

INTRO

PRE

LCDM Model

Interacting Models

Summary

■ Interacting models

□ List of what to make clear

□ BASIC

Evolution of energy density for $Q_c = \xi H \rho_c$, constant ξ , constant w , and $\xi \neq -3w$

$$\Omega_m = \Omega_{m0} (1+z)^{3-\xi}$$

$$\Omega_d = \left(\Omega_{d0} + \frac{\xi}{3w + \xi} \Omega_{m0} \right) (1+z)^{3(1+w)} + \frac{-\xi}{3w + \xi} \Omega_m \equiv \Omega_{d0}' (1+z)^{3(1+w)} + \frac{-\xi}{3w + \xi} \Omega_m$$

Evolution of energy density for $Q_c = \xi H \rho_d$, constant ξ , constant w , and $\xi \neq -3w$

$$\Omega_m = \left(\Omega_{m0} + \frac{\xi}{\xi + 3w} \Omega_{d0} \right) (1+z)^3 + \frac{-\xi}{\xi + 3w} \Omega_d \equiv \Omega_{m0}' (1+z)^3 + \frac{-\xi}{\xi + 3w} \Omega_d$$

$$\Omega_d = \Omega_{d0} (1+z)^{3(1+w)+\xi}$$

So in the two cases, coupling constant has two effects:

1. Amplifies the curve of deceleration parameter.

2. Energy flow between DE and DM.

■ Interacting model $Q_c = \xi H \rho_c$ with constant ξ and constant EoS w .

Derived from (transition redshift, Ω_{m0}) plane, the allowed region for coupling constant ξ is $(-1.28, -0.46)$ with a center at -0.88 , i.e., $-0.88^{+0.42}_{-0.40}$, taken the case that the universe is flat, and choose the EoS parameter $\{w=-1\}$.

Derived from the (transition redshift, $\frac{\Omega_{m0}}{\Omega_{d0}}$) plane, the allowed region of coupling constant ξ is $(-1.25, -0.47)$ with a center at -0.88 , i.e., $-0.88^{+0.41}_{-0.37}$.

There is a bit difference between the two answers. One possible reason is the second method doesn't assume a flat universe, while the first one supposes the universe is flat.

The full table of fitting results are shown below. The light purple element are the final results.

In[315]:=

tabξFinaltICC

$Q_c = \xi H \rho_c$, constant ξ , constant $w = -1$: Results for ξ			
Ω_{m0}/Ω_{d0} :Transition	$z_t = 0.376$	$z_t = 0.426$	$z_t = 0.508$
$r = 0.358$	-1.25282	-0.965436	-0.617444
$r = 0.378$	-1.15011	-0.875189	-0.542347
$r = 0.398$	-1.05453	-0.791252	-0.472561

Out[315]=

To check the consistency of the two methods ((Transition, Ω_{m0}) plane fitting and (Transition, Ω_{m0}/Ω_{d0}) plane fitting), we find out the fitting results of coupling constant ξ for a flat universe, i.e., $r = \frac{\Omega_{m0}}{1-\Omega_{m0}}$ in the (Transition, Ω_{m0}) plane, applying the data from (Transition, Ω_{m0}/Ω_{d0}) plane.

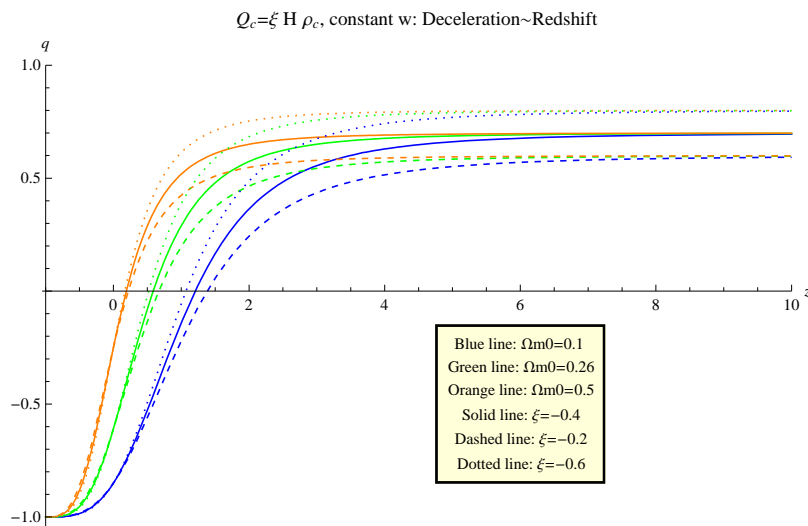
By solving out Ω_{m0} , we get $\Omega_{m0} = \frac{r}{1+r}$ (this is a monotonic function) in this case. Thus if we use the constrain that $r \in (0.358, 0.398)$ with a center value 0.378, the value of Ω_{m0} is (0.263623, 0.284692), centered at 0.274311. Use this set of value of Ω_{m0} as the constrain, we have the fitting results in (Transition, Ω_{m0}) plane, which is $-0.88^{+0.41}_{-0.37}$. This result is exactly the same as the result directly derived from (transition redshift, $\frac{\Omega_{m0}}{\Omega_{d0}}$) plane. The same has been done to $Q_c = \xi H \rho_d$ with ξ constant and w constant model, and the result is that the two methods are also consistent.

The plots of deceleration parameter are shown below. At the limit $z \rightarrow \infty$, the deceleration parameter is degenerate for different Ω_{m0} in this constant ξ and constant w model.

Theoretically, this limit is determined by the interaction coupling constant ξ , which is $\frac{(1-\xi)}{2}$, with $3w + \xi < 0$.

In[316]:=

pldecICCSum



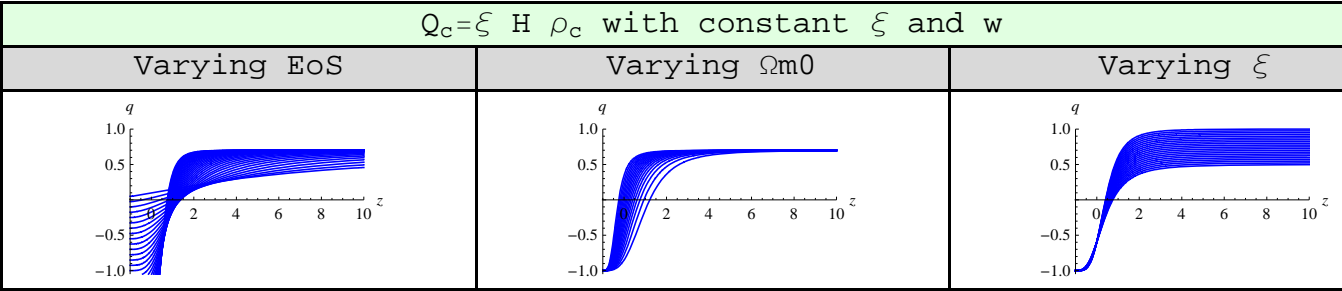
Out[316]=

Check the effect of different parameters on deceleration parameter.

Interaction ξ changes the value of deceleration parameter at $z \rightarrow \infty$ limit. EoS changes the the whole shape. Matter fraction determines how fast q varies, but just in a small time scale.

In[317]:=

varyingICCSum

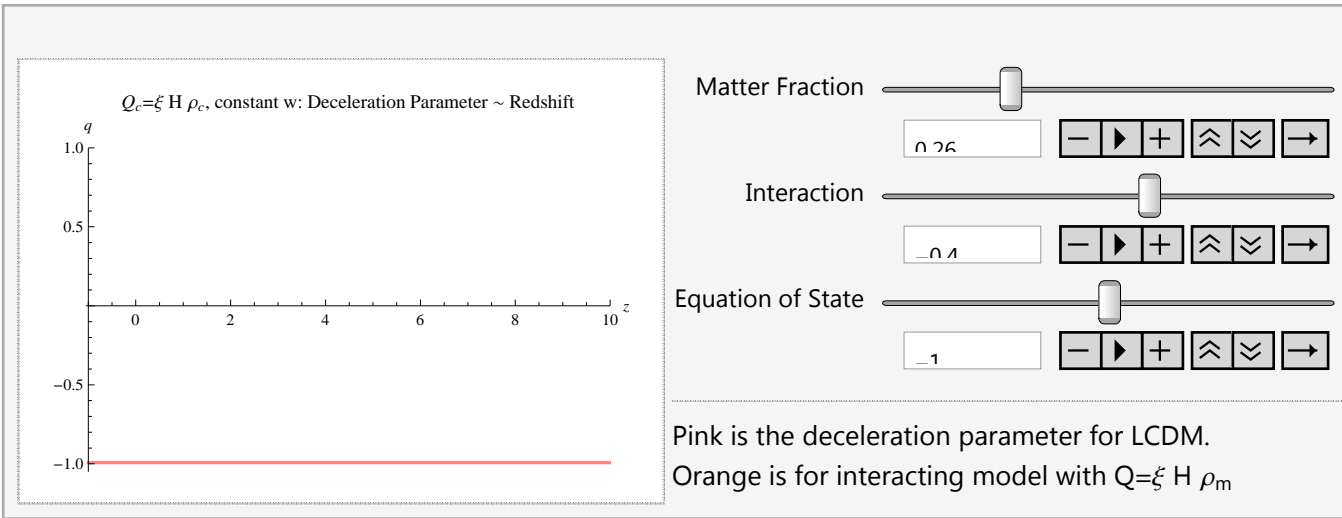


Out[317]=

A toy to play with is also provided. Slide the bars to view the effects of different parameters on deceleration parameter.

In[318]:=

pldecICCSum

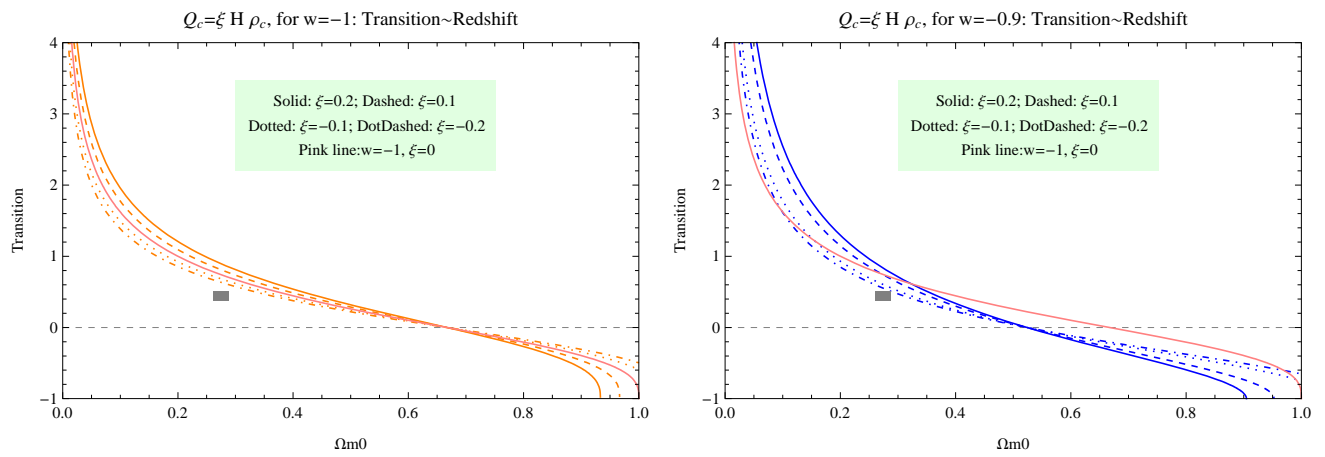


Out[318]=

It can be inferred from the expression for Ω_m and Ω_d that if the transition happens before $z=0$, increasing coupling ξ will bring forward the transition and if it happens after $z=0$, increasing coupling ξ will delay the emergence of transition. The following figure shows this result. Gray rectangle is the region given by Riess (References, Data From, 2).

- Orange for $w=-1$
- Blue for $w=-0.9$
- Solid line: $\xi=0.2$
- Dashed line: $\xi=0.1$
- Dotted line: $\xi=-0.1$
- DotDashed line: $\xi=-0.2$
- Pink solid line: $w=-1, \xi=0$

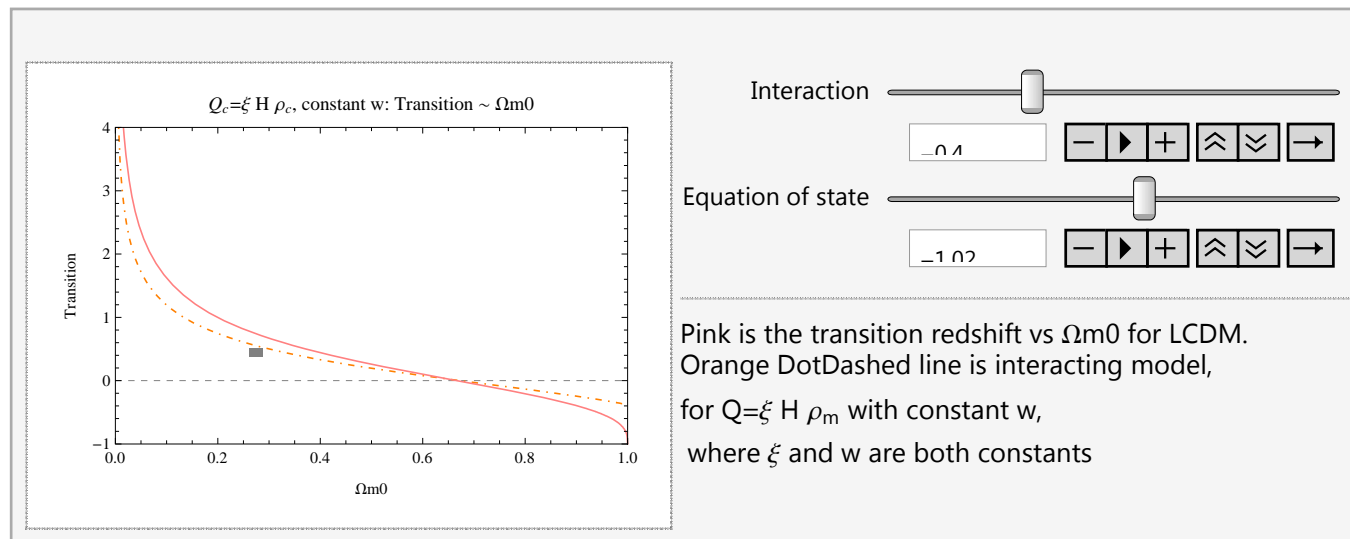
In[319]:=

plztrvsΩm0ICCSum

Out[319]=

This can also be seen clearly from the following toy. Gray rectangle is the region given by Riess (References, Data From, 2) .

In[320]:=

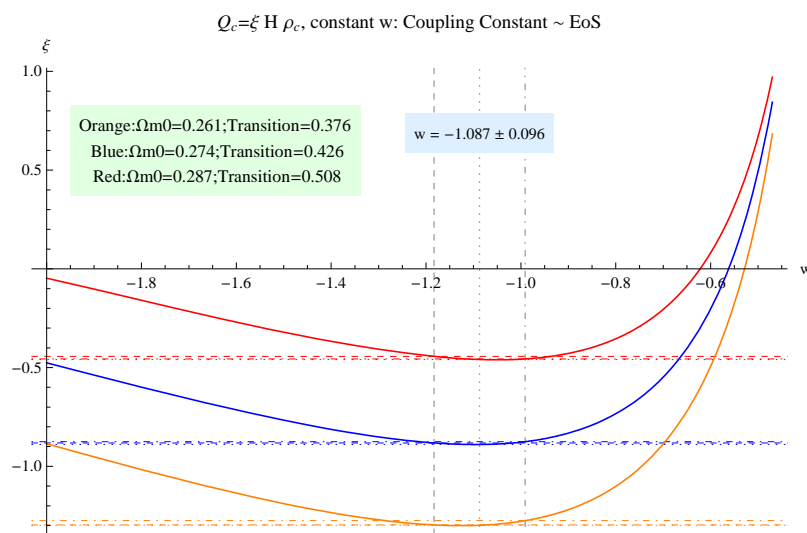
plztrICCMansum

Out[320]=

The fitting results of coupling constant ξ can also be dynamic.

In[323]:=

pltξvwExamICC



Out[323]=

Or just casually use the following parameters.

(-1 within 5%: (-1.05, -0.95))

$w=-1$ (-1.279, -0.457) center: -0.878

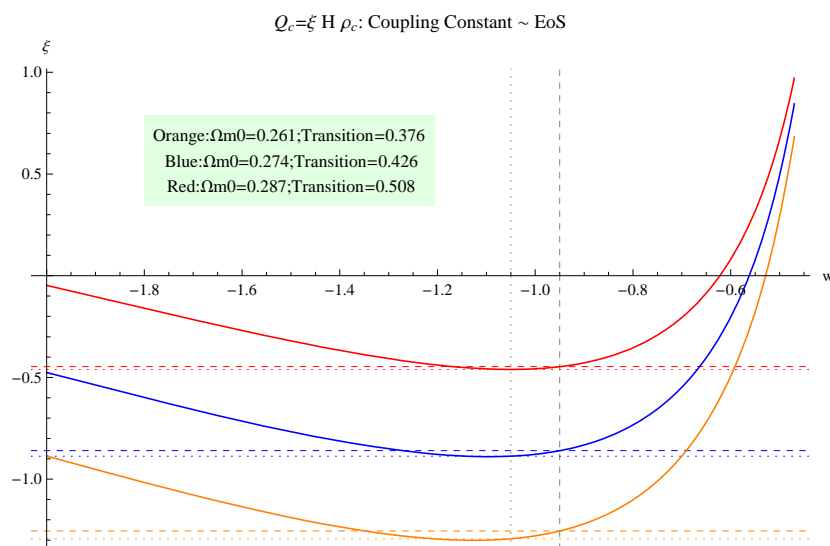
$w=-1.05$ (-1.293, -0.461) center: -0.887

$w=-0.95$ (-1.255, -0.447) center: -0.860

The following graph show how do ξ changes with EoS. The grid lines are the results of $w = -1 \pm 0.05$. Two verticle lines are -1.05 and -0.95 respectively. Horizontal lines are their intersections with the $\xi \sim w$ lines. EoS does not monotonically change ξ . And the minima of these line occurs at a larger w with an increasing Ω_{m0} .

In[324]:=

pltξvwExam1ICC



Out[324]=

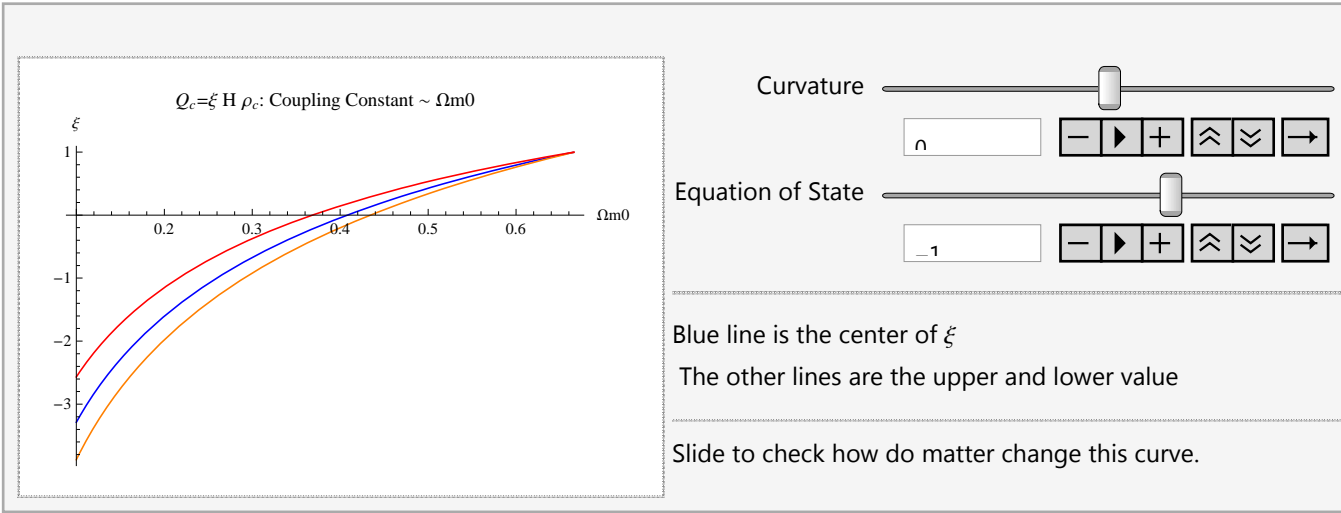
Now we assume we do not have the observed Ω_{m0} data, how do this Ω_{m0} change the result of ξ . In other words, if the observed Ω_{m0} data float around some value, then how is the fitting result? We also consider the curvature.

In the figure below, it seems that there is a point where three lines converge. This has something to do with the phenomena that

In[325]:=

pltξvΩm0ICCMansum

Out[325]=



Some data for flat Λ CDM universe. The following data shows how $\Omega m0$ change our results for ξ if we already have transition redshift data {0.426|0.376,0.508}.

If $\Omega m0$ varies 5 percent from 0.274,

In[326]:=

tabξICCSum

Out[326]=

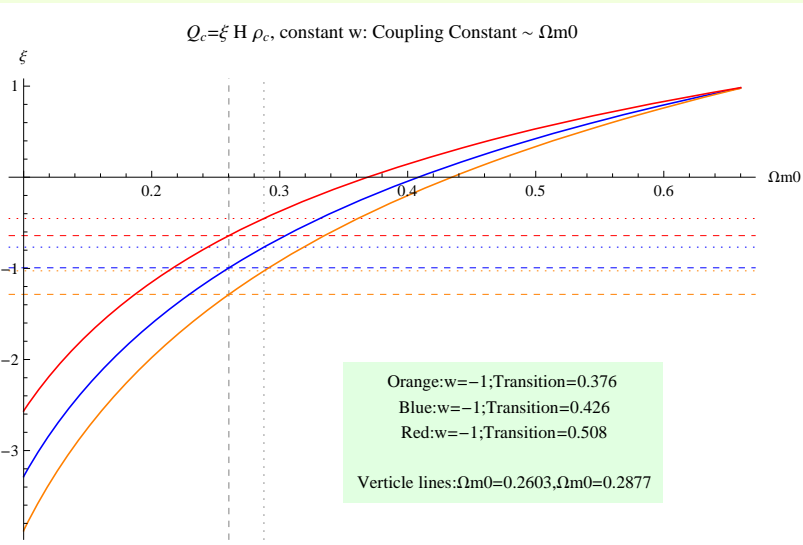
For $\Omega m0 \in 0.274 (1 \pm 0.05)$			
Table of ξ for different $\Omega m0 \sim$ Transition combination			
$\Omega m0 \sim$ Transition	0.426	0.376	0.508
0.2603	-0.994339	-1.28571	-0.641508
0.274	-0.877755	-1.15303	-0.544482
0.2877	-0.767582	-1.02756	-0.452892

Monotonic line.

In[327]:=

pltξvΩm0ICCSum

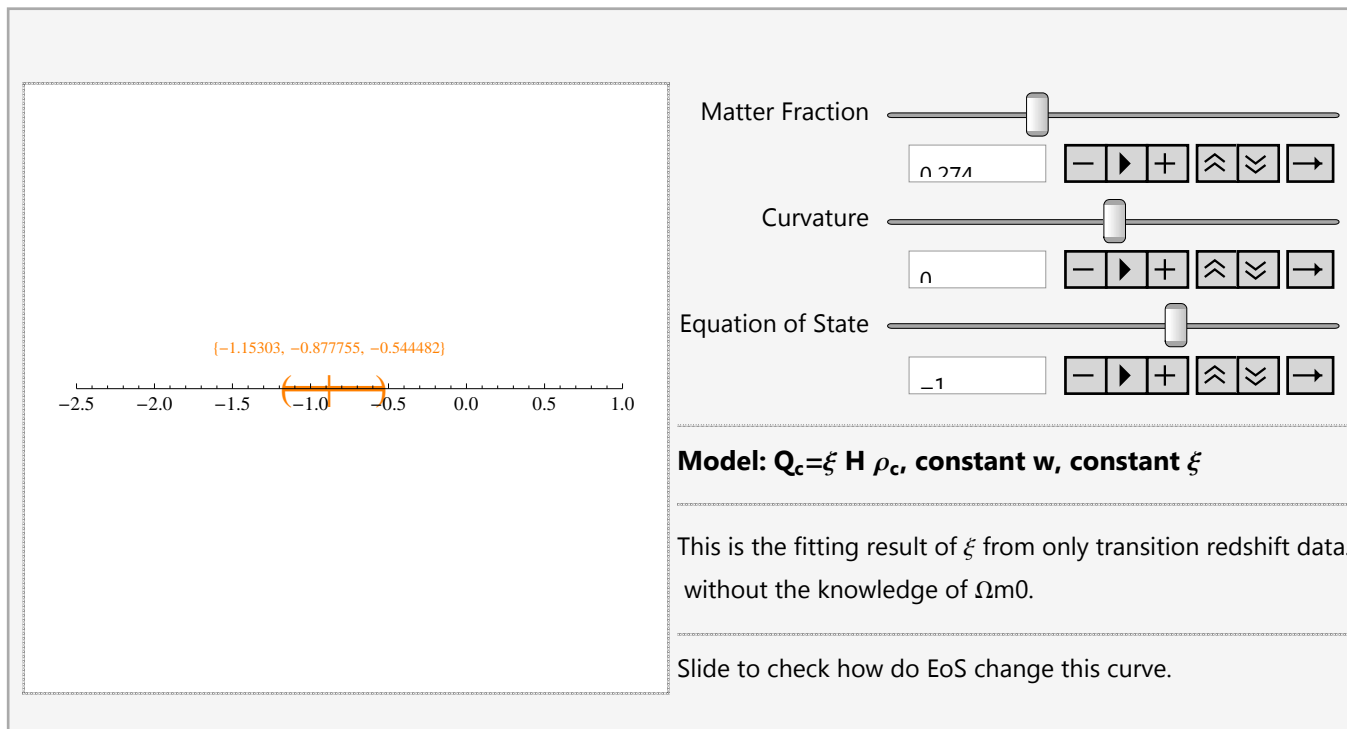
Out[327]=



Besides $\Omega m0$, we can also find out the effects of Curvature, EoS. Assuming we have a constrain of Transition redshift (0.376,0.508) with a center at 0.426.

In[328]:=

fitξ2ICCMansum



Out[328]=

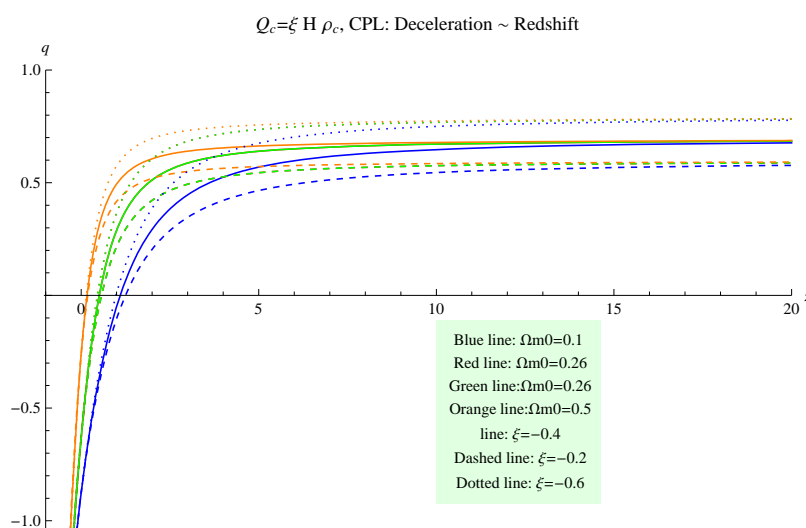
- Interacting model $Q_c = \xi H \rho_c$ with constant ξ and CPL parameterized EoS $w = w_0 + w_1 \frac{z}{1+z}$.

For a flat universe, choose the parameters $\{w_0=-1.02, w_1=0.6\}$, the region for interaction constant ξ should be $(-1.04, -0.21)$ with a center at -0.64 , i.e., $-0.64^{+0.42}_{-0.40}$, derived from the (transition redshift, Ω_{m0}) plane, while a result of $(-1.01, -0.23)$ with a center at -0.63 , i.e., $-0.63^{+0.40}_{-0.38}$, derived from (transition redshift, $\frac{\Omega_{m0}}{\Omega_{d0}}$) plane.

Deceleration parameter is shown below. Behaves similar to the constant ξ constant w situation.

In[329]:=

pldecICCPShowSum

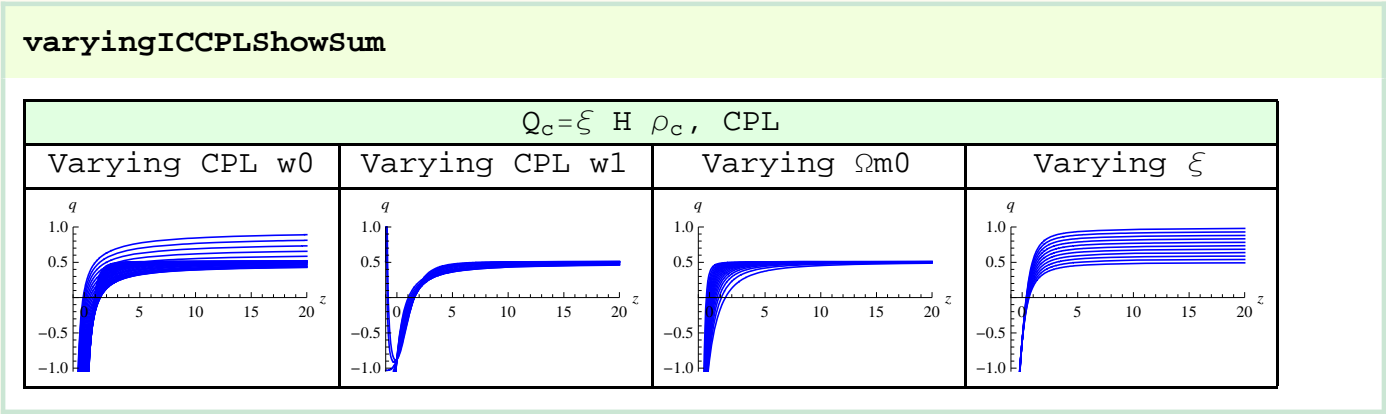


Out[329]=

The following plots show the effect of different parameters. Each plot shows how the deceleration parameter vs redshift line changes under uniformly distributed w_0, w_1, Ω_{m0} or ξ . w_0 moves the line up or down, but not monotonously; w_1 changes the late time behavior;

Ωm_0 changes the slope;
 ξ has moves the line up or down;

In[330]:=



Out[330]=

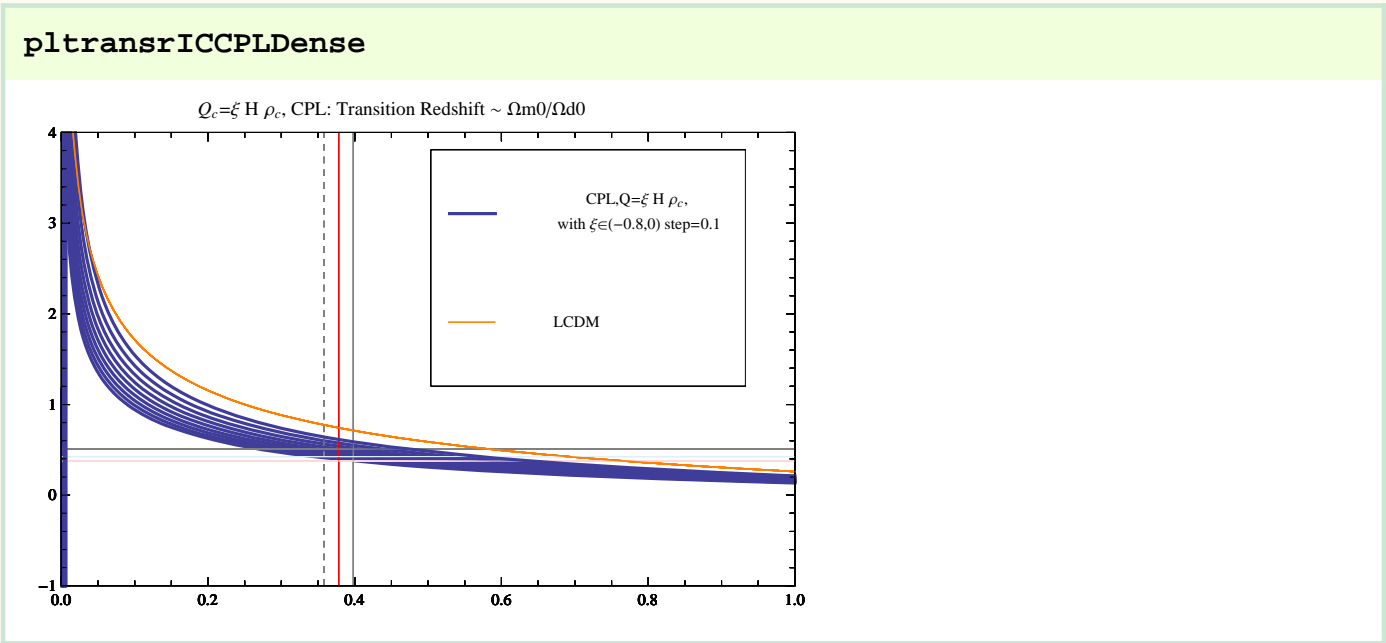
A toy to play with the $q \sim z$ plot. When $\xi=0$, $w_0=-1$, $w_1=0$, the curve reduced to LCDM curve.

In[331]:=



A plot shows how bad it is to use transition redshift to constrain interacting model. This is a CPL parameterized example. For $\xi \in (-0.8, 0)$, the line just stays near the allowed region constrained by Riess's results (References, Data From, 2).

In[332]:=



Out[332]=

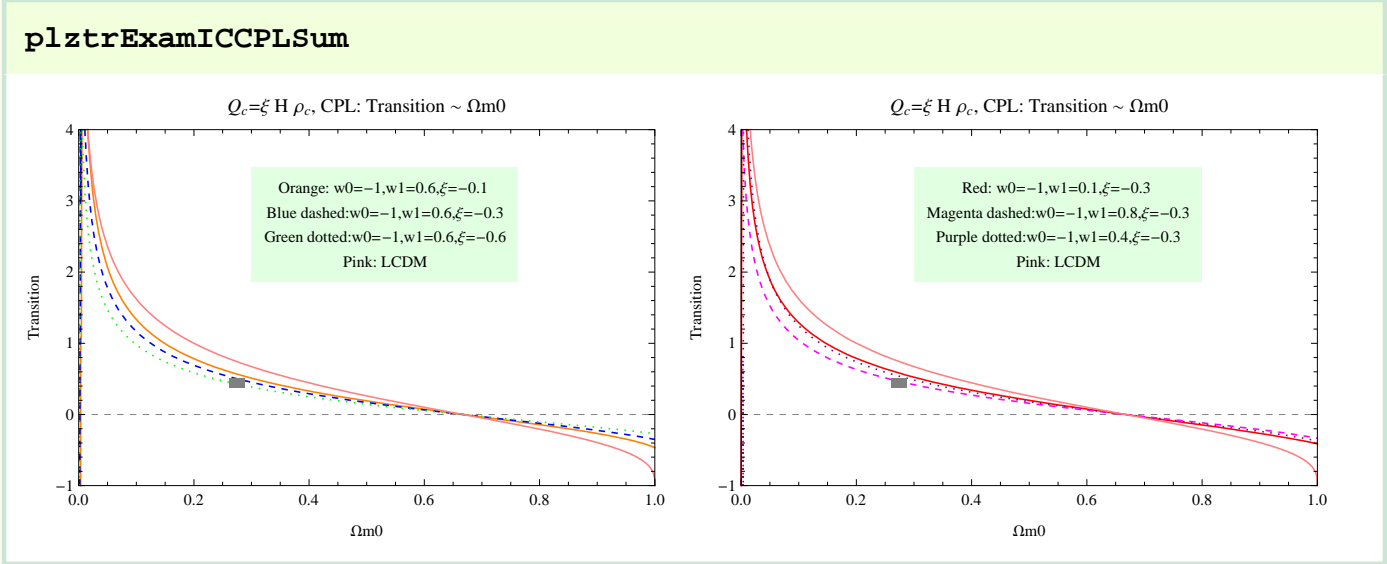
From the manipulate below, we find for $w_0=-1$, there is a point on this transition $\sim \Omega m_0$ curve do not change with coupling constant ξ and w_1 . (Well, what's the use of that...)

In[333]:=



An explicit proof of this statement.

In[334]:=



Out[334]=

A manipulate of the fitting results.

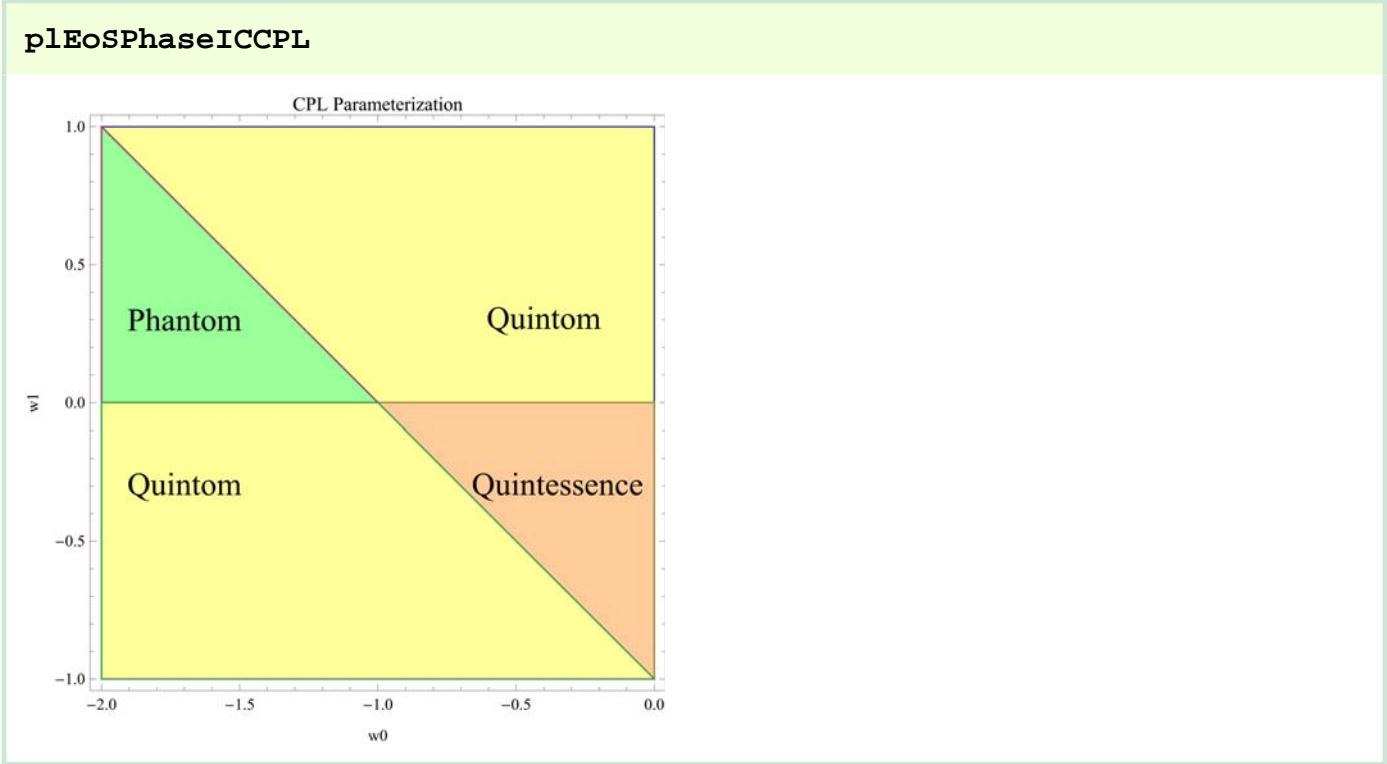
In[335]:=



Category

For different w0 and w1 in its EoS equation,

In[336]:=



Out[336]=

Quintom

Color illustrations for the following two figures.

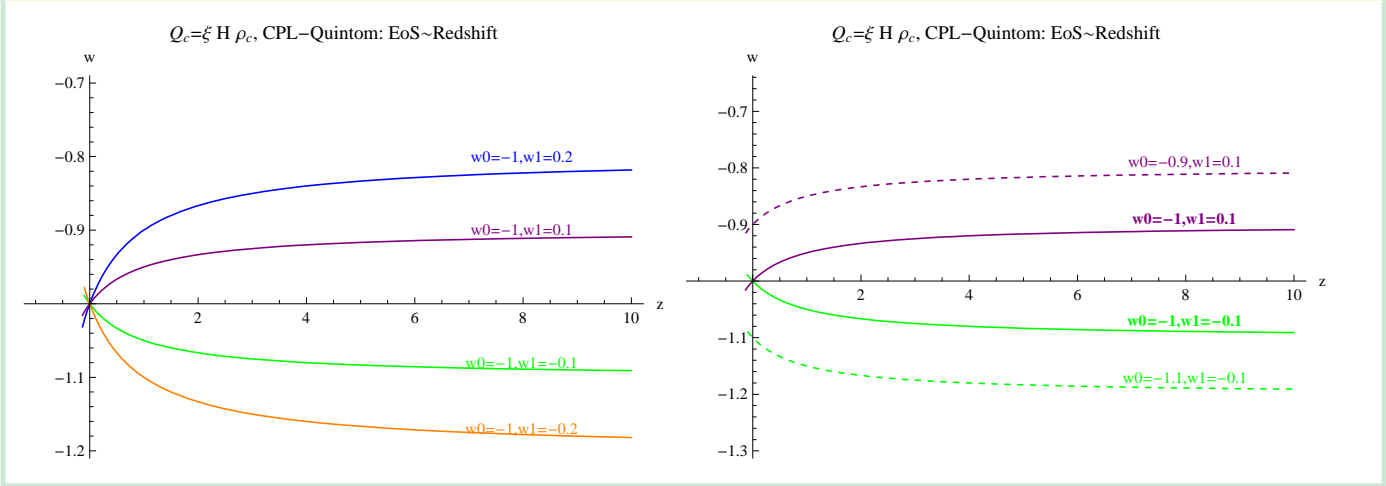
“Purple line:(-1,0.1). Blue line: (-1,0.2) Green line:(-1,-0.1)\n. Pink line:LCDM”

“Purple Dashed:(-0.9,0.1) Green Dashed:(-1.1,0.1)”

In[337]:=

plEoSICCPLQuintomSum

Out[337]=

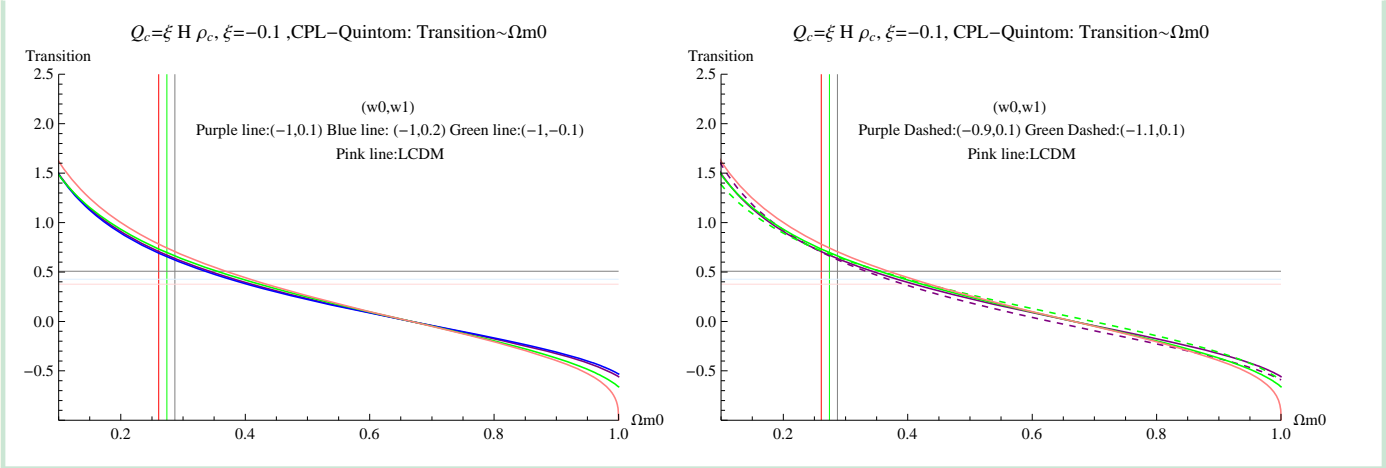


Plots of Transition redshift vs Ω_{m0} .
Legends are shown on the plots. Hard to distinguish from each other.

In[338]:=

plztrICCPLQuintomSum

Out[338]=



For **different EoS**, the fitting results are different. The following table and plots show how do w_0 and w_1 change the results.

In[339]:=

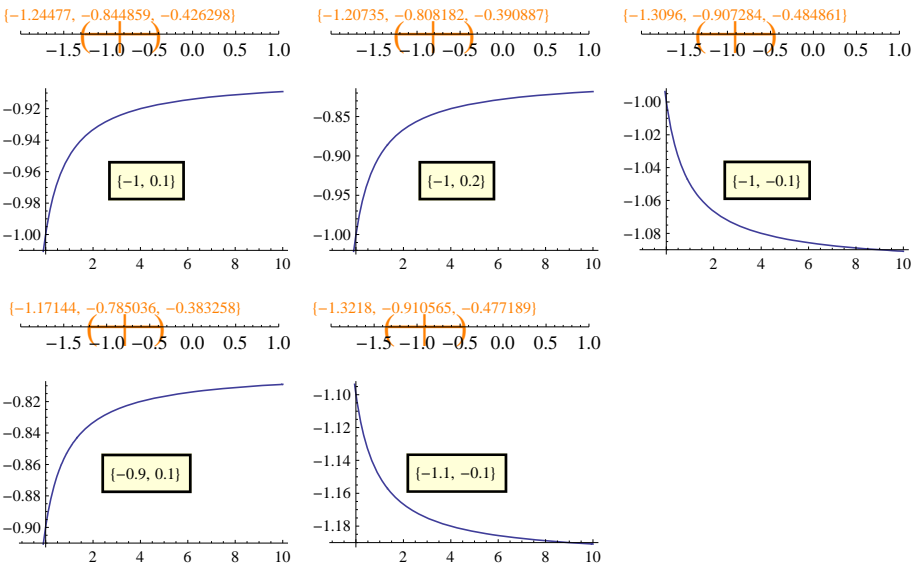
tab\xvwExamICCPLQuintom

Out[339]=

ξ results for $Q_c = \xi H \rho_d$, CPL, Quintom.			
$\{w_0, w_1\}$	Center	Lower	Upper
$\{-1, -0.1\}$	-0.907284	-1.3096	-0.484861
$\{-1, 0\}$	-0.877755	-1.27874	-0.457448
$\{-1, 0.1\}$	-0.844859	-1.24477	-0.426298
$\{-0.9, 0.1\}$	-0.785036	-1.17144	-0.383258
$\{-1.1, -0.1\}$	-0.910565	-1.3218	-0.477189

In[340]:=

pltfitξICCPLeQuintomSum

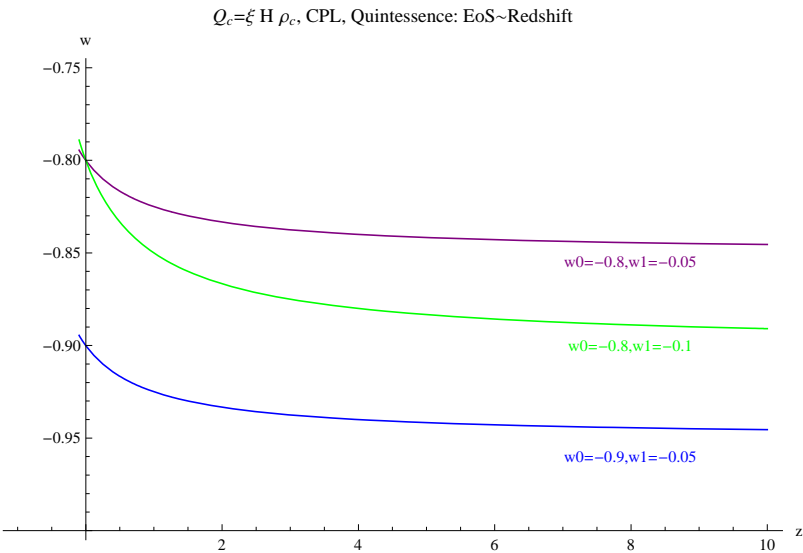


Out[340]=

▣ Quintessence

In[341]:=

plEoSICCPLeQuintessenceSum

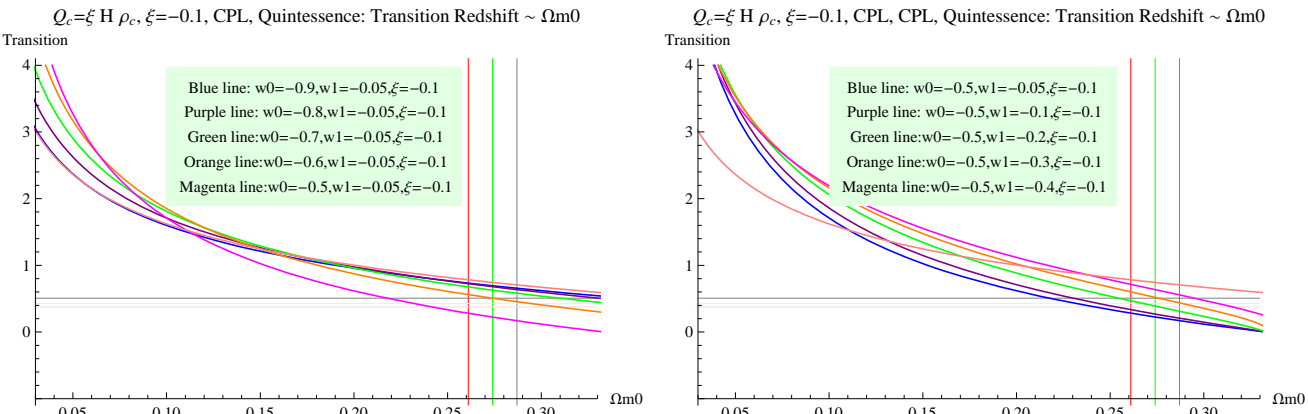


Out[341]=

Different from constant w results,

In[342]:=

plztrICCPLeQuintessenceSum



Out[342]=

Some ξ fitting results are shown below. This shows how do w_0 and w_1 change ξ results.

In[343]:=

`tabξvwExamICCPLQuintessence`

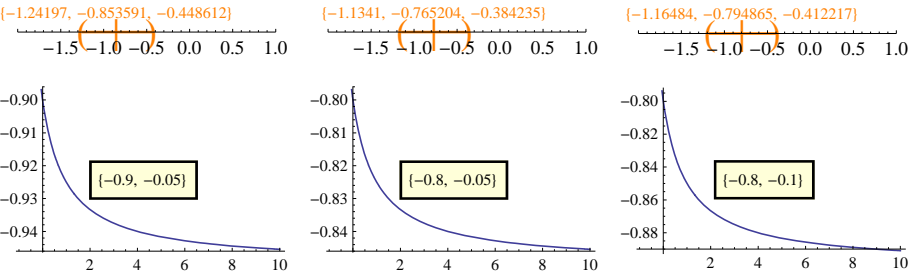
Out[343]=

ξ results for $Q_c=\xi H \rho_d$, CPL,Quintessence.			
$\{w_0,w_1\}$	Center	Lower	Upper
$\{-0.9,-0.05\}$	-0.853591	-1.24197	-0.448612
$\{-0.8,-0.05\}$	-0.765204	-1.1341	-0.384235
$\{-0.8,-0.1\}$	-0.794865	-1.16484	-0.412217

In[344]:=

`pltfitξICCPLQuintessenceSum`

Out[344]=

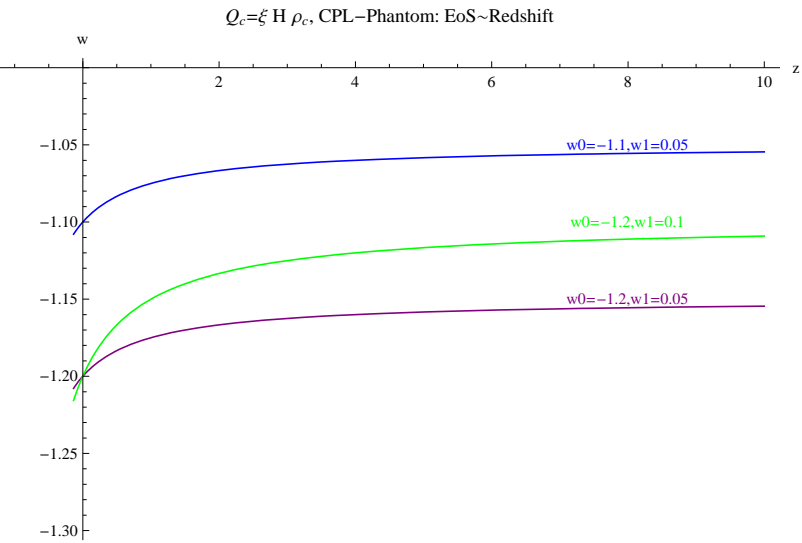


▣ Phantom

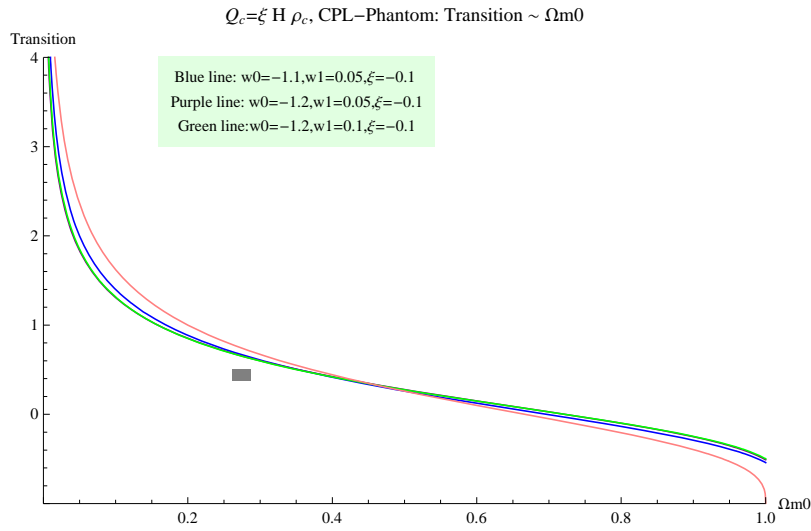
In[345]:=

`plEoSICCPLPhantomSum`

Out[345]=



In[346]:=

plztrICCPLPhantomSum

Out[346]=

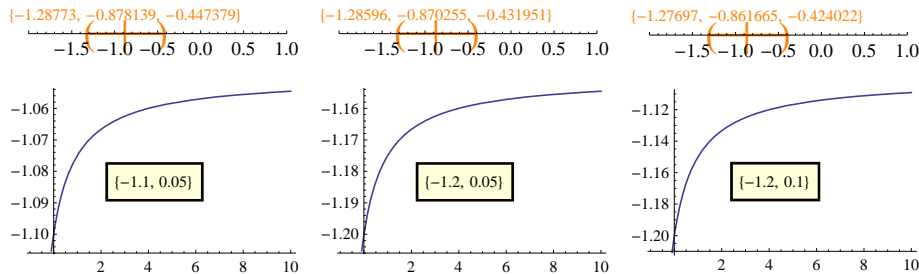
In[347]:=

tabxiwExamICCPLPhantom

ξ results for $Q_c = \xi H \rho_d$, CPL, Phantom.			
$\{w_0, w_1\}$	Center	Lower	Upper
$\{-1.1, 0.05\}$	-0.878139	-1.28773	-0.447379
$\{-1.2, 0.05\}$	-0.870255	-1.28596	-0.431951
$\{-1.2, 0.1\}$	-0.861665	-1.27697	-0.424022

Out[347]=

In[348]:=

pltfitxiICCPLPhantomSum

Out[348]=

■ Interacting model $Q_c = \xi H \rho_d$ with constant ξ and constant EoS w .

Derived from (transition redshift, Ω_{m0}) plane, the allowed region for coupling constant ξ is $(-1.06, -0.42)$ with a center at -0.76 , i.e., $-0.76^{+0.34}_{-0.30}$, taken the case that the universe is flat, and choose the EoS parameter $\{w=-1\}$.

Derived from the (transition redshift, $\frac{\Omega_{m0}}{\Omega_{d0}}$) plane, the allowed region of coupling constant ξ is $(-1.07, -0.41)$ with a center at -0.76 , i.e., $-0.76^{+0.35}_{-0.31}$.

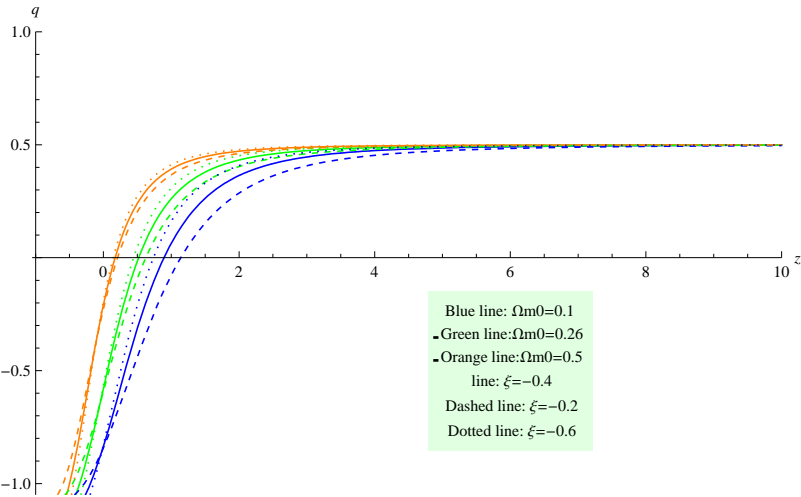
The plots of deceleration parameter are shown below. At the limit $z \rightarrow \text{Infinity}$, the deceleration parameter ALL goes to $\frac{1}{2}$.

Theoretically, this limit is $\frac{1}{2}$ which is not related to any parameters, with $3w + \xi < 0$.

In[349]:=

pldecI2CCShowSum

$Q_c=\xi\,H\,\rho_d$, constant w : Deceleration \sim Redshift

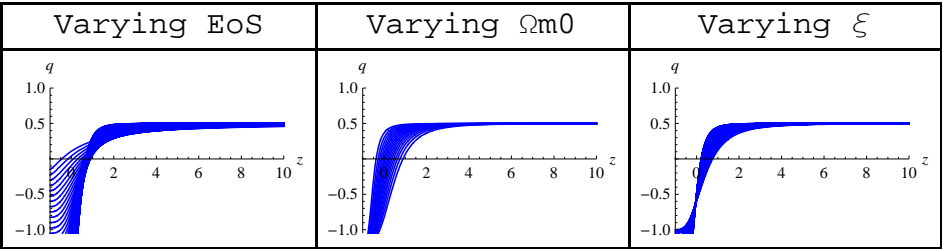


Out[349]=

To check the effect of different parameters, another plot is shown.

In[350]:=

varyingI2CCShowSum



Out[350]=

A toy to play with deceleration vs z curve is also provied

In[351]:=

pldecI2CCManSum

If the transition happens before $z=0$, increasing couling ξ will delay the transition.
If the it happens after $z=0$, increasing coupling ξ will bring forward the emergence of transition.

The following figure shows this result.

Gray rectangle is the region given by Riess.

Orange for $w=-1$
Blue for $w=-0.9$

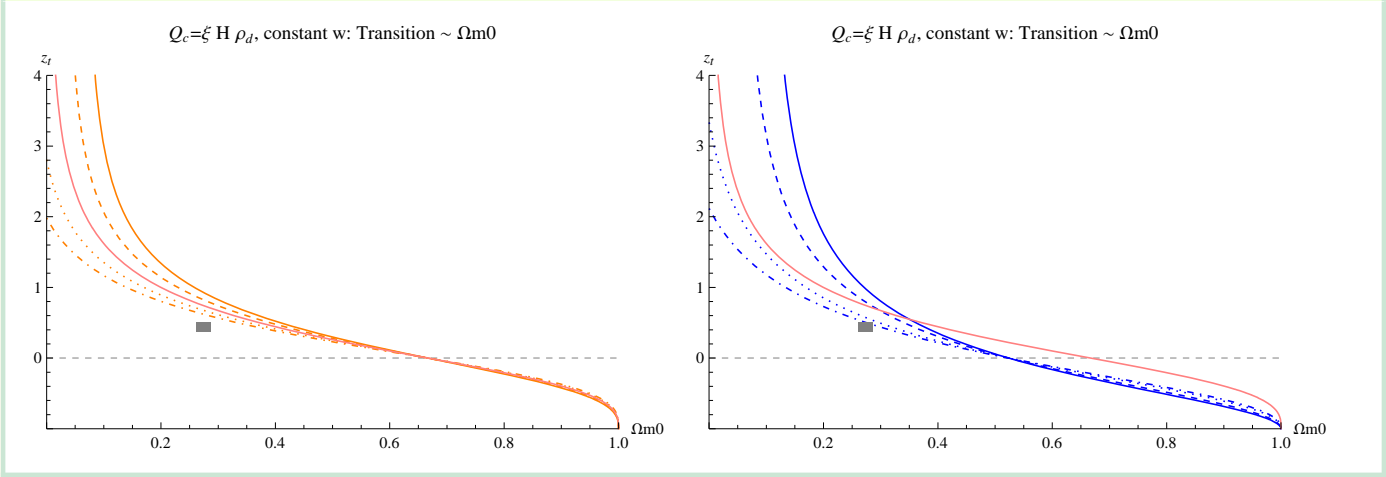
line: $\xi=0.2$
Dashed: $\xi=0.1$
Dotted: $\xi=-0.1$
DotDashed: $\xi=-0.2$

Pink line: $w=1$, $\xi=0$

In[352]:=

plztrvsΩm0I2CCSum

Out[352]=

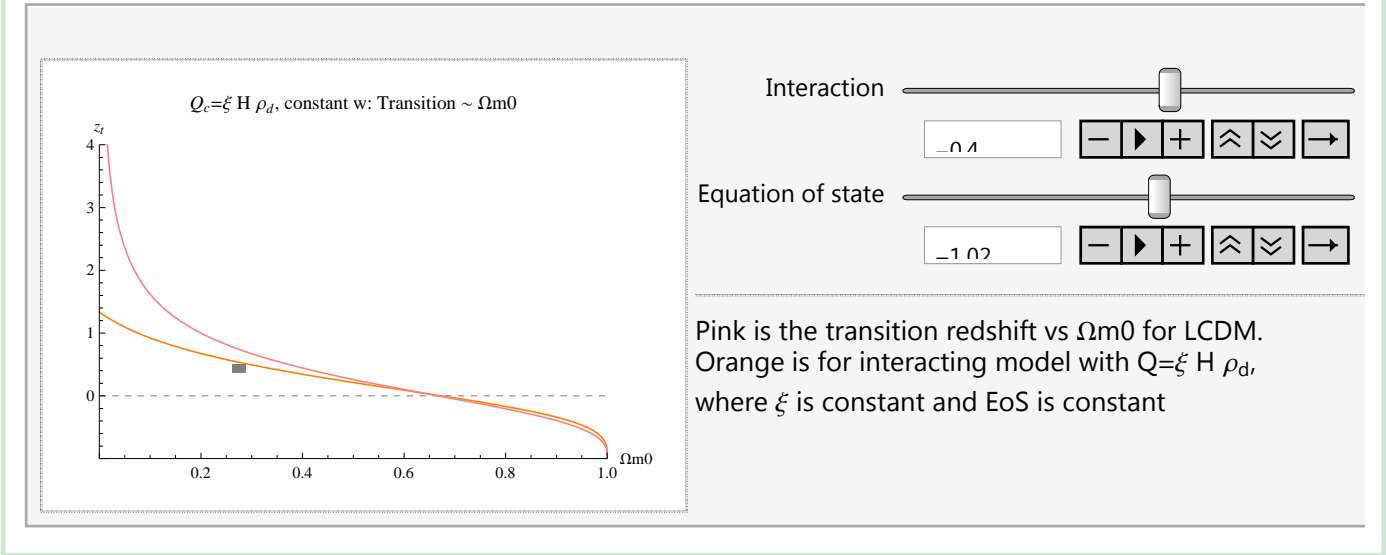


A toy of transition redshift. Gray rectangle is the allowed region of Ω_{m0} ~Transition redshift

In[353]:=

plztrI2CCManSum

Out[353]=



The fitting results of coupling constant ξ is

Or we can use some fitting results from WMAP etc. Take the example of $w=-1.087\pm0.096$.

In[356]:=

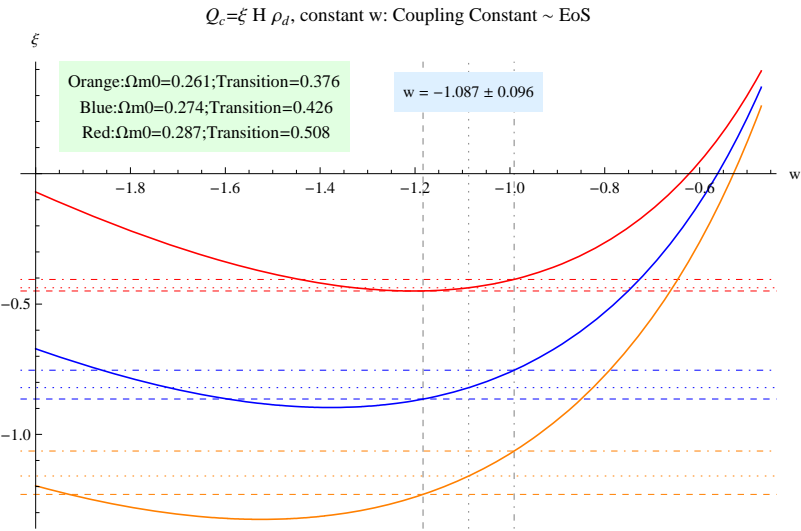
tabξvwExamI2CC

$Q_c=\xi H \rho_d, \text{Constant } w. \text{ (Data used:Data From, 2)}$			
w	Center	Lower	Upper
-1.183	-0.864289	-1.22984	-0.449552
-1.087	-0.820486	-1.15946	-0.437339
-0.991	-0.753634	-1.06346	-0.405262

Out[356]=

In[357]:=

pltξvwExamI2CC

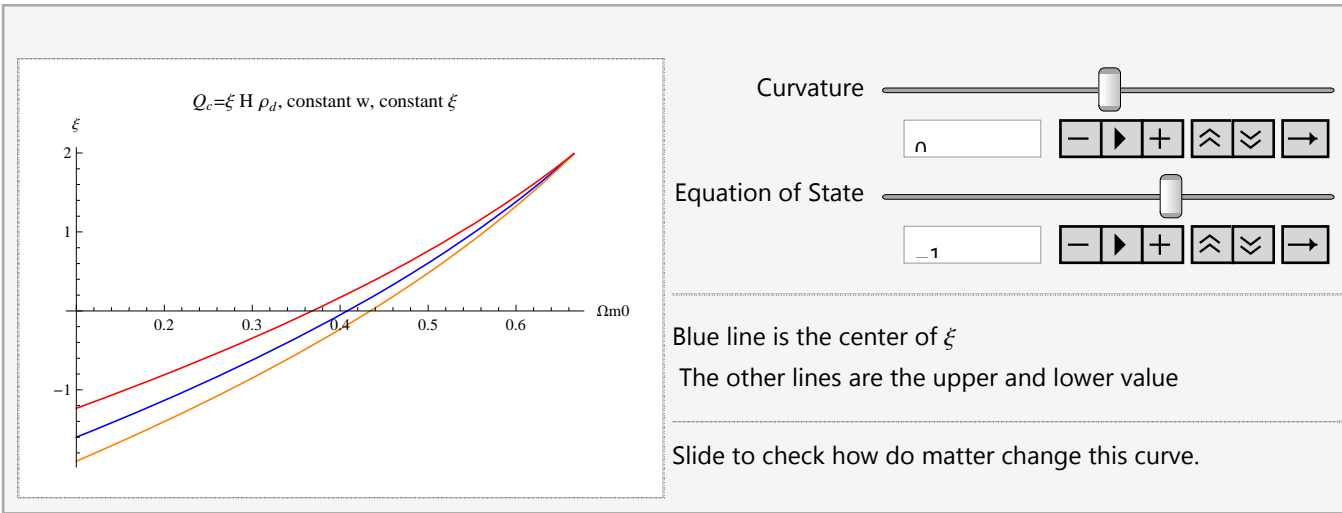


Out[357]=

Now we assume we do not have the observed Ω_{m0} data, how do this Ω_{m0} change the result. In other words, if the observed Ω_{m0} data float around some value, then how is the fitting result? We also consider the curvature.

In[358]:=

pltξvΩm0I2CCManSum



Out[358]=

If Ω_{m0} varies 0.05 percent from 0.274,

In[359]:=

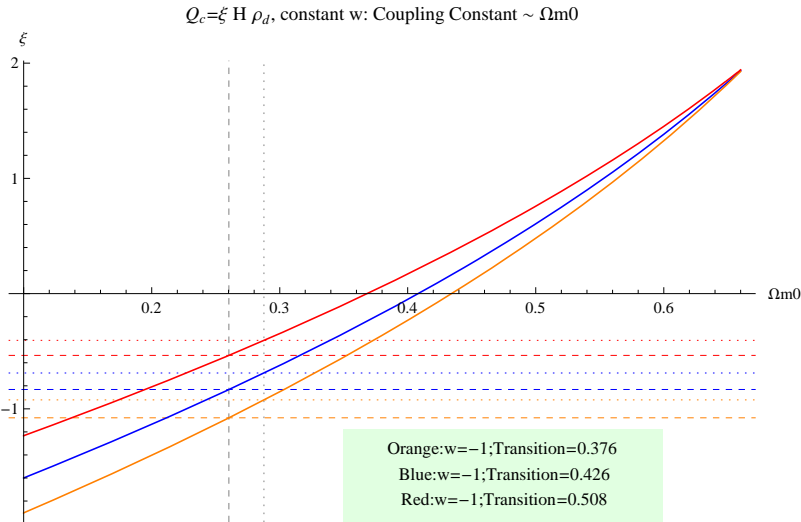
tabξI2CCSum

For $\Omega_{m0} \in 0.274 (1 \pm 0.05)$			
Table of ξ for different Ω_{m0} ~Transition combination			
Ω_{m0} ~Transition	0.426	0.376	0.508
0.2603	-0.832284	-1.07758	-0.53584
0.274	-0.760999	-1.00068	-0.471298
0.2877	-0.688664	-0.922602	-0.40585

Out[359]=

In[360]:=

pltξvΩm0I2CCSum



Out[360]=

In addition, we can also find out the effects of Curvature, EoS. Assuming we have a constrain of Transition redshift (0.376,0.508) with a center at 0.426.

In[361]:=

fitξ2I2CCManSum

$(-1.00068, -0.760999, -0.471298)$

Matter Fraction

Curvature

Equation of State

$Q_c = \xi H \rho_d$, constant w , constant ξ .

This is the fitting result of ξ from only transition redshift data without the knowledge of Ω_{m0} .

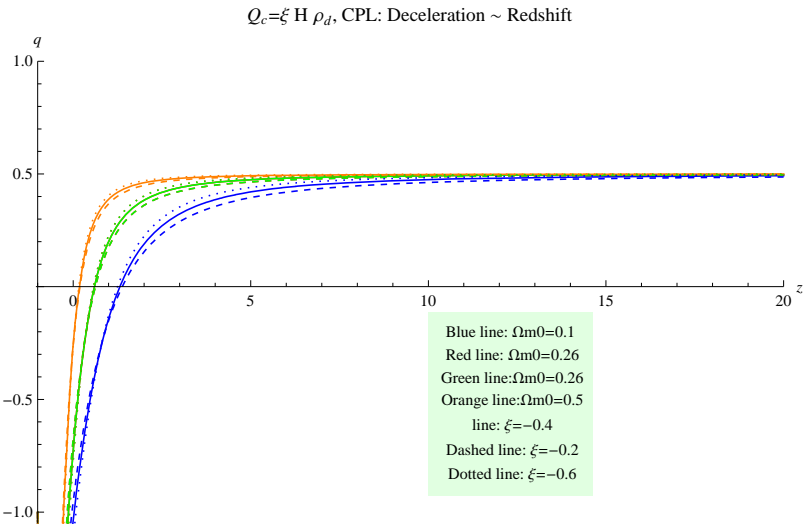
Slide to check how do EoS change this curve.

Out[361]=

■ Interacting model $Q_c = \xi H \rho_d$ with constant ξ and CPL parameterization.

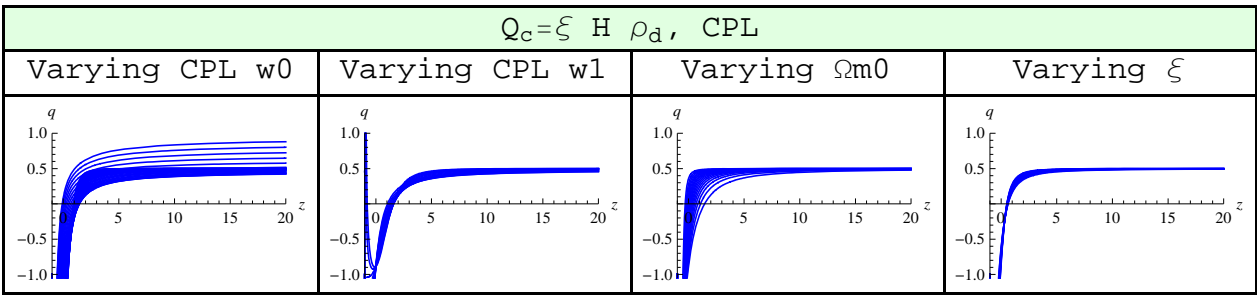
In[362]:=

pldecI2CCPLShowSum



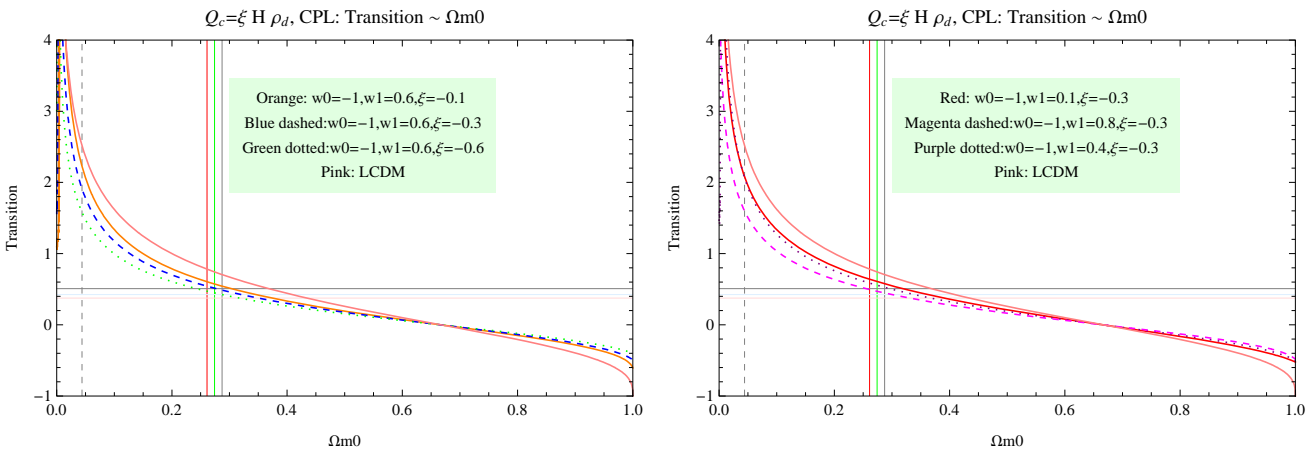
In[363]:=

varyingI2CCPLShowSum



In[364]:=

plztrExamI2CCPLSum

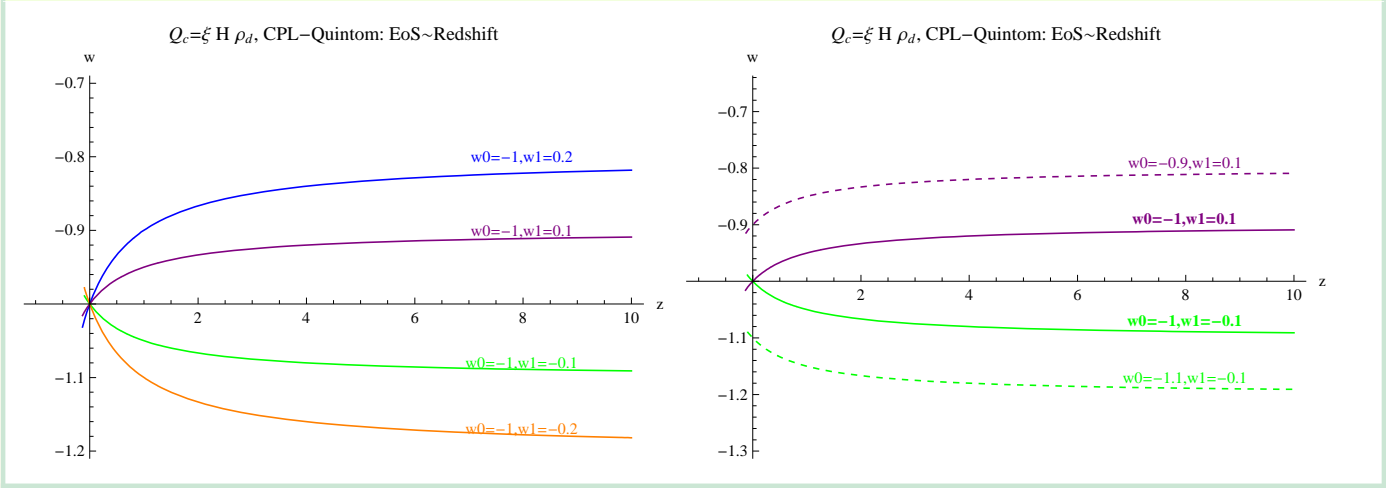


▣ Quintom

In[365]:=

plEoSII2CCPLQuintomSum

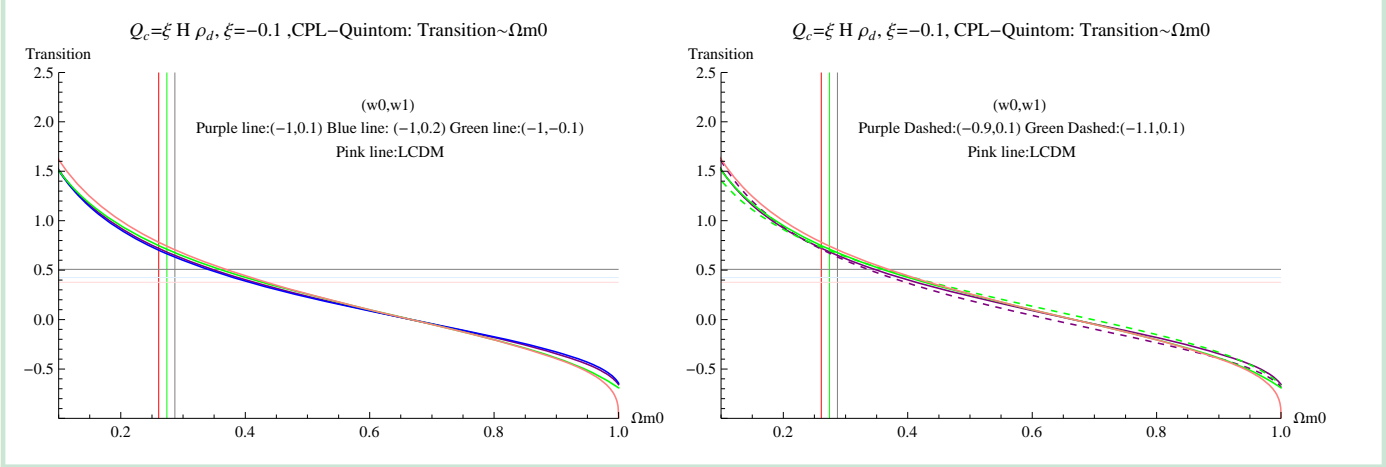
Out[365]=



In[366]:=

plztrII2CCPLQuintomSum

Out[366]=



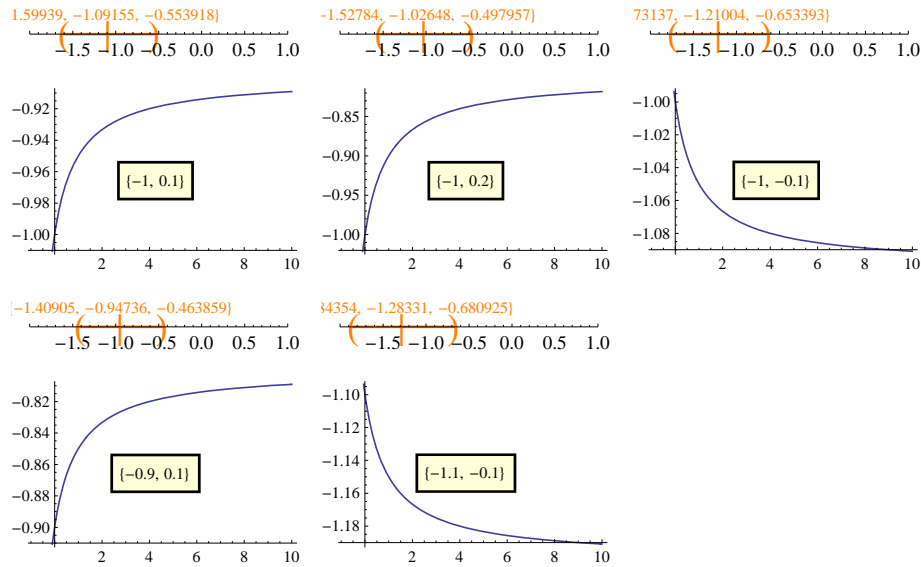
In[367]:=

tabξvwExamI2CCPLQuintom

Out[367]=

ξ results for $Q_c = \xi H \rho_d$, CPL, Quintom.			
{w0,w1}	Center	Lower	Upper
{-1, -0.1}	-1.21004	-1.73137	-0.653393
{-1, 0}	-1.15265	-1.66715	-0.605615
{-1, 0.1}	-1.09155	-1.59939	-0.553918
{-0.9, 0.1}	-0.94736	-1.40905	-0.463859
{-1.1, -0.1}	-1.28331	-1.84354	-0.680925

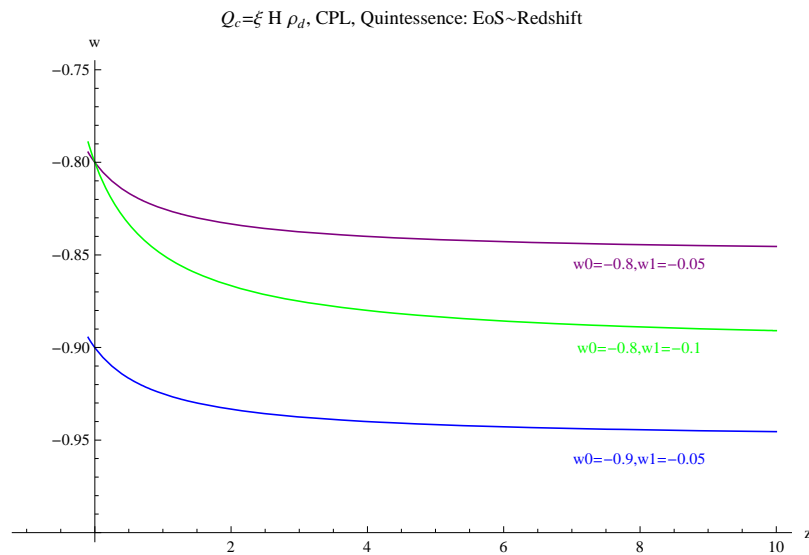
In[368]:=

pltfitξI2CCPLQuintomSum

Out[368]=

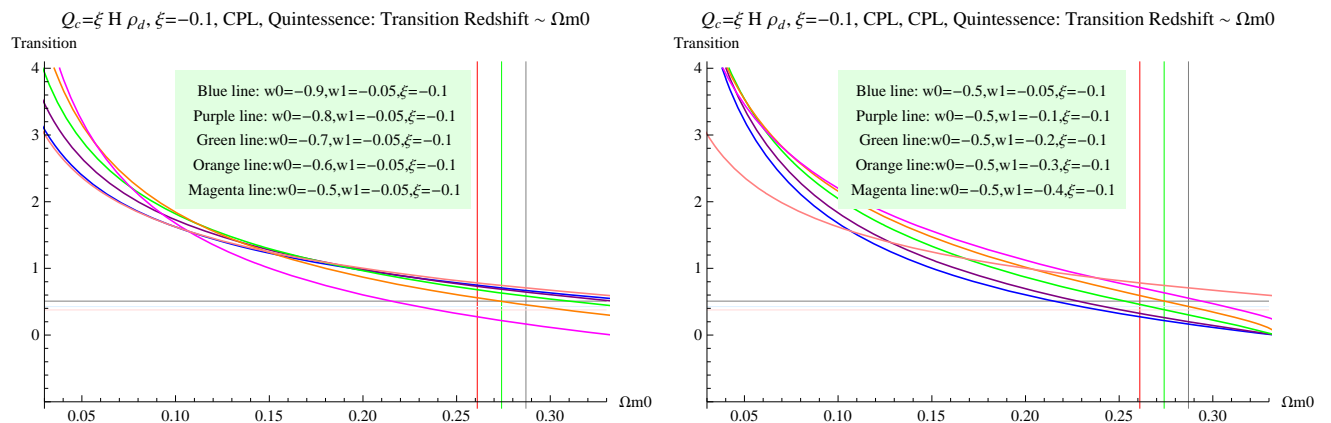
□ Quintessence

In[369]:=

plEoS I2CCPLQuintessenceSum

Out[369]=

In[370]:=

plztr I2CCPLQuintessenceSum

Out[370]=

In[371]:=

tabξvwExamI2CCPLQuintessence

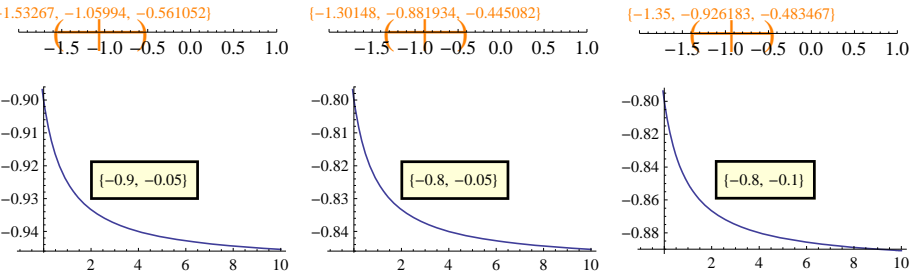
Out[371]=

ξ results for $Q_c=\xi$ H ρ_d , CPL,Quintessence.			
{w0,w1}	Center	Lower	Upper
{-0.9, -0.05}	-1.05994	-1.53267	-0.561052
{-0.8, -0.05}	-0.881934	-1.30148	-0.445082
{-0.8, -0.1}	-0.926183	-1.35	-0.483467

In[372]:=

pltfitξI2CCPLQuintessenceSum

Out[372]=

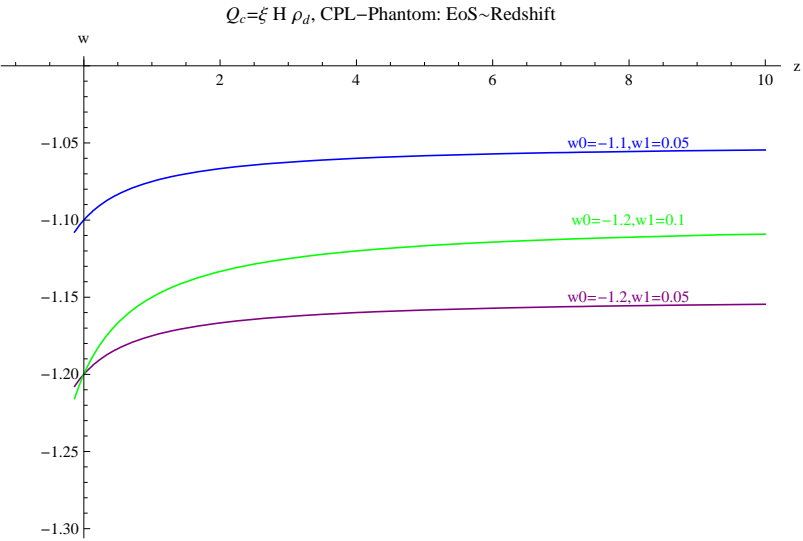


▣ Phantom

In[373]:=

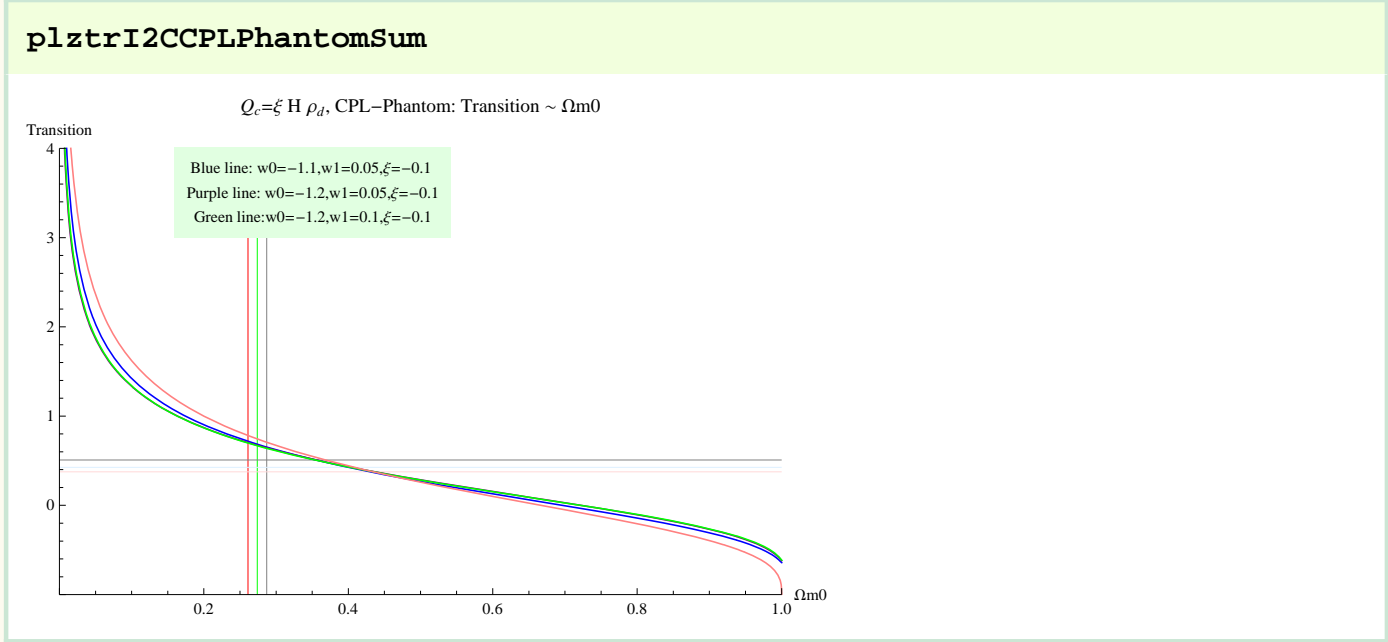
plEoS I2CCPLPhantomSum

Out[373]=



In[374]:=

Out[374]=



In[375]:=

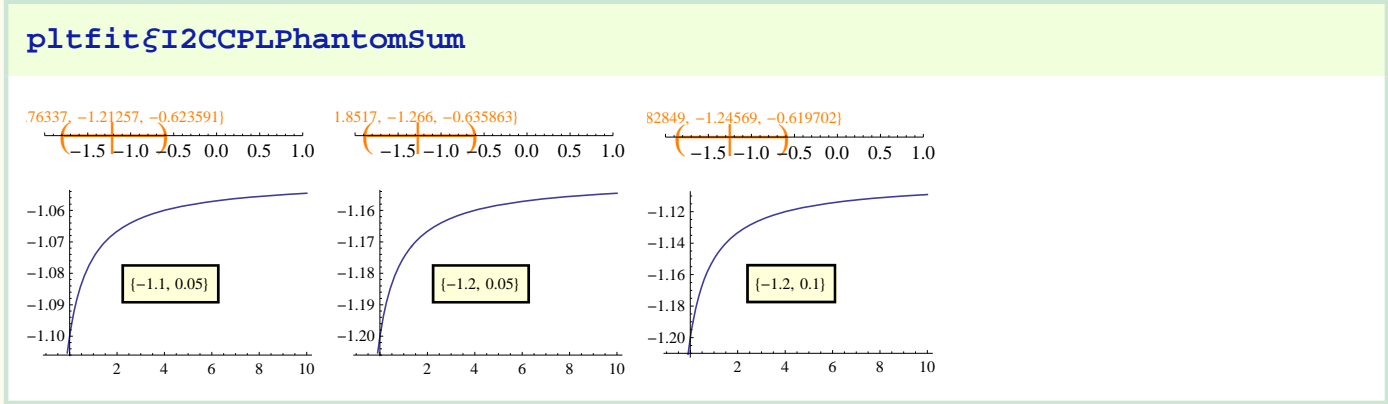
Out[375]=

tabxivwExamI2CCPLPhantom

ξ results for $Q_c = \xi H \rho_d$, CPL, Phantom.			
$\{w_0, w_1\}$	Center	Lower	Upper
$\{-1.1, 0.05\}$	-1.21257	-1.76337	-0.623591
$\{-1.2, 0.05\}$	-1.266	-1.8517	-0.635863
$\{-1.2, 0.1\}$	-1.24569	-1.82849	-0.619702

In[376]:=

Out[376]=



■ References

▣ Data From

1. CPL data

Combining SN1a, BAO 3, WMAP5, $H(z)$ (From arXiv:0909.0596)

$\Omega_{m0} = 0.269^{+0.017}_{-0.008}$, $w_0 = -0.97^{+0.12}_{-0.07}$, $w_1 = 0.03^{+0.26}_{-0.75}$.

2. LCDM

From WMAP: $\Omega_{m0}=0.265$
arXiv:astro-ph/0611572, Riess et al :
arXiv:1205.4688 : $\Omega_{m0}=0.247$ (+0.013, -0.013) and Transition 0.426 (+0.082, -0.050)

3. Equation of state

From arXiv:1202.0545v1

$w = -1.087 \pm 0.096$.