# NUR Assignment 2

#### Christiaan van Buchem - s1587064

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#### Abstract

In this document I will be giving my answers to the questions of the second assignment for the Numerical Recipes for Astrophysics course. For each question I will give a short introduction, write out any non-coded answers that may be required, produce the print statements and the plots, and finally I will show the script used to produce the results.

## 1 Normally distributed pseudo-random numbers

### 1.1 RNG

For exercise 1 we were tasked with writing a random number generator that returns a random floating point number between 0 and 1. At minimum we had to use some combination of an MWC and a 64-bit XOR-shift. The plots made to test the quality of the RNG can be seen in Figures 1(a), 1(b), and 2.

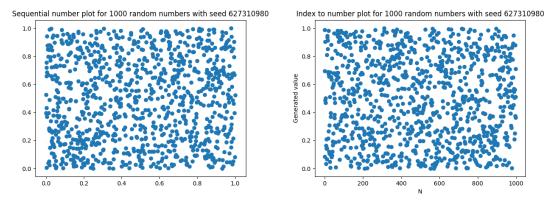


Figure 1: Left: Sequential number plot showing that it appears that each number is independent of its predecessor. Right: Index to number plot showing that there does not appear to be a relation between the index of a number and its value.

#### 1.2 Box-Muller method

Using the Box-Muller method we had to generate 1000 normally distributed random numbers. In order to check if they follow the expected distribution we make a histogram with an over-plotted Gaussian. The results can be seen in Figure 3.

#### 1.3 KS-test

For this exercise we tested whether or not our function is consistent with the normal distribution. The resulting plot can be seen in Figure 4. The slight difference between the two may be attributed to the fact that in the self written KS-test the following approximation was used:

$$P_{KS}(z) \approx \begin{cases} \frac{\sqrt{2\pi}}{z} [(e^{-\pi^2/(8z^2)}) + (e^{-\pi^2/(8z^2)})^9 + (e^{-\pi^2/(8z^2)})^2 5], & (z < 1.18) \\ 1 - 2[(e^{-2z^2}) - (e^{-2z^2})^4 + (e^{-2z^2})^9], & (z \ge 1.18) \end{cases}$$

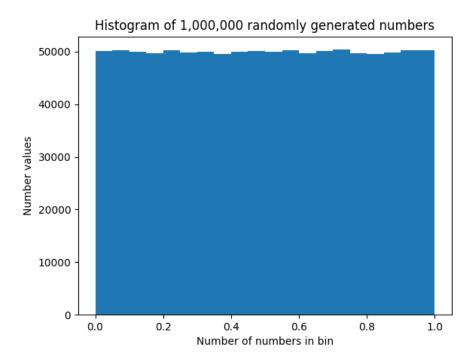


Figure 2: This histogram places the random number generator under a sharper knife, allowing us to see that there are some fluctuations between the bins. Overal it appears to be quite unbiased.

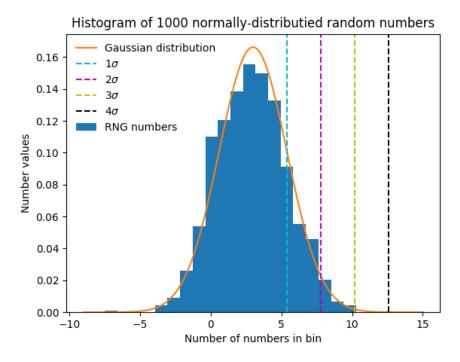


Figure 3: In this figure we can see that numbers generated using the Box-Muller method do indeed follow the Gaussian distribution.

## 1.4 Kuiper's-test

The same as for the KS-test except that we had to use Kuiper's test. Results can be seen in Figure 5.

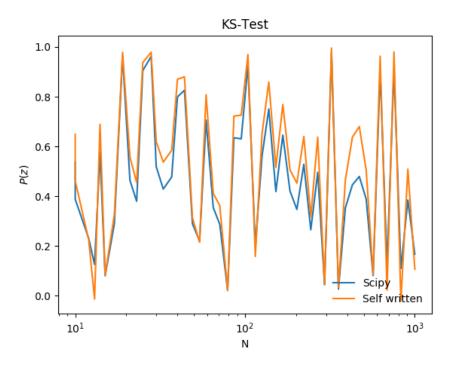


Figure 4: Here we see that the 'self-written' KS-test follows the Scipy KS-test results almost exactly.

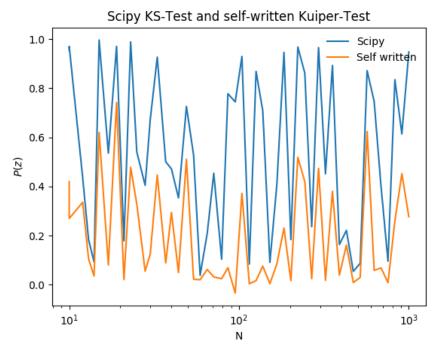


Figure 5: Here we compare the Kuipers test.

## 1.5 Analysing a dataset

In this exercise we were tasked with analysing a giving data set using either the KS-test or Kuipers test. The results can be seen in Figure 6.

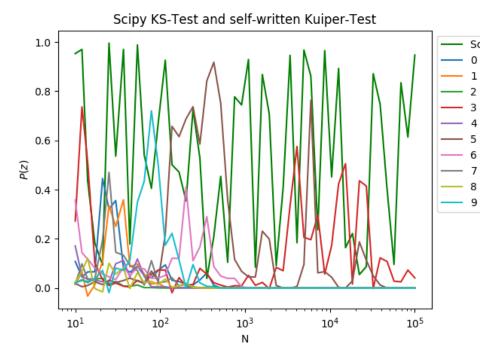


Figure 6: Analysing the different datasets.

## 1.6 Scripts

Here we can see the terminal output of the script used for this exercise:

```
--- Exercise 1 ---
Original seed: 627310980
Generated plots/la.png
Generated plots/lb.png
Generated plots/lc.png
Generated plots/lc.png
Generated plots/ld.png
Generated plots/lf.png
Generated plots/lf.png
Generated plots/lf.png
Generated plots/lf.png
```

 $a2\_1.txt$ 

Here is the script used to produce these results:

```
# a2_1
  import numpy as np
  import sys
  from matplotlib import pyplot as plt
  from scipy import stats
  import os
  # --- Functions and classes ---
  class rng(object):
      # Rng object that is initiated with a give seed
      a1, a2, a3 = np.int64(21), np.int64(35), np.int64(4)
      a \, = \, 4294957665
14
      def __init__(self, seed):
16
           self.state = np.int64(seed)
17
19
       def MWC(self):
          # Multiply with carry generator
```

```
x = np.int64 (self.state)
22
            self.state = self.a*(x&(2**32-1))+(x>>32)
23
        def XOR_shift(self):
24
            # XOR-shift generator
25
            x = np.int64 (self.state)
26
            x = x ^ x >> self.a1
x = x ^ x << self.a2
x = x ^ x >> self.a3
27
28
29
            self.state = np.int64(x)
30
       #end XOR-shift()
31
32
        def rand_num(self,l,min=0,max=1):
33
            # Generates 'l' random numbers between min and max
34
35
            output = []
            for i in range(l):
36
                 self.XOR_shift()
37
                 self.MWC()
38
                 self.XOR_shift()
39
                 \verb"output.append" (self.state")
40
41
            output = np.array(output)/sys.maxsize
            return min+(output*(max-min))
42
43
       #end rand_num()
  #end rng()
44
4.5
   def box\_muller(u1,u2,mu,sigma):
       # Implementation of the Box Muller transform
47
       x1 = (-2*np.\log(u1))**0.5*np.\sin(2*np.pi*u2)
48
       x2 = (-2*np.log(u1))**0.5*np.cos(2*np.pi*u2)
       return x1*sigma+mu, x2*sigma+mu
50
  #end box_muller
51
  \begin{array}{ll} \textbf{def} & \texttt{central\_diff}(\,f\,,h\,,x\,): \end{array}
53
       # Calculates the central difference\n",
       return (f(x+h)-f(x-h))/(2*h)
55
  #end central_diff()
56
   def ridders_diff(f,x):
58
       #Differentiates using Ridder's method
59
60
       m = 10
       D\,=\,np\,.\,z\,eros\,(\,(m,\textcolor{red}{len}\,(\,x\,)\,)\,)
61
62
       d = 2
       h = 0.001
63
       for i in range(m):
64
            D\_new\,=\,D
            for j in range (i+1):
66
67
                 if j == 0:
                      D_{new}[j] = central_diff(f,h,x)
68
69
                      D_{new}[j] = (d**(2*(j+1))*D[j-1]-D_{new}[j-1])/(d**(2*(j+1))-1)
70
            D = D_new
71
            h = h/d
72
73
       return D[m-1]
  #end ridders_diff()
74
75
   def comp_trapezoid(f,a,b,n):
76
       # Composite trapezoid rule used in romber_int()
       h \, = \, 1/(\,2\,{*}\,{*}\,(\,n\!-\!1)\,)\,{*}\,(\,b\!-\!a\,)
78
       sum = 0
79
       for i in range(1,2**(n-1)):
80
            sum += f(a+i*h)
        return (h/2.)*(f(a)+2*sum+f(b))
82
  #end comp_trapezoid()
83
   def romber_int(f,a,b):
85
       # Integrates from a to b up to an accuracy of 6 decimals
87
        for n in range (1,10):
            S_new = np.zeros((n))
88
89
            S_{new}[0] = comp\_trapezoid(f,a,b,n)
            for j in range (2,n+1):
90
```

```
S_{\text{new}}[j-1] = (4**(j-1)*S_{\text{new}}[j-2]-S[j-2])/(4**(j-1)-1)
                                           S = S_new
  92
                                           if n > 3:
  93
                                                           \hspace{.1in} \hspace{.1
  94
                                                                           return S[-1]
  95
                            return S[-1]
  96
           #end romber_int()
  97
  98
            \begin{array}{lll} \textbf{def} & KS\_Kuip\_test \, (\, sample \, , f \, , mu, sig \, , Kuip\!\!=\!\!False \, ) \, : \end{array}
  99
                           # Implementation of the Kalgorov-Smirnov test
100
                           N = len(sample)
                           x = np.linspace(mu-5*sig,mu+5*sig,1000)
                           F, Fn = np.zeros(len(x)), np.zeros(len(x))
                           Dmin = 0
                           Dmax = 0
                            for i in range(len(Fn)):
106
                                           Fn\left[\:i\:\right] \: = \: \underset{}{\text{len}}\left(\:np\:.\:where\left(\:sample {<\!\!=\!\!}x\left[\:i\:\right]\right)\:\left[\:0\:\right]\right)/N
                                          F[i] = romber_int(f, x[0], x[i])
108
                                           Dn = F[i] - Fn[i]
                                           if Dn > Dmin:
                                                          \mathrm{Dmin}\,=\,\mathrm{Dn}
                                          Dn = Fn[i] - F[i]
113
                                            if Dn > Dmax:
                                                           Dmax = Dn
114
                           # Determine the manner in which D is calculated
                            if Kuip:
116
                                        D = Dmin+Dmax
118
                            else:
                                         D = np.max((Dmin,Dmax))
                           # Calculate the probability
120
                           z = (N**0.5+0.12+0.11*N**(-0.5))*D
121
                            if z < 1.18:
                                         P = (2*np.pi)**0.5*((np.exp(-1*np.pi**2/(8*z**2))) + (np.exp(-1*np.pi**2/(8*z**2))) + (np.exp(-1*
                           )**9+(np.exp(-1*np.pi**2/(8*z**2)))**25)
                                         return D,1-P
                            else:
125
                                          P = 1 - 2*((np.exp(-2*z**2)) - (np.exp(-2*z**2)) **4 + (np.exp(-2*z**2)) **9)
126
                                           return D,1-P
           #end KS_test()
128
129
           def random_field_generator(n,N,rng,mu=0):
130
                           # Prepares a random field in Fourier space
                            print(f'Generating a random field with n = \{n\} of dimension \{N\}x\{N\} (mu = \{mu\})')
                           df = np.zeros((N,N), dtype=complex)
133
                           # Setting values of top half of the field
134
                           for j in range ((N//2)+1):
135
                           # Determining the value of k_y
136
                                           k_y = j*2*np.pi/N
137
                                            for i in range(N):
138
                                                           # Determining the value of k_x and sigma_x
139
                                                           if i <= (N//2):
140
                                                                          k_{-}x = (i)*2*np.pi/N
141
142
                                                            else:
                                                                          k_x = (-N+i)*2*np.pi/N
143
                                                          # Avoid dividing by 0
144
                                                            if i != 0 or j != 0:
145
                                                                          sig = ((k_x **2 + k_y **2) **0.5) **(n/2)
146
                                                            else:
147
                                                                          sig = 0
148
                                                          # Drawing a random number from normal distrib
149
                                                           \#df[j][i] = np.random.normal(0, sig) + 1j*np.random.normal(0, sig)
150
                                                           rand = box_muller(rng.rand_num(1), rng.rand_num(1), mu, sig)
                                                           df[j][i] = rand[0] + 1j*rand[1]
                           # Setting values of points who need to equal their own conjugates
                           df[0][0] = 0
                           df[0][N//2] = (df[0][N//2].real)**2
155
156
                            df[N//2][0] = (df[N//2][0].real)**2
                           df[N//2][N//2] = (df[N//2][N//2].real)**2
                           # Setting values of bottom half of the field using conjugates
                            for j in range ((N//2)+1):
```

```
for\ i\ in\ range\left(N\right):
160
                                   df[-j][-i] = df[j][i].conjugate()
161
                 return df
       #end random_field generator()
163
164
       # — Commands, prints and plots — if __name__ = '__main__':
165
166
                 print ('--- Exercise 1 --
167
                seed = 627310980
168
                rng = rng(seed)
169
                print('Original seed:', seed)
171
                #---- 1.a -
                # MWC and XOR-Shift
                N = 1000
174
175
                rand = rng.rand_num(N)
                # Sequential number plot
                 plt. scatter (rand[:(len(rand)-1)], rand[1:])
                plt.title('Sequential number plot for {} random numbers with seed {}'.format(1000,
178
                seed))
179
                 plt.savefig('plots/1a.png')
                 plt.close()
180
181
                 print('Generated plots/1a.png')
                # Index to number plot
182
                 plt.scatter(np.arange(0,N,1),rand)\\
183
                 plt.title('Index to number plot for {} random numbers with seed {}'.format(1000, seed
                ))
                plt.xlabel('N')
185
                plt.ylabel('Generated value')
186
                 plt.savefig('plots/1b.png')
187
188
                 plt.close()
                 print('Generated plots/1b.png')
189
                # Histogram
190
191
                N = 1000000
                rand = rng.rand_num(N)
192
                 plt.hist(rand,bins=20,range=(0,1))
193
                 plt.title('Histogram of 1,000,000 randomly generated numbers'.format(1000,seed))
194
                 plt.xlabel('Number of numbers in bin')
195
                 plt.ylabel ('Number values')
196
197
                 plt.savefig('plots/1c.png')
                 plt.close()
198
                 print('Generated plots/1c.png')
199
200
                #--- 1.b --
201
                # Box-Muller method
202
                N = 1000
203
204
                mu, sig = 3, 2.4
                rand = box_muller(rng.rand_num(N),rng.rand_num(N),mu,sig)
205
                 {\tt gauss = lambda \ x, mu, sig : 1/(2*np.pi*sig**2)**0.5*np.exp(-0.5*(x-mu)**2/sig**2)}
206
                x = np. linspace (mu - (sig *5), mu + (sig *5), 1000)
207
                 plt.hist(rand[0],bins=20,label='RNG numbers',density=1)
208
                 plt.\,plot\left(x,gauss\left(x,mu,sig\right),label='Gaussian\ distribution'\right)
209
210
                 plt.title('Histogram of {} normally-distributied random numbers'.format(1000))
                 plt.xlabel('Number of numbers in bin')
211
                 plt.ylabel ('Number values')
                 plt.axvline(x=mu+sig, label='$1\sigma$', color='c', linestyle='--')
213
                plt.axvline(x=mu+2*sig, label='$2\sigma$',color='m',linestyle='--')
plt.axvline(x=mu+3*sig, label='$3\sigma$',color='y',linestyle='--')
plt.axvline(x=mu+4*sig, label='$4\sigma$',color='k',linestyle='--')
214
215
216
                 plt.legend(frameon=False)
217
                 plt.savefig('plots/1d.png')
218
                 plt.close()
219
                 print('Generated plots/1d.png')
220
221
                         - 1.c.
                # KS-test
223
                # Setting parameters
224
                mu, sig = 0,1
225
                rand = box_muller(rng.rand_num(N), rng.rand_num(N), mu, sig)
                gauss = \frac{1}{2} = \frac{1}{2}
227
```

```
n = np. \log pace(np. \log 10(10), np. \log 10(1000), dtype=int)
228
        # Preparing arrays
229
        P, P_s = np. zeros(len(n)), np. zeros(len(n))
230
        d, d_s = np.zeros(len(n)), np.zeros(len(n))
        # Running test for different values of N
232
        for i in range(len(n)):
233
234
            rand = box_muller(rng.rand_num(n[i]), rng.rand_num(n[i]), mu, sig)
            d[i],P[i] = KS_Kuip_test(rand[0], gauss, mu, sig)
235
             d_s[i], P_s[i] = stats.kstest(rand[0], 'norm')
236
        # Plotting
237
        plt.plot(n, P_s, label='Scipy')
        plt.plot(n,P,label='Self written')
230
        plt.title('KS-Test')
240
        plt.ylabel('$P(z)$')
241
        plt.xlabel('N')
242
        plt.xscale('log')
243
        plt.legend(loc = 'lower right', frameon=False)
        plt.savefig('plots/1e.png')
245
        plt.close()
246
        print('Generated plots/1e.png')
247
248
        #---1.d-
249
250
        # Kuipers test
251
        # Preparing arrays
        kuip\_P , kuip\_P\_s = np. zeros(len(n)), np. zeros(len(n))
252
        kuip_d, kuip_d_s = np. zeros(len(n)), np. zeros(len(n))
253
        # Running test for different values of N
254
        for i in range(len(n)):
255
            rand = box\_muller(rng.rand\_num(n[i]), rng.rand\_num(n[i]), mu, sig)
256
             \label{eq:kuip_def} \begin{array}{ll} \text{kuip\_P[i]} = \text{KS\_Kuip\_test(rand[0], gauss,mu, sig, Kuip=True)} \end{array}
257
             kuip_d_s[i], kuip_P_s[i] = stats.kstest(rand[0], 'norm')
258
        # Plotting
259
        plt.plot(n, kuip_P_s, label='Scipy')
plt.plot(n, kuip_P, label='Self written')
260
261
        plt.title('Scipy KS-Test and self-written Kuiper-Test')
262
        plt.ylabel('$P(z)$')
263
        plt.xlabel('N')
264
        plt.xscale('log')
265
        plt.legend(loc = 'upper right', frameon=False)
266
267
        plt.savefig('plots/1f.png')
        plt.close()
268
        print('Generated plots/1f.png')
269
270
        #---1.e---
271
        # Testing on given random numbers
        filename = 'randomnumbers.txt
273
        url = 'https://home.strw.leidenuniv.nl/~nobels/coursedata/'
274
        if not os.path.isfile(filename):
275
            print(f'File not found, downloading {filename}')
os.system('wget '+url+filename)
27
        random_num = np.genfromtxt(filename, delimiter=' ', skip_footer=1)
278
279
        n = np.logspace(np.log10(10), np.log10(len(random_num)), dtype=int)
        test_P, test_D = np. zeros((10, len(n)), dtype=list), np. zeros((10, len(n)), dtype=list)
281
        # Applying Kuipers test
282
283
        for i in range (10):
             for j in range(len(n)):
284
                 rand = np. array(random_num[:n[j],i])
285
                 test_D[i][j], test_P[i][j] = KS_Kuip_test(rand, gauss, mu, sig, Kuip=True)
286
        # Plotting
287
        plt.plot(n, kuip_P_s, label='Scipy (KS)', color = 'g')
        for i in range(10):
289
             plt.plot(n,test_P[i],label = i)
290
        plt.title('Scipy KS-Test and self-written Kuiper-Test')
291
        plt.ylabel('$P(z)$')
292
        plt.xlabel('N')
293
        plt.xscale('log')
294
        plt.legend(loc=2, bbox_to_anchor=(1,1))
295
        plt.savefig('plots/1g.png')
        plt.close()
297
```

 $a2_1.py$ 

## 2 Making an initial density field

For this exercise we were asked to generate a Gaussian random field. The field is generated in Fourier Space. The complex Fourier amplitudes are given by  $\tilde{Y} = |\tilde{Y}exp(i\phi)|$  where phi is a random phase. The power spectrum has the following form:

$$P(k) \propto k^n \tag{1}$$

In Figure 7 the generated Gaussian random fields are given for different n values.

Choose a minimum physical size and explain how this impacts the maximum physical size, the minimum k and maximum k.

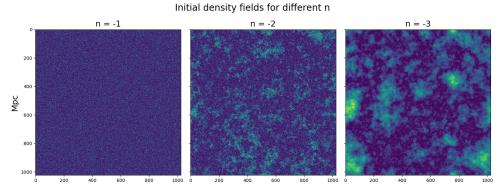


Figure 7: Gaussian random fields for different n values. Notice the clear presence of larger structure when the spectrum is more peaked (lower n).

### 2.1 Scripts

Here we can see the terminal output of the script used for this exercise:

 $a2\_2.txt$ 

Here is the script used to produce these results:

```
# a2-2
import numpy as np
import sys
from matplotlib import pyplot as plt
from scipy import stats
import os
from a2-1 import rng, box_muller

# --- Functions and classes ---

def random_field_generator(n,N,rng,mu=0):
    # Prepares a random field in Fourier space
    print(f'Generating a random field with n = {n} of dimension {N}x{N} (mu = {mu})')

df = np.zeros((N,N),dtype=complex)
    # Setting values of top half of the field
```

```
for j in range ((N//2)+1):
       # Determining the value of k_y
            k_y = j*2*np.pi/N
18
            for i in range(N):
19
                 # Determining the value of k_x and sigma_x
20
                 if i <= (N//2):
21
22
                      k_x = (i) *2*np.pi/N
23
                  else:
                      k_-x = (-N+i)*2*np.pi/N
24
                 # Avoid dividing by 0
25
                  if i != 0 or j != 0:
26
                      \mathrm{sig} \; = \; (\,(\,k_{-}x \!*\!\!*\!\! 2 \!+\! k_{-}y \,\!*\!\!*\!\! 2\,) \,\!*\!\!*\!\! 0\,.5\,) \,\!*\!\!*\!\! (\,n\,/\,2\,)
27
28
                      sig = 0
29
                 # Drawing a random number from normal distrib
30
                 \#df\left[\,j\,\right]\left[\,i\,\right] \;=\; np.\,random\,.\,normal\left(\,0\,\,,\,s\,i\,g\,\right) +\; 1\,j*np\,.\,random\,.\,normal\left(\,0\,\,,\,s\,i\,g\,\right)
31
                 rand = box\_muller(rng.rand\_num(1), rng.rand\_num(1), mu, sig)
32
                  df[j][i] = rand[0] + 1j*rand[1]
33
       # Setting values of points who need to equal their own conjugates
34
        df \, [\, 0\, ] \, [\, 0\, ] \; = \; 0
35
36
        df[0][N//2] = (df[0][N//2].real)**2
        df[N//2][0] = (df[N//2][0].real)**2
37
        df[N//2][N//2] = (df[N//2][N//2].real)**2
38
       # Setting values of bottom half of the field using conjugates
39
        for j in range ((N//2)+1):
40
             for i in range(N):
41
                 df[-j][-i]= df[j][i].conjugate()
42
        return df
43
  #end random_field generator()
45
  # — Commands, prints and plots — if __name__ = '__main__':
46
47
        print ('--- Exercise 2 --
48
        seed = 627310980
       rng = rng(seed)
50
        print('Original seed:', seed)
51
52
       # Making initial density fields for different n values
53
       N\,=\,1024
54
55
        df1 = random_field_generator(-1,N,rng)
        df1_inft = np.fft.ifft2(df1)
56
57
        df2 = random\_field\_generator(-2,N,rng)
        df2_i nft = np. fft. ifft2 (df2)
58
        df3 = random_field_generator(-3,N,rng)
59
        df3_inft = np.fft.ifft2(df3)
       # Plotting fields
61
        fig \;,\;\; ((ax1,ax2,ax3)) \;=\; plt.\, subplots (1,\;\; 3,sharex='col',\;\; sharey='row',figsize=(15,15))
62
        ax1.set_title('n = -1', size = 18)
63
        ax1.imshow(np.abs(df1_inft))
64
65
       ax1.set_ylabel('Mpc', size=18)
       ax1.invert_yaxis()
66
       ax2.set\_title("n = -2", size=18)
67
        ax2.imshow(np.abs(df2\_inft))
        ax3.set_title('n = -3', size=18)
69
        ax3.imshow(np.abs(df3_inft))
70
        fig.suptitle('Initial density fields for different n', y=0.7, size=20)
71
        fig.tight_layout()
72
        plt.savefig('plots/2.png',bbox_inches='tight',pad_inches = 0)
73
74
        plt.close()
        print ('Generated plots/2.png')
```

a2\_2.py

### 3 Linear Structure Growth

The evolution of density perturbations in the initial universe evolves according to the following equation:

$$\frac{\partial^2 \delta}{\partial t^2} + 2 \frac{\dot{a}}{a} \frac{\partial \delta}{\partial t} = \frac{3}{2} \Omega_0 H_0^2 \frac{\delta}{a^3} \tag{2}$$

In the early Universe we can separate the density perturbation as having a spatial part and a temporal part:  $\delta = D(t)\Delta(x)$ . In the case of a second order equation we have two growth factors. This means that the above partial differential equation becomes:

$$\frac{d^2D}{dt^2} + 2\frac{\dot{a}}{a}\frac{dD}{dt} = \frac{3}{2}\Omega_0 H_0^2 \frac{D}{a^3}$$
 (3)

We were asked to look at a Einstein-de Sitter Universe where  $\Omega_m = 1$  and the scale factor is given by:

$$a(t) = (\frac{3}{2}H_0t)^{2/3} \tag{4}$$

The density growth equation for this Universe is the following:

$$\frac{d^2D}{dt^2} = \frac{-4}{3t}\frac{dD}{dt} + \frac{2}{3t^2}D\tag{5}$$

For this exercise we were to calculate the numerical solutions for three different sets of initial conditions. These results were then to be compared with the analytical solutions of the ODE.

In Table ?? we can see the different cases and their analytical solutions.

	D(1)	D'(2)	D(t)
case 1	3	2	$3t^{2/3}$
${\rm case}\ 2$	10	-10	$10t^{-1}$
case3	5	0	$(3t^{5/3}+2)t^{-1}$

Table 1: The three different sets of initial conditions.

In Figure 8 we can see the numerical and analytical solutions for the 3 different cases. Mention why they do not match.

#### 3.1 Scripts

Here we can see the terminal output of the script used for this exercise:

```
Original seed: 627310980
Generated plots/la.png
Generated plots/lb.png
Generated plots/lc.png
Generated plots/ld.png
Generated plots/le.png
Generated plots/le.png
Generated plots/lf.png
Generated plots/lf.png
Generated plots/lf.png
Exercise 3 —
```

 $a2_3.txt$ 

Here is the script used to produce these results:

```
# a2_3
import numpy as np
import sys
import matplotlib.pyplot as plt
from a2_1 import rng, box_muller

def k_calc(h,f,t,x1,x2,xn):
    # Likely soruce of error: What do we do with the second variable when calculating k?
    # To-do: Try to solve the problem by applying it on a simpler function
    k1 = h * f(t,x1,x2)
```

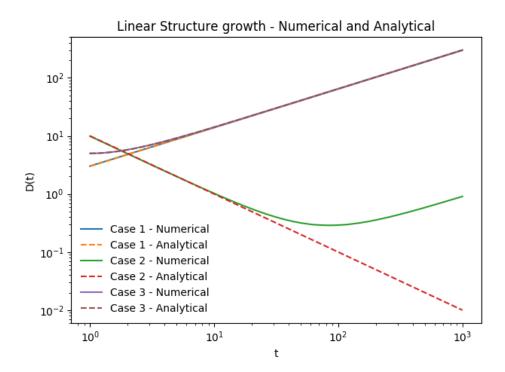


Figure 8: Analytical and numerical solutions to the partial differential equations given in this question.

```
k2 = h * f(t+0.5*h, x1+0.5*k1, x2+0.5*k1)
       k3 = h * f(t+0.5*h, x1+0.5*k2, x2+0.5*k2)
12
       {\bf k4} \; = \; {\bf h} \; * \; \; {\bf f} \left( \, {\bf t} {+} {\bf h} \, , {\bf x1} {+} {\bf k3} \, , {\bf x2} {+} {\bf k3} \, \right)
13
        return xn+1/6*k1+1/3*k2+1/3*k3+1/6*k4
14
   def runge_kutta(x0, y0, f, xmax, h=0.0001):
       #Implementation of the Runge-Kutta method for ODE integration
17
       # x0, y0 are the starting values and f is the ode()
       xn\;,yn\;=\;x0\;,y0
19
       y_out , x_out = [],[]
20
21
        while xn < xmax:
            yn_new = k_calc(h, f, xn, yn)
22
             y_out.append(yn_new)
23
24
            xn += h
            x_{\text{out}}. append (xn)
25
26
            yn = yn_new
27
        plt.plot(x_out,y_out)
28
29
        return np.sum(y_out)*h
30
31
   def k_{-}calc2nd(h, f, g, t, x1, x2):
32
       # Support function for runge_kutta method (for 2nd order ODEs)
33
       k1 = h * f(t, x1, x2)
34
        11 = h * g(t, x1, x2)
35
       k2 = h * f(t+0.5*h, x1+0.5*k1, x2+0.5*l1)
36
       12 = h * g(t+0.5*h, x1+0.5*k1, x2+0.5*l1)
37
       k3 = h * f(t+0.5*h, x1+0.5*k2, x2+0.5*k2)
38
       13 = h * g(t+0.5*h, x1+0.5*k2, x2+0.5*k2)
39
40
       k4 = h * f(t+h, x1+k3, x2+k3)
       14 = h * g(t+h, x1+k3, x2+k3)
41
42
43
       x1\_new = x1+1/6*(k1+2*k2+2*k3+k4)
       x2_new = x2+1/6*(11+2*12+2*13+14)
44
        return x1_new, x2_new
  #end k_calc2nd()
```

```
def runge_kutta2nd(x1_0, x2_0, t0, t, f, g, h=0.01):
       #Implementaiton of the Runge-Kutta method for ODE integration
50
       \# x0,y0 are the starting values and f is the ode()
51
       t = np.arange(t0, t+h,h)
52
       #print(t)
53
       x1n\,,x2n\ =\ x1\_0\,\,,x2\_0
55
       x1_out = np.zeros(len(t))
       for i in range(len(t)):
56
57
            x1n, x2n = k_{calc} 2nd(h, f, g, t[i], x1n, x2n)
            x1_out[i] = x1n
58
       return np.sum(x1_out)*h,x1_out
50
  # — Commands, prints and plots -
if __name__ = '__main__':
    print('— Exercise 3 — ')
61
62
63
       seed = 627310980
64
       rng = rng(seed)
       print('Original seed:', seed)
66
67
       f = lambda t, x1, x2: x2
       g = lambda t, x1, x2: -4/(3*t)*x2 + 2/(3*t**2)*x1
69
70
       case1, yt1 = runge_kutta2nd(3,2,1,1000,f,g)
       case2, yt2 = runge_kutta2nd (10, -10,1,1000, f, g)
71
       case3, yt3 = runge_kutta2nd(5,0,1,1000,f,g)
       print(f'case1: {case1}, case2: {case2}, case3: {case3}')
74
       f = lambda t, x1, x2 : x2
75
       g = lambda t, x1, x2 : x1*6-x2
76
77
       D1 = lambda t : 3*t**(2/3)
       D2 = lambda t : 10/t
       D3 = lambda t : (3*t**(5/3)+2)/t
80
       t = np.arange(1,1000+0.01,0.01)
       plt.plot(t,yt1,label='Case 1 - Numerical')
plt.plot(t,D1(t),linestyle='--',label='Case 1 - Analytical')
82
83
       plt.plot(t,yt2,label='Case 2 - Numerical')
       plt.plot(t,D2(t),linestyle='--',label='Case 2 - Analytical')
       plt.plot(t,yt3, label='Case 3 - Numerical')
plt.plot(t,D3(t),linestyle='--',label='Case 3 - Analytical')
       plt.legend(frameon=False)
       plt.xlabel('t')
       plt.ylabel('D(t)')
90
       plt.title('Linear Structure growth - Numerical and Analytical')
91
       plt.xscale('log')
       plt.yscale('log')
93
       plt.tight_layout()
       plt.savefig('plots/3.png')
       plt.close()
96
       print('Generated plots/3.png')
```

a2\_3.py

## 4 Zeldovich approximation

In this exercise we will be looking at the Zeldovich approximation.

### 4.1 Calculating the linear growth factor to a given redshift.

Our first task was to integrate the linear growth factor up to a redshift of z = 50. The integral to be solved is the following:

$$D(z) = \frac{5\Omega_m H_0^2}{2} H(z) \int_z^{\infty} \frac{1+z'}{H^3(z')} dz'$$
 (6)

Where

$$H(z)^{2} = H_{0}^{2}(\Omega_{m}(1+z)^{3} + \Omega_{\Lambda})$$
(7)

In order to avoid having to integrate up to  $\infty$  we will be substituting  $z = \frac{1}{a} - 1$ . This gives us the following equations:

$$D(a) = \frac{5\Omega_m H_0^2}{2} H(a) \int_0^a \frac{1}{a^3 H^3(a')} da'$$
 (8)

Where

$$H(a)^2 = H_0^2 \left(\frac{\Omega_m}{a^3} + \Omega_\Lambda\right) \tag{9}$$

The resulting value is: D(1/51) = 0.0196. The exact number and the way that it was calculated can be found in the print output below.

### 4.2 Calculating the derivative of the linear growth factor at a given redshift

In order to accomplish this task we had to analytically derive the value of  $\dot{D}(t)$ . One can calculate this indirectly using the following equation:

$$\dot{D}(t) = \frac{dD}{da}\dot{a}\tag{10}$$

Where

$$\dot{a} = \frac{H_0}{\sqrt{a}} \tag{11}$$

If we use the chain rule we get:

$$\frac{dD}{da} = \frac{5\Omega_m H_0^2}{2} \left[ \frac{dH(a)}{da} I + \frac{dI}{da} H(a) \right] \tag{12}$$

Where

$$I = \int_0^a \frac{1}{a^3 H(a)^3} da \tag{13}$$

Which gives us:

$$\dot{D}(a) = \frac{5\Omega_m H_0^3}{2\sqrt{a}} \left[ \frac{-3\Omega_m}{2\sqrt{a^5(\Omega_m + \Omega_\Lambda a^3)}} \int_0^a \frac{1}{a^3 H(a)^3} da + \frac{1}{a^3 H(a)^3} H_0 \sqrt{\frac{\Omega_m}{a^3} + \Omega_\Lambda} \right]$$
(14)

The resulting value is:  $\dot{D}(1/51) = 1239$  REQUIRE UNITS . The exact number and the way that it was calculated can be found in the print output below.

## 4.3 Evolution of a volume in 2D