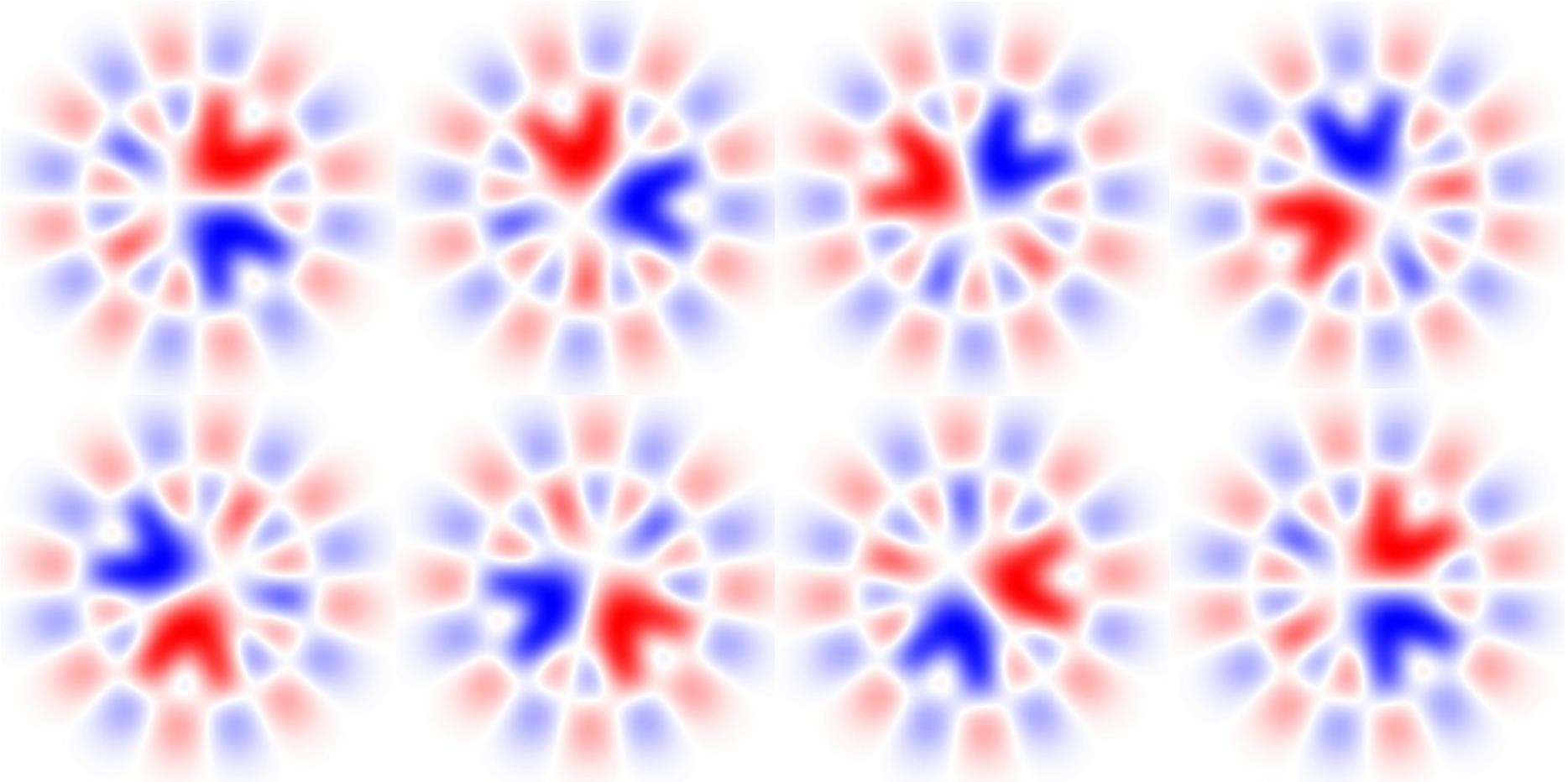




# Galaxy Image Modeling Using Shapelets and Sparse Techniques

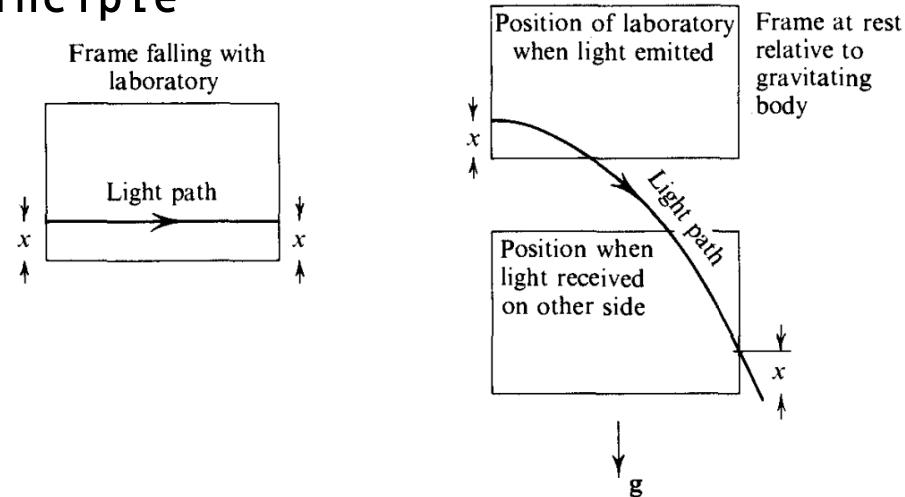


Student:  
Andrija Kostić

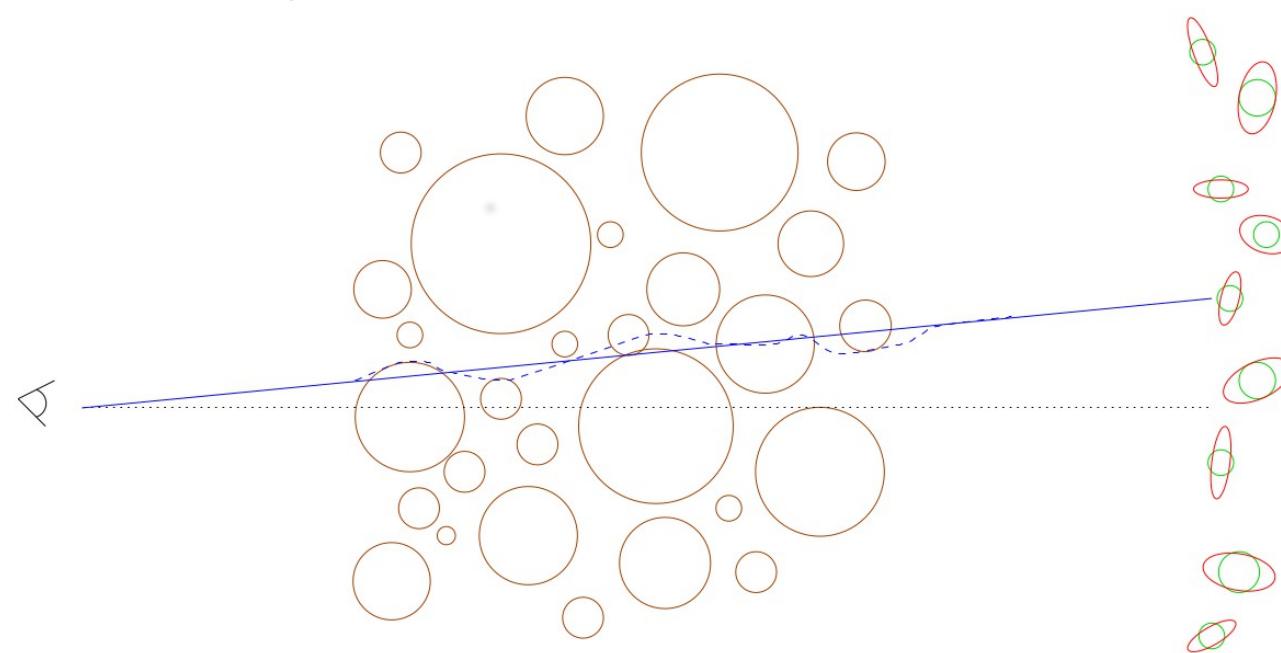
Supervisor:  
Arun Kannawadi

# Motivation

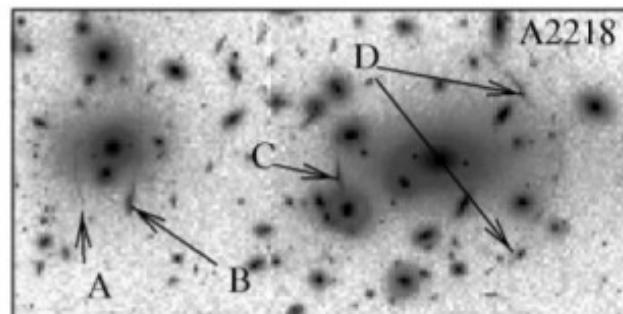
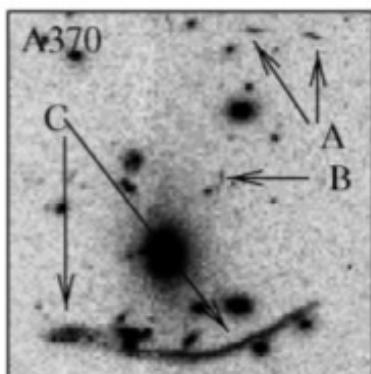
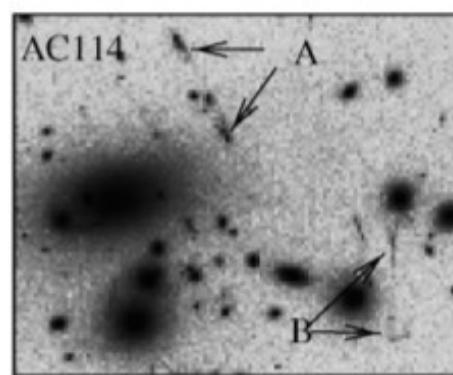
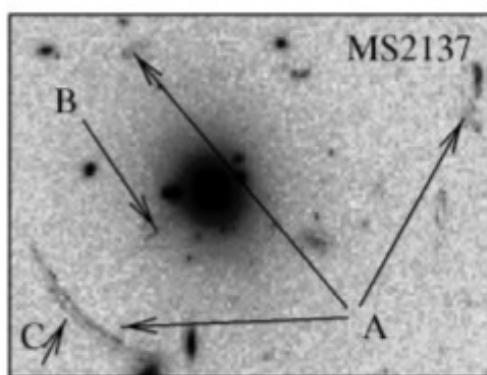
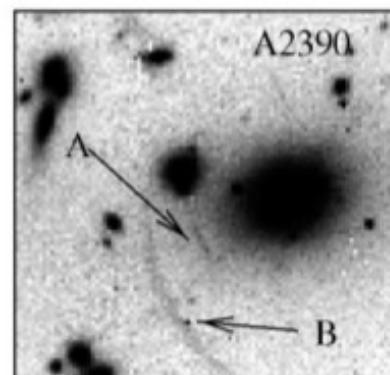
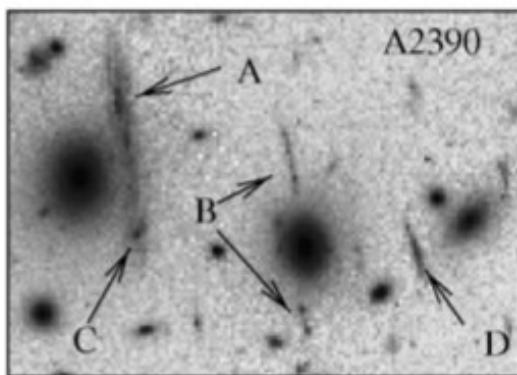
## Equivalence principle



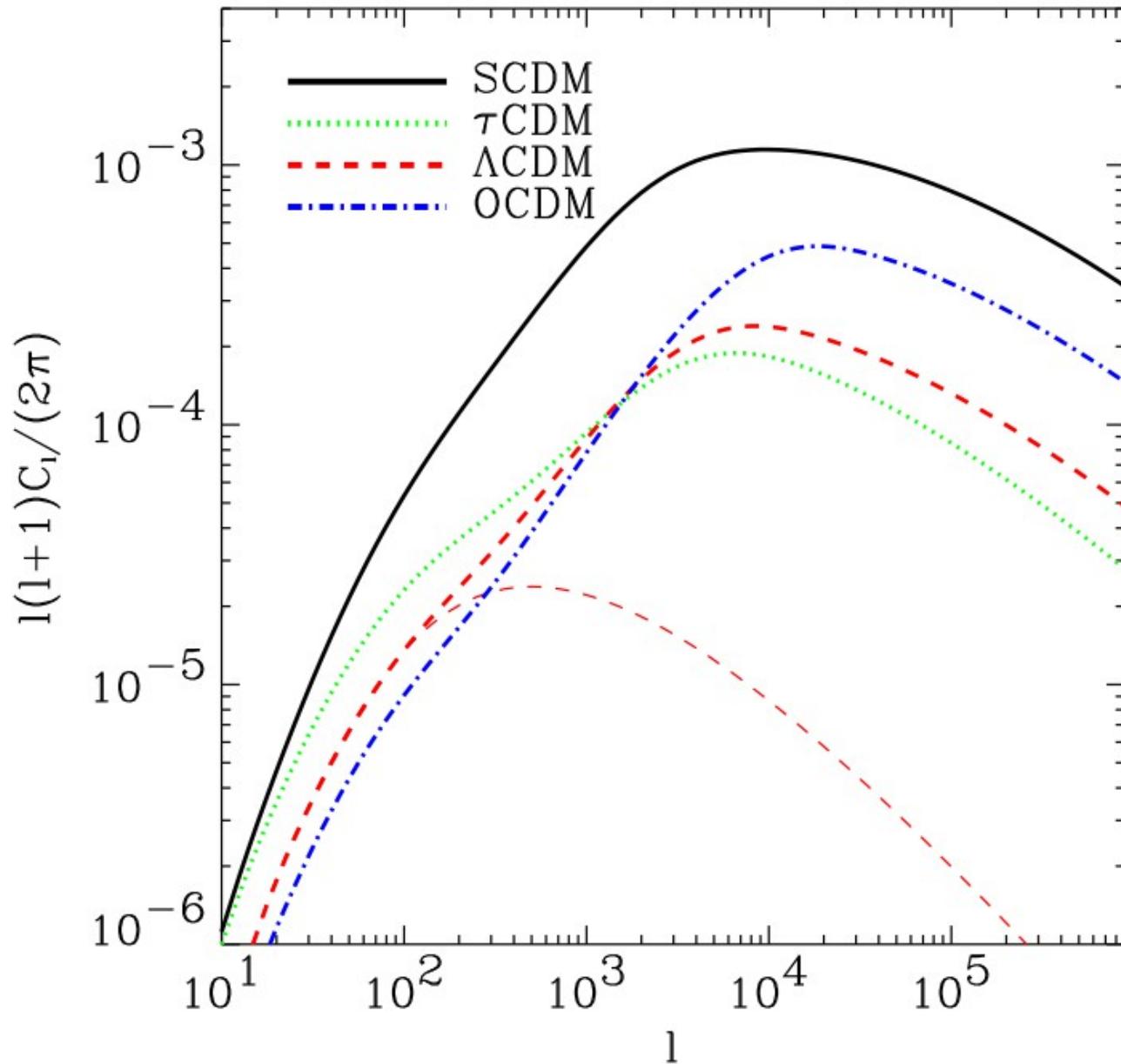
## Applied to cosmological models



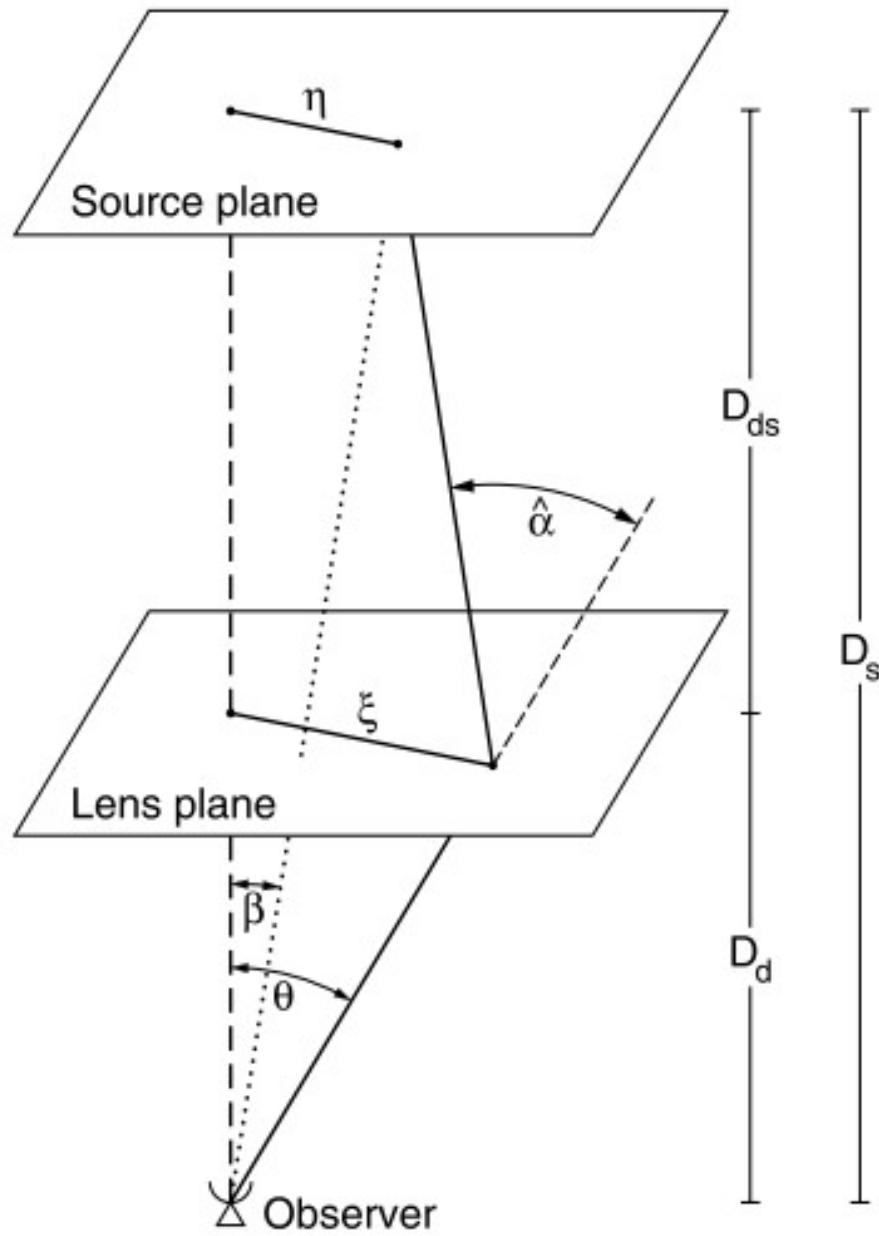
# Motivation



# Motivation



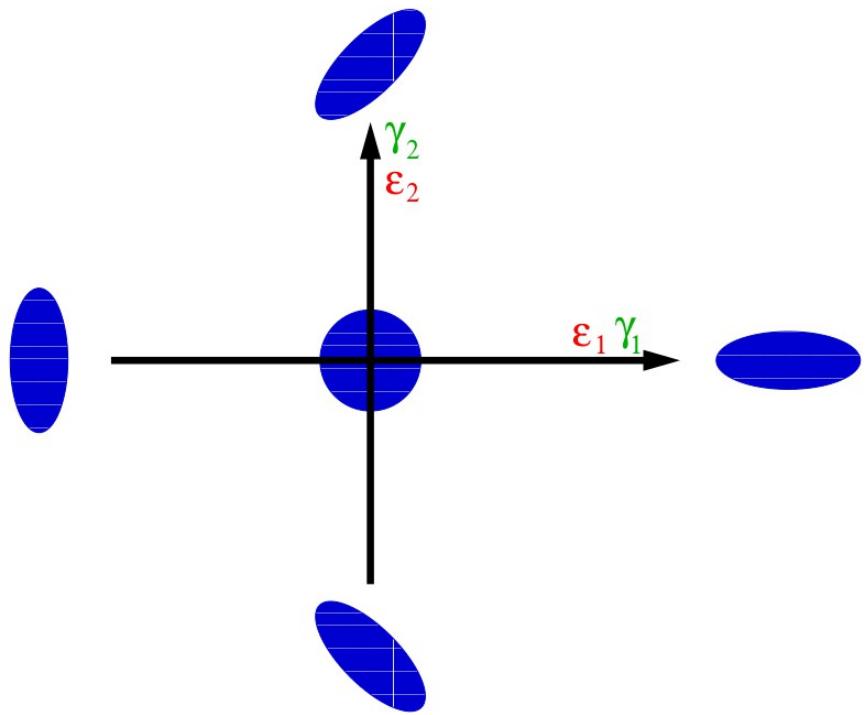
# More precisely



Surface brightness  
 $I(\theta) = I^{(s)}[\beta(\theta)]$

Jacobian of transformation

$$\mathcal{A}(\theta) = \frac{\partial \beta}{\partial \theta} = \left( \delta_{ij} - \frac{\partial^2 \psi(\theta)}{\partial \theta_i \partial \theta_j} \right) = \begin{pmatrix} 1 - \kappa - \gamma_1 & -\gamma_2 \\ -\gamma_2 & 1 - \kappa + \gamma_1 \end{pmatrix}$$

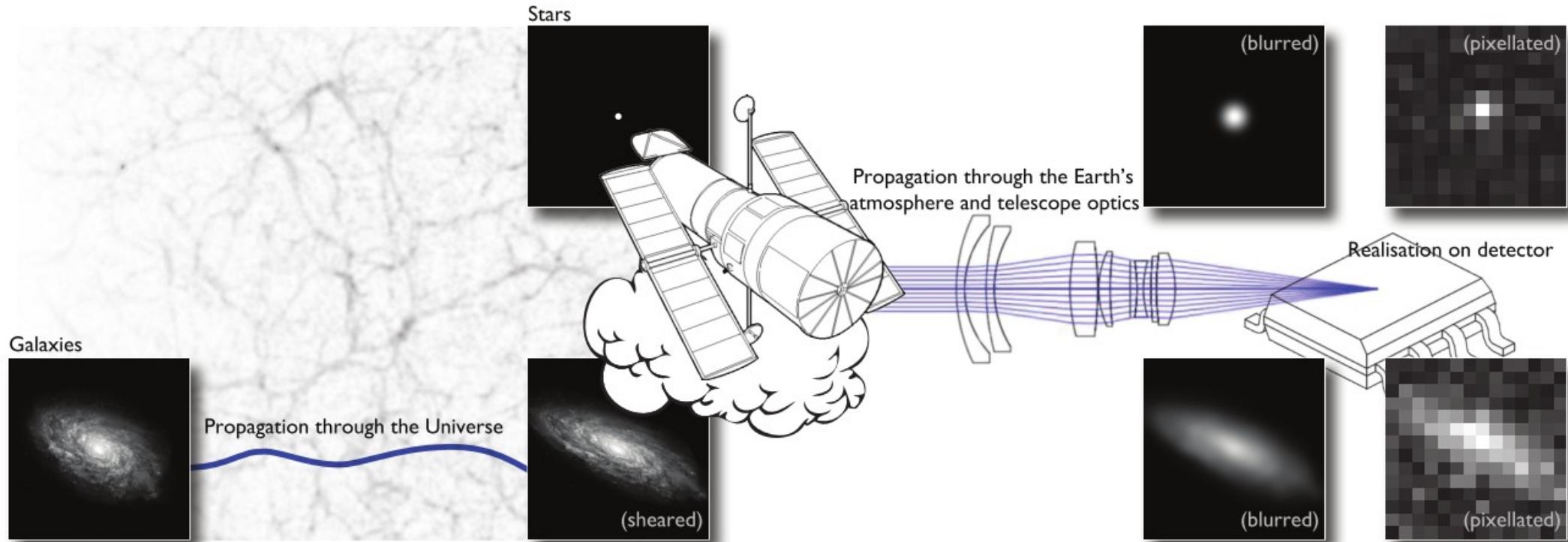


A2218 cluster at  $z = 0.175$  - HST



It's the ellipticity averaged over the ensemble of galaxies that is important for the tidal gravitational field

# Weak lensing - difficulties



- Signal is very weak, hard to distinguish from the observational distortions
- Ellipticities are nonlinear functions of the measured brightness, and hence biased

**Solution is doing simulations!**

Real galaxy

Model



Shapelets are very good at capturing irregularly shaped galaxies

Real galaxy

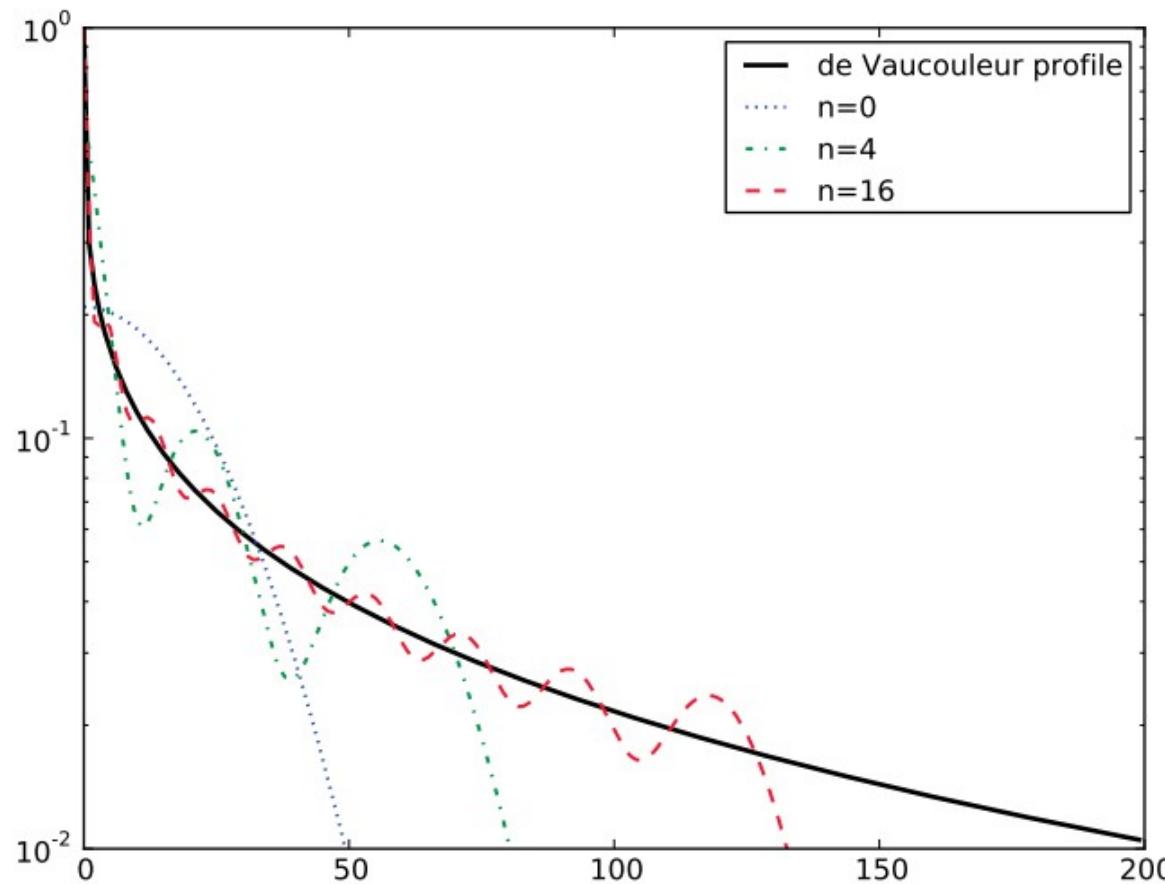
Model



Real galaxy

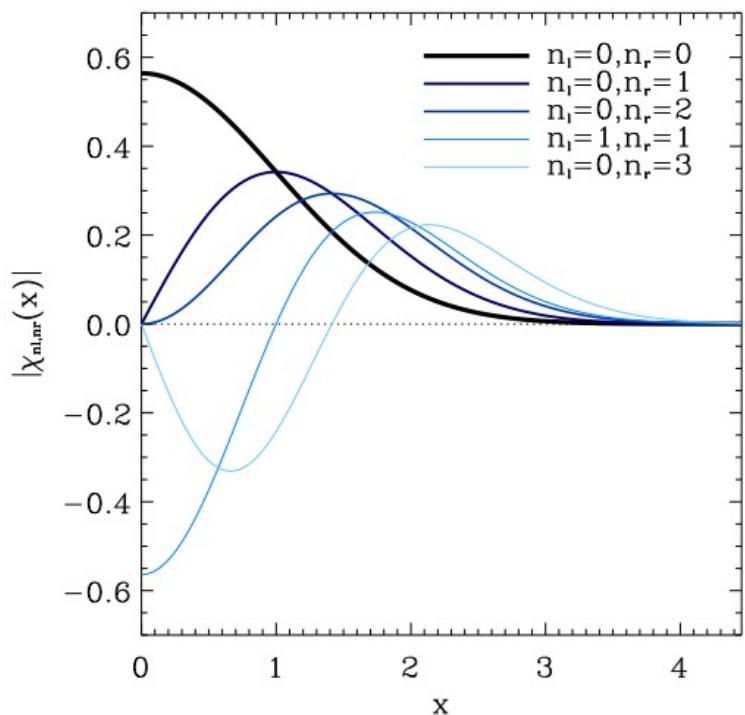


Usually **Sersic** profiles are used for modeling



# Shapelets - theory

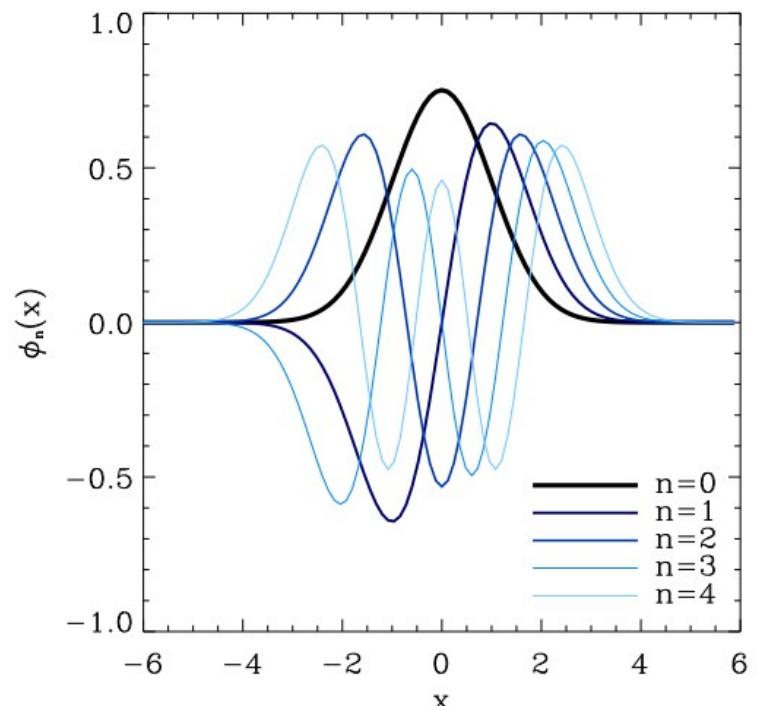
**Polar**



$$\chi_{n_l, n_r}(x, \varphi) = [\pi n_l! n_r!]^{-\frac{1}{2}} H_{n_l, n_r}(x) e^{-x^2/2} e^{i(n_r - n_l)\varphi}$$

$$A_{n_l, n_r}(x, \varphi; \beta) = \beta^{-1} \chi_{n_l, n_r}(\beta^{-1} x, \varphi)$$

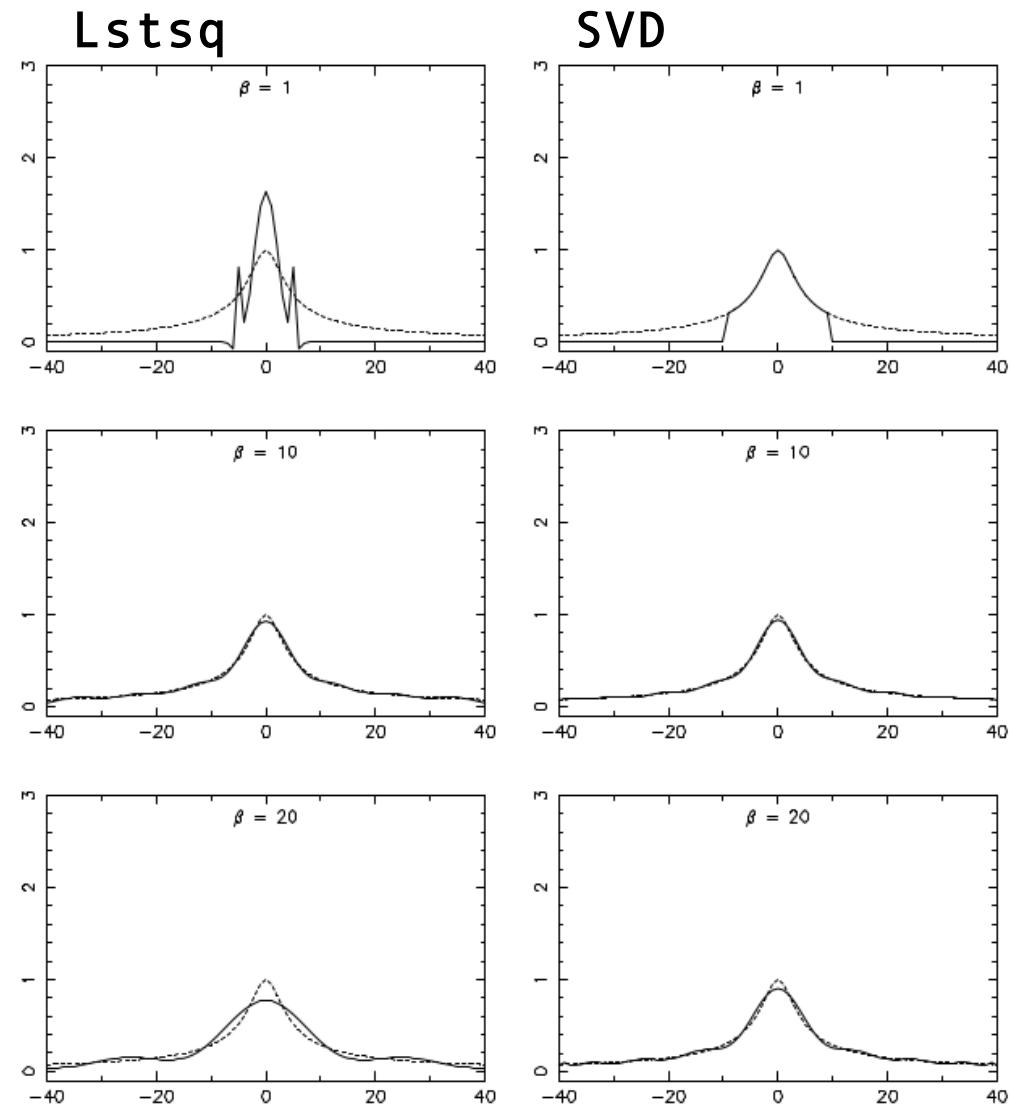
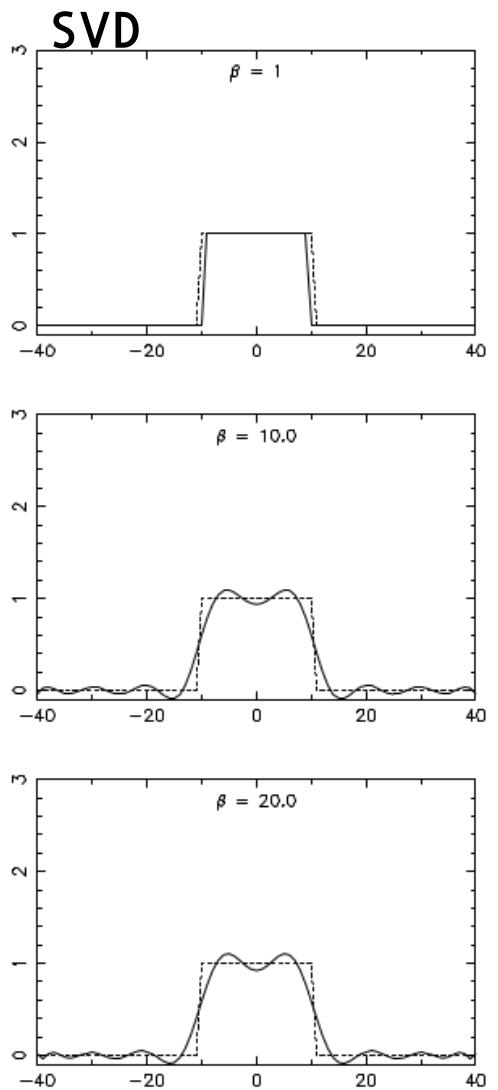
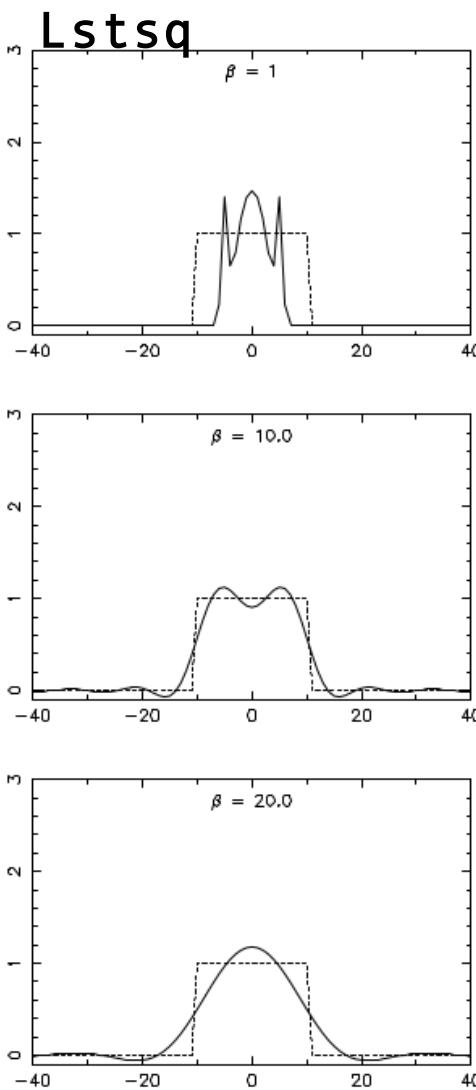
**Cartesian**

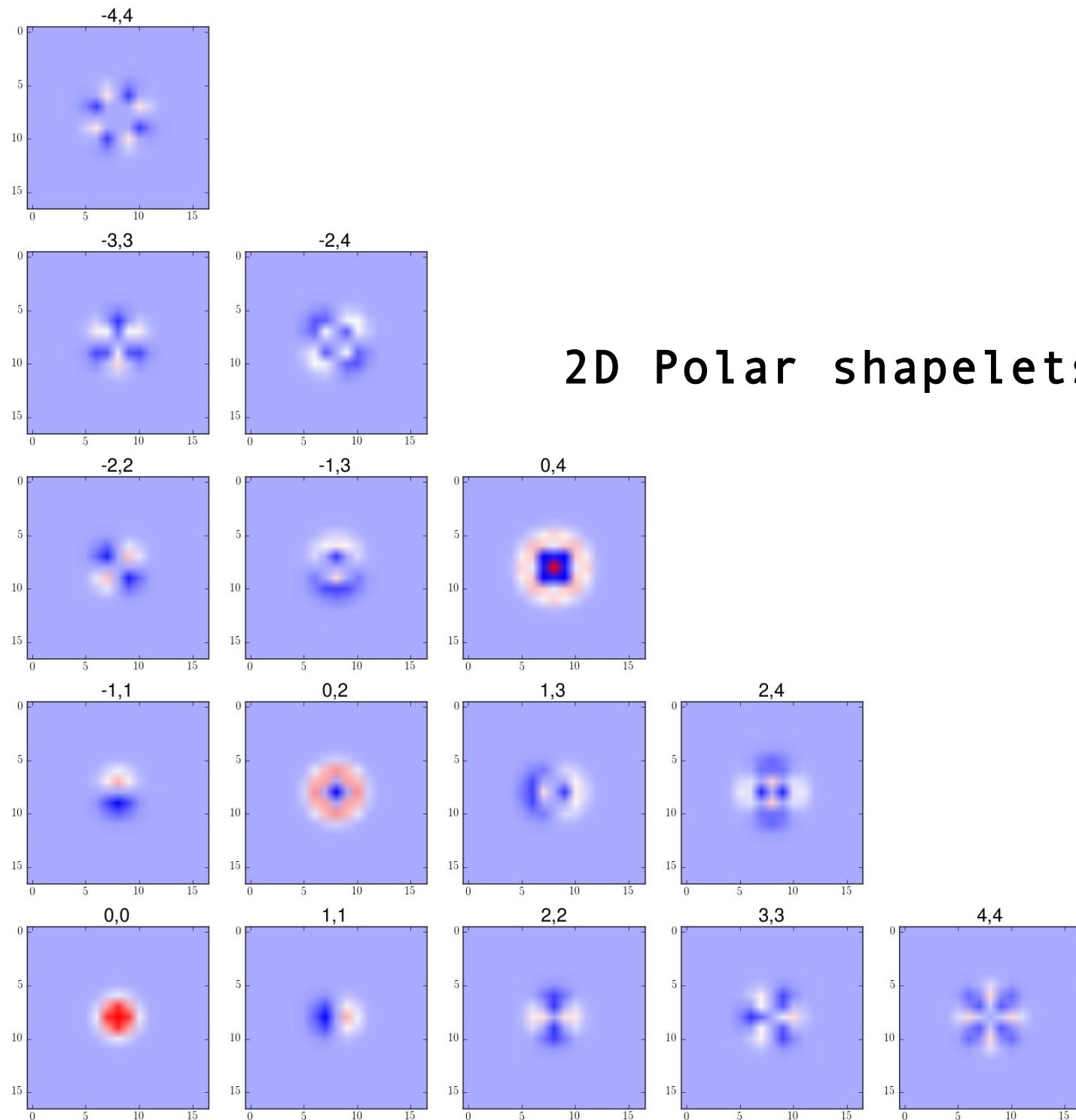


$$\phi_n(x) \equiv \left[ 2^n \pi^{\frac{1}{2}} n! \right]^{-\frac{1}{2}} H_n(x) e^{-\frac{x^2}{2}}$$

$$B_n(x; \beta) \equiv \beta^{-\frac{1}{2}} \phi_n(\beta^{-1} x)$$

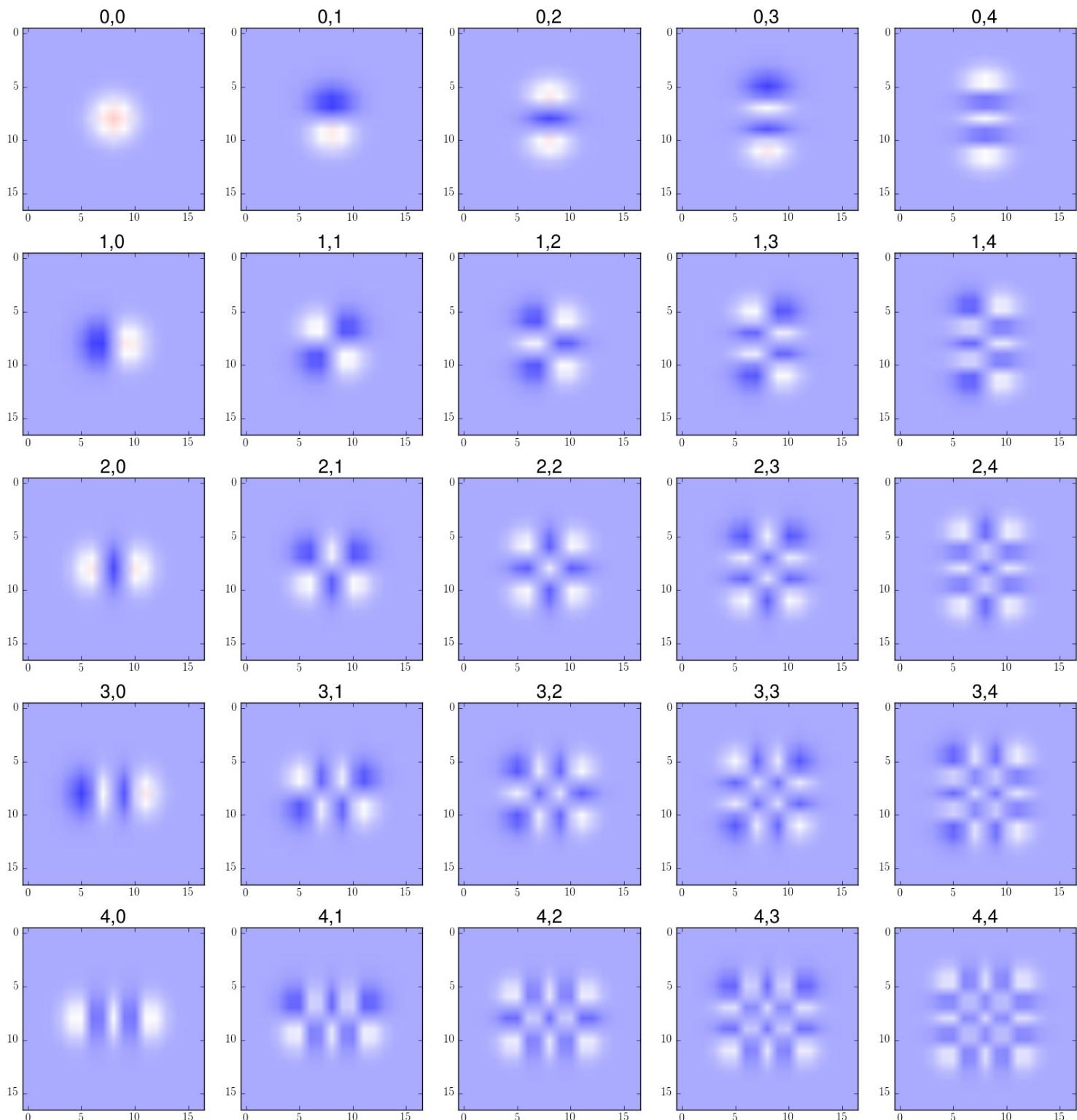
# Example of 1D shapelets decomposition





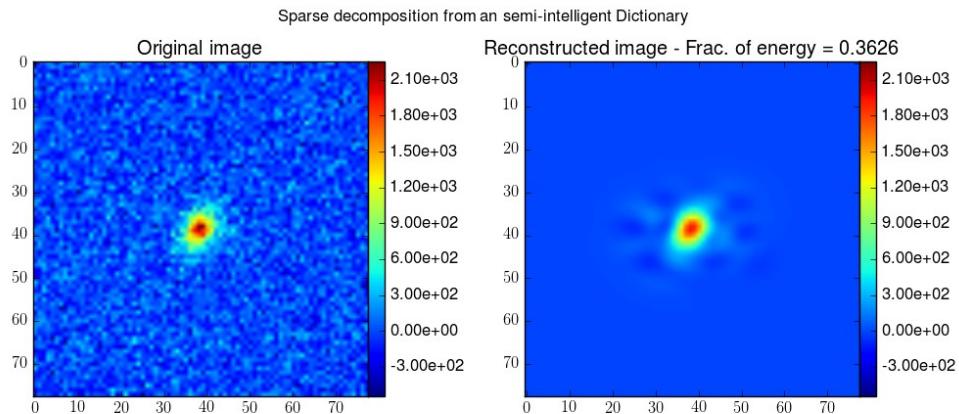
## 2D Polar shapelets

## 2D Cartesian shapelets

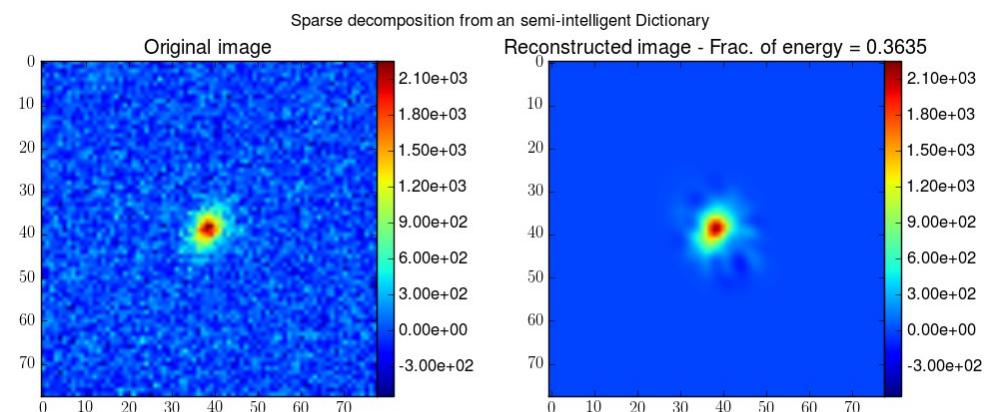


# Example of 2D shapelet decomposition

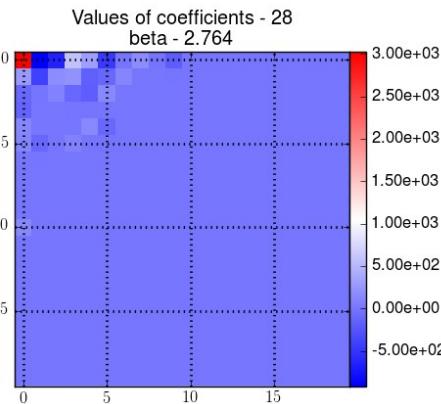
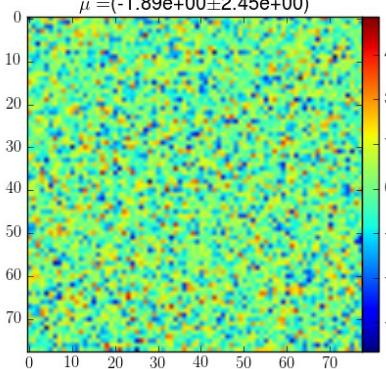
28 shapelets - OMP - Cartesian



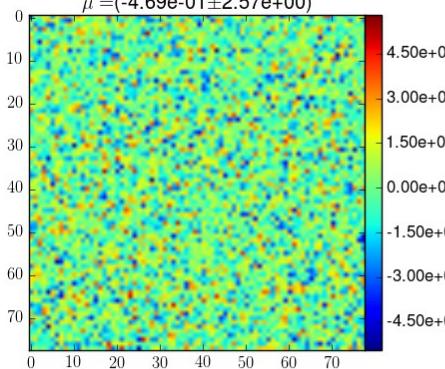
28 shapelets - OMP - Polar



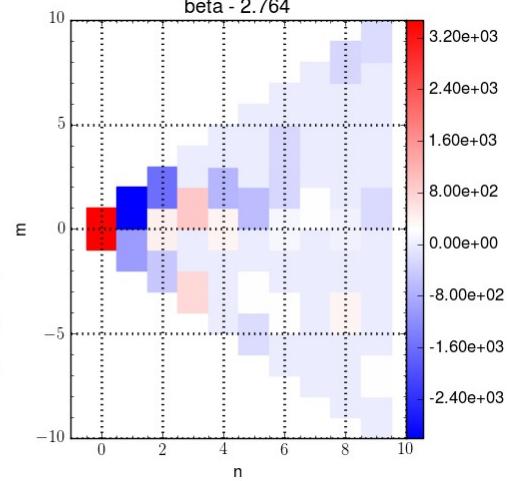
Residual image - frac. of energy = 0.6386  
 $\sigma = (1.63e+02 \pm 2.00e+00)$   
 $\mu = (-1.89e+00 \pm 2.45e+00)$



Residual image - frac. of energy = 0.6365  
 $\sigma = (1.63e+02 \pm 2.10e+00)$   
 $\mu = (-4.69e-01 \pm 2.57e+00)$



Values of coefficients - 28  
 $\beta = 2.764$



# Shapelets are easily transformed

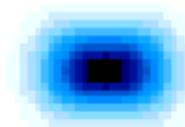
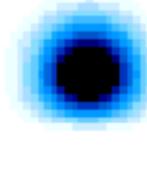
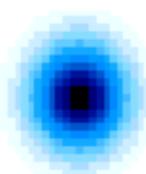
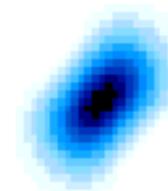
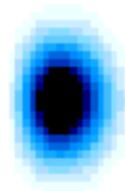
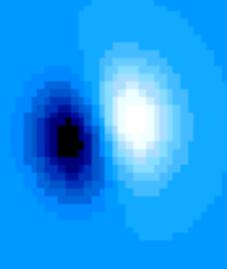
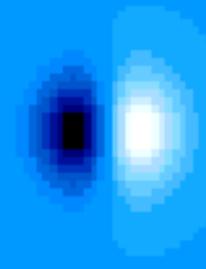
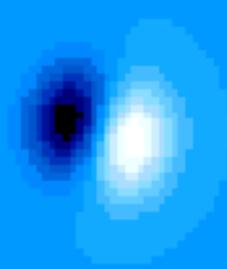
$$\mathbf{x} \rightarrow \mathbf{x}' = (1 + \Psi)\mathbf{x} + \epsilon \quad \Psi = \begin{pmatrix} \kappa + \gamma_1 & \gamma_2 - \rho \\ \gamma_2 + \rho & \kappa - \gamma_1 \end{pmatrix}$$

$$\downarrow \quad f(\mathbf{x}) = \sum_{i=0}^{\infty} \frac{\mathbf{x}' - \mathbf{x}}{i!} \nabla(f(\mathbf{x}))|_{\mathbf{x}=\mathbf{x}'}$$

$$\begin{aligned} \hat{R} &= -i(\hat{x}_1\hat{p}_2 - \hat{x}_2\hat{p}_1) = \hat{a}_1\hat{a}_2^\dagger - \hat{a}_1^\dagger\hat{a}_2 \\ \hat{K} &= -i(\hat{x}_1\hat{p}_1 + \hat{x}_2\hat{p}_2) = 1 + \frac{1}{2}(\hat{a}_1^{\dagger 2} + \hat{a}_2^{\dagger 2} - \hat{a}_1^2 - \hat{a}_2^2) \\ \hat{S}_1 &= -i(\hat{x}_1\hat{p}_1 - \hat{x}_2\hat{p}_2) = \frac{1}{2}(\hat{a}_1^{\dagger 2} - \hat{a}_2^{\dagger 2} - \hat{a}_1^2 + \hat{a}_2^2) \\ \hat{S}_2 &= -i(\hat{x}_1\hat{p}_2 + \hat{x}_2\hat{p}_1) = \hat{a}_1^\dagger\hat{a}_2^\dagger - \hat{a}_1\hat{a}_2 \\ \hat{T}_j &= -i\hat{p}_j = \frac{1}{\sqrt{2}}(\hat{a}_j^\dagger - \hat{a}_j), \quad j = 1, 2. \end{aligned}$$

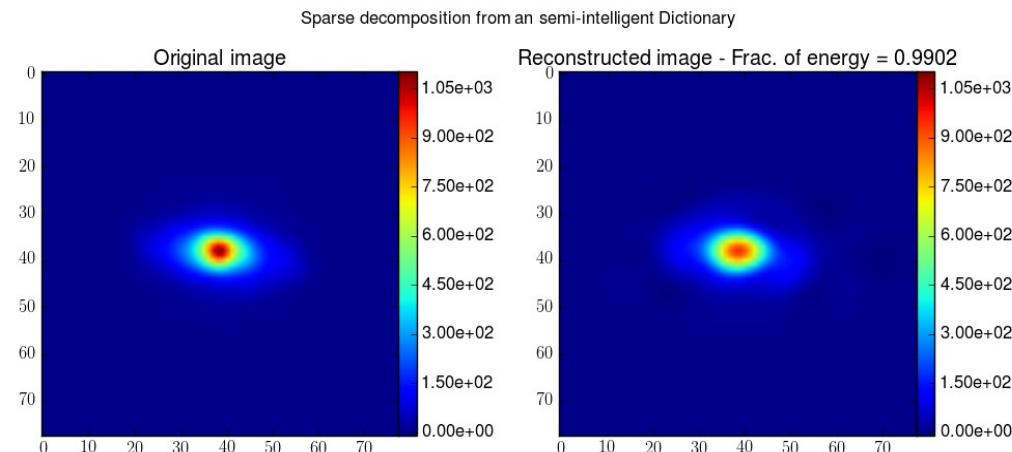
$$\downarrow$$

$$f' \simeq (1 + \rho\hat{R} + \kappa\hat{K} + \gamma_j\hat{S}_j + \epsilon_i\hat{T}_i)f$$

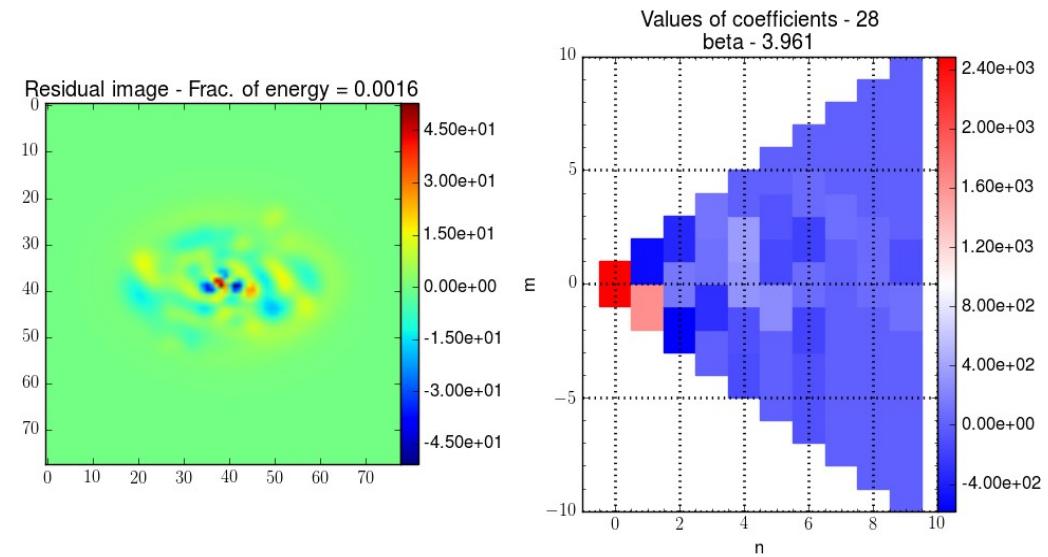
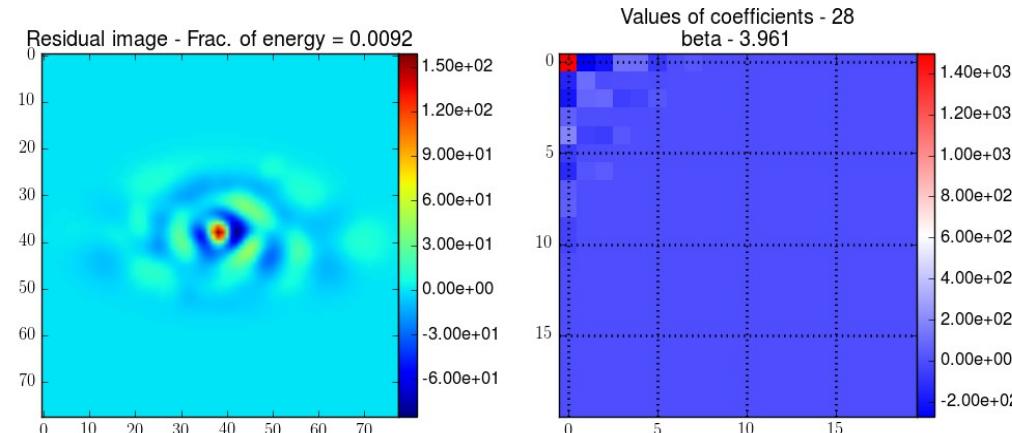
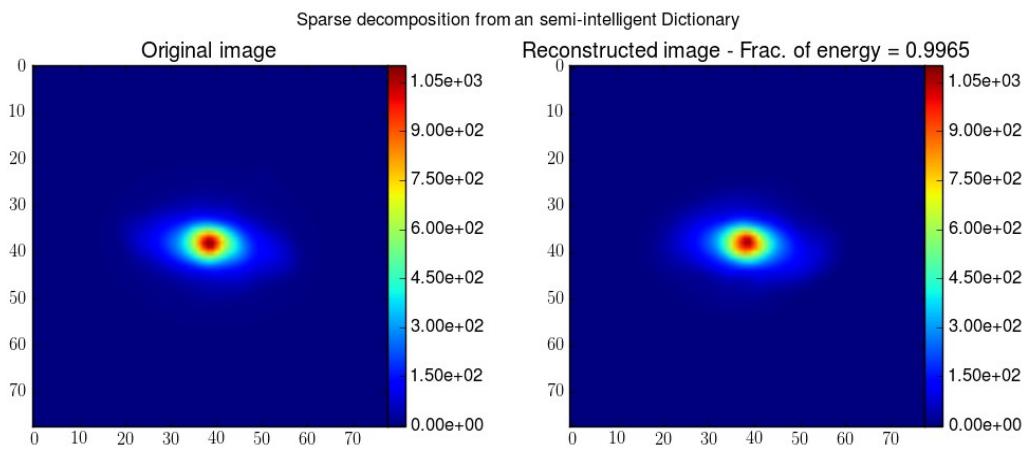
$|00\rangle$  $(1+\epsilon_2 T_2)|00\rangle$  $(1+\gamma_1 S_1)|00\rangle$  $(1+\epsilon_1 T_1)|00\rangle$  $(1+\kappa K)|00\rangle$  $(1+\gamma_2 S_2)|00\rangle$  $(1-\rho R)|10\rangle$  $|10\rangle$  $(1+\rho R)|10\rangle$ 

# Capturing the profile well

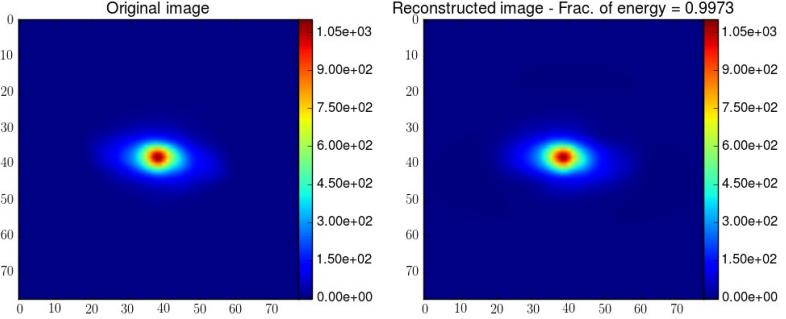
28 shapelets - OMP - Cartesian



28 shapelets - OMP - Polar



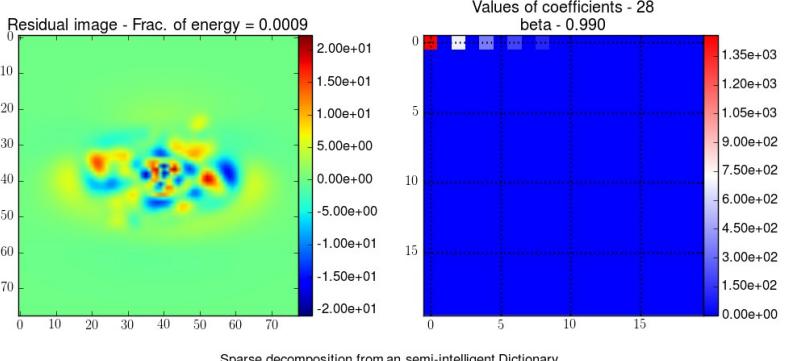
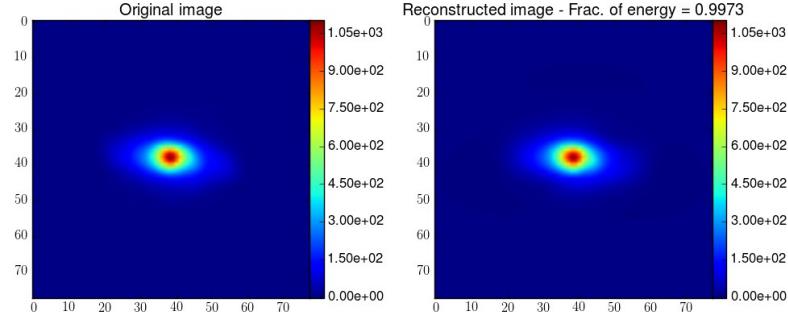
Sparse decomposition from an semi-intelligent Dictionary



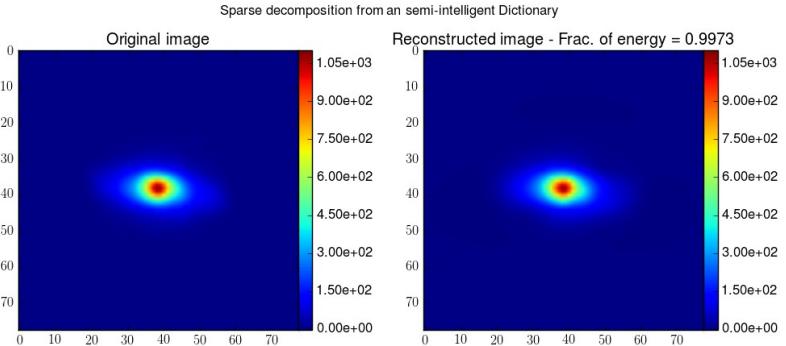
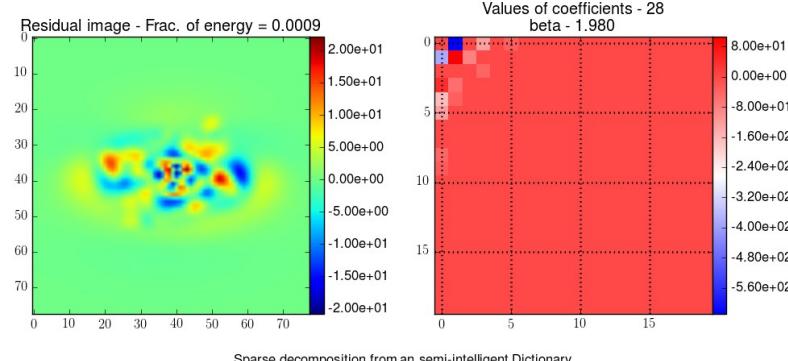
## Compound XY

$$\beta = 0.99$$

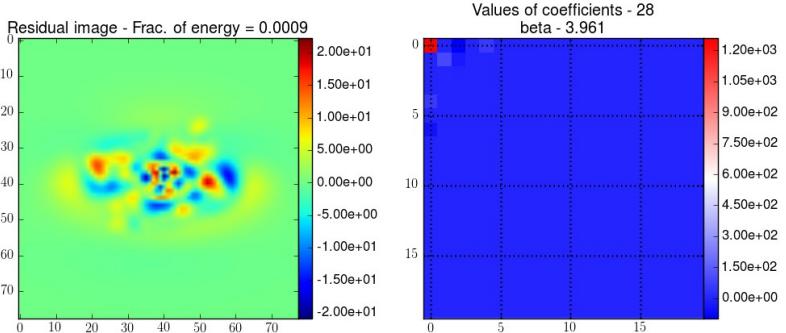
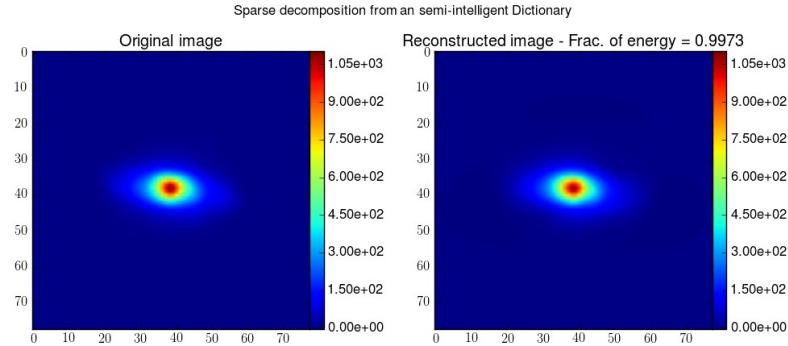
Sparse decomposition from an semi-intelligent Dictionary



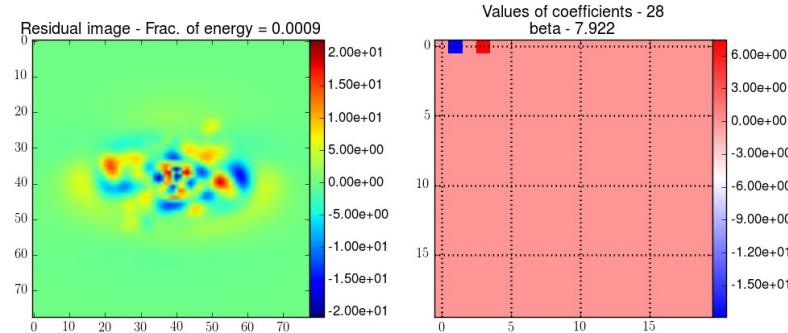
$$\beta = 1.98$$



$$\beta = 3.97$$

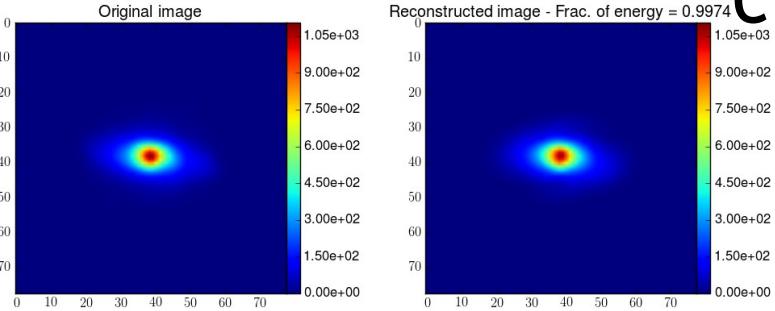


$$\beta = 7.92$$

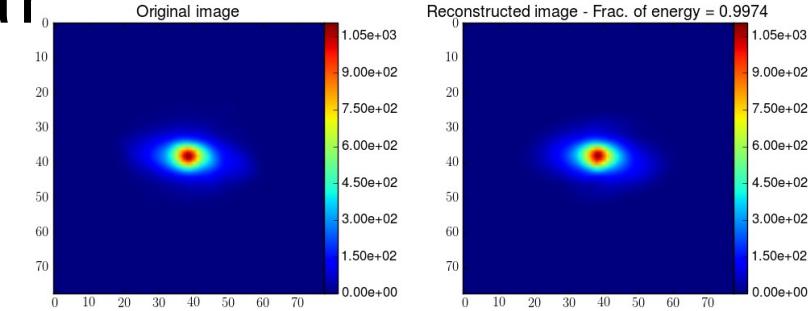


# Compound Polar

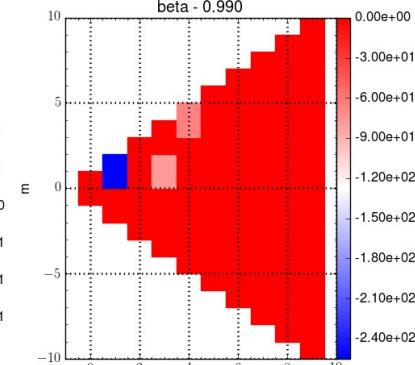
Sparse decomposition from an semi-intelligent Dictionary



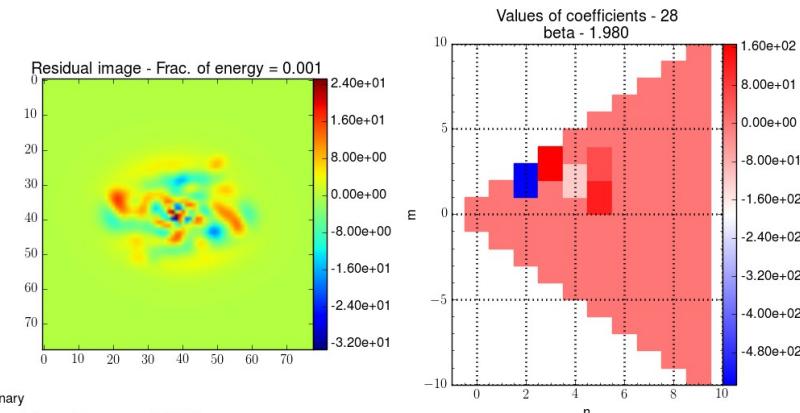
Sparse decomposition from an semi-intelligent Dictionary



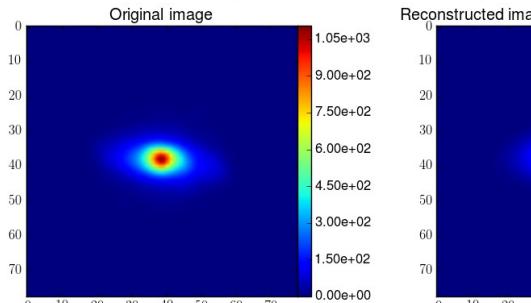
Values of coefficients - 28



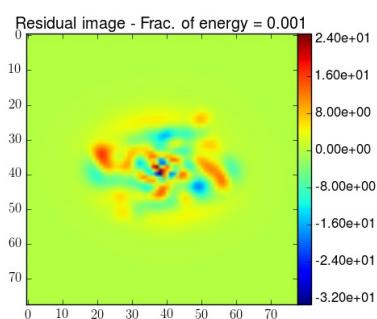
$\beta = 1.98$



Sparse decomposition from an semi-intelligent Dictionary

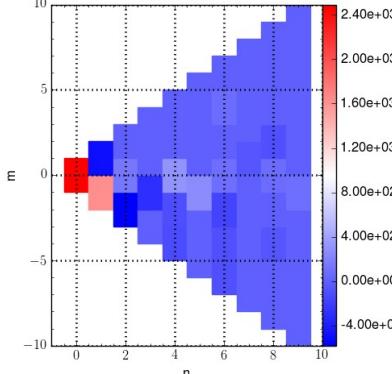


Residual image - Frac. of energy = 0.001



Values of coefficients - 28

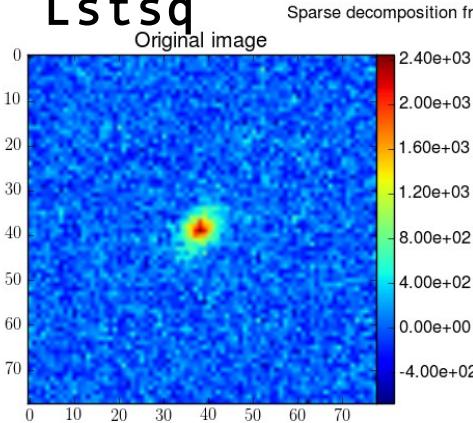
$\beta = 3.961$



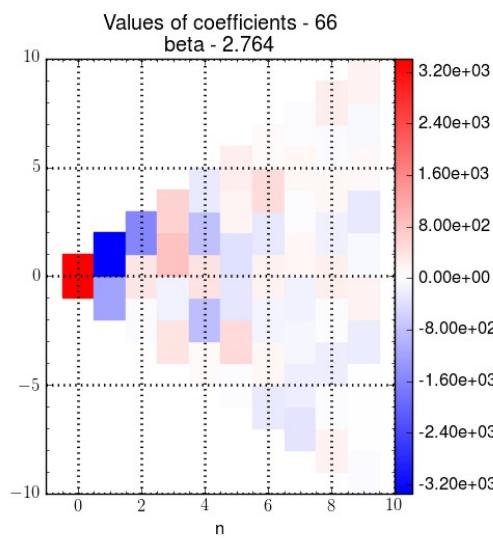
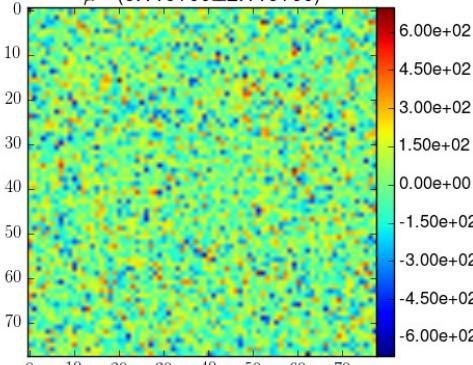
# Need to use sparse solvers

- Use of Compound basis asks for solution restriction
- Solvers considered:
  - Without free parameters:
    - Least squares
$$\min \|y - Xw\|_2^2$$
    - Singular value decomposition (SVD)
  - Sparse solvers:
    - Lasso - regression method
$$\min \frac{1}{2N} \|y - Xw\|_2^2 + \lambda \|w\|_1$$
    - Orthogonal matching pursuit (OMP)
$$\min \|y - Xw\|_2^2 \text{ s.t. } \|w\|_0 \leq L$$

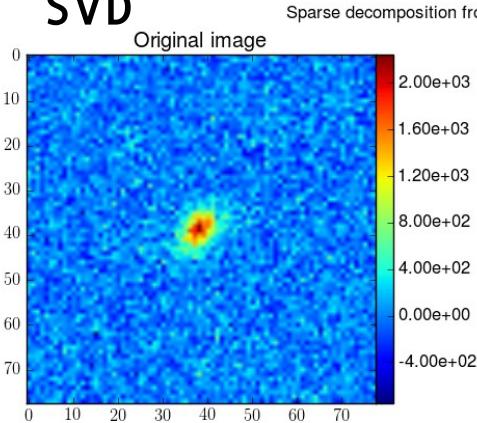
**Lstsq**



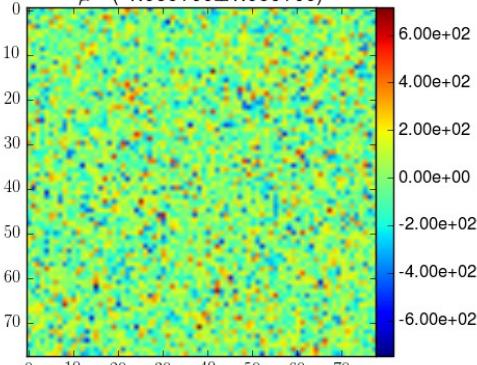
Residual image - frac. of energy = 0.6941  
 $\sigma = (1.86e+02 \pm 1.73e+00)$   
 $\mu = (5.11e+00 \pm 2.11e+00)$



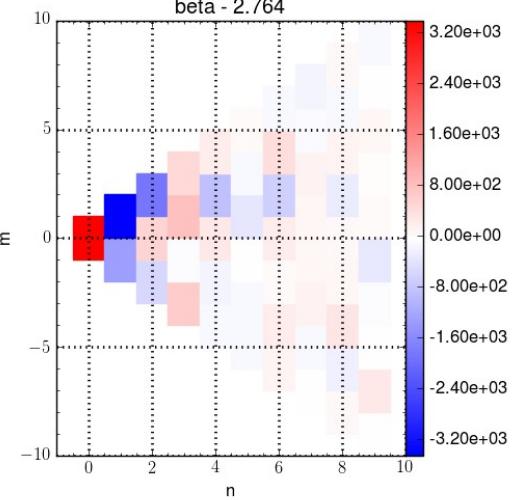
**SVD**



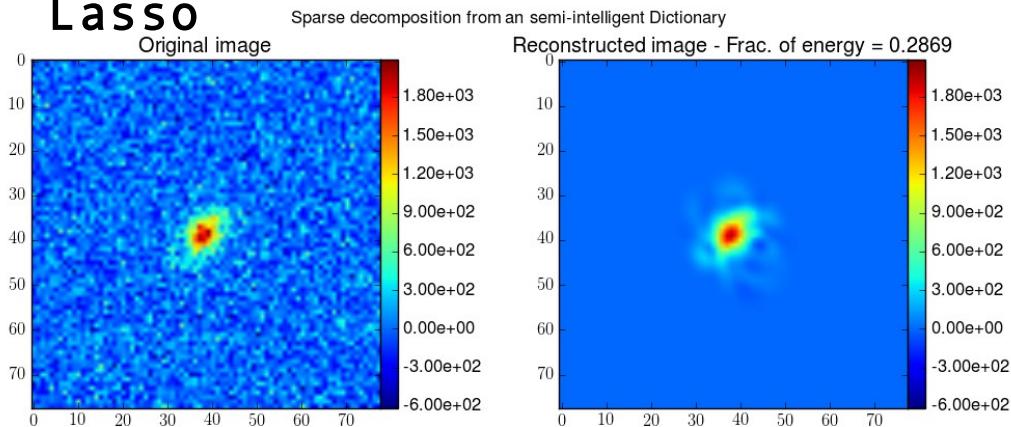
Residual image - frac. of energy = 0.6952  
 $\sigma = (1.84e+02 \pm 1.59e+00)$   
 $\mu = (-1.98e+00 \pm 1.95e+00)$



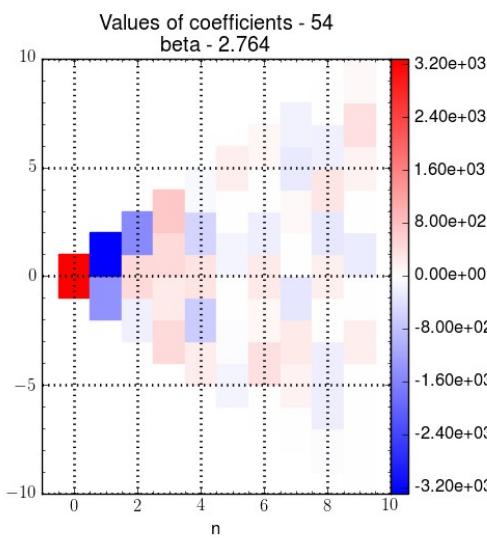
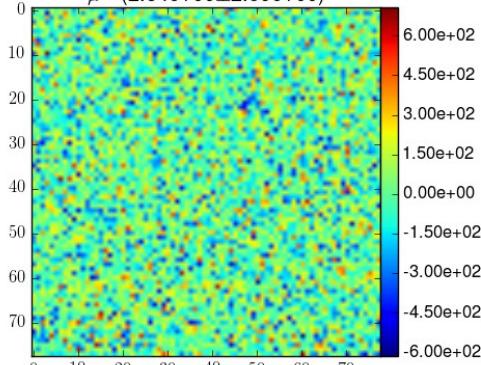
Values of coefficients - 66  
 $\beta = 2.764$



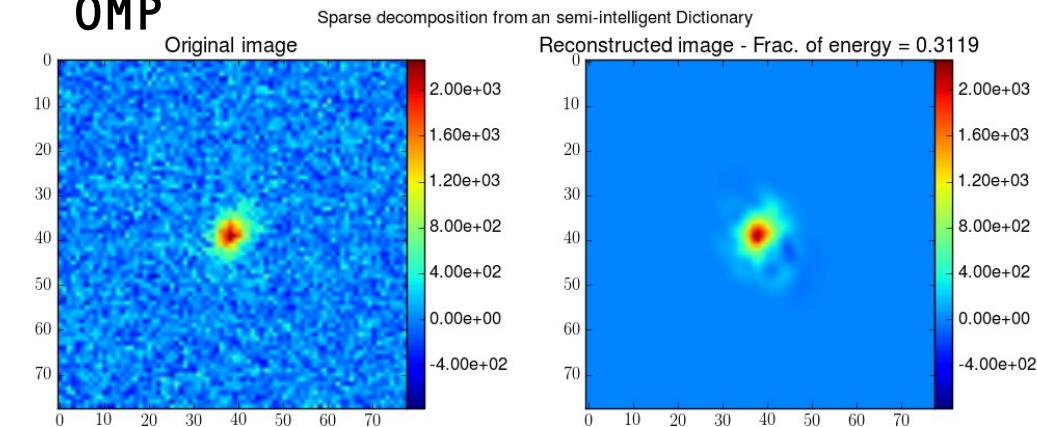
## Lasso



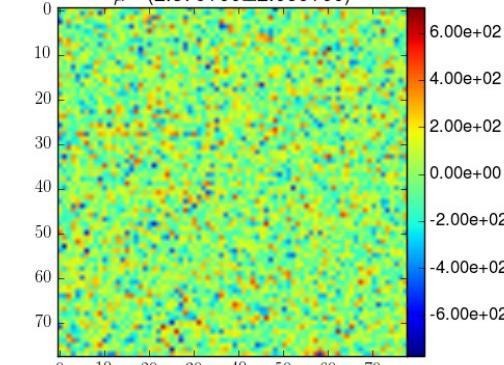
Residual image - frac. of energy = 0.7044  
 $\sigma = (1.84e+02 \pm 2.12e+00)$   
 $\mu = (2.84e+00 \pm 2.59e+00)$



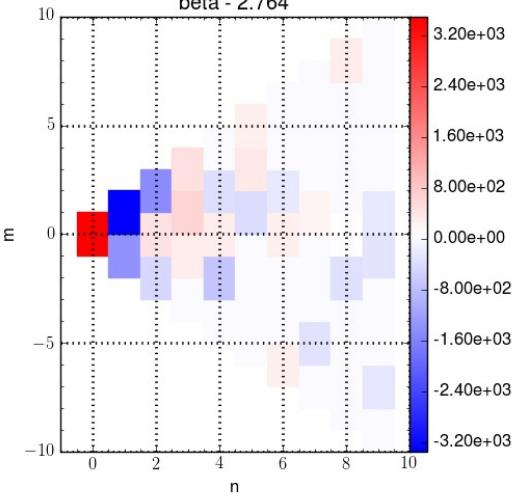
## OMP



Residual image - frac. of energy = 0.6871  
 $\sigma = (1.82e+02 \pm 1.70e+00)$   
 $\mu = (2.37e+00 \pm 2.08e+00)$

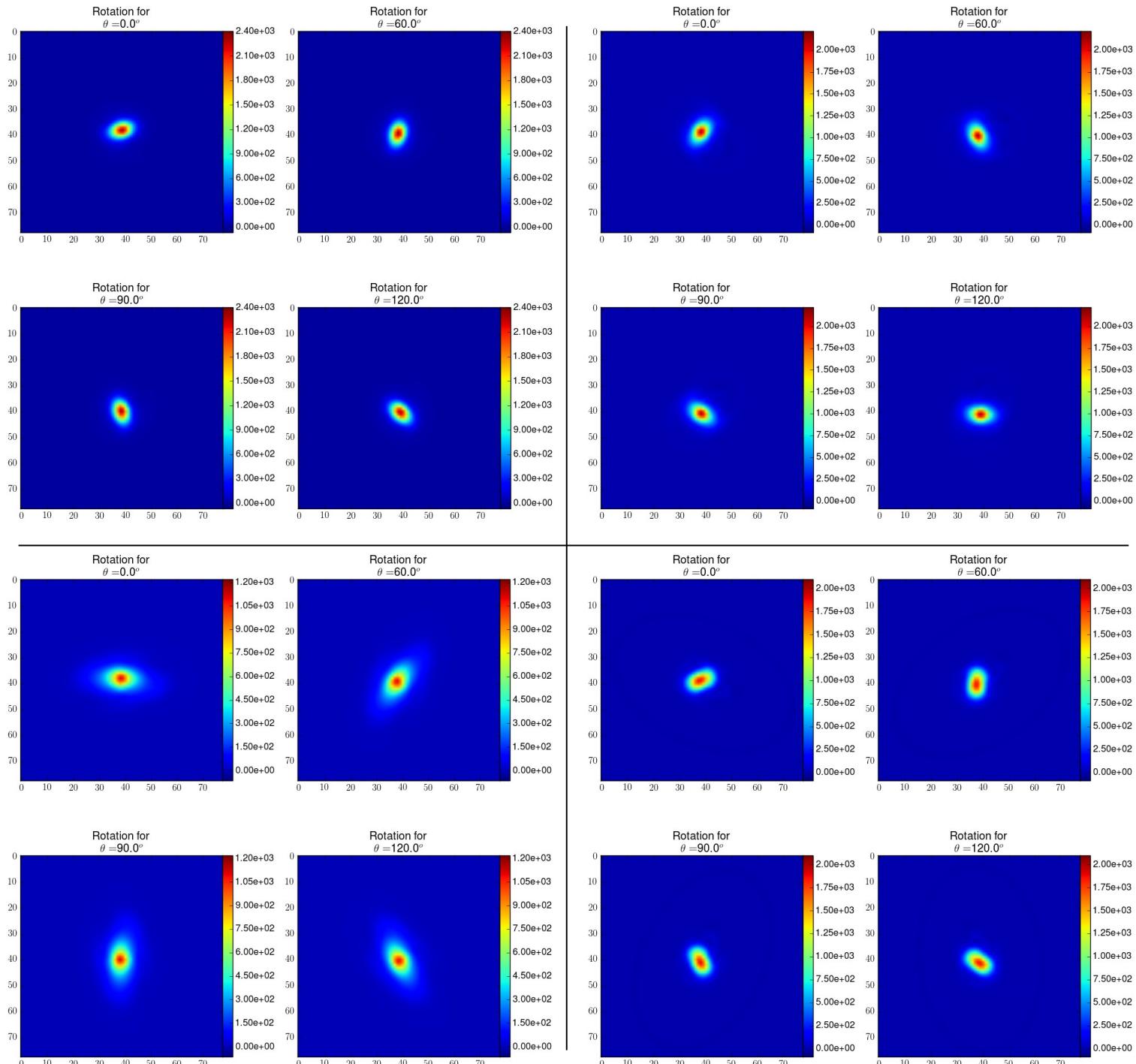


Values of coefficients - 28  
beta - 2.764

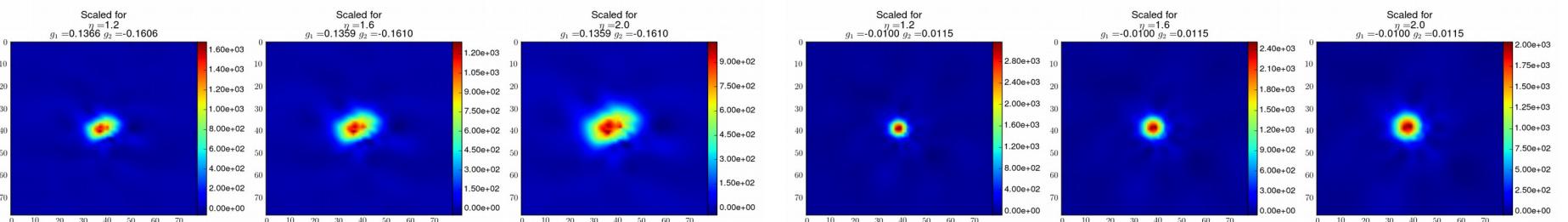
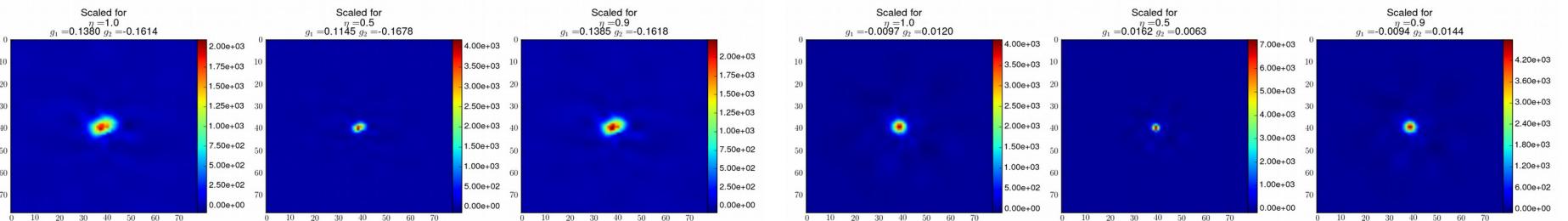
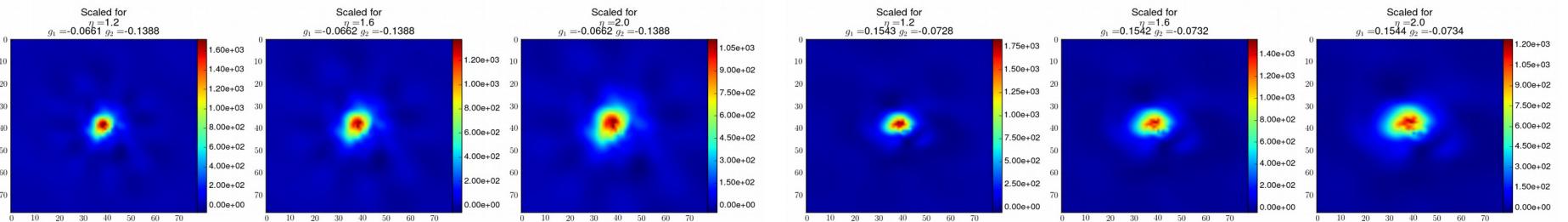
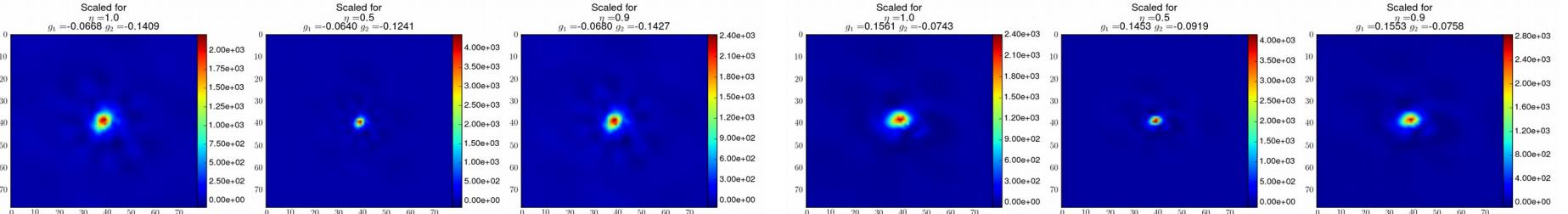


# Creating mock galaxies

Rotation

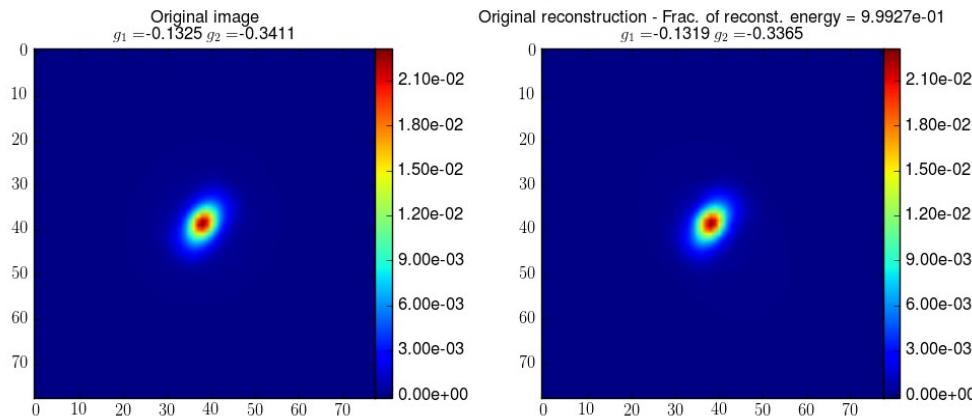


# Scaling

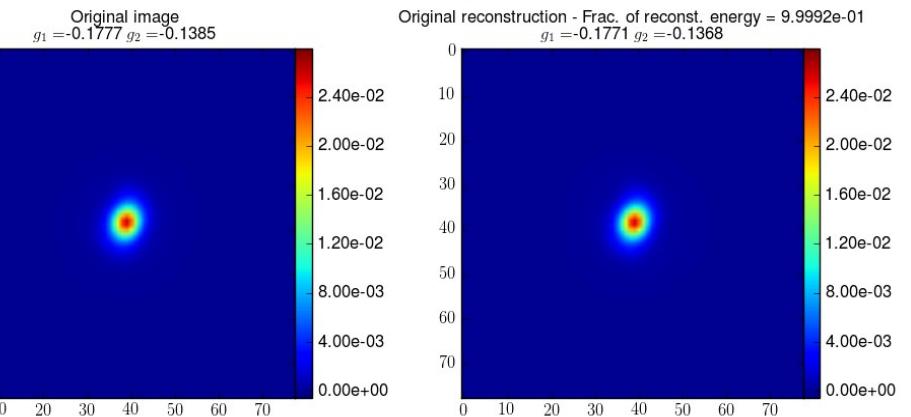


# Perturbing the coefficients

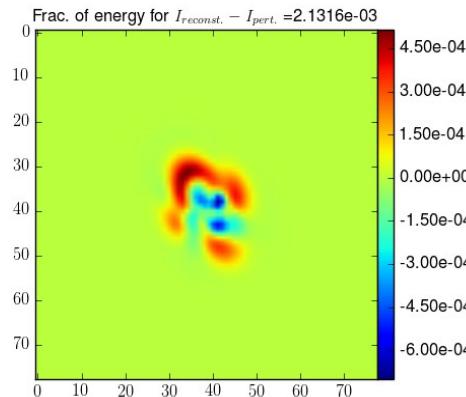
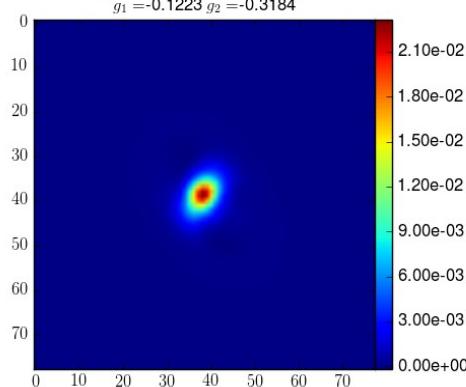
Mock galaxy image by perturbation



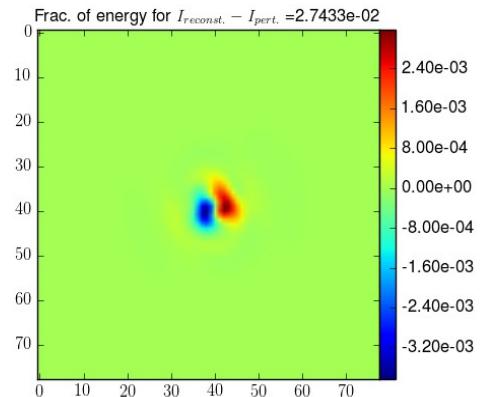
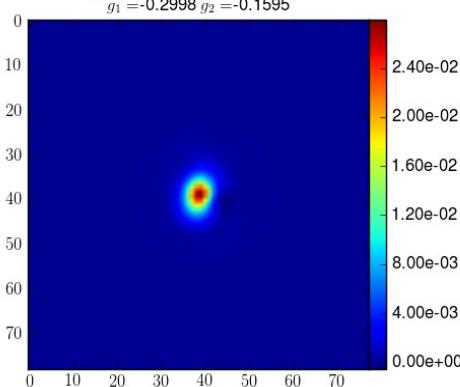
Mock galaxy image by perturbation



Perturbed image - Frac. of reconst. energy = 1.0226e+00

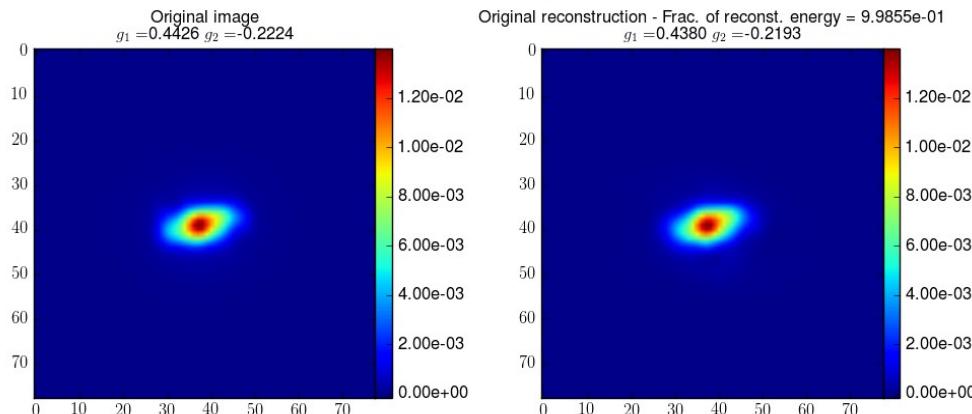


Perturbed image - Frac. of reconst. energy = 1.0981e+00

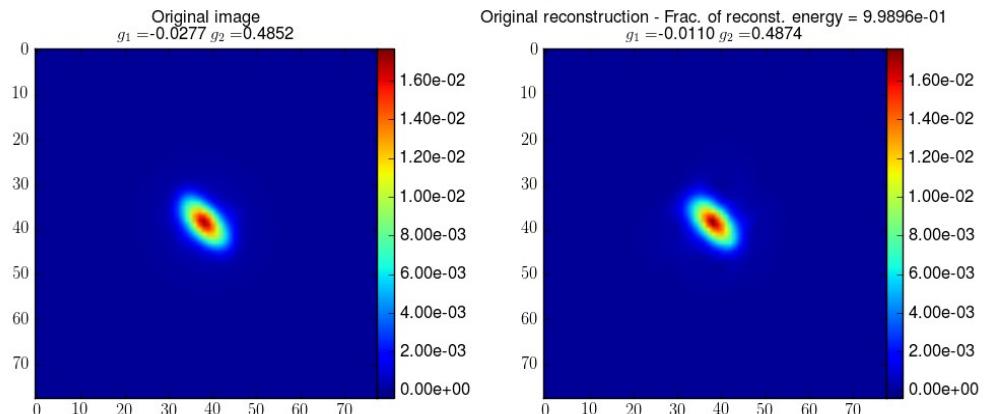


# Perturbing the coefficients

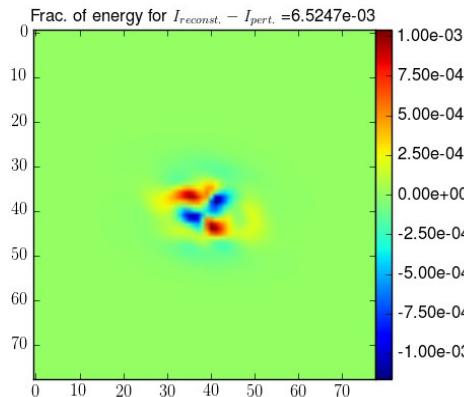
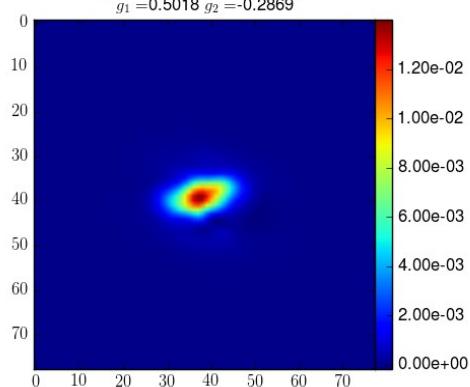
Mock galaxy image by perturbation



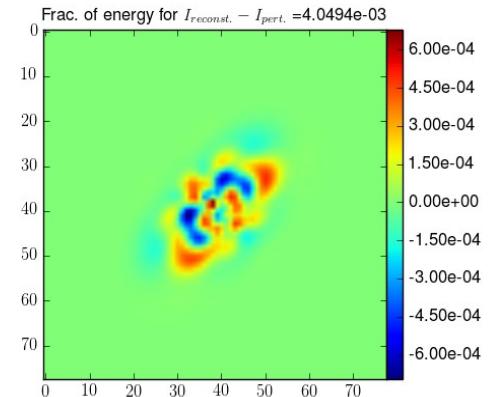
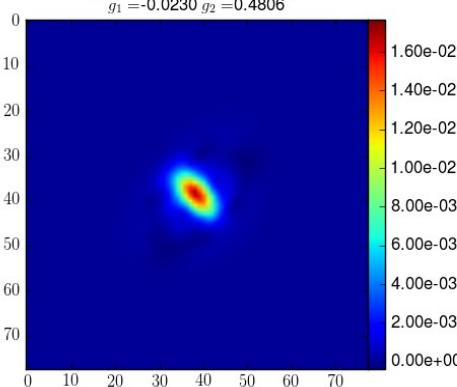
Mock galaxy image by perturbation



Perturbed image - Frac. of reconst. energy = 1.0223e+00

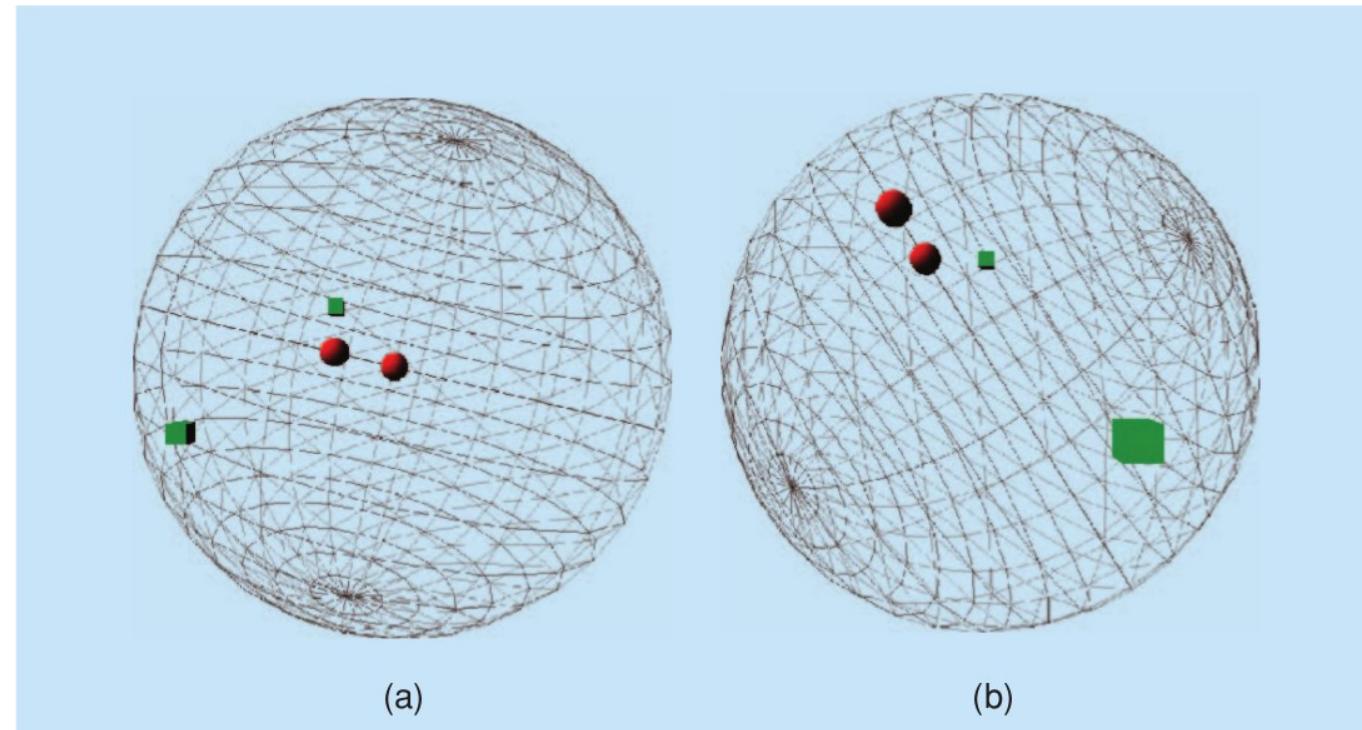


Perturbed image - Frac. of reconst. energy = 9.8119e-01

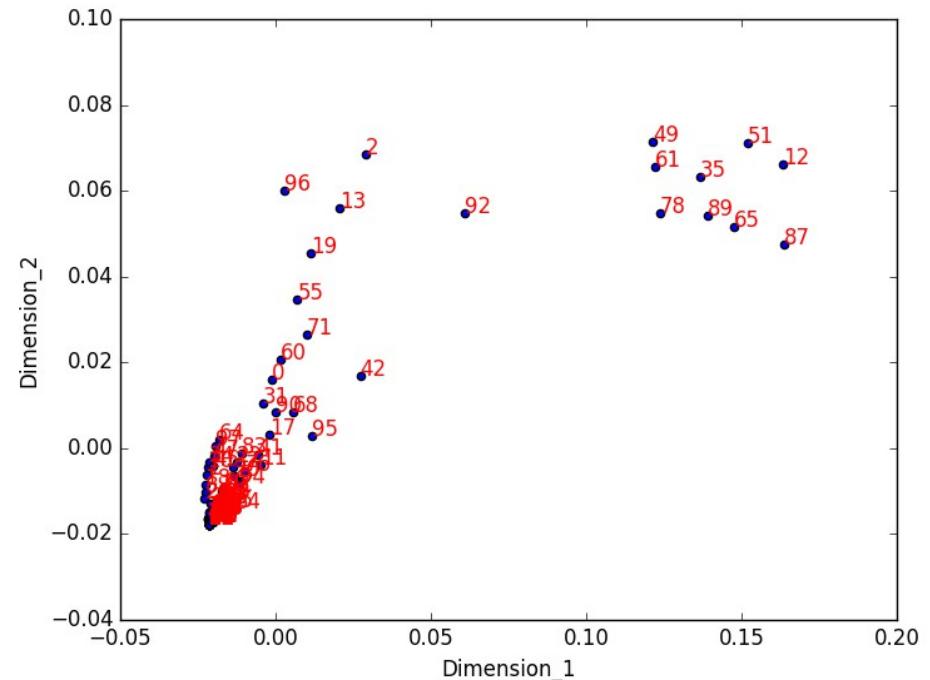
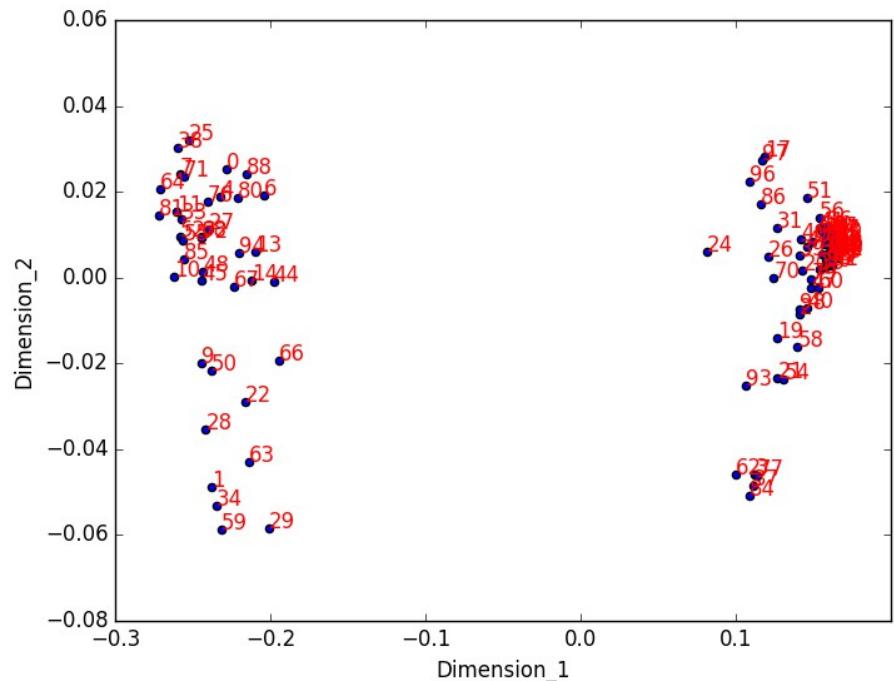


# Motivation for clustering

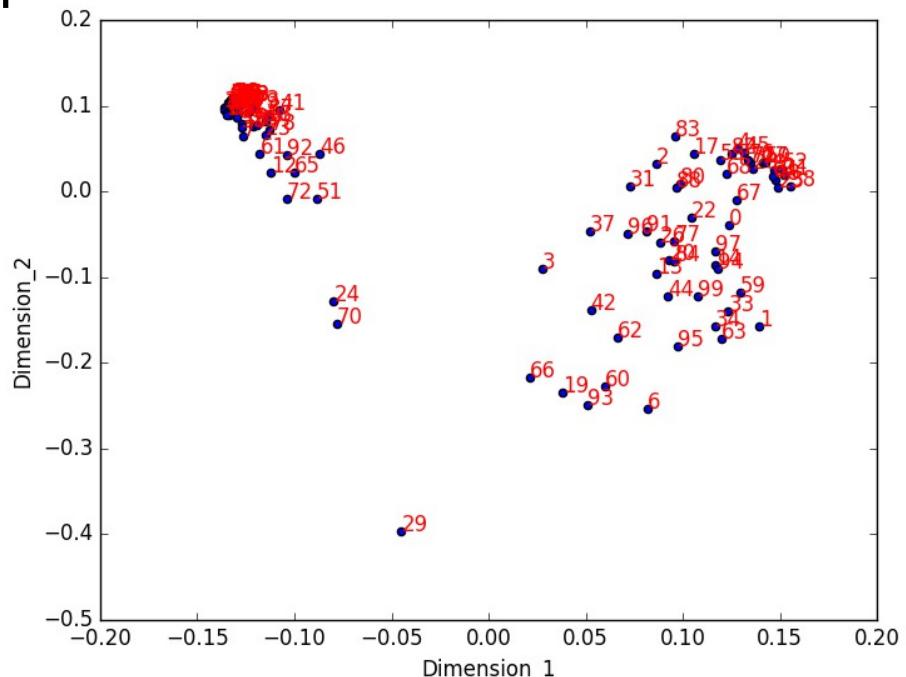
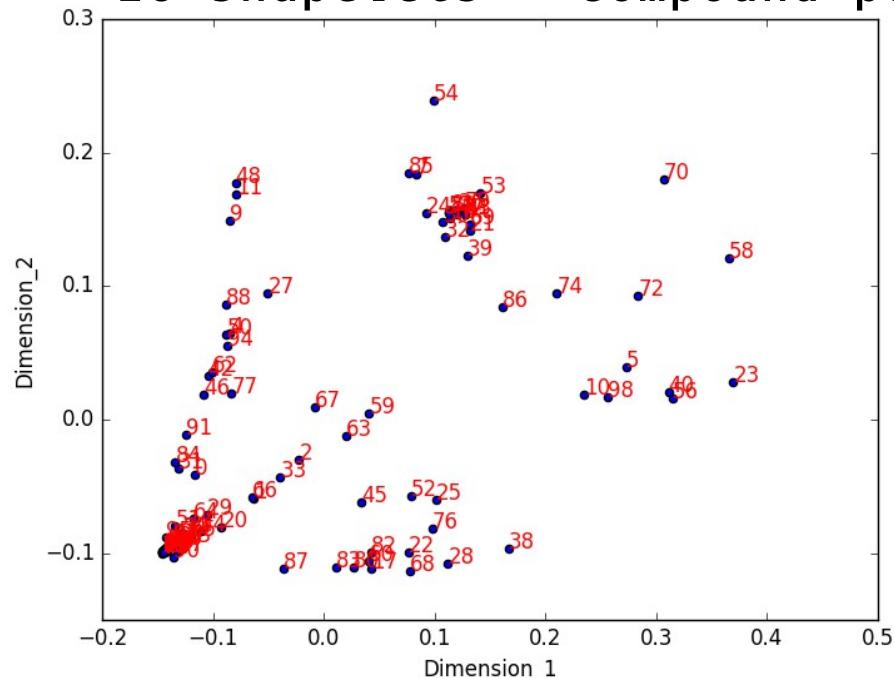
- We want to do re-sampling of the high dimensional distribution
- Generally one needs to project down the problem



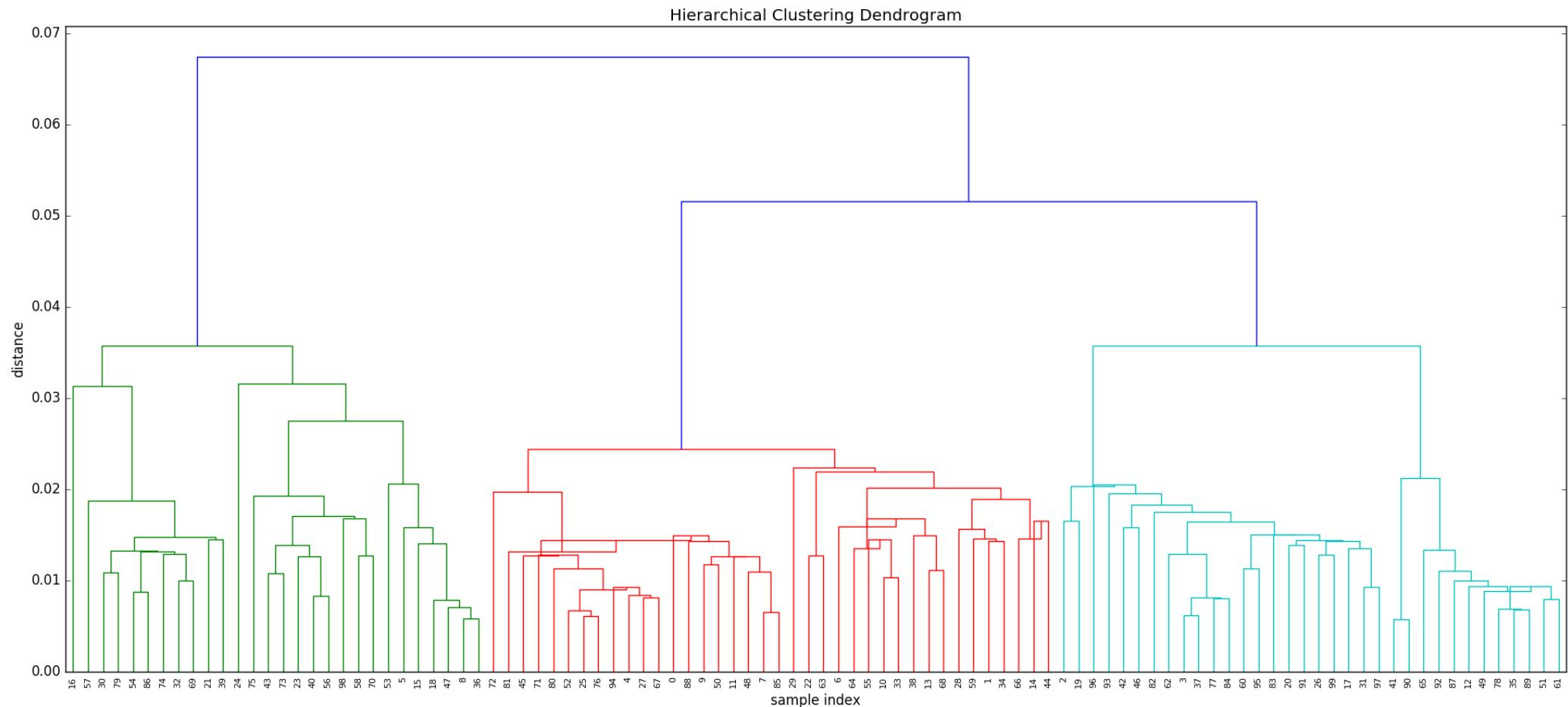
## 28 shapelets - Compound XY



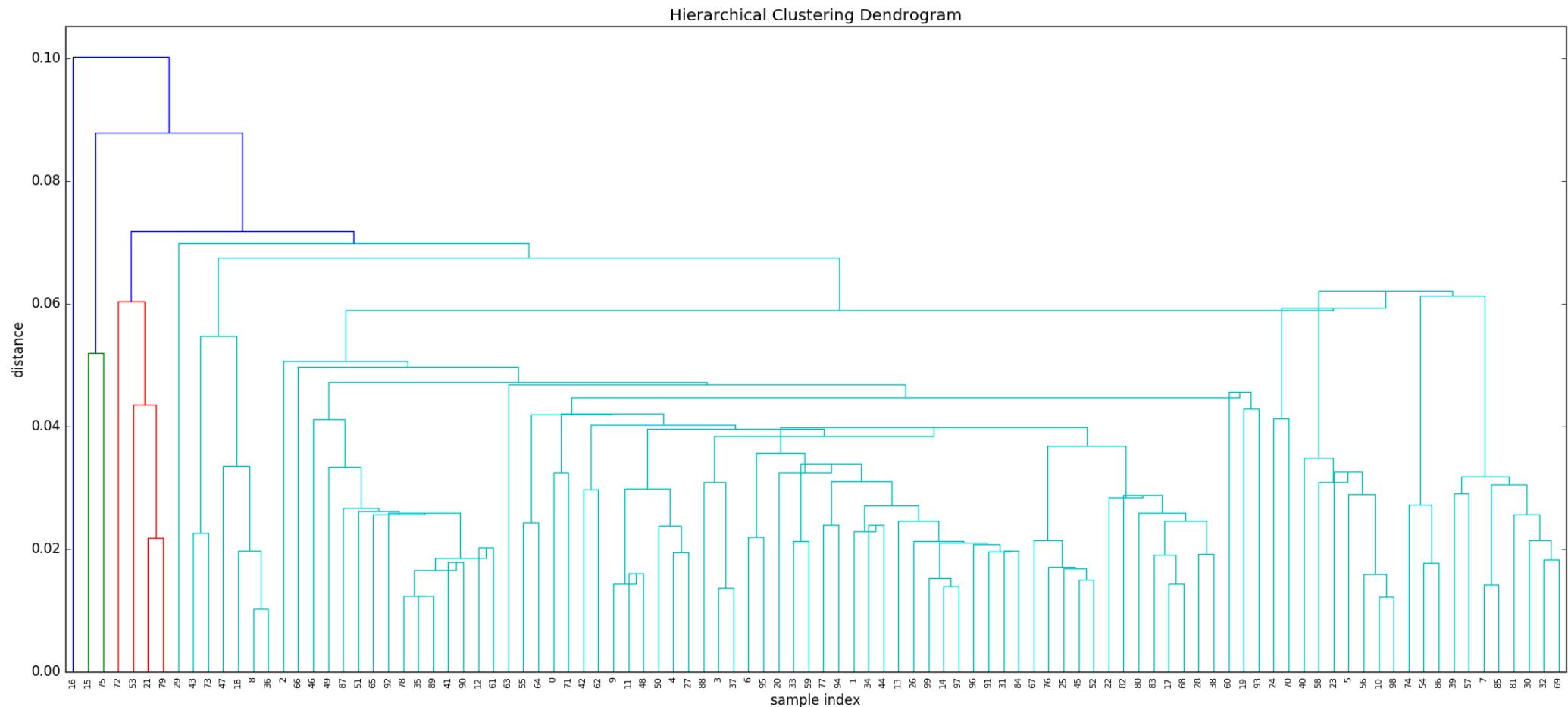
## 28 shapelets - Compound polar



# Finding clustering

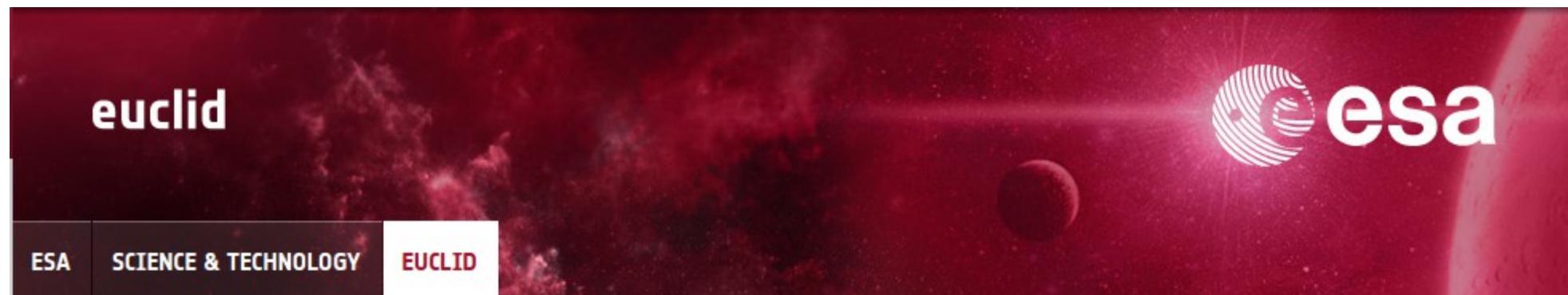


# Finding clustering



# Conclusion and Future work

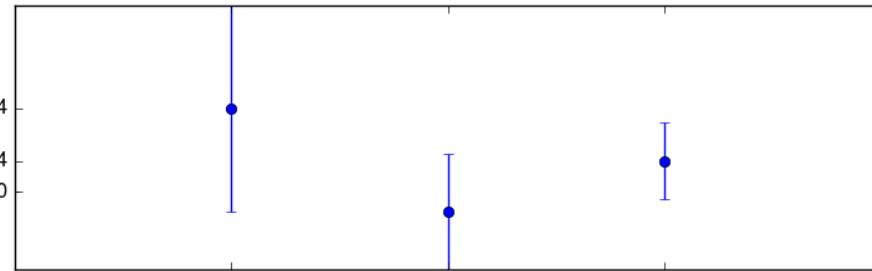
- Shapelets are good at capturing the profile and easily manipulated
- Potential of creating good mock catalogues with high enough details for a good bias estimate  
*(Self organizing maps)*
- Need to asses the problem of PSF deconvolution
- Especially useful for the upcoming missions → Euclid mission
- It would be good to see how well the bias can be reduced with shapelets



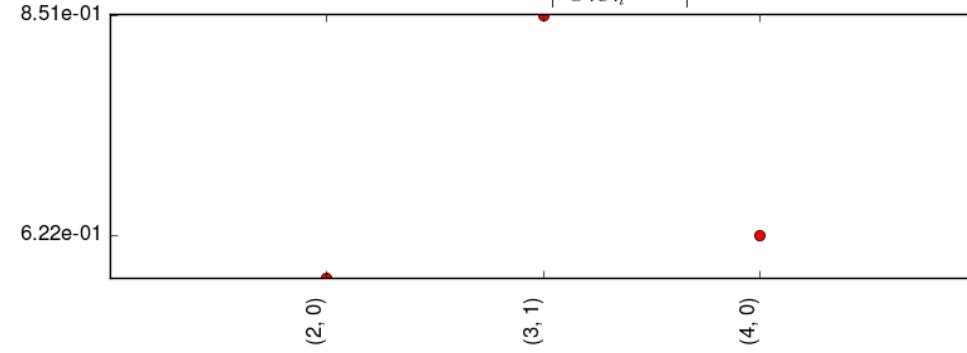
Thank you for your  
attention.

# Stability tests

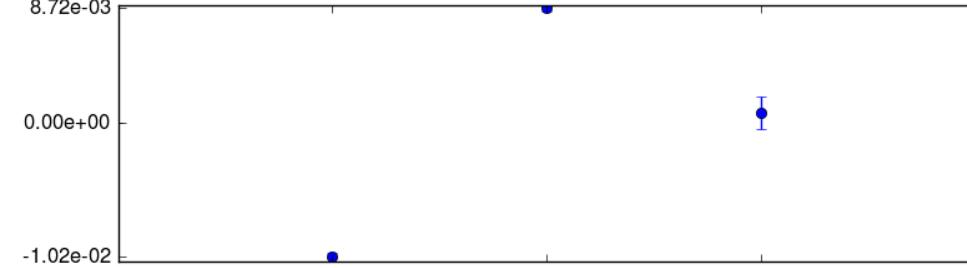
Scatter plot of  $\langle N.C_i \rangle$  for 3 biggest  $\langle O.C_i \rangle$  coeffs



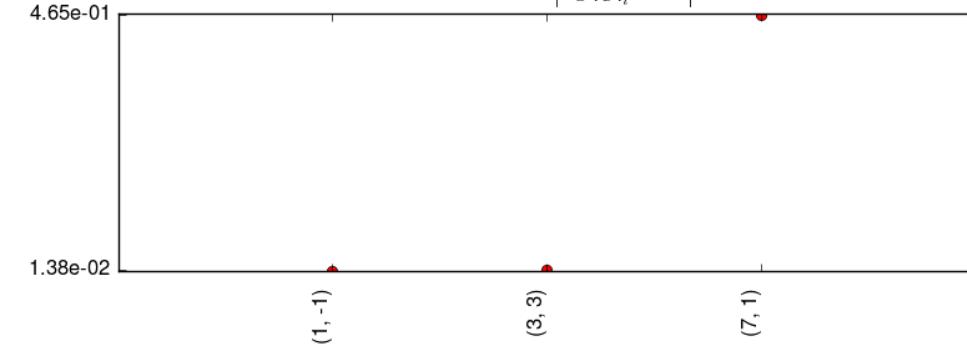
Scatter plot of  $\left| \frac{\langle N.C_i \rangle}{O.C_i} - 1 \right|$



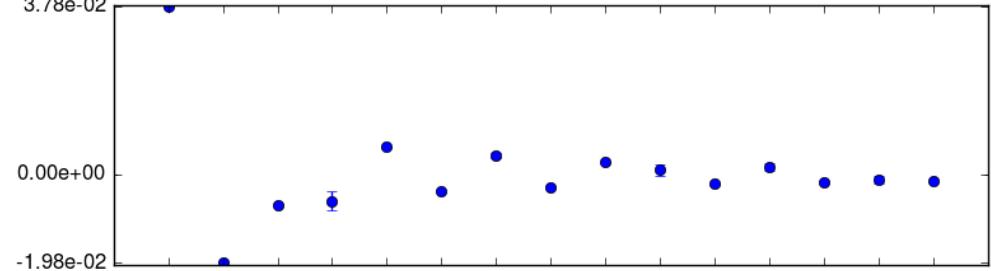
Scatter plot of  $\langle N.C_i \rangle$  for 3 biggest  $\langle O.C_i \rangle$  coeffs



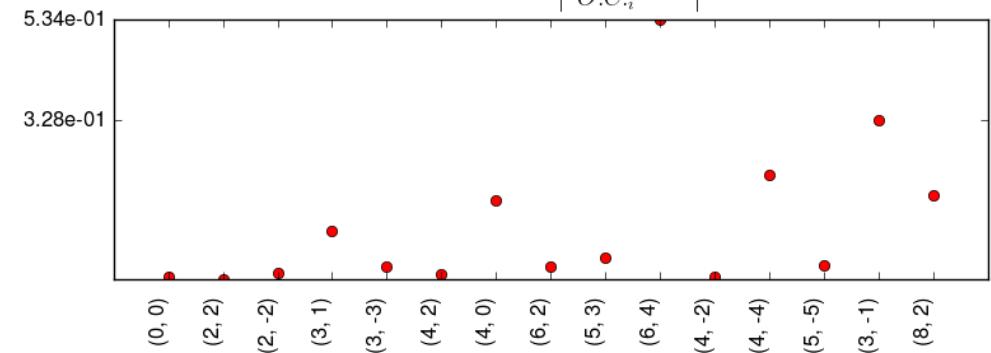
Scatter plot of  $\left| \frac{\langle N.C_i \rangle}{O.C_i} - 1 \right|$



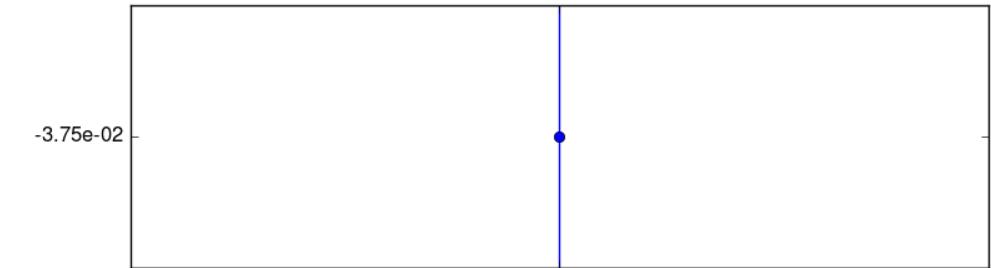
Scatter plot of  $\langle N.C_i \rangle$  for 15 biggest  $\langle O.C_i \rangle$  coeffs



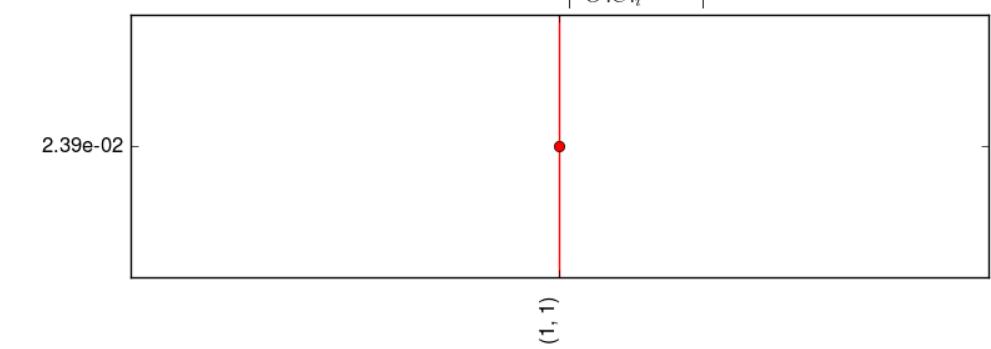
Scatter plot of  $\left| \frac{\langle N.C_i \rangle}{O.C_i} - 1 \right|$



Scatter plot of  $\langle N.C_i \rangle$  for 1 biggest  $\langle O.C_i \rangle$  coeffs

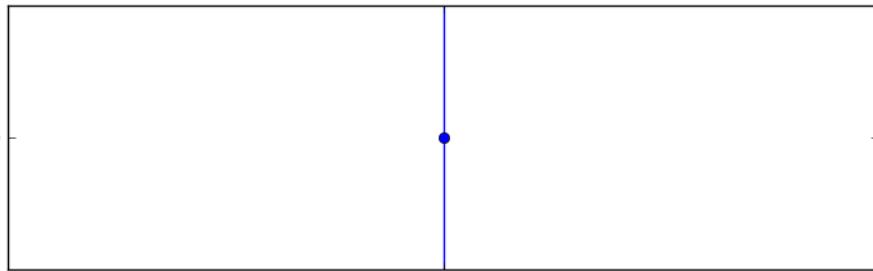


Scatter plot of  $\left| \frac{\langle N.C_i \rangle}{O.C_i} - 1 \right|$

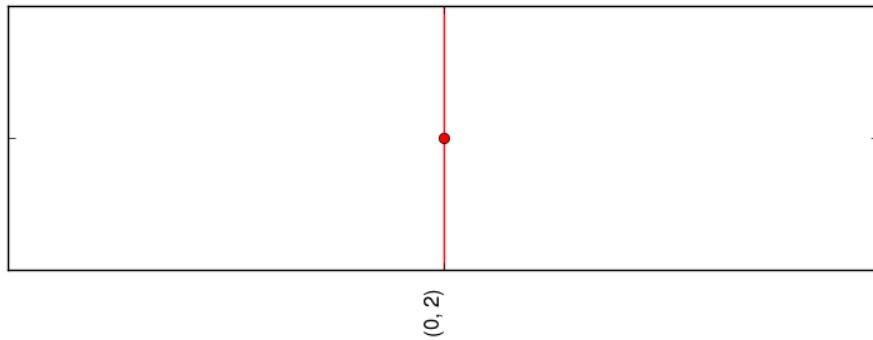


# Stability tests

Scatter plot of  $\langle N.C_i \rangle$  for 1 biggest  $\langle O.C_i \rangle$  coeffs

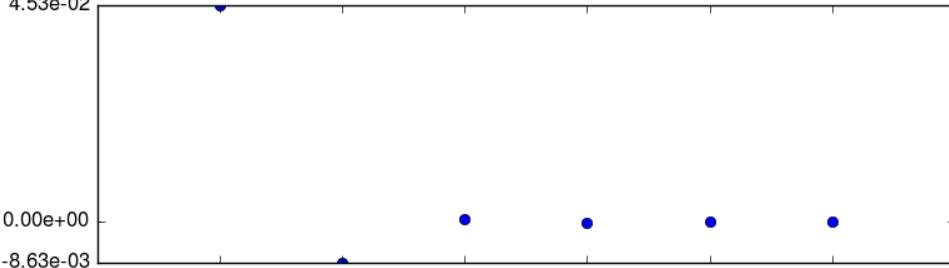


Scatter plot of  $\left| \frac{\langle N.C_i \rangle}{\langle O.C_i \rangle} - 1 \right|$

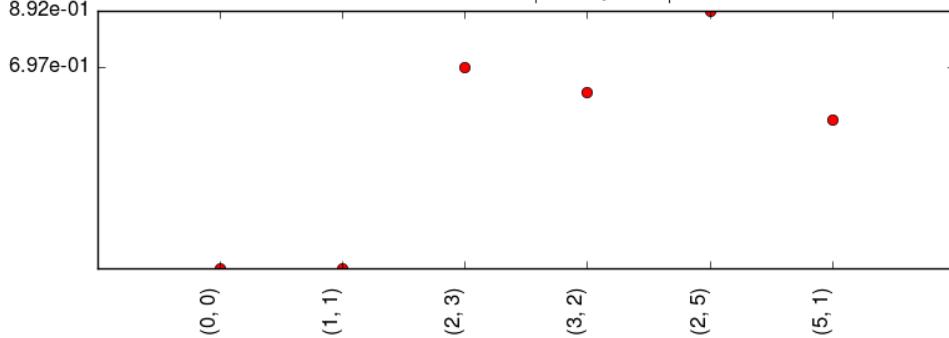


(0, 2)

Scatter plot of  $\langle N.C_i \rangle$  for 6 biggest  $\langle O.C_i \rangle$  coeffs



Scatter plot of  $\left| \frac{\langle N.C_i \rangle}{\langle O.C_i \rangle} - 1 \right|$



(0, 0)

(1, 1)

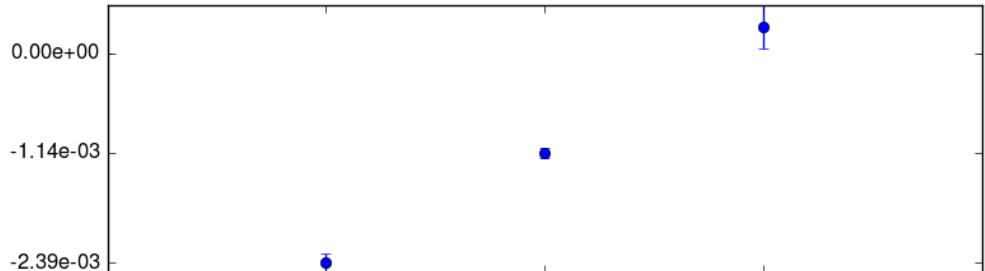
(2, 3)

(3, 2)

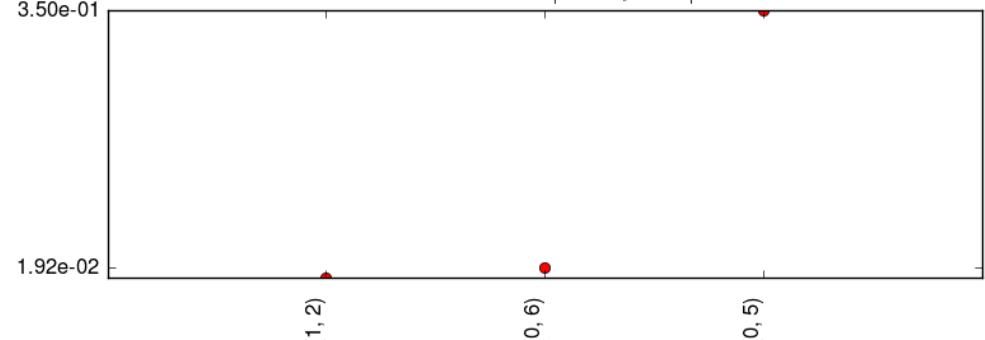
(2, 5)

(5, 1)

Scatter plot of  $\langle N.C_i \rangle$  for 3 biggest  $\langle O.C_i \rangle$  coeffs



Scatter plot of  $\left| \frac{\langle N.C_i \rangle}{\langle O.C_i \rangle} - 1 \right|$

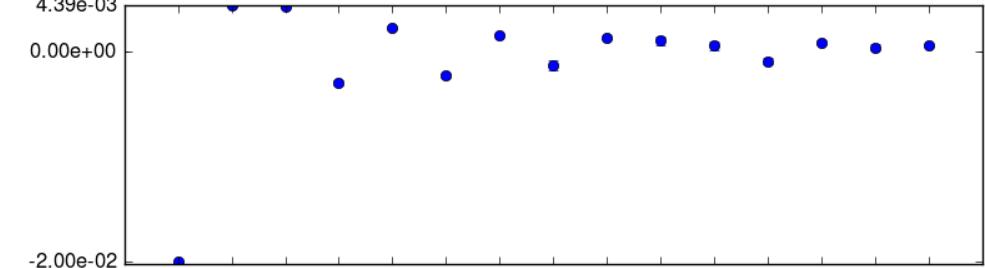


(1, 2)

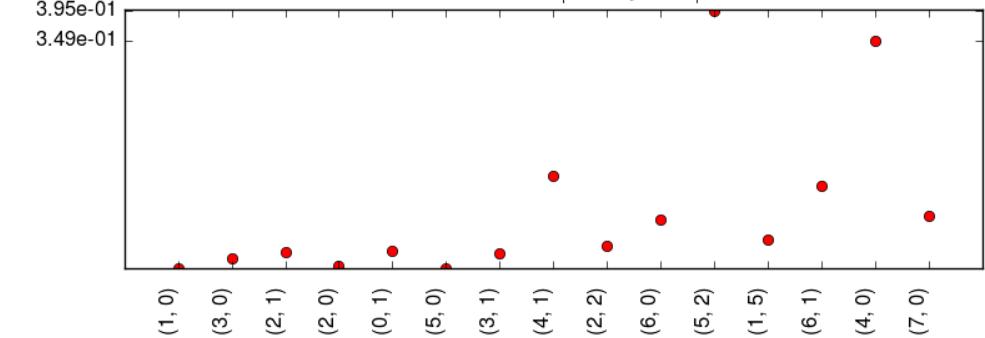
(0, 6)

(0, 5)

Scatter plot of  $\langle N.C_i \rangle$  for 15 biggest  $\langle O.C_i \rangle$  coeffs



Scatter plot of  $\left| \frac{\langle N.C_i \rangle}{\langle O.C_i \rangle} - 1 \right|$



(1, 0)

(3, 0)

(2, 1)

(2, 0)

(0, 1)

(5, 0)

(3, 1)

(4, 1)

(2, 2)

(6, 0)

(5, 2)

(1, 5)

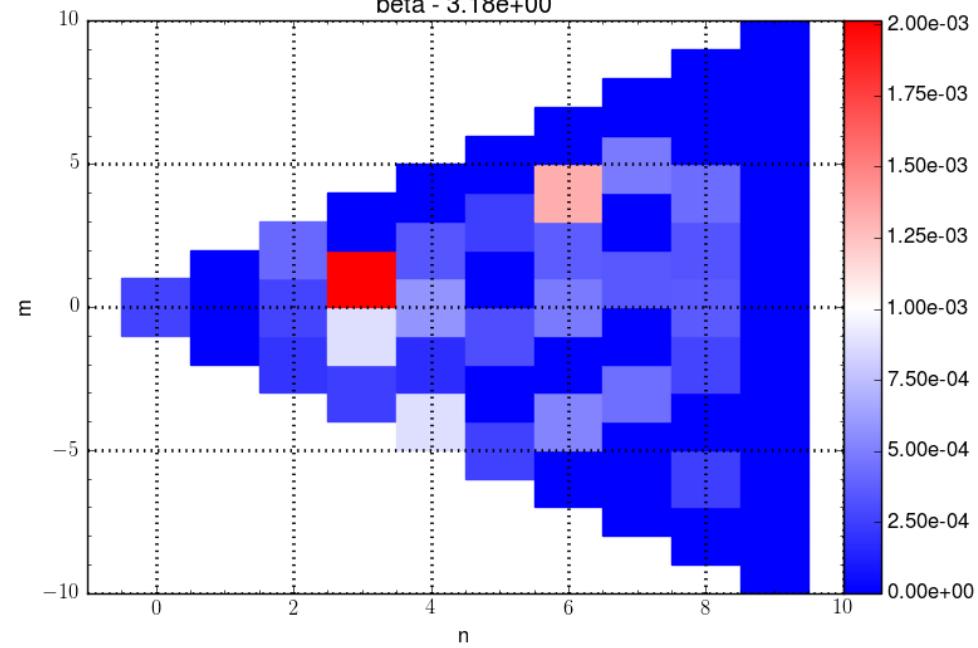
(6, 1)

(4, 0)

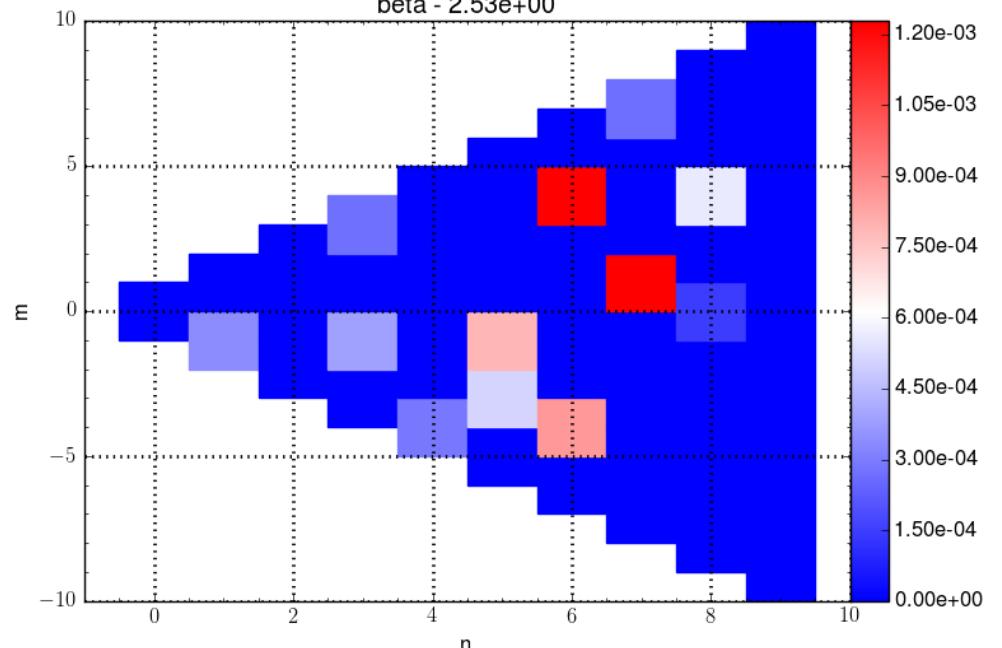
(7, 0)

# Stability tests

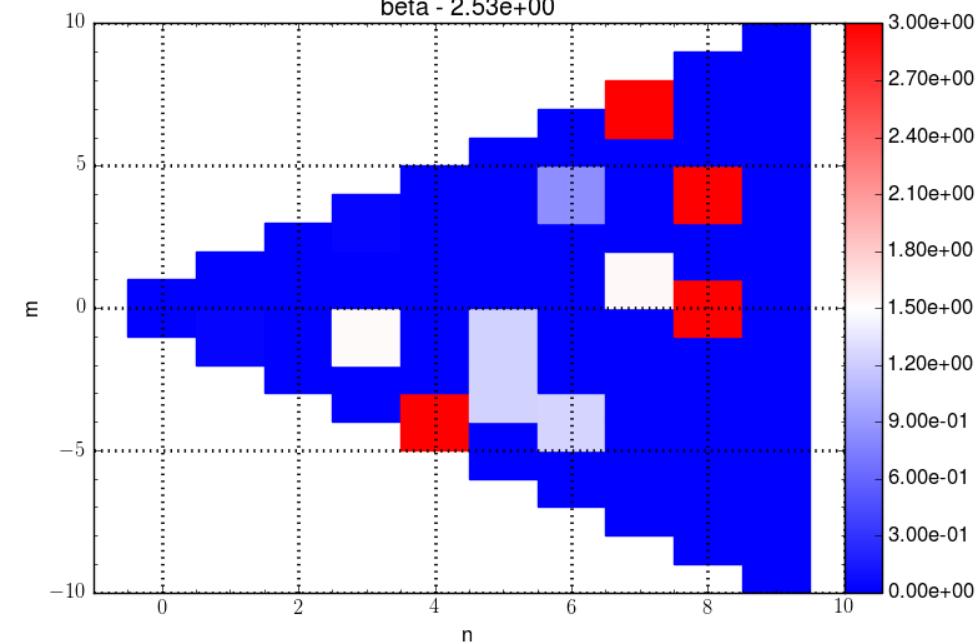
$\sigma$  matrix  $\sigma(N.C_i)$   
 S/N =  $5.104e+01$   
 beta -  $3.18e+00$



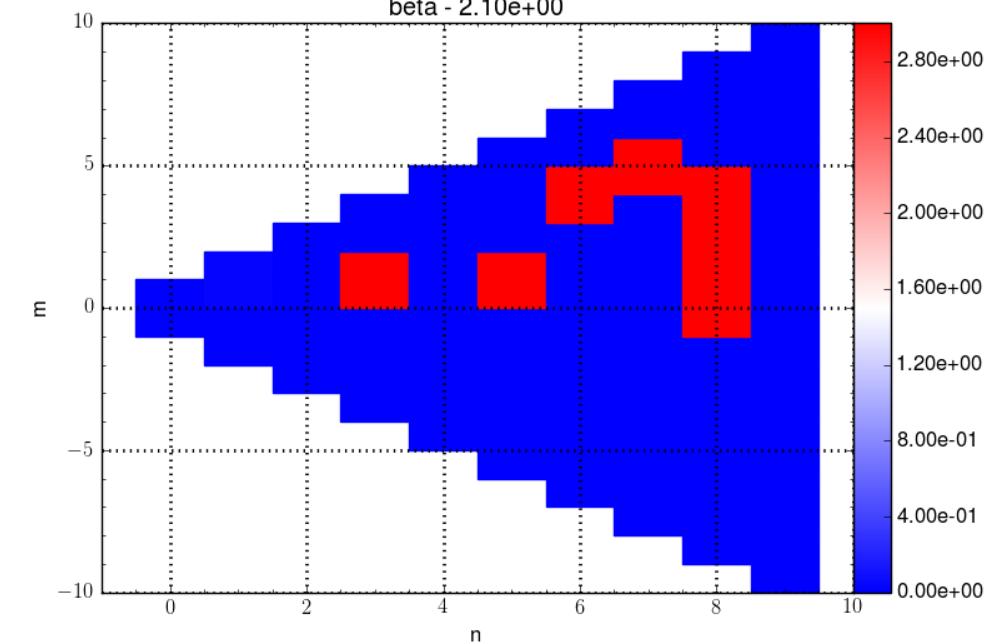
$\sigma$  matrix  $\sigma(N.C_i)$   
 S/N =  $5.104e+01$   
 beta -  $2.53e+00$



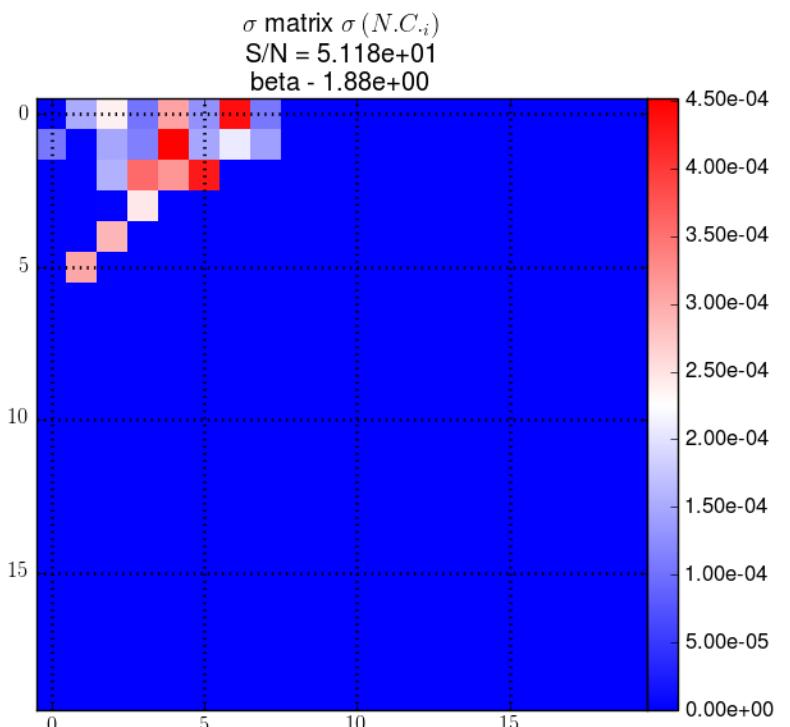
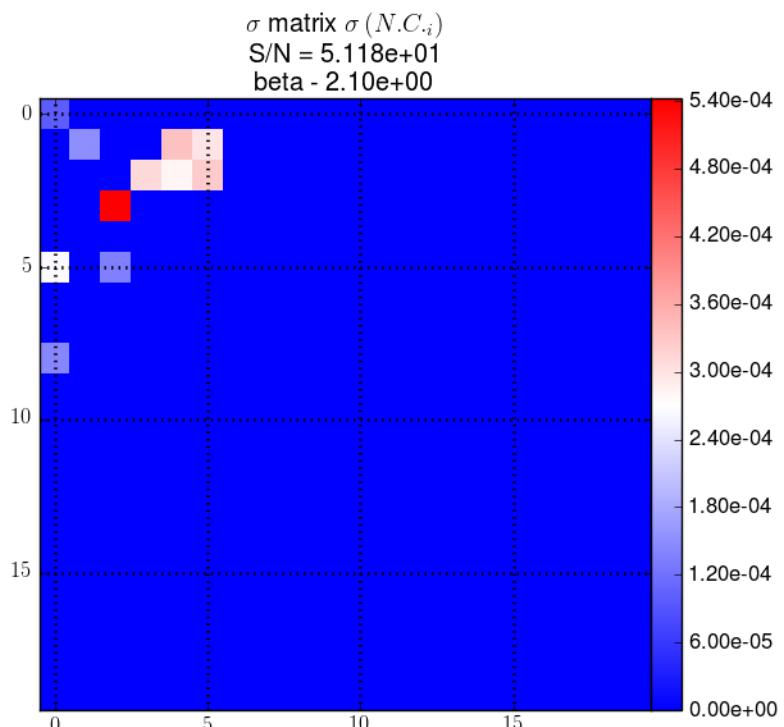
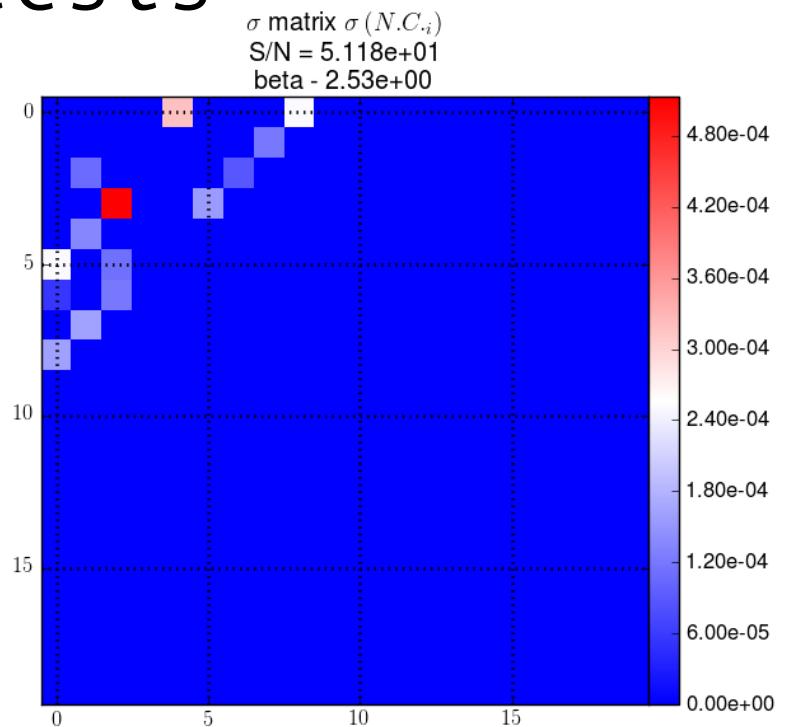
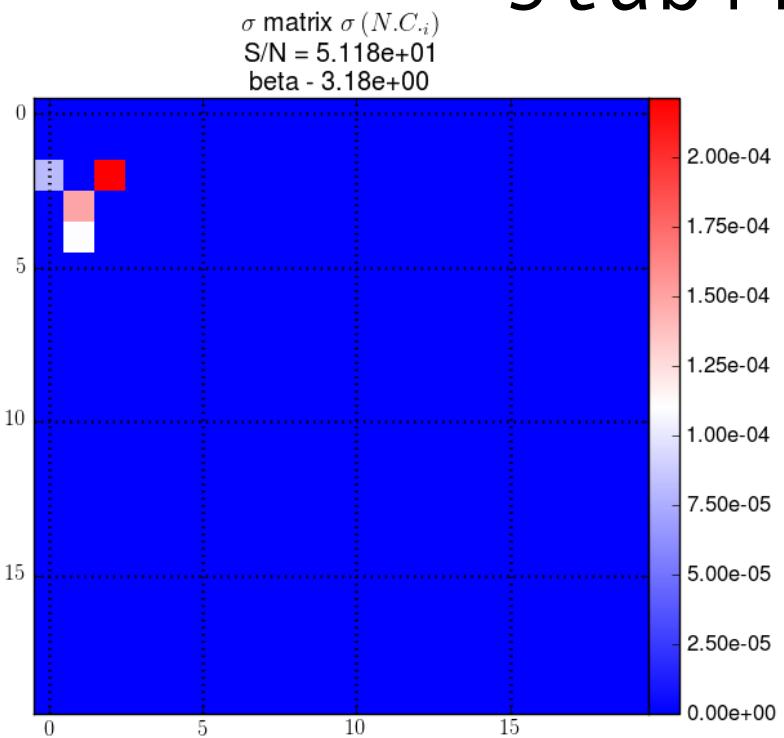
Rel.  $\sigma$  matrix  $\sigma(N.C_i) / |\langle N.C_i \rangle|$   
 S/N =  $5.104e+01$   
 beta -  $2.53e+00$



Rel.  $\sigma$  matrix  $\sigma(N.C_i) / |\langle N.C_i \rangle|$   
 S/N =  $5.104e+01$   
 beta -  $2.10e+00$



# Stability tests



# A little bit more on the algorithms used

**Task:** Approximate the solution of  $(P_0)$ :  $\min_{\mathbf{x}} \|\mathbf{x}\|_0$  subject to  $\mathbf{Ax} = \mathbf{b}$ .

**Parameters:** We are given the matrix  $\mathbf{A}$ , the vector  $\mathbf{b}$ , and the error threshold  $\epsilon_0$ .

**Initialization:** Initialize  $k = 0$ , and set

- The initial solution  $\mathbf{x}^0 = \mathbf{0}$ .
- The initial residual  $\mathbf{r}^0 = \mathbf{b} - \mathbf{Ax}^0 = \mathbf{b}$ .
- The initial solution support  $S^0 = Support\{\mathbf{x}^0\} = \emptyset$ .

**Main Iteration:** Increment  $k$  by 1 and perform the following steps:

- **Sweep:** Compute the errors  $\epsilon(j) = \min_{z_j} \|\mathbf{a}_j z_j - \mathbf{r}^{k-1}\|_2^2$  for all  $j$  using the optimal choice  $z_j^* = \mathbf{a}_j^T \mathbf{r}^{k-1} / \|\mathbf{a}_j\|_2^2$ .
- **Update Support:** Find a minimizer,  $j_0$  of  $\epsilon(j)$ :  $\forall j \notin S^{k-1}$ ,  $\epsilon(j_0) \leq \epsilon(j)$ , and update  $S^k = S^{k-1} \cup \{j_0\}$ .
- **Update Provisional Solution:** Compute  $\mathbf{x}^k$ , the minimizer of  $\|\mathbf{Ax} - \mathbf{b}\|_2^2$  subject to  $Support\{\mathbf{x}\} = S^k$ .
- **Update Residual:** Compute  $\mathbf{r}^k = \mathbf{b} - \mathbf{Ax}^k$ .
- **Stopping Rule:** If  $\|\mathbf{r}^k\|_2 < \epsilon_0$ , stop. Otherwise, apply another iteration.

**Output:** The proposed solution is  $\mathbf{x}^k$  obtained after  $k$  iterations.

# A little bit more on the algorithms used

