

**restrict:**

$\text{coarse}(i,j) = \text{fine}(2*i-1, 2*j-1)$

**prolongate:**

$\text{fine}(i,j) = (\text{coarse}(\text{floor}((i+1)/2.), \text{floor}((j+1)/2.)) \&$   
 $\quad + \text{coarse}(\text{floor}((i+1)/2.), \text{ceiling}((j+1)/2.)) \&$   
 $\quad + \text{coarse}(\text{ceiling}((i+1)/2.), \text{floor}((j+1)/2.)) \&$   
 $\quad + \text{coarse}(\text{ceiling}((i+1)/2.), \text{ceiling}((j+1)/2.)) / 4.$

# Conservation equations

- No rotation
- No heat sources

Conservation of momentum (Navier-Stokes equation)

$$\frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \nabla \vec{v} = -\frac{1}{\rho} \nabla p + \nu \nabla^2 \vec{v} + \cancel{2\vec{\Omega} \times \vec{v}} + g\beta T \vec{e}_y$$

Conservation of energy (thermodynamic equation)

$$\frac{\partial T}{\partial t} + \vec{v} \cdot \nabla T = \kappa \cdot \nabla^2 T + \cancel{Q}$$

Conservation of mass (continuity equation)

$$\nabla \cdot \vec{v} = 0$$

$\vec{v}$ :	Velocity vector (u,v)
$\rho$ :	Density
$p$ :	Pressure
$\nu$ :	Kinematic viscosity
$\Omega$ :	Angular velocity
$g$ :	Gravitational acceleration
$\beta$ :	Thermal expansion coefficient
$T$ :	Temperature
$\vec{e}_y$ :	Unit vector in y direction
$\kappa$ :	Diffusion coefficient
$Q$ :	Heating rate

# Nondimensionalization

## Advantages:

- Less parameters
- Easier to recognize the dynamic regime
- Easier to compare systems of different spatial / temporal scales and similar dynamics

$$\tilde{x} = \frac{x}{L} \quad \tilde{T} = \frac{T}{\Delta T} \quad \tilde{t} = \frac{t}{L^2/\kappa} \quad \tilde{v} = \frac{v}{\kappa/L} \quad \tilde{p} = \frac{p}{\rho\nu\kappa/L^2}$$

$\rho$ : density    $\nu$ : kinematic viscosity    $\kappa$ : diffusion coefficient    $L$ : characteristic length

# Navier-Stokes equation

$$\frac{\kappa^2}{L^3} \cdot \left( \frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \nabla \vec{v} \right) = -\frac{\nu \kappa}{L^3} \nabla p + \nu \frac{\kappa/L}{L^2} \nabla^2 \vec{v} + \Delta T g \beta T \vec{e}_y \quad \left| \cdot \frac{L^3}{\nu \kappa} \right.$$

$$\underbrace{\frac{\kappa}{\nu}}_{\frac{1}{\text{Pr}}} \cdot \left( \frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \nabla \vec{v} \right) = -\nabla p + \nabla^2 \vec{v} + \underbrace{\frac{\Delta T g \beta L^3}{\nu \kappa}}_{\text{Ra}} T \vec{e}_y \quad \left| \text{Pr} \rightarrow \infty \right.$$

$$-\nabla p + \nabla^2 \vec{v} = -\text{Ra} \cdot T \cdot \vec{e}_y \quad \xrightarrow{\text{Exercise 6}}$$

Poisson equations

$$\nabla^2 \psi = -\omega$$

$$-\nabla^2 \omega = \text{Ra} \cdot \frac{\partial T}{\partial x}$$

Pr: Prandtl number Ra: Rayleigh number

## Thermodynamic equation

$$\frac{\Delta T}{L^2/\kappa} \cdot \left( \frac{\partial T}{\partial t} + \vec{v} \cdot \nabla T \right) = \frac{\Delta T}{L^2} \kappa \cdot \nabla^2 T \quad \left| \cdot \frac{L^2}{\kappa \Delta T} \right.$$

$$\frac{\partial T}{\partial t} + \vec{v} \cdot \nabla T = \nabla^2 T$$

## Exercise 7

The following dimensionless conservation equations describe convection in a fluid with high viscosity (e.g. Earth's mantle):

$$\frac{\partial T}{\partial t} + \vec{v} \cdot \nabla T = \nabla^2 T \quad (1)$$

$$-\nabla^2 \omega = \text{Ra} \cdot \frac{\partial T}{\partial x} \quad (2)$$

$$\nabla^2 \psi = -\omega \quad (3)$$

Your task in this exercise is to write a Fortran program that solves these equations. For this purpose, combine the Poisson solver from Exercise 6 with the advection-diffusion program from Exercise 5.

## Exercise 7

### 1. Read in the input variables from a namelist:

- Number of grid points **nx** and **ny**
- Integration time **total\_time**
- Time step constants **a\_adv** and **a\_diff**
- **NEW**: Convergence criterion for Poisson solver **max\_err**
- **NEW**: Rayleigh number **Ra**
- **NEW**: Temperature initialization **T\_ini\_type**: **random** or **cosine**

### 2. Initialize the variables:

- Temperature **T** with random numbers or a cosine function.  
Cosine e.g.  $T(x) = 0.5 \cdot \left(1 + \cos\left(3\pi \cdot \frac{x}{x_{max}}\right)\right)$
- Stream function and vorticity **S=W=0** (only at the beginning)
- Grid spacing **h=1. / (ny - 1. )**

## Exercise 7

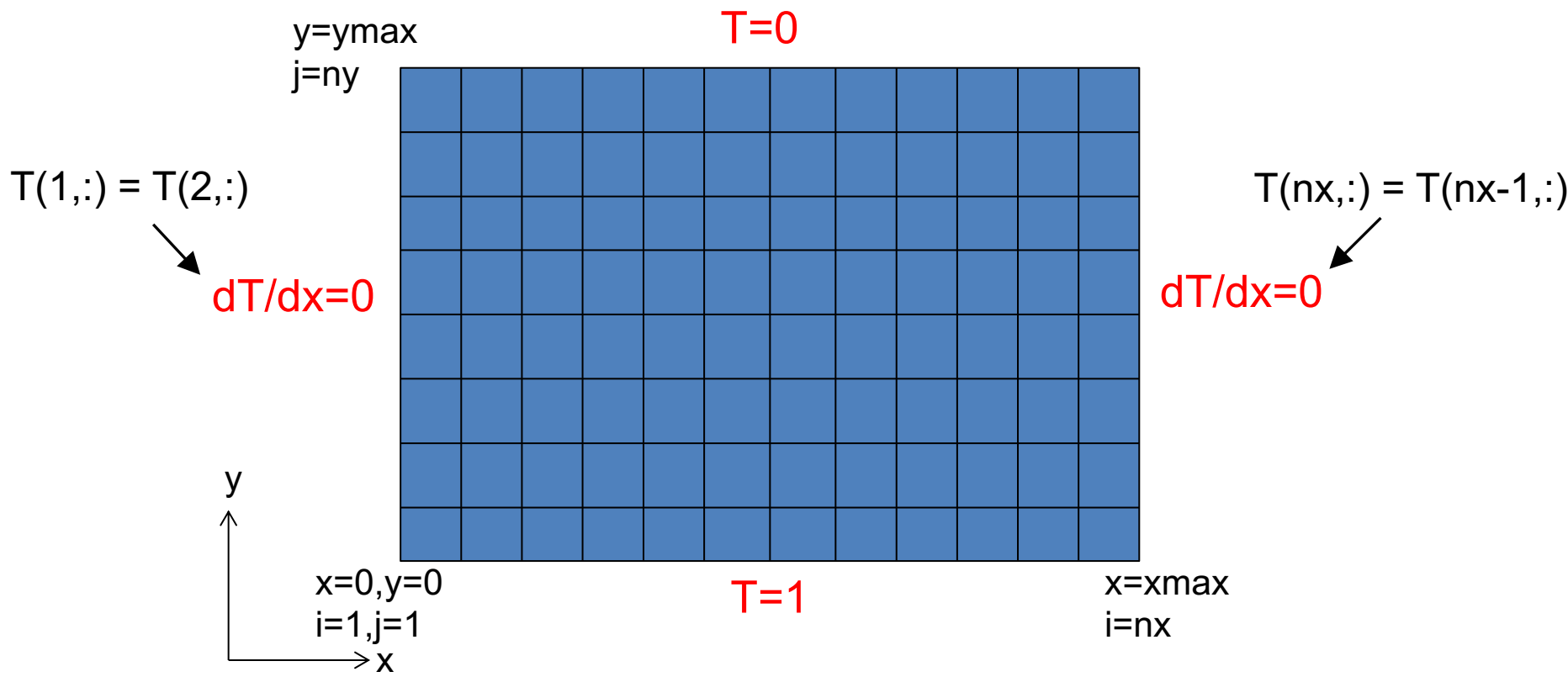
3. Perform several time steps until the integration time is reached:

- Compute  $Ra \cdot \frac{\partial T}{\partial x}$
- Determine  $\omega$  from  $Ra \cdot \frac{\partial T}{\partial x}$  using the Poisson solver (Equation 2)
- Determine  $\psi$  from  $\omega$  using the Poisson solver (Equation 3)
- Compute the wind speeds  $u$  and  $v$  from  $\psi$
- Compute the time step from  $a\_adv$ ,  $a\_diff$  and the maximum wind speed in the model domain
- Compute  $\nabla^2 T$  and  $\vec{v} \cdot \nabla T$  using the subroutines from Exercise 5
- Integrate forward in time:
  - $T_{i,j}^{n+1} = T_{i,j}^n + \dots$  (Equation 1)
  - $t = t + \Delta t$



# Exercise 7

Boundary conditions:



## Exercise 7

4. Test the program with the following two namelists, and save the output in binary files.

```
1  &INPUTS
2    nx=257
3    ny=65
4    a_diff=0.23
5    a_adv=0.4
6    total_time=0.1
7    max_err=1.E-3
8    Ra=1.E5
9    T_ini_type='cosine'
10 /
```

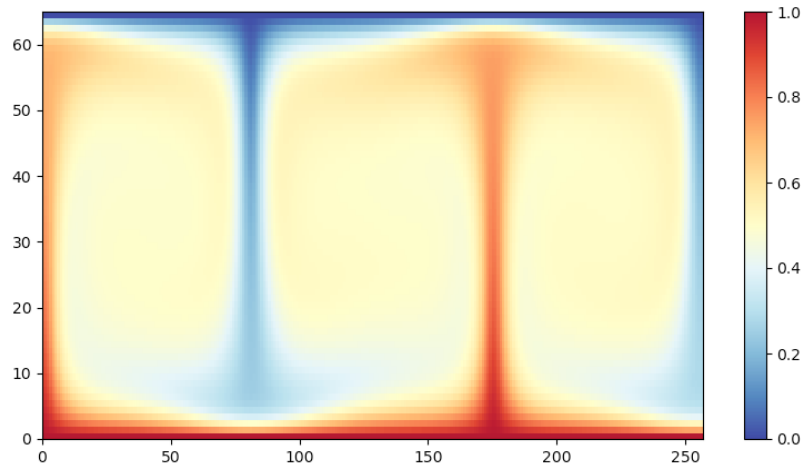
```
1  &INPUTS
2    nx=1025
3    ny=65
4    a_diff=0.23
5    a_adv=0.4
6    total_time=0.1
7    max_err=1.E-3
8    Ra=1.E5
9    T_ini_type='random'
10 /
```

5. Plot the temperature fields with your favorite plotting tool.

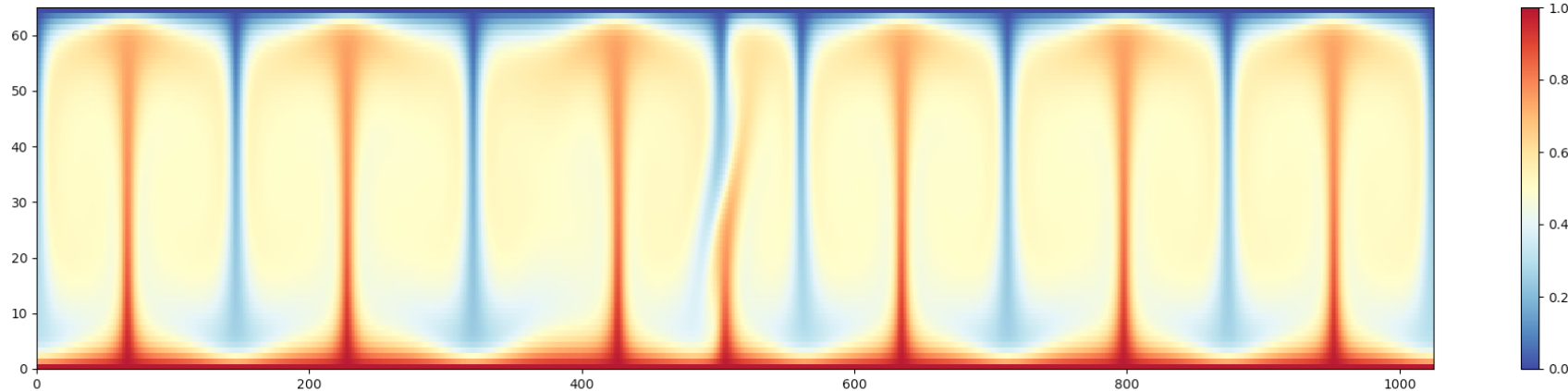
## Resulting temperature fields

Should look like this (qualitatively)

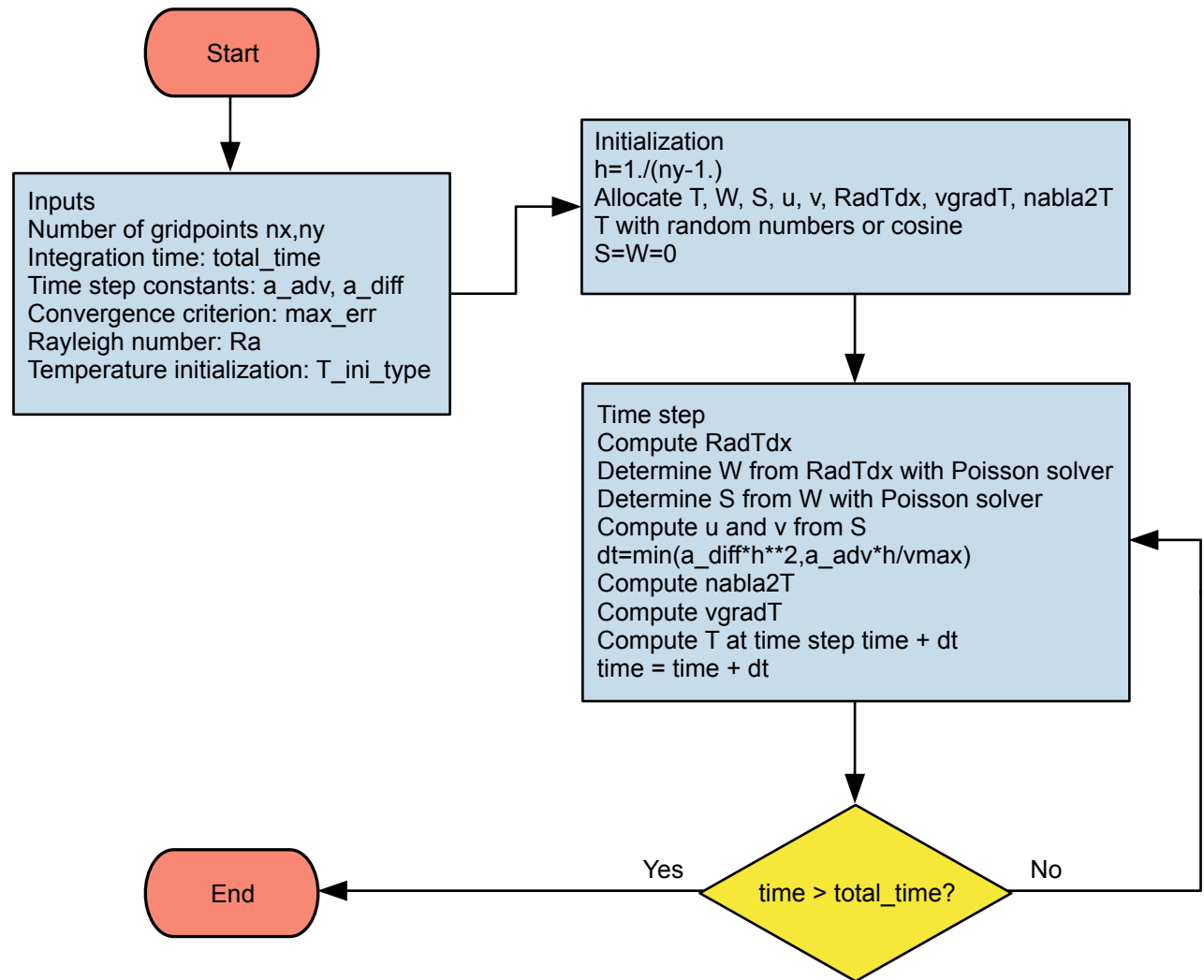
Namelist 1



Namelist 2



# Flowchart for Exercise 7



## Exercise 7

Deadline:

- Please hand in your solutions (.f90 files and plots T) no later than Tuesday, **14 May 2024, 23:59**.

Questions?

- Email me ([marina.duetsch@univie.ac.at](mailto:marina.duetsch@univie.ac.at))  
Or pass by my office (UZA II, 2G551).