

Power Spectrum & Its Evolution

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Figure 1 shows several ratios of the Hubble distance.

Figure 2 and figure 3 are the growth functions.

Figure 4 is the power spectra.

Figure 5 shows the normalization factors, which are defined as

$$Q^2 = \frac{\text{Power spectra of DE model with EoS } w}{\text{Power spectra of LCDM model}}, \text{ at some wavenumber } k. \quad (1)$$

XXXX

XXXX

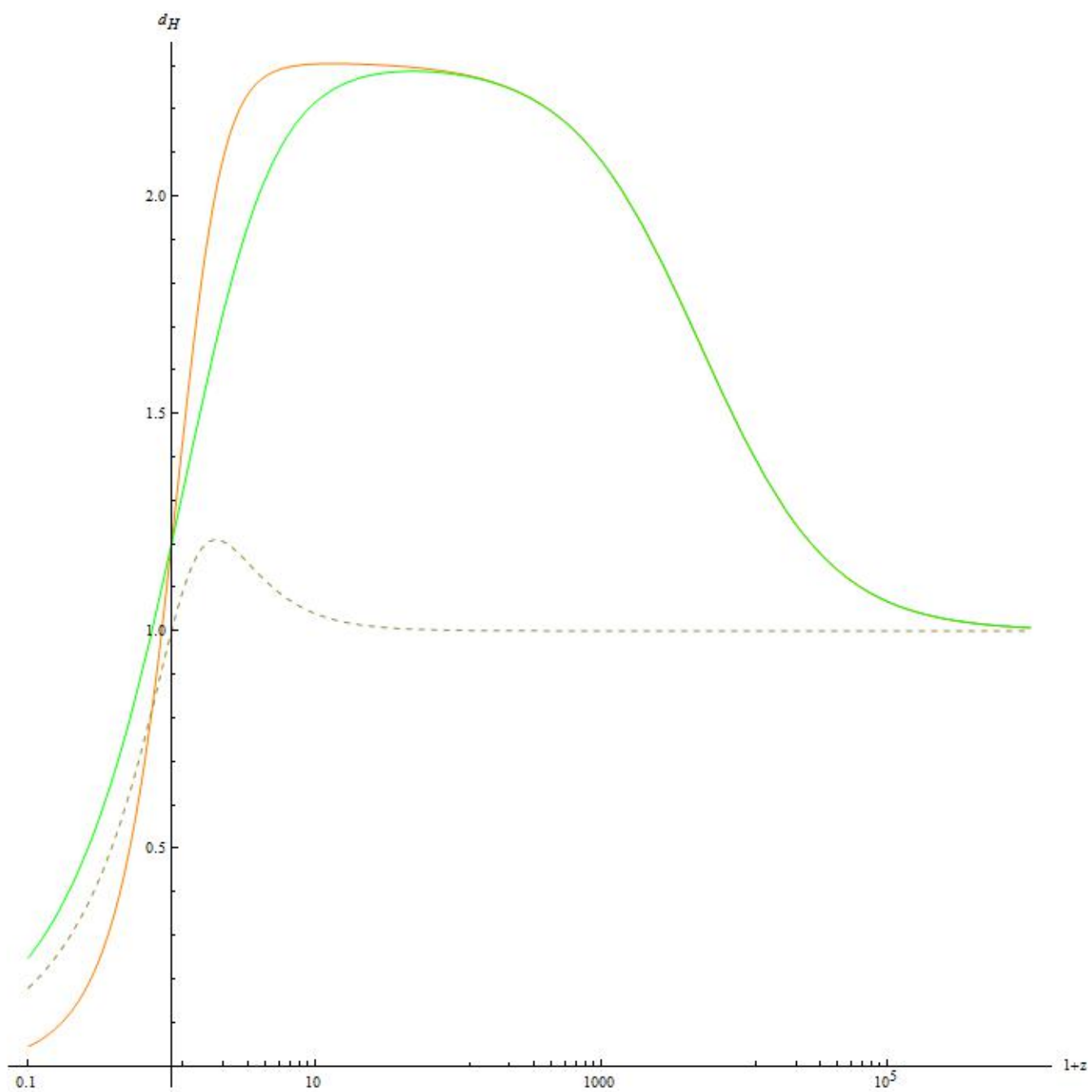


Figure 1: Hubble distance with a legend

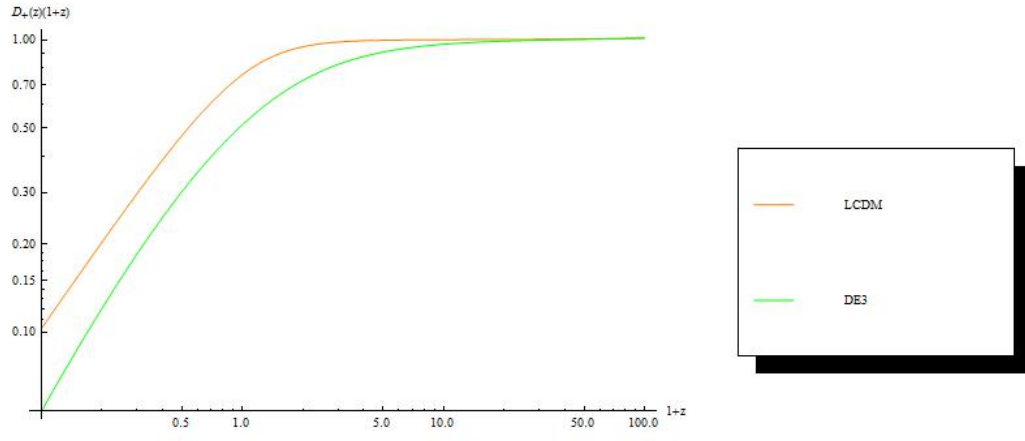


Figure 2: Growth function

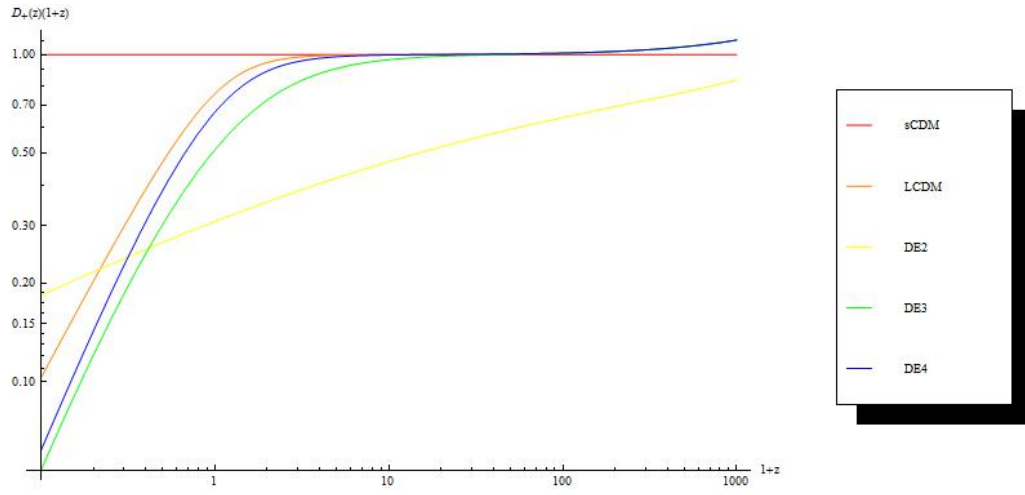


Figure 3: Growth function

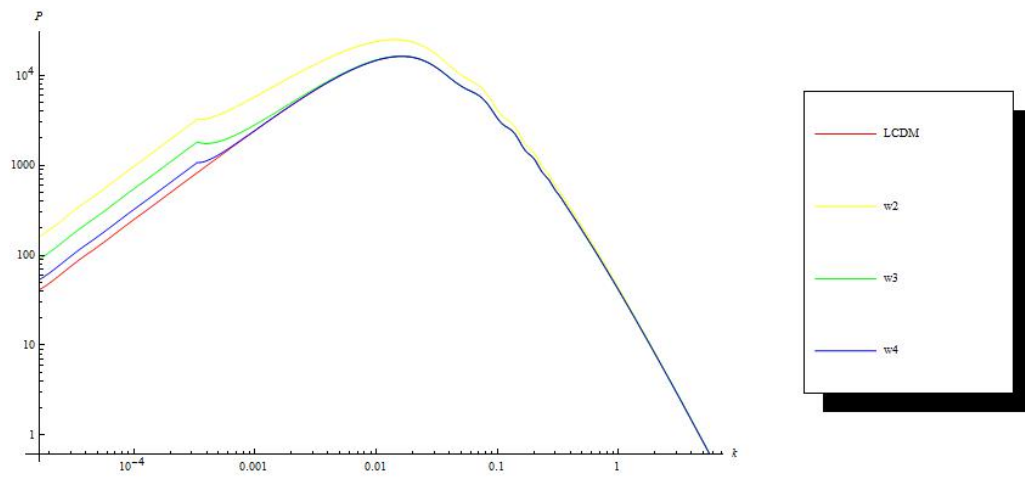


Figure 4: Power spectrum

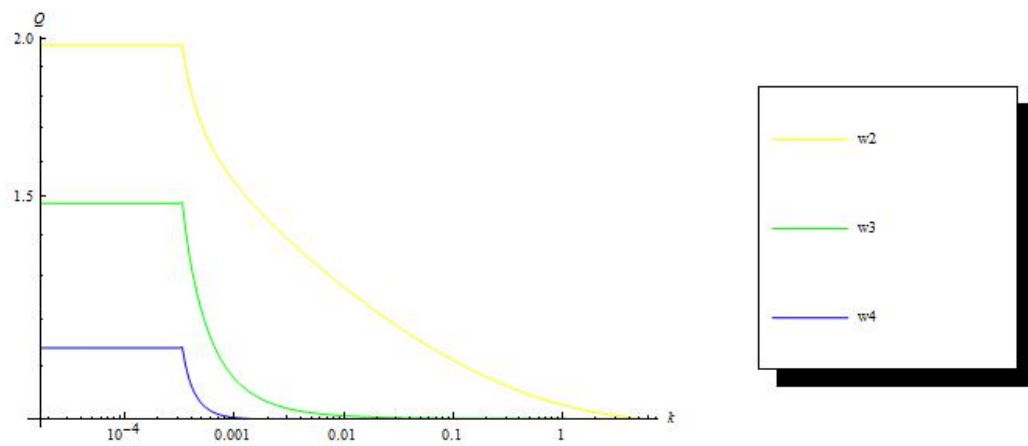


Figure 5: Q factors