INTRO

- \heartsuit This is a program for χ^2 fitting, using SN1a data.
- ♡ SN data are taken from http://supernova.lbl.gov/Union/

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PRE

Settings

Include the packages needed.

```
Needs["ErrorBarPlots`"]
<< PhysicalConstants`</pre>
```

Set work directory. Modify it to your own directory before runing this program. DATA folder are placed in this directory. DATA folder should contain the SN data files.

```
SetDirectory[
   "E:\\Nutshare\\Store\\Projects\\Growth\\git\\CoChiSquare"];
```

■ Conventions, parameters, etc

```
Line element: ds^2 = -dt^2 + a[t]^2 \left( \frac{dr^2}{1 - kr^2} + r^2 (d\theta^2 + Sin[\theta]^2 d\phi^2) \right)
```

 Ω m0: matter fraction

Ωd0: dark energy/cosmological constant fraction

 Ω r0: radiation fraction

 $\Omega \text{m0w:}$ matter fraction given by WMAP

 Ω d0w: dark energy/cosmological constant fraction given by WMAP

Ωr0w: radiation fraction given WMAP

```
\Omega m0w = 0.27;
\Omega d0w = 0.73;
\Omega r0w = 8.09 * 10^{-5};
```

■ Basic Equations in LCDM

H0: Hubble constant in unit of km/s/Mpc

c: speed of light

```
H0 = 71;

c = SpeedOfLight * Second

1000 Meter

gra = GravitationalConstant

149 896 229

500
```

$$\frac{6.67428 \times 10^{-11} \; \text{Meter}^2 \; \text{Newton}}{\text{Kilogram}^2}$$

 $hubble[\Omega m0_,\Omega d0_,\Omega k0_,z_] : Hubble \ functions \ in \ LCDM.$

H[z]: an example of hubble function with given parameters.

$$\begin{aligned} & \text{hubble}[\Omega \text{m0}_? \text{NumberQ}, \ \Omega \text{d0}_? \text{NumberQ}, \ \Omega \text{k0}_? \text{NumberQ}, \ \textbf{z}_] = \\ & \text{H0} \ \sqrt{\Omega \text{m0} \ (1+z)^3 + \Omega \text{d0} + \Omega \text{k0} \ (1+z)^2} \ ; \\ & \text{H[z}_] = & \text{hubble}[\Omega \text{m0w}, \ \Omega \text{d0w}, \ 0, \ z]; \end{aligned}$$

 $q[\Omega m0_,\Omega d0_,\Omega k0_,z_]: Deceleration parameter. \ DEF: \ q[\Omega m0_,\ \Omega d0_,\ \Omega k0_,\ z_] = \frac{-1}{H[z]} \ \frac{D[D[a[t],t],t]}{D[a[t],t]}$

$$q[\Omega m0_{,\Omega d0_{,\Omega k0_{,z}}] = \frac{(1+z)}{hubble[\Omega m0, \Omega d0, \Omega k0, z]} D[hubble[\Omega m0, \Omega d0, \Omega k0, z];$$

(General) Friedmann equations are

$$\frac{3(a[t]' + k)}{a[t]^2} = 8\pi G \rho$$

$$2\frac{a[t]''}{a[t]} + \frac{a[t]' + k}{a[t]^2} = -8\pi G p$$

Since "zt" will be used as a parameter in most cases, we define "ztr" as a name for transition redshift, "ztrr" for transition redshift with a parameter r.

ztr[Ω m0_, Ω d0_,z_]: Transition redshift in LCDM with Ω k0 not zero with full parameters

ztrr[r_,z_]: with
$$r = \frac{\Omega \text{m0}}{\Omega \text{d0}}$$

ztr[
$$\Omega m0_{-}$$
, $\Omega d0_{-}$] = $\left(2 \frac{\Omega d0}{\Omega m0}\right)^{1/3} - 1$;
ztrr[r_{-}] = $\left(\frac{2}{r}\right)^{1/3} - 1$;

Regard "zt" as a paramter, solve out Ω d0 and Ω k0=1- Ω m0- Ω d0

```
\Omega d0 = \frac{1}{2} \Omega m0 (1 + zt)^{3}
\Omega k0 = 1 - \Omega m0 - \frac{1}{2} \Omega m0 (1 + zt)^{3}
(1)
```

CAL

- Plot SN data with error bar
- Import and display "SCPUnion2_AllSNe.tex"

Read the data file of SCPUnion2_mu_vs_z.txt

```
dataunionfulltmp =
   Import["DATA\\UNION2\\SCPUnion2_AllSNe.tex", "Table"];
```

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Length of dataunionfulltmp

```
lenfull = Length[dataunionfulltmp];
```

Delete the & simbols in dataunionfulltmp

 $V \hat{z} \not = \mathring{S} \mathring{y} \mathring{o} \mathring{S} \mathring{A}_{z} \not = \mathring{o} \mathring{A}_{z} \not = \mathring{o} \mathring{A}_{z} \mathring{p} \mathring{p}_{z} \mathring{h} \mathring{o} \mathring{o} \mathring{c} \mathring{c} \mathring{a}_{z} \mathring{a$

Import and display "SCPUnion2_mu _vs _z _adjust.txt"

Import dataunion adjusted data. This data includes name, redshift, distance, error.

```
dataunion =
   Import["DATA\\UNION2\\SCPUnion2_mu_vs_z_adjust.txt", "Table"];
datauniontable = Grid[Join[{{Name, Redshift, Distance, Error}},
   dataunion], Alignment → Left, Spacings → {2, 1}, Frame → All,
   ItemStyle → "Text", Background → {{Gray, None}, {LightGray, None}}];
```

Show dataunion table.

```
datauniontable;
```

Length of this dataunion table

```
len = Length[dataunion];
```

Simplify dataunion table and call it du

```
du = Table[{dataunion[[i, 2]],
          dataunion[[i, 3]], dataunion[[i, 4]]}, {i, 1, len}];
dutable = Grid[Join[{{Redshift, Distance, Error}}, du],
    Alignment \rightarrow Left, Spacings \rightarrow {2, 1}, Frame \rightarrow All,
    ItemStyle \rightarrow "Text", Background \rightarrow {{Gray, None}, {None, None}}];
```

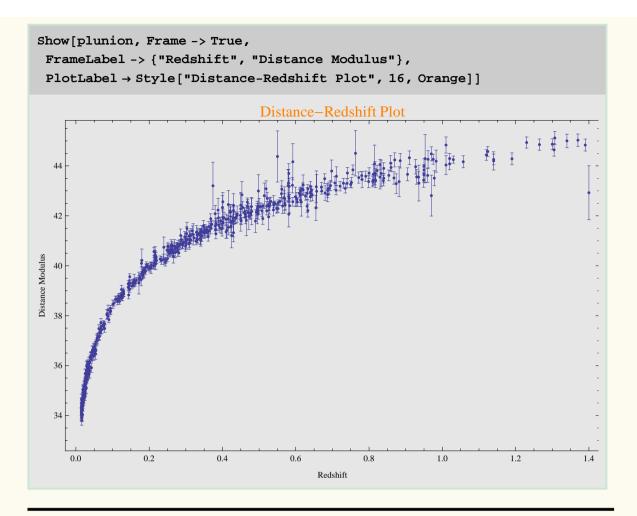
Show du table

```
dutable;
```

Visualize data.

Plot du table. That is SN

```
plunion = ErrorListPlot[du, PlotRange → All];
```



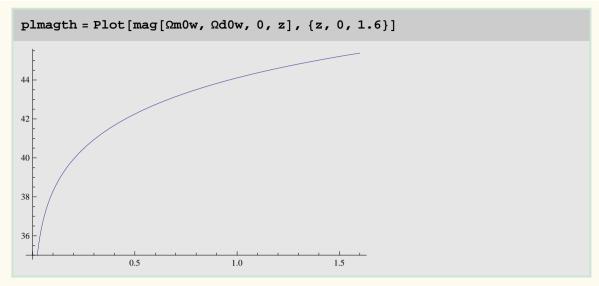
- Transition redshift, deceleration parameter, theoretically.
- Theoretical Function and Values of distance/mag at different redshift.

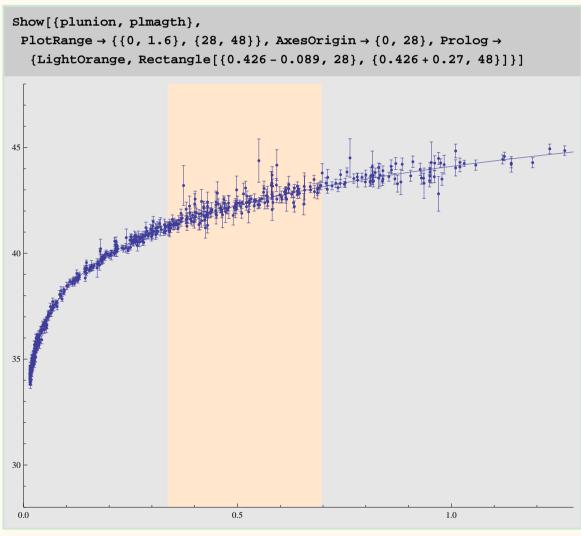
Distance at redshift z

```
 \begin{split} &\text{dz}\left[\Omega\text{m0}_?\text{NumberQ},\ \Omega\text{d0}_?\text{NumberQ},\ \Omega\text{k0}_?\text{NumberQ},\ z\_?\text{NumberQ}\right] := \\ & (1+z)\ \text{NIntegrate}\Big[\frac{1}{\text{hubble}\left[\Omega\text{m0},\ \Omega\text{d0},\ \Omega\text{k0},\ \text{tmp}\right]},\ \{\text{tmp, 0, z}\}\Big]; \end{split}
```

Magnitude calculated from distance. Theoretically of cause.

```
\label{eq:log_constraint} $$ mag[\Omega m0_?NumberQ, \Omega d0_?NumberQ, z_?NumberQ] := $$ $$ Log[10, c*dz[\Omega m0, \Omega d0, \Omega k0, z]] + 25;
```





Deceleration parameter

Deceleration parameter can be plotted with respect to redshift z. Using flat FRW model. That is $\Omega k0=0$.

```
pldec[Ωm0v_, color_] := Plot[q[Ωm0v, 1 - Ωm0v, 0, z],
    {z, -1, 5}, PlotRange → {{-1.05, 5}, {-1.05, 0.55}},
    PlotStyle → color, AxesOrigin → {-1, 0}]

Show[{pldec[0.1, Red], pldec[0.26, Green], pldec[0.5, Blue],
    pldec[1, Orange], Plot[0, {z, -1, 5}, PlotStyle → Thick]}]

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Manipulate[pldec[Ωm0v, {Orange, Thick}],
```

Plot deceleration parameter with respect to r. This is not related to Ω k0, i.e., the curvature of the spacetime won't affect this result.

```
pldecr = Plot[ztrr[r], {r, 0, 1}, PlotRange → {{0, 1}, {-1, 4}},
    PlotStyle → {Thick, Orange}, AxesOrigin → {0, -1}, Frame → True,
    GridLines → {{{0.358, Dashed}, {0.378, Directive[Red]}, 0.398},
    {{0.376, LightRed}, {0.426, LightBlue}, {0.508, Gray}}}]
```

In a flat universe, Ω k0=0. Then Ω d0=1- Ω m0.

• chi square fitting. Using mathematica's functions.

Definations etc

chi square defination with error only.

```
chi2unions1[\Omegam0_?NumberQ, zt_?NumberQ] := 

ParallelSum[\(\du[[i, 2]] - \mag[\Om0, \frac{1}{2}\Om0 (1 + zt)^3,\)
1 - \Omegam0 - \frac{1}{2}\Om0 (1 + zt)^3, \du[[i, 1]]\)\(\rangle^2 \sqrt{du[[i, 3]]^2, {i, 1, len}}\)\)
```

Define a vector/table of "theory-observation".

thobtab[
$$\Omega m0_{,}$$
 zt_] = ParallelTable[du[[i, 2]] - mag[$\Omega m0_{,}$]
$$\frac{1}{2}\Omega m0_{,} (1+zt)^{3}, 1-\Omega m0_{,} -\frac{1}{2}\Omega m0_{,} (1+zt)^{3}, du[[i, 1]], \{i, 1, len\}];$$

Import covariance matrix data. "covm" is data with systematic and "covmns" is data without sys.

```
covm = Import["DATA/UNION2/SCPUnion2_covmat_sys.txt", "Table"];
icovm = Inverse[covm];

covmns = Import["DATA/UNION2/SCPUnion2_covmat_nosys.txt", "Table"];
icovmns = Inverse[covmns];
```

Testing the imported file.

```
du[[1]][[3]]<sup>-2</sup>
icovmns[[1, 1]]

19.5538
```

Define chi2 with and without systematic.

```
chi2union[Ωm0_?NumberQ, zt_?NumberQ] :=
  thobtab[Ωm0, zt].icovm.thobtab[Ωm0, zt]

chi2unionns[Ωm0_?NumberQ, zt_?NumberQ] :=
  thobtab[Ωm0, zt].icovmns.thobtab[Ωm0, zt]
```

Test the calcuation speed.

```
chi2union[0.27, 0.5] // Timing
```

{8.236, 560.12}

chi2unionns[0.27, 0.5] // Timing

{8.128, 735.117}

chi2unions1[0.27, 0.5] // Timing

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```
{0.156, 735.121}
```

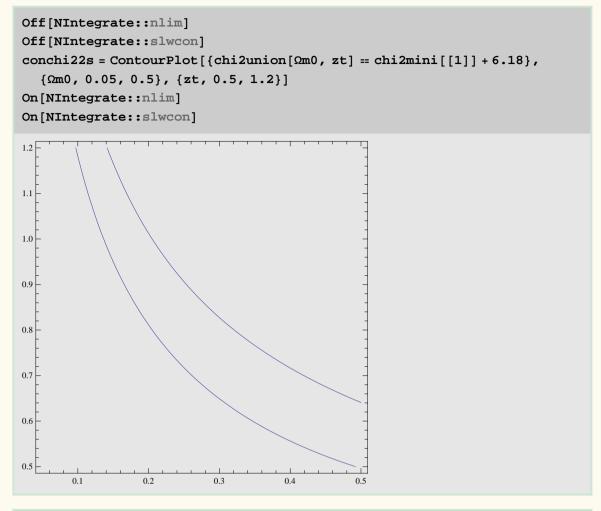
- Find the minimum chi2 for chi2 def. without sys.. AND Export the results to folder DATA/2/
- With Systematic in data. Find the minimum of chi2.

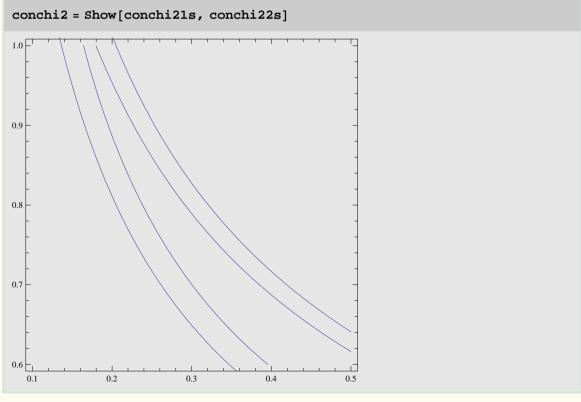
Export["OUTPUT\\2\\chi2union_minimum.dat", chi2mini]

OUTPUT\2\chi2union_minimum.dat

Plot the contour with confidential level 1σ and 2σ . This is a 2 parameter model. $\Delta \chi^2(1\sigma) = 2.30$ && $\Delta \chi^2(2\sigma) = 6.18$ && $\Delta \chi^2(3\sigma) = 11.83$.

With sys.





```
Export["OUTPUT\\2\\conchi21s.eps", conchi21s]
Export["OUTPUT\\2\\conchi22s.eps", conchi22s]
Export["OUTPUT\\2\\conchi2.eps", conchi2]

OUTPUT\2\\conchi21s.eps

OUTPUT\2\\conchi22s.eps

OUTPUT\2\\conchi22s.eps
```

- With only error, not the covariance matrix.
- Perform a grid method