Non-interacting fermions

Two class of experimental systems

- Electrons are Fermions
- He3 are Fermions

Bosons and Fermions

- There are two types of particles in nature.
- In quantum mechanics, energies are discrete variables. We can label them by a subscript.
- For Fermions (Bosons), for any energy e_i , at most one (no limit) particles can have that energy.

Non-interacting Fermions

- For g the grand canonical potential,
- $e^{-g\beta} = \sum_{[n]} \exp{\{\beta[n_1(e_1-\mu) + n_2(e_2-\mu) + ...]\}} = \sum_{n_1} \exp{[\beta n_1(e_1-\mu)]} \sum_{n_2} \exp{[\beta n_2(e_2-\mu)]} ...$
- $\sum_{n} e^{-\beta(e-\mu)} = 1 + e^{-\beta(e-\mu)}$.
- Hence $e^{-g\beta}=\Pi_i$ [1+ $e^{-\beta(e_i-\mu)}$].
- $g = -\sum_{i} \ln [1 + e^{-\beta(e_i \mu)}] / \beta$.

•
$$e=\frac{p^2}{2m}$$
, Define $\alpha=e^{\beta\mu}$

•
$$g = -kTV \int d^3p \ln \left| 1 + \alpha e^{-\beta p^2/2m} \right| /h^3$$

- h³ is from quantum mechanics.
- The density $\rho = \frac{\langle N \rangle}{V} = -\partial \beta g/\partial (\beta \mu)/V$

•
$$\rho = \int d^3p \alpha e^{-\beta p^2/2m} / \left[1 + \alpha e^{-\beta p^2/2m} \right] / h^3$$

$$\bullet = \int d^3p / \left[1 + e^{\left(\frac{\beta p^2}{2m} - \mu\right)} \right]$$

•
$$e = \frac{p^2}{2m}$$
,

•
$$\rho = \int d^3p / h^3 / [1 + e^{(e-\mu)}]$$

- $f(e) = 1/[1 + e^{(e-\mu)}]$ is called the Fermi distribution function.
- $d^3p/h^3 = \frac{4\pi p^2 dp}{h^3} = g(e)de$, g(e) is called the density of states

•
$$g(e) = \frac{2\pi(2m)^{3/2}e^{1/2}}{h^3}$$

•
$$\rho = \int de \ g(e) f(e)$$



Thermodynamics

Fermion Hamiltonian Mean-field approximation

Thermodynamics

Mean occupancy Free electrons Zero temperature Chemical potential Estimated value Low temperatures Fermi delta function Integrals Sommerfeld expansion Chemical potential Expansions Final chemical potential Heat capacity Internal energy Final internal energy Heat capacity

Predicted behavior

 We found that the thermodynamic properties of a system of non-interacting fermions can be obtained from the Landau potential

$$\Phi = -k_{\rm B}T \log Z_G$$

$$= -k_{\rm B}T \sum_{\alpha} \log\{1 + \exp[\beta(\mu - \varepsilon_{\alpha})]\}.$$



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■ For example, the mean number of particles and the mean energy are

$$\langle N \rangle = \sum_{\alpha} \langle n_{\alpha} \rangle, \qquad \langle E \rangle = \sum_{\alpha} \langle n_{\alpha} \rangle \varepsilon_{\alpha},$$

in which the mean number of particles occupying the single-particle state α is

$$\langle n_{\alpha} \rangle = \frac{1}{\exp[\beta(\varepsilon_{\alpha} - \mu)] + 1}.$$



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$$\langle n_{\alpha} \rangle = \frac{1}{\exp[\beta(\varepsilon_{\alpha} - \mu)] + 1}.$$

■ Since the exponential function is never negative, we have

$$\langle n_{\alpha} \rangle \leq 1$$
,

which is crucially different from the boson case.



Mean occupancy

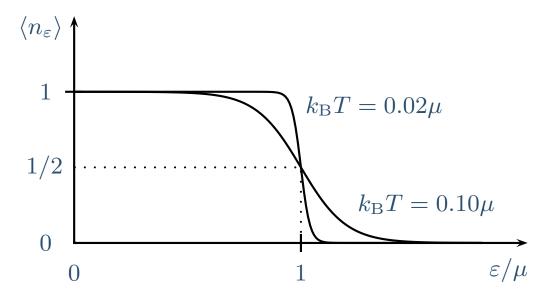
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The Fermi–Dirac occupancy function $\langle n_{\varepsilon} \rangle$ is plotted below for two different temperatures:





Mean occupancy

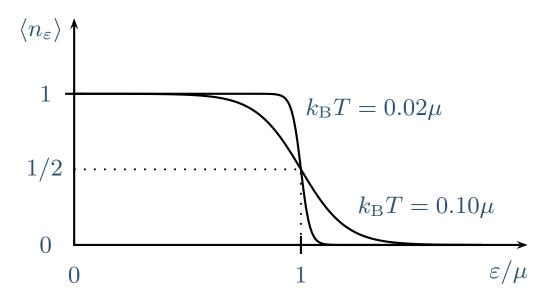
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Predicted behavior

The Fermi–Dirac occupancy function $\langle n_{\varepsilon} \rangle$ is plotted below for two different temperatures:



- lacksquare Note that the chemical potential μ depends on T.
- The value of μ is usually fixed by the particle-number condition

$$\langle N \rangle = \int_0^\infty \langle n_\varepsilon \rangle g(\varepsilon) \, d\varepsilon, \qquad \langle n_\varepsilon \rangle = \frac{1}{\exp[\beta(\varepsilon - \mu)] + 1}.$$



Free electrons

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Predicted behavior

For free fermions of spin s, the density of states is

$$g(\varepsilon) = (2s+1)\frac{V}{(2\pi)^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} \varepsilon^{1/2}.$$



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- The case of free electrons (s = 1/2) is particularly relevant for the study of metals.
- The particle-number condition is then

$$\langle N \rangle = \frac{V}{2\pi^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} \int_0^\infty \frac{\varepsilon^{1/2} d\varepsilon}{\exp[\beta(\varepsilon - \mu)] + 1}.$$



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As an exercise, you will show that the density $n=\langle N \rangle/V$ satisfies

$$n\Lambda^3 = (2s+1)f_{3/2}(z), \qquad z \equiv \exp \beta \mu,$$

where $f_{\nu}(z)$ is the Fermi–Dirac integral

$$f_{\nu}(z) = \frac{1}{\Gamma(\nu)} \int_0^\infty \frac{x^{\nu-1} dx}{z^{-1} \exp x + 1} = \sum_{k=1}^\infty (-1)^{k-1} \frac{z^k}{k^{\nu}}.$$



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How do free electrons behave at very low temperatures?



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Predicted behavior

- How do free electrons behave at very low temperatures?
- Let us suppose that when $T \to 0$, the chemical potential has a well defined limit:

$$\lim_{T\to 0}\mu(T)=\mu_0.$$



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■ The number of electrons at T=0 is then

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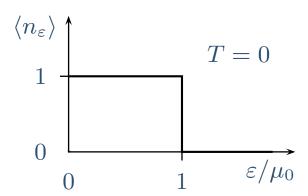
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From this, we see that the occupancy function $\langle n_{\varepsilon} \rangle$ is a step function with a discontinuity at $\varepsilon = \mu_0$:





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The number of electrons is therefore

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lacktriangle Upon solving for the chemical potential μ_0 , we get

$$\mu_0 = \frac{\hbar^2}{2m} (3\pi^2 n)^{2/3}, \qquad n \equiv \frac{\langle N \rangle}{V}.$$



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■ To estimate the magnitude of μ_0 , we can use atomic units:

$$\mu_0 = \frac{\hbar^2}{me^2} \frac{e^2}{2} (3\pi^2 n)^{2/3}$$
$$= a_0 \frac{e^2}{2} (3\pi^2 n)^{2/3}$$
$$= \frac{e^2}{2a_0} (3\pi^2 n a_0^3)^{2/3}.$$



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■ The conduction electron density in typical metals is

$$n \sim 10^{22} - 10^{23} \text{ cm}^{-3}$$
.



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Since the Bohr radius is $a_0 = 0.529 \times 10^{-8}$ cm, we have $a_0^3 \approx 1.5 \times 10^{-25}$ cm³ and

$$3\pi^2 n a_0^3 \lesssim O(1).$$



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■ The low-temperature chemical potential is therefore

$$\mu_0 = \frac{e^2}{2a_0} \times O(1)$$
 $\sim 13.6 \text{ eV}$
 $= 1.58 \times 10^5 \text{ K}.$

■ This number is huge on the scale of $T \approx 300 \text{ K}$ or $k_{\rm B}T \approx 0.0259 \text{ eV}$.



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- This number is huge on the scale of $T \approx 300 \text{ K}$ or $k_{\rm B}T \approx 0.0259 \text{ eV}$.
- As T increases, $\mu(T)$ decreases, eventually reaching the classical limit when $k_{\rm B}T\gg\mu_0$.
- Unlike the boson case, nothing interesting happens when $\mu(T) \approx 0$.



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How exactly does μ vary with temperature in the physically interesting range $k_{\rm B}T\ll\mu_0$?



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- To determine this, it helps to begin by examining the temperature dependence of the Fermi–Dirac distribution $\langle n_{\varepsilon} \rangle$.



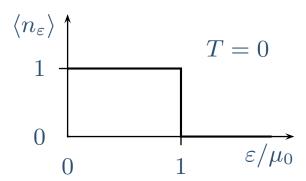
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- \blacksquare At T=0, we have already noticed the step-function behavior



■ The derivative of $\langle n_{\varepsilon} \rangle$ is therefore a Dirac delta function:

$$\frac{\partial}{\partial \varepsilon} \langle n_{\varepsilon} \rangle = -\delta(\varepsilon - \mu_0) \qquad (T = 0).$$



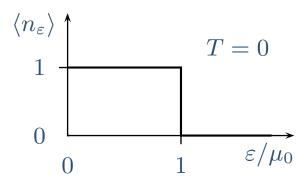
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For T > 0, we can also define a "Fermi delta function"

$$\Delta(\varepsilon) \equiv -\frac{\partial}{\partial \varepsilon} \langle n_{\varepsilon} \rangle.$$



Fermi delta function

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The Fermi delta function is given explicitly by

$$\Delta(\varepsilon) = -\frac{\partial}{\partial \varepsilon} \left(\frac{1}{\exp[\beta(\varepsilon - \mu)] + 1} \right)$$

$$= \left(\frac{1}{\exp[\beta(\varepsilon - \mu)] + 1} \right)^2 \beta \exp[\beta(\varepsilon - \mu)]$$

$$= \frac{1}{4k_B T} \frac{1}{\cosh^2[\beta(\varepsilon - \mu)/2]}.$$



Fermi delta function

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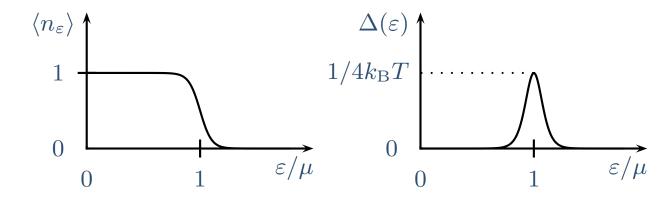
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This is an exponentially localized function with unit area, height $1/4k_{\rm B}T$, and width $\sim 4k_{\rm B}T$.



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■ We often need to evaluate integrals of the form

$$\int_0^\infty y(\varepsilon)\Delta(\varepsilon)\,\mathrm{d}\varepsilon,$$

where $y(\varepsilon)$ is a smooth function of ε .



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For T>0, it is convenient to expand $y(\varepsilon)$ about $\varepsilon=\mu(T)$:

$$y(\varepsilon) = y(\mu) + (\varepsilon - \mu)y'(\mu) + \frac{1}{2}(\varepsilon - \mu)^2 y''(\mu) + \cdots$$



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$$y(\varepsilon) = y(\mu) + (\varepsilon - \mu)y'(\mu) + \frac{1}{2}(\varepsilon - \mu)^2 y''(\mu) + \cdots$$

- If $k_{\rm B}T\ll\mu_0$, $\Delta(\varepsilon)$ has a sharp peak at $\varepsilon=\mu$, so we can extend the lower limit of integration to $-\infty$ with negligible error.
- All odd powers of $(\varepsilon \mu)$ then integrate to zero, because $\Delta(\varepsilon)$ is an even function of $(\varepsilon \mu)$.



Sommerfeld expansion

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For the even powers of $(\varepsilon - \mu)$, we get¹

$$\int_{-\infty}^{\infty} (\varepsilon - \mu)^n \Delta(\varepsilon) d\varepsilon = 2(k_B T)^n n! (1 - 2^{-n+1}) \zeta(n) \qquad (n \text{ even}).$$



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■ In this way, we obtain the Sommerfeld expansion

$$\int_0^\infty y(\varepsilon)\Delta(\varepsilon)\,\mathrm{d}\varepsilon = y(\mu) + \frac{\pi^2}{6}(k_\mathrm{B}T)^2y''(\mu) + \frac{7\pi^4}{360}(k_\mathrm{B}T)^4y^{(4)}(\mu) + \cdots.$$



Sommerfeld expansion

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■ For the even powers of $(\varepsilon - \mu)$, we get¹

$$\int_{-\infty}^{\infty} (\varepsilon - \mu)^n \Delta(\varepsilon) d\varepsilon = 2(k_{\rm B}T)^n n! (1 - 2^{-n+1}) \zeta(n) \qquad (n \text{ even}).$$

■ In this way, we obtain the Sommerfeld expansion

$$\int_0^\infty y(\varepsilon)\Delta(\varepsilon)\,\mathrm{d}\varepsilon = y(\mu) + \frac{\pi^2}{6}(k_\mathrm{B}T)^2y''(\mu) + \frac{7\pi^4}{360}(k_\mathrm{B}T)^4y^{(4)}(\mu) + \cdots.$$

We can also write this as

$$\Delta(\varepsilon) = \delta(\varepsilon - \mu) + \frac{\pi^2}{6} (k_{\rm B}T)^2 \delta''(\varepsilon - \mu) + \frac{7\pi^4}{360} (k_{\rm B}T)^4 \delta^{(4)}(\varepsilon - \mu) + \cdots,$$

with the understanding that this expression is to be used only inside an integral.

¹R. Kubo, Statistical Mechanics (North-Holland, 1967), pp. 231-233.



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- We can now use this approach to find the chemical potential $\mu(T)$ for $k_{\rm B}T\ll\mu_0$.
- lacktriangle As we found above, the condition that determines $\mu(T)$ is

$$\langle N \rangle = \frac{V}{2\pi^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} \int_0^\infty \varepsilon^{1/2} \langle n_\varepsilon \rangle \, \mathrm{d}\varepsilon.$$



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But we also know that

$$\langle N \rangle = \frac{V}{3\pi^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} \mu_0^{3/2}.$$

■ Hence, the first integral can be written as

$$\mu_0^{3/2} = \frac{3}{2} \int_0^\infty \varepsilon^{1/2} \langle n_\varepsilon \rangle \, \mathrm{d}\varepsilon.$$



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■ Hence, the first integral can be written as

$$\mu_0^{3/2} = \frac{3}{2} \int_0^\infty \varepsilon^{1/2} \langle n_\varepsilon \rangle \, \mathrm{d}\varepsilon.$$

■ Upon integrating by parts, we find

$$\mu_0^{3/2} = -\frac{3}{2} \int_0^\infty \left(\frac{2}{3} \varepsilon^{3/2}\right) \frac{\partial \langle n_\varepsilon \rangle}{\partial \varepsilon} \, \mathrm{d}\varepsilon = \int_0^\infty \varepsilon^{3/2} \Delta(\varepsilon) \, \mathrm{d}\varepsilon.$$



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$$\mu_0^{3/2} = \int_0^\infty \varepsilon^{3/2} \Delta(\varepsilon) d\varepsilon$$

$$= \mu^{3/2} + \frac{\pi^2}{6} (k_B T)^2 \left(\frac{3}{4} \mu^{-1/2}\right) + \frac{7\pi^4}{360} (k_B T)^4 \left(\frac{9}{16} \mu^{-5/2}\right) + \cdots$$

$$= \mu^{3/2} \left[1 + \frac{\pi^2}{8} \left(\frac{k_B T}{\mu}\right)^2 + \frac{7\pi^4}{640} \left(\frac{k_B T}{\mu}\right)^4 + \cdots\right].$$



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■ Upon bringing $\mu^{3/2}$ over to the left side, we find

$$\mu = \mu_0 \left[1 + \frac{\pi^2}{8} \left(\frac{k_{\rm B} T}{\mu} \right)^2 + \frac{7\pi^4}{640} \left(\frac{k_{\rm B} T}{\mu} \right)^4 + \cdots \right]^{-2/3}.$$



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■ A power series expansion then gives

$$\mu = \mu_0 \left[1 - \frac{\pi^2}{12} \left(\frac{k_{\rm B}T}{\mu} \right)^2 + \frac{\pi^4}{720} \left(\frac{k_{\rm B}T}{\mu} \right)^4 + \cdots \right].$$



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■ Finally, we can iterate this expression to obtain

$$\mu = \mu_0 \left[1 - \frac{\pi^2}{12} \left(\frac{k_{\rm B}T}{\mu_0} \right)^2 - \frac{\pi^4}{80} \left(\frac{k_{\rm B}T}{\mu_0} \right)^4 - \cdots \right],$$

which gives μ as an explicit function of T.



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which gives μ as an explicit function of T.

- How much does $\mu(T)$ differ from μ_0 ?
- For typical metals, we have

$$\mu_0 \sim 10^4 - 10^5 \text{ K}.$$



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- How much does $\mu(T)$ differ from μ_0 ?
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$$\mu_0 \sim 10^4 - 10^5 \text{ K}.$$

lacktriangle Hence, at room temperature $(T\sim300~{
m K})$, the fractional change is

$$\frac{\Delta\mu}{\mu_0} \sim \left(\frac{300}{10^4 - 10^5}\right)^2 \sim \frac{10^5}{10^8 - 10^{10}}$$
$$\sim 10^{-3} - 10^{-5},$$

which is very small.



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Let us now use these results to find the heat capacity of a system of free fermions (often used as a primitive model for a metal).



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- Let us now use these results to find the heat capacity of a system of free fermions (often used as a primitive model for a metal).
- By definition, the heat capacity is

$$C_V = \left(\frac{\partial U}{\partial T}\right)_{V,N},$$

where the internal energy is

$$U = \langle E \rangle = \int_0^\infty \varepsilon g(\varepsilon) \langle n_\varepsilon \rangle d\varepsilon$$
$$= \frac{V}{2\pi^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} \int_0^\infty \varepsilon^{3/2} \langle n_\varepsilon \rangle d\varepsilon.$$



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$$= \frac{V}{2\pi^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} \int_0^\infty \varepsilon^{3/2} \langle n_\varepsilon \rangle d\varepsilon.$$

Here we can divide by $\langle N \rangle$ (i.e., N in thermodynamic notation) to get rid of some constants:

$$\frac{U}{N} = \frac{3}{2} \mu_0^{-3/2} \int_0^\infty \varepsilon^{3/2} \langle n_\varepsilon \rangle \, \mathrm{d}\varepsilon.$$



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$$\frac{U}{N} = \frac{3}{2} \mu_0^{-3/2} \int_0^\infty \left(\frac{2}{5} \varepsilon^{5/2}\right) \left(-\frac{\partial \langle n_\varepsilon \rangle}{\partial \varepsilon}\right) d\varepsilon$$
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$$= \frac{3}{5} \mu_0^{-3/2} \int_0^\infty \varepsilon^{5/2} \Delta(\varepsilon) d\varepsilon.$$

■ The Sommerfeld expansion then gives

$$\frac{U}{N} = \frac{3}{5}\mu_0^{-3/2} \int_0^\infty \varepsilon^{5/2} \left[\delta(\varepsilon - \mu) + \frac{\pi^2}{6} (k_{\rm B}T)^2 \delta''(\varepsilon - \mu) + \cdots \right] d\varepsilon$$

$$= \frac{3}{5}\mu_0^{-3/2} \left[\mu^{5/2} + \frac{\pi^2}{6} (k_{\rm B}T)^2 \left(\frac{5}{2} \right) \left(\frac{3}{2} \right) \mu^{1/2} + O(k_{\rm B}T)^4 \right].$$



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$$= \frac{3}{5}\mu_0^{-3/2} \left[\mu^{5/2} + \frac{\pi^2}{6} (k_{\rm B}T)^2 \left(\frac{5}{2} \right) \left(\frac{3}{2} \right) \mu^{1/2} + O(k_{\rm B}T)^4 \right].$$

 \blacksquare At T=0, the energy per particle is simply

$$\frac{U}{N} = \frac{3}{5}\mu_0.$$

The ground-state energies of electron systems are therefore enormous $(\mu_0 \sim 10^5 \ {\rm K})$, which is important in the stability of matter.



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For the case T > 0, our previous result for the chemical potential

$$\frac{\mu}{\mu_0} \approx 1 - \frac{\pi^2}{12} \left(\frac{k_{\rm B}T}{\mu_0} \right)^2$$

gives

$$\left(\frac{\mu}{\mu_0}\right)^{5/2} \approx 1 - \frac{5}{2} \left(\frac{\pi^2}{12}\right) \left(\frac{k_{\rm B}T}{\mu_0}\right)^2.$$



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■ The internal energy then becomes

$$\frac{U}{N} \approx \frac{3}{5}\mu_0 \left[1 - \frac{5\pi^2}{24} \left(\frac{k_{\rm B}T}{\mu_0} \right)^2 + \frac{5\pi^2}{8} \left(\frac{k_{\rm B}T}{\mu_0} \right)^2 \right]
\approx \frac{3}{5}\mu_0 \left[1 + \frac{5\pi^2}{12} \left(\frac{k_{\rm B}T}{\mu_0} \right)^2 \right].$$



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\approx \frac{3}{5}\mu_0 \left[1 + \frac{5\pi^2}{12} \left(\frac{k_{\rm B}T}{\mu_0} \right)^2 \right].$$

Continuing the expansion one term further, we get

$$\frac{U}{N} = \frac{3}{5}\mu_0 \left[1 + \frac{5\pi^2}{12} \left(\frac{k_{\rm B}T}{\mu_0} \right)^2 - \frac{\pi^4}{16} \left(\frac{k_{\rm B}T}{\mu_0} \right)^4 + \cdots \right].$$



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■ To connect with experiment, we can now calculate the heat capacity

$$C_{V} = \left(\frac{\partial U}{\partial T}\right)_{V,N} = \frac{k_{\rm B}}{\mu_{0}} \left(\frac{\partial U}{\partial (k_{\rm B}T/\mu_{0})}\right)_{V,N}$$

$$= \frac{3}{5} N k_{\rm B} \frac{\partial}{\partial (k_{\rm B}T/\mu_{0})} \left[1 + \frac{5\pi^{2}}{12} \left(\frac{k_{\rm B}T}{\mu_{0}}\right)^{2} - \frac{\pi^{4}}{16} \left(\frac{k_{\rm B}T}{\mu_{0}}\right)^{4} + \cdots\right]$$

$$= \frac{3}{5} N k_{\rm B} \left[\frac{5\pi^{2}}{12} (2) \left(\frac{k_{\rm B}T}{\mu_{0}}\right) - \frac{\pi^{4}}{16} (4) \left(\frac{k_{\rm B}T}{\mu_{0}}\right)^{3} + \cdots\right].$$



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$$= \frac{3}{5} N k_{\rm B} \frac{\partial}{\partial (k_{\rm B}T/\mu_{0})} \left[1 + \frac{5\pi^{2}}{12} \left(\frac{k_{\rm B}T}{\mu_{0}}\right)^{2} - \frac{\pi^{4}}{16} \left(\frac{k_{\rm B}T}{\mu_{0}}\right)^{4} + \cdots\right]$$

$$= \frac{3}{5} N k_{\rm B} \left[\frac{5\pi^{2}}{12} (2) \left(\frac{k_{\rm B}T}{\mu_{0}}\right) - \frac{\pi^{4}}{16} (4) \left(\frac{k_{\rm B}T}{\mu_{0}}\right)^{3} + \cdots\right].$$

We can write this as

$$C_V = \frac{3}{2} N k_{\rm B} \left[\frac{\pi^2}{3} \left(\frac{k_{\rm B} T}{\mu_0} \right) - \frac{\pi^4}{10} \left(\frac{k_{\rm B} T}{\mu_0} \right)^3 + \cdots \right],$$

where $(3/2)Nk_{\rm B}$ is the classical result.

 \blacksquare Hence, at small T, the classical heat capacity is reduced by a factor of

$$\frac{\pi^2}{3} \left(\frac{k_{\rm B} T}{\mu_0} \right) \ll 1.$$



Predicted behavior

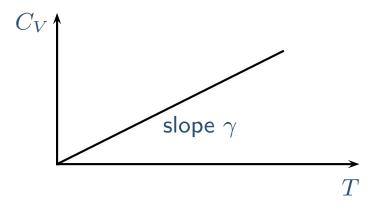
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Thus, we now have a prediction for the low-temperature $(k_BT \ll \mu_0)$ heat capacity of free and independent electrons:

$$C_V = \gamma T + O(T^3), \qquad \gamma = \frac{3}{2} N k_{\rm B} \left(\frac{\pi^2}{3} \frac{k_{\rm B}}{\mu_0} \right).$$



■ How does this compare with experiment?