

## Mecânica Celeste

## Momento Angular

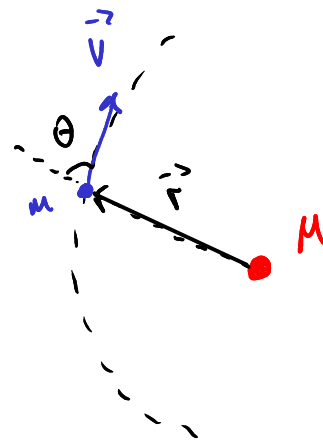
O momento angular é uma grandeza vetorial associada ao movimento de rotação.

$$\vec{L} = \vec{r} \times \vec{p}$$

( Para órbitas:

$v$  em módulo

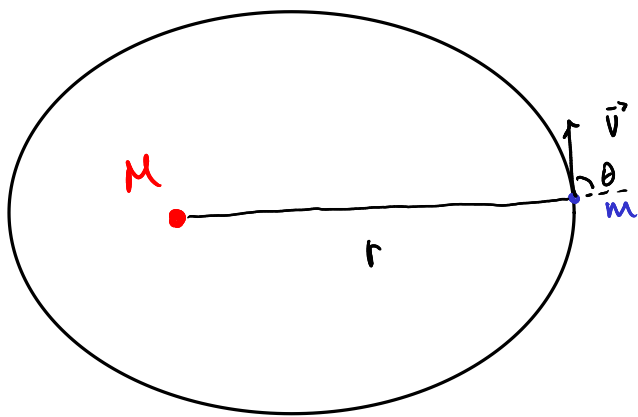
$$L = m v r \sin \theta$$



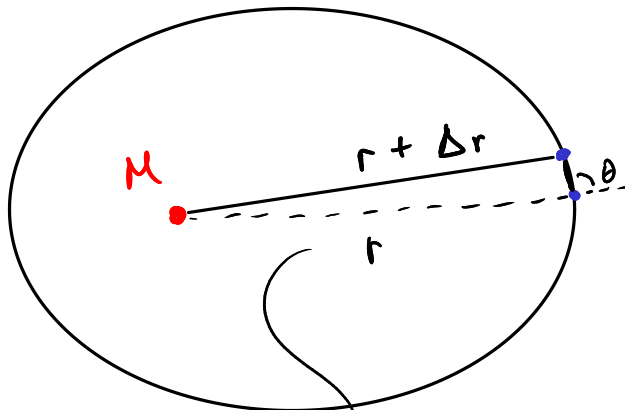
Em órbitas, o momento angular se conserva

## 2ª Lei de Kepler: Demonstração

Partindo desse fato podemos uma demonstração para a 2ª Lei de Kepler:



Para uma  
passagem de tempo  
muito pequena



A área do triângulo  
é:

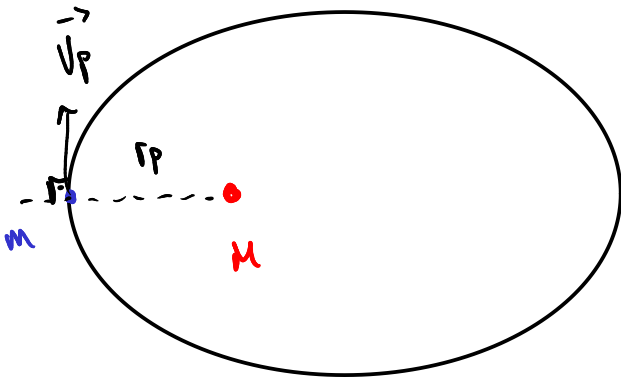
$$\Delta A = \frac{r \cdot v \Delta t \sin \theta}{2} \quad \therefore \quad \frac{\Delta A}{\Delta t} = \frac{r v \sin \theta}{2}$$

$$\frac{\Delta A}{\Delta t} = \frac{m r v \sin \theta}{2m}$$

$$L = m r v \sin \theta \quad \therefore \quad \boxed{\frac{\Delta A}{\Delta t} = \frac{L}{2m}}$$

constante

## Momento Angular de uma Órbita Elíptica



$$h = m v_p r_p \sin \theta_p$$

$$\sin \theta_p = 1$$

$$\therefore h = m v_p r_p$$

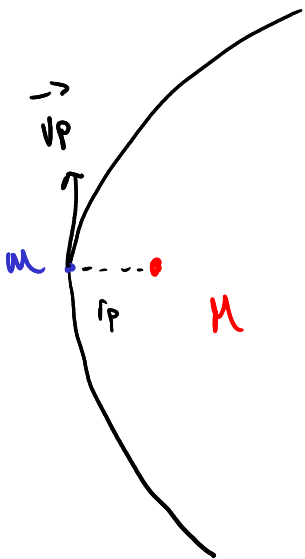
$$h = m \sqrt{\frac{GM}{a} \frac{1+e}{1-e}} a(1-e)$$

$$L = m \sqrt{\frac{GM}{a} \frac{1+e}{1-e} a^2 (1-e)^2}$$

$$h = m \sqrt{GM a (1+e)(1-e)}$$

$$\therefore h = m \sqrt{GM a (1-e^2)}$$

## Momento Angular de uma Órbita Hiperbólica



$$r_p = a(e-1)$$

$$v_p = \sqrt{GM \left( \frac{2}{r_p} - \frac{1}{a} \right)}$$

$$\therefore v_p = \sqrt{\frac{GM}{a} \frac{e+1}{e-1}}$$

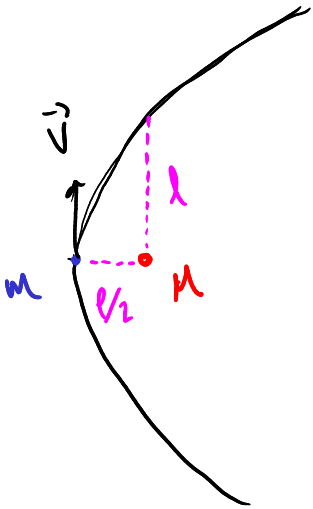
$$\therefore L = m v_p r_p \sin \theta$$

$$L = m \sqrt{\frac{GM}{a} \frac{e+1}{e-1}} a(e-1)$$

$$L = m \sqrt{\frac{GM}{a} \frac{e+1}{e-1} a^2 (e-1)^2} \therefore L = m \sqrt{GMa(e+1)(e-1)}$$

$$L = m \sqrt{GMa(e^2 - 1)}$$

Momento Angular de uma Órbita Parabólica



$$v = \sqrt{\frac{2GM}{l/2}} \therefore v = \sqrt{\frac{4GM}{l}}$$

$$h = m v \frac{l}{2} \sin \theta$$

$$\therefore L = 2m \sqrt{\frac{GM}{l}} \frac{l}{2}$$

$$L = m \sqrt{\frac{GM}{l} l^2}$$

$$\therefore L = m \sqrt{GMl}$$