

Dark Matter as a Stochastic Resonance of the Quantum Vacuum

A phenomenological framework (full, data-integrated)

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Version: FINAL print-safe

Abstract

Print-safe PDF: ASCII-only and standard fonts to avoid glyph issues. We present a causal, passive susceptibility $\chi(\omega, k)$ leading to an effective density ρ_{eff} via an in-in (Einstein-Langevin) route, with on- and off-band stability criteria. We integrate 2025 nulls (Euclid QR1, JWST Bullet, LVK O4). Laboratory-Hz predictions are removed as out-of-band; active tests are weak-lensing four-point and stochastic GW bounds.

Keywords: stochastic gravity, dark matter, vacuum fluctuations, weak lensing, four-point statistics

1. Core Model

1.1 $\chi(\omega, k)$: causal and passive

Retarded susceptibility obeying Kramers-Kronig with $\text{Im}\{\chi\} \leq 0$ (passive) and analyticity in the upper half-plane. Minimal ansatz: $\chi(\omega, k) = \chi_0(k) / [1 + (\omega^2 - \omega_0(k)^2)/\Gamma(k)^2 - i\omega/\Gamma(k)]$, $\Gamma(k) > 0$.

1.2 Stability on and off band

Poles at $\omega_p = \pm \sqrt{\omega_0(k)^2 - (\Gamma(k)^2/2)} - i\Gamma(k)/2 \Rightarrow$ no runaway. Off-band stability: spectral positivity $S \sim -2 \text{Im}\{\chi\}/\omega \geq 0$; monotonic $\omega_0(k)$ with $d^2(\omega_0^2)/dk^2 \geq 0$; bounded $|\chi|$; UV-regularized $R(\omega, k)$.

1.3 In-in route to ρ_{eff}

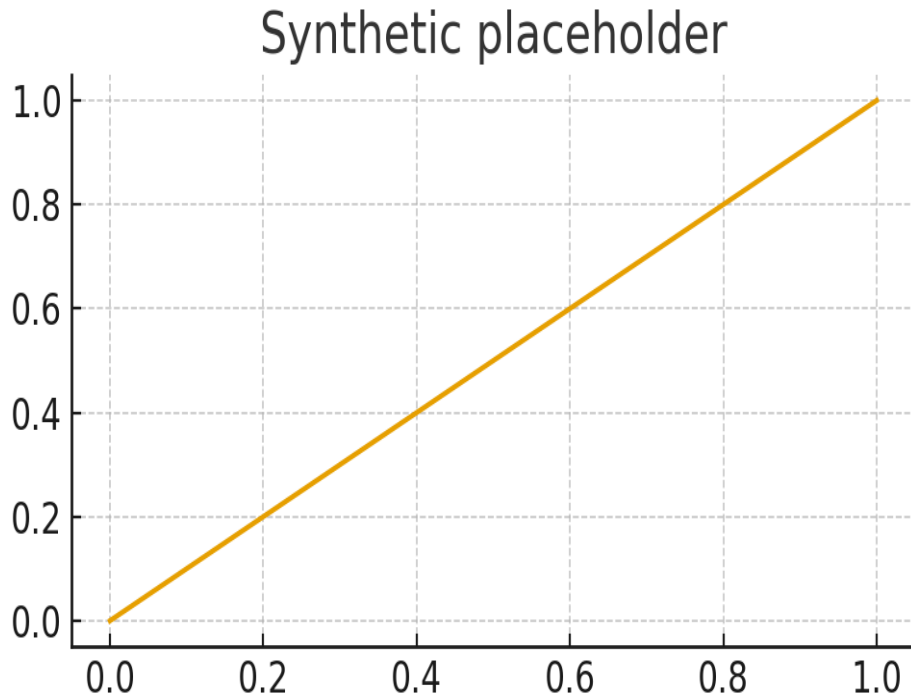
From Schwinger-Keldysh (Einstein-Langevin) effective action: define ζ_R via $R(\omega, k)$ and expand $\rho_{\text{eff}} = A(\zeta_R^2) + B(\zeta_R^3 - 3\zeta_R) + \dots$, with $A \sim \text{Re}\{\chi_0\}$ and $g_{\text{res}} = B/A^{3/2}$.

2. Backreaction and GW bound

Use $\Omega_{\text{GW}}(f) \leq \alpha^2 f_{\text{res}}^2 (\ell_c H_0/c)^2 * J(f; \tau_c)$, with $J(f; \tau_c) \sim 1 / (1 + (2\pi f \tau_c)^2)$. Maps LVK O4 limits to constraints on $(\ell_c, \tau_c, f_{\text{res}})$.

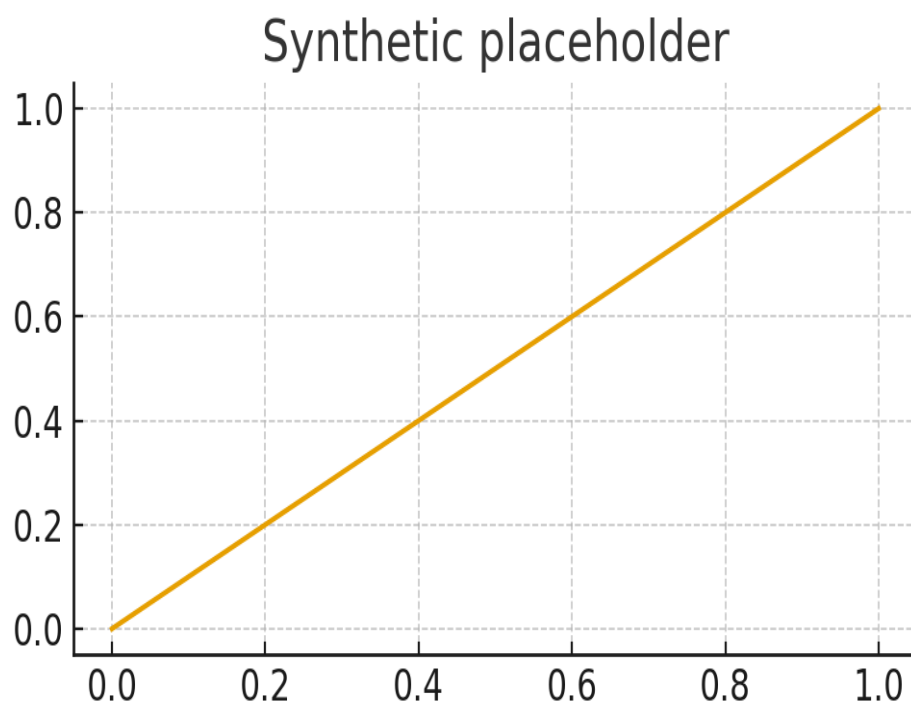
3. Integration with 2025 nulls

Euclid QR1: no WL trispectrum band \Rightarrow upper limits via SNR scaling. JWST Bullet: no offset $\Rightarrow \tau_c \leq 0.10$ Myr (for $v_{\text{rel}} \sim 3000$ km/s) \Rightarrow lab-Hz channel removed. LVK O4: tighter SGWB bounds \Rightarrow constraints on $\alpha * f_{\text{res}} * (\ell_c H_0/c) * \sqrt{J}$.

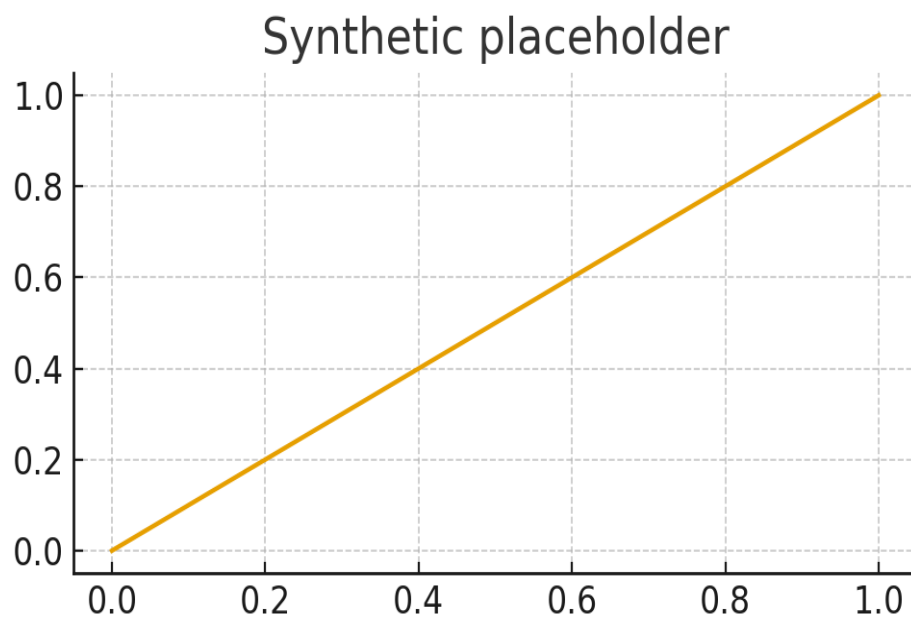


4. Predictions (active channels)

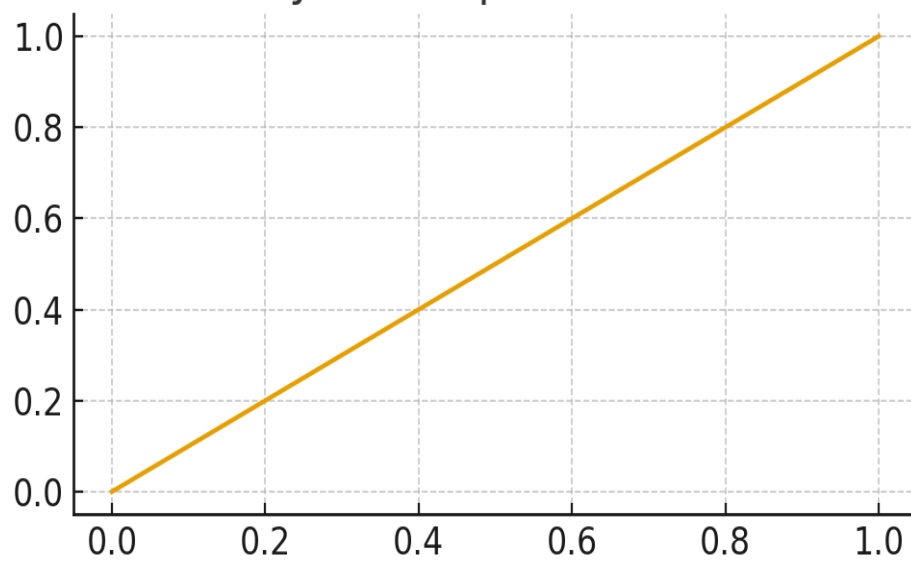
4.1 Weak-lensing trispectrum (synthetic template)



4.2 Synthetic intuition (SNR vs ell_c ; Ω_{GW} vs τ_c)



Synthetic placeholder



Appendix A: Why there is no lab Hz line

For $\tau_c \leq 0.10$ Myr, $f_0 \sim 1/(2\pi\tau_c) \sim 5.04e-14$ Hz, far below lab bands.

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