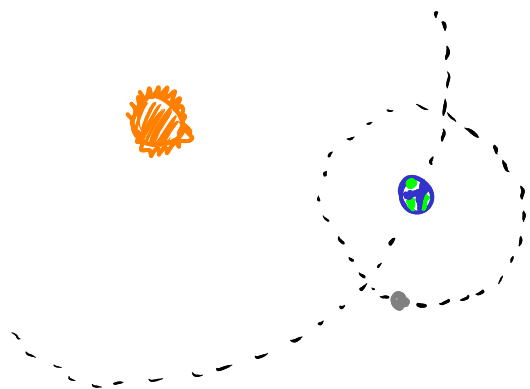


Órbitas



A Terra e os demais planetas **orbitam** o Sol.

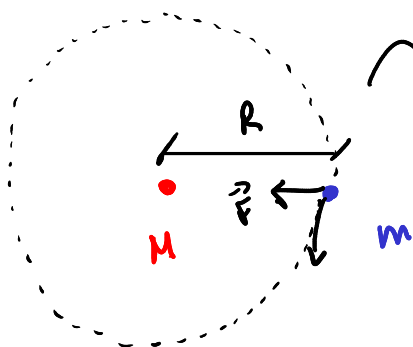
Assim como as luas **orbitam** seus planetas

↳ Formato das órbitas

↳ Parâmetros importantes

↳ Cálculo da velocidade, posição, etc.

↳ Órbitas Circulares



Movimento Circular Uniforme (MCU)

↳ P: Período

↳ R: Raio

↳ V: Velocidade

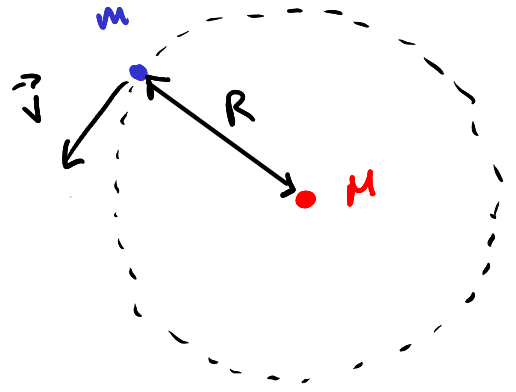
↳ ω : Velocidade Angular

↳ Velocidades

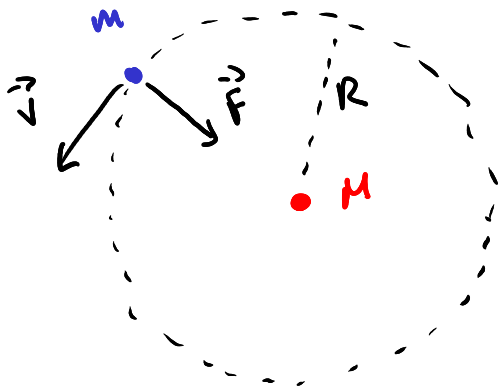
$$v = \frac{\Delta s}{\Delta t} \therefore v = \frac{2\pi R}{p}$$

$$\omega = \frac{\Delta \theta}{\Delta t} \therefore \omega = \frac{2\pi}{p}$$

$$v = \omega R$$



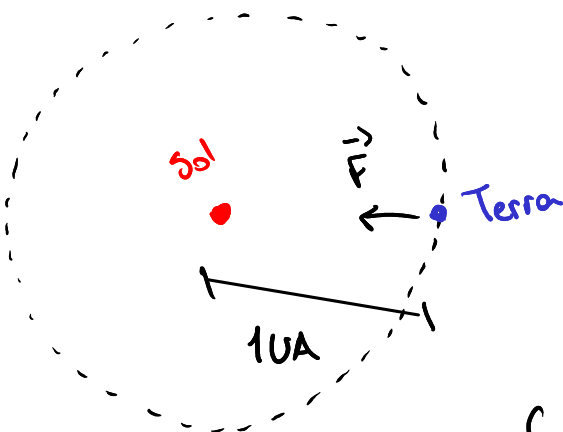
↳ Força Centrípetra



$$\left. \begin{aligned} F &= \frac{mV^2}{R} \\ F &= \frac{GMm}{R^2} \end{aligned} \right\} \frac{\cancel{GMm}}{\cancel{R^2}} = \frac{\cancel{mV^2}}{\cancel{R}}$$

$$V = \sqrt{\frac{GM}{R}}$$

Exemplo: Calcule a Velocidade da Terra



Órbita da Terra \approx circular

$$\left. \begin{aligned} F &= \frac{mV^2}{R} \\ F &= \frac{GMm}{R^2} \end{aligned} \right\} V = \sqrt{\frac{GM}{R}}$$

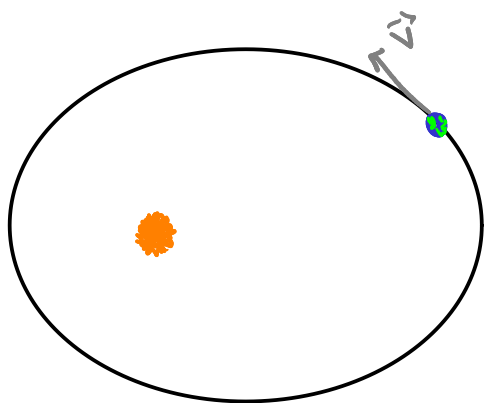
$$\begin{aligned} G &= 6,67 \cdot 10^{-11} \text{ N m}^2/\text{kg}^2 \\ R &= 1\text{UA} = 1,496 \cdot 10^{11} \text{ m} \\ M &= 1,99 \cdot 10^{30} \text{ kg} \end{aligned}$$

$$\boxed{V \approx 29,8 \text{ km/s}}$$

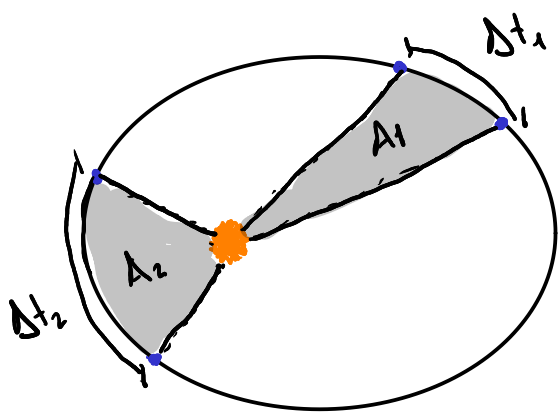
Leis de Kepler

As Leis de Kepler descrevem aspectos fundamentais das órbitas.

1ª Lei: As órbitas dos planetas são **elipses**, com o Sol em um dos focos.



2ª Lei: A "área varrida" pelo corpo por unidade de tempo é constante.



$$\frac{A_1}{\Delta t_1} = \frac{A_2}{\Delta t_2}$$

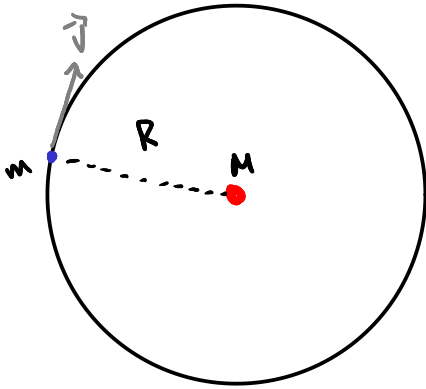
$$\frac{A_1}{\Delta t_1} = \frac{A_2}{\Delta t_2} = \text{cte}$$

3ª Lei: O quadrado do período orbital dividido pelo cubo do raio da órbita é uma constante.

$$\frac{P^2}{R^3} = K$$

$$V = \sqrt{\frac{GM}{R}}$$

$$V = \frac{2\pi R}{P}$$



$$\hookrightarrow \sqrt{\frac{GM}{R}} = \frac{2\pi R}{P}$$

$$\frac{GM}{R} = \frac{4\pi^2 R^2}{P^2}$$

$$\boxed{\frac{P^2}{R^3} = \frac{4\pi^2}{GM}}$$