# Chapter 14 Notes - MC

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#### 14 Vector Functions

# 14.1 Vector Functions and Space Curves

• vector-valued functions/vector functions - a function whose domain is a set of real numbers and whose range is a set of vectors. Written in terms of its components as a parametric equation:

$$\mathbf{r}(t) = \langle f(t), g(t), h(t) \rangle = f(t)\mathbf{i} + g(t)\mathbf{j} + h(t)\mathbf{k}$$

• Limits of a vector function: If  $\mathbf{r}(t) = \langle f(t), g(t), h(t) \rangle$ , then

$$\lim_{t \to a} \mathbf{r}(t) = \left\langle \lim_{t \to a} f(t), \lim_{t \to a} g(t), \lim_{t \to a} h(t) \right\rangle$$

provided the limits of the component functions exist.

• Space curves: suppose that f, g, and h are continuous real-valued functions on an interval I. Then the space curve is the set C of all points (x, y, z) in space, where

$$x = f(t)$$
  $y = g(t)$   $z = h(t)$ 

and t varies throughout the interval I.

# 14.2 Derivatives and Integrals of Vector Functions

• derivative of a vector-valued function:

$$\frac{d\mathbf{r}}{dt} = \mathbf{r}'(t) = \lim_{h \to 0} \frac{\mathbf{r}(t+h) - \mathbf{r}(t)}{h}$$

if the limit exists.

• Theorem 2: If  $\mathbf{r}(t) = \langle f(t), g(t), h(t) \rangle = f(t)\mathbf{i} + g(t)\mathbf{j} + h(t)\mathbf{k}$ , where f, g, and h are differentiable functions, then

$$\mathbf{r}'(t) = \langle f'(t), g'(t), h'(t) \rangle = f'(t)\mathbf{i} + g'(t)\mathbf{j} + h'(t)\mathbf{k}$$

• Theorem 3: Suppose  $\mathbf{u}$  and  $\mathbf{v}$  are differentiable vector functions, c is a scalar, and f is a real valued function. Then:

$$\frac{d}{dt}[\mathbf{u}(t) + \mathbf{v}(t)] = \mathbf{u}'(t) + \mathbf{v}'(t)$$

$$\frac{d}{dt}[c\mathbf{u}(t)] = c\mathbf{u}'(t)$$

$$\frac{d}{dt}[f(t)\mathbf{u}(t)] = f'(t)\mathbf{u}(t) + f(t)\mathbf{u}'(t)$$

$$\frac{d}{dt}[\mathbf{u}(t) \cdot \mathbf{v}(t)] = \mathbf{u}'(t) \cdot \mathbf{v}(t) + \mathbf{u}(t) \cdot \mathbf{v}'(t)$$

$$\frac{d}{dt}[\mathbf{u}(t) \times \mathbf{v}(t)] = \mathbf{u}'(t) \times \mathbf{v}(t) + \mathbf{u}(t) \times \mathbf{v}'(t)$$

$$\frac{d}{dt}[\mathbf{u}(t) \times \mathbf{v}(t)] = \mathbf{u}'(t) \times \mathbf{v}(t) + \mathbf{u}(t) \times \mathbf{v}'(t)$$

$$\frac{d}{dt}[\mathbf{u}(f(t))] = f'(t)\mathbf{u}'(f(t))$$
(chain rule)

• Integral of a vector function:

$$\int_{a}^{b} \mathbf{r}(t)dt = \left(\int_{a}^{b} f(t)dt\right)\mathbf{i} + \left(\int_{a}^{b} g(t)dt\right)\mathbf{j} \left(\int_{a}^{b} h(t)dt\right)\mathbf{k}$$

### 14.3 Arc Length and Curvature

• Length of a curve in 3D space:

$$L = \int_{a}^{b} \sqrt{[f'(t)]^{2} + [g'(t)]^{2} + [h'(t)]^{2}} dt = \int_{a}^{b} \sqrt{\left(\frac{dx}{dt}\right)^{2} + \left(\frac{dy}{dt}\right)^{2} + \left(\frac{dx}{dt}\right)^{2}} dt = \int_{a}^{b} |\mathbf{r}'(t)| dt$$

• curvature of a curve is defined as:

$$\kappa = \left| \frac{d\mathbf{T}}{ds} \right|$$

where T is the unit tangent vector.

 $\kappa(t) = rac{|\mathbf{T}'(t)|}{|\mathbf{r}'(t)|}$ 

ullet Theorem 10: The curvature of the curve given by the vector function  ${f r}$  is

$$\kappa(t) = \frac{|\mathbf{r}'(t) \times \mathbf{r}''(t)|}{|\mathbf{r}'(t)|^3}$$

• Equations for unit tangent, unit normal and binormal vectors, and curvature:

$$\mathbf{T}(t) = \frac{\mathbf{r}'(t)}{|\mathbf{r}'(t)|} \qquad \mathbf{N}(t) = \frac{\mathbf{T}'(t)}{|\mathbf{T}'(t)|} \qquad \mathbf{B}(t) = \mathbf{T}(t) \times \mathbf{N}(t)$$

$$\kappa = \left| \frac{d\mathbf{T}}{ds} \right| = \frac{|\mathbf{T}'(t)|}{|\mathbf{r}'(t)|} = \frac{|\mathbf{r}'(t) \times \mathbf{r}''(t)|}{|\mathbf{r}'(t)|^3}$$

# 14.4 Motion in Space: Velocity and Acceleration

• Velocity:

$$\mathbf{v}(t) = \mathbf{r}'(t)$$

- $\bullet\,$  speed is the magnitude of velocity.
- Parametric equations of trajectory:

$$x = (v_0 \cos \alpha)t$$
  $y = (v_0 \sin \alpha)t - \frac{1}{2}gt^2$ 

• Tangential and normal components of acceleration:

$$\mathbf{a} = v'(\mathbf{T}) + \kappa v^2 \mathbf{N}$$

- Kepler's laws:
  - A planet revolves around the sun in an elliptical orbit with the sun at one focus.
  - The line joining the sun to a planet sweeps out equal areas in equal times.
  - The square of the period of revolution of a planet is proportional to the cube of the length of the major axis of orbit.