

Project 1: Constraining modified gravity with large scale structure

Project Lead: Eva Mueller

Program per session (2 hours each , total 8 hours):

- 1 Intro to CosmoMC, modified gravity and growth rate measurements
- 2 Implementation of a new growth rate likelihood
- 3 Intro to MGCAMB, implementation of new model
- 4 Constrain modified gravity model, make plots etc.

Requirements (software, bibliography):

CosmoMC, getdist (comes with CosmoMC, needs Python 2.7+ or 3.4+), MGCAMB
(CosmoMC comes with CAMB and I want the students to implement the MG files so this doesn't need to be 'installed' only downloaded)

Are there any concepts (cosmology/statistics) that you would like the students review previous to your sessions?

Basic MCMC concepts, basic large scale structure

1. Gabriela Bárcenas
2. Bryan Sagredo
3. Cristhian García
4. Juan Carlos Ruelas Vázquez
5. Juan Flores
6. César Hernández Aguayo
7. Francisco Linares
8. Rebeca Martínez Carrillo

Project 2: Learning Galaxy Formation from Large-scale Structure and Weak Lensing

Project Lead: Ying Zu

Github page: https://github.com/nye17/macss2017_quenching

Program per session (2 hours each , total 8 hours):

1. Fitting Halo Mass to Weak Lensing Profiles.

- i. Introduce the weak lensing measurement of Mandelbaum+2016 on the average halo mass of galaxies at fixed stellar mass split by colors, $\langle M_h | M_* \rangle_{\text{red}}$ vs. $\langle M_h | M_* \rangle_{\text{blue}}$. (~20 mins)
- ii. Build a model for NFW profiles at fixed M_h and concentration. (~30 mins)
- iii. Vary parameters in the NFW profile on a grid and overplot the results against the measurements from i. (~20 mins)
- iv. Fit individual weak lensing profiles using simple curve fitting. (~20 mins)
- v. Fit individual weak lensing profiles using MCMC sampling, and compare the results to iv. (~20 mins)

2. Predicting Halo Mass from Theory.

- i. Interpret the results from session 1 and describe the steps for predicting them from theory. (~20 mins)
- ii. Measure halo mass function (HMF) and stellar-to-halo mass relation (SHMR) from simulated halos and mock galaxies (provided by me) (~30 mins)
- iii. Predict HMF and SHMR from theory (plenty of HMF code online!). (~30 mins)
- iv. Convolve HMF and SMHR to compute $\langle M_h | M_* \rangle$. (~30 mins)

3. Constructing Galaxy Quenching Models.

- i. Introduce the halo quenching model of Zu & Mandelbaum (2016). (~20 mins)
- ii. Implement a quenching model into the $\langle M_h | M_* \rangle$ code from session #2, and predict $\langle M_h | M_* \rangle_{\text{red}}$ vs. $\langle M_h | M_* \rangle_{\text{blue}}$. (~30 mins)
- iii. Measure $\langle M_h | M_* \rangle_{\text{red}}$ vs. $\langle M_h | M_* \rangle_{\text{blue}}$ from the mock catalog. (~20 mins)
- iv. Vary parameters in the quenching model on a grid and overplot the results against the measurements from iii. (~30 mins)

4. Model inference.

- i. Write the likelihood functions and run MCMC inferences (~ 30 mins).
- ii. Try different quenching models and repeat the analysis. (~ 30 mins)
- iii. Comment on model selection, quenching of satellite population, extra constraints from correlation functions, etc.

Requirements (software, bibliography):

software: Python, numpy, scipy, h5py, your favorite HMF and cosmology code (type "halo mass function" into github).

biblio:

Mandelbaum et al. 2016 (<http://adsabs.harvard.edu/abs/2016MNRAS.457.3200M>)

Zu and Mandelbaum 2015 (<http://adsabs.harvard.edu/abs/2015MNRAS.454.1161Z>)

Zu and Mandelbaum 2016 (<http://adsabs.harvard.edu/abs/2016MNRAS.457.4360Z>)

Are there any concepts (cosmology/statistics) that you would like the students review previous to your sessions?

None. Just ask questions during the sessions.

1. José Carlos Carvajal García
2. Hiram Kalid Herrera Alcantar
3. Osvaldo Rosales Pérez
4. Gabriel Karim Miranda Carrion
5. Mariana Jaber
6. Oleg Burgueño Gerardo
7. Fidel Sosa Nuñez
8. Brenda Izamar Tapia Benavies
9. Cynthia Cotto Parraguirre

Project 3: 21-cm Cosmology

Project Lead: Jordan Mirocha

Program per session (2 hours):

1. Background theory of 21-cm background and galactic foreground, development of simple phenomenological models for each.
2. The likelihood function for signal, foreground, and (perhaps) instrument in multiple sky regions.
3. Uncertainties in 21-cm measurements and priors on fitting parameters, leading to running of first MCMC fits of synthetic data.
4. Analysis of MCMC data. Parameter inference, relationship with independent measurements (e.g., tau, luminosity functions, 21-cm power spectrum, etc.)

Requirements (software, bibliography):

emcee

Are there any concepts (cosmology/statistics) that you would like the students review previous to your sessions?

Basic MCMC <http://dan.iel.fm/emcee/current/>

1. José Abraham Arvizu Valenzuela
2. David Florencia Cruz
3. Pablo Andrés Lizardo Romo
4. Hector Rios Hernandez
5. ALEJANDRA GUTIÉRREZ AGUIRRE
6. Josué Ely Molina Becerra
7. Elizabeth Moreno Hilario
8. Iliana Mairén Fernández-Roldán

Plenary Project: MCMC from scratch

Project Lead: Josue De Santiago and Juan Carlos Hidalgo

Program per session (2 hours):

1 Cosmology basics

- a) Type Ia supernovae as standard candles.
- b) SNIa datasets.
- c) Luminosity distance.
- d) Definition and computation of χ^2

2 Sampling algorithm

- a) Complementary theory
- b) Steps for the Metropolis method
- c) Implementation of the method
- d) Execution of a single walk

3 Extras

- a) Convergence criteria
- b) Other sampling methods
- c) Extra data
- d) Extra models

Requirements (software, bibliography):

- Python
- Python libraries: Scipy, numpy and corner
- SNIa data from the following links:
 - http://supernova.lbl.gov/Union/figures/SCPUnion2_mu_vs_z.txt
 - http://supernova.lbl.gov/Union/figures/SCPUnion2_covmat_sys.txt

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| José Abraham | Arvizu Valenzuela | 3 | |
| José Carlos | Carvajal García | 2 | 1 |
| Hiram Kalid | Herrera Alcantar | 2 | 1 |
| Osvaldo | Rosales Pérez | 2 | 1 |
| Gabriela | Bárcenas | 1 | 3 |
| Francisco | Cervantes | | |
| Josué Ely | Molina Becerra | 2 | 3 |
| Gabriel Karim | Miranda Carrion | 2 | 1 |
| David | Florencia Cruz | 3 | |
| Bryan | Sagredo | 1 | 3 |

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| Cristhian | García | 1 | 2 |
| Pablo Andrés | Lizardo Romo | 3 | |
| Hector | Rios Hernandez | 3 | |
| Juan Carlos | Ruelas Vázquez | 1 | 2 |
| Brenda Izamar | Tapia Benavies | 1 | 2 |
| Mariana | Jaber | 2 | 1 |

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| Oleg | Burgueño Gerardo | 2 | 1 |
| Juan | Flores | 1 | 2 |
| César | Hernández Aguayo | 1 | 2 |
| Francisco | Linares | 1 | 2 |
| Rebeca | Martínez Carrillo | 1 | 2 |
| ALEJANDRA | GUTIÉRREZ AGUIRRE | 3 | 2 |
| Irvin | Martínez | - | - |
| Elizabeth | Moreno Hilario | 3 | 2 |

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| Fidel | Sosa Nuñez | 2 | 1 |
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| Cynthia | Cotto Parraguirre | 1 | 2 |
| Iliana Mairén | Fernández-Roldán | 3 | 2 |
| Ma. Rafaela | Gonzalez Grana | | |

Proyecto 3

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| José Abraham | Arvizu Valenzuela |
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| David | Florencia Cruz |
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| Pablo Andrés | Lizardo Romo |
| Hector | Rios Hernandez |

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| ALEJANDRA | GUTIÉRREZ AGUIRRE |
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| Josué Ely | Molina Becerra | |
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| Elizabeth | Moreno Hilario |
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| Iliana Mairén | Fernández-Roldán |
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Proyecto 2

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| José Carlos | Carvajal García | 2 |
| Hiram Kalid | Herrera Alcantar | 2 |
| Osvaldo | Rosales Pérez | 2 |

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| Gabriel Karim | Miranda Carrion | 2 |

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| Mariana | Jaber | 2 |
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| Index | Code Name | 2 |
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| Brenda Izamar | Tapia Benavies | 2 |
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Proyecto 1

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| Gabriela | Bárcenas | 1 |
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| Bryan | Sagredo | 1 |
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| Cristhian | García | 1 |
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| Juan Carlos | Ruelas Vázquez | 1 |
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| Juan | Flores | 1 |
| César | Hernández Aguayo | 1 |
| Francisco | Linares | 1 |
| Rebeca | Martínez Carrillo | 1 |

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| Cynthia | Cotto Parraguirre | 1 |
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