

Title

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1 Objectives

For LCDM, interacting models, and CPL, calculate

- ξ range for varying EoS while fixing Ωm_0
- ξ range for varying Ωm_0 or r , while fixing ω
- Does $\xi < 0$ means energy transfer to dark energy in this method?

2 Background

Deceleration parameter reads

$$q(z) = -1 + \frac{1+z}{H} \frac{dH}{dz} \quad (1)$$

For interaction models, the Friedmann equations,

$$\dot{\rho}_c + 3H\rho_c = Q_c \quad (2a)$$

$$\dot{\rho}_d + 3H(1+w)\rho_d = -Q_c \quad (2b)$$

$Q_c = \xi H \rho_c$ Background equations,

$$\Omega m = \Omega m_0 (1+z)^{3-\xi} \quad (3a)$$

$$\Omega d = (\Omega d_0 + \frac{\xi}{3w+\xi} \Omega m_0) (1+z)^{3(1+w)} + \frac{-\xi}{\xi+3w} \Omega m = \Omega \bar{d}_0 (1+z)^3 + \frac{-\xi}{\xi+3w} \Omega m \quad (3b)$$

$Q_c = \xi H \rho_d$

$$\Omega m = (\Omega m_0 + \frac{\xi}{\xi+3w} \Omega d_0) (1+z)^3 + \frac{-\xi}{\xi+3w} \Omega d = \omega \bar{m}_0 (1+z)^3 + \frac{-\xi}{\xi+3w} \Omega d \quad (4a)$$

$$\Omega d = \Omega d_0 (1+z)^{3(1+w)+\xi} \quad (4b)$$

Eqn 3 and eqn 4 shows that the coupling constant has two effects,

1. Change the amplitude of the evolution of matter or dark energy energy density.
2. Transfer energy between DE and DM.

2.1 Some definitions

1. For short

$$r = \frac{\Omega_{m0}}{\Omega_{d0}}$$

3 Data & Method

3.1 Data

LCDM Parameters From WMAP, $\Omega_{m0} = 0.265$

Constraints $\Omega_{m0} = 0.247(+0.013, -0.013)$; Transition redshift $0.426(+0.082, -0.050)$.(arXiv:1205.4688, arXiv:astro-ph/0611572).

In $(\Omega_{m0}, \text{Transition redshift})$ plane, allowed region is a rectangle centred at $(0.274, 0.426)$ with two diagonal points $(0.261, 0.376)$ and $(0.287, 0.508)$.

CPL $\Omega_{m0} = 0.269(+0.017, -0.008)$, $w_0 = -0.97(+0.12, -0.07)$, $w_1 = 0.03(+0.26, -0.75)$

4 Results

Check the files in files folder.

INTRO

PRE

LCDM Model

Interacting Models

Summary

■ Interacting models

□ List of what to make clear

□ BASIC

Evolution of energy density for $Q_c = \xi H \rho_c$, constant ξ , constant w , and $\xi \neq -3w$

$$\Omega_m = \Omega_{m0} (1+z)^{3-\xi}$$

$$\Omega_d = \left(\Omega_{d0} + \frac{\xi}{3w + \xi} \Omega_{m0} \right) (1+z)^{3(1+w)} + \frac{-\xi}{3w + \xi} \Omega_m \equiv \Omega_{d0}' (1+z)^{3(1+w)} + \frac{-\xi}{3w + \xi} \Omega_m$$

Evolution of energy density for $Q_c = \xi H \rho_d$, constant ξ , constant w , and $\xi \neq -3w$

$$\Omega_m = \left(\Omega_{m0} + \frac{\xi}{\xi + 3w} \Omega_{d0} \right) (1+z)^3 + \frac{-\xi}{\xi + 3w} \Omega_d \equiv \Omega_{m0}' (1+z)^3 + \frac{-\xi}{\xi + 3w} \Omega_d$$

$$\Omega_d = \Omega_{d0} (1+z)^{3(1+w)+\xi}$$

So in the two cases, coupling constant has two effects:

1. Amplifies the curve of deceleration parameter.
2. Energy flow between DE and DM.

■ Interacting model $Q_c = \xi H \rho_c$ with constant ξ and constant EoS w .

Derived from (transition redshift, Ω_{m0}) plane, the allowed region for coupling constant ξ is $(-1.28, -0.46)$ with a center at -0.88 , i.e., $-0.88^{+0.42}_{-0.40}$, taken the case that the universe is flat, and choose the EoS parameter $\{w=-1\}$.

Derived from the (transition redshift, $\frac{\Omega_{m0}}{\Omega_{d0}}$) plane, the allowed region of coupling constant ξ is $(-1.25, -0.47)$ with a center at -0.88 , i.e., $-0.88^{+0.41}_{-0.37}$.

There is a bit difference between the two answers. One possible reason is the second method doesn't assume a flat universe, while the first one supposes the universe is flat.

The full table of fitting results are shown below. The light purple element are the final results.

In[315]:=

tabξFinaltICC

$Q_c = \xi H \rho_c$, constant ξ , constant $w = -1$: Results for ξ			
Ω_{m0}/Ω_{d0} : Transition	$z_t = 0.376$	$z_t = 0.426$	$z_t = 0.508$
$r = 0.358$	-1.25282	-0.965436	-0.617444
$r = 0.378$	-1.15011	-0.875189	-0.542347
$r = 0.398$	-1.05453	-0.791252	-0.472561

Out[315]=

To check the consistency of the two methods ((Transition, Ω_{m0}) plane fitting and (Transition, Ω_{m0}/Ω_{d0}) plane fitting), we find out the fitting results of coupling constant ξ for a flat universe, i.e., $r = \frac{\Omega_{m0}}{1-\Omega_{m0}}$ in the (Transition, Ω_{m0}) plane, applying the data from (Transition, Ω_{m0}/Ω_{d0}) plane.

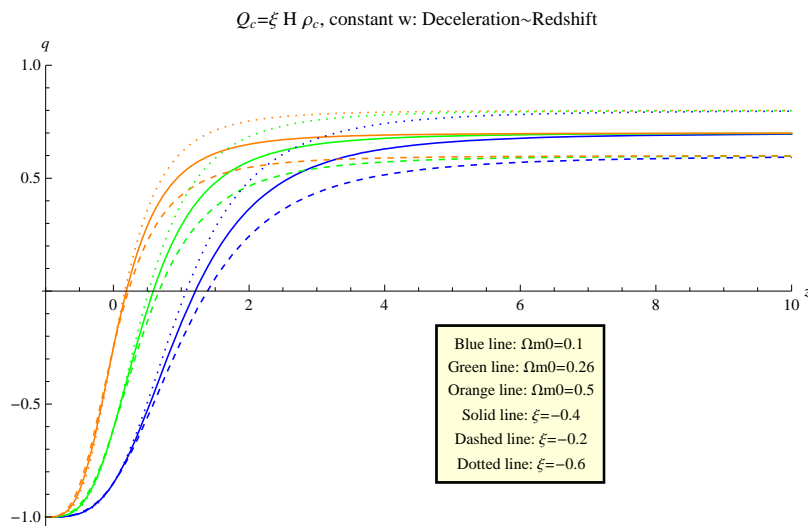
By solving out Ω_{m0} , we get $\Omega_{m0} = \frac{r}{1+r}$ (this is a monotonic function) in this case. Thus if we use the constrain that $r \in (0.358, 0.398)$ with a center value 0.378, the value of Ω_{m0} is (0.263623, 0.284692), centered at 0.274311. Use this set of value of Ω_{m0} as the constrain, we have the fitting results in (Transition, Ω_{m0}) plane, which is $-0.88^{+0.41}_{-0.37}$. This result is exactly the same as the result directly derived from (transition redshift, $\frac{\Omega_{m0}}{\Omega_{d0}}$) plane. The same has been done to $Q_c = \xi H \rho_d$ with ξ constant and w constant model, and the result is that the two methods are also consistent.

The plots of deceleration parameter are shown below. At the limit $z \rightarrow \infty$, the deceleration parameter is degenerate for different Ω_{m0} in this constant ξ and constant w model.

Theoretically, this limit is determined by the interaction coupling constant ξ , which is $\frac{(1-\xi)}{2}$, with $3w + \xi < 0$.

In[316]:=

pldecICCSum



Out[316]=

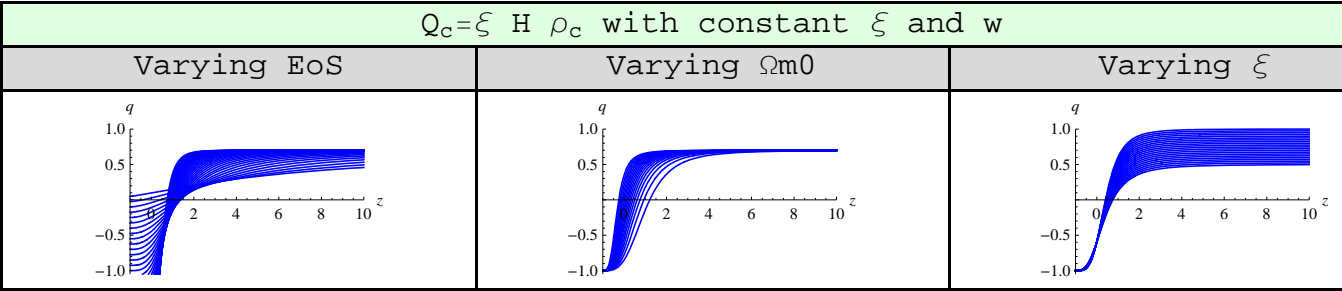
Check the effect of different parameters on deceleration parameter.

Interaction ξ changes the value of deceleration parameter at $z \rightarrow \infty$ limit. EoS changes the the whole shape. Matter fraction determines how fast q varies, but just in a small time scale.

In[317]:=

varyingICCSum

Out[317]=

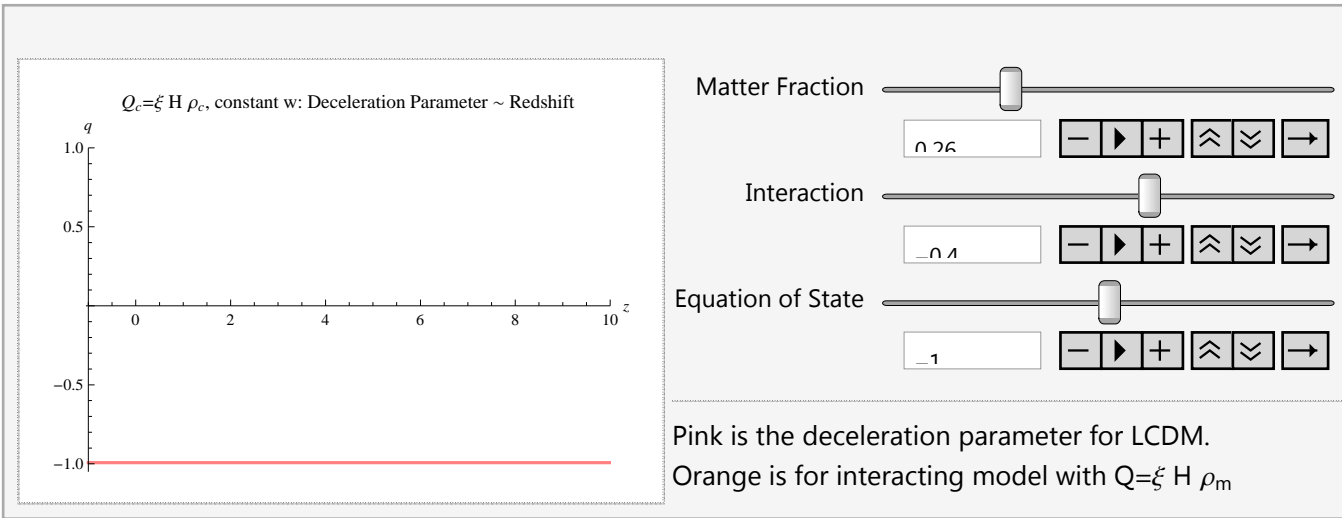


A toy to play with is also provided. Slide the bars to view the effects of different parameters on deceleration parameter.

In[318]:=

pldecICCSum

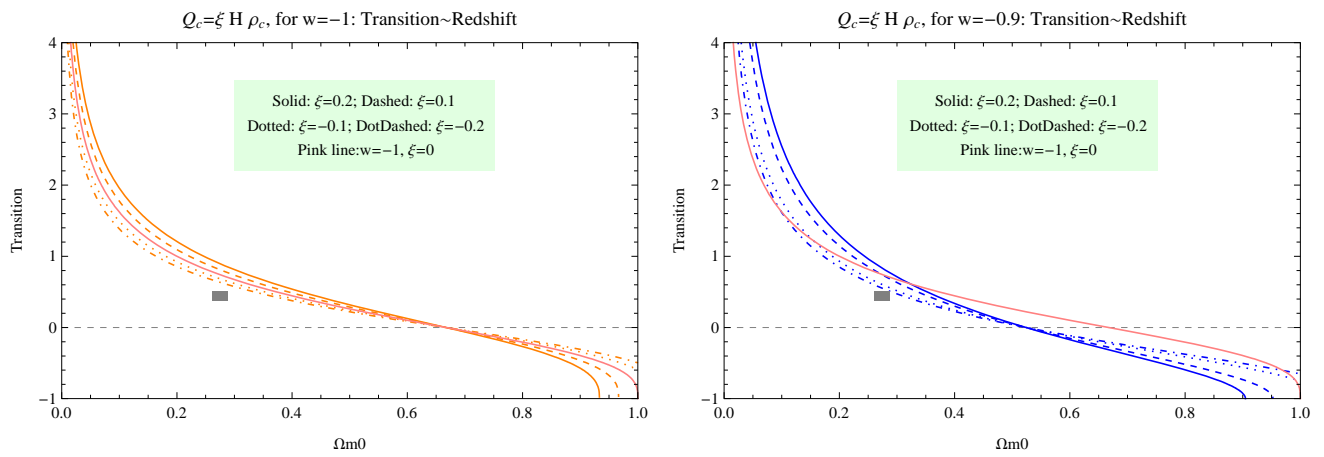
Out[318]=



It can be inferred from the expression for Ω_m and Ω_d that if the transition happens before $z=0$, increasing coupling ξ will bring forward the transition and if it happens after $z=0$, increasing coupling ξ will delay the emergence of transition. The following figure shows this result. Gray rectangle is the region given by Riess (References, Data From, 2).

- Orange for $w=-1$
Blue for $w=-0.9$
- Solid line: $\xi=0.2$
Dashed line: $\xi=0.1$
Dotted line: $\xi=-0.1$
DotDashed line: $\xi=-0.2$
- Pink solid line: $w=-1, \xi=0$

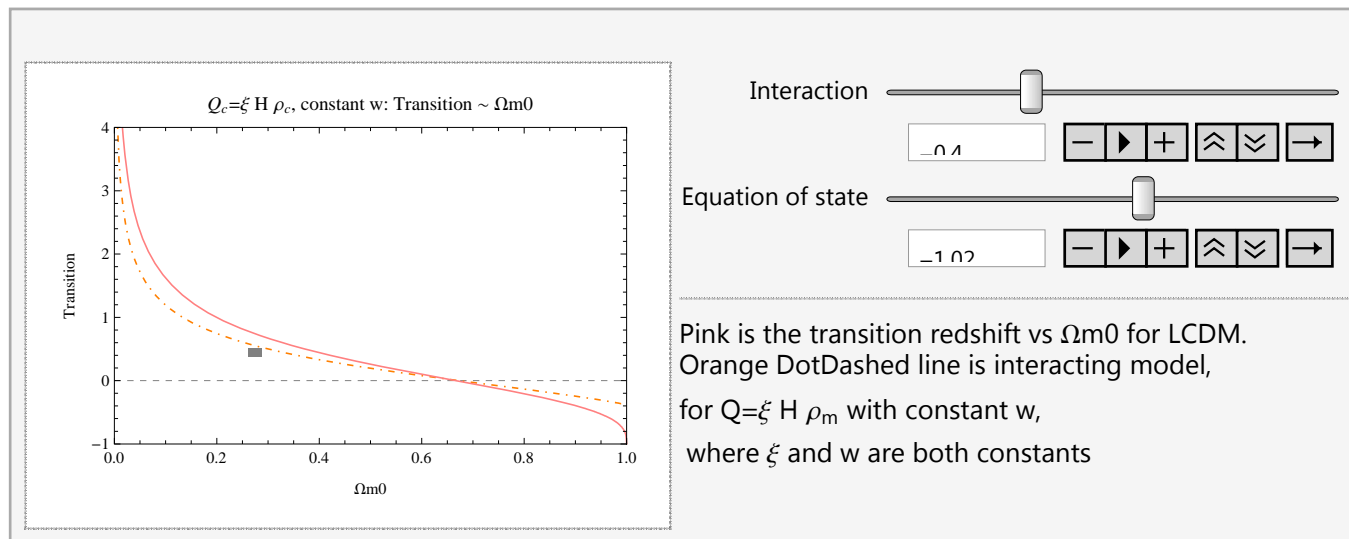
In[319]:=

plztrvsΩm0ICCSum

Out[319]=

This can also be seen clearly from the following toy. Gray rectangle is the region given by Riess (References, Data From, 2) .

In[320]:=

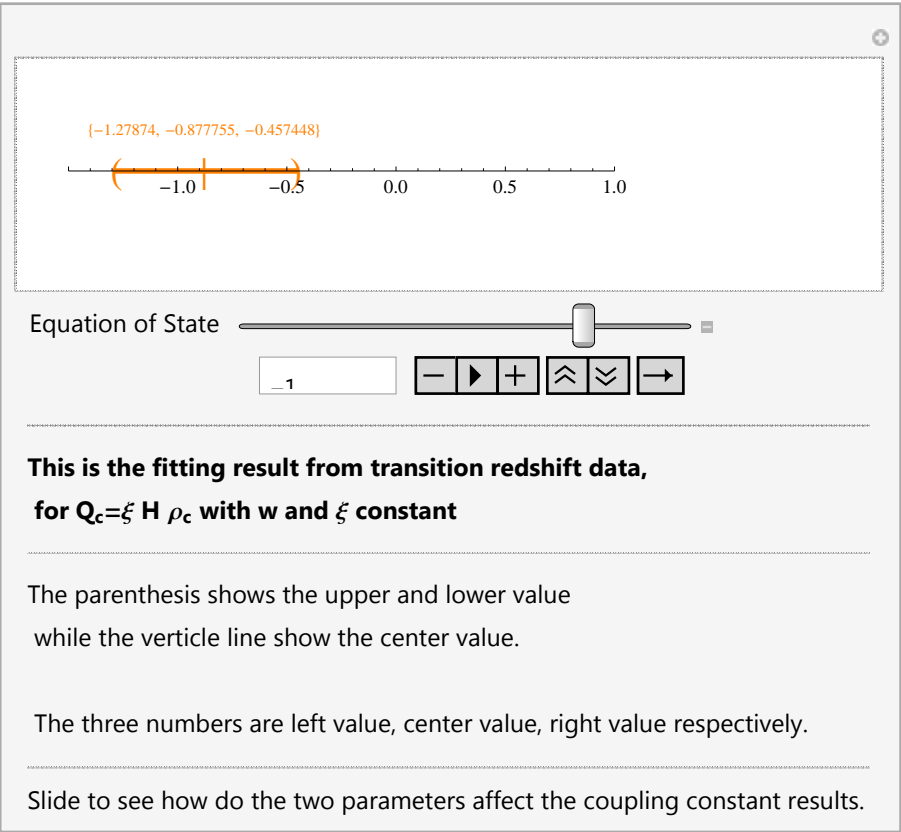
plztrICCMansum

Out[320]=

The fitting results of coupling constant ξ can also be dynamic.

In[321]:=

fitξICCMansum



Out[321]=

For different constant EoS, the fitting results using $\Omega_{m0} \in (0.261, 0.287)$ with a center value 0.274 and Transition redshift $\in (0.376, 0.508)$ with a center value 0.426. When EoS is very small, the line crosses zero. But that is not so useful.

Some data points are derived using $w = -1.087 \pm 0.096$ (from Reference, Data From, 3).

In[322]:=

tabξvwExamICC

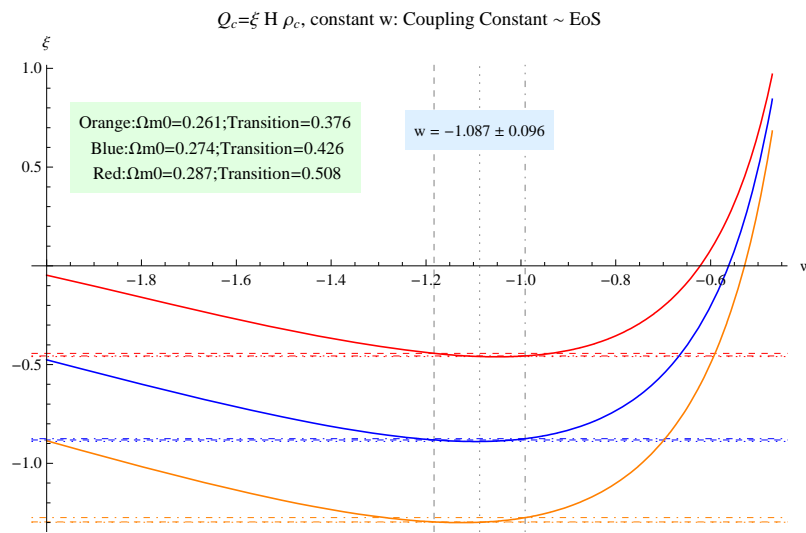
ξ results for $Q_c = \xi H \rho_c$ (Fitting data: Data From, 2)			
w	Center	Lower	Upper
-1.183	-0.881565	-1.29687	-0.443589
-1.087	-0.88948	-1.29859	-0.459135
-0.991	-0.875238	-1.27522	-0.456176

Out[322]=

A plot showing these data points and the curves of $\xi \sim w$.

In[323]:=

pltξvwExamICC



Out[323]=

Or just casually use the following parameters.

(-1 within 5%: (-1.05, -0.95))

$w=-1$ (-1.279, -0.457) center: -0.878

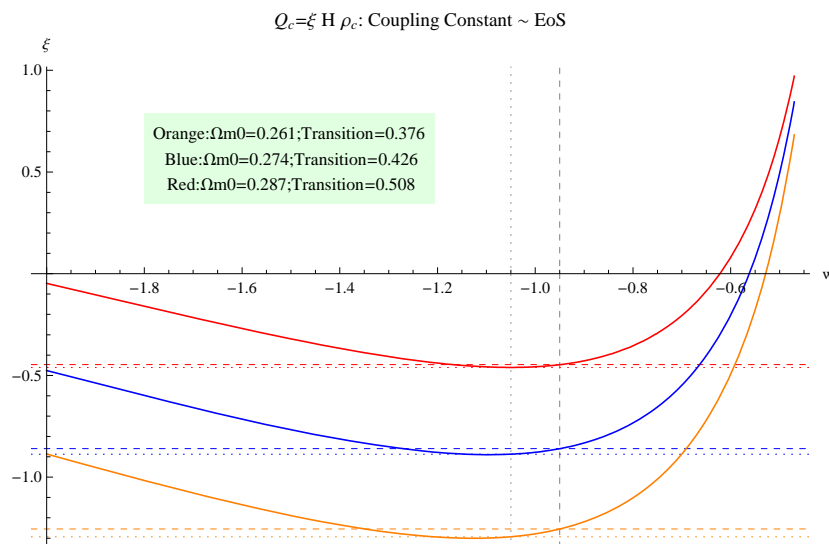
$w=-1.05$ (-1.293, -0.461) center: -0.887

$w=-0.95$ (-1.255, -0.447) center: -0.860

The following graph show how do ξ changes with EoS. The grid lines are the results of $w = -1 \pm 0.05$. Two verticle lines are -1.05 and -0.95 respectively. Horizontal lines are their intersections with the $\xi \sim w$ lines. EoS does not monotonically change ξ . And the minima of these line occurs at a larger w with an increasing Ω_{m0} .

In[324]:=

pltξvwExam1ICC



Out[324]=

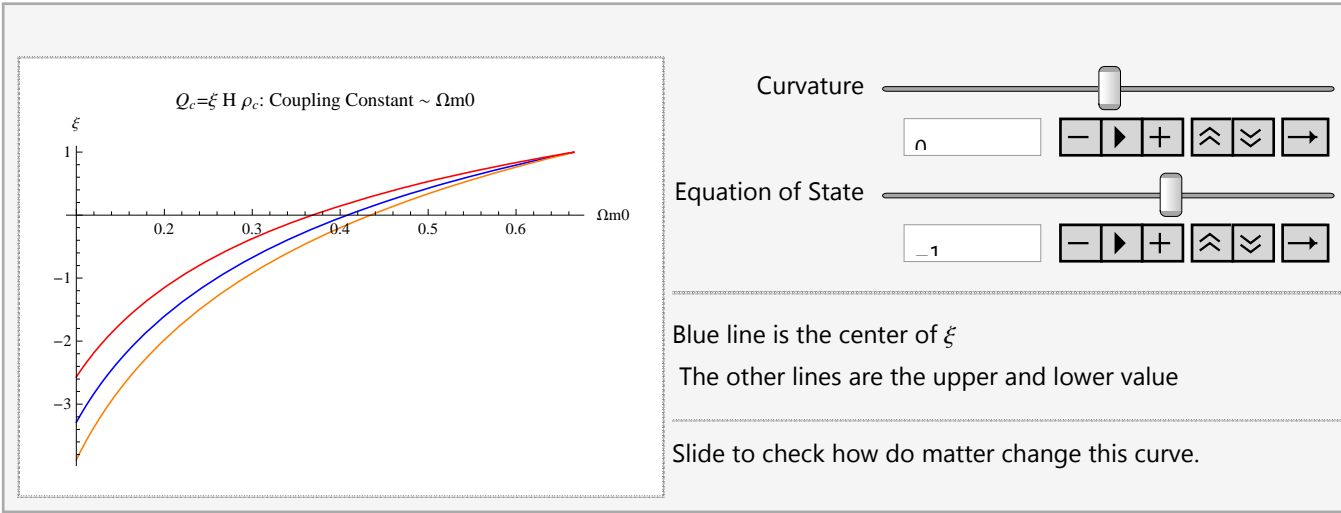
Now we assume we do not have the observed Ω_{m0} data, how do this Ω_{m0} change the result of ξ . In other words, if the observed Ω_{m0} data float around some value, then how is the fitting result? We also consider the curvature.

In the figure below, it seems that there is a point where three lines converge. This has something to do with the phenomena that

In[325]:=

```
pltξvΩm0ICCMansum
```

Out[325]=



Some data for flat Λ CDM universe. The following data shows how Ω_{m0} change our results for ξ if we already have transition redshift data {0.426|0.376,0.508}.

If Ω_{m0} varies 5 percent from 0.274,

In[326]:=

```
tabξICCSum
```

Out[326]=

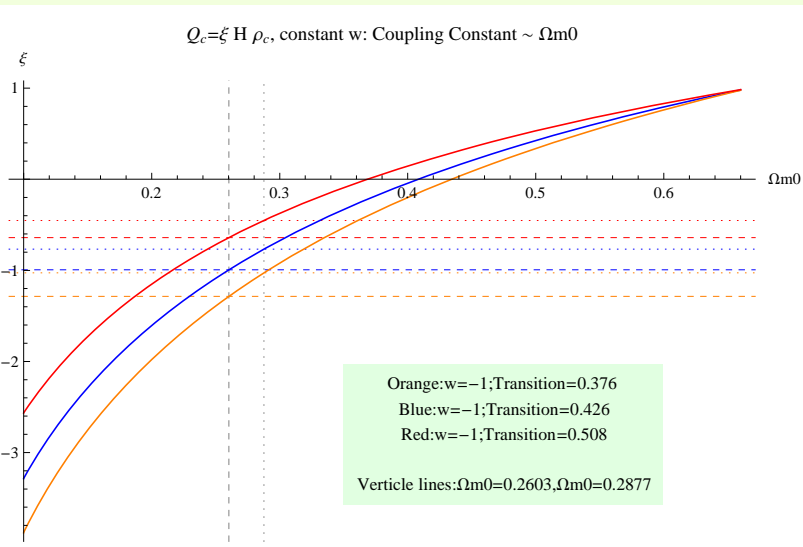
For $\Omega_{m0} \in 0.274 (1 \pm 0.05)$			
Table of ξ for different Ω_{m0} ~Transition combination			
Ω_{m0} ~Transition	0.426	0.376	0.508
0.2603	-0.994339	-1.28571	-0.641508
0.274	-0.877755	-1.15303	-0.544482
0.2877	-0.767582	-1.02756	-0.452892

Monotonic line.

In[327]:=

```
pltξvΩm0ICCSum
```

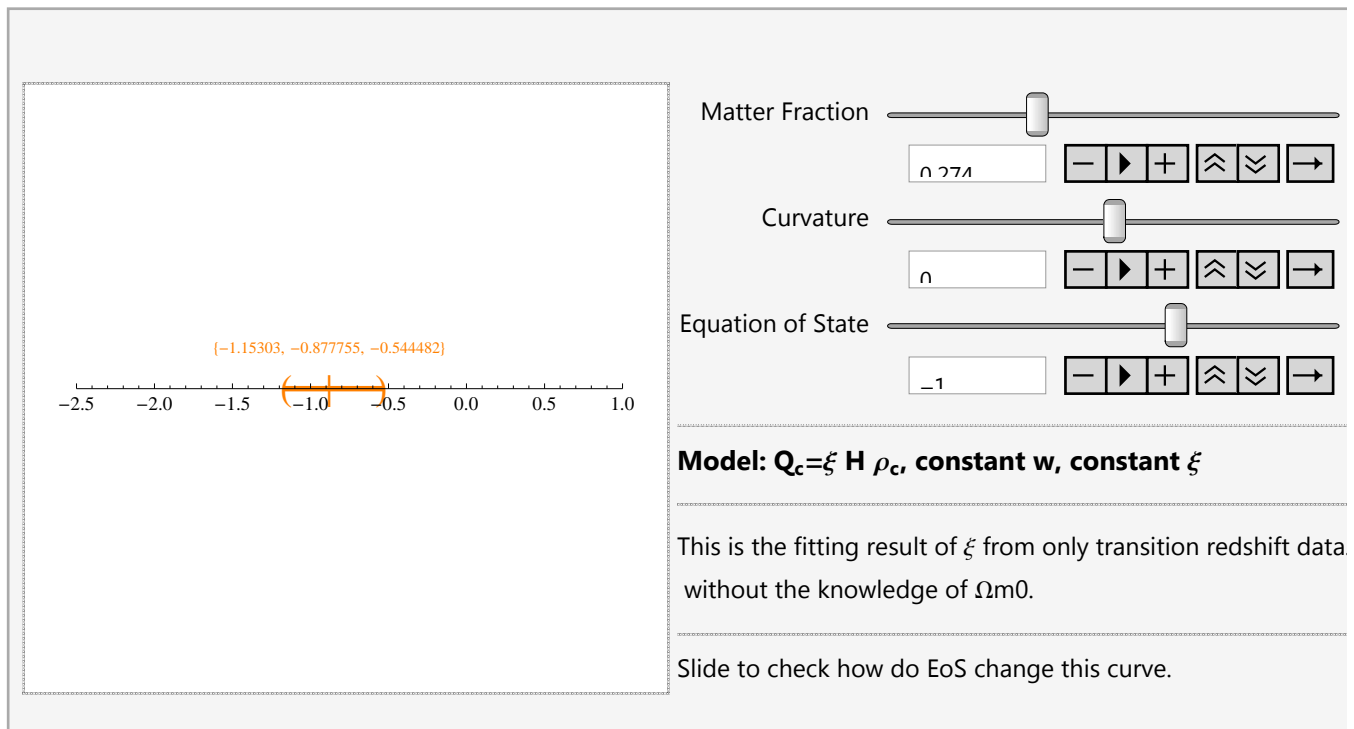
Out[327]=



Besides Ω_{m0} , we can also find out the effects of Curvature, EoS. Assuming we have a constrain of Transition redshift (0.376,0.508) with a center at 0.426.

In[328]:=

fitξ2ICCMansum



Out[328]=

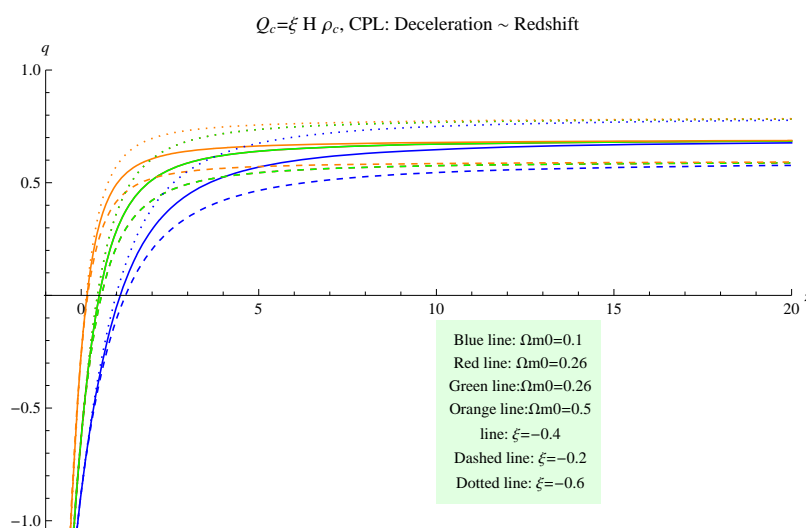
- Interacting model $Q_c = \xi H \rho_c$ with constant ξ and CPL parameterized EoS $w = w_0 + w_1 \frac{z}{1+z}$.

For a flat universe, choose the parameters $\{w_0 = -1.02, w_1 = 0.6\}$, the region for interaction constant ξ should be $(-1.04, -0.21)$ with a center at -0.64 , i.e., $-0.64^{+0.42}_{-0.40}$, derived from the (transition redshift, Ω_{m0}) plane, while a result of $(-1.01, -0.23)$ with a center at -0.63 , i.e., $-0.63^{+0.40}_{-0.38}$, derived from (transition redshift, $\frac{\Omega_{m0}}{\Omega_{d0}}$) plane.

Deceleration parameter is shown below. Behaves similar to the constant ξ constant w situation.

In[329]:=

pldecICCPShowSum

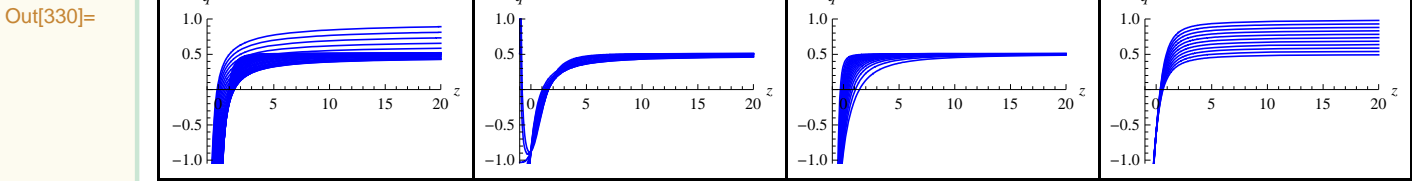


Out[329]=

The following plots show the effect of different parameters. Each plot shows how the deceleration parameter vs redshift line changes under uniformly distributed w_0, w_1, Ω_{m0} or ξ . w_0 moves the line up or down, but not monotonously; w_1 changes the late time behavior;

Ωm_0 changes the slope;
 ξ has moves the line up or down;

```
In[330]:= varyingICCPLShowSum
```

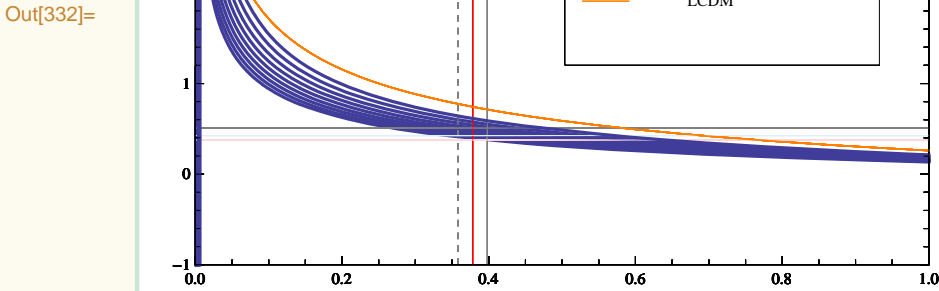


A toy to play with the $q \sim z$ plot. When $\xi=0$, $w_0=-1$, $w_1=0$, the curve reduced to LCDM curve.

```
In[331]:= pldecICCPLManSum
```

A plot shows how bad it is to use transition redshift to constrain interacting model. This is a CPL parameterized example. For $\xi \in (-0.8, 0)$, the line just stays near the allowed region constrained by Riess's results (References, Data From, 2).

```
In[332]:= pltransrICCPLDense
```

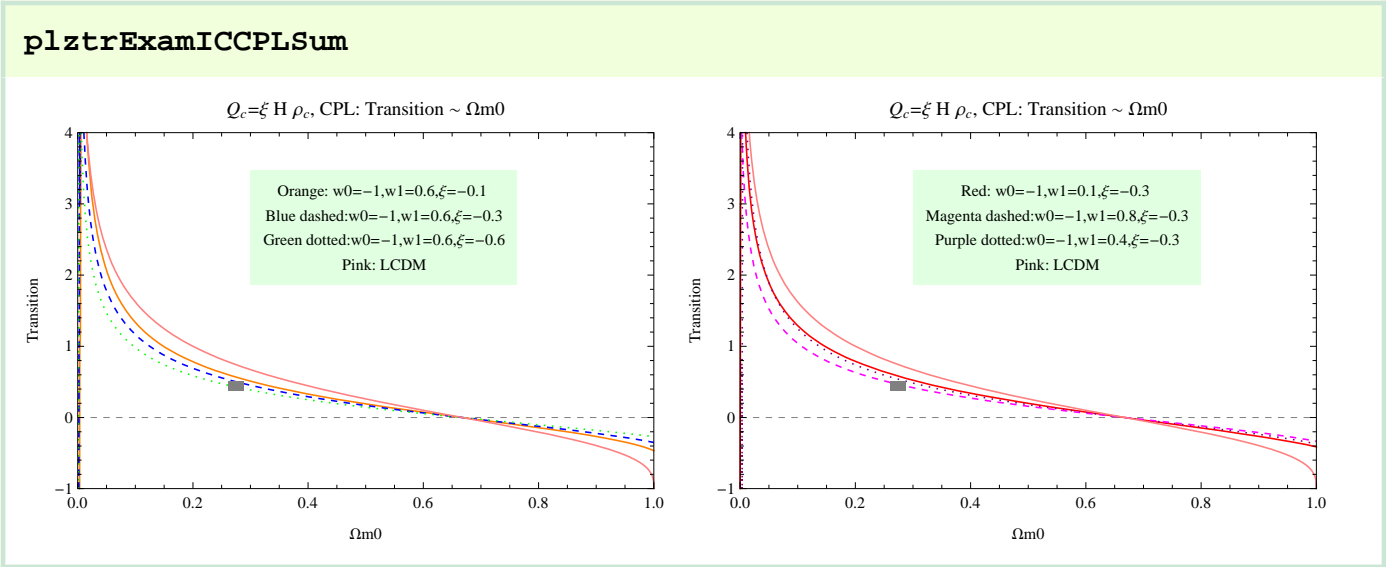


From the manipulate below, we find for $w_0=-1$, there is a point on this transition $\sim \Omega m_0$ curve do not change with coupling constant ξ and w_1 . (Well, what's the use of that...)

```
In[333]:= plztrICCPLManSum
```

An explicit proof of this statement.

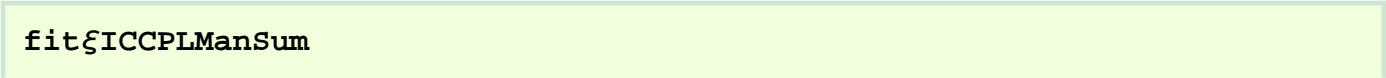
In[334]:=



Out[334]=

A manipulate of the fitting results.

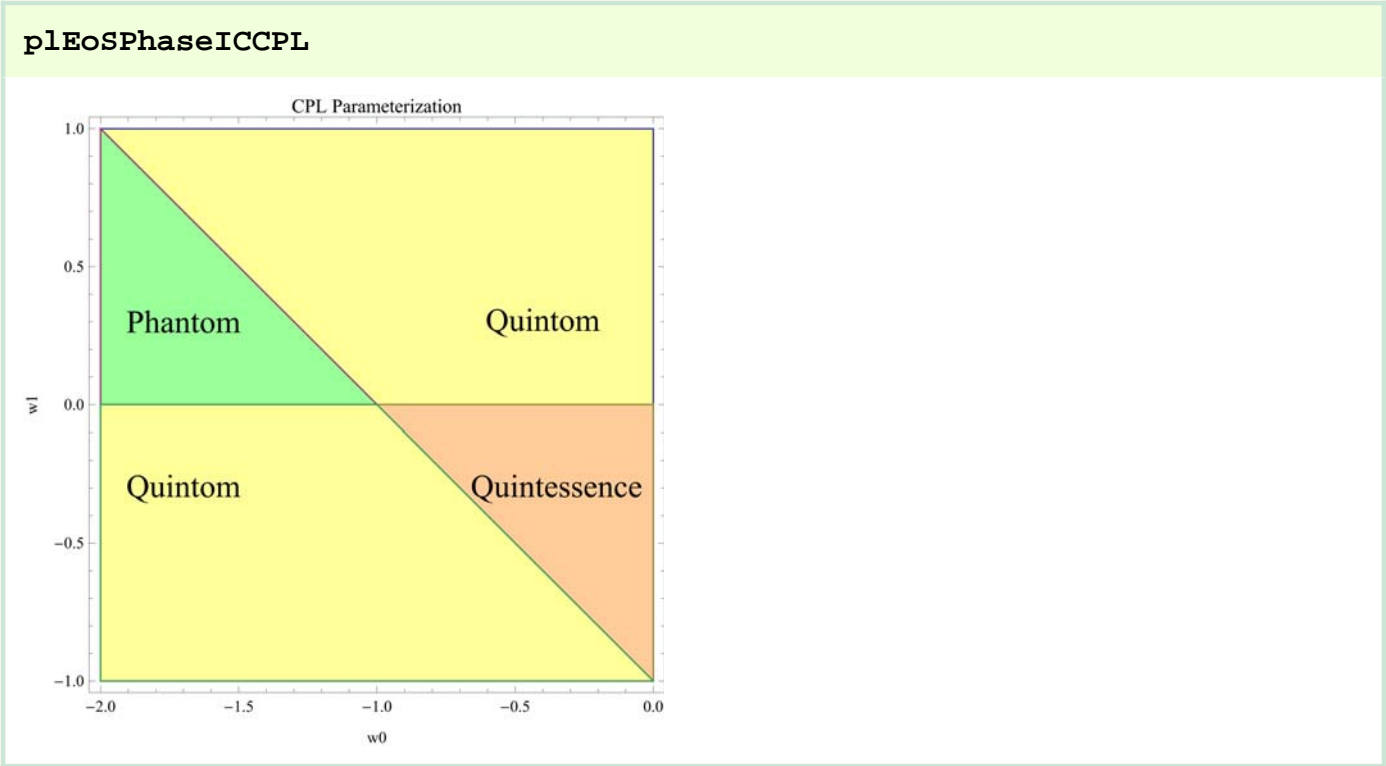
In[335]:=



Category

For different w0 and w1 in its EoS equation,

In[336]:=



Out[336]=

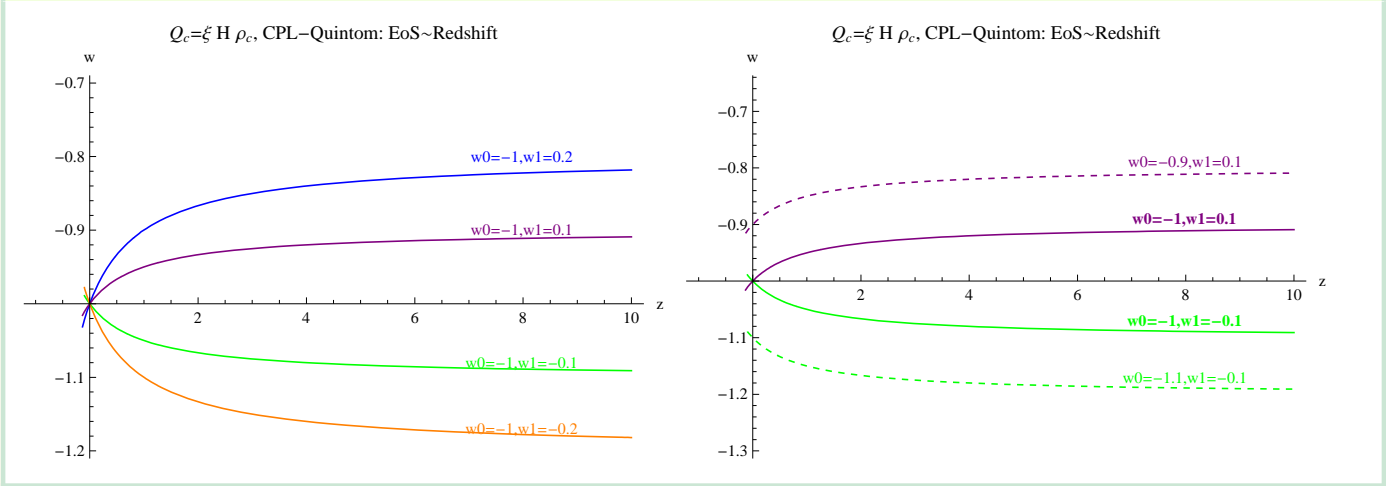
Quintom

Color illustrations for the following two figures.
“Purple line:(-1,0.1). Blue line: (-1,0.2) Green line:(-1,-0.1)\n. Pink line:LCDM”
“Purple Dashed:(-0.9,0.1) Green Dashed:(-1.1,0.1)”

In[337]:=

plEoSICCPLQuintomSum

Out[337]=

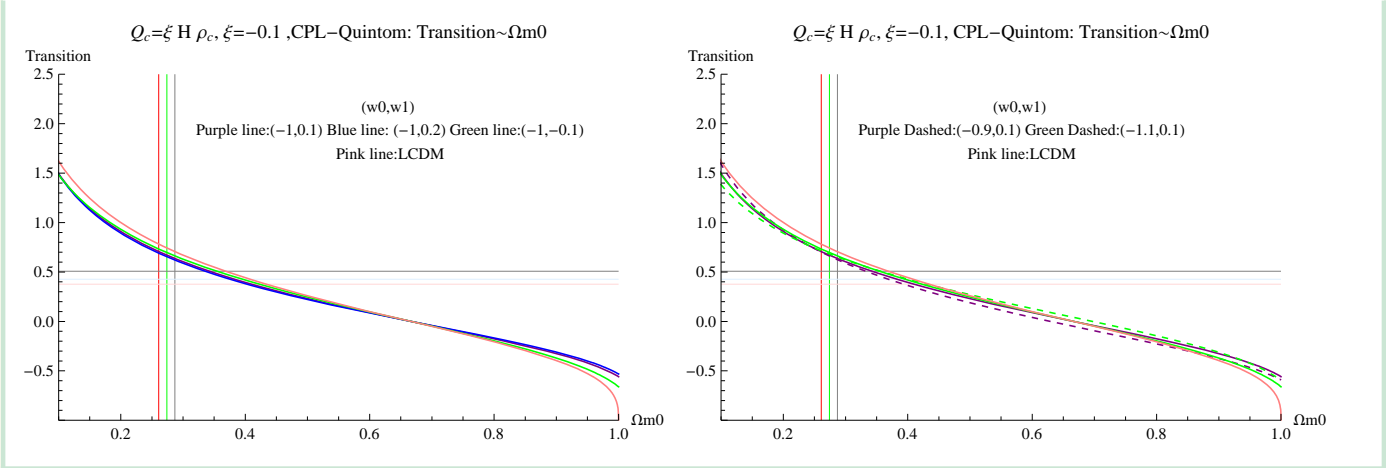


Plots of Transition redshift vs Ω_{m0} .
Legends are shown on the plots. Hard to distinguish from each other.

In[338]:=

plztrICCPLQuintomSum

Out[338]=



For **different EoS**, the fitting results are different. The following table and plots show how do w_0 and w_1 change the results.

In[339]:=

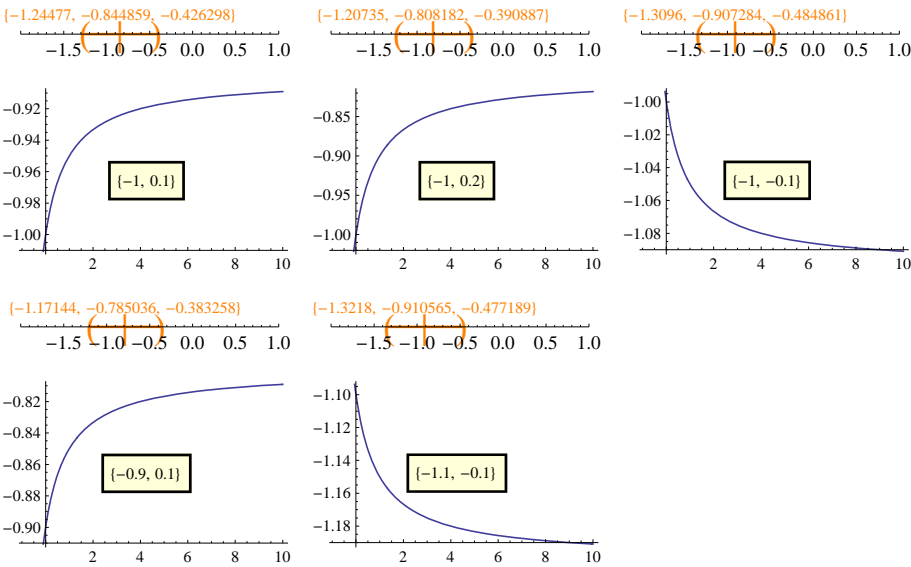
tab\xvwExamICCPLQuintom

Out[339]=

ξ results for $Q_c = \xi H \rho_d$, CPL, Quintom.			
$\{w_0, w_1\}$	Center	Lower	Upper
$\{-1, -0.1\}$	-0.907284	-1.3096	-0.484861
$\{-1, 0\}$	-0.877755	-1.27874	-0.457448
$\{-1, 0.1\}$	-0.844859	-1.24477	-0.426298
$\{-0.9, 0.1\}$	-0.785036	-1.17144	-0.383258
$\{-1.1, -0.1\}$	-0.910565	-1.3218	-0.477189

In[340]:=

pltfitξICCPLeQuintomSum

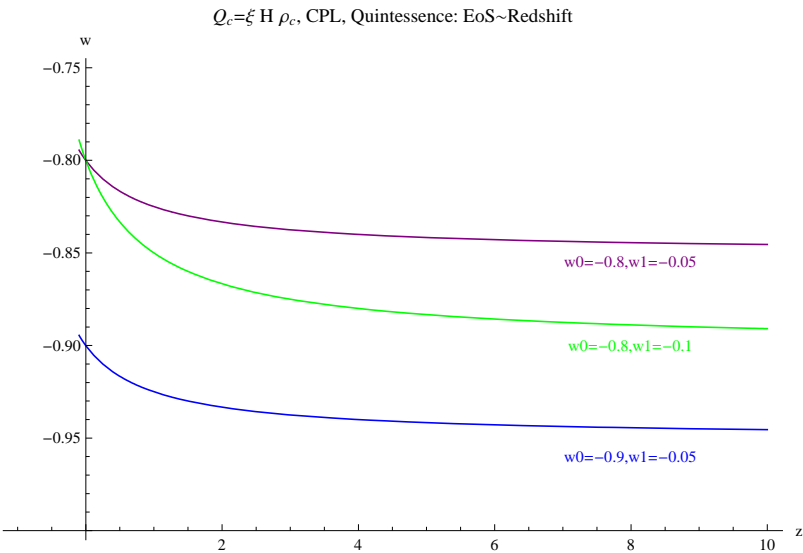


Out[340]=

▣ Quintessence

In[341]:=

plEoSICCPLeQuintessenceSum

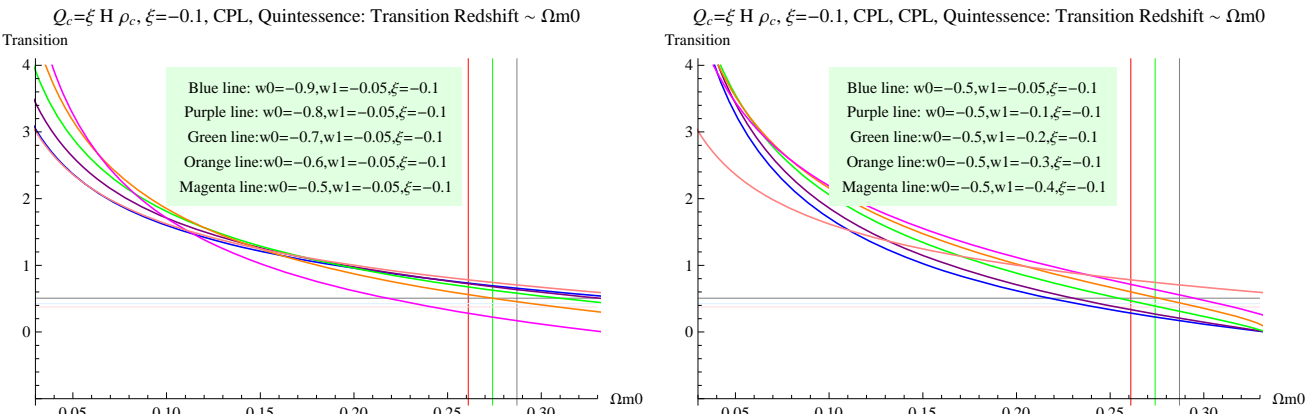


Out[341]=

Different from constant w results,

In[342]:=

plztrICCPLeQuintessenceSum



Out[342]=

Some ξ fitting results are shown below. This shows how do w_0 and w_1 change ξ results.

In[343]:=

`tabξvwExamICCPLQuintessence`

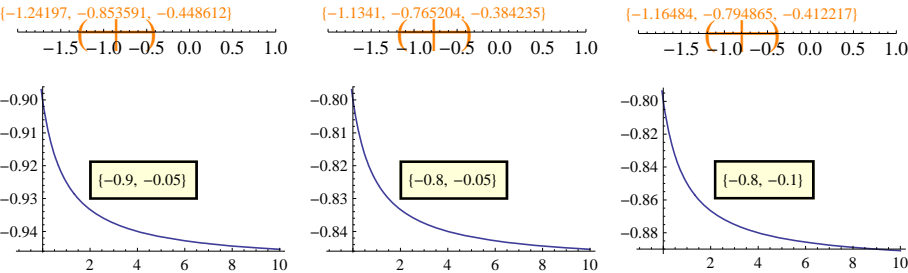
Out[343]=

ξ results for $Q_c=\xi H \rho_d$, CPL,Quintessence.			
$\{w_0,w_1\}$	Center	Lower	Upper
$\{-0.9,-0.05\}$	-0.853591	-1.24197	-0.448612
$\{-0.8,-0.05\}$	-0.765204	-1.1341	-0.384235
$\{-0.8,-0.1\}$	-0.794865	-1.16484	-0.412217

In[344]:=

`pltfitξICCPLQuintessenceSum`

Out[344]=

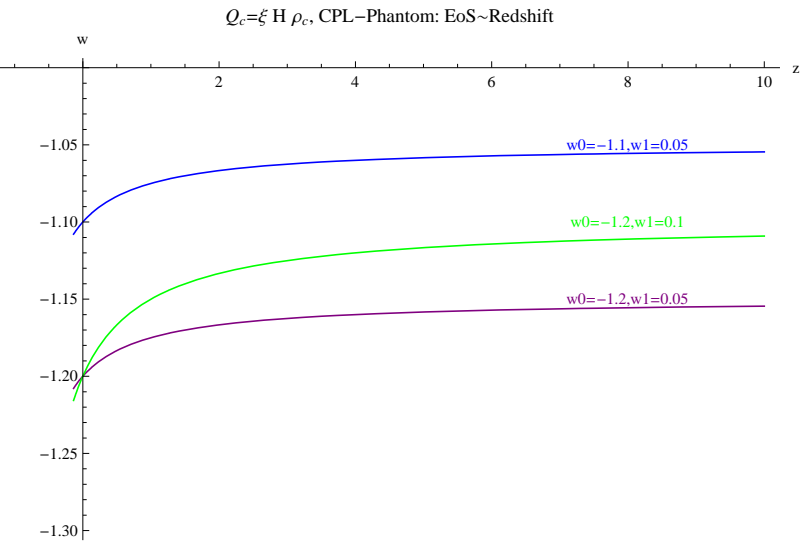


▣ Phantom

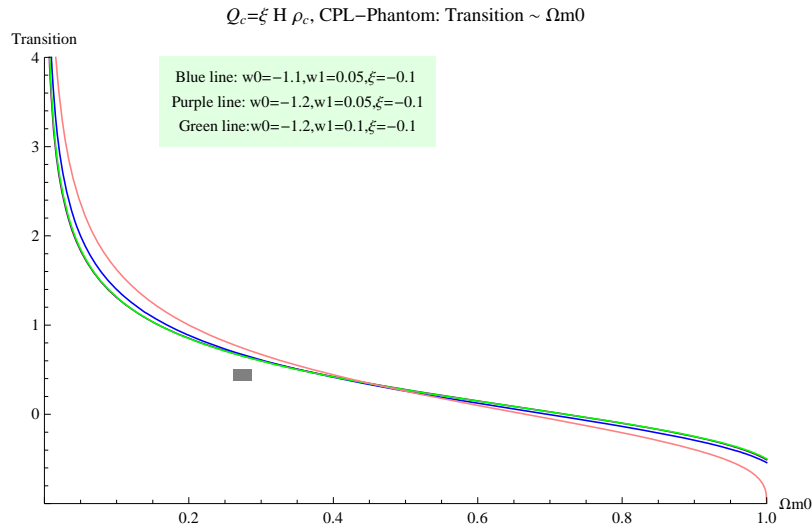
In[345]:=

`plEoSICCPLPhantomSum`

Out[345]=



In[346]:=

plztrICCPLPhantomSum

Out[346]=

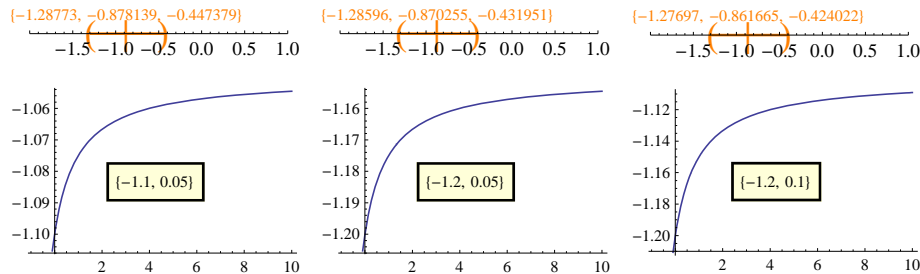
In[347]:=

tabξvwExamICCPLPhantom

ξ results for $Q_c = \xi H \rho_d$, CPL, Phantom.			
$\{w_0, w_1\}$	Center	Lower	Upper
$\{-1.1, 0.05\}$	-0.878139	-1.28773	-0.447379
$\{-1.2, 0.05\}$	-0.870255	-1.28596	-0.431951
$\{-1.2, 0.1\}$	-0.861665	-1.27697	-0.424022

Out[347]=

In[348]:=

pltfitξICCPLPhantomSum

Out[348]=

■ Interacting model $Q_c = \xi H \rho_d$ with constant ξ and constant EoS w .

Derived from (transition redshift, Ω_{m0}) plane, the allowed region for coupling constant ξ is $(-1.06, -0.42)$ with a center at -0.76 , i.e., $-0.76^{+0.34}_{-0.30}$, taken the case that the universe is flat, and choose the EoS parameter $\{w=-1\}$.

Derived from the (transition redshift, $\frac{\Omega_{m0}}{\Omega_{d0}}$) plane, the allowed region of coupling constant ξ is $(-1.07, -0.41)$ with a center at -0.76 , i.e., $-0.76^{+0.35}_{-0.31}$.

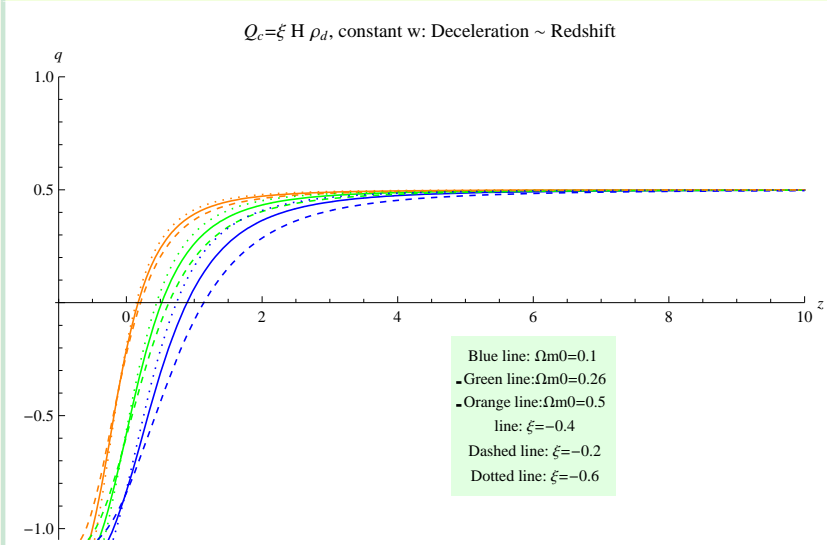
The plots of deceleration parameter are shown below. At the limit $z \rightarrow \text{Infinity}$, the deceleration parameter ALL goes to $\frac{1}{2}$.

Theoretically, this limit is $\frac{1}{2}$ which is not related to any parameters, with $3w + \xi < 0$.

In[349]:=

pldecI2CCShowSum

Out[349]=

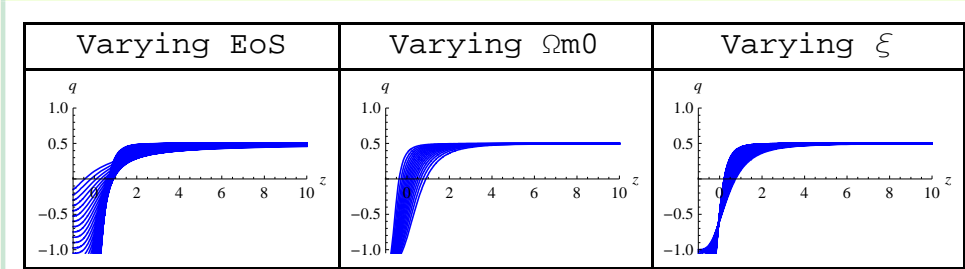


To check the effect of different parameters, another plot is shown.

In[350]:=

varyingI2CCShowSum

Out[350]=



A toy to play with deceleration vs z curve is also provided

In[351]:=

pldecI2CCManSum

If the transition happens before $z=0$, increasing coupling ξ will delay the transition.
If the it happens after $z=0$, increasing coupling ξ will bring forward the emergence of transition.

The following figure shows this result.

Gray rectangle is the region given by Riess.

Orange for $w=-1$
Blue for $w=-0.9$

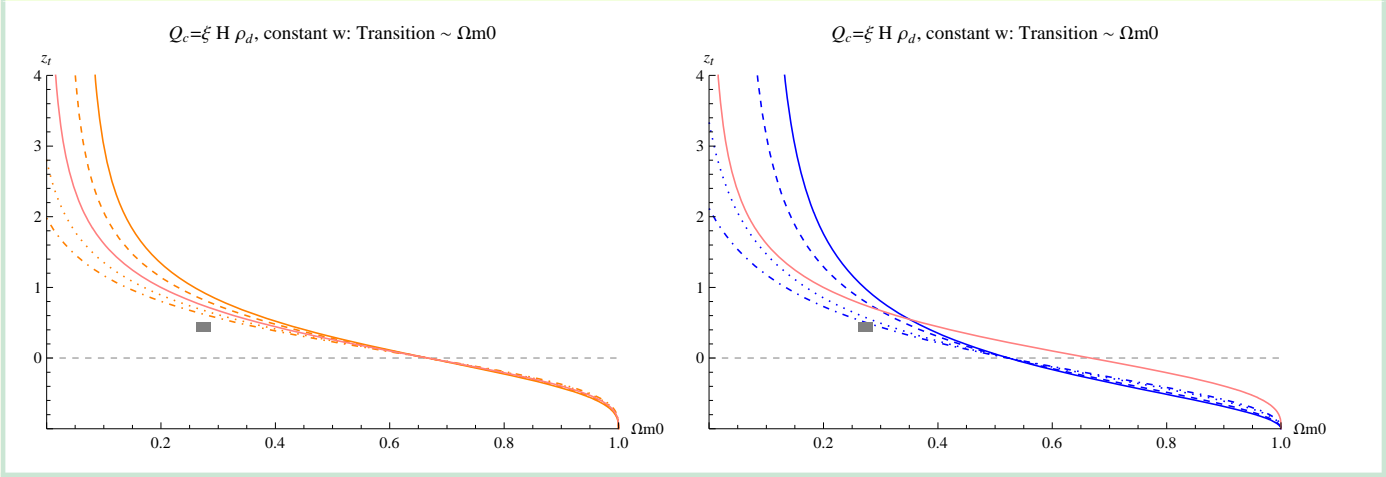
line: $\xi=0.2$
Dashed: $\xi=0.1$
Dotted: $\xi=-0.1$
DotDashed: $\xi=-0.2$

Pink line: $w=1$, $\xi=0$

In[352]:=

plztrvsΩm0I2CCSum

Out[352]=

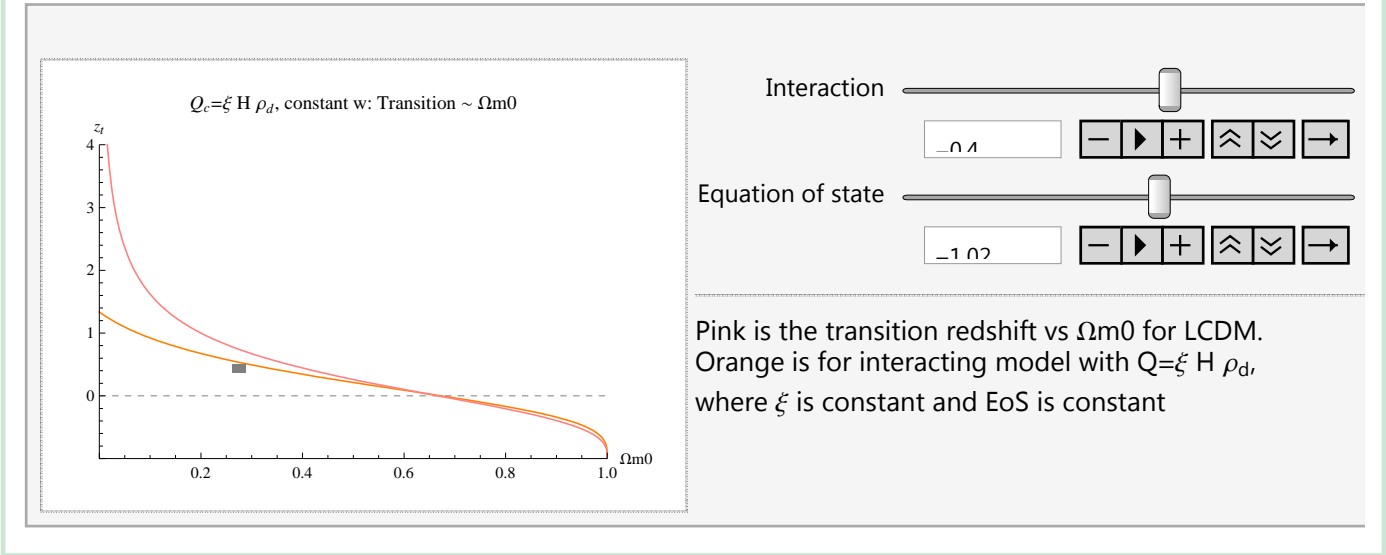


A toy of transition redshift. Gray rectangle is the allowed region of $\Omega_{m0} \sim$ Transition redshift

In[353]:=

plztrI2CCManSum

Out[353]=



The fitting results of coupling constant ξ is

Or we can use some fitting results from WMAP etc. Take the example of $w=-1.087\pm0.096$.

In[356]:=

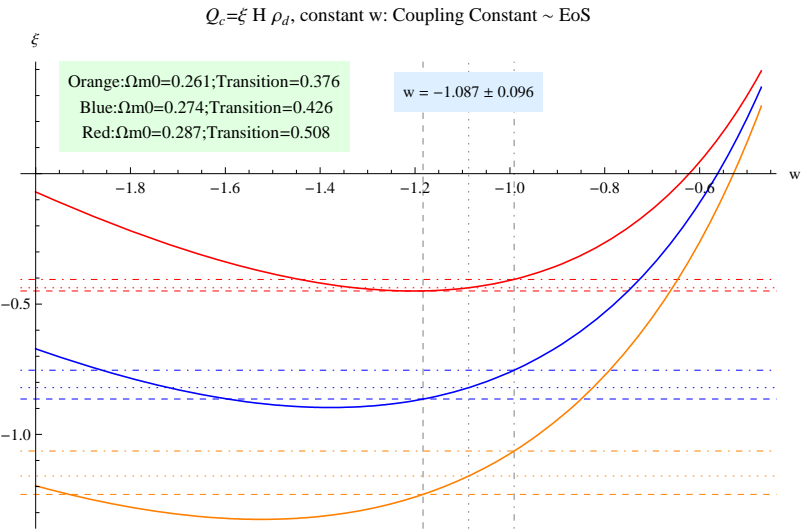
tabξvwExamI2CC

$Q_c=\xi\ H\ \rho_d, \text{Constant } w. \text{ (Data used:Data From, 2)}$			
w	Center	Lower	Upper
-1.183	-0.864289	-1.22984	-0.449552
-1.087	-0.820486	-1.15946	-0.437339
-0.991	-0.753634	-1.06346	-0.405262

Out[356]=

In[357]:=

pltξvwExamI2CC

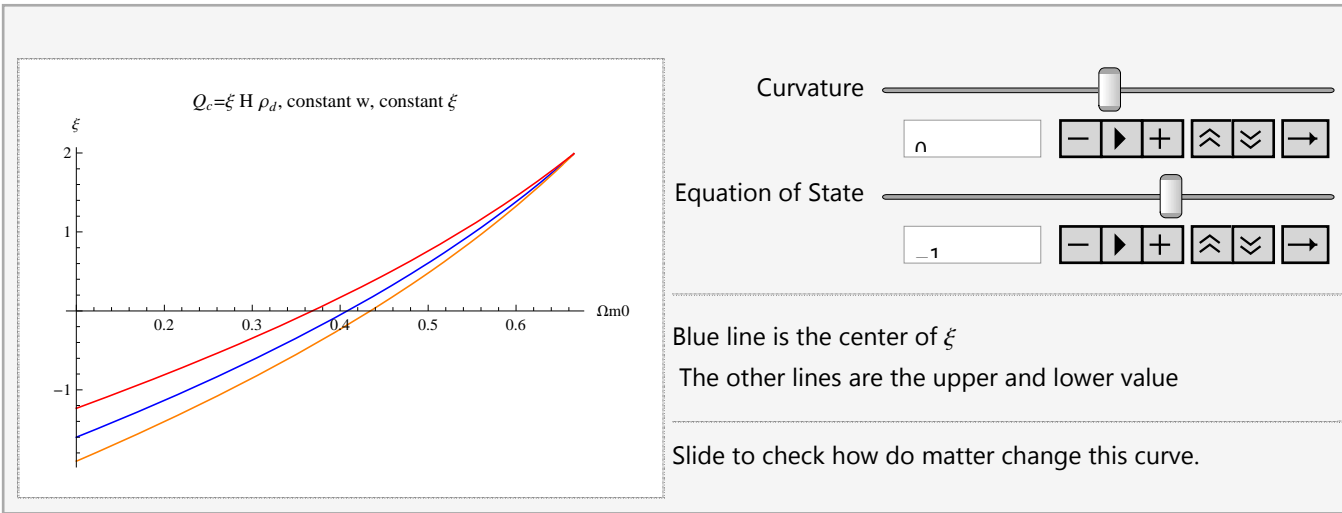


Out[357]=

Now we assume we do not have the observed Ω_{m0} data, how do this Ω_{m0} change the result. In other words, if the observed Ω_{m0} data float around some value, then how is the fitting result? We also consider the curvature.

In[358]:=

pltξvΩm0I2CCManSum



Out[358]=

If Ω_{m0} varies 0.05 percent from 0.274,

In[359]:=

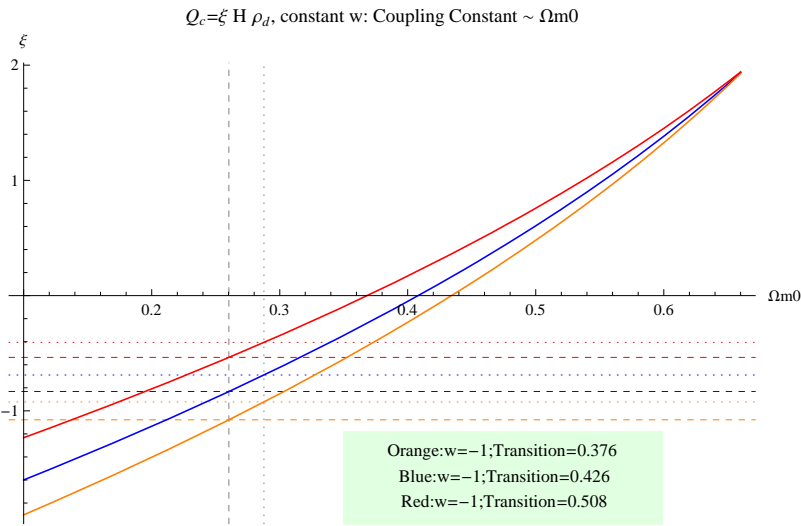
tabξI2CCSum

For $\Omega_{m0} \in 0.274 (1 \pm 0.05)$			
Table of ξ for different Ω_{m0} ~Transition combination			
Ω_{m0} ~Transition	0.426	0.376	0.508
0.2603	-0.832284	-1.07758	-0.53584
0.274	-0.760999	-1.00068	-0.471298
0.2877	-0.688664	-0.922602	-0.40585

Out[359]=

In[360]:=

pltξvΩm0I2CCSum



Out[360]=

In addition, we can also find out the effects of Curvature, EoS. Assuming we have a constrain of Transition redshift (0.376,0.508) with a center at 0.426.

In[361]:=

fitξ2I2CCManSum

Matter Fraction

Curvature

Equation of State

$Q_c = \xi H \rho_d$, constant w , constant ζ .

This is the fitting result of ξ from only transition redshift data without the knowledge of Ω_{m0} .

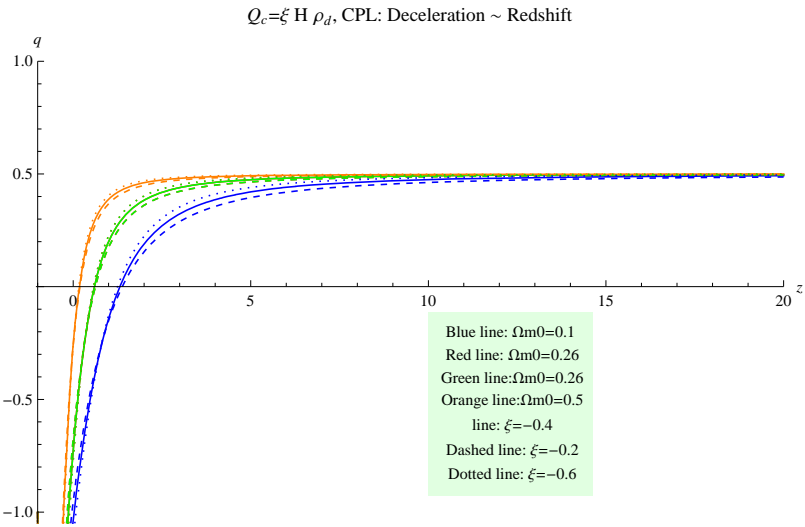
Slide to check how do EoS change this curve.

Out[361]=

■ Interacting model $Q_c = \xi H \rho_d$ with constant ξ and CPL parameterization.

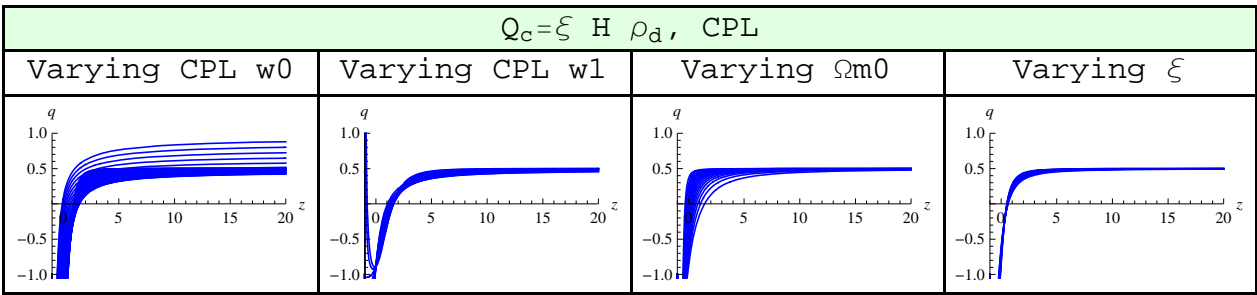
In[362]:=

pldecI2CCPLShowSum



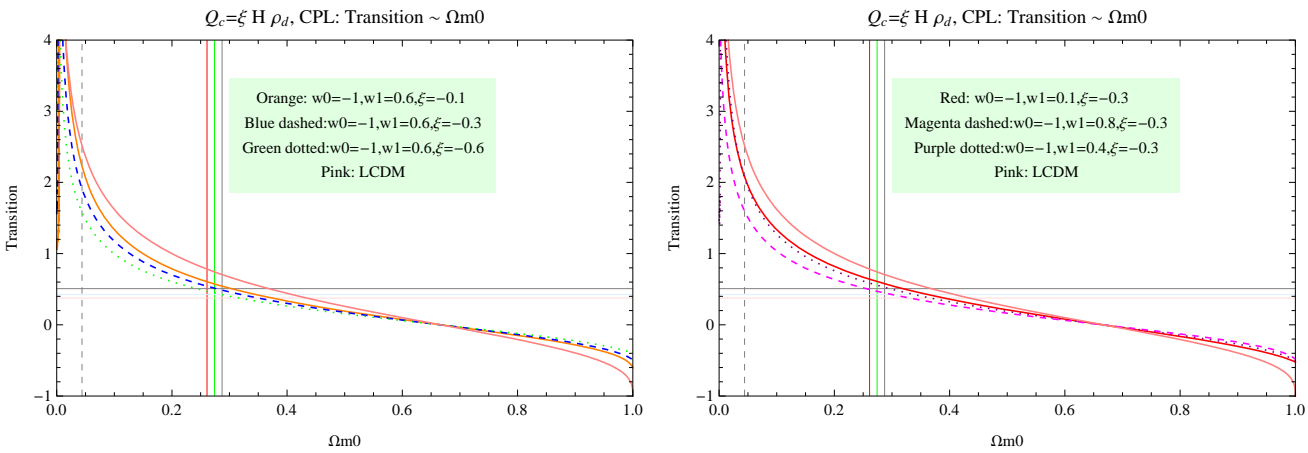
In[363]:=

varyingI2CCPLShowSum



In[364]:=

plztrExamI2CCPLSum

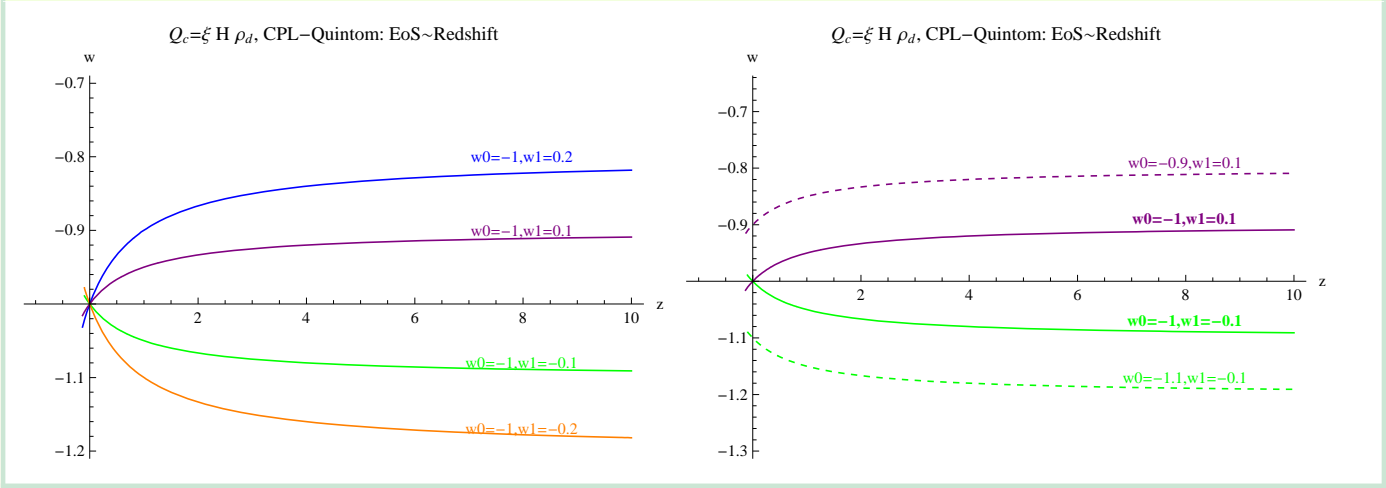


▣ Quintom

In[365]:=

plEoSII2CCPLQuintomSum

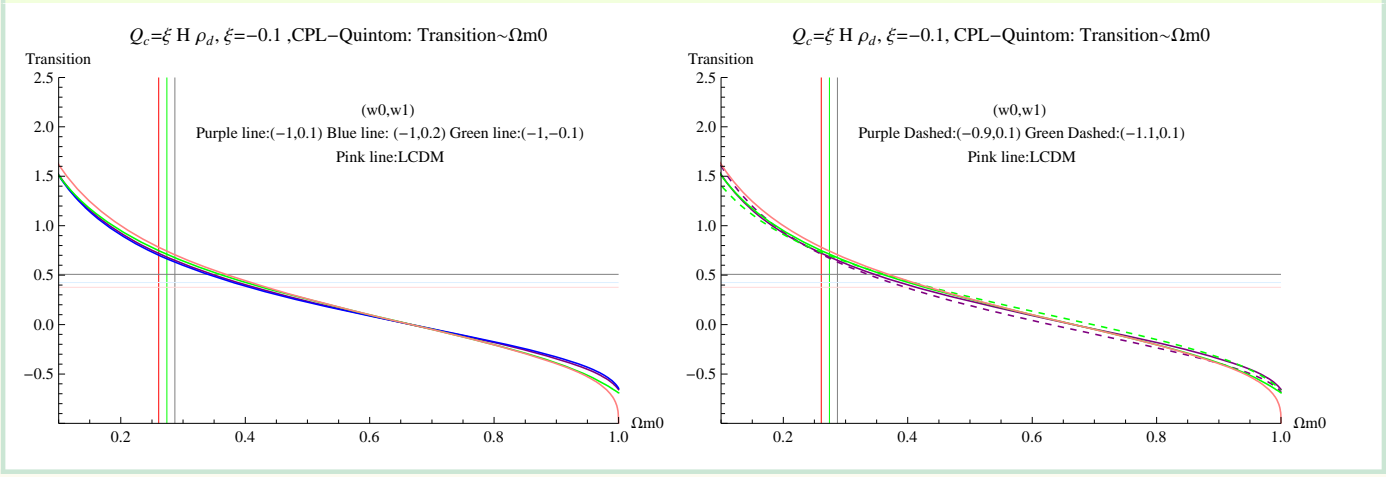
Out[365]=



In[366]:=

plztrII2CCPLQuintomSum

Out[366]=



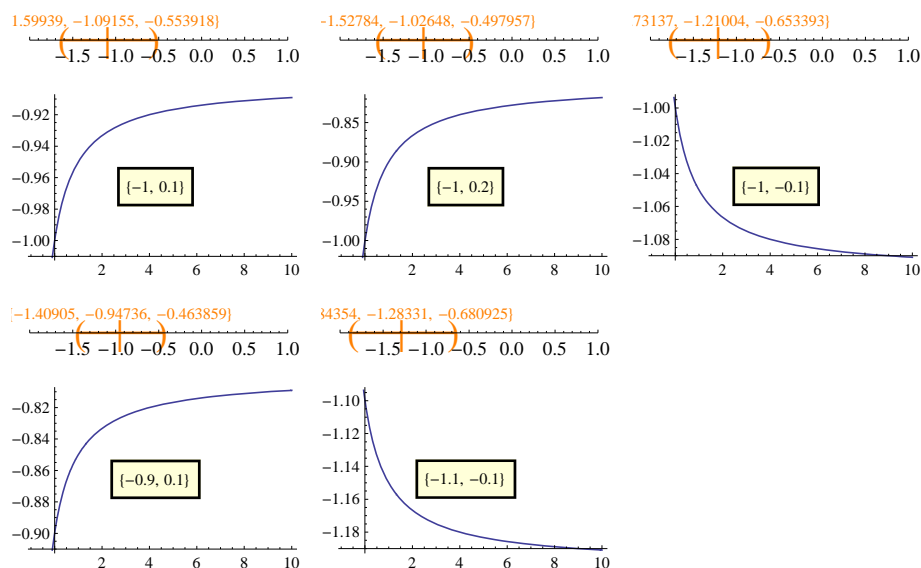
In[367]:=

tab\xvwExamI2CCPLQuintom

Out[367]=

ξ results for $Q_c = \xi H \rho_d$, CPL, Quintom.			
$\{w_0, w_1\}$	Center	Lower	Upper
$\{-1, -0.1\}$	-1.21004	-1.73137	-0.653393
$\{-1, 0\}$	-1.15265	-1.66715	-0.605615
$\{-1, 0.1\}$	-1.09155	-1.59939	-0.553918
$\{-0.9, 0.1\}$	-0.94736	-1.40905	-0.463859
$\{-1.1, -0.1\}$	-1.28331	-1.84354	-0.680925

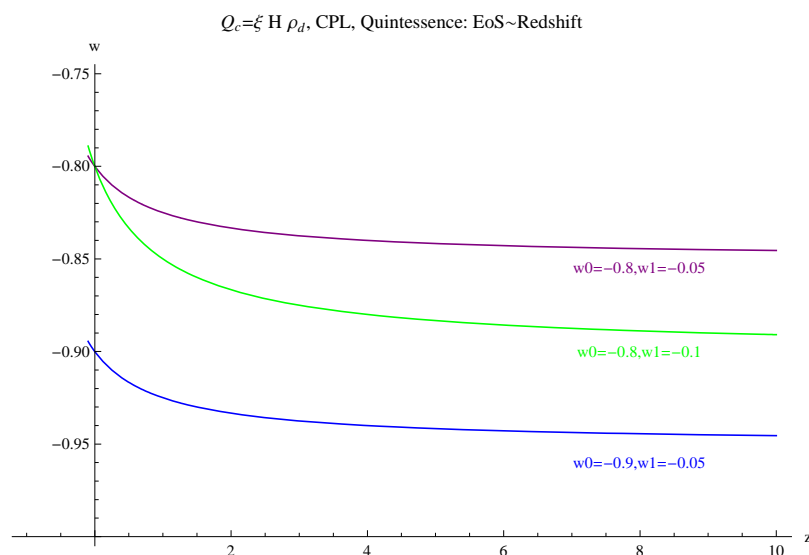
In[368]:=

pltfitξI2CCPLQuintomSum

Out[368]=

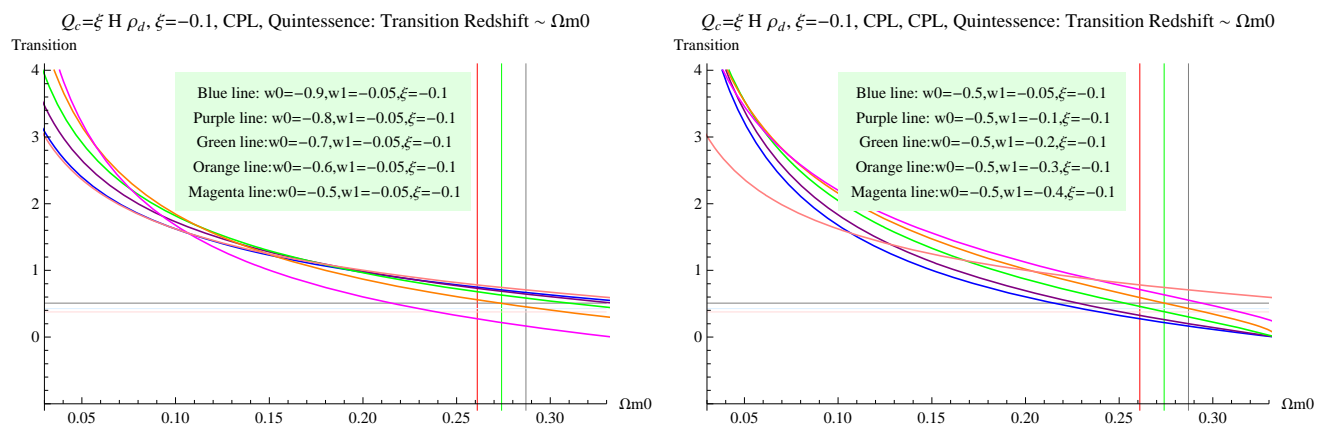
□ Quintessence

In[369]:=

plEoS12CCPLQuintessenceSum

Out[369]=

In[370]:=

plztrI2CCPLQuintessenceSum

Out[370]=

In[371]:=

tabξvwExamI2CCPLQuintessence

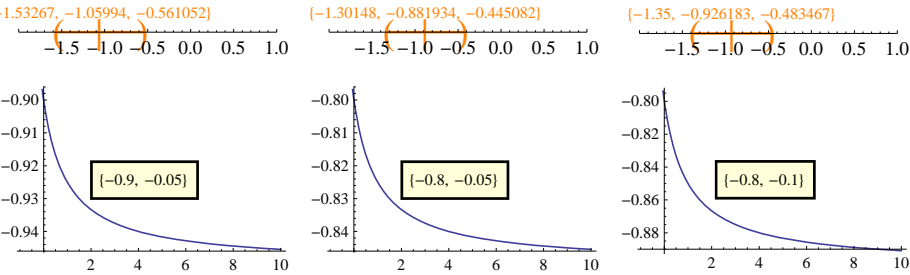
Out[371]=

ξ results for $Q_c=\xi$ H ρ_d , CPL,Quintessence.			
{w0,w1}	Center	Lower	Upper
{-0.9, -0.05}	-1.05994	-1.53267	-0.561052
{-0.8, -0.05}	-0.881934	-1.30148	-0.445082
{-0.8, -0.1}	-0.926183	-1.35	-0.483467

In[372]:=

pltfitξI2CCPLQuintessenceSum

Out[372]=

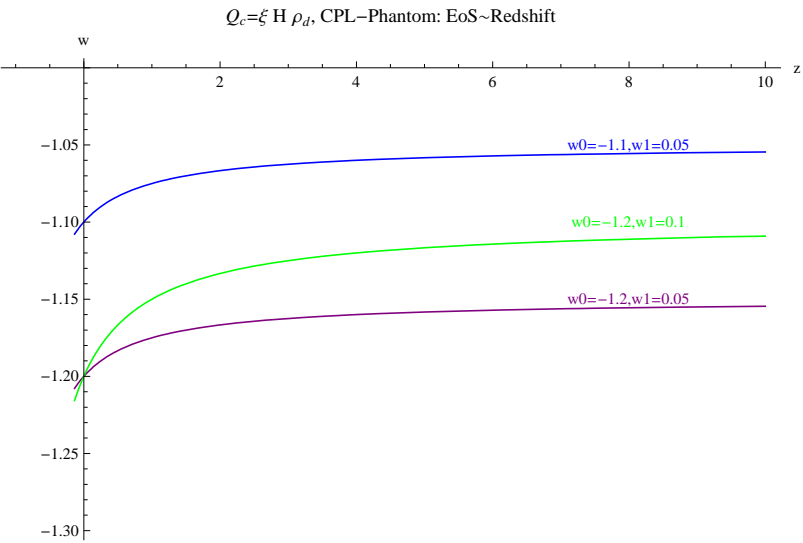


▣ Phantom

In[373]:=

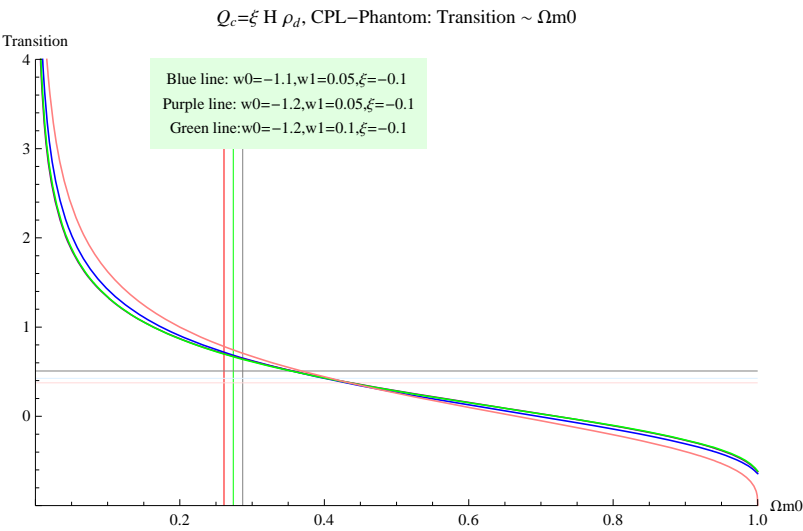
plEoSII2CCPLPhantomSum

Out[373]=



In[374]:=

plztrI2CCPLPhantomSum



Out[374]=

In[375]:=

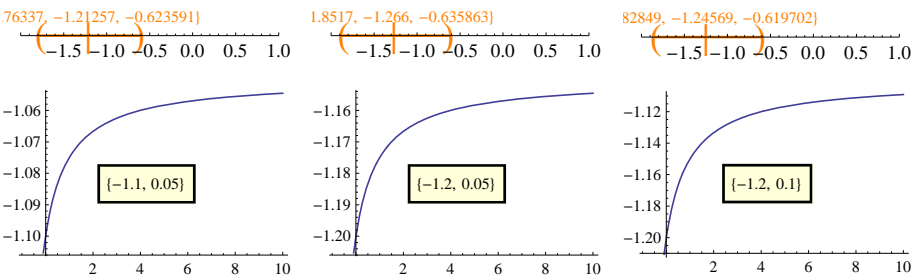
tabξvwExamI2CCPLPhantom

ξ results for $Q_c = \xi H \rho_d$, CPL, Phantom.			
$\{w_0, w_1\}$	Center	Lower	Upper
$\{-1.1, 0.05\}$	-1.21257	-1.76337	-0.623591
$\{-1.2, 0.05\}$	-1.266	-1.8517	-0.635863
$\{-1.2, 0.1\}$	-1.24569	-1.82849	-0.619702

Out[375]=

In[376]:=

pltfitξI2CCPLPhantomSum



Out[376]=

■ References

▣ Data From

1. CPL data

Combining SN1a, BAO 3, WMAP5, $H(z)$ (From arXiv:0909.0596)

$\Omega_{m0} = 0.269^{+0.017}_{-0.008}, w_0 = -0.97^{+0.12}_{-0.07}, w_1 = 0.03^{+0.26}_{-0.75}.$

2. LCDM

From WMAP: $\Omega_{m0}=0.265$
arXiv:astro-ph/0611572, Riess et al :
arXiv:1205.4688 : $\Omega_{m0}=0.247 (+0.013, -0.013)$ and Transition $0.426 (+0.082, -0.050)$

3. Equation of state

From arXiv:1202.0545v1

$w = -1.087 \pm 0.096$.

INTRO

PRE

ΛCDM Model

Interacting Models

Summary

Supplementary

■

1. Why does smaller ξ lead to earlier transition in $Q=\xi H \rho_c$ model?

$\dot{\rho}_c + 3 H \rho_c = \xi H \rho_c$, positive ξ means energy flow from DE to DM, negative ξ means DM to DE.

This can be explained with the equations in the file I sent last time. I'll retype them here.

Evolution of energy density for $Q_c = \xi H \rho_c$, constant ξ , constant w , and $\xi \neq -3w$

$$\Omega_m = \Omega_{m0} (1+z)^{3-\xi} \quad (2)$$

$$\Omega_d = \left(\Omega_{d0} + \frac{\xi}{3w+\xi} \Omega_{m0} \right) (1+z)^{3(1+w)} + \frac{-\xi}{3w+\xi} \Omega_m \equiv \Omega_{d0} (1+z)^{3(1+w)} + \frac{-\xi}{3w+\xi} \Omega_m \quad (3)$$

Evolution of energy density for $Q_c = \xi H \rho_d$, constant ξ , constant w , and $\xi \neq -3w$

$$\Omega_m = \left(\Omega_{m0} + \frac{\xi}{\xi+3w} \Omega_{d0} \right) (1+z)^3 + \frac{-\xi}{\xi+3w} \Omega_d \equiv \Omega_{m0} (1+z)^3 + \frac{-\xi}{\xi+3w} \Omega_d \quad (4)$$

$$\Omega_d = \Omega_{d0} (1+z)^{3(1+w)+\xi} \quad (5)$$

So in the two cases, coupling constant has two effects:

1. Amplifies the curve of deceleration parameter / energy density.

2. Energy flow between DE and DM.

These solutions have the same Ω_d and Ω_m values at $z=0$. That means energy flow of DM to DE would yield a larger energy density of DM at $z>0$ than ΛCDM. That means there would be less DM for a larger ξ if the transition happens before $z=0$.

This is also the reason of that amplification effect mentioned above. A smaller ξ will decrease the factor

$\left(\Omega_{d0} + \frac{\xi}{3w+\xi} \Omega_{m0} \right) (1+z)^{3(1+w)}$ in Ω_d . I will plot the evolution of Ω_d in terms of redshift.

In[65]:=

```
Omegadfunctest[Ωd0_, Ωm0_, w_, ξ_, z_] :=  
  (Ωd0 + (ξ / (3w + ξ)) Ωm0) (1 + z)^(3(1+w)) + (-ξ / (3w + ξ)) (Ωm0 (1 + z)^(3-ξ));  
Omegamfunctest[Ωm0_, ξ_, z_] := Ωm0 (1 + z)^(3-ξ);
```

```

In[67]:=
Omegadplottest[ $\Omega d_0$ _,  $\Omega m_0$ _,  $w$ _,  $\xi$ _,  $color$ _] :=
  Plot[Omegadfunctest[ $\Omega d_0$ ,  $\Omega m_0$ ,  $w$ ,  $\xi$ ,  $z$ ], { $z$ , -0.9, 10}, PlotStyle →  $color$ ];
Omegamplottest[ $\Omega m_0$ _,  $\xi$ _,  $color$ _] :=
  LogPlot[Omegamfunctest[ $\Omega m_0$ ,  $\xi$ ,  $z$ ], { $z$ , -0.9, 10}, PlotStyle →  $color$ ];

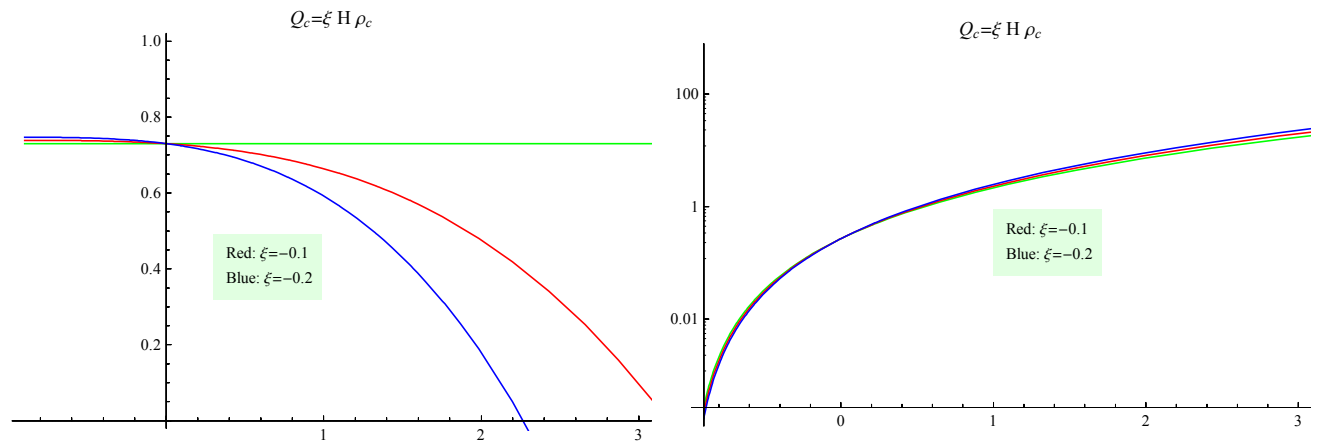
```

```

In[86]:=
Grid[
  {{Show[Omegadplottest[0.73, 0.27, -1, 0, Green],
    Omegadplottest[0.73, 0.27, -1, -0.1, Red],
    Omegadplottest[0.73, 0.27, -1, -0.2, Blue], PlotRange → {{-0.9, 3}, {0, 1}},
    PlotLabel → " $Q_c = \xi H \rho_c$ ",
    Epilog → Inset[Framed[Style["Red:  $\xi = -0.1$ \n Blue:  $\xi = -0.2$ ", 10],
      Background → LightGreen, FrameStyle → None], {0.3, 0.5}, {Left, Top}],
    ImageSize → 400], Show[Omegamplottest[0.27, 0, Green],
    Omegamplottest[0.27, -0.1, Red], Omegamplottest[0.27, -0.2, Blue],
    PlotRange → {{-0.9, 3}, Automatic}, PlotLabel → " $Q_c = \xi H \rho_c$ ",
    Epilog → Inset[Framed[Style["Red:  $\xi = -0.1$ \n Blue:  $\xi = -0.2$ ", 10],
      Background → LightGreen, FrameStyle → None], {1, 0.01}, {Left, Top}],
    ImageSize → 400]}}]

```

Out[86]=

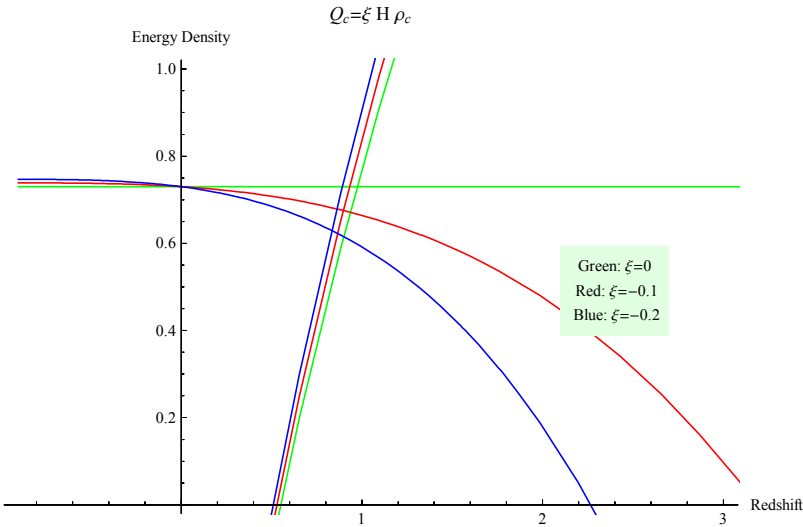


The following figure shows that if the transition happens before $z=0$, a larger.

In[87]:=

```
Grid[
  {{Show[Omegadplottest[0.73, 0.27, -1, 0, Green], Omegamplottest[0.27, 0, Green],
    Omegadplottest[0.73, 0.27, -1, -0.1, Red], Omegamplottest[0.27, -0.1, Red],
    Omegadplottest[0.73, 0.27, -1, -0.2, Blue], Omegamplottest[0.27, -0.2, Blue],
    PlotRange -> {{-0.9, 3}, {0, 1}}, AxesLabel -> {"Redshift", "Energy Density"},
    PlotLabel -> " $Q_c = \xi H \rho_c$ ",
    Epilog -> Inset[Framed[Style["Green:  $\xi=0$ \n Red:  $\xi=-0.1$ \n Blue:  $\xi=-0.2$ ", 10],
      Background -> LightGreen, FrameStyle -> None], {2.1, 0.6}, {Left, Top}],
    ImageSize -> 500]}}
```

Out[87]=

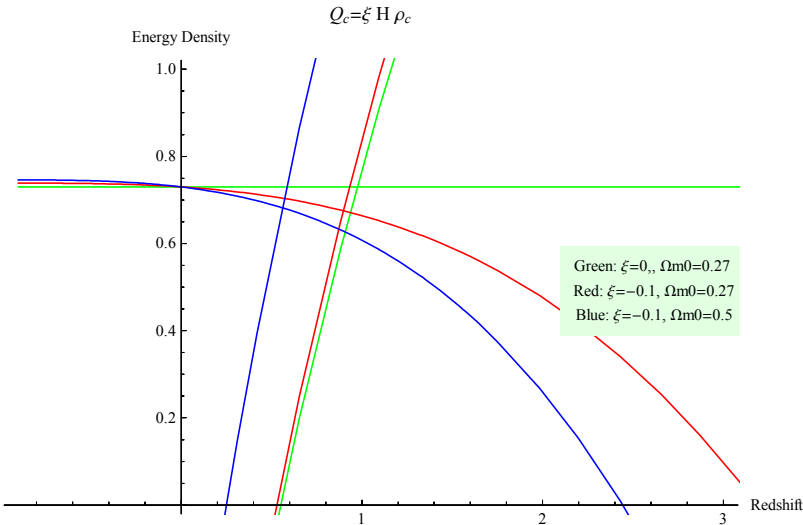


Ω_{m0} here is the energy density of matter today. If we have more matter today, the transition happens nearer to $z=0$. Since the energy density of DE and DM varies less with a smaller redshift, the effect of ξ would be reduced.

In[88]:=

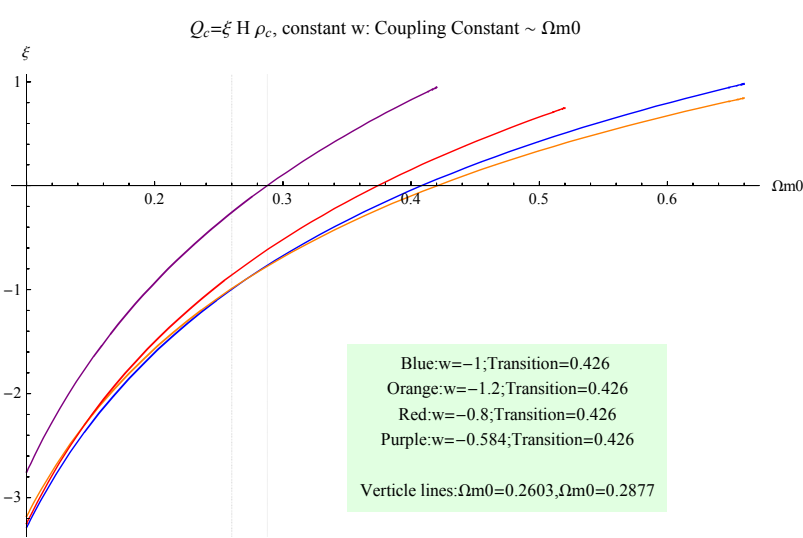
```
Grid[
  {{Show[Omegadplottest[0.73, 0.27, -1, 0, Green], Omegamplottest[0.27, 0, Green],
    Omegadplottest[0.73, 0.27, -1, -0.1, Red], Omegamplottest[0.27, -0.1, Red],
    Omegadplottest[0.73, 0.5, -1, -0.1, Blue], Omegamplottest[0.5, -0.1, Blue],
    PlotRange -> {{-0.9, 3}, {0, 1}}, AxesLabel -> {"Redshift", "Energy Density"},
    PlotLabel -> "Q_c=ξ H ρ_c",
    Epilog ->
      Inset[
        Framed[
          Style[
            "Green: ξ=0,, Ωm0=0.27\n Red: ξ=-0.1, Ωm0=0.27\n Blue: ξ=-0.1, Ωm0=0.5",
            10], Background -> LightGreen, FrameStyle -> None], {2.1, 0.6}, {Left, Top}],
        ImageSize -> 500]}}}]
```

Out[88]=



▢ Interacting model $Q_c = \xi H \rho_c$ with constant ξ and constant EoS w .

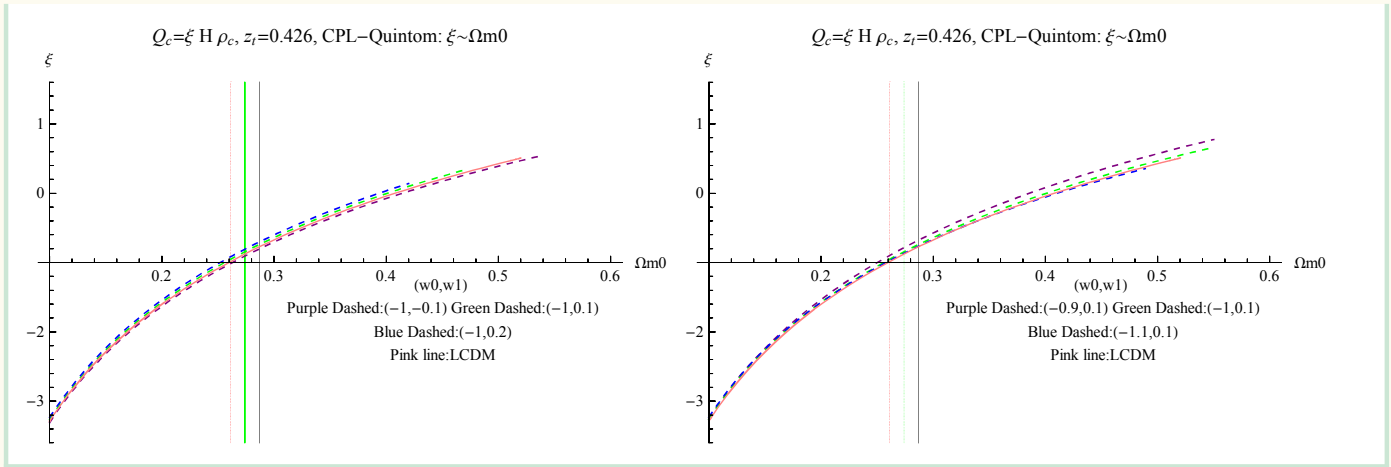
Out[1]=



EoS value when $\xi=0$	
w	Transition 0.426
$\Omega_{m0}=0.2877$	-0.58406

Result: $w \in (-0.58406, -0.45064)$ if we constrain $\Omega_{m0} \in (0.2603, 0.2877)$ and transition redshift 0.426.

- ▣ Interacting model $Q_c = \xi H \rho_c$ with constant ξ and CPL parameterized EoS $w = w_0 + w_1 \frac{z}{1+z}$.



Caption of the following figures:

Figure on the left:

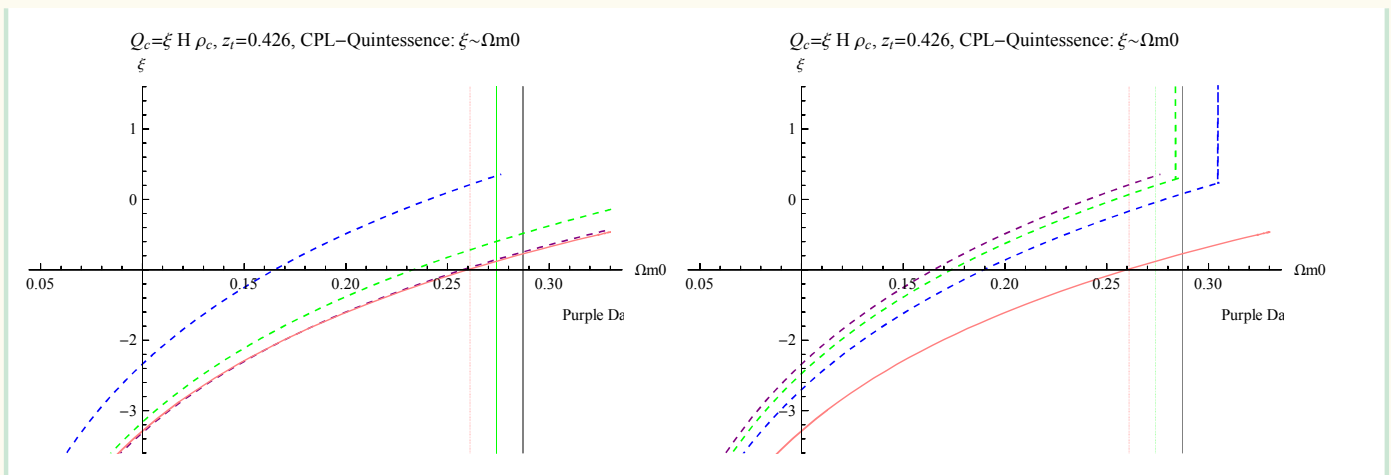
(w_0, w_1) value of the lines:

Purple Dashed: $(-0.9, -0.05)$; Green Dashed: $(-0.7, -0.05)$; Blue Dashed: $(-0.5, -0.05)$; Pink line: LCDM

Figure on the right:

(w_0, w_1) value:

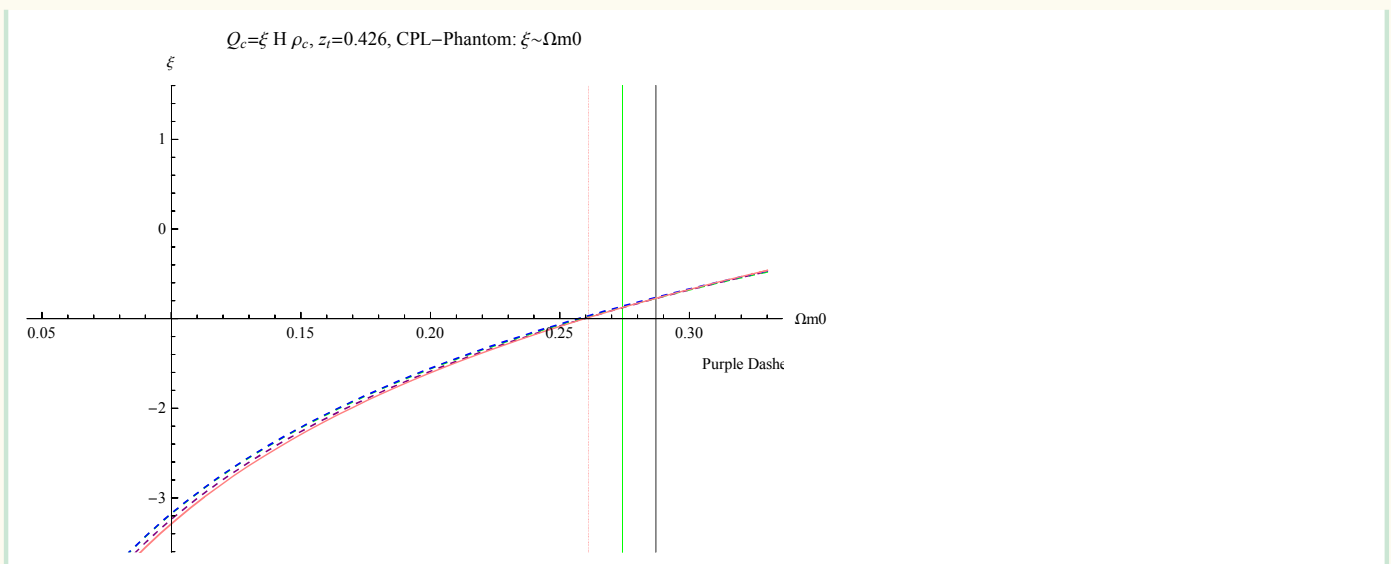
Purple Dashed: $(-0.5, -0.05)$; Green Dashed: $(-0.5, -0.1)$; Blue Dashed: $(-0.5, -0.2)$; Pink line: LCDM



Caption for the following figure:

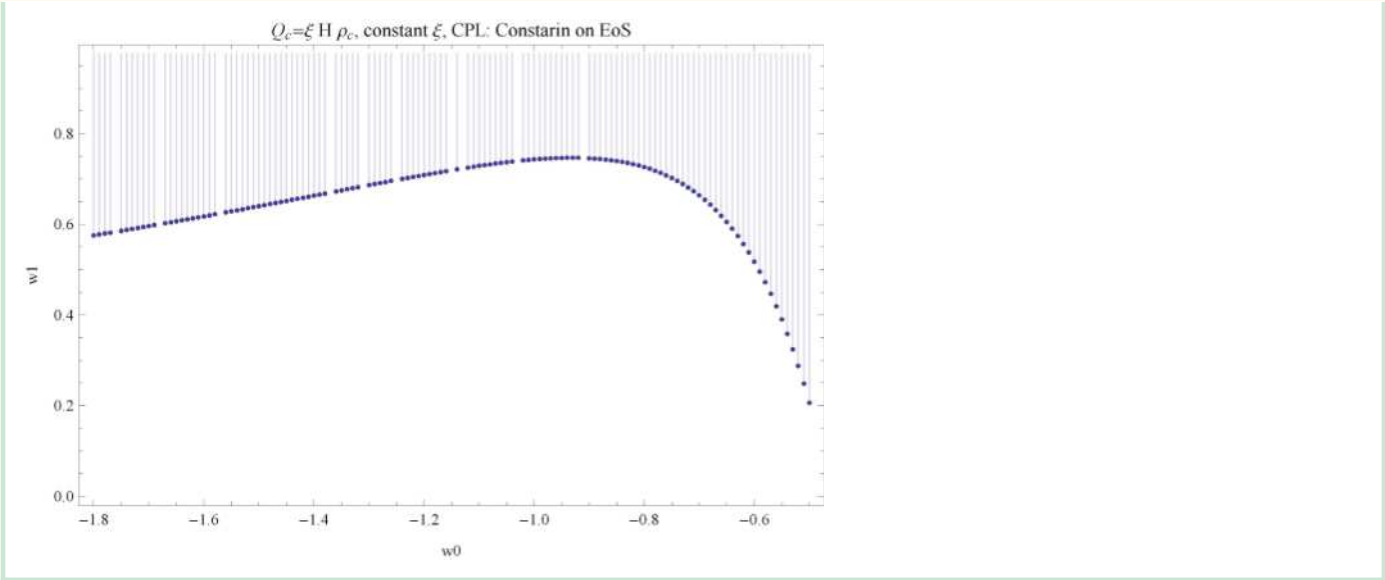
(w_0, w_1) value:

Purple Dashed : $(-1.1, 0.05)$; Green Dashed : $(-1.2, 0.05)$; Blue Dashed : $(-1.2, 0.1)$; Pink line : LCDM

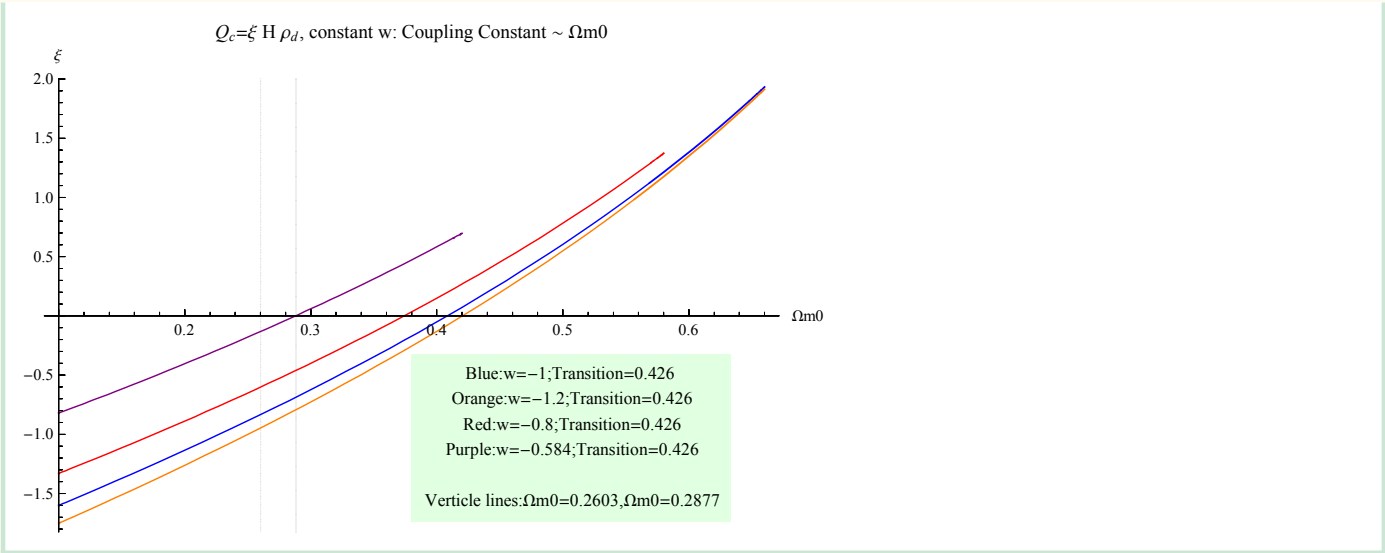


Lower boundary of the the allowed parameters. Since lower boundary is calculated when $\xi=0$, this remains the same for different models we investigated here.

There should be a upper boundary. But I haven't calculated it yet. The upper boundary should vary for different models.



□ Interacting model $Q_c = \xi H \rho_d$ with constant ξ and constant EoS w .



$Q_c = \xi H \rho_d$, constant w : EoS value when $\xi=0$	
\therefore	Transition 0.426
$\Omega_{m0}=0.2877$	—

Figure on the left

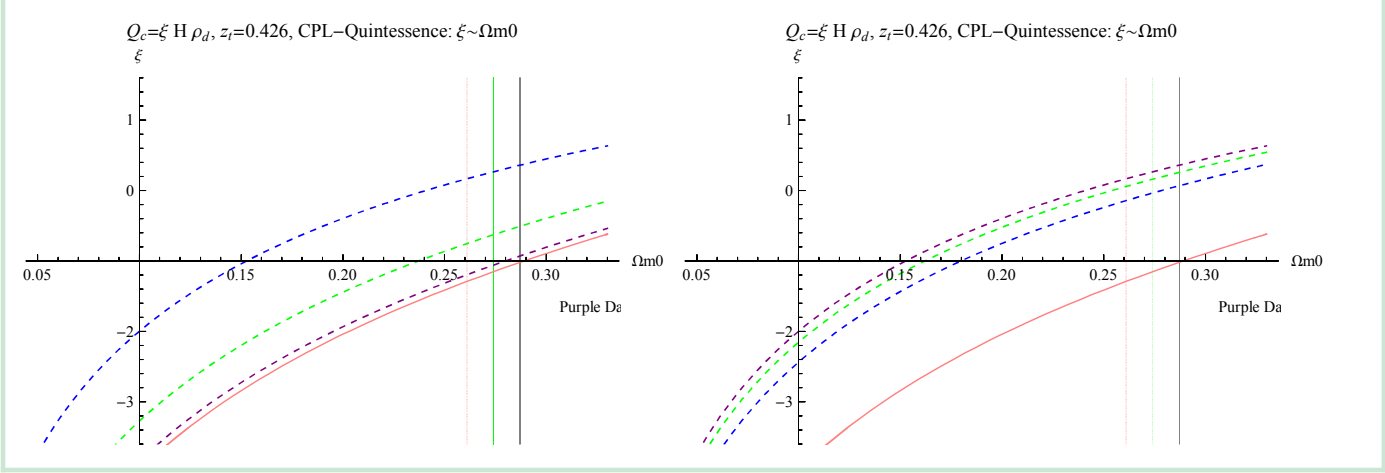
(w0, w1) :

Purple Dashed : (-0.9, -0.05); Green Dashed : (-0.7, -0.05);
Blue Dashed : (-0.5, -0.05); Pink line : LCDM

Figure on the right

(w0, w1) :

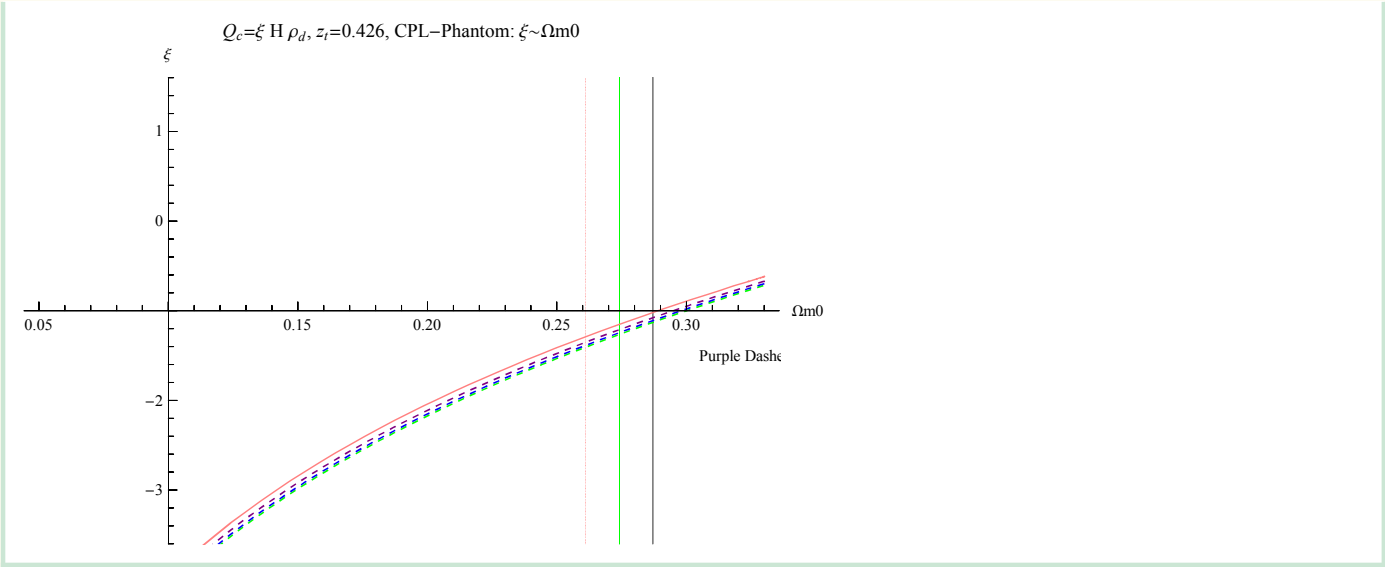
Purple Dashed : (-0.5, -0.05); Green Dashed : (-0.5, -0.1);
Blue Dashed : (-0.5, -0.2); Pink line : LCDM



The following figure:

(w0,w1):

Purple Dashed:(-1.1,0.05); Green Dashed:(-1.2,0.05); Blue Dashed:(-1.2,0.1); Pink line:LCDM



Here I recalculated the lower boundary in this model. This is exactly the same with the one I plotted previously.

