## Computational Physics Homework 6

Cristopher Cerda Puga

## Richardson Extrapolation using the Euler-Maclaurin formula

It is of our interest to develop a method to extrapolate the value of the series:

$$\sum_{k=1}^{\infty} \frac{1}{k^s}$$

The partial sums of such series behave as:

$$\sum_{k=1}^{n} \frac{1}{k^{s}} \sim \zeta(s) - \frac{1}{(s-1)n^{s-1}} + \frac{1}{2n^{s}} - \sum_{i=1}^{n} \frac{B_{2i}}{(2i)!} \frac{(s+2i-2)!}{(s-1)!n^{s+2i-1}}$$

where the  $B_{2i}$  correspond to the Bernoulli numbers. The strategy to solve this problem is to write a system of equations for multiple partial sums:

$$A_{N} = q_{0} + \frac{q_{1}}{N^{s-1}} + \frac{q_{1}}{N^{s}} - \sum_{i=1}^{\infty} \frac{B_{2i}}{(2i)!} \frac{(s+2i-2)!}{(s-1)!N^{s+2i-1}}$$

$$A_{N+1} = q_{0} + \frac{q_{1}}{(N+1)^{s-1}} + \frac{q_{1}}{(N+1)^{s}} - \sum_{i=1}^{\infty} \frac{B_{2i}}{(2i)!} \frac{(s+2i-2)!}{(s-1)!(N+1)^{s+2i-1}}$$
...
$$A_{N+p} = q_{0} + \frac{q_{1}}{(N+p)^{s-1}} + \frac{q_{1}}{(N+p)^{s}} - \sum_{i=1}^{\infty} \frac{B_{2i}}{(2i)!} \frac{(s+2i-2)!}{(s-1)!(N+p)^{s+2i-1}}$$

Or in matrix notation:

$$\begin{pmatrix} A_{N} \\ A_{N+1} \\ \vdots \\ A_{N+p} \end{pmatrix} = \begin{pmatrix} 1 & 1/N^{s-1} & 1/N^{s} & \dots \\ 1 & 1/(N+1)^{s-1} & 1/(N+1)^{s} & \dots \\ \vdots & \vdots & \vdots & \vdots \\ 1 & 1/(N+p)^{s-1} & 1/(N+p)^{s} & \dots \end{pmatrix} \begin{pmatrix} q_{0} \\ q_{1} \\ \vdots \\ q_{p} \end{pmatrix}$$

We will see in a moment how to get the dot columns in the matrix. For now, the strategy is as follows: we have a matrix equation

$$A = MQ$$

Where the only thing that is of our interest is the first entry of Q, which corresponds to the approximation of the series. Now, one way to get such vector is by calculating the inverse  $M^{-1}$  of M and apply it to both sides of the matrix equation:

$$M^{-1}A = O$$

However, it is difficult to find and explicit formula for the inverse, specially because since the fourth column the entries take a form related to the ith dependent terms of the Euler-Maclaurin formula. So, we must calculate the partial sums  $A_i$ , i = N, ..., N + p, so increasing the number of p implies increasing the size of the matrix M. We will do it numerically using mpmath, first up to  $A_{N+3}$  and then up to  $A_{N+4}$ .

Now we can talk about the ith dependent terms of the matrix, for the N+3 case:

$$M = \begin{pmatrix} 1 & 1/N^{s-1} & 1/N^s & 1/N^{s+1} \\ 1 & 1/(N+1)^{s-1} & 1/(N+1)^s & 1/(N+1)^{s+1} \\ 1 & 1/(N+2)^{s-1} & 1/(N+2)^s & 1/(N+2)^{s+1} \\ 1 & 1/(N+3)^{s-1} & 1/(N+3)^s & 1/(N+3)^{s+1} \end{pmatrix}$$

Now we can start working on the problem

```
1
           import mpmath as mp
   2
  3
           mp.dps = 100
  4
  5
            def partial(nmax, s):
  6
                          return mp.nsum(lambda n: mp.power(n, -s), [1, nmax])
  7
  8
            A = mp.matrix([0,0,0,0])
10
           k = 100 # here k corresponds to N in our previous analysis
            s = 3/2 # the value of s
11
12
13
           for i in range(len(A)):
14
                          \# add the values of the partial function to the matrix A
15
                          nmax = i+k
16
                          A[i] = partial(nmax, s)
17
          # define the matrix M
18
           M = mp.matrix([[1,1/k**(s-1),1/k**s,1/k**(s+1)],[1,1/(k+1)**(s-1),1/(k+1)**s,1/(k+1) \leftrightarrow (s-1),1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k+1)**s,1/(k
19
                         **(s+1)],[1,1/(k+2)**(s-1),1/(k+2)**s,1/(k+2)**(s+1)],[1,1/(k+3)**(s-1),1/(k+3)**\longleftrightarrow
                        s,1/(k+3)**(s+1)]])
20
21
           # calculate the inverse of M
22
          Minv = mp.inverse(M)
23
24
         # calculate Q matrix
25
        Q = Minv*A
          print(Q)
          \Rightarrow [ 2.61237534844416] \Rightarrow The approximaton of the Riemann zeta function at s = 3/2
27
28
                           [-1.9999999491331]
29
                           [ 0.499999550103022]
30
                           [-0.124979896719848]
```

We can see that the approximation is pretty good. For a better approximation we can increase the size of our matrix and in consequence of the vector containing the partial sums, as well as select a good choose of k:

```
1
   import mpmath as mp
 2
3
   mp.dps = 100
   k = 50
 4
   s = 3/2
5
 6
7
   def partial(nmax, s):
8
        return mp.nsum(lambda n: mp.power(n, -s), [1, nmax])
9
10
   B = mp.matrix([0,0,0,0,0])
11
   for i in range(len(B)):
12
        nmax = i+k
13
        B[i] = partial(nmax, s)
14
   T = mp.matrix([[1, 1/k**(s - 1), 1/k**(s), 1/k**(s + 1), 1/k**(s + 3)], [1, 1/(k + 1) \leftrightarrow (1, 1/k)
15
       **(s - 1), 1/(k + 1)**(s), 1/(k + 1)**(s + 1), 1/(k + 1)**(s + 3)],[1, 1/(k + 2)\leftarrow
       **(s - 1), 1/(k + 2)**(s), 1/(k + 2)**(s + 1),1/(k + 2)**(s + 3)], [1, 1/(k + 3) \leftarrow
       **(s - 1), 1/(k + 3)**(s), 1/(k + 3)**(s + 1), 1/(k + 3)**(s + 3)], [1, 1/(k + 4) \leftarrow
       **(s-1), 1/(k+4)**(s), 1/(k+4)**(s+1), 1/(k+4)**(s+3)]]
16
   Tinv = mp.inverse(T)
17
18
   Q2 = Tinv*B
19
   print(Q2)
20 \gg [2.61237534868775] - The approximation of the Riemann zeta function at s = 3/2
```

```
21  [ -2.000000001121]

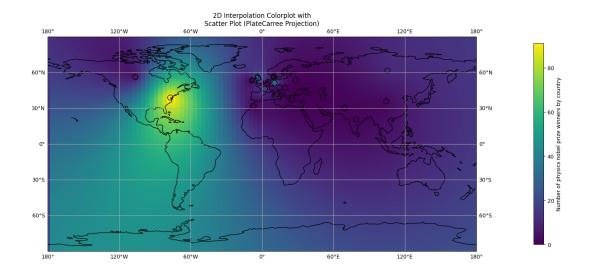
22  [ 0.500000005620954]

23  [-0.125000118835117]

24  [0.0182315027087133]
```

So we can see that for a very slow value of p, like 4 or 5, we get a pretty good approximation of the series. Increasing the size of the matrix (i.e. increasing p) gives us a better approximation.

## Radial basis functions interpolation.



```
import numpy as np
   import matplotlib.pyplot as plt
3
   import cartopy.crs as ccrs
4
   from scipy.interpolate import Rbf
5
6
   lats = [38.895,51.165691,55.378051,46.227638,61.52401,36.204824,52.132633,56.130366,
7
            41.87194,60.128161,47.516231,46.818188,56.26392,35.86166,31.791702,53.41291,
8
            20.593684,30.375321,50.503887,53.709807,47.162494]
9
   lons = [-77.0366,10.451526,-3.435973,2.213749,105.318756,138.252924,5.291266,
10
            -106.346771,12.56738,18.643501,14.550072,8.227512,9.501785,104.195397,
11
            -7.09262, -8.24389, 78.96288, 69.345116, 4.469936, 27.953389, 19.503304]
12
   data = [96,28,25,16,11,10,9,5,5,4,4,4,3,2,1,1,1,1,1,1,1]
13
   print(len(lats), len(lons), len(data))
14
15
   # Create a meshgrid for interpolation
16
   xi = np.linspace(-180, 180, 100)
17
   yi = np.linspace(-90, 90, 100)
18
   xi, yi = np.meshgrid(xi, yi)
19
   # Different functions for the spherical interpolation:
```

```
21 # 'linear': Linear radial basis function
22 # 'multiquadric': Multiquadric radial basis function
23 # 'inverse': Inverse radial basis function
24 # 'gaussian': Gaussian radial basis function
25 # 'cubic': Cubic radial basis function
26 # 'quintic': Quintic radial basis function
27 # 'thin_plate': Thin-plate spline radial basis function
28
29 # Perform 2D interpolation using Rbf for spherical interpolation
30 rbf = Rbf(lons, lats, data, function='linear')
31 zi = rbf(xi, yi)
32
33 # Create a Cartopy PlateCarree projection centered at lon = 0
34 ax = plt.axes(projection=ccrs.PlateCarree(central_longitude=0))
35
36 # Add coastlines and gridlines
37 ax.coastlines()
38 ax.gridlines(draw_labels=True)
39
40 # Set extent of the plot
41 ax.set_extent([-180, 180, -90, 90], crs=ccrs.PlateCarree())
42
43 # Create color plot
44 plt.pcolormesh(xi, yi, zi, shading='auto', cmap='viridis')
45
46\, # Create scatter plot with data as color
47 scatter = ax.scatter(lons, lats, c=data, cmap='viridis', s=100,\
48
                        edgecolors='black', linewidths=1,\
49
                        transform=ccrs.PlateCarree())
50
51 # Add colorbar
52 cbar = plt.colorbar(scatter, label='Number of physics nobel prize winners by country↔
       ', ax=ax, shrink=0.7, pad = .1)
53
  cbar.mappable.set_clim(0, 91)
54
55 # Add title
56 plt.title('2D Interpolation Colorplot with\nScatter Plot (PlateCarree Projection)')
57
58 plt.show()
```