**Estimation of Cloud Fraction over the South-West Indian Ocean using BSRN Surface Observations**

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# Abstract

Cloudiness is the key to understand the variability of surface solar radiation, and to improve its predictability, which is particularly important for photovoltaic (PV) electricity production. This study estimates Cloud Fraction (CF) with observations from the BSRN (Baseline Solar Radiation Network) station at the island of La Réunion (55.5°E, 20.9°S) in the South-West Indian Ocean, using a XGBoost machine-learning model. Besides the statistical XGBoost model, a physical model known as the Automatic Partial Cloud Amount Detection Algorithm (APCADA) is also applied as a benchmark to estimate CF using LWD, surface air temperature and relative humidity data. CF data are obtained from collocated and coincident observations from a SkyCamVision all-sky camera within the UV-Indien network. After testing and validating the quality of the data with a set of quality control (QC) procedures derived from the BSRN recommendations, a XGBoost model is applied to estimate CF. A 2-year dataset (September 2019-February 2021) with a 5-min time step is used for this study, with 90% of it used for training and 10% for validation and test. After feature selections and model optimizations, a final score of 92.02% is achieved, which is comparable to the results from recent studies. Compared with APCADA, XGBoost performs better in terms of correlation (0.93 vs 0.85), Root Mean Square Error (0.12 vs 0.31) and Mean Absolute Error (0.08 vs 0.21). The statistical Machine Learning model, like XGBoost, could be potentially used to extend the temporal coverage of available CF observations, at a better accuracy compared to the traditional physical model such as APCADA.

Keywords: cloud fraction, longwave radiation, BSRN, XGBoost, Reunion Island, APCADA

# Introduction

## Importance of cloud:

Clouds play an important role in Earth's energy balance by reflecting incoming solar radiation back to space and trapping outgoing longwave radiation from the Earth's surface. With the strong attenuation effect of Surface Solar Radiation (SSR), clouds can have a significant impact on the output of solar photovoltaic (PV) systems, and thus responsible for a large part of the intermittency of solar energy resource at various temporal and spatial scales ([Maimó-Far et al., 2020](#_ENREF_17)), which challenges one of the primary goals of electric utilities, which is to balance supply and demand.

Clouds also play an important role in the Earth's water cycle, as they transport and distribute water vapor around the globe. This can have important implications for the availability of water resources, as well as the frequency and intensity of precipitation events.

Understanding the formation, evolution of cloud and its complex meteorological and energetic effect of clouds is the fundamental for weather forecasting, climate change, and optimizing solar energy technologies, which calling the needs of cloud dataset of large spatial and temporal coverage and of high quality.

## Cloud observation

### History of CF observation:

Cloudiness could be measured as Cloud Cover or Cloud Fraction (CF), as the fraction of the sky covered clouds which could be derived from human observation, meteorological satellites and surface remote observation [Kim and Cha, 2020](#_ENREF_12). Human observation refers to a regular observer recording cloud cover data at either 1, 3, or 6 h interval. To minimize human errors, ground- and aircraft-based measurements have been introduced to timely record 1-D cloud snapshots [Qian et al., 2012](#_ENREF_22). In the global observation trend, remote sensing has been applied to evaluate global cloud properties and estimate cloud cover parameters over years from visible and infrared radiations [Qian et al., 2012](#_ENREF_22).

Use of camera-based devices in cloud coverage observation improved spatiotemporal resolution, and image quality with a fish-eye lens (180o field of view), and ability to provide continuous valuable cloud datasets. Even though data quality depends on the capacity of the camera, cloud cover detection with sunlight showed great performance during the day. Responding to the absence of visible light during nighttime, infrared sensors and filters have been developed to detect radiance from clouds [Kim and Cha, 2020](#_ENREF_12).

Presence of long-term continuous ground-based measurements has attracted/created statistical and machine learning (ML) research interests. Forsythe

CF, short for fraction of area occupied by clouds, found to have a significant correlation with air temperature, precipitation, and topography [Didier, 2015](#_ENREF_5); [Forsythe et al., 2015](#_ENREF_10). After several studies of ML techniques in atmospheric radiative transfer studies, due to its ability to address nonlinear complexity problems, some studies were done in CF estimation (ref).

### CF from Climate model

In cloud-resolving models (CRMs), neural networks-based cumulus parameterization accurately diagnosed both precipitation and CF with little instability for multiple time steps prediction (brenowitz2018prognostic)).

By learning from CRMs simulations, deep neural networks achieved a more accurate convection and rapid execution on accelerators like Graphical Processing Units (GPUs) (ukkonen2020accelerating).

Statistically, Yunfeng Cao (2022) - with random forests (RF), extremely randomized trees (ERT), and categorical boosting (CatBoost) - found a high dependence of atmospheric temperature and water vapour on the Clear-sky longwave downward radiation.

Our case study is the tropical island of La Réunion (21°S; 55.5°E) in the South-West Indian Ocean (SWIO), where SE is one key in the transition to an electricity system 100% renewable by 2030 ([Selosse et al., 2018](#_ENREF_23), more refs). Despite the high solar resource availability over the island ([Mialhe et al., 2020](#_ENREF_18); Tang 2023; Morel; Bessafi), the development of SE there, could be hindered by drawbacks associated with the large variability of cloudiness over the island ([Badosa et al., 2013](#_ENREF_2); [Durand et al., 2021](#_ENREF_7); [Vérèmes et al., 2019](#_ENREF_26)).

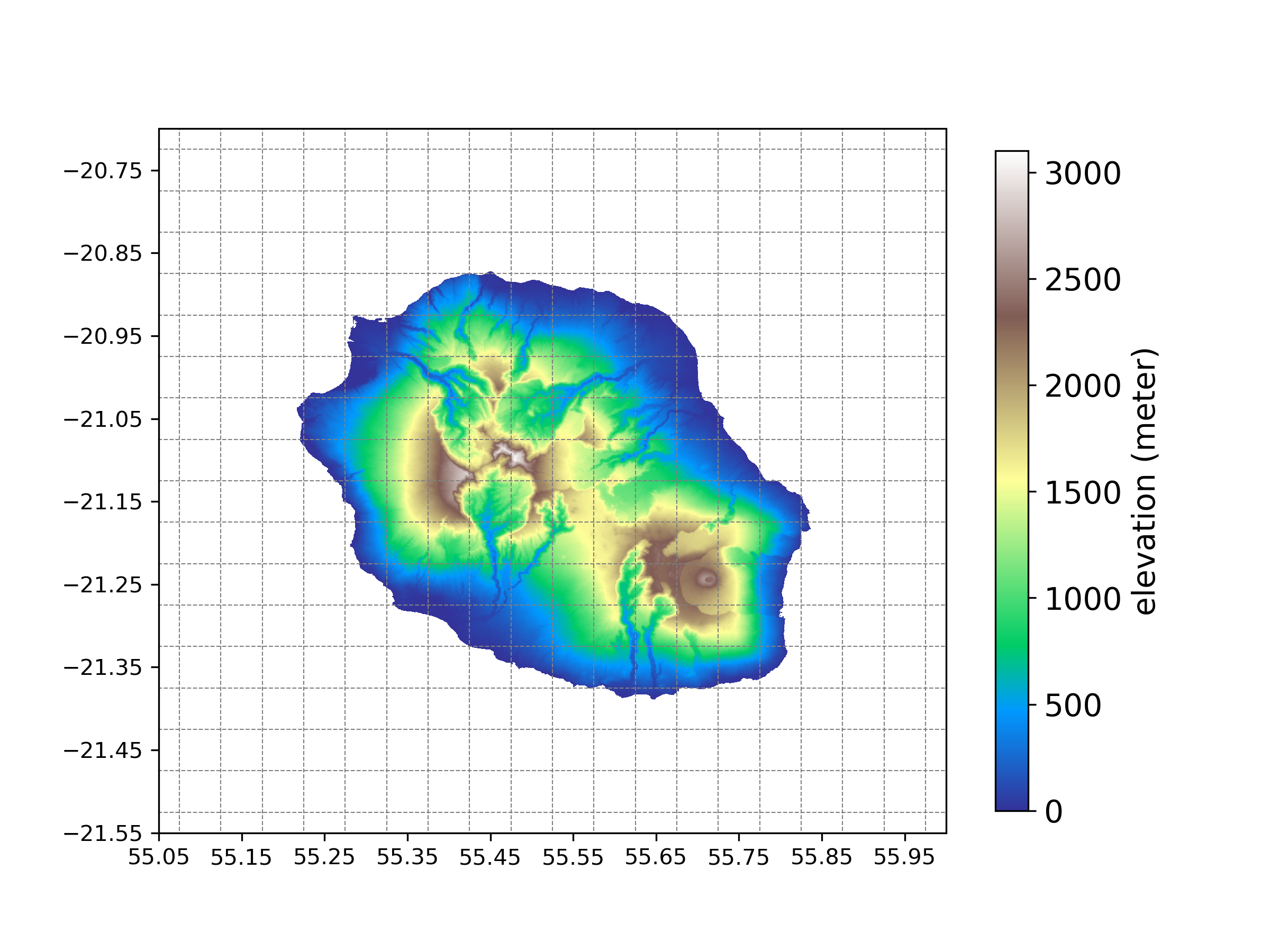


Fig. 1 Topography of La Réunion (in m) on a 0.05-degree grids (dotted gray lines). The map is based on the ASTER Global Digital Elevation Model from NASA Jet Propulsion Laboratory (see [ASTER (2019)](#_ENREF_1)).

Located in the South-West Indian Ocean, La Réunion island is mainly under an east-southeast trade wind flow (reunion mau circulation ref), which changes in direction according to the season: easterly in summer and more southeasterly in winter Because of the topographical characteristics of the island (Fig. 1), this main flow is more a “flow around” than a “flow over '' regime ([Mialhe et al., 2020](#_ENREF_18)). The main cloud regimes over La Réunion are (a) orographic clouds generated by the local topography and (b) clouds caught on the island driven by synoptic systems [Badosa et al., 2013](#_ENREF_2).

chao: shall we put some sentences of climate of Reunion somewhere (here or data, where we show the target variable from LACy (data) ?),

Todo: especially the information linked to cloud, e.g., the wet dry seasons, etc. to interpret the results which may be season / time -dependent.

### LW correlation with CF

In 2013, while estimating daytime downward longwave radiation under both cloudy and clear skies, Facundo Carmona found a high correlation between CF and longwave with multiple linear regression. Several studies have been done estimating longwave with cloud fractions as input, but no work done yet the other way around with ML.

However, to our best knowledge, no previous study has dealt with CF estimation based on LW.

In this paper, we estimated CF with the Reunion new station of the Baseline Surface Radiation Network (BSRN). Two methods were used: XGBoost, a ML model that uses a tree and a sequential neural network, and a physical model: Automatic Partial Cloud Amount Detection Algorithm (APCADA). The results were compared by using Mean Absolute Error (MAE), Root Mean Square Error (RMSE) and coefficient of correlation (r).

# Methods and data

## Data

### Reunion BSRN-station data

The Reunion BSRN station is in the main campus of the University of Reunion Island (21.00° S, 55.45° E) in the city of Saint-Denis, on the northern coast of the island (see Fig. 1). The station is managed by the ENERGY-Lab.

As one of the BSRN global network, data from the Reunion BSRN station, passed an automatic quality control (AQC) as described in [Long and Shi (2008)](#_ENREF_16) and visual quality control (VQC) before publishing. The Reunion station records several variables at 1-minute interval: global horizontal shortwave irradiance (GSW, in W/m2), diffuse horizontal shortwave irradiance (SWDif, in W/m2), direct normal shortwave irradiance (SWDir, in W/m2), downward longwave irradiance (LWD, in W/m2), temperature (T, in °C), relative humidity (RH, in %), and pressure (P, in hPa). Unshaded and shaded CMP22 pyranometer measures GSW, and SWDif respectively, CHP11 pyrheliometer is used to measure SWDir, CGR4 pyrgeometer for LWD, and WXT530 weather transmitter for T, RH, and P (see Table 1).

Table 1 Summary of data used in this study. The temporal resolution of all these variables is 5-minute.

|  |  |  |  |  |  |
| --- | --- | --- | --- | --- | --- |
| **Variable** | **Full name** | **unit** | **data source (device)** | **XGBoost** | **APACADA** |
| T | near surface air temperature | ° C | BSRN station | predictor | input |
| (WXT530 weather transmitter) |
| GSW | global horizontal shortwave irradiance | W/m2 | BSRN station | predictor | - |
| (unshaded CMP22 pyranometer) |
| SWDir | direct normal shortwave irradiance | W/m2 | BSRN station | predictor | - |
| (CHP11 pyrheliometer) |
| SWDif | diffuse horizontal shortwave irradiance | W/m2 | BSRN station | predictor | - |
| (shaded CMP22 pyranometer) |
| LWD | downward longwave irradiance | W/m2 | BSRN station | predictor | input |
| (CGR4 pyrgeometer) |
| RH | relative humidity | % | BSRN station | predictor | input |
| (WXT530 weather transmitter) |
| P | pressure | hPa | BSRN station | predictor | - |
| (WXT530 weather transmitter) |
| CF | cloud fraction | % | UV-Indien network | predictand | output |
| (allsky camera) |

In addition to the QCs, the LWD data are also compared to CERES (Clouds and Earth’s Radiant Energy Systems) satellite-based product, SYN1deg-1Hour [Smith et al., 2011](#_ENREF_25). The LWD of CERES is retrieved from the nearest pixel of its 1°x1° grid. All available data from Reunion BSRN station (June 2019 - February 2021) was used for the comparison, and RMSE, MBE, and correlation efficient r are calculated at hourly, daily, and monthly timescales. As shown in Fig. 2 a correlation of 76, 90, and 98%; RMSE of 20.57, 11.36, and 5.76 W/m2 and MAE of 16.21, 9.04 and 5.10 W/m2 between LWD of the BSRN station and CERES is found at hourly, daily, and monthly timescales respectively. These statistics are comparable to previous works on CERES LWD data, such as the regional study [Yan et al. (2011)](#_ENREF_27), [Sheng et al. (2009)](#_ENREF_24), [dos Santos Nascimento et al. (2018)](#_ENREF_6) and the global validation of CERES data quality summary ([NASA, 2021](#_ENREF_20)).

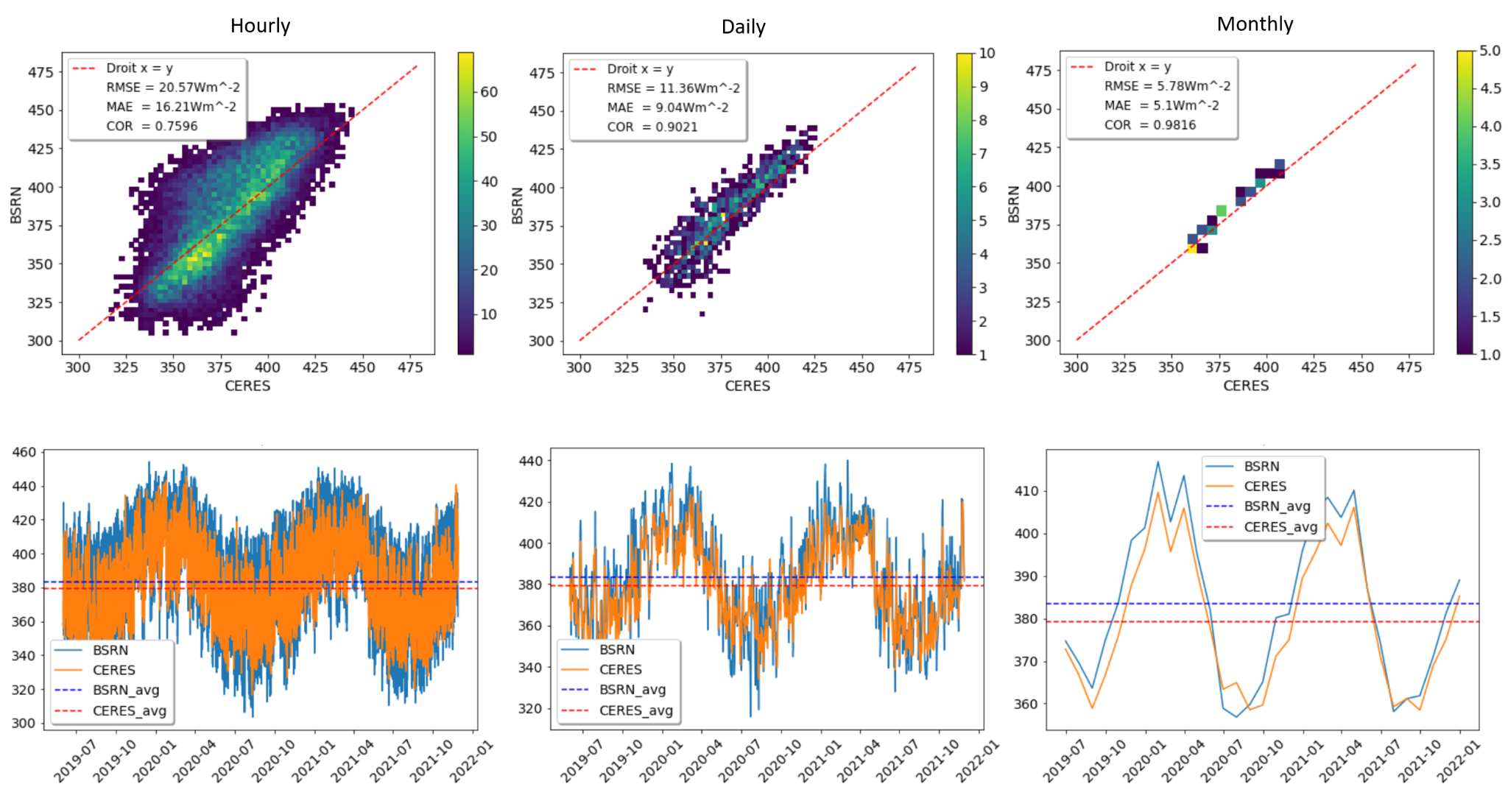


Fig. 2 Comparison of LWD from Reunion BSRN station and CERES, at timescales of hourly, daily, and monthly timescales in the period of June 2019 - February 2022.

### CF measurements from the UV-Indien network

CF data are obtained from collocated and coincident observations from a SkyCamVision all-sky camera (<http://www.reuniwatt.com/>) within the UV-Indien network [Lamy et al., 2021](#_ENREF_13). This camera is located less than 100 meters from the Reunion BSRN station. It equipped with a fisheye-lens CMOS Sensor of 1600 x 1200 pixels resolution and acquires 1-minute hemispherical images in the visible range (380-440 nm). From images, pixels are classified into clear sky; sun; thick cloud; or thin cloud with cloud segmentation algorithm. CF is computed from pixels with clear and cloudy sky, geometrically calibrated image [Cadet et al., 2020](#_ENREF_3). As the camera is not mounted with infrared sensors, the station does not record night-time CF. In this study, only daytime CF observations were used in the period of **2019-09-13 to 2021-02-01**.

CF from 2016-12 to present (2021-02)

LW from 2019-06 to present.

#### To-do: seasonal data test.

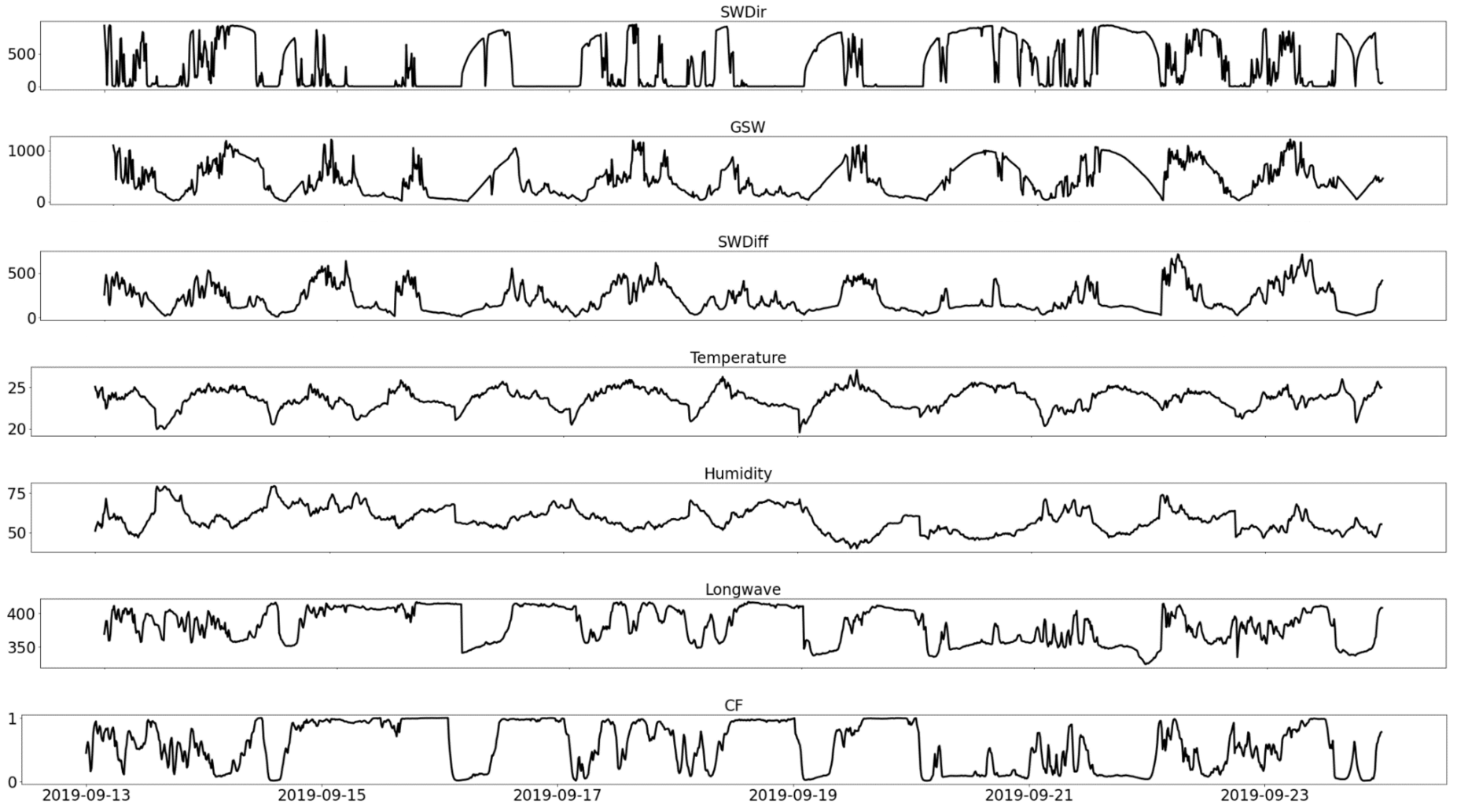


Fig. 3 One-week example of time series of the measurements used in this study.

## Automatic Partial Cloud Amount Detection Algorithm (APCADA)

A method of Automatic Partial Cloud Amount Detection Algorithm (APCADA) is developed by [Dürr and Philipona (2004)](#_ENREF_8) to estimate the cloud amount without high clouds directly from longwave downward radiation (LWD), air temperature, and humidity.

The determination of partial cloud amount according to APCADA is based on two parameters: the cloud-free index (CFI) and the variability of longwave downward radiation (STD LWD). The CFI, a ratio of apparent emittance of the sky to the empirical apparent cloud free emittance, is used to distinguish between clear or cloudy conditions. It is calculated as:

Eq. 1

where is the Stefan–Boltzmann constant, is the air temperature in Kelvin, and the emissivity of a cloud-free sky, which is defined as:

Eq. 2

where is altitude-dependent emittance of dry atmosphere, is a location-dependent coefficient, is air temperature (in K), and is water vapor pressure (Pa). In this study the emissivity of a cloud-free sky, i.e., is defined by fitting Eq. 2 in the clearsky condition selected by the method of [Long and Ackerman (2000)](#_ENREF_15) in Reunion during the period of September 2019 - February 2021 (Fig. 4).

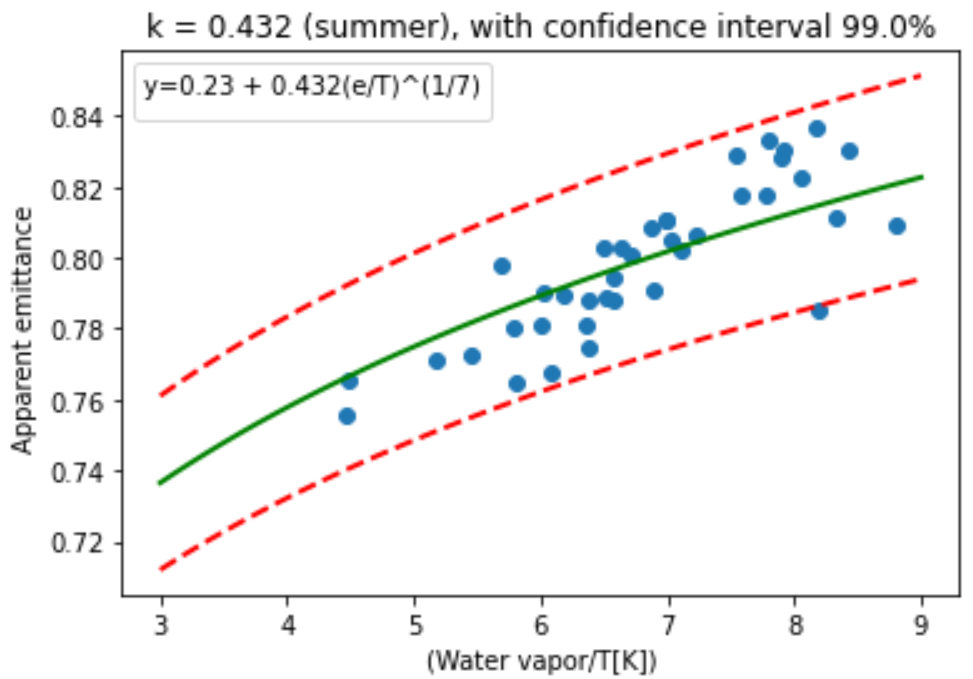


Fig. 4 Apparent emittance vs ratio of water vapor pressure and temperature at Reunion for September 2019 - February 2021 period.

According to [Dürr and Philipona (2004)](#_ENREF_8), besides CFI, variability of LWD is also used, as it allows the distinction between cloud fraction types: broken clouds strongly influence the variability signal, while overcast and cloudless skies lead to a low variability. Since LWD measured at the Earth’s surface is marginally affected by high clouds because of large distance and cold emittance temperature, as a drawback, APCADA can detect only total cloud amounts without high clouds (hereinafter referred to as partial cloud amount, PCA). A dataset on frequency of 5-min from September 2019 to February 2021 is used in this study, according to the availability of CF measurements.

## XGBoost model

### Variable selection

See:

Math of XGBoost see: [Zhang et al., 2018](#_ENREF_28).

Extreme gradient boosting (XGBoost) regression is an ensemble machine learning algorithm that is widely used in data mining with excellent performance ([Chen et al., 2019](#_ENREF_4)).In contrast to some other machine learning models such as RF, XGBoost has a more complex structure and introduces regularization items in loss function to control against overfitting so that it can better handle complex data. As a boosting method, XGBoost proceeds by iteratively adding new trees that predict the residuals or errors of prior trees, then all trees are combined to make the final prediction.

In this study, the XGBoost is applied on a 2-year dataset (September 2019-February 2022) with a 5-min temporal resolution, where 90% for training and 10% for validation and test. The predictors in the model are then the LWD and SWD fluxes and the additional meteorological parameters measured at the BSRN station (Table 1), and the predictand is the CF measurements from the UV-Indien network. Cross validation is performed for training. Learning curves based on RMSE are used to evaluate the model during the training and the validation.

see [Fan et al., 2020](#_ENREF_9) and [Liu et al., 2022](#_ENREF_14).

## Error metrics

To evaluate the performance of XGBoost, APCADA, root mean square error (RMSE), mean absolute error (MAE), and correlation coefficient (r) were calculated.

# Results

## APCADA performance

APCADA-estimated PCA (in octas) was first converted to CF (continuous values) and then compared with observations in Reunion from September 2019 to February 2021. APCADA was able to estimate dynamics of the observed CF, a correlation of 0.84 is achieved, with RMSE of 0.23 and MAE of 0.31.

We need to add some other results to increase analysis: (1) showing the level of estimating three classes: CF [0 - 0.25, 0.25 - 0.75, and 0.75 - 1] → to ‘better’ evaluate under/overestimation. (2) plot 1-day results.

here

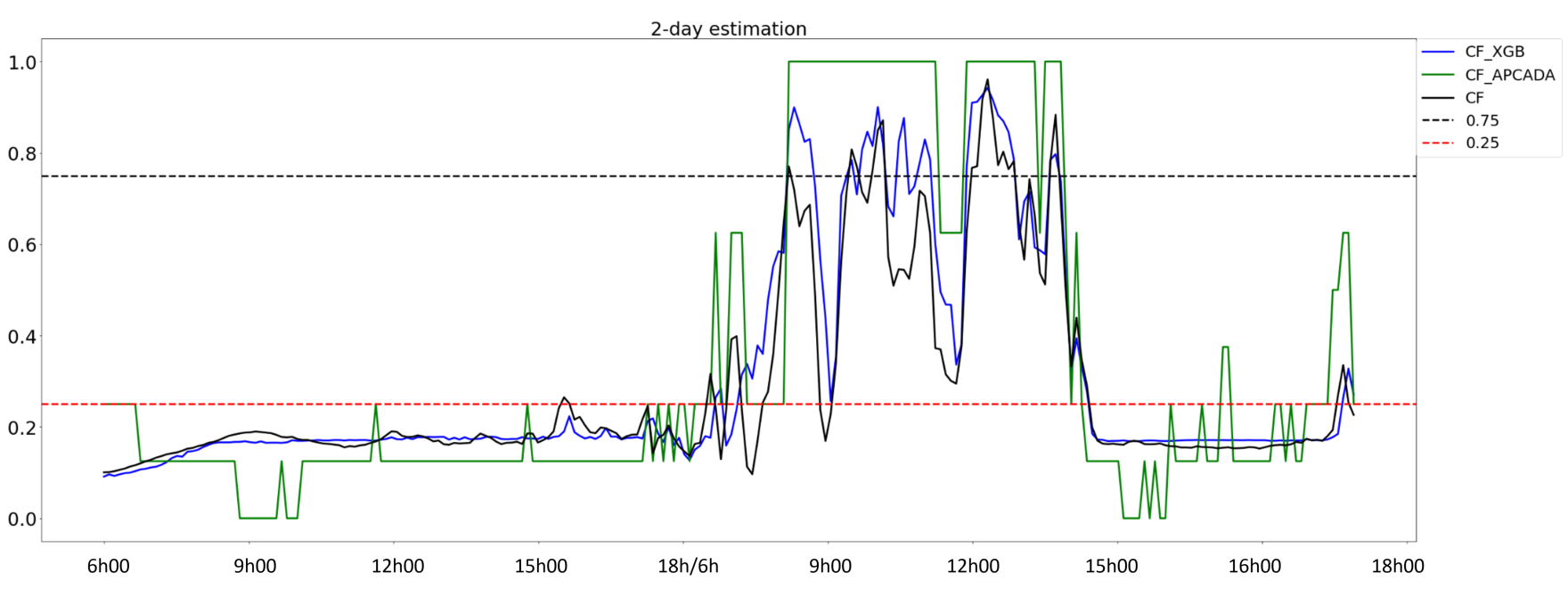


Table 2: Model performance in estimating CF different levels (low, medium, and high cloudy)

|  |  |  |  |
| --- | --- | --- | --- |
| Model\I | 0.00 - 0.25 | 0.25 - 0.75 | 0.75 - 1.00 |
| XGBoost | 0.08 | 0.13 | 0.14 |
| APCADA | 0.22 | 0.31 | 0.14 |

Comment: All values are greater than 80%, except for KWA, where the minimum score is about 70%.The lowest score rates are found for KWA ([Table 5](https://agupubs.onlinelibrary.wiley.com/doi/10.1029/2003JD004182#jgrd11026-tbl-0005)). Investigations showed that the percentage of broken PCA (2–6 octas) in KWA is over 60% compared to about 30% in PAY, which decreases the precision of APCADA. This is a small island next to Australia, [link](https://www.google.com/maps/place/Kwajalein+Atoll,+RMI/@8.716688,167.698314,13z/data=!4m5!3m4!1s0x65aeb9f09b73c9a5:0xacb7ba68a691d2a3!8m2!3d8.716667!4d167.733333?hl=en-GB)

## XGBoost performance

See turning of XGBoost in [Mo et al., 2019](#_ENREF_19).

### Variable selection

see [Huang et al., 2021](#_ENREF_11)

The variable selection step is important in constructing machine learning models. We used the random forest algorithm to select data variables (Zeng et al., 2020). Normalized daily data were used to construct and train the random forest model and to calculate the model’s importance. The data preprocessing experiment was intended to verify the importance of variables in a given model and to analyze the impact of changes in the variables on the model’s predictive performance. The experiment proceeded as follows:

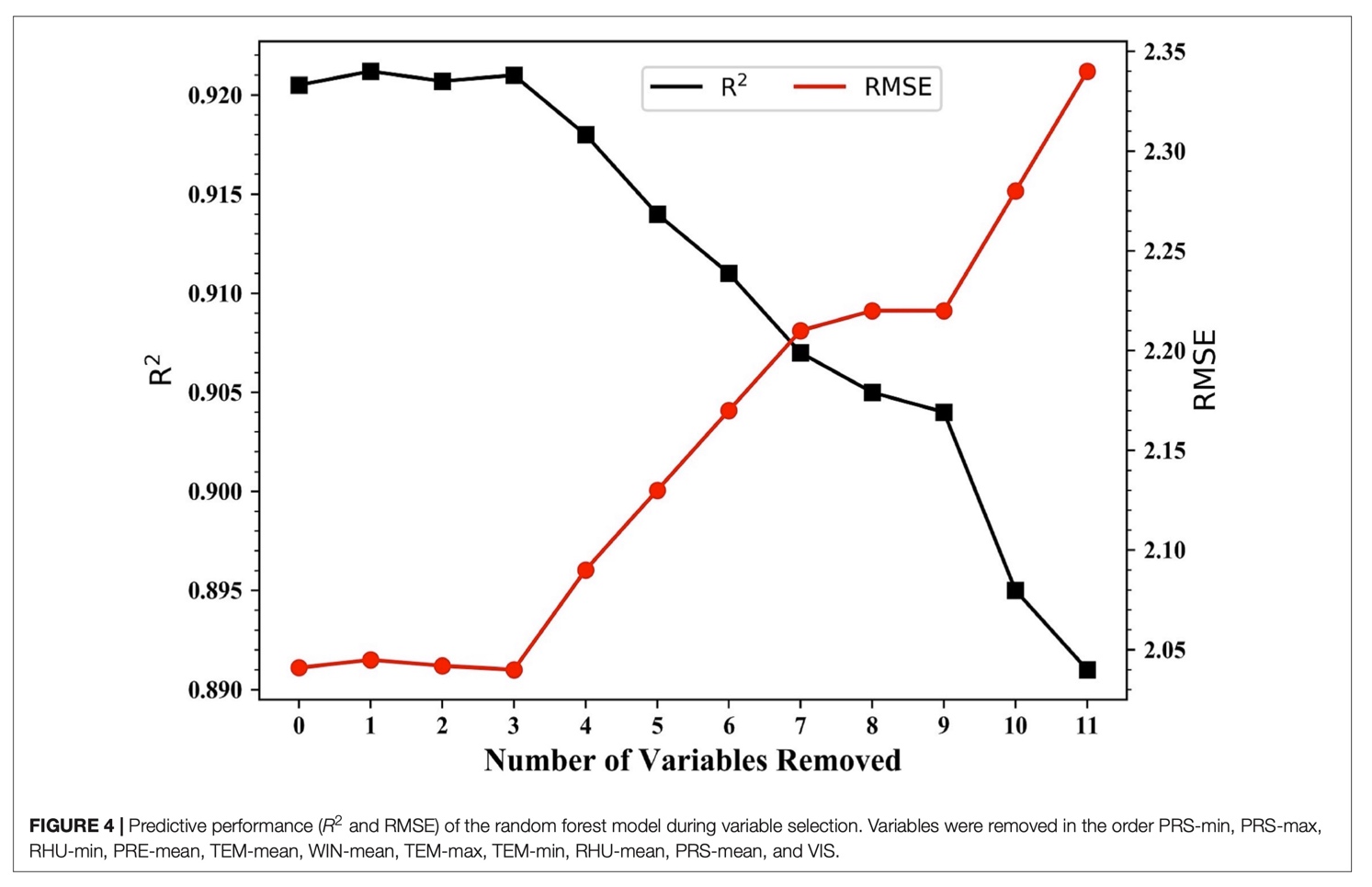
(1) divide the dataset into a training set and test set after completing the data quality control process;

(2) use the training set to train and save the model, then calculate the correlation coefficient (*R*2) and the root mean square error (RMSE) of the saved model;

(3) based on the order of importance of the variables in the model, eliminate the least important variable;

(4) repeat steps(2) and (3) until only two variables remain (the minimum required for calculation).

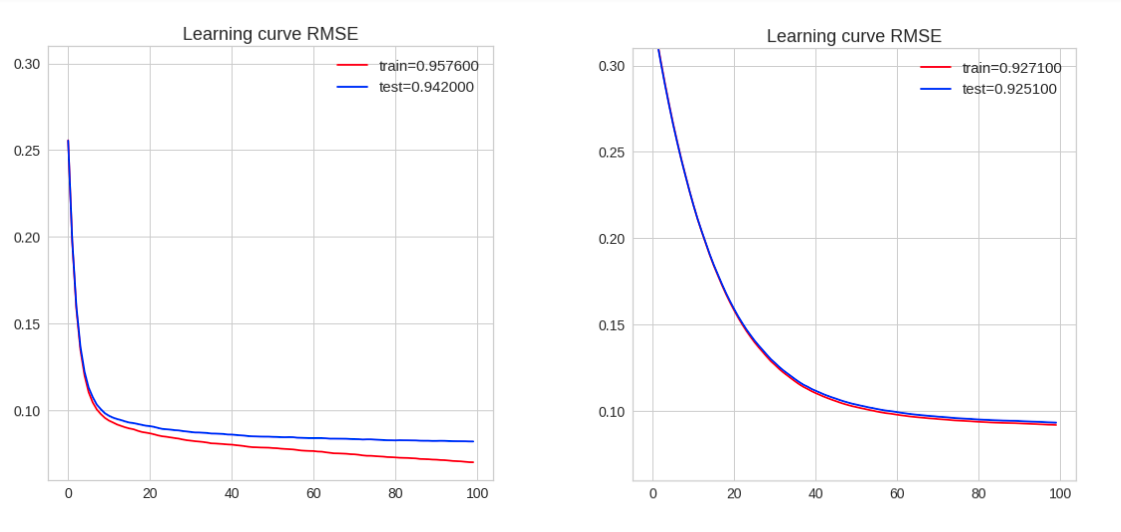
**Figure 4** shows that when the model contained <10 variables, *R*2 tended to decrease and the RMSE tended to increase. Between 12 and 10 variables, *R*2 reached 0.921 and the RMSE was 2.042 MJ/m2. With four variables, *R*2 decreased sharply from 0.904 to 0.895 and the RMSE decreased from 2.19 to 2.28 MJ/m2. Therefore, the prediction of solar radiation can achieve the best performance when using 10 variables, then the subsequent model experiments were trained with these 10 variables.



The first model fitting with the XGBoost default parameter values returns a model with an overfitting issue. To overcome this problem, several tests were performed with different optimization methods such as Principal Component Analysis (PCA), Recursive Feature Elimination (RFE) and feature importance. Each applied method was followed by tuning of hyperparameters. Results after PCA are not satisfactory as the model performance degrades. In contrast, in the case of RFE, results in three of the tests performed show an increase in the model’s performance. Finally, feature importance gives results similar to the best results with RFE. The figure (\ref3)) shows learning curves before and after tuning hyperparameters. And the table (\ref2)) shows optimal hyperparameters.

Chao: I suppose in the table you will show the final optimized parameters. while the optimization is not ONLY to handle overfitting. The final purpose of ML is generalization, there should be other considerations when turning the hyperparameters. So better to make people think the turning is not only for overfitting.

After feature selections and model optimizations, a final score of 92.02% is achieved, which is comparable to the results from recent studies ([Dürr and Philipona, 2004](#_ENREF_8)) do we have other similar studies). Figure (\ref5)) shows CF estimated time series with observations, with RMSE and MAE of 0.12 and 0.08 respectively (Table(\ref2))). We find that the surface pressure is not necessary for this application since it makes the XGBoost model more complex, Figure(\ref4)).



(a) (b)

Fig 3: Learning curve (RMSE) before (a) and (b) after tuning hyper-parameters of XGBoost model.

Table 2: XGBoost model optimized hyperparameters

|  |  |
| --- | --- |
| Hyper-parameter | Values |
| N estimators | 100 |
| Learning rate | 0.05 |
| Max depth | 6 |
| Subsample | 1 |
| Colsample bylevel | 0.8 |
| Colsample bytree | 1.0 |
| Colsample bynodes | 0.8 |
| Min child weight | 4 |

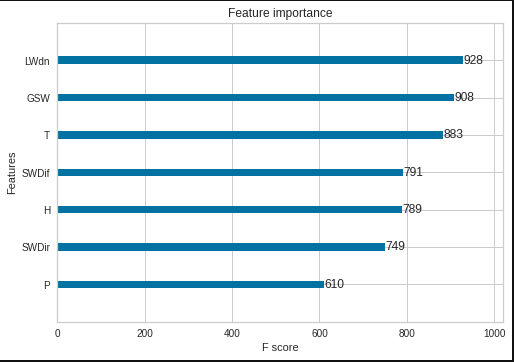


Fig 4: F-score of XGBoost model features to CF estimation.

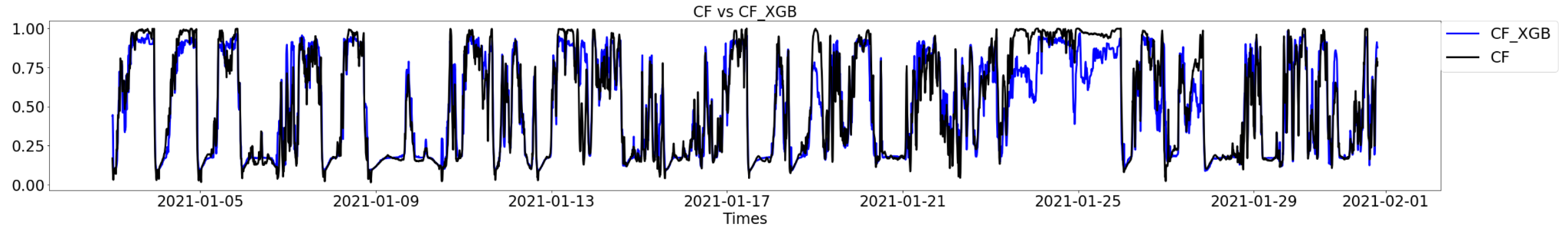


Fig 5: XGBoost CF estimation and observations at Reunion for January 2021 - February 2021 period. (plot modification)

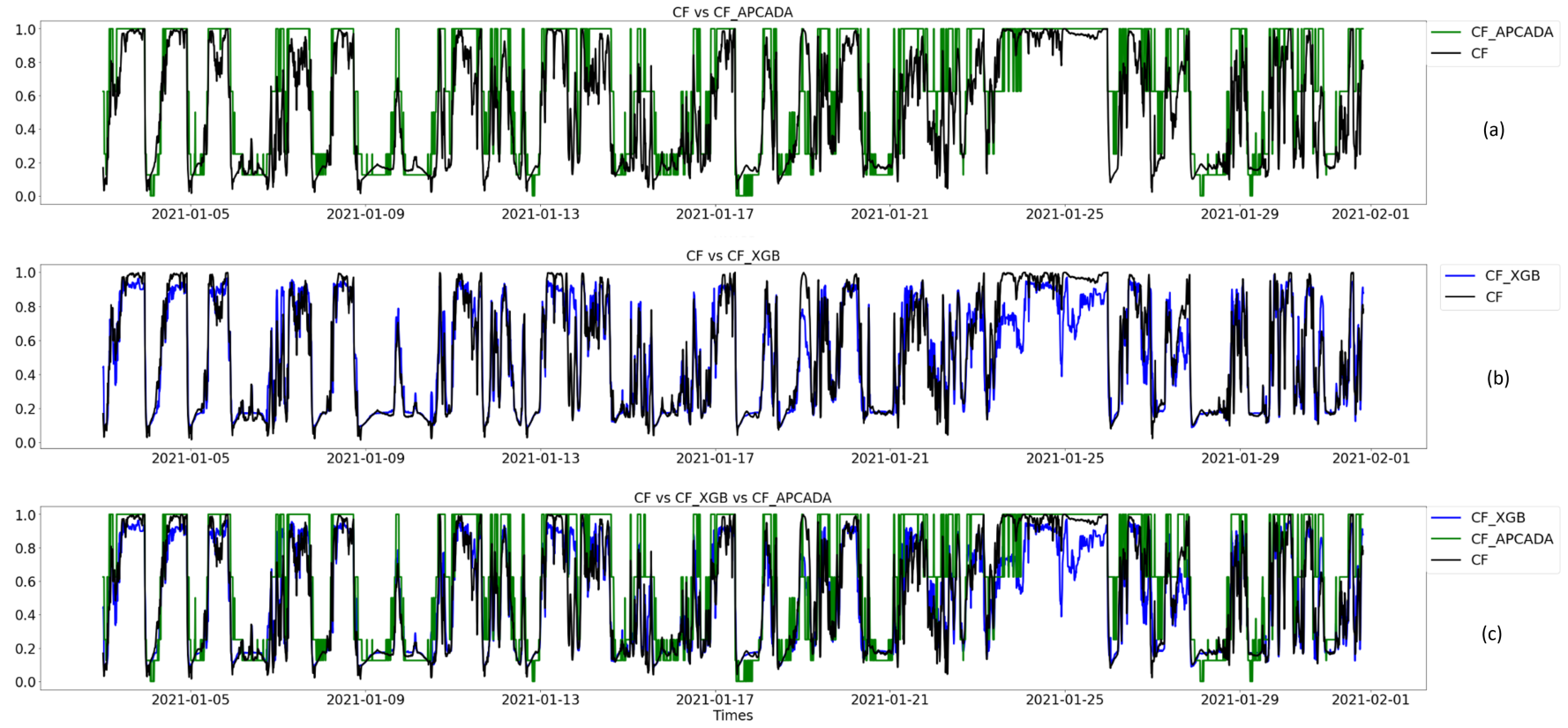


Fig 6: Comparison of observed and estimated CF, (a) observation, (b) XGBoost, and (c) APCADA together with CFI.

## CF estimations

Statistical and physical methods were compared, Table (\ref2), and \ref3)). The correlation of 85.9% was observed between XGBoost and APCADA estimations. In addition, RMSE and MAE of 0.27 and 0.21 were observed between the two methods.

Chao: is the r Pearson correlation? if yes, 1) usually, the correlation coefficient varies between +1 and -1. 2) The Pearson correlation is less appreciated if the two variables (CF\_APCADA, converted from octas for example) are not gaussian, and need to be checked. 3) it’s better that we keep the consistency of the precision of the matrix (r, RMSE, etc) 1 -> 1.00, 0.931 -> 0.93, that’s important but not easy to explain by text here.

Table 3: Correlation matrix of the estimated CF by XGBoost, and APCADA with observed CF.

|  |  |  |  |
| --- | --- | --- | --- |
| r | CF | CF\_XGB | CF\_APCADA |
| CF | 1.00 | 0.93 | 0.84 |
| CF\_XGB | 0.93 | 1. | 0.89 |
| CF\_APCADA | 0.84 | 0.89 | 1. |

Table 4: Error metrics between CF estimations and observation

|  |  |  |
| --- | --- | --- |
| Estimations | RMSE | MAE |
| CF vs CF\_XGB | 0.12 | 0.08 |
| CF vs CF\_APACADA | 0.23 | 0.16 |
| CF\_XGB vs CF\_APCADA | 0.21 | 0.16 |

Discussion and perspectives

This study shows the possibility to estimate cloud fraction by a machine learning method from meteorological variables, where the longwave radiation plays a crucial role because of its strong relationship with clouds. The XGBoost model is successfully applied to estimate CF by LW radiation (along with other meteorological variables). It could be better to highlight the role of LW for the XGBoost model, we could add this later. The performance of XGBoost is proven superior to the traditional physical empirical method (e.g., APCADA). While it is worth noting that, XGBoost estimates CF in continuous values, whereas APCADA produces discrete data, which may potentially impact the statistical comparison of these two methods. High-quality and abundant CF measurements is the key to apply a Machine Learning approach (e.g., XGBoost), a future application could benefit from the increasing availability of CF data.

Only daytime CF, from the visible camera, is used to train the XGBoost model in this study. Similar experiments could be done during the night with nighttime CF measurement such as that from an infrared camera. With performed accuracy compared to traditional physical methods, the XGBoost model could be applied to extend the temporal coverage of existing CF observations from about 1.5 year in this study to more than 2.5 years with available meteorological variables.

## Limitations of APCADA at Reunion

Presence of dynamic low-level clouds in ocean regions mainly from the near surface moisture hindered the performance of APCADA. For the same reason, a correlation of below 70% was observed at Kwajalein (8.43°N, 167.44°E) (durr2004automatic)). other information on this?

* Underestimation of lower cloud fraction
* few clear sky days over the region → dynamics of cloud formation over oceanic region
* number of clear sky days tends to decrease from morning to evening
* data dependence estimation

## Limitation of XGBoost model

Chao, I don’t have an idea on this part

* underestimating high values, [link](https://www.mdpi.com/2072-4292/13/9/1848)
* Overestimation?
* huge amount of data dependence

# Conclusion

nd its representation in Numerical Weather Prediction (NWP) models still has a long way to go (heidinger1996finite)). The current uncertainties lie from misrepresentation of brokenness (inhomogeneity) of clouds in atmospheric radiative transfer calculations, which has been studied over 20 years for solar radiation in general and longwave radiation in particular (aida1977reflection), and ellingson1982effects)).

In addition to emissions, cloud brokenness effects to longwave radiation consider the side face shading of the broken cloud blocks. (more information on CF in NWP \*11 May\*)

Even though clouds play a key role in the Earth's radiation budget, there is still a gap in cloud fraction (CF) estimation for renewable energy (RE), photovoltaic (PV) production in particular. Different studies showed a strong radiative cooling effect from low level clouds, and negligible effect from the high clouds corresponding to the distance and emittance temperature ([Dürr and Philipona, 2004](#_ENREF_8), [Park and Shin, 2019](#_ENREF_21)). Presence of these low-level clouds, with abundant water droplets, is high in ocean regions mainly from near surface moisture. In addition to cloud vertical structure, and cloud optical depth, CF as a fraction of cloud occupied area has been used as a measure of cloud quantification and in the Earth’s energy budget ([Qian et al., 2012](#_ENREF_22)), and [Zhang et al., 2011](#_ENREF_29)). With the rise of CF measurement, different CF estimation methodologies have been developed.

As CF ground-based measurement is getting to maturity stage, radiation transfer uncertainties in general circulation models (GCMs), and physical models have started being addressed. Accuracy of long-term ground-based measurements have been hindering CF simulation and estimation with GCMs and physical models (qian2012evaluation)). In different regions of the world, use of both ground-based cameras and aircraft measurements produces 1-D cloud snapshots. For global observations, with remote sensing, global cloud properties and cloud cover parameters have been developed over years from visible and infrared radiations ([Qian et al., 2012](#_ENREF_22)). With increase of data, statistical studies have been conducted and found significant correlations of CF with precipitation, air temperature, and

With recent technology of using aerosol lidar and thermographic cameras, continuous valuable cloud datasets are being produced and welcoming machine learning methods.

Estimation of the cloud brokenness has been using different approximation methods, which requires a significant computational effort. Statistical methods - correlation analysis - have been used on a large amount whereas significant correlations have been found between cloud fraction and precipitation, air temperature, and topography (didier2015comparison), and forsythe2015detailed)). With the recent rise of machine learning and availability of data, different studies showed a large correlation between cloud fraction and longwave radiation while estimating the downward longwave (ref). (more information on relationship of CF with other variables -physically and statistically \*12 May\*)

have been used to evaluate the results, and comparison of the longwave data with Clouds and the Earth’s Radiant Energy System (CERES) data. And to evaluate the machine learning models, learning curves were used.

Before using the model, we need to check if the model work well. By using leaning curve we can evaluate the performance of the model. Learning curve show 2 curves, one is for training and the other for validation. Overall, during the training both curves are supposed to decrease together and the generalization gap between both curves is necessary low. Generalization gap is the difference between both curves. Three cases can be appears: over-fitting, under-fitting and good fitting. Over-fitting mean when the model is good during the training, but bad for the validation. Under-fitting mean the model doing great for the validation and bad for the training. Good fitting when the generalization gap is lower and both curves are nearly overlapping. The tuning hyper-parameters are necessary to solve over-fitting or under-fitting issues.

The next figures show the learning curves of the model before and after tuning hyper-parameteres:

Feature importance is an usefull method used to enhence a model machine learning, it allows to select the most important feature to the model. The feature importance result is presented to the following figure :

Taking into account the result of feture iumportance, several test have been carried out to improve the performance of the model. The results are shown in the following figures :

Result about CF estimation with XGBoost

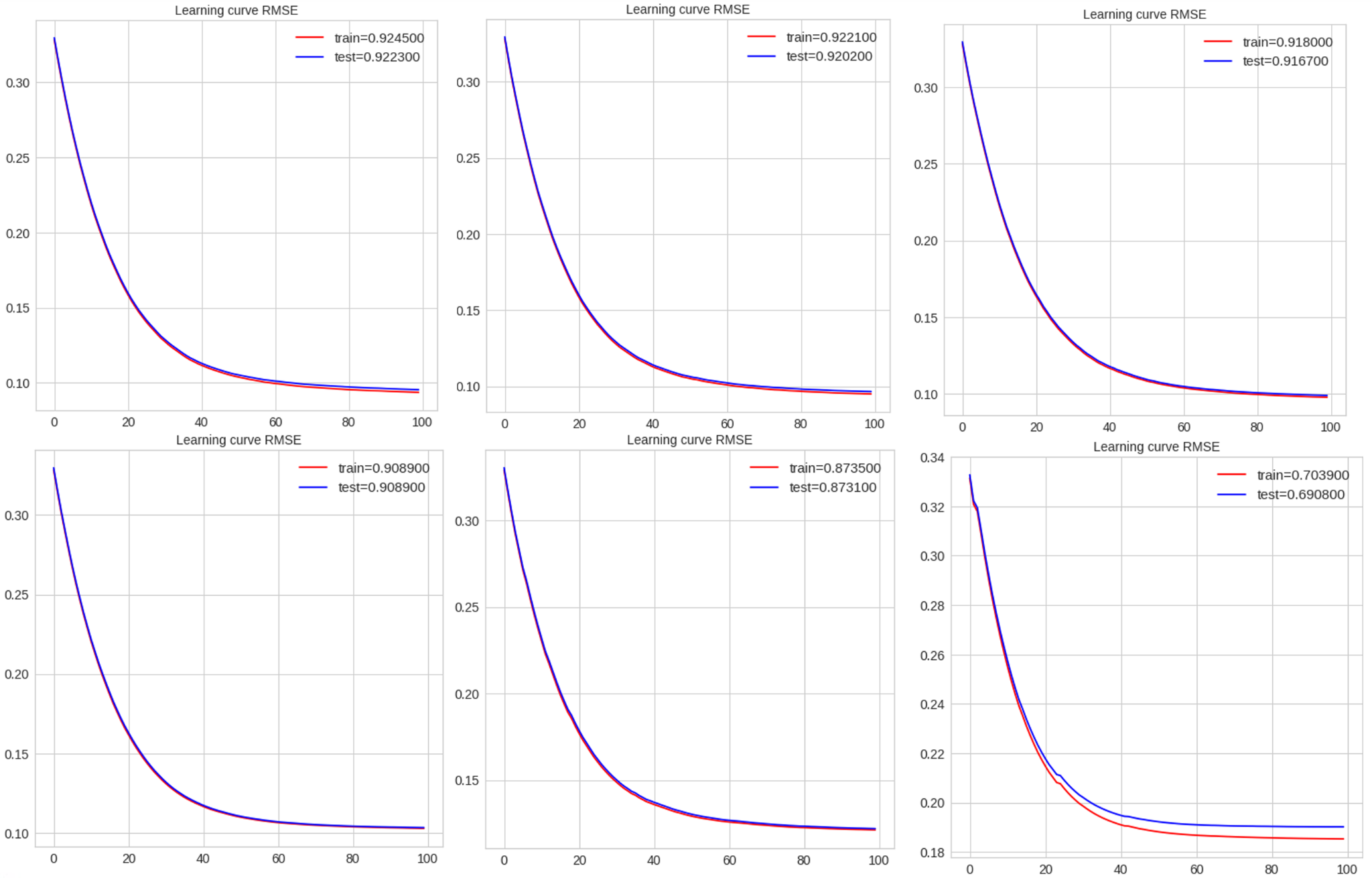
An estimation of the cloud fraction was made after selecting the best characteristics for the model and the following figure is a sample of the result:

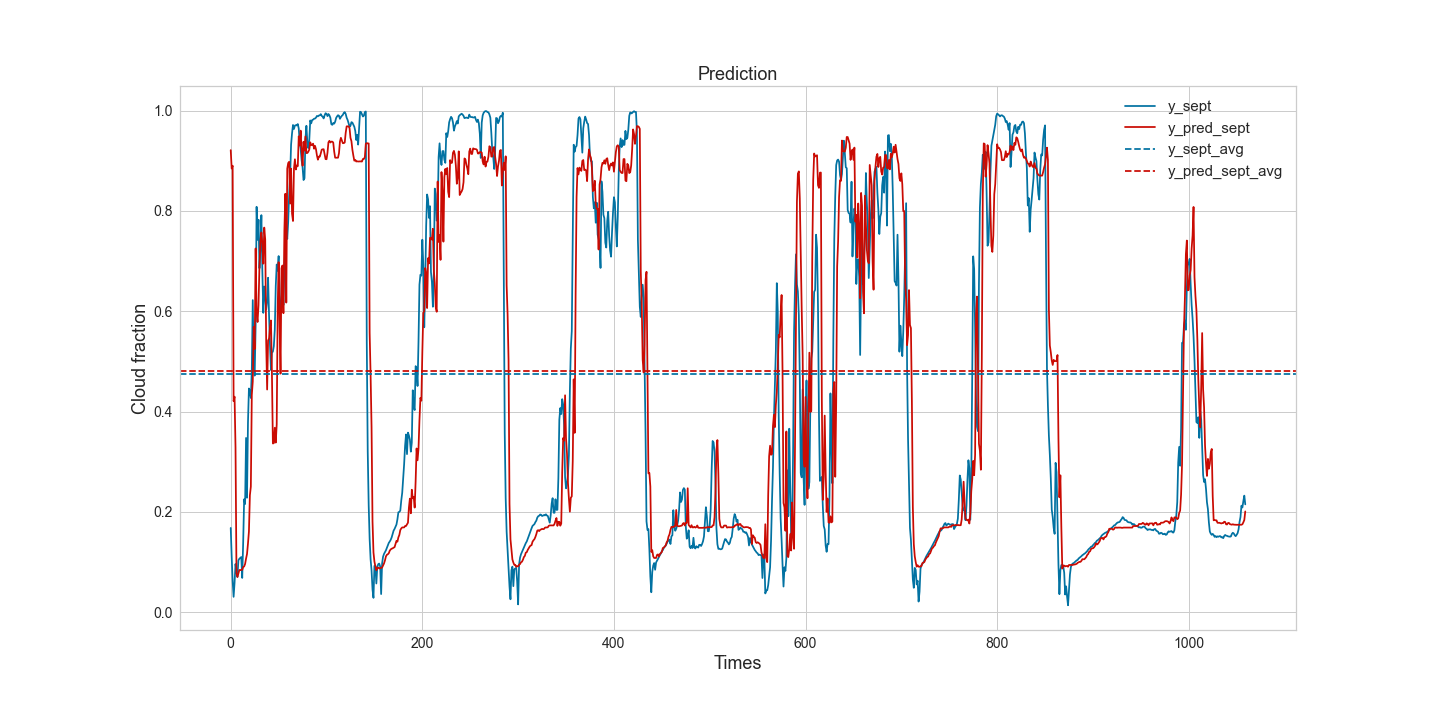
APCADA results :

Physical model called APCADA was used as a reference to assess the results obtained with the model XGBoost. In the following figure, CF\_octat is the real observation, CF\_XGB\_octat represent the result from the model XGBoost, PCA is the partial cloud amount in octat from APCADA, PCA/8 represente the result from the model physique, CFI cloud fraction index, CF\_XGB represent the cloud fraction fom XGBoost model and CF is the real observation.

|  |  |  |  |
| --- | --- | --- | --- |
| r | RMSE | MAE | Corr |
| CF\_octat vs PCA | 2.47 | 1.77 | 0.77 |
| CF\_octat vs CF\_XGB\_octat | 1.06 | 0.66 | 0.92 |
| PCA vs CF\_XGB\_octat | 2.13 | 1.65 | 0.84 |

|  |  |  |  |
| --- | --- | --- | --- |
| Model XGBoost | | | |
| Data source | Input | Output | Period |
| BSRN and UV-Indien network | GSW, SWDif, SWDir, LWD, T, RH, P | CF\_XGB, CF | 2019-09-01 to 2021-03-01 |
| Model APCADA | | | |
| Data source | Input | Output | Period |
| BSRN Reunion, LACy | LWD, T, RH | CF\_APCADA, CF | 2019-11-01 to 2021-05-31 |





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