

# Portable **U**niversity **M**odel of the **A**tmosphere (PUMA)

**Dynamical core of the PlanetSimulator (University Hamburg)**

## **Original version**

- Global spectral general circulation model of the dry atmosphere
- Numerical solution of hydrostatic EULER equations for an ideal gas on a rotating sphere
- Diabatic, dissipative processes: only Newtonian Cooling and Rayleigh friction

## **Simple physics extension**

- Inclusion of a vapor transport equation
- Boundary layer and condensation scheme by Reed and Jablonowski (2011)



## Summary of the governing PUMA equations

$$\text{Vorticity: } \frac{\partial \eta}{\partial t} = \nabla_h \cdot (\mathbf{k} \times \mathbf{F}_{\mathbf{v}_h}) + D_\eta$$

$$\text{Divergence: } \frac{\partial D}{\partial t} = -\nabla_h \cdot \mathbf{F}_{\mathbf{v}_h} + \nabla_h^2 \left( \frac{\mathbf{v}_h^2}{2} \cos^2(\varphi) + \phi + RT \ln(p_s) \right) + D_D$$

$$\text{Temperature: } \frac{\partial T}{\partial t} = -\nabla_h \cdot (\mathbf{v}_h T) + DT - \dot{\sigma} \frac{\partial T}{\partial \sigma} + \frac{\kappa T}{p} \frac{Dp}{Dt} + D_T$$

$$\text{Water vapor: } \frac{\partial}{\partial t} (p_s q) = -\nabla_h \cdot (\mathbf{v}_h p_s q) - \frac{\partial}{\partial \sigma} (\dot{\sigma} p_s q) + D_q$$

$$\text{Surface pressure: } \frac{\partial \ln p_s}{\partial t} + \int_0^1 \frac{1}{p_s} \nabla_h \cdot (p_s \mathbf{v}_h) d\sigma = 0$$

$$\text{Geopotential: } \frac{\partial \phi}{\partial \sigma} = -\frac{RT}{\sigma}$$

$$\text{Vertical velocity: } \dot{\sigma} = \sigma \int_0^1 \frac{1}{p_s} \nabla_h \cdot (p_s \mathbf{v}_h) d\sigma - \int_0^\sigma \frac{1}{p_s} \nabla_h \cdot (p_s \mathbf{v}_h) d\sigma$$

$\sigma = p/p_s$  is the vertical coordinate,  $D_x$  irreversible Processes



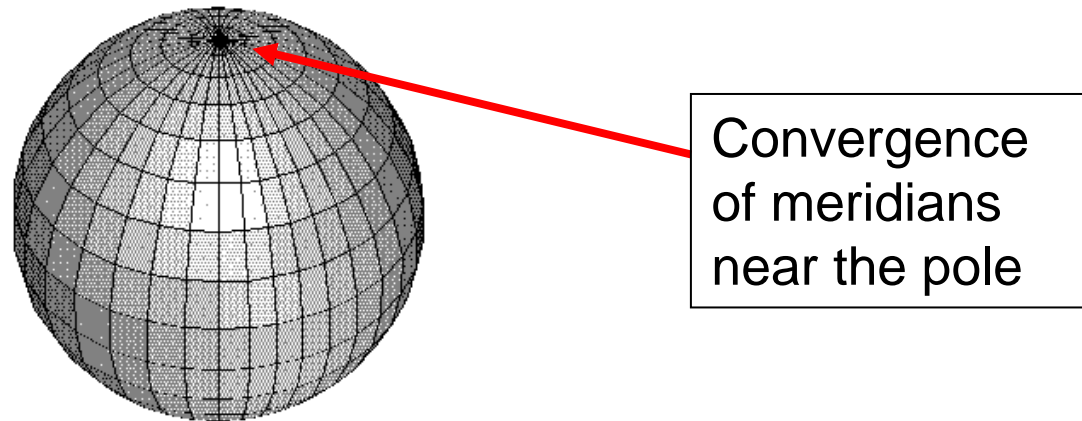
## Level arrangement for a five layer model ( $\Delta\sigma = 1/N = 1/5 = 0.2$ )

$\sigma$ -value	Index	Variable
$\sigma=0$	$\frac{1}{2}$	$\dot{\sigma}_{\frac{1}{2}} = 0$
$\sigma = \frac{1}{2} \Delta\sigma = 0.1$	1	$\mathbf{v}_{h1}, \phi_1, T_1$
$\sigma = \Delta\sigma = 0.2$	$1\frac{1}{2}$	$\dot{\sigma}_{1+\frac{1}{2}}$
$\sigma = 1\frac{1}{2} \Delta\sigma = 0.3$	2	$\mathbf{v}_{h2}, \phi_2, T_2$
$\sigma = 2\Delta\sigma = 0.4$	$2\frac{1}{2}$	$\dot{\sigma}_{2+\frac{1}{2}}$
$\sigma = 2\frac{1}{2} \Delta\sigma = 0.5$	3	$\mathbf{v}_{h3}, \phi_3, T_3$
$\sigma = 3\Delta\sigma = 0.6$	$3\frac{1}{2}$	$\dot{\sigma}_{3+\frac{1}{2}}$
$\sigma = 3\frac{1}{2} \Delta\sigma = 0.7$	4	$\mathbf{v}_{h4}, \phi_4, T_4$
$\sigma = 4\Delta\sigma = 0.8$	$4\frac{1}{2}$	$\dot{\sigma}_{4+\frac{1}{2}}$
$\sigma = 5\frac{1}{2} \Delta\sigma = 0.9$	$N=5$	$\mathbf{v}_{h5}, \phi_5, T_5$
$\sigma=1$	$N+ \frac{1}{2}$	$\dot{\sigma}_{5+\frac{1}{2}} = 0$



## Spectral method (Hoskins and Simmons 1975)

PUMA is based on a longitude-latitude grid. However, with spherical coordinates singularities occur at the poles. These singularities induce a collapse of coordinate lines into a single point.



Source: <http://www.personal.umich.edu/~cjablono/project.html>

With the spectral method this problem can be solved. All fields are represented by the expansion

$$G(\lambda, \varphi, t) = \sum_{n=0}^{\infty} \sum_{m=-n}^{m=n} Y_n^m(\lambda, \mu(\varphi)) \hat{G}_n^m(t)$$

Linear terms are evaluated in spectral space while for nonlinear terms the spectral transform method is applied.



## Spectral triangular truncation

The condition that at least  $(3M+1)$  longitudes must be used leads to typical values for the spectral truncation wavenumber.

The fastest performance of a Fast FOURIER Transform results when  $2^n$  ( $n$  integer) longitudes are given in the model.

Longitudes	Truncation	DCMIP denotation	Resolution at equator [km]
32	T10		1250.9
64	T21		625.5
128	T42		312.7
256	T85	LOW	156.4
512	T170	MEDIUM	78.2
1024	T341	HIGH	39.1
2048	T682	ULTRA	19.6



## Semi implicit time integration scheme

- Linear part of PUMA can be solved without spectral transform method.
- This part describes high frequency gravity wave oscillations.

With the semi-implicit scheme the tendency is also a function of the “unknown” new state vector  $\mathbf{X}^{q+1}$

$$\mathbf{X}^{q+1} = \mathbf{X}^{q-1} - i\Delta t \mathbf{A} \cdot (\mathbf{X}^{q-1} + \mathbf{X}^{q+1})$$

where  $\mathbf{A}$  is the linear tendency operator of the linear part.  
This scheme gives unconditional numerical stability

## Hyperdiffusion

To dampen small-scale spatial noise a hyperdiffusion term of the form

$$(-1)^{n_h-1} k_h \nabla_h^{2n_h} F$$

is added to the vorticity, divergence and temperature equations  
where  $n_h$  is the order of hyperdiffusion and  $k_h$  the coefficient.



# Flow diagram of the PUMA model code

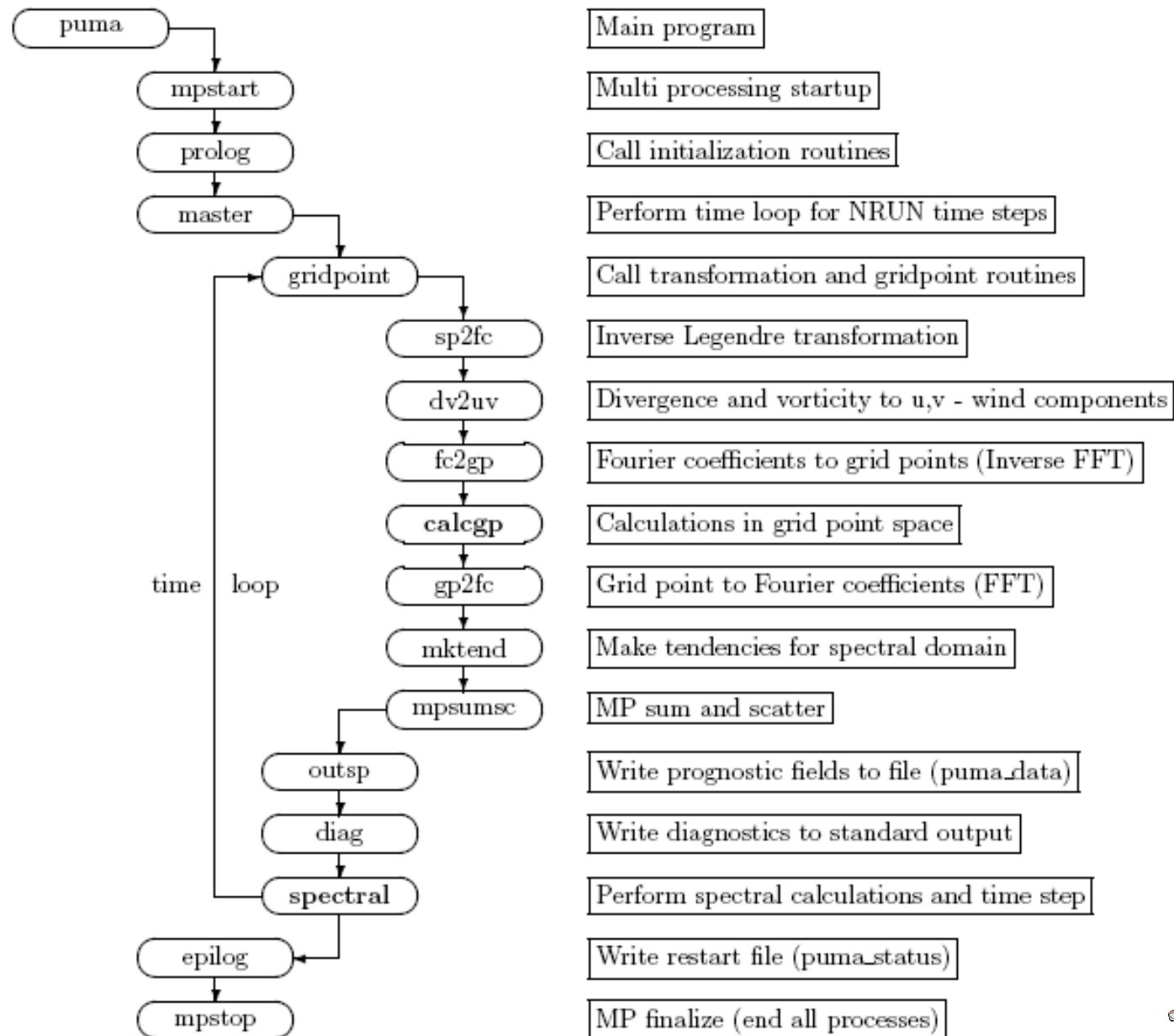


Figure 8.4: Flow diagram of main routines

