

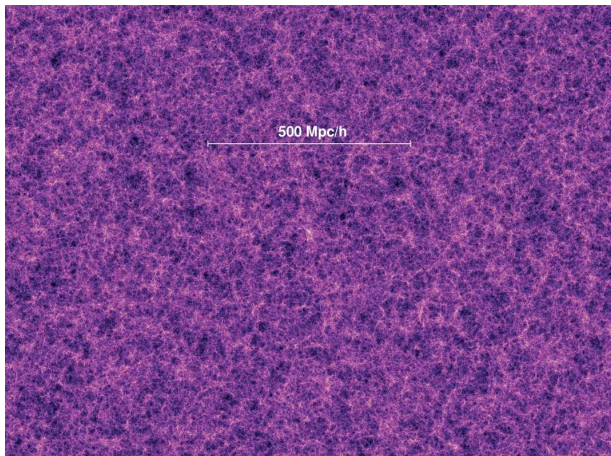
Comprehensive models of galaxy formation and evolution

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@violegp



The Millennium simulation



- Simulation:

<http://virgodb.cosma.dur.ac.uk:8080/Millennium/>
Millimillennium box size = $62.5 \text{ Mpc} h^{-1}$

Mass of each dark matter particle = $8.6 \cdot 10^8 M_{\odot} h^{-1}$

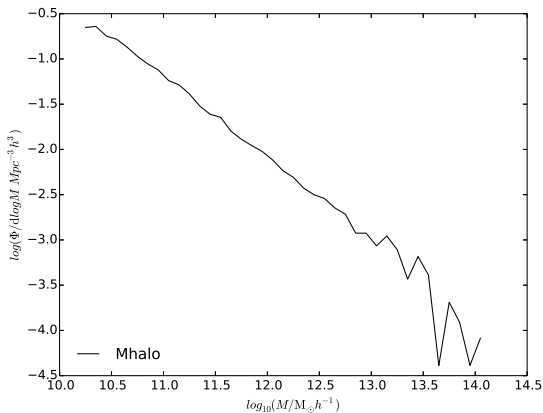
- Data: <http://www.astro.ljmu.ac.uk/~ikb/research/gama-gsmf-paper.html>

Query for getting the dark matter mass function

```
select .1*(.5+floor((log10(np*0.86)+9.)/.1)) as mass,  
log10(count(*)/power(62.5,3.)/.1) as phi  
from millimil..MPAHalo  
where snapnum=63  
group by .1*(.5+floor((log10(np*0.86)+9.)/.1))  
order by mass
```

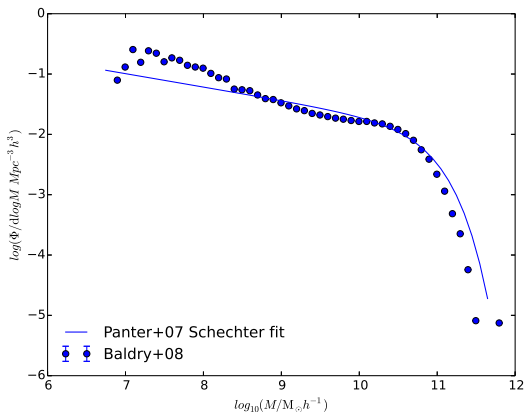
The mass function

Plotting directly the output from the query: mass vs phi



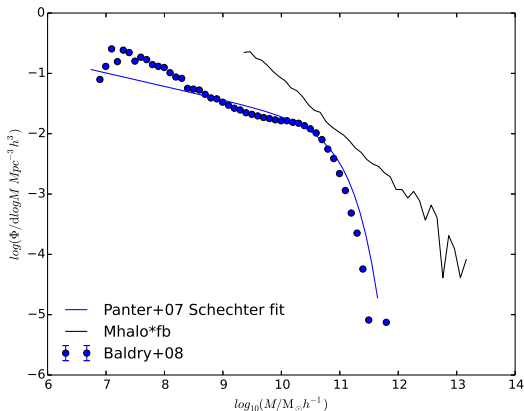
The galaxy stellar mass function

Plotting the galaxy stellar mass function from observations (Baldry et al. 2008) together with the Schechter function derived by Panter et al. 2007:



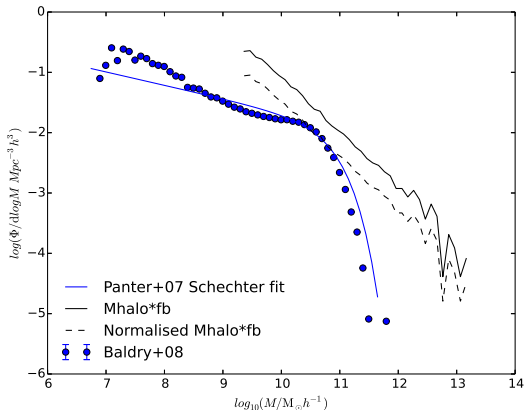
The galaxy stellar mass function

Let's compare the observed the galaxy stellar mass function with that derived from the dark matter mass function (mass and phi from the query) by multiplying it by the baryonic fraction ($M_* = M_{\text{halo}} \cdot f_b$)



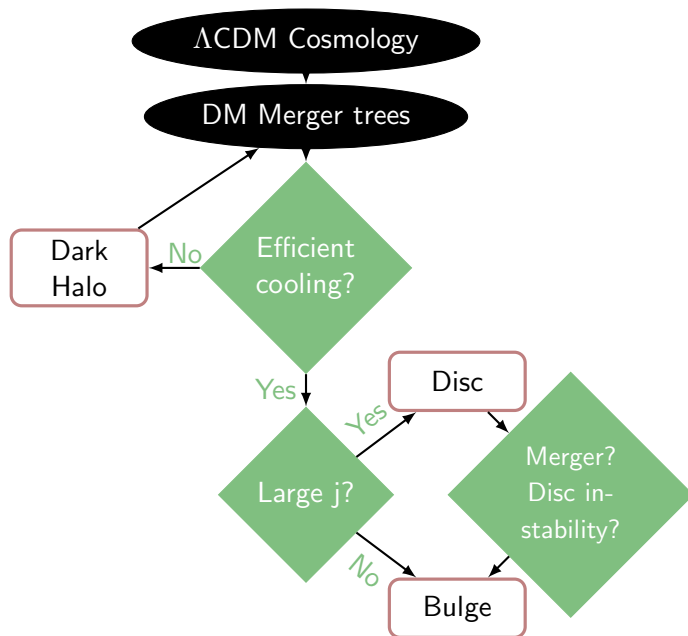
The galaxy stellar mass function

Let's normalized the galaxy stellar mass function around the knee of the observed one ($M_* = N \cdot M_{\text{halo}} \cdot f_b$):

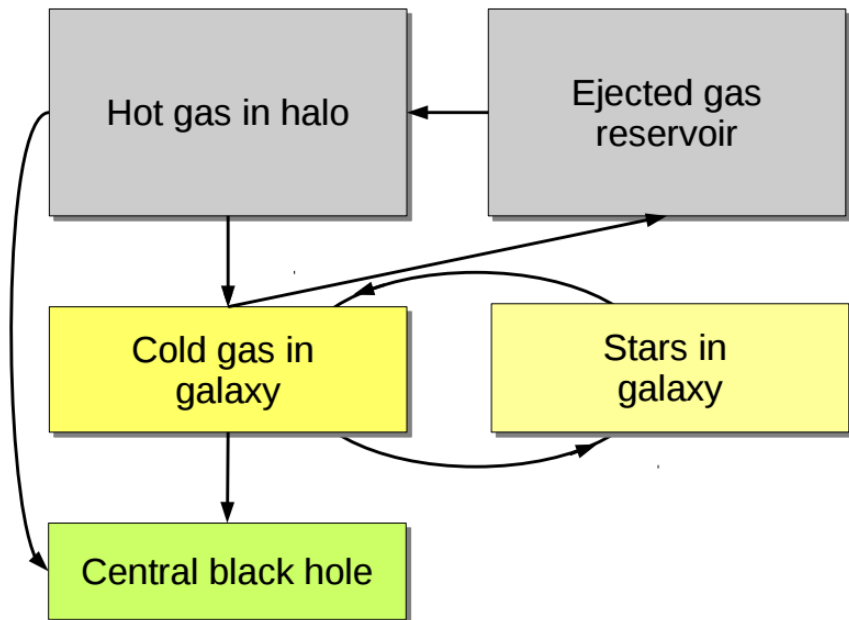


The shapes are still different! We need a better model to connect the luminous matter to the dark one.

Making galaxies in the computer with a semi-analytical model



The model baryonic component in a halo



The semi-analytical approach



Credits: Volker Springel, John Helly, Daniel Farrow, Joel Snape, Andrew Moore, Carlton Baugh, Richard Bower, Cedric Lacey and Carlos

Query for getting the stellar mass function

The following query gets the galaxy stellar mass function from the De Lucia et al. 2006 model, which is a comprehensive model of galaxy formation and evolution:

```
select .1*(.5+floor((log10(stellarMass)+10.)/.1)) as  
mass,  
log10(count(*)/power(62.5,3.)/.1) as phi  
from millimil..DeLucia2006a  
where snapnum = 63 and stellarMass > 0  
group by .1*(.5+floor((log10(stellarMass)+10.)/.1))  
order by mass
```

The galaxy stellar mass function

Plotting the result of the previous query, together with the other results:

