^2 p^2} + \frac{\pi^2}{6} + O(\varepsilon) \right) \quad A.F. \text{ just have differences in constants.}$$

$$I_h = \frac{P(1+\varepsilon)}{Q^2} \left(\cancel{\frac{1}{\varepsilon^2}} + \cancel{\frac{1}{\varepsilon} \ln \frac{\mu^2}{Q^2}} + \cancel{\frac{1}{2} \ln^2 \frac{\mu^2}{Q^2}} - \cancel{\frac{\pi^2}{6}} \right)$$

$$I_c = \frac{P(1+\varepsilon)}{Q^2} \left(-\cancel{\frac{1}{\varepsilon^2}} - \cancel{\frac{1}{2} \ln \frac{\mu^2}{p^2}} - \cancel{\frac{1}{2} \ln^2 \frac{\mu^2}{p^2}} + \cancel{\frac{\pi^2}{6}} \right)$$

$$I_{\bar{c}} = \frac{P(1+\varepsilon)}{Q^2} \left(-\cancel{\frac{1}{\varepsilon^2}} - \cancel{\frac{1}{2} \ln \frac{\mu^2}{L^2}} - \cancel{\frac{1}{2} \ln^2 \frac{\mu^2}{L^2}} + \cancel{\frac{\pi^2}{6}} \right)$$

$$I_s = \frac{P(1+\varepsilon)}{Q^2} \left(\cancel{\frac{1}{\varepsilon^2}} + \cancel{\frac{1}{2} \ln \frac{\mu^2 Q^2}{L^2 p^2}} + \cancel{\frac{1}{2} \ln^2 \frac{\mu^2 Q^2}{L^2 p^2}} + \cancel{\frac{\pi^2}{6}} \right)$$

UV-divergence

cancelled.

IR-divergence

$$\Rightarrow I = I_h + I_c + I_{\bar{c}} + I_s \xrightarrow{\text{Mathematically}} \frac{1}{Q^2} \left[\frac{\pi^2}{3} + \frac{1}{2} \left[(\ln \mu^2 - \ln Q^2)^2 - (\ln \mu^2 - \ln p^2)^2 - (\ln \mu^2 - \ln L^2)^2 + (\ln \mu^2 - \ln Q^2 - \ln p^2 - \ln L^2)^2 \right] \right]$$

Imagine its expanding form.

$\ln^2 \mu^2, \ln^2 p^2, \ln^2 Q^2$ terms vanished, $\ln \mu^2 \cdot \ln x^2$ terms were also cancelled

$$= \frac{1}{Q^2} \left[\frac{\pi^2}{3} + \frac{1}{2} (2 \ln^2 \mu^2 - 2 \ln Q^2 \ln p^2 - 2 \ln Q^2 \ln L^2 + 2 \ln^2 \mu^2) \right]$$

$$\underline{\text{Factorize}} \quad \frac{1}{Q^2} \left(\frac{\pi^2}{3} + \ln \frac{\mu^2}{L^2} \ln \frac{\mu^2}{p^2} + O(1) \right)$$

3. Massive Sudakov Problem (the virtual particle has mass of "m")

In this case, soft contribution vanishes.

That is:

$$I_s = i\bar{\pi}^{\frac{d}{2}} \mu^{4-d} \int d^d k \frac{1}{(k^2 m^2 + i\epsilon) [(k+l)^2 + i\epsilon] [(k+p)^2 + i\epsilon]} \quad (\text{where } k^2 \ll m^2)$$

$$\stackrel{LO}{=} -i\bar{\pi}^{\frac{d}{2}} \mu^{4-d} \int d^d k \frac{1}{m^2 (2k \cdot l + l^2) (2k \cdot p + p^2)} \rightarrow \text{scaleless.}$$

Since soft region vanishes and hard region remains its former form, if the (anti-)collinear regions also remain unchanged, the divergence could no more be cancelled. So we should introduce new regulators ——"Analytic regulator".

$$I = i\bar{\pi}^{\frac{d}{2}} \mu^{4-d} \int d^d k \frac{(-v^2)^\alpha}{(k^2 m^2 + i\epsilon) [(k+l)^2 + i\epsilon] [(k+p)^2 + i\epsilon]^{\alpha}}$$

For simplicity, we let the two outer legs be onshell and a mass of "0" $\Rightarrow l^2 = p^2 = 0$

$$I_c = i\bar{\pi}^{\frac{d}{2}} \mu^{4-d} \int d^d k \frac{(-v^2)^\alpha}{(k^2 m^2 + i\epsilon) [2k \cdot l + i\epsilon] [(k+p)^2 + i\epsilon]^{\alpha}} \quad \text{Feynman Parameter 3.}$$

$$= i\bar{\pi}^{\frac{d}{2}} \mu^{4-d} \int_0^\infty dx dy \int dk \frac{(2+\alpha)(1+\alpha)(-v^2)^\alpha}{[(1+x)(k^2 + 2k \cdot (\frac{y+P}{1+x}) - \frac{xm^2}{1+x})]^{2+\alpha}}$$

$$= i\bar{\pi}^{\frac{d}{2}} \mu^{4-d} \frac{(-1)^{3+\alpha} \cdot \pi^{\frac{d}{2}} \Gamma(3\alpha - \frac{d}{2}) v^d}{\Gamma(1+\alpha)} \int_0^\infty dx dy \left(\frac{xm^2}{1+x} + \frac{2y \cdot l \cdot P}{(1+x)^2} \right)^{\frac{d}{2}-3-\alpha} \cdot \frac{1}{(1+x)^{3+\alpha}}$$

$$= \frac{(-1)^\alpha \mu^{\frac{d}{2}} P(1+\alpha+\epsilon) (-v^2)^\alpha}{P(1+\alpha)} \int_0^\infty \frac{1}{(1+x)^{1-\alpha-\epsilon}} \left(m^2 x(x+1) + \vec{Q} \cdot \vec{y} \right)^{-1-\alpha-\epsilon} dx dy.$$

Integrate out y .

$$\int_0^\infty dy \frac{1}{(A+By)^{1+\alpha+\epsilon}} = \frac{1}{(\alpha+\epsilon)} \frac{1}{B \cdot A^{\alpha+\epsilon}}$$

$$= \frac{\mu^{\frac{d}{2}} P(1+\alpha+\epsilon) v^\alpha}{P(1+\alpha)} \int_0^\infty dx \frac{1}{(1+x)^{1-\alpha-\epsilon}} \frac{1}{(\vec{Q}^2 \cdot (m^2)^{\alpha+\epsilon} \cdot x^{\alpha+\epsilon} \cdot (x+1)^{\alpha+\epsilon})}$$

$$= \frac{P(\alpha+\epsilon)}{\alpha^2 P(1+\alpha)} \left(\frac{m^2}{m^2} \right)^\epsilon \left(\frac{v^2}{m^2} \right)^\alpha \int_0^\infty \frac{1}{(1+x)^{1-\epsilon}} \frac{1}{x^{\alpha+2\epsilon}}$$

$$= \frac{1}{\alpha^2} \left(\frac{m^2}{m^2} \right)^\epsilon \left(\frac{v^2}{m^2} \right)^\alpha \cdot \frac{P(\alpha+\epsilon) \cdot P(\alpha) \cdot P(1-\alpha-\epsilon)}{P(1+\alpha) \cdot P(1-\epsilon)}$$

$$= \frac{P(1+\epsilon)}{\alpha^2} \left(\frac{m^2}{m^2} \right)^\epsilon \left(\frac{v^2}{m^2} \right)^\alpha \cdot \left(\frac{1}{\alpha\epsilon} - \frac{1}{2^2} + \frac{\pi^2}{3} \right)$$

$$= \frac{P(1+\epsilon)}{\alpha^2} \left(-\frac{1}{2^2} + \frac{1}{\alpha\epsilon} + \frac{1}{\alpha^2} \cancel{\ln \frac{m^2}{m^2}} + \frac{1}{2} \ln \frac{v^2}{m^2} + \ln \frac{m^2}{m^2} \ln \frac{v^2}{m^2} - \frac{1}{2} \ln^2 \frac{m^2}{m^2} + \frac{\pi^2}{3} \right)$$

$$\begin{aligned}
I_{\bar{c}} &= i \bar{x}^{\frac{d}{2}} \mu^{4-d} \int d^d k \frac{(-v^2)^{\alpha}}{(k^2 m^2 + i\varepsilon) [(k+t)^2 + i\varepsilon] [(k+p)^2 + i\varepsilon]} \\
&= i \bar{x}^{\frac{d}{2}} \mu^{2\alpha} \frac{(-v^2)^{\alpha}}{\Gamma(1+\alpha)} \frac{i \bar{x}^{\frac{d}{2}} \Gamma(1+\alpha+\varepsilon) (v^2)^{\alpha}}{\int_0^\infty dx dy} \int_0^\infty dx dy \frac{(x+y)^{2+2\alpha+2\varepsilon}}{(x+y)^{3+\alpha} [(2k \cdot l_t + m^2 x) y + Q^2]^{1+\alpha+\varepsilon}} \\
&\stackrel{t=\frac{y}{x}}{=} \frac{\mu^{2\alpha} v^{2\alpha} \Gamma(1+\alpha+\varepsilon)}{\Gamma(1+\alpha)} \int_0^\infty dt dy \frac{y^{-1+\varepsilon}}{(1+t)^{1-\alpha-2\varepsilon} [m^2 t (1+t) y + Q^2]^{1+\alpha+\varepsilon}} \\
&\quad \int_0^\infty dy \frac{1}{y^{1+\varepsilon} (A y + B)^{1+\alpha+\varepsilon}} = \frac{\Gamma(1+\alpha) \Gamma(\varepsilon)}{\Gamma(1+\alpha+\varepsilon)} \frac{1}{A^\varepsilon B^{1+\alpha}} \\
&= \frac{1}{Q^2} \frac{\mu^{2\alpha}}{m^{2\alpha}} \frac{v^{2\alpha}}{Q^{\alpha}} \frac{1}{\Gamma(\varepsilon)} \int_0^\infty dt \cdot \frac{1}{(1+t)^{1-\alpha-\varepsilon} t^\varepsilon} \\
&= \frac{1}{Q^2} \left(\frac{\mu^2}{m^2} \right)^\varepsilon \left(\frac{v^2}{Q^2} \right)^\alpha \frac{\Gamma(\varepsilon) \cdot \Gamma(-\alpha) \Gamma(1+\varepsilon)}{\Gamma(1-\alpha-\varepsilon)} \\
&= \frac{\Gamma(1+\varepsilon)}{Q^2} \left(\frac{\mu^2}{m^2} \right)^\varepsilon \left(\frac{v^2}{Q^2} \right)^\alpha \left(-\frac{1}{\alpha\varepsilon} + \frac{\pi^2}{6} \right) \\
&= \frac{\Gamma(1+\varepsilon)}{Q^2} \left(-\frac{1}{\alpha\varepsilon} - \frac{1}{\alpha} \ln \frac{\mu^2}{m^2} - \frac{1}{\varepsilon} \ln \frac{v^2}{Q^2} - \ln \frac{\mu^2}{m^2} \ln \frac{v^2}{Q^2} + \frac{\pi^2}{6} \right)
\end{aligned}$$

$$\begin{aligned}
I_{c+\bar{c}} &= I_c + I_{\bar{c}} = \frac{\Gamma(1+\varepsilon)}{Q^2} \left(-\frac{1}{\varepsilon^2} - \frac{1}{\varepsilon} \ln \frac{\mu^2}{Q^2} + \ln \frac{\mu^2}{m^2} \ln \frac{Q^2}{m^2} - \frac{1}{2} \ln^2 \frac{\mu^2}{m^2} + \frac{\pi^2}{2} \right) \\
&= \frac{\Gamma(1+\varepsilon)}{Q^2} \left(-\cancel{\frac{1}{\varepsilon^2}} - \cancel{\frac{1}{\varepsilon} \ln \frac{\mu^2}{Q^2}} - \ln \frac{\mu^2}{m^2} \ln \frac{Q^2}{m^2} + \frac{1}{2} \ln^2 \frac{\mu^2}{m^2} + \frac{\pi^2}{2} \right)
\end{aligned}$$

Our previous conclusion:

$$\begin{aligned}
I_h &= \frac{\Gamma(1+\varepsilon)}{Q^2} \left(\cancel{\frac{1}{\varepsilon^2}} + \cancel{\frac{1}{\varepsilon} \ln \frac{\mu^2}{Q^2}} + \frac{1}{2} \ln^2 \frac{\mu^2}{Q^2} - \frac{\pi^2}{6} \right) \\
\Rightarrow I &= I_{c+\bar{c}} + I_h = \frac{1}{Q^2} \left(\frac{\pi^2}{3} + \frac{1}{2} \ln^2 \frac{\mu^2}{Q^2} - \ln \frac{\mu^2}{m^2} \ln \frac{m^2}{Q^2} + \frac{1}{2} \ln^2 \frac{\mu^2}{m^2} \right) \\
&= \frac{1}{Q^2} \left(\frac{\pi^2}{3} + \frac{1}{2} \ln^2 \frac{\mu^2}{Q^2} \right)
\end{aligned}$$

— Collinear Anomaly (Factorization Anomaly).

If we didn't introduce the " α "? We would have a divergence in the integral for the Feynman Parameter!

$$\begin{aligned}
* \quad I_c &= i \bar{x}^{\frac{d}{2}} \mu^{4-d} \int d^d k \frac{1}{(k^2 m^2 + i\varepsilon) [(k+t)^2 + i\varepsilon] [(k+p)^2 + i\varepsilon]} \\
&= i \bar{x}^{\frac{d}{2}} \mu^{4-d} \int d^d k \frac{1}{(k^2 m^2) (2k \cdot l_t) (k^2 + 2k \cdot p)} \quad \text{— Feynman Parameter 2} \\
&= i \bar{x}^{\frac{d}{2}} \mu^{4-d} \int d^d k \int_0^\infty dx dy \frac{2}{[(1+y)k + 2k \cdot (x l_t + y p) - m^2]^3}
\end{aligned}$$

$$\begin{aligned}
&= \mu^{\omega} \int_{(1+\varepsilon)}^{\infty} dx dy \frac{1}{(1+y)^3 \left[\frac{m^2}{1+y} + \frac{xy Q^2}{(1+y)^5} \right]^{\frac{1}{1-\varepsilon}}} \\
&= \frac{1}{m^2} \left(\frac{\mu^{\omega}}{m^2} \right) \int_0^{\infty} dx dy \frac{1}{(1+y)^{1-\varepsilon} (1+y + xy)^{\frac{1}{1-\varepsilon}}} \\
&\quad \int_0^{\infty} dx \frac{1}{(A+Bx)^{\frac{1}{1-\varepsilon}}} = \frac{1}{A^{\varepsilon} B^{\cdot \varepsilon}} \\
&= \frac{1}{m^2} \left(\frac{\mu^{\omega}}{m^2} \right)^2 \int_0^{\infty} dy \frac{1}{(1+y)^{\frac{1}{1-\varepsilon} \cdot \text{yr} \cdot \varepsilon}} \rightarrow \text{divergent when } y \rightarrow 0
\end{aligned}$$

so extra regulator is necessary!