The Central Limit Theorem And the Limits of the Central Limit Theorem

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Gaussian and Non-Gaussian

Two true statements:

Averaging over non-Gaussian random variables is an effective Gaussianization.

Non-linear functions of Gaussian random variables follow non-Gaussian distributions.

Which is your case?



Descriptive statistics

- Noncentral moments: $m_n = \langle x^n \rangle$
- Cumulants: $\kappa_n = m_n \sum_{k=1}^{n-1} \binom{n-1}{k-1} \kappa_k m_{n-k}$ (combinatorical superposition of moments)
- Gaussian: has infinitely many even moments, but only the first and second cumulant

Moment-generating function:

$$m_{x}(t) = \langle e^{tx} \rangle \tag{1}$$

Cumulant-generating function:

$$c_x(t) = \log(m_x(t)) \tag{2}$$

Factors *i* can appear when a solution in the complex plane exists, but not on the real axis.

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Take *n* iid random variables x_i , meaning $x_i \sim p(x) \ \forall i$.

Now take the mean

$$\bar{x} = \frac{1}{n} \sum_{i}^{n} x_i. \tag{3}$$

p(x) must have a finite mean μ and variance σ^2 ... but is not further specified and can be non-Gaussian.

Then, the sample average \bar{x} for $n \to \infty$ will be

$$\bar{x} \sim \mathcal{G}(\mu, \sigma^2/n),$$
 (4)

where $\mathcal{G}(\mu, \sigma^2/n)$ is the Gaussian of mean μ and variance σ^2/n . Equivalently, the whitened variable Y follows

$$Y = \frac{\bar{x} - \mu}{\sigma / \sqrt{n}} \sim \mathcal{N}(0, 1). \tag{5}$$

Proof: we show that the higher order cumulants die out.

We whiten the data points

$$z_i = \frac{x_i - \mu}{\sigma} \,\forall i,\tag{6}$$

such that by definition

$$\langle z_i \rangle = 0, \quad \langle z_i^2 \rangle = 1.$$
 (7)

In terms of the z_i , the variable Y is then

$$Y = \frac{1}{\sqrt{n}} \sum_{i=1}^{n} z_i,\tag{8}$$

and its moment-generating function is

$$m_Y(t) = \langle e^{tY} \rangle = \left\langle \exp\left(\frac{tz_i}{\sqrt{n}}\right) \right\rangle^n.$$
 (9)

Now expand the average into a power series and use linearity of the average:

$$\left\langle \exp\left(\frac{tz_i}{\sqrt{n}}\right) \right\rangle = \left\langle 1 + \frac{tz_i}{\sqrt{n}} + \frac{t^2 z_i^2}{2n} + \frac{t^3 z_i^3}{3! n^{3/2}} + \dots \right\rangle,$$
 (10)

the second term is zero due to $\langle z_i \rangle = 0$, and in the third term $\langle z_i^2 \rangle = 1$. Plug back into the power-n:

$$m_Y(t) = \left[1 + \frac{t^2}{2n} + \frac{t^3 \langle z_i^3 \rangle}{3! n^{3/2}} + \dots \right]^n$$

$$= \left[1 + \frac{1}{n} \left(\frac{t^2}{2} + \frac{t^3 \langle z_i^3 \rangle}{3! n^{1/2}} + \dots \right)\right]^n.$$
(11)

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Now introduce a shorthand for the power series

$$u = \frac{t^2}{2} + \frac{t^3 \langle z_i^3 \rangle}{3! n^{1/2}} + \dots,$$
 (12)

such that

$$m_Y(t) = \left[1 + \frac{u}{n}\right]^n. \tag{13}$$

For averaging over ever more samples, we have $n \to \infty$, and in this limit have

$$u \to \frac{t^2}{2},\tag{14}$$

because the other fractions will be supressed by the powers of n in their denominator. But each of these fractions contains a moment $\langle z_i^n \rangle$ and hence the higher moments die out, if we average over increasingly more samples n.

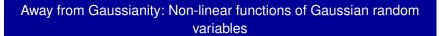
At the same time, for $n \to \infty$, we have

$$\lim_{n \to \infty} \left[1 + \frac{u}{n} \right]^n = e^u. \tag{15}$$

Consequently, in the limit of $n \to \infty$, the moment-generating function of Y is

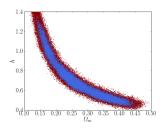
$$m_Y(t) = e^{\frac{t}{2}},\tag{16}$$

which is the Laplace transform of the standard normal distribution \mathcal{N} . Hence, the central limit theorem is proven. Most importantly, we also see the limit in which it arises: if n is large but finite, higher moments of the initial distribution p(x) can survive – if they exist.



Non-Gaussianity

Posteriors of Gaussian data will still remain forever non-Gaussian, if you have perfect parameter degeneracies no matter how many new data you get.



→ Uncertainty on parameters can be driven by the model, instead of by the data. ⇒ Discussion about degeneracy breaking in cosmology.

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Non-linear functions

Non-linear functions of Gaussian random variables follow non-Gaussian distributions, and the CLT agrees with this.

- the ratio of two Gaussian rv follows a Cauchy distribution
- the exponential of a Gaussian rv is log-normally distributed
- the absolute value of Fourier modes from a Gaussian random field follow a Rayleigh distribution
- many more ⇒ Exercises: Are astronomical magnitudes
 Gaussianly distributed? And what is the distribution of |x| if x is
 Gaussianly distributed?

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