Input and output files Basic running

./class can take two input files *.ini and *.pre:

>./class my_model.ini some_precision.pre

But one is enough. Syntax:

```
h = 0.7
T_cmb = 2.726 # comment
output = tCl, pCl
more comments, ignored because there is no equal sign
# comment with an =, still ignored thanks to the sharp
```

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>./class my_model.ini some_precision.pre
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more comments, ignored because there is no equal sign
# comment with an =, still ignored thanks to the sharp
```

- Order of lines doesn't matter at all.
- All parameters not passed fixed to default, i.e. the most reasonable or minimalistic choice
- All possible input parameters and details on the syntax explained in explanatory.ini
- This is only a reference file; we advise you *never* to modify it, but rather to copy it and reduce it to a shorter and more friendly file.
- For basic usage: explanatory.ini = full documentation of the code

For instance, we can create a very short file lcdm.ini:

```
**********
* CLASS input parameter file *
********
---> background parameters:
HO = 72.
omega_b = 0.0266691
omega\_cdm = 0.110616
----> thermodynamics parameters:
z reio = 10.
----> define primordial perturbation spectra:
A s = 2.3e-9
n_s = 1.
----> define which perturbations should be computed:
output = tCl, pCl
----> parameters for the output spectra:
1 \text{ scalar max} = 2500
```

Try to run the code with an even smaller input file nut.ini:

output = tCl
output_verbose=1



Run with

./class nut.ini

Check that C_l 's have been written in output/nut00_cl.dat

more output/nut00_cl.dat

Essential input parameters (1/2)

Essential input parameters controlling the output (see details in explanatory.ini):

```
modes = s,t
ic = ad, cdi, bi, nid, niv
lensing = yes
non linear = halofit
output = tCl, pCl, 1Cl, mPk, mTk, vTk, nCl, sCl
1_{max_scalars=2500}
1 max tensors=500
l_max_lss = 1000
P_k_max_h/Mpc = 0.2
#P_k_max_1/Mpc =
z_pk = 0 \# or 1,2,10
root = output/test_ #default: output/<ini_file>##_
headers = [yes/no]
format = [class/camb]
```

Essential input parameters (2/2)

Essential input parameters controlling the output (see details in explanatory.ini):

```
write background = [yes/no]
write thermodynamics = [yes/no]
k_output_values = 0.01, 0.1, 0.0001
write primordial = [yes/no]
write parameters = [yes/no]
write warnings = [yes/no]
input_verbose = 1  # or 0, 2, 3,...
background_verbose = 1
thermodynamics_verbose = 1
perturbations_verbose = 1
transfer verbose = 1
primordial_verbose = 1
spectra_verbose = 1
nonlinear verbose = 1
lensing_verbose = 1
output_verbose = 1
```

Following files created (or not) automatically (here we assume that root=test_), depending on the content of the input fields:

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
lensing= [yes, no]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
# ... and more, e.g. for nCl, sCl redshift bins
```

ullet test_cl.dat total unlensed C_l 's

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
lensing= [yes, no]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
# ... and more, e.g. for nCl, sCl redshift bins
```

- ullet test_cl.dat total unlensed C_l 's
- ullet test_cl_lensed.dat total lensed C_l 's

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
lensing= [yes, no]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
# ... and more, e.g. for nCl, sCl redshift bins
```

- test_cl.dat total unlensed C_l's
- ullet test_cl_lensed.dat total lensed C_l 's
- ullet test_cls.dat scalar C_l 's when two modes
- lacktriangle test_clt.dat tensor C_l 's when two modes

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
lensing= [yes, no]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
# ... and more, e.g. for nCl, sCl redshift bins
```

- test_cl.dat total unlensed C_l's
- ullet test_cl_lensed.dat total lensed C_l 's
- lacktriangle test_cls.dat scalar C_l 's when two modes
- lacktriangledown test_clt.dat tensor C_l 's when two modes
- test_cl_ad.dat, test_cl_cdi.dat, test_cl_ad_cdi.dat etc. when different i.c. requested

Following files created (or not) automatically (here we assume that root=test_), depending on the content of the input fields:

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
lensing= [yes, no]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
# ... and more, e.g. for nCl, sCl redshift bins
```

- test_cl.dat total unlensed C_l's
- ullet test_cl_lensed.dat total lensed C_l 's
- lacktriangle test_cls.dat scalar C_l 's when two modes
- lacktriangledown test_clt.dat tensor C_l 's when two modes
- test_cl_ad.dat, test_cl_cdi.dat, test_cl_ad_cdi.dat etc. when different i.c. requested

Number of columns in these files can vary a lot depending on input parameters. Always indicated in the header.



Run with root=output/test1_, with either output=tCl or tCl,pCl,lCl, and either format = camb or class. In each case, look at headers in output/test_cl.dat

Following files created (or not) automatically (here we assume that root=test_), depending on the content of the input fields:

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
non linear= [none, halofit]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
z_pk = [list of values]
```

• test_pk.dat matter power spectrum

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
non linear= [none, halofit]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
z_pk = [list of values]
```

- test_pk.dat matter power spectrum
- test_pk_nl.dat non-linear matter power spectrum

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
non linear= [none, halofit]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
z_pk = [list of values]
```

- test_pk.dat matter power spectrum
- test_pk_nl.dat non-linear matter power spectrum
- test_pk_ad.dat, test_pk_cdi.dat, test_pk_ad_cdi.dat etc. when different i.c. requested

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
non linear= [none, halofit]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
z_pk = [list of values]
```

- test_pk.dat matter power spectrum
- test_pk_nl.dat non-linear matter power spectrum
- test_pk_ad.dat, test_pk_cdi.dat, test_pk_ad_cdi.dat etc. when different i.c. requested
- test_tk.dat density and/or velocity transfer functions

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
non linear= [none, halofit]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
z_pk = [list of values]
```

- test_pk.dat matter power spectrum
- test_pk_nl.dat non-linear matter power spectrum
- test_pk_ad.dat, test_pk_cdi.dat, test_pk_ad_cdi.dat etc. when different i.c. requested
- test_tk.dat density and/or velocity transfer functions
- test_tk_ad.dat, test_tk_cdi.dat, test_tk_ad_cdi.dat etc. when different i.c. requested

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
non linear= [none, halofit]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
z_pk = [list of values]
```

- test_pk.dat matter power spectrum
- test_pk_nl.dat non-linear matter power spectrum
- test_pk_ad.dat, test_pk_cdi.dat, test_pk_ad_cdi.dat etc. when different i.c. requested
- test_tk.dat density and/or velocity transfer functions
- test_tk_ad.dat, test_tk_cdi.dat, test_tk_ad_cdi.dat etc. when different i.c. requested
- if pk or tk requested at different redshift, several files, with extra suffix _z0, _z1, etc.

Following files created (or not) automatically (here we assume that root=test_), depending on the content of the input fields:

```
output= [tCl,pCl,lCl,nCl,sCl,mPk,dTk,vTk]
non linear= [none, halofit]
modes = [s,v,t]
ic = [ad, bi, cdi, nid, niv]
z_pk = [list of values]
```

- test_pk.dat matter power spectrum
- test_pk_nl.dat non-linear matter power spectrum
- test_pk_ad.dat, test_pk_cdi.dat, test_pk_ad_cdi.dat etc. when different i.c. requested
- test_tk.dat density and/or velocity transfer functions
- test_tk_ad.dat, test_tk_cdi.dat, test_tk_ad_cdi.dat etc. when different i.c. requested
- if pk or tk requested at different redshift, several files, with extra suffix _z0, _z1, etc.



Run with root=output/test2_, output=mPk, and without or with the extra line z_pk=0,0.4,0.8. In each case, look at output file names and headers,

Following files created (or not) automatically (here we assume that root=test_), depending on the content of the input fields:

```
write background = [yes,no]
write thermodynamics = [yes,no]
write primordial = [yes,no]
k_output_values = [list of values]
modes = [s,v,t]
```

• test_background.dat background quantities versus time and redshift

```
write background = [yes,no]
write thermodynamics = [yes,no]
write primordial = [yes,no]
k_output_values = [list of values]
modes = [s,v,t]
```

- test_background.dat background quantities versus time and redshift
- test_thermodynamics.dat thermodynamical quantities versus redshift

```
write background = [yes,no]
write thermodynamics = [yes,no]
write primordial = [yes,no]
k_output_values = [list of values]
modes = [s,v,t]
```

- test_background.dat background quantities versus time and redshift
- test_thermodynamics.dat thermodynamical quantities versus redshift
- test_primordial.dat primordial spectra (may follow from inflation simulation)

```
write background = [yes,no]
write thermodynamics = [yes,no]
write primordial = [yes,no]
k_output_values = [list of values]
modes = [s,v,t]
```

- test_background.dat background quantities versus time and redshift
- test_thermodynamics.dat thermodynamical quantities versus redshift
- test_primordial.dat primordial spectra (may follow from inflation simulation)
- test_perturbations_k*_s[vt].dat evolution of perturbations versus time

Following files created (or not) automatically (here we assume that root=test_), depending on the content of the input fields:

```
write background = [yes,no]
write thermodynamics = [yes,no]
write primordial = [yes,no]
k_output_values = [list of values]
modes = [s,v,t]
```

- test_background.dat background quantities versus time and redshift
- test_thermodynamics.dat thermodynamical quantities versus redshift
- test_primordial.dat primordial spectra (may follow from inflation simulation)
- test_perturbations_k*_s[vt].dat evolution of perturbations versus time



Run with root=output/test3_, output=tCl, k_output_values = 0.001,0.01.

Look at output file names and headers.

Exercise 0



The exercise text is located in the Dropbox, in the CLASS_exercise/ folder.

0

Use the information sent by ${
m CLASS}$ to the standard output (= displayed in the terminal) to check the characteristics of the Planck best-fit model presented in the Planck2013 Cosmological Parameter paper.