Python 3

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2a.) The hydrostatice equilibrium formula is affroximately dP = GMS dr GO P = SdP = GMS where <math>S = M so $P \propto M^2$ R^2 R^3 R^4

3.9.) Using the ideal gas Law P=gkt, and subbing it in the equilibrium formula, we get: um+

 $\frac{g_{KT} = GMg}{um_{H}} = \frac{GMg}{g}$ $\Rightarrow \frac{g_{KT} = GM}{um_{H}} = \frac{g_{KT}}{g} =$

4.01) Density is in units K3/m3, so:

P X M B3

6.) Given that ROFM from my slope, we can subthis proportionality in for R in the proportionalities found in 2a, 3a and 4a:

Pc or M2 => Pc or M4/3 => Pc or 1 R4 M8/3 M4/3

TC X M => TC X M => TC X M 1/3

R: M2/3

POLM => POLM => POLI R3 M2 M 7.) Using Po, 4=0.62 and T~1.44x10 K like in 10.2, we find the ideal gas pressure to be:

 $P_{i} = \frac{90 \text{ KT}}{400 \text{ M}}$ $P_{i} = \frac{(11410)(1.38 \times 10^{-23})(1.44 \times 10^{7})}{(0.62)(1.673 \times 10^{-27})}$ $P_{i} = 2.7 \times 10^{14} \text{ N/m}^{2}$

The radiation pressure using the same temperature is,

 $P_{r} = \frac{1}{3}a_{7} + \frac{1}{3}(7.9658 \times 10^{-16})(1.44 \times 10^{7})^{4}$ $P_{r} = \frac{1}{3}(63 \times 10^{12} \text{ N/m}^{2})^{16}$

() m

Find the ratio between Pr and the total pressure gives:

 $\frac{Pr}{Pr+Pi} = \frac{3.63\times10^{12}}{2.7363\times10^{14}} = 0.013 = 1.3\%$

Therefore, the radiation pressure is only 1.3% of the total pressure making it almost negligable.

We can say Pr becomes important when Pr ≥ Pg, thereore

 $aT^4 > \overline{g}_{0}KT$ > Subbing in scaling relations for T with M, we get: $a(m's)^3 \ge g_{0}K$ $aT^3 > g_{0}K$ am_{H} am_{H}

M = 3(1410)(1.38×10-23)
(7.5658×10-16)(0.62)(1.673×10-24)

=7.44×1022