## flares

January 21, 2023

## 1 FORECASTING MONTHLY SUN SPOTS

#### 1.1 Overview

The sun is composed of a moving electrically charged hot gas that generates a powerful magnetic field that goes through a cycle called the **solar cycle**. A solar cycle is a cycle that the sun's magnetic field makes at approximately every 11 years. At every 11 years or so, this magnetic field flips completely with the north pole switching places with the south pole of the sun. In another 11 years, they switch back places to their original state. Solar cycle affects the activity on the Sun's surface, one of which is the sunspots which are caused by the sun's magnetic fields. As the magnetic fields change, the amount of activity on the Sun's surface does the same.

To track the solar cycle, one way is to count the **number of sunspots**. At the beginning of a solar cycle is the solar minimum when the sun has the least sunspots. Overtime, the solar activity as well as the number of sunspots increases. At the middle of the cycle is the solar maximum where the number of sunspots are the highest and as the cycle ends, the number of sunspots fades back to the solar minimum and then a new cycle begins. Apart from the number of sunspots, other solar activities such as giant eruptions of the sun such as solar flares and coronal mass ejections also increase during the solar cycle, releasing powerful bursts of energy and material into space.

The Earth can also be affected by solar activities. For example, lights in the sky called aurora, impact on radio communications, and electricity grids. It can also cause a huge effect on satellite electronics and limit their lifespan and as well, can affect astronauts that work in International Space Stations. Some cycles can have minimums with lots of sunspots and activity while other cycles can have very few subspots and little activity. As a result of the effects of solar activity on Earth, it is important to predict the strength and duration of solar cycles. This will enable scientists forecast these solar conditions (space weather) and help them to protect radio communications on Earth, help keep satellites of space stations (eg. NASA) and keep astronauts safe by delaying their spacewalks.

In this project, we will forecast the monthly average number of sunspots for 2023. We will also forecast the average yearly sunspots for the next 11 years.

## 1.2 Dataset

The dataset contains daily sunspots observed from Jan. 1818 till Dec. 2022. There are days when no sunspots were recorded; these were replaced with -1 indicating missing values.

## 1.3 Data Visualisation and Analysis

In this section, we will start with some data visualisation and analysis to understand the data better. But firstly, we will import the dataset and some helper packages and libraries to aid in wrangling, analysis, and visualisation

We will look at the following research questions

#### 1.3.1 Research Questions

- What is the average yearly sunspots over time (1818-2022)?
- What is the summary statistics of monthly sunspots recorded between 1818-2022?
- What month of the year did we have the least and highest number of sunspots?
- Are there months in the year when sunspots are more?
- Is there any difference in the monthly sunspots recorded in the last 50 years?
- Forecasting average monthly sunspots for the next 12 months in 2023
- Forecasting average yearly sunspots to be observed for the next 11 years (2023-2033)

#### 1.3.2 Data Preprocessing

```
[1]: import pandas as pd
import numpy as np
import seaborn as sns
import matplotlib.pyplot as plt
import altair as alt
from scipy import stats
```

```
[2]: from pmdarima.arima import auto_arima, ndiffs from statsmodels.tsa.seasonal import seasonal_decompose from statsmodels.tsa.stattools import adfuller from statsmodels.graphics.tsaplots import plot_acf, plot_pacf
```

```
[3]: %matplotlib inline
```

Data Preprocessing steps - Data Integrity: Checking for missing rows and duplicate values - Replace -1 with np.nan to indicate missing values and drop them (From data source, we were told that -1 represents missing values) - Create the date variable - Selecting relevant variables

```
[5]: # data description
data.describe()
```

```
[5]:
                                                         date_year_frac
                                                                           sunspot_num
                                   month
                     year
                                                           74875.000000
                                                                          74875.000000
     count
            74875.000000
                           74875.000000
                                          74875.000000
                                             15.729416
                                                             1920.500267
                                                                              78.617242
     mean
             1920.000267
                                6.523005
                                              8.800057
                                                                              77.204256
     std
               59.178240
                                3.448705
                                                               59.178944
     min
             1818.000000
                                1.000000
                                              1.000000
                                                             1818.001000
                                                                              -1.000000
     25%
                                                             1869.249500
             1869.000000
                                4.000000
                                              8.000000
                                                                              15.000000
     50%
             1920.000000
                               7.000000
                                              16.000000
                                                             1920.501000
                                                                              58.000000
     75%
             1971.000000
                              10.000000
                                             23.000000
                                                             1971.750500
                                                                             123.000000
             2022.000000
                              12.000000
                                             31.000000
                                                             2022.999000
                                                                             528.000000
     max
             sunspot_num_std
                                      n_{obs}
                                                    status
               74875.000000
                              74875.000000
                                             74875.000000
     count
                    6.926375
                                   4.981008
                                                  0.997543
     mean
     std
                    4.805783
                                   9.064097
                                                  0.049512
     min
                   -1.000000
                                   0.000000
                                                  0.000000
     25%
                    3.000000
                                   1.000000
                                                  1.000000
     50%
                    6.500000
                                   1.000000
                                                  1.000000
     75%
                   10.000000
                                                  1.000000
                                   1.000000
                   77.700000
                                  69.000000
                                                  1.000000
     max
[6]: data = data.replace(-1, np.nan) # replace -1 with np.nan
     data['date'] = (data.iloc[:, :3]
                      .apply(lambda x: '-'.join([str(i).zfill(2) for i in x]), axis=1)
                      .apply(pd.to_datetime))
[7]: # are there any missing values
     data.isna().any().sum()
[7]: 2
[8]: # Any duplicates?
     sum(data.duplicated())
[8]: 0
[9]: missing = data.isnull().sum()
     missing[missing > 0]
[9]: sunspot_num
                         3247
     sunspot_num_std
                         3247
     dtype: int64
    There are 3247 days where no sunspots were recorded or the data were missing. We will drop these
    rows and select relevant features
```

[10]: # dropping missing values

data.dropna(inplace=True) # dropping missing rows

```
# selecting relevant variables
data = data[['date', 'sunspot_num', 'year', 'month']]
```

## [11]: data.head()

```
[11]:
              date
                    sunspot_num year month
      7 1818-01-08
                           65.0 1818
      12 1818-01-13
                           37.0 1818
                                            1
      16 1818-01-17
                           77.0 1818
                                           1
      17 1818-01-18
                           98.0 1818
                                           1
      18 1818-01-19
                           105.0 1818
                                           1
```

## 1.3.3 Data Visualisation and Analysis

### Average Sunspots over time (Per Year)

```
[12]: av_yearly_sunspots = data.resample('Y', on='date').sunspot_num.mean()
```

C:\Users\chigozie\DS\_projects\DS\_env\lib\site-packages\altair\utils\core.py:317: FutureWarning: iteritems is deprecated and will be removed in a future version. Use .items instead.

for col\_name, dtype in df.dtypes.iteritems():

#### [13]: alt.Chart(...)

From the plot above, we see that - there's an evidence of solar cycles approximately every 11 years. - Between Dec. 1818 till Dec. 2022, 18 solar cycles have happened. - The highest average number of sunspots per cycle was seen between 1954-1964, with a over 240 sunspots seen at its solar maximum (1957) around that time - Since the last three cycles, the solar maximum have decreased.

## Distribution of average monthly sunspots

```
[14]: av_monthly_sunspots = data.resample('M', on='date').sunspot_num.mean().fillna(0)
```

```
[15]: print('The average total sunspots per year is: {:,.0f}'.format(data.

resample('1Y', on='date').sunspot_num.sum().mean()))

print(f'The average of sunspots per month per year is: {av_monthly_sunspots.

mean():.0f}')
```

```
The average total sunspots per year is: 28,730 The average of sunspots per month per year is: 83
```

```
[16]: # when did we get the highest monthly sunspot on average
      print(av_monthly_sunspots[av_monthly_sunspots == av_monthly_sunspots.max()])
     date
     1957-10-31
                   359.387097
     Freq: M, Name: sunspot_num, dtype: float64
        • Over 28,000 sunspots are observed per year with an average of 83 sunspots observed per
        • Over 359 sunspots were observed per day in Oct. 1957
[17]: av_monthly_sunspots.describe()
[17]: count
               2460.000000
     mean
                 82.987526
      std
                 68.413782
                  0.000000
     min
      25%
                 24.274194
      50%
                 69.916667
      75%
                125.951613
                359.387097
      max
      Name: sunspot_num, dtype: float64
[18]: (
          alt.Chart(av_monthly_sunspots.reset_index())
          .mark_boxplot(size=100)
          .encode(alt.Y('sunspot_num', title='Monthly Number of sunspots', axis=alt.
       →Axis(grid=True)))
          .properties(height=300, width=400, title='Average Monthly sunspots_
```

C:\Users\chigozie\DS\_projects\DS\_env\lib\site-packages\altair\utils\core.py:317: FutureWarning: iteritems is deprecated and will be removed in a future version. Use .items instead.

for col\_name, dtype in df.dtypes.iteritems():

## [18]: alt.Chart(...)

- The maximum sunspot number recorded per month between 1818 to 2022 is about 359, with a minimum value of zero (no sunspot observed in that month)
- The average sunspot recorded is about 82 per month.
- The average monthly is right-skewed with some values as outliers.
- Most of the monthly sunspots recorded were above the median value, with fewer number recorded below the median value.

#### What months of the year did we record the highest and lowest number of sunspots?

- The maximum sunspot number recorded per month between 1818 to 2022 is about 359, with a minimum value of zero (no sunspot observed in that month)
- The average sunspot recorded is about 82 per month.
- The average monthly is right-skewed with some values as outliers.
- Most of the monthly sunspots recorded were above the median value, with fewer number recorded below the median value.

```
[ 1 2 3 4 5 6 7 8 9 10 11 12]
[1822 1823 1824 1855 1856 1867 1878 1879 1901 1902 1912 1913 2009]
```

- All the months of the year have records of no sunspots
- $\bullet$  These records were observed in 1822-1824, 1855-1856, 1867,1878-1879, 1901-1902, 1912-1913, and 2009

## Are there months in the year when sunspots are least or greatest?

C:\Users\chigozie\DS\_projects\DS\_env\lib\site-packages\altair\utils\core.py:317: FutureWarning: iteritems is deprecated and will be removed in a future version. Use .items instead.

for col\_name, dtype in df.dtypes.iteritems():

## [21]: alt.Chart(...)

- The number of observed sunspots is relatively stable all through the months of the year.
- There seems to be no seasonality in observed sunspots per month of the year

#### Is there any difference between number of observed sunspots last 50 years

```
[22]: last_50_years = av_yearly_sunspots.reset_index().assign(
    last_50 = 1*(av_yearly_sunspots.index.year > 2022-50)
```

```
[23]: last_50_years.groupby('last_50').sunspot_num.mean()
[23]: last 50
      0
           83.554005
      1
           80.685155
      Name: sunspot_num, dtype: float64
[24]: # get sunspots recorded the past 50 years
      df = last_50_years.groupby('last_50').sunspot_num.apply(lambda x: x.values)
      df
[24]: last_50
           [52.93896713615023, 38.53413654618474, 24.2321...
      0
           [54.12602739726027, 49.21643835616438, 22.4986...
      1
      Name: sunspot_num, dtype: object
[25]:
     stats.ttest_ind(df[0], df[1])
```

[25]: Ttest\_indResult(statistic=0.275540722249262, pvalue=0.7831811162296609)

• From the T-test result, there's no statistical difference in the average number of observed sunspots last 50 years, hence, we will fail to reject the null and conclude that at 5% significance level, there's no change in the average sunspots observed in the last 50 years and theose in pre-historic times

#### 1.3.4 Forecasting

Here, we will forecast the average monthly sunspots for 2023 and the average yearly sunspots for the next 11 years. We will use the average monthly and yearly sunspots used in the visualisation section.

These are the following steps to be taken in this section: Investigate the monthly and yearly sunspots for - Stationarity by performing time-series statistical tests such as the ADFuller test. - Decompose monthly and yearly sunspots into trend, seasonality and residual. - Plot the Autocorrelation and Partial Correlation plots to determine the optimal p,d,q values for ARIMA model - Split data into train and test sets: Train data will contain historical data for months before 2022 while the test set will contain sunspots observed in all the months in 2022. - Fit an ARIMA model on train data - Evaluate Peformance on test set. Peformance will be evaluated based on the root mean squared error (RMSE), mean absolute error (MAE) and mean absolute percentage error (MAPE) - Fit a time-series model on all data and forecast sunspots for the months in 2023 (monthly data) and for the next 11 years from 2023-2033 (yearly data)

#### Stationarity

This is a term used to describe a time series whose properties do not depend on the time the series was observed. That is, it has no trend or seasonality. This includes a time series with a cyclic behaviour but with no trend or seasonality, which the data is one. Cycles do not have a fixed

length and hence before it is observed, we are not sure where the peaks and troughs of the cycles will be.

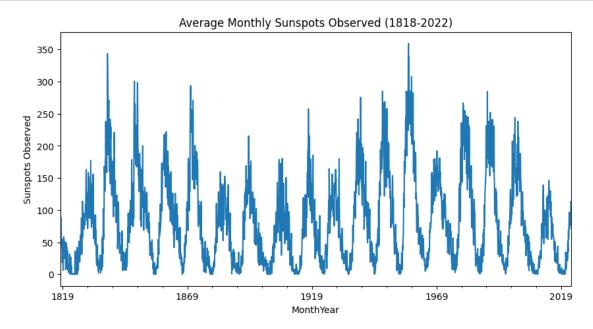
In general, a stationary time series will have no predictable pattern in the long-term and the time plots will show a series that is roughly horizontal and with constant variance.

We will visualise the monthly sunsplots and yearly sunspots to determine its stationarity. Similarly, the Augmented Dickey-Fuller (ADFuller) test for stationarity will be used. Another is the Kwiatkowski-Phillips-Schmidt-Shin (KPSS) test. We will also confirm this stationarity by decomposing the time series into trend, seasonality and residual values.

Methods to make time series stationary. One is differencing which involves computing differences between consecutive observations. A second-order or more differencing may be applied to make it stationary. This can stabilise the mean of the series by removing changes in the time series, thus eliminating or reducing trend and seasonality. Similarly, various transformations (example, log transformations) can be applied to stabilise the variance.

## Monthly Sunspots

```
[26]: # stationarity
av_monthly_sunspots.plot(figsize=(10,5))
plt.title('Average Monthly Sunspots Observed (1818-2022)')
plt.xlabel('MonthYear')
plt.ylabel('Sunspots Observed');
```



H0: The Series is non-stationary

H1: The Series is stationary

alpha = 0.05

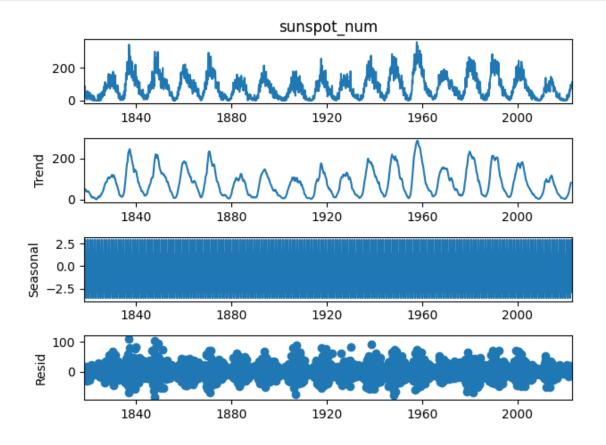
```
[27]: # ADFuller Test
stat, pval = adfuller(av_monthly_sunspots)[:2]
stat, pval

[27]: (-9.720023042852699, 9.599130204913231e-17)

[28]: # returns the number of differencing using a specified test
ndiffs(av_monthly_sunspots, test='kpss')

[28]: 0

[29]: ax = seasonal_decompose(av_monthly_sunspots).plot()
```



- We see that the average monthly number of sunspots is stationary
- Zero differencing was returned by the KPSS test indicating that no-differencing should be applied
- The seasonal decompose plot also shows us that there is neither any trend or seasonality in the data

## **AutoCorrelation and Partial Autocorrelation Plots**

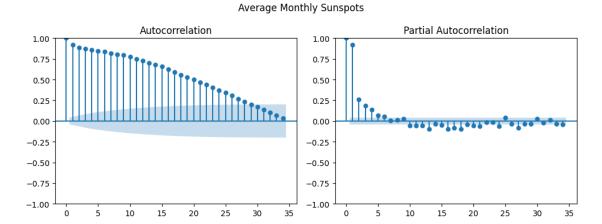
Autocorrelation tells us the relationship between a time point and the time point after it. A time

point may also be related to time points far away from it, but it is unsure if this correlation is because of the time points connecting them rather than the inherent contribution of that time point far away to the forecast. To overcome this problem, partial autocorrelation is used to find the relationship between a time point and n-timepoints by removing the effect of the time points in between.

Now, we will plot the autocorrelation and partial autocorrelation plots for our data to understand lags which may be informative in forecasting. An ARIMA model is composed of three orders: p,d,q, where the p indicates the number of autoregressive values that add information to forecasting the next value while the q indicates the number of moving averages. Autoregression is a type of regression of a variable against itself. The Autoregressive model uses a linear combination of past values of the variable to forecast the next value while a moving average model uses past forecast errors in a regression-like model to forecast the next value.

The Autocorrelation plot enables us to choose the number of autoregressions (lags) of order p while the partial autocorrelation plot enables us to choose the number of past forecast errors of order q.

```
[30]: fig, ax = plt.subplots(1,2,figsize=(10,4))
    plot_acf(av_monthly_sunspots,ax=ax[0])
    plot_pacf(av_monthly_sunspots,ax=ax[1], method='ywm')
    plt.suptitle('Average Monthly Sunspots', fontsize=12)
    plt.tight_layout()
    plt.show()
```



- From the ACF and PACF plots, we see that the monthly sunspots follow both an Autoregressive and moving average model.
- From ACF, the autocorrelation begins to decrease as the number of lags increase.
- From PACF, the correlation between one lag with the other begins to decrease up to the 4th value in the series

To select the optimal p,d,q model, we will automatically do it using the AutoARIMA model from pmdarima library

**Model Development and Evaluation** In this part, we will split the average daily number of sunspots observed per month of every year. We will use the average daily sunspots observed in 2022 to evaluate the model's performance.

We will select the p,d,q values automatically

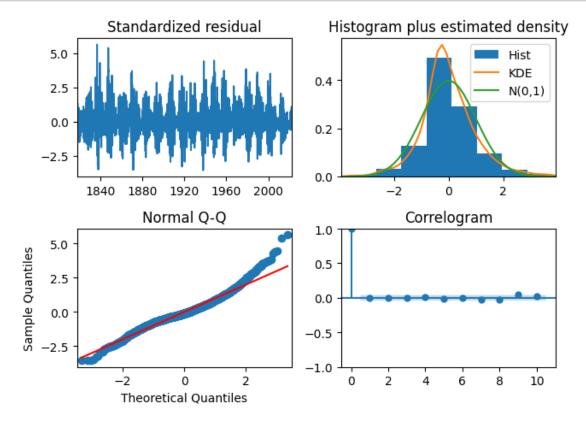
```
[31]: def mean_squared_error(y, preds):
         return np.mean((y-preds)**2)
     def mean_absolute_error(y, preds):
         return np.mean(np.abs(y-preds))
     def eval_metrics(y, ypreds):
         rmse = np.sqrt(mean_squared_error(y, ypreds))
         mae = mean_absolute_error(y, ypreds)
         print(f'MAE: {mae:.4f}\tRMSE: {rmse:.4f}')
[32]: train = av_monthly_sunspots[av_monthly_sunspots.index < '2022-01-31']
     test = av_monthly_sunspots[av_monthly_sunspots.index >= '2022-01-31']
[33]: train.shape, test.shape
[33]: ((2448,), (12,))
     Model Development (Forecasting)
[34]: model1 = auto_arima(train, seasonal=False, stationary=True, stepwise=False,
       →maxiter=100)
[35]: print(model1.summary())
                                   SARIMAX Results
     _____
     Dep. Variable:
                                           No. Observations:
                                                                            2448
                                           Log Likelihood
     Model:
                         SARIMAX(3, 0, 2)
                                                                      -11280.412
     Date:
                         Fri, 20 Jan 2023
                                           AIC
                                                                       22574.824
     Time:
                                 22:19:31
                                           BIC
                                                                       22615.445
     Sample:
                               01-31-1818
                                           HQIC
                                                                       22589.587
                             - 12-31-2021
     Covariance Type:
                                                    P>|z|
                                                               [0.025]
                                                                          0.975
                     coef
                             std err
     intercept
                   0.1319
                               0.018
                                         7.133
                                                    0.000
                                                               0.096
                                                                           0.168
     ar.L1
                                                    0.000
                   2.3486
                               0.023
                                       103.157
                                                               2.304
                                                                           2.393
     ar.L2
                  -1.7125
                               0.045
                                       -37.768
                                                    0.000
                                                              -1.801
                                                                          -1.624
                                                    0.000
     ar.L3
                   0.3623
                               0.023
                                        15.944
                                                               0.318
                                                                           0.407
     ma.L1
                  -1.8139
                               0.014
                                      -127.163
                                                    0.000
                                                              -1.842
                                                                          -1.786
                                        57.966
     ma.L2
                   0.8293
                               0.014
                                                    0.000
                                                               0.801
                                                                           0.857
```

sigma2	588.0926	12.000	49.007	0.000	564.573	611.613
===		=======	=======	=========	:======:	========
Ljung-Box	(L1) (Q):		0.02	Jarque-Bera	(JB):	
696.28			0.00	Decah (ID)		
Prob(Q): 0.00			0.88	Prob(JB):		
Heterosked	lasticity (H):		0.78	Skew:		
0.57						
Prob(H) (t 5.35	cwo-sided):		0.00	Kurtosis:		
=======		=======		.=======		========
===						

## Warnings:

[1] Covariance matrix calculated using the outer product of gradients (complex-step).

```
[36]: model1.plot_diagnostics()
    plt.tight_layout()
    plt.show()
```



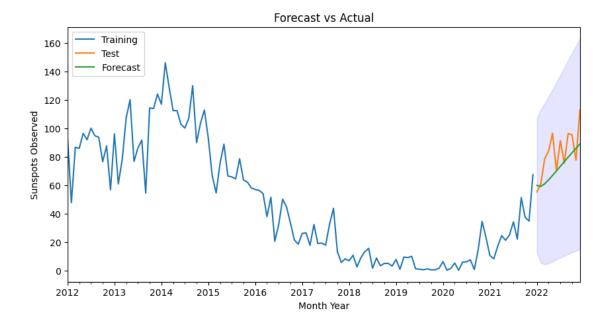
• From the model parameters, we see that the optimal p,d,q orders with least error is 3,0,2.

- We see that coefficients are statistically significant at 5% significance level.
- The model predictions for samples in the train data resembles a normal distribution with skewness of 0.57 with deviations at the end of the distribution (from the Normal Q-Q plot)

#### Model Evaluation

MAE: 12.9170 RMSE: 15.9218

- The model seems not be doing badly in its predictions.
- On average, the absolute difference between the actual and predicted values, given by the MAE, is about 13 sunspots while its predictions deviate from each other, given by the RMSE, is about 16 on average.



## Forecasting average monthly sunspots to be observed in 2023

[41]: model1 = auto\_arima(av\_monthly\_sunspots, seasonal=False, stationary=True, stepwise=False, maxiter=100)

[42]: print(model1.summary())

# SARIMAX Results

Dep. Variable: No. Observations: 2460 Model: SARIMAX(3, 0, 2)Log Likelihood -11331.884 Date: 22677.768 Fri, 20 Jan 2023 AIC Time: 22:20:44 BIC 22718.424 Sample: 01-31-1818 HQIC 22692.540

- 12-31-2022

Covariance Type: opg

	coef	std err	z	P> z	[0.025	0.975]
intercept	0.1324	0.018	7.211	0.000	0.096	0.168
ar.L1	2.3476	0.023	103.424	0.000	2.303	2.392
ar.L2	-1.7103	0.045	-37.837	0.000	-1.799	-1.622
ar.L3	0.3612	0.023	15.946	0.000	0.317	0.406
ma.L1	-1.8136	0.014	-127.412	0.000	-1.841	-1.786
ma.L2	0.8289	0.014	58.067	0.000	0.801	0.857
sigma2	586.2520	11.919	49.186	0.000	562.891	609.613

===

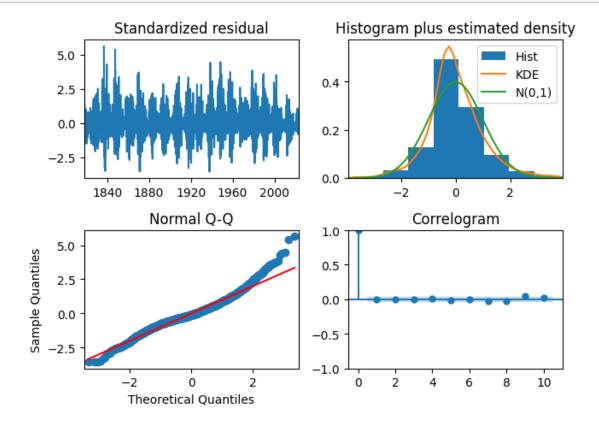
```
Ljung-Box (L1) (Q):
                                       0.02
                                               Jarque-Bera (JB):
702.13
Prob(Q):
                                               Prob(JB):
                                       0.89
0.00
Heteroskedasticity (H):
                                       0.78
                                               Skew:
0.57
Prob(H) (two-sided):
                                       0.00
                                               Kurtosis:
```

5.36

## Warnings:

[1] Covariance matrix calculated using the outer product of gradients (complexstep).

[43]: model1.plot\_diagnostics() plt.tight\_layout()



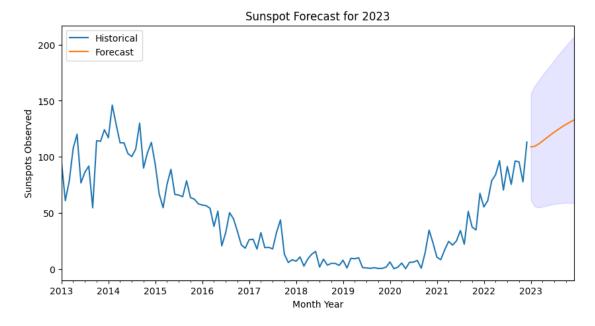
[44]: # evaluating eval\_metrics(av\_monthly\_sunspots, model1.fittedvalues())

MAE: 17.6788 RMSE: 24.2203

```
[45]: # forecasting for 2023
forecast, conf_int = model1.predict(12, return_conf_int=True)

conf_int = pd.DataFrame(conf_int, index=forecast.index, columns=['lower', upper'])
```

```
fig, ax = plt.subplots(figsize=(10,5))
av_monthly_sunspots[-120:].plot(ax=ax, label='Historical')
forecast.plot(ax=ax, label='Forecast')
ax.fill_between(conf_int.index, y1=conf_int.lower, y2=conf_int.upper,u
color='b', alpha=.1)
ax.legend(loc='upper left')
ax.set(title='Sunspot Forecast for 2023', ylabel='Sunspots Observed',u
cxlabel='Month Year');
```



• Our model forecasts that the average daily number of sunspots to be observed, per month, in 2023 will increase. This makes sense since, the last solar cycle ended in 2019 and a new cycle begun from 2020

**Yearly sunspots** Next, we will try to forecast the average monthly sunspots to be observed per year in the next 5 years. We will repeat the steps in the previous section

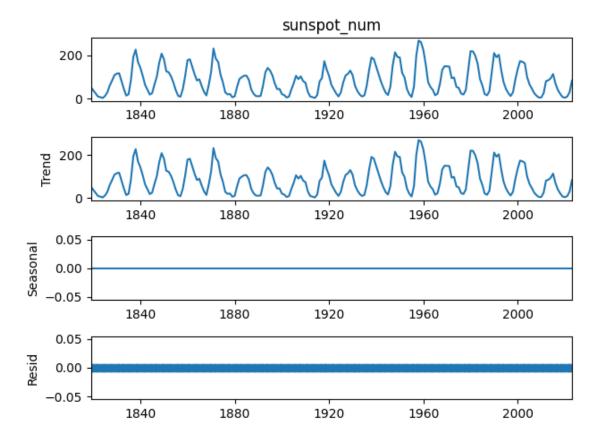
```
[47]: # ADFuller Test
stat, pval = adfuller(av_yearly_sunspots)[:2]
stat, pval
```

```
[47]: (-2.4955836887520406, 0.1165525621184536)
```

```
[48]: # returns the number of differencing using a specified test ndiffs(av_yearly_sunspots, test='kpss')
```

[48]: 0

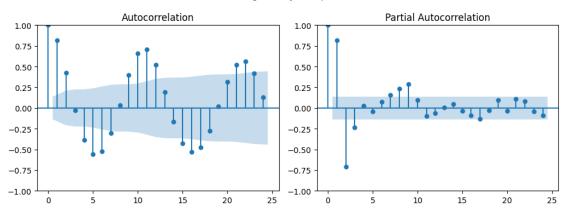
```
[49]: ax = seasonal_decompose(av_yearly_sunspots).plot()
```



- The ADFuller test returns that the data is non-stationary (pvalue > 0.05), however, using the kpss test to check the number of difference orders to apply, we see that no-differencing is to be applied.
- Just like in the monthly data, there is no seasonality or trend in our data

```
[50]: fig, ax = plt.subplots(1,2,figsize=(10,4))
    plot_acf(av_yearly_sunspots,ax=ax[0])
    plot_pacf(av_yearly_sunspots,ax=ax[1], method='ywm')
    plt.suptitle('Average Yearly Sunspots', fontsize=12)
    plt.tight_layout()
    plt.show()
```

#### Average Yearly Sunspots



- We see from the ACF plot that there's a cyclical pattern in the yearly average sunspots.
- We also see that between the 10th and 11th year, the cycle seems to peak at those times

## Model Development (Forecasting)

```
[51]: # using data for the last 5 years as test and others preceding them for training
train = av_yearly_sunspots[av_yearly_sunspots.index < '2013-01-31']
test = av_yearly_sunspots[av_yearly_sunspots.index >= '2013-01-31']
```

[52]: train.shape, test.shape

[52]: ((195,), (10,))

[53]: model1 = auto\_arima(train, seasonal=False, stationary=False, stepwise=False)

[54]: print(model1.summary())

#### SARIMAX Results

Dep. Variable:	у	No. Observations:	195
Model:	SARIMAX(3, 0, 2)	Log Likelihood	-903.131
Date:	Fri, 20 Jan 2023	AIC	1818.261
Time:	22:21:22	BIC	1837.899
Sample:	12-31-1818	HQIC	1826.213

- 12-31-2012

Covariance Type: opg

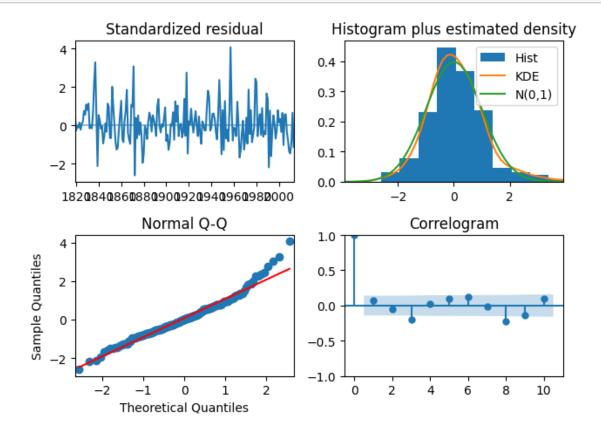
	coef	std err	Z	P> z	[0.025	0.975]
ar.L1	2.6191	0.029	90.040	0.000	2.562	2.676
ar.L2	-2.5653	0.051	-50.713	0.000	-2.664	-2.466
ar.L3	0.9430	0.024	40.017	0.000	0.897	0.989
ma.L1	-1.5676	0.074	-21.295	0.000	-1.712	-1.423

ma.L2 sigma2	0.7062 599.9209	0.077 51.801	9.194 11.581	0.000 0.556 0.857 0.000 498.393 701.449
========	==========			
===				
Ljung-Box	(L1) (Q):		0.92	Jarque-Bera (JB):
33.56				
Prob(Q):			0.34	Prob(JB):
0.00				
Heterosked	asticity (H):		1.05	Skew:
0.66				
Prob(H) (t	wo-sided):		0.84	Kurtosis:
4.55				
=======		=======	=======	

===

Warnings: [1] Covariance matrix calculated using the outer product of gradients (complexstep).

[55]: model1.plot\_diagnostics() plt.tight\_layout();



- From the model parameters, we see that the optimal p,d,q orders with least error is 3,0,2.
- We see that coefficients are statistically significant at 5% significance level.
- The model predictions for samples in the train data resembles a normal distribution with skewness of 0.66 with slight deviations at the end of the distribution (from the Normal Q-Q plot)

#### Model Evaluation

```
[56]: # Evaluation on train data
    eval_metrics(train, model1.fittedvalues())

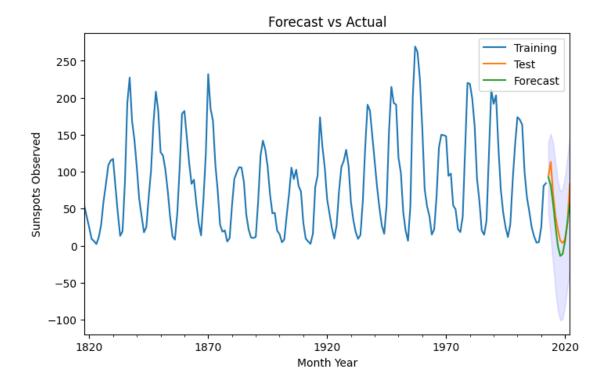
MAE: 18.6194    RMSE: 24.7802

[57]: test_preds, conf_int = model1.predict(len(test), return_conf_int=True)
    conf_int = pd.DataFrame(conf_int, index=test_preds.index, columns=['lower', u'upper'])

[58]: # evaluation on test data
    eval_metrics(test, test_preds)

MAE: 14.9751    RMSE: 18.0323
```

- The model seems not be doing badly in its predictions.
- On average, the absolute difference between the actual and predicted values, given by the MAE, is about 15 sunspots while its predictions deviate from each other, given by the RMSE, is about 18 on average.



• From the plot, we see that the model seems to model the movement of the cycle. However, the confidence of the predictions tend to decrease from the 3rd-4th year (lower confidence interval)

## Forecasting average yearly sunspots to be observed in the next 11 years

```
[60]: model1 = auto_arima(av_yearly_sunspots, seasonal=False, stationary=False, used)

stepwise=False, maxiter=100)
```

## [61]: print(model1.summary())

#### SARIMAX Results

Dep. Variable:	у	No. Observations:	205
Model:	SARIMAX(3, 0, 2)	Log Likelihood	-946.245
Date:	Fri, 20 Jan 2023	AIC	1904.489
Time:	22:21:46	BIC	1924.427
Sample:	12-31-1818	HQIC	1912.554

- 12-31-2022

Covariance Type: opg

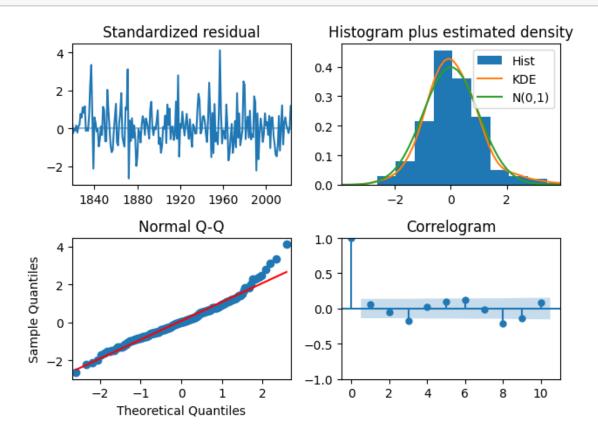
========						=======
	coef	std err	Z	P> z	[0.025	0.975]
ar.L1	2.6193	0.029	91.894	0.000	2.563	2.675
ar.L2	-2.5649	0.050	-51.495	0.000	-2.663	-2.467

ar.L3	0.9428	0.023	40.540	0.000	0.897	0.988
ma.L1	-1.5661	0.071	-22.037	0.000	-1.705	-1.427
ma.L2	0.7004	0.074	9.472	0.000	0.555	0.845
sigma2	582.1283	48.607	11.976	0.000	486.861	677.396
========		=======	=======	=========		========
===						
Ljung-Box (	(L1) (Q):		0.77	Jarque-Bera	(JB):	
36.15						
Prob(Q):		0.38	Prob(JB):			
0.00						
Heteroskeda	Heteroskedasticity (H):			Skew:		
0.64						
<pre>Prob(H) (two-sided):</pre>			1.00	Kurtosis:		
4.61						

## Warnings:

[1] Covariance matrix calculated using the outer product of gradients (complex-step).

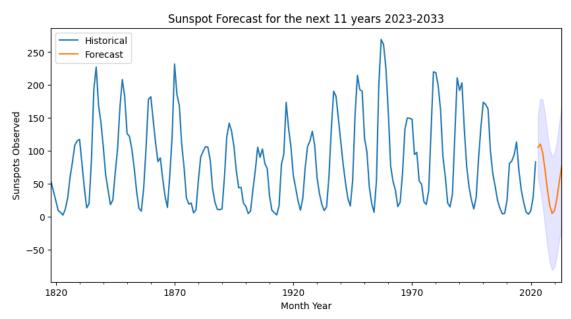
[62]: model1.plot\_diagnostics()
plt.tight\_layout()



```
[63]: # evaluating
      eval_metrics(av_yearly_sunspots, model1.fittedvalues())
     MAE: 18.2473
                     RMSE: 24.4045
[64]: # forecasting from 2023-2033
      forecast, conf_int = model1.predict(11, return_conf_int=True)
      conf_int = pd.DataFrame(conf_int, index=forecast.index, columns=['lower',_

¬'upper'])
[65]: with plt.ion():
          fig, ax = plt.subplots(figsize=(10,5))
          av_yearly_sunspots.plot(ax=ax, label='Historical')
          forecast.plot(ax=ax, label='Forecast')
          ax.fill_between(conf_int.index, y1=conf_int.lower, y2=conf_int.upper,_
       ⇔color='b', alpha=.1)
          ax.legend(loc='upper left')
          ax.set(title='Sunspot Forecast for the next 11 years 2023-2033', __

¬ylabel='Sunspots Observed', xlabel='Month Year');
```



- Our model seems to go with the cyclical nature of sunspots
- It forecasts that the current solar cycle reach its maximum peak in 2024 and then completes its cycle in 2030