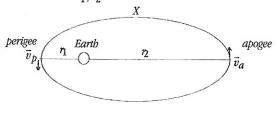
Universal Gravity

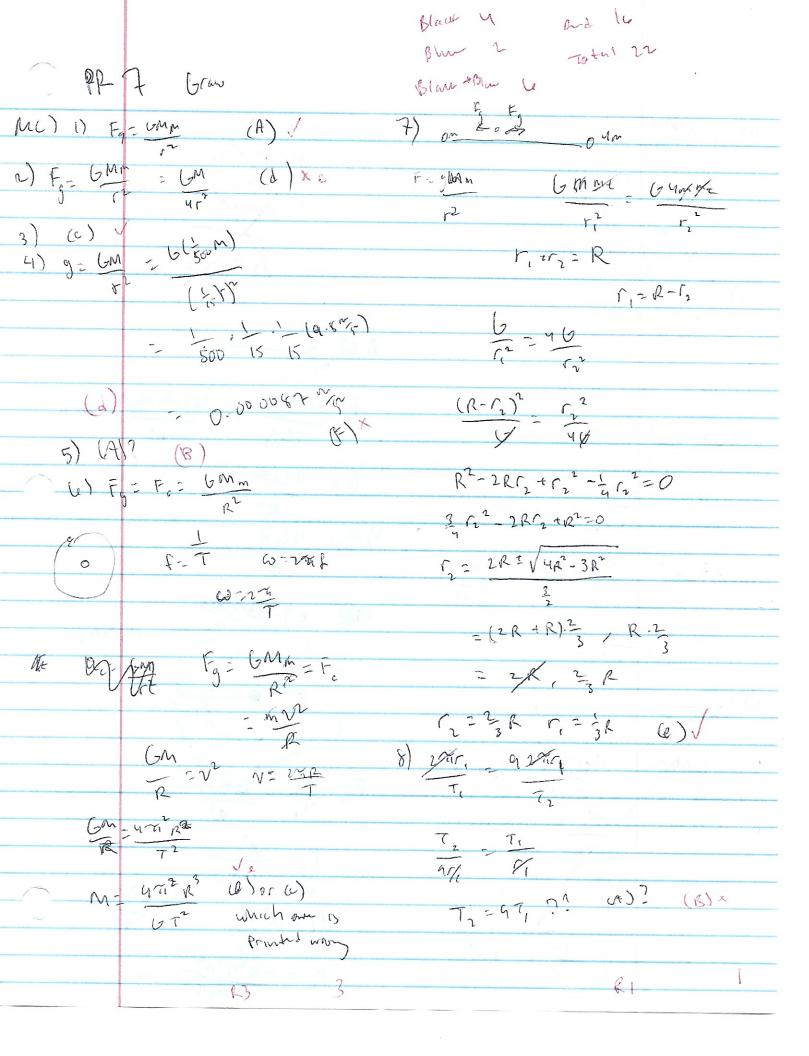
- If a distance be between two point particles is doubled, then the gravitational force between them.
 - A. decreases by a factor of 4
 - B. decreases by a factor of 2
 - C. increases by a factor of 2
 - D. increases by a factor of 4
 - E. Cannot be determined without knowing the masses
- On the surface of the earth, an object of mass m has weight w. If this object is transported to an altitude that's twice the radius of the earth, at the new location, its mass and weight are A. m/2, w/2 B. m, w/2 C. m/2, w/4 D. m, w/4 E. m, w/9 F.
- A moon of mass M orbits a planet of mas 100 M. If the strength of the gravitational force exerted by the planet on the moon is F_P and the strength of the gravitational force exerted by the moon on the planet is F_m , which statement is true?
 - A. $F_P = 100 \, F_m$
 - B. $F_P = 10 \, F_m$ D. $F_P = F_m / 10$ E. $F_P = F_m / 100$

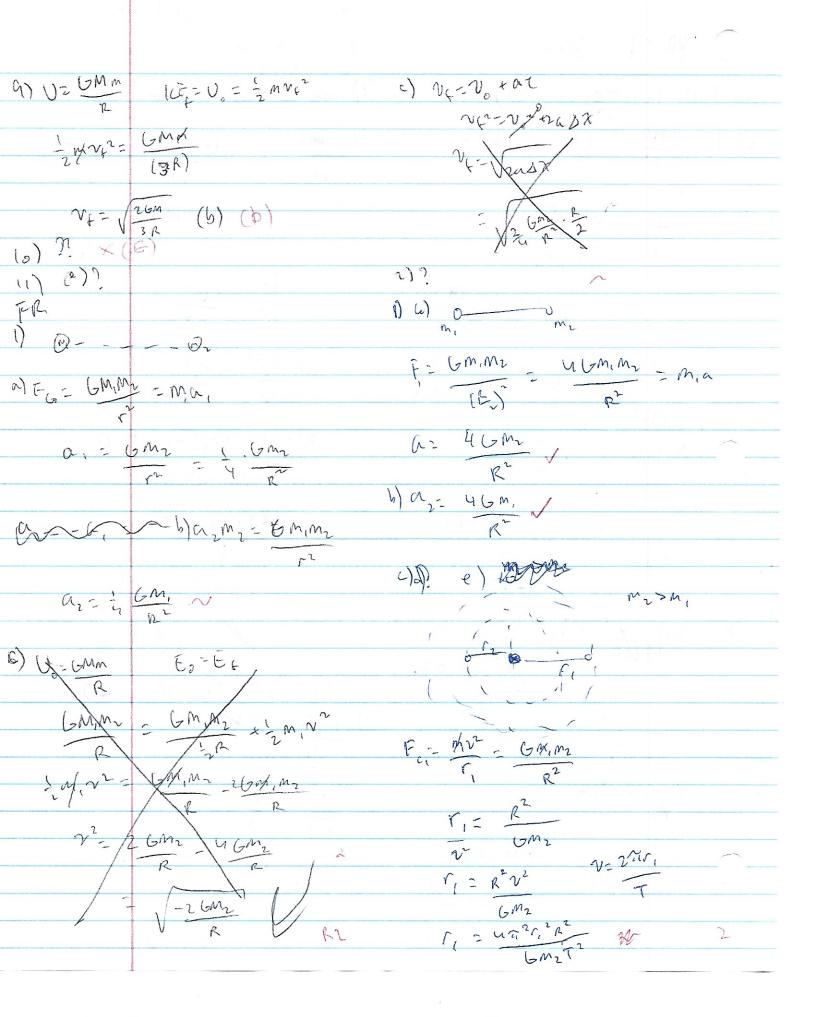
- Pluto has 1/500 the mass and 1/15 the radius of Earth. What is the value of g on its surface?
 - A. $0.3 \, m/s^2$ B. $1.6 \, m/s^2$ C. $2.4 \, m/s^2$ D. $4.5 \, m/s^2$ E. $7.1 \, m/s^2$ F
- 5. A satellite in a near circular orbit of radius R around Earth has kinetic energy K. When the satellite moves to a new orbit of radius 2R, what will be its new kinetic energy? A. K/4 B. K/2 C/K D. 2K E. 4K
- A moon of Jupiter has a nearly circular orbit of radius R and a period of T. What is the mass of Jupiter?
- B. $4\pi^2 R/T^2$
- C. $2\pi R^3/(GT^2)$
- D. $4\pi R^2/(GT^2)$ E. $2\pi R^3/(GT^2)$
- Two large bodies, A of mass m and B of mass 4m, are seperated by a distance R. At what distance from A, along the line joining the bodies, would teh gravitational force on an object be equal to zero?
 - A. R/16 B. R/8 C. R/5 D. R/4 E. R/3
- The mean distance from Saturn to the Sun is 9 times greater tan the mean distance from Earth to the Sun. How long is a Saturn year in Earth years?
 - A. 18 B. 27 C. 81 D. 243 E. 729 F.
- The Moon has mass M and radius R. A small object is dropped from a distance 3R from the Moon's center. The object's impact speed when it strikes the surface of the Moon is equal to

- A planet orbits the Sun in an elliptical orbit of eccentricity e. 10. What is the ratio of the planet's speed at perihelionto its speed at aphelion?
 - A. 1/(1-e) B. e/(1-e) C. 1/(1+e) D. e/(1+e) E. (1+e)/(1-e) F
- Two satellits A and B orbit a planet in circular orbits with radii R_A and $R_B = 3R_A$. What is the relation between the velocities?
 - A. $v_B = v_A$ B. $v_B = 3v_A$ C. $v_B = 9v_A$ D. $v_B = \sqrt{3}v_A$ E. $v_B = \frac{v_A}{\sqrt{3}}$

- Donsider two uniform spherical bodies in deep space with masses m_1 and m_2 . Starting from rest from a distance R apart, they are gravitationally attracted to each other.
 - Calculate the acceleration of m_1 when the spheres are a distance R/2 apart.
 - B. Calculate the acceleration of m_2 when the spheres are a distance R/2 apart.
 - Calculate the speed of m_1 when the spheres are a distance R/2 apart.
 - D. Calculate the speed of m_2 when the spheres are a distance R/2 apart.
 - Now assume that the spheres orbit their center of mass with the same orbital period T. Determine the radii of their orbits.
- A satellite of mass m is in an elliptical orbit.
 - Determine the speed v_D of the satellite at perigee (the point on the ellips closest to the Earth) in terms of η , η ,
 - Determine the speed v_a of the satellite at apogee (the point on the ellips fartest from the Earth) in terms of η , r_2 , M and G.
 - Express the ratio v_{p}/v_{a} in the simplest form.
 - What is the satellite's angular momentum when it is at
 - Ē. Determine the speed of the satellite when it is at the midpoint between perigee and apogee marked as X in in terms of η , η , M and G.
 - Determine the period of the satellite's orbit. F.
 - What is the eccentricity of the satellite's orbit, in terms of







1-42-116 r = Cm2 T2 Fr= min - 6 mins 12 = 2 - 6MI Cz= VRZ V= 24/2 52 - 40 12 P2 Gm, T2 1 = 47, 2 R2 5- 6M, 72 Mi) 2) man Fg = 6MM - 6 MM - 4 6 MM

i) (-e) 2) 4 Min, +m, v, =0 m, v, -m, v, =0 v, - M; v, (M2 = 1 m, v, 2 + 1 m2 (m, v) $\left(\frac{m_{1}}{R} - \frac{1}{2}\left(\frac{m_{1}+m_{2}\left(\frac{m_{1}}{m_{2}}\right)^{2}}{\sqrt{2}}\right)V_{1}^{2}\right)$ $= \sqrt{1 - m_{1}}\sqrt{\frac{2}{m_{1}}}\sqrt{\frac{m_{1}}{m_{2}}}$ S) 2 - Mi Vi - Mi Miz Vimer JMir - Mi V26 (Mither) Mir e) mini - 6 minz - m2 v2 m1 2/2/2 - m2 /2/2/2 (1, +C)3 = 4(1) =) (1, +C)=3 (6(m,+m)) T) (1,+C)3 = 6(m,+m) T) (1,+C)3 = 6(m,+m) T) $\Gamma_1 + \Gamma_2 = \frac{3}{3} \frac{b(m_1 + m_2)}{a_1 + b_2} = \frac{1}{3} \frac{4m_1}{m_2} \frac{\Gamma_1}{3}$ $100 + \Gamma_1 = \left(\frac{m_2}{m_1 + m_2}\right) \left(\frac{b(m_1 + m_2)}{a_2 + b_2}\right) \frac{3}{3}$ 3

2) n) = -6mm 20-5, tr f) Th = 47 > 7 = 120 and all KI = E - U = - 6 mm + 6 mm - CMM. FZ 5 mm, 2 - 6 mm . 52 VI JEM TO = (2-51) 14- E-1/2= - 6 mm = 6 mm - your 12 22 / CAMEN C) vi (ren) v V 1) 12 = CIMVI = CIM /26MC1 - m /2 6 Mr. (2) el a control 2000, to Xx = E-OX -- PWW + PWW 2 x 26M 5

CHAPTER 7 MC

上在 長以か→ア→2r, 長→上長

ZE R>3R F> JF

3. C Action-Reaction Func = - F

4- D J= GMP2 => J= GMP2

ON = G 500ME 2 = 225 G ME = 4.5 M

5-B === 6 mm = moz

SR-22 => KE-> KE-> KE

6 = F = F → 6 Mm = m 22 = m (2712)2 → M = 4722

G mm = G mm m 2 22

1n = 4m => (R-10)2=402

12= (1±2) x

ス=一里,星

If the m is between ABB

21 = 3.

83 T2X23 (see 6 above)

2 D C.E. KE+Vi=KEp+Vp

 $0-G\frac{Mm}{3R}=\frac{1}{2}mv_{2}^{2}-G\frac{Mm}{2}$ $\Rightarrow v_{1}=\frac{1}{4}GM$

10, € Co. Z: muz 2=mur > 2= 1

Ia Fa= E => m, q= G m, m, 9=46 m2 toward 2

≥ 1002 in a above 012=49 m1 toward 1

We external force on the two-sphio system > Pr is conserved. mr. +m ~ =0

m, v, -m, v= > v= m, v,

C.F. Kithi-Kythe

0-6 mm = 1 mur + 1 mur - 6 mm

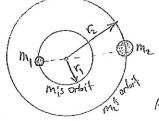
 $G\frac{m_2}{2} = \frac{1}{2}m_1v_1^2 + \frac{1}{2}m_1(\frac{m_1}{m_1}v_1^2)^2$

 $\Rightarrow G \frac{m_2}{R} = \frac{1}{2} \left(m_1 + m_2 \left(\frac{m_1}{m_2} \right)^2 \right) o_1^2$ $v_1^2 = m_2 \sqrt{\frac{2G}{(m_1 + m_2)m_1 R}}$

(1) from (1)

12= m1 0= m1 m2 26 (m+m) Rm ~= m1 (29) (m+m) Rm

(e) I. Gravitational is the source of the centripetal motion.



 $\frac{R}{M_{\text{A}}} = \frac{R}{R} \frac{(R+R)_{\text{S}}}{(R+R)_{\text{S}}}$

Action-reaction

m2 = m 2

 $\Rightarrow \frac{M_1}{K} \left(\frac{2\pi K}{T} \right)^2 = \frac{M_2}{K} \left(\frac{2\pi K}{T} \right)^2 ,$

> MIG = M25

(1)

See derivation of Keplert 3rd law: Substitute from above (1) $\Sigma = \frac{m_1}{m_2}$ (1) $\Gamma = \frac{m_1}{m_2}$ (1) $\Gamma = \frac{m_1}{m_2}$ $r_1 + \frac{m_1}{m_2} r_1 = \frac{3}{3}$ " $\Rightarrow \sqrt{-\left(\frac{m_2}{m_1+m_1}\right)\left(\frac{G(m_1+m_2)^{-2}}{1/172}\right)^3}$ $K_1 = t - U_1 = -6 \frac{mM}{r + r} - \left(-6 \frac{mM}{r}\right)$ K= GMM E K= GMM T => 12 = 2GM T & From a 86 $\frac{V_1}{V_2} = \frac{\sqrt{\frac{2GMr_2}{m_1}r_1}}{\sqrt{\frac{2GM}{m_1}r_2}} = \frac{r_1}{r_1}$ or use conservation of angular man $MNN = MV2 = \frac{V}{V} = \frac{V}{V}$

 $\frac{T^{2}}{(\eta+r_{2})^{3}} = \frac{4\pi^{2}}{G(m_{1}+m_{2})} \times r_{1}+r_{2} = \sqrt[3]{\frac{G(m_{1}+m_{1})T^{2}}{4\pi^{2}}} = \frac{1}{2} = r_{2} + r_{1} + r_{2} + r_{3} + r_{4} + r_{5} +$ = K to c. of Earth a= 1 (vitra) Kx = E-Ux = - 6 mM - (- 6 mM) Kx= G mM , Kx= 2 mn/2 => ~= [2 BM] $r_2 = \frac{m_1}{m_2} r_1 = \left(\frac{m_1}{m_1 + m_2}\right) \left(\frac{G(m_1 + m_2)}{24\pi r^2} - 2\right)^{1/3}$ $f = \frac{m_1}{m_2} r_1 = \left(\frac{m_1}{m_1 + m_2}\right) \left(\frac{G(m_1 + m_2)}{24\pi r^2} - 2\right)^{1/3}$ $f = \frac{m_1}{m_2} r_1 = \left(\frac{m_1}{m_1 + m_2}\right) \left(\frac{G(m_1 + m_2)}{24\pi r^2} - 2\right)^{1/3}$ $f = \frac{m_1}{m_2} r_1 = \left(\frac{m_1}{m_1 + m_2}\right) \left(\frac{G(m_1 + m_2)}{24\pi r^2} - 2\right)^{1/3}$ $f = \frac{m_1}{m_2} r_1 = \left(\frac{m_1}{m_1 + m_2}\right) \left(\frac{G(m_1 + m_2)}{24\pi r^2} - 2\right)^{1/3}$ $f = \frac{m_1}{m_2} r_1 = \left(\frac{m_1}{m_1 + m_2}\right) \left(\frac{G(m_1 + m_2)}{24\pi r^2} - 2\right)^{1/3}$ $f = \frac{m_1}{m_2} r_1 = \frac{m_1}{m_1 + m_2} \left(\frac{G(m_1 + m_2)}{24\pi r^2} - 2\right)^{1/3}$ $f = \frac{m_1}{m_2} r_1 = \frac{m_1}{m_1 + m_2} \left(\frac{G(m_1 + m_2)}{24\pi r^2} - 2\right)^{1/3}$ $\frac{2}{3} = \frac{4\pi^2}{6M}$, $\frac{1}{2} = \frac{4\pi^2}{6M} = \frac{1}{2} = \frac{4\pi^2}{6M} = \frac{1}{2} = \frac$ T=T1/(1+42)3] $\frac{1}{2}mv_{1}^{2} = G\frac{mM}{(r_{1}+r_{2})}\frac{r_{2}}{r_{1}} \Rightarrow v_{1}^{2} = \sqrt{\frac{26M}{r_{1}}\frac{r_{2}}{r_{2}}} \qquad e = \frac{c}{a} = \frac{a-r_{1}}{a} = \frac{\frac{1}{2}(r_{1}+r_{2})-r_{1}}{\frac{1}{2}(r_{1}+r_{2})} = \frac{r_{2}-r_{1}}{r_{2}+r_{1}}$