

# **DOING PHYSICS WITH PYTHON**

## **COMPUTATIONAL OPTICS**

### **RAYLEIGH-SOMMERFELD 1**

#### **DIFFRACTION INTEGRAL CIRCULAR**

#### **APERTURE:**

### **HERMITE-GAUSS BEAM PROPAGATION**

**Ian Cooper**

Please email me any corrections, comments, suggestions or

additions: **matlabvisualphysics@gmail.com**

[maitreyee.roy@unsw.edu.au](mailto:maitreyee.roy@unsw.edu.au)

#### **DOWNLOAD DIRECTORIES FOR PYTHON CODE**

**Google drive**

**GitHub**

**emRSHGXY.PY**

Propagation of a Hermite-Gauss beam from a circular aperture. The irradiance in an observation XY plane is computed by solving the RS1 diffraction integral.

Links → **Hermite polynomials and Hermite-Gauss functions**  
**RS1 diffraction integral**

**Hermite–Gaussian (HG) beams** have significant potential to improve the capacity of free-space optical communication. Free-space optical communication has the advantages of high bandwidth, high data rate, low energy consumption and high confidentiality and is considered a promising technology for next-generation communication.

So, this article will discuss the propagation of a HG beam from a circular aperture. A higher-order HG beam can form a multiple-spot pattern for the irradiance rather than a single spot as generated by lowest-order Gaussian beam. Also, a HG beam has unique orthogonal spatial modes and well-preserved irradiance distribution in free-space propagation.

**Hermite-Gauss circular aperture functions** and are given by

$$E_{mn}(x_Q, y_Q) = E_0 H_m \left( \frac{\sqrt{2} x_Q}{w_0} \right) H_n \left( \frac{\sqrt{2} y_Q}{w_0} \right) \exp \left( -\frac{x_Q^2 + y_Q^2}{2 w_0^2} \right)$$

where  $m$  and  $n$  represent the transverse mode numbers,  $H$  are the Hermite-Gauss polynomials and  $w_0$  is a characteristic mode width.

For a circular aperture  $a$

$$\sqrt{x_Q^2 + y_Q^2} > a \Rightarrow E(x_Q, y_Q) = 0$$

In particular, for  $m = n = 0$ , the HG beam is reduced a **Gaussian beam** (the lowest-order TEM<sub>00</sub> beam). For the Gaussian beam,  $w_0$  denotes the TEM<sub>00</sub> beam waist at the aperture and  $w$  denotes the TEM<sub>00</sub> spot size in the XY observer place

$$w(z) = w_0 \sqrt{1 + \left( \frac{\lambda z}{\pi w_0^2} \right)^2} \quad w(0) = w_0$$

## SIMULATIONS

The Python Code **emRSHGXY.py** is used for the simulations to calculate the XY irradiance pattern at a given distance between the aperture and observation plane for different values of the orders  $m$  and  $n$ . A summary of the simulation parameters is displayed in the Console Window.

**nQ = 201   nP = 201**

**wavelength wL = 633 nm   aperture radius a = 0.400 mm**

**m = 1   n = 1**

**zP = 1.00 m**

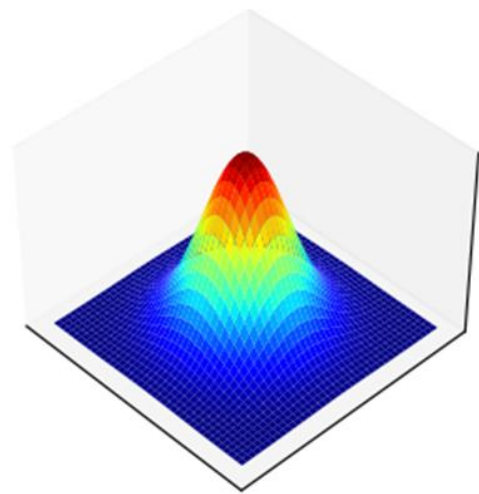
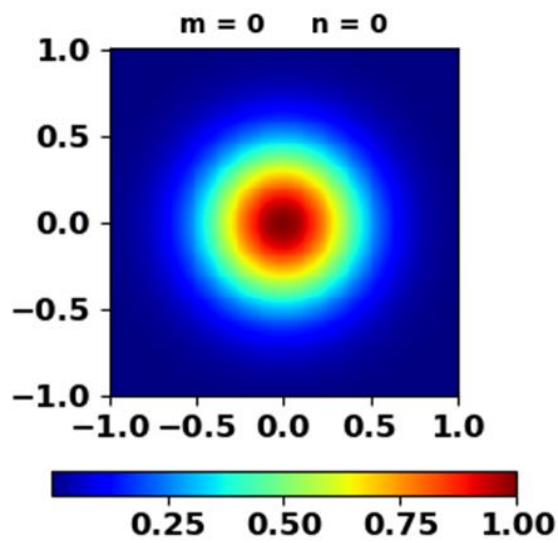
**Imax = 1.06e+24**

**Execution time 114 s**

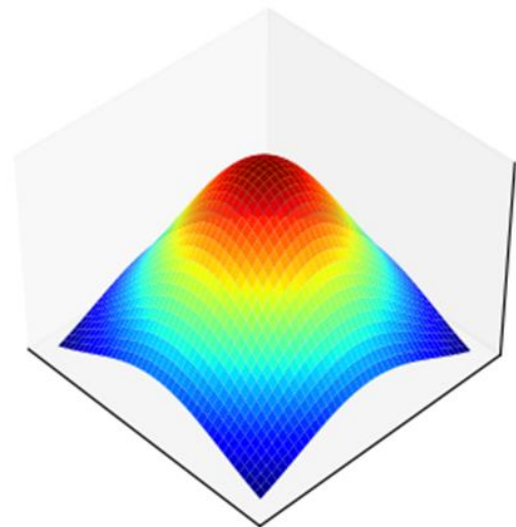
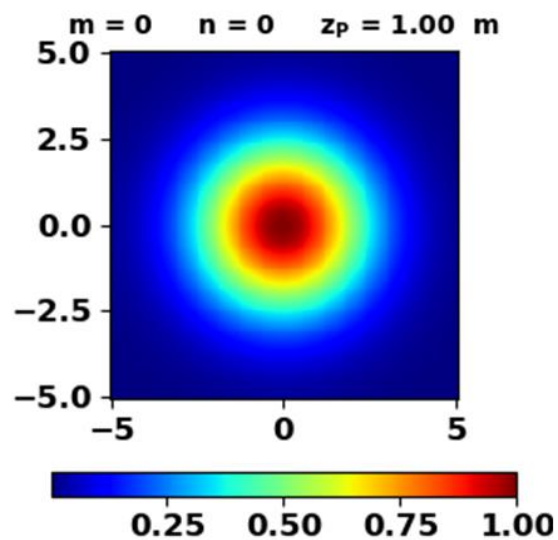
For the  $\text{TEM}_{mn}$  plots below, the irradiance patterns have been scaled  $(I_{XY})^f$   $0 < f \leq 1$ , so that the weak minor maxima are more visible.

## TEM<sub>00</sub>

### APERTURE SPACE

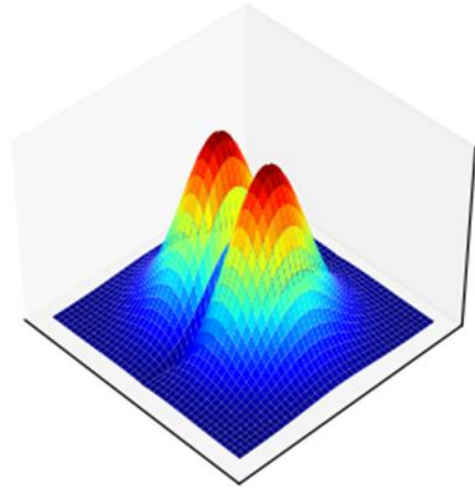
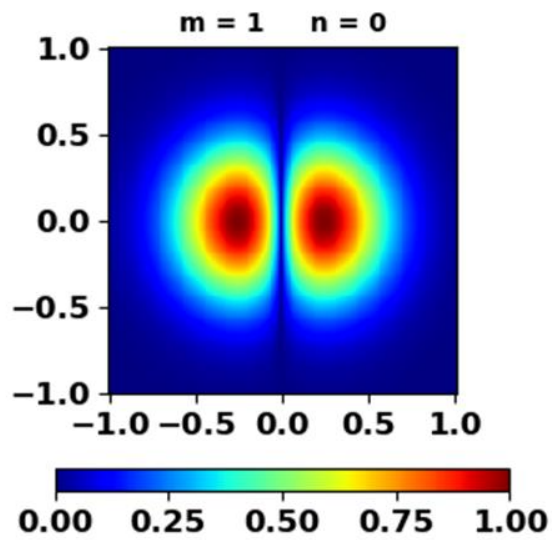


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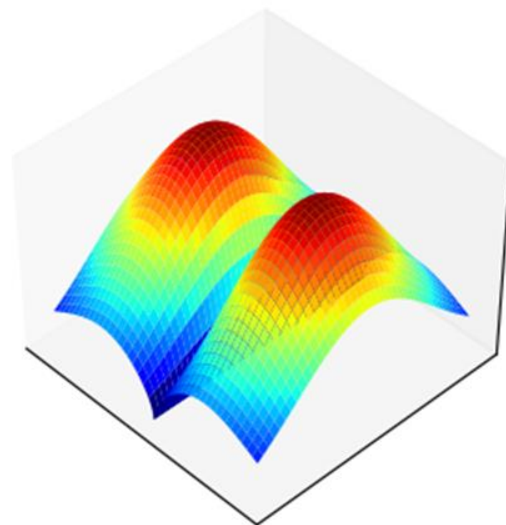
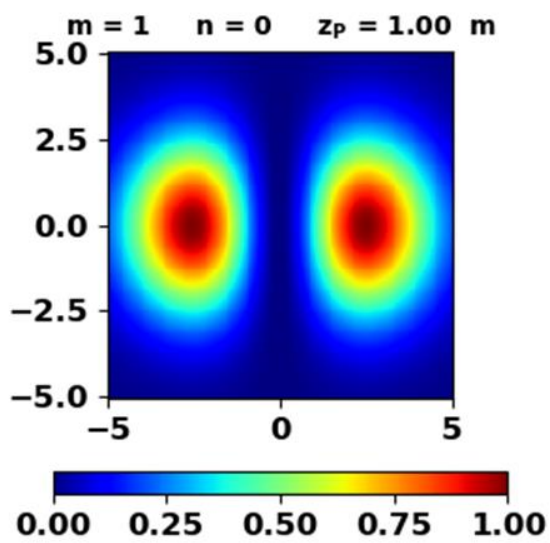


## TEM<sub>10</sub>

### APERTURE SPACE

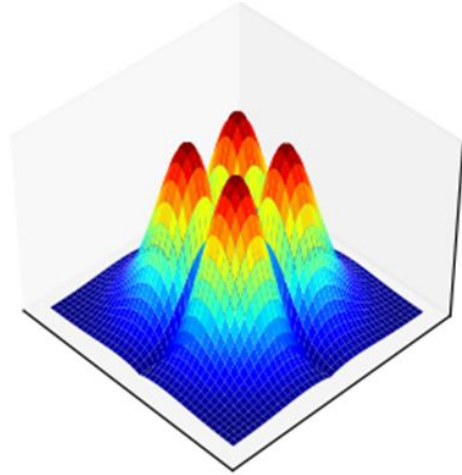
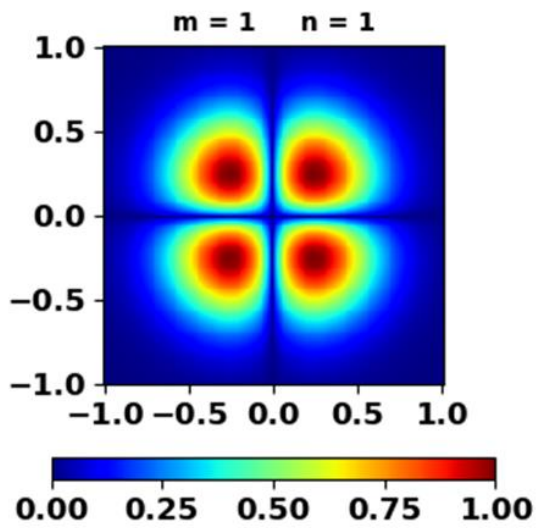


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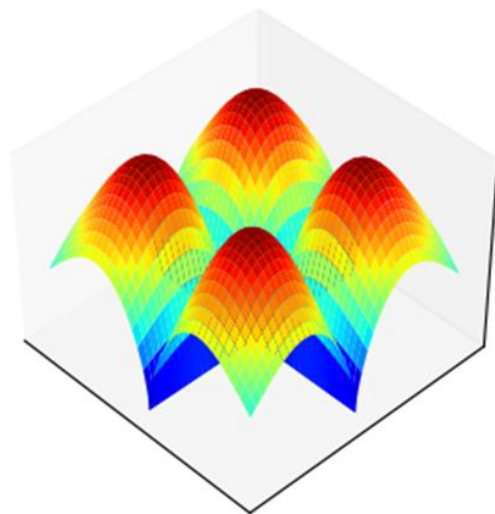
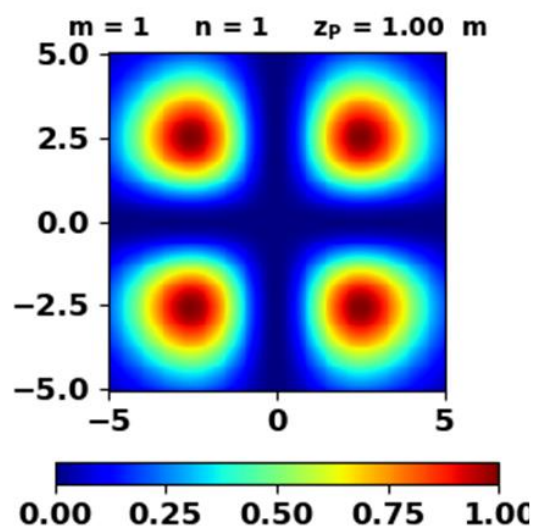


## TEM<sub>11</sub>

### APERTURE SPACE

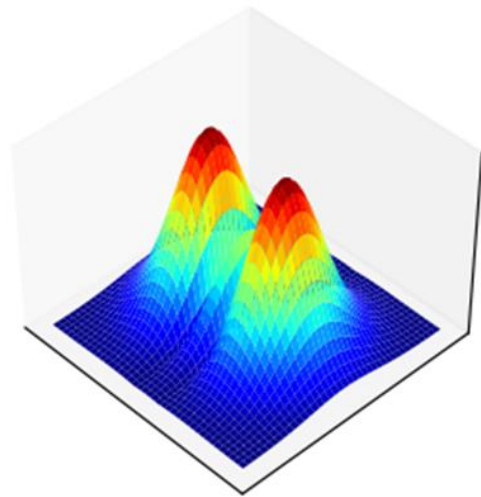
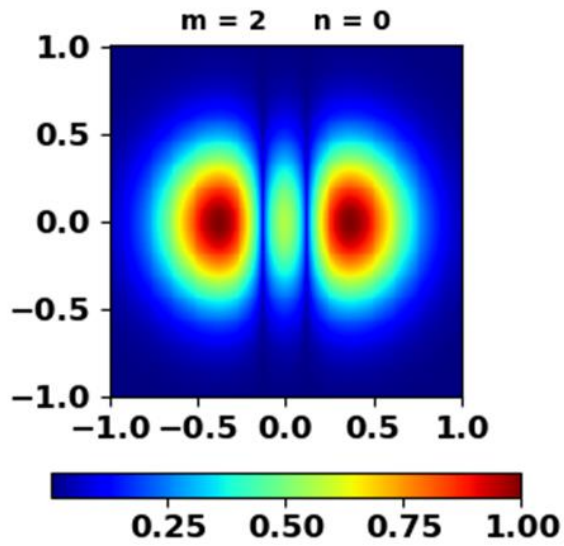


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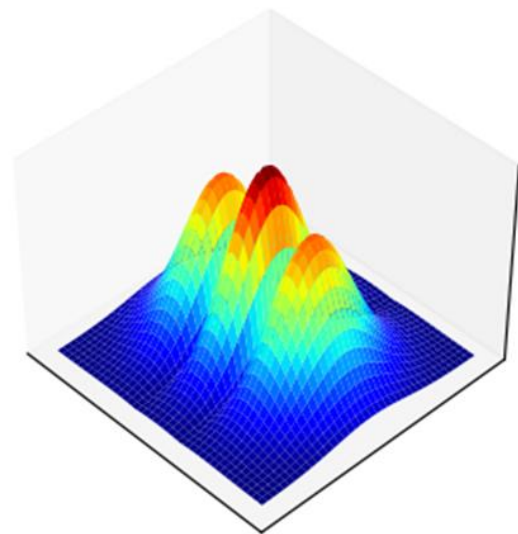
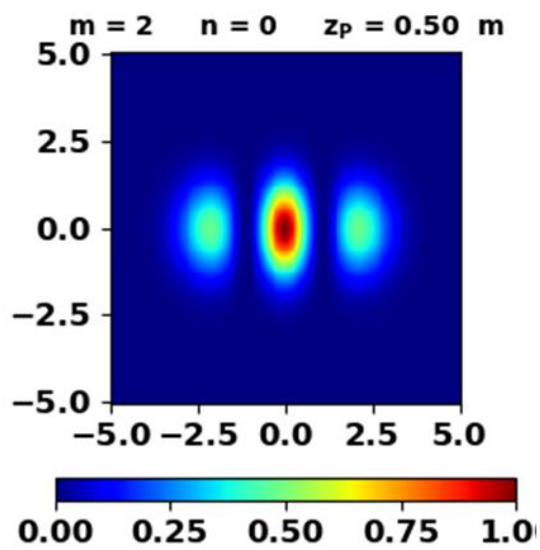


## TEM<sub>20</sub>

### APERTURE SPACE



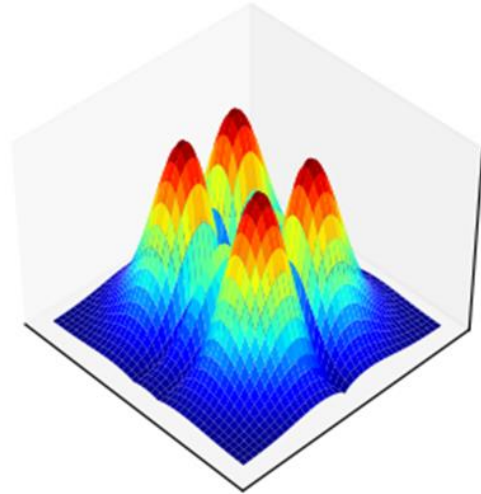
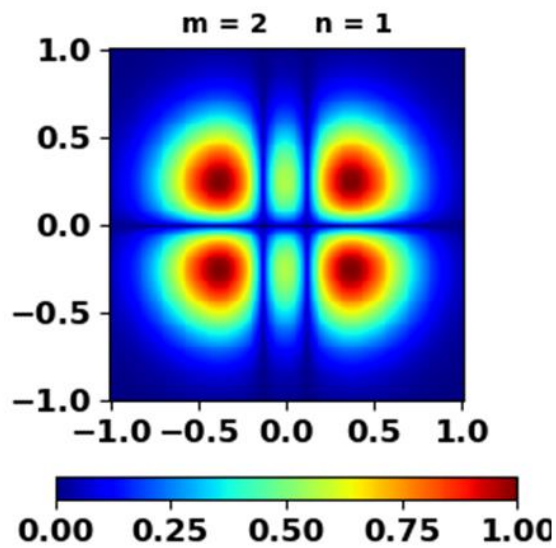
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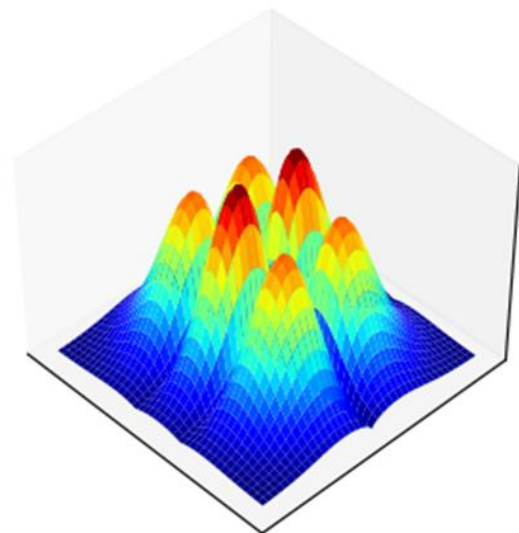
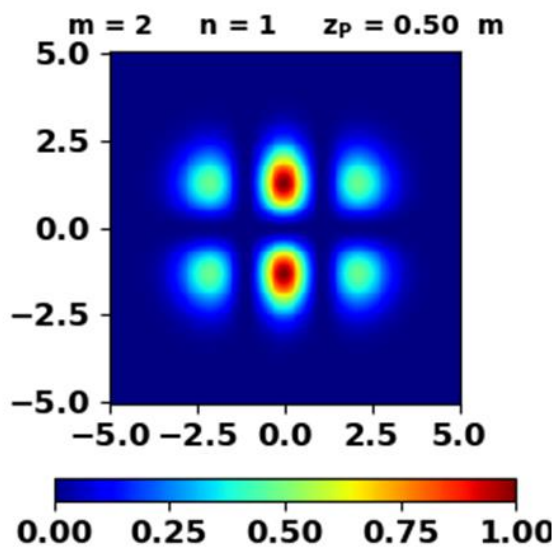


## TEM<sub>21</sub>

### APERTURE SPACE



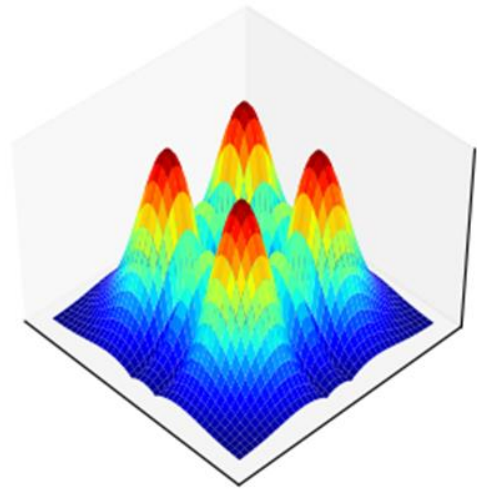
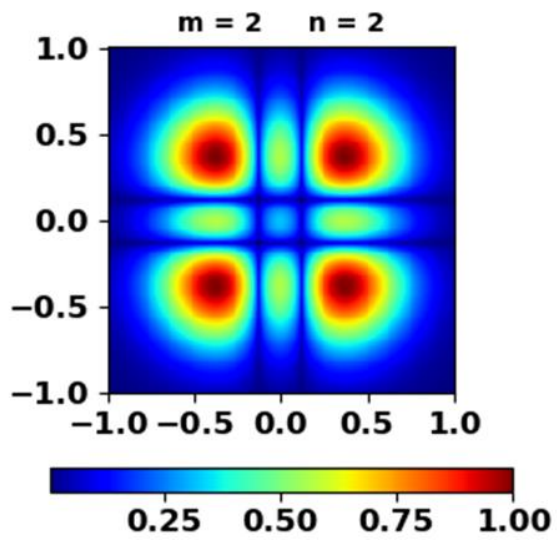
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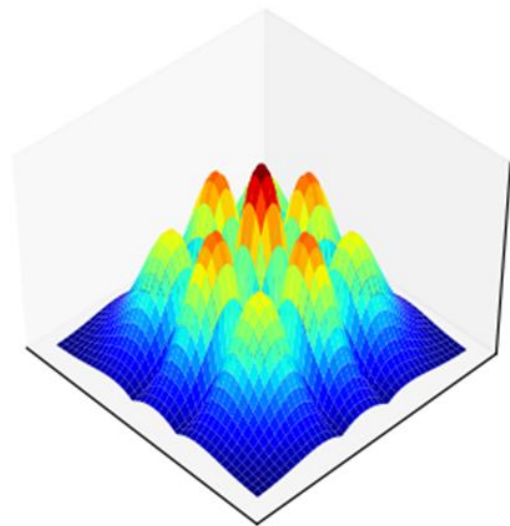
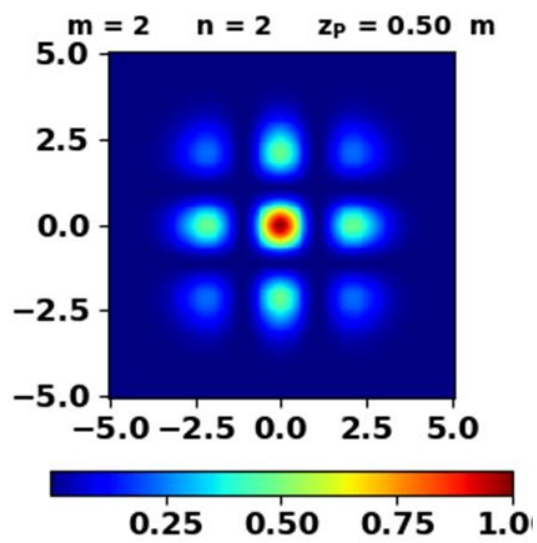


## TEM<sub>22</sub>

### APERTURE SPACE

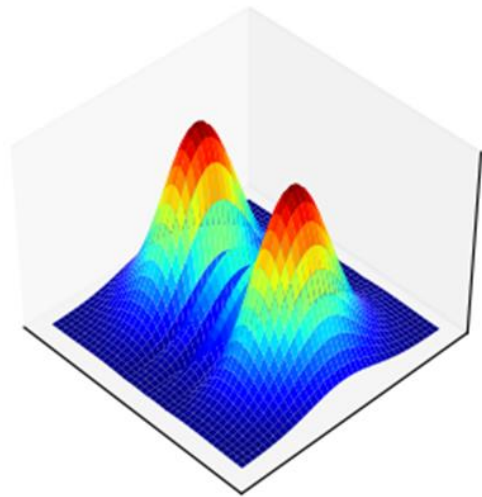
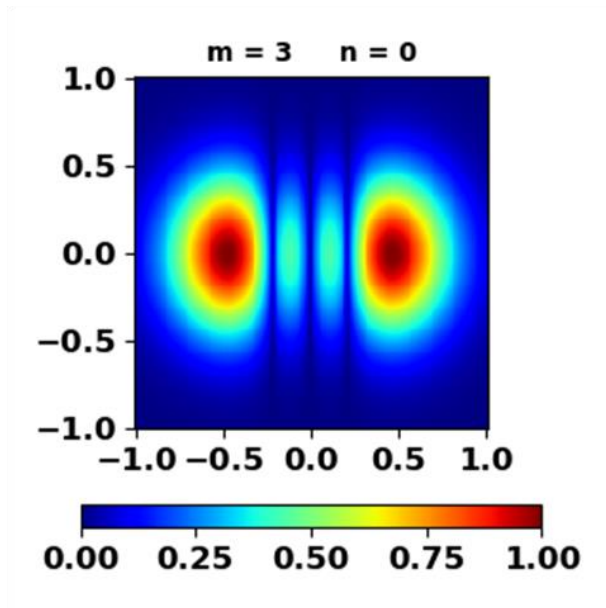


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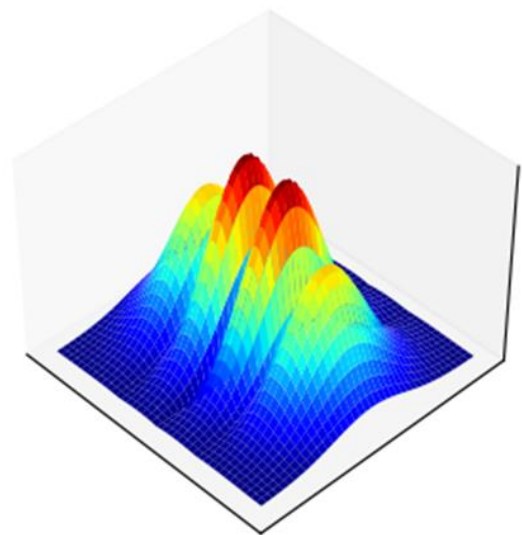
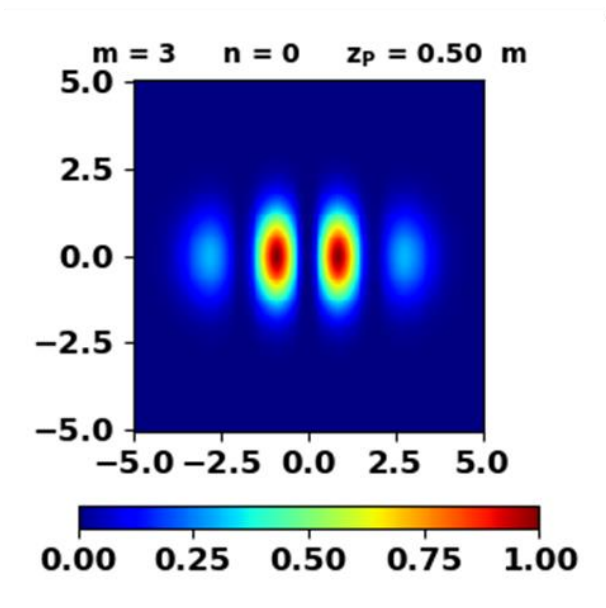


## TEM<sub>30</sub>

### APERTURE SPACE

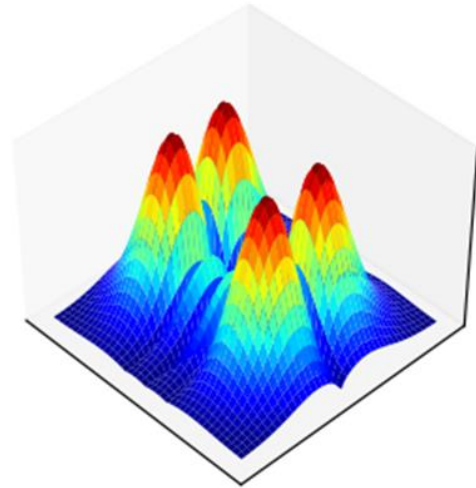
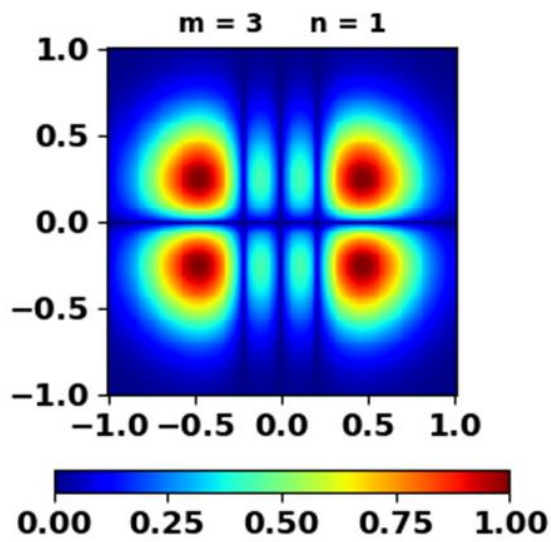


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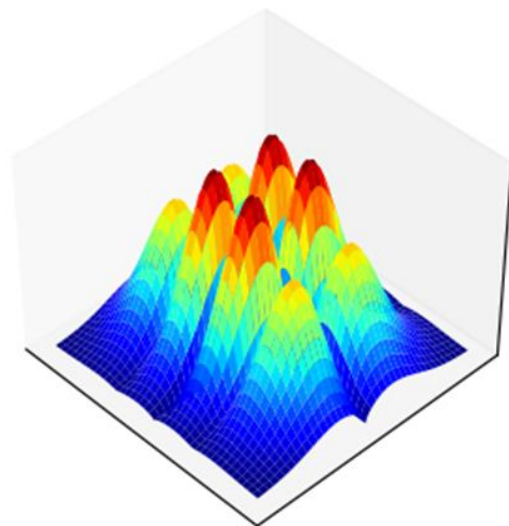
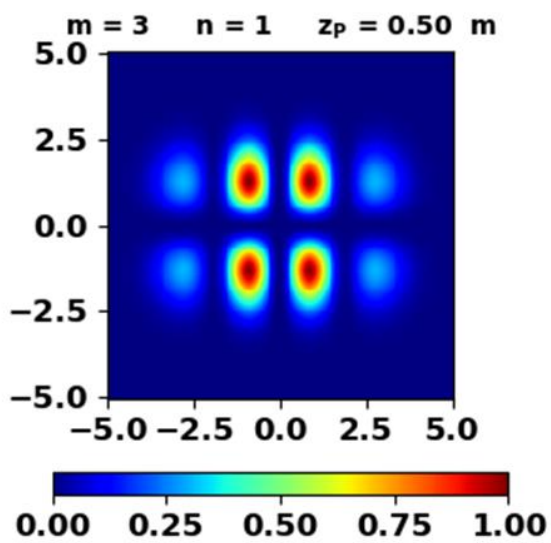


## TEM<sub>31</sub>

### APERTURE SPACE

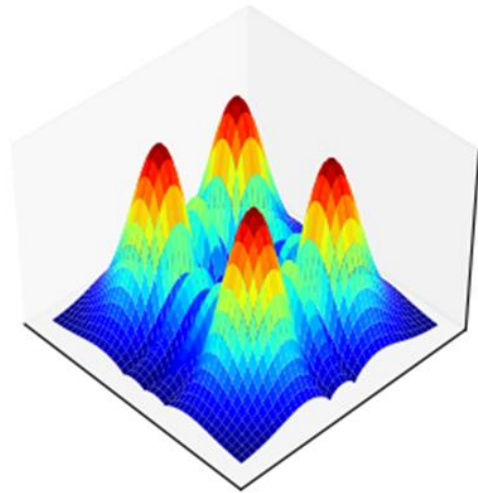
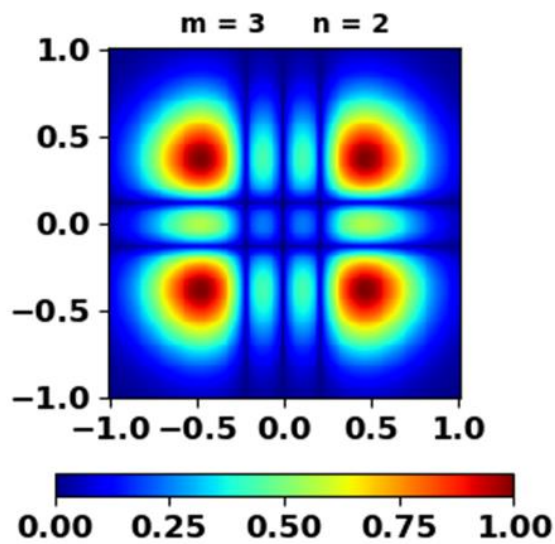


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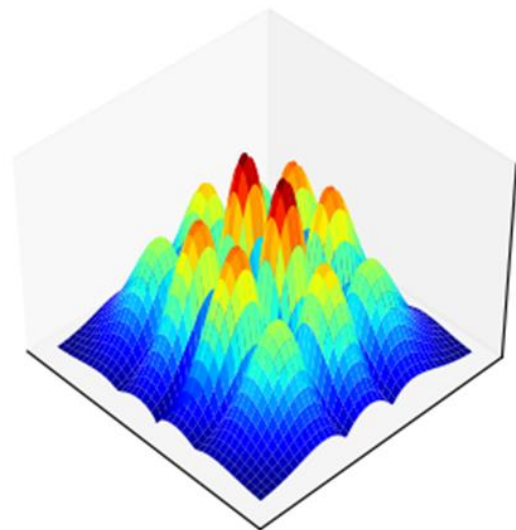
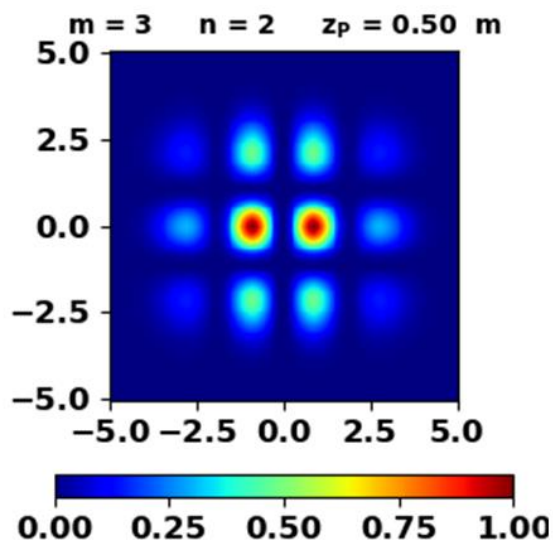


## TEM<sub>32</sub>

### APERTURE SPACE

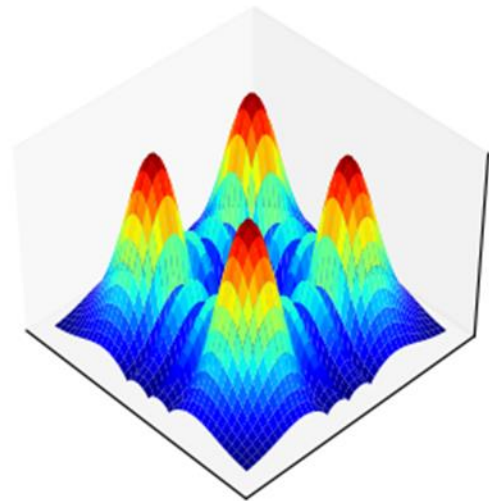
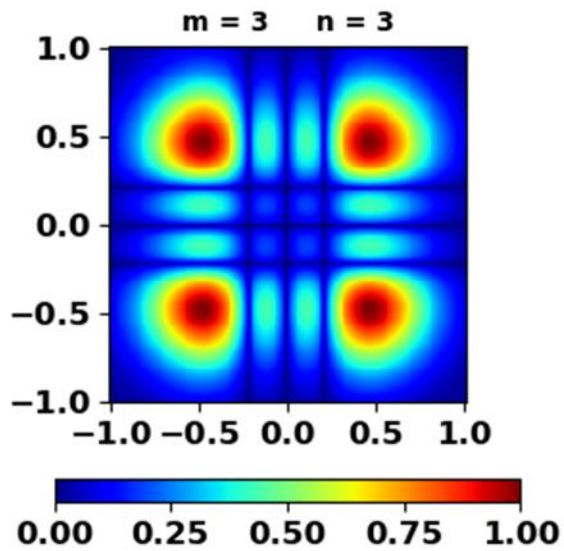


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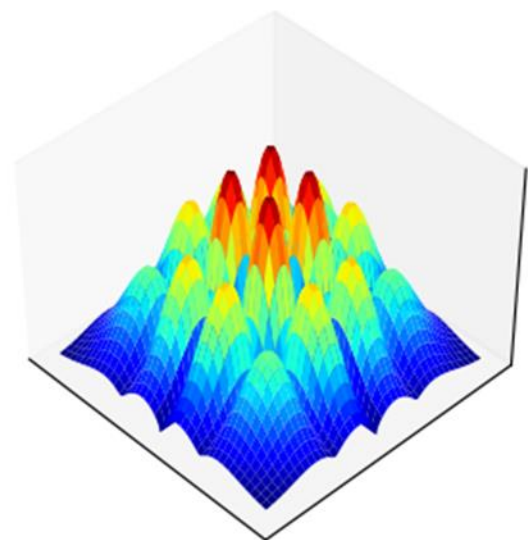
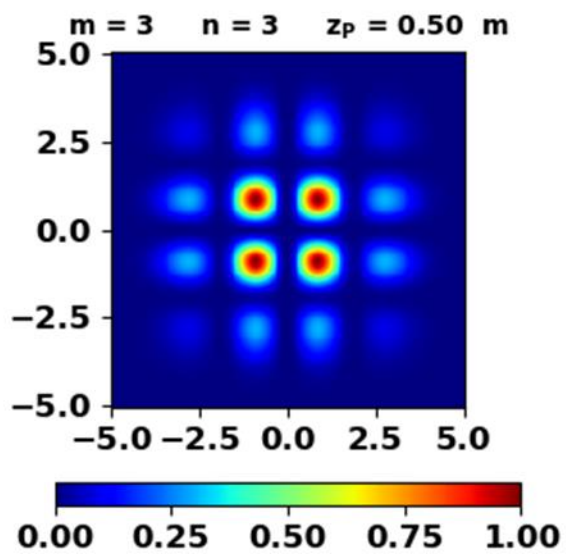


## TEM<sub>33</sub>

### APERTURE SPACE



### OBSERVATION SPACE

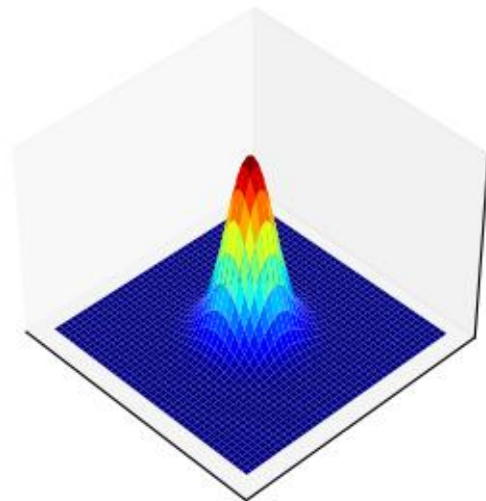
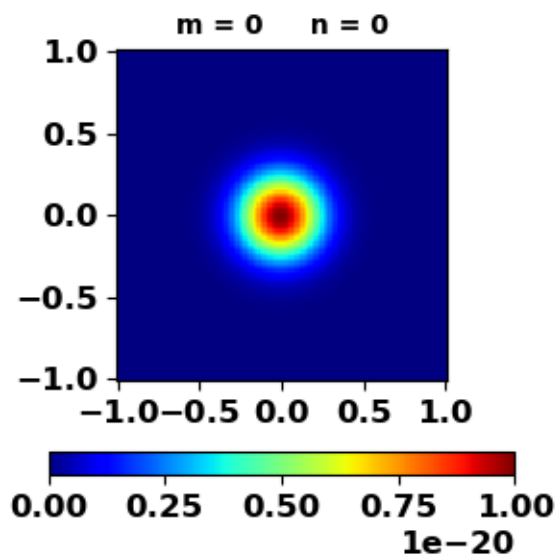
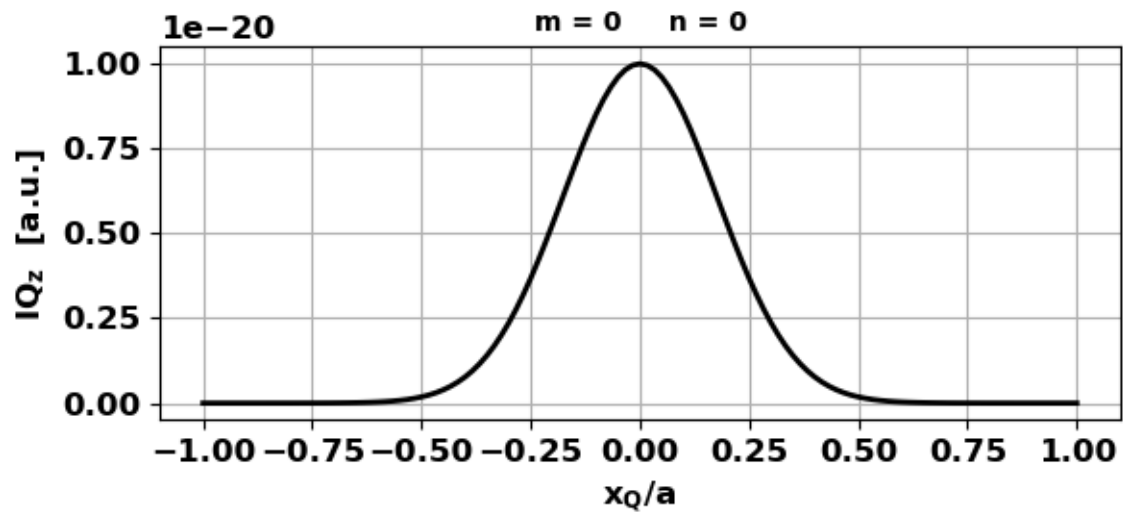




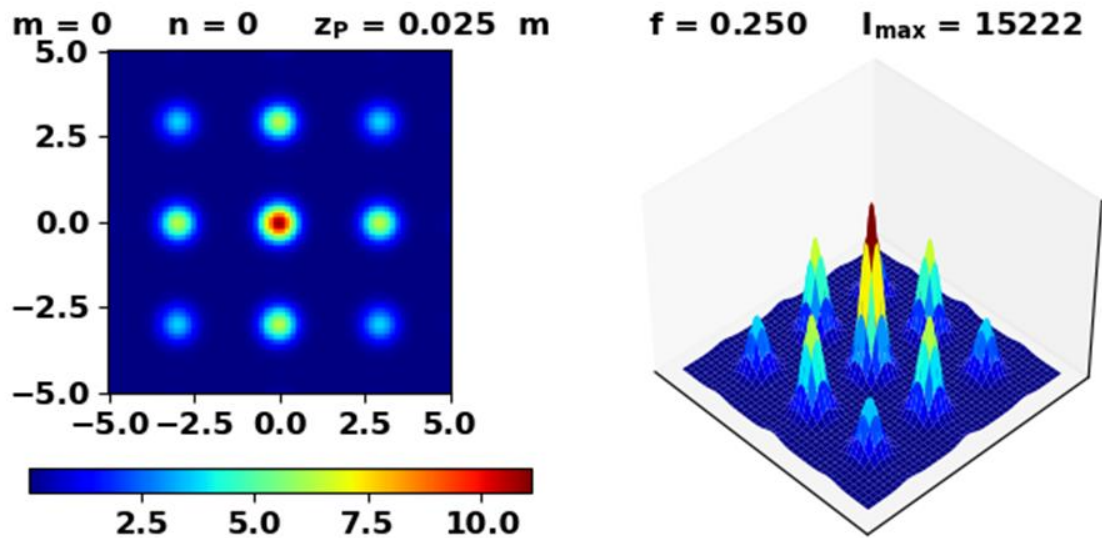
## VARIATION OF IRRADIANCE PATTERN WITH $z_P$

TEM<sub>00</sub> Gaussian beam

APERTURE SPACE

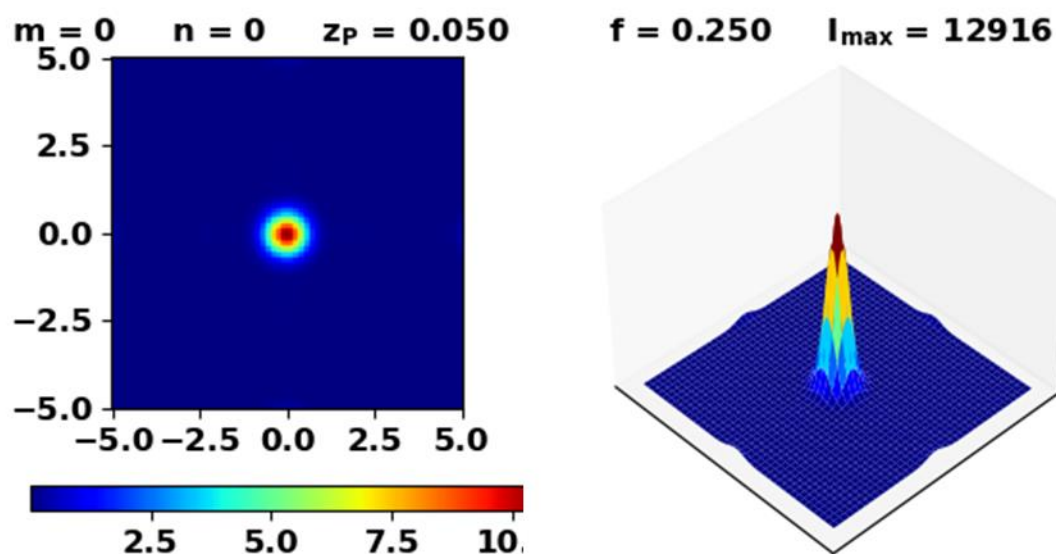


## OBSERVATION SPACE



When the observation plane is very close to the aperture, many peaks in the irradiance occur because of the constructive and destructive interference of the light from different parts of the aperture (Fresnel diffraction).

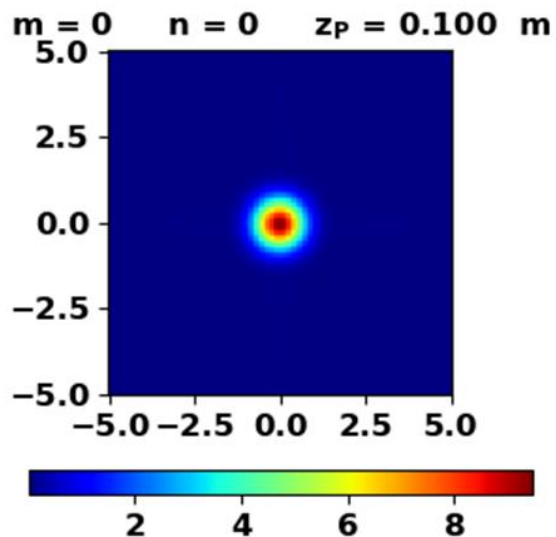
## OBSERVATION SPACE



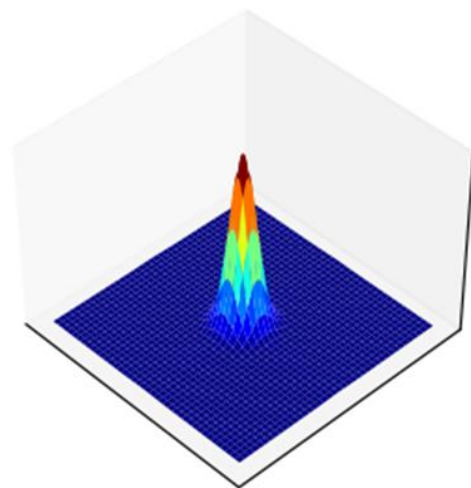


For distances between the aperture and observation plane greater than about 0.05 m ( $z_p > \sim 0.5$  m), the irradiance profile is to a good approximation Gaussian.

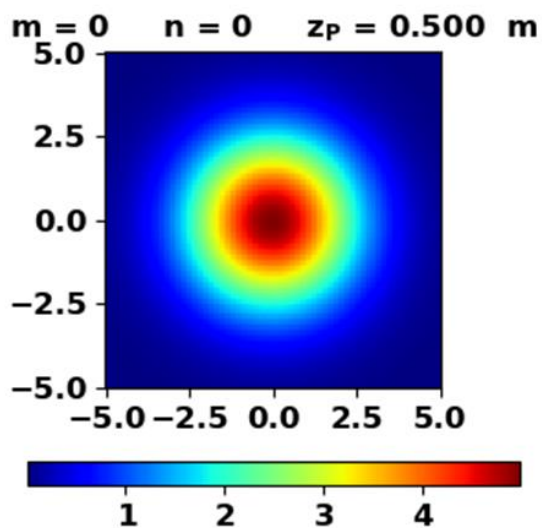
### OBSERVATION SPACE



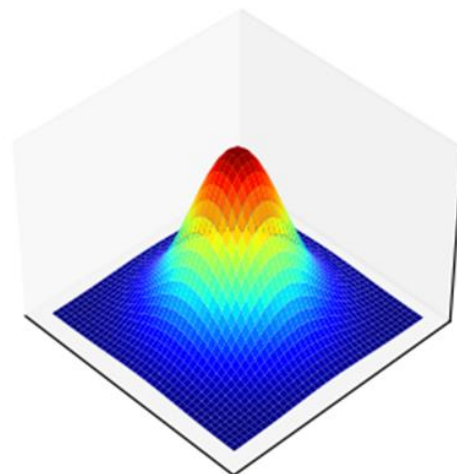
$f = 0.250$   $I_{\max} = 8038$



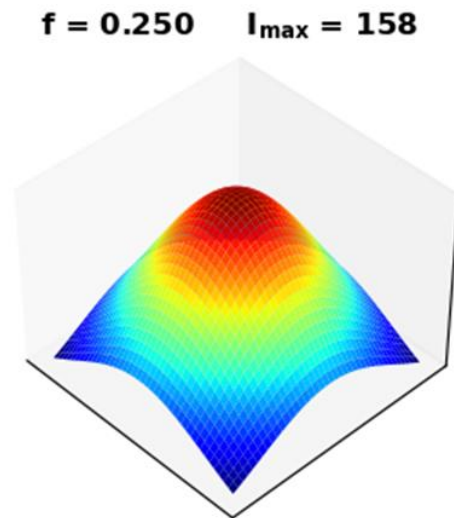
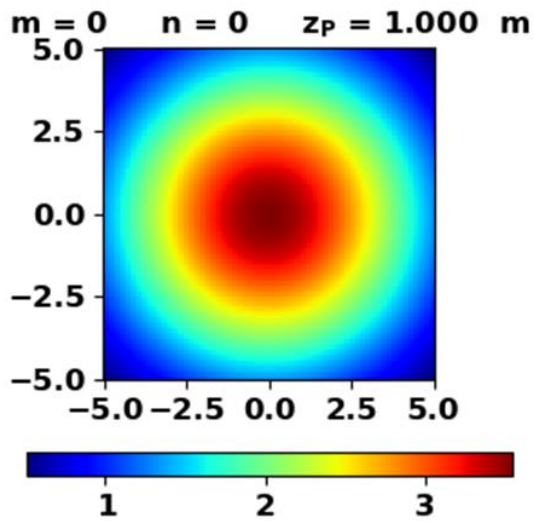
### OBSERVATION SPACE



$f = 0.250$   $I_{\max} = 614$



## OBSERVATION SPACE

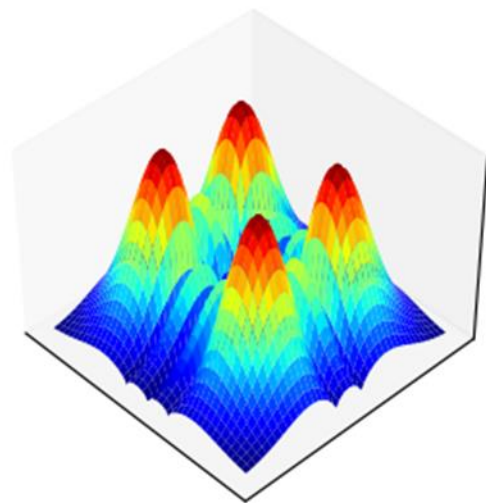
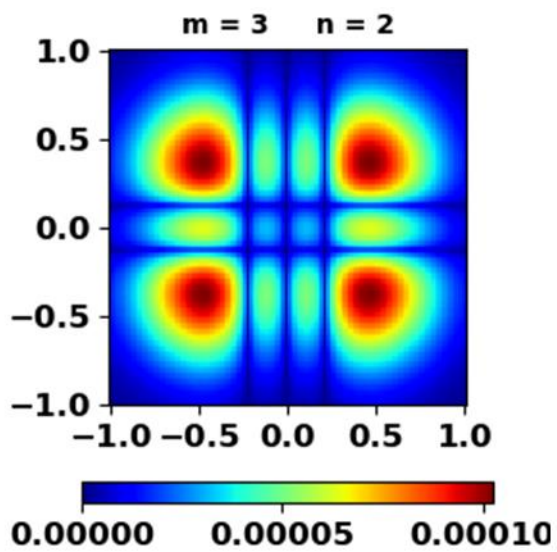
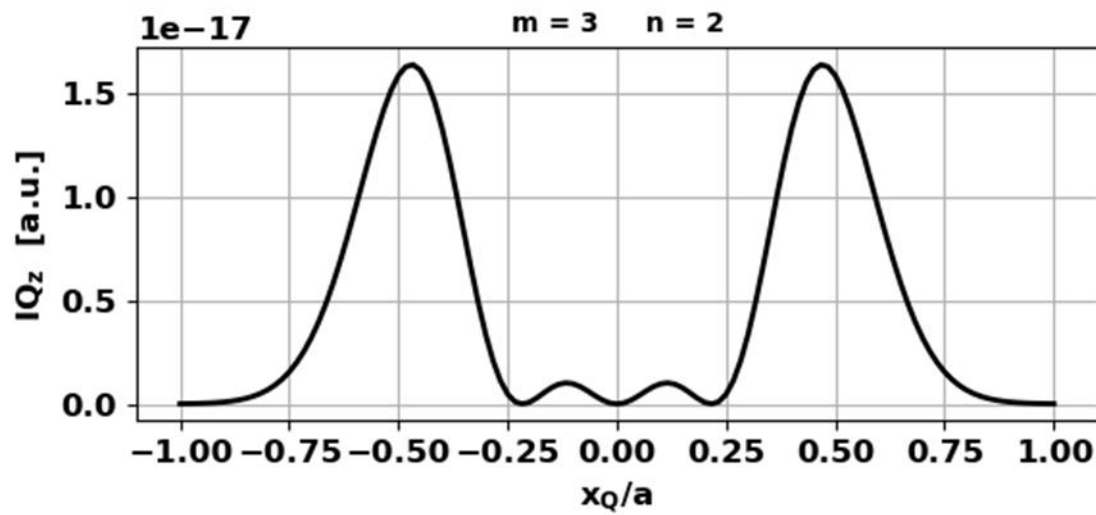


As the observation screen moves further away from the aperture, the Gaussian profile of the irradiance is maintained as it spreads. For large distances between the aperture and observation screen, the aperture behaves like a point source and the inverse square law is valid.

$z_p$ [m]	0.50	1.00	2.00
$I_{\max}$ [a.u.]	614	158	40
		$614/4 = 153$	$158/4 = 40$

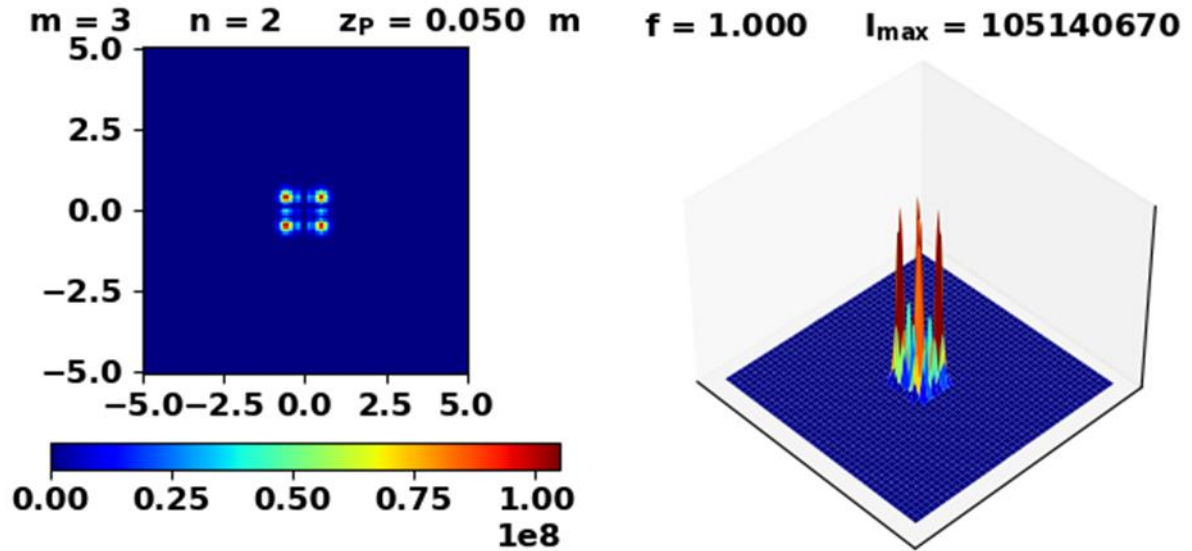
## HIGHER ORDER MODE $\text{TEM}_{32}$

### APERTURE SPACE



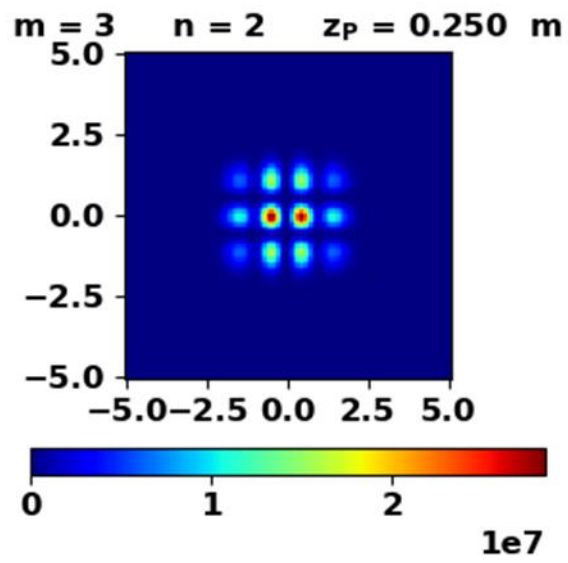
Scaling factor  $f = 0.25$ .

## OBSERVATION SPACE

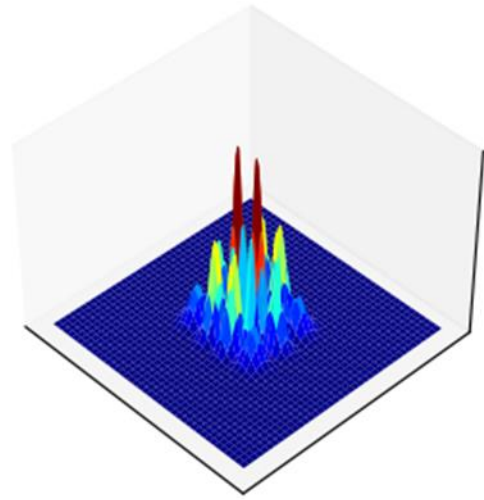


There are four peaks in the observation plane and the irradiance profile is similar to the aperture irradiance profile. However, as the distances between the aperture and observation plane increases, two dominant peaks are found along the X axis.

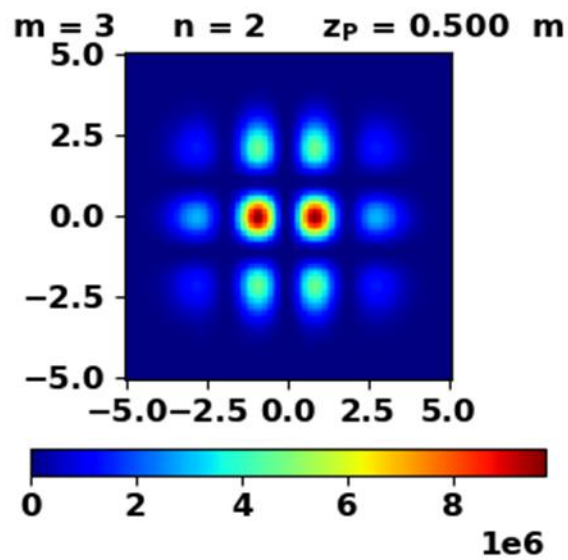
## OBSERVATION SPACE



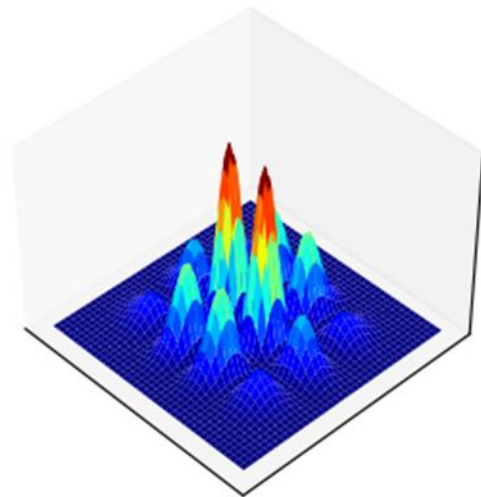
$f = 1.000$     $I_{\max} = 28439386$



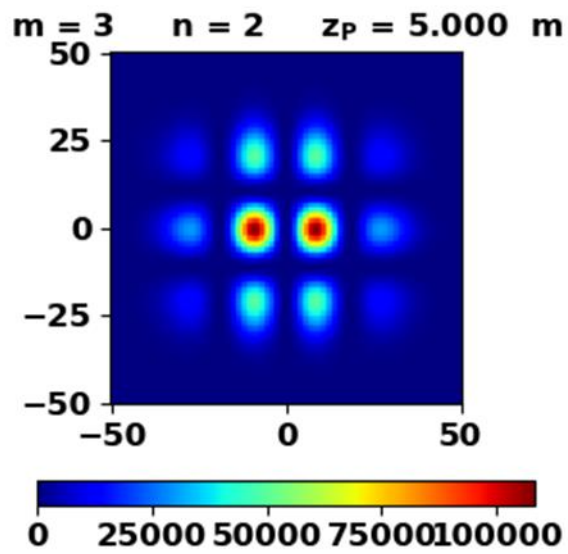
## OBSERVATION SPACE



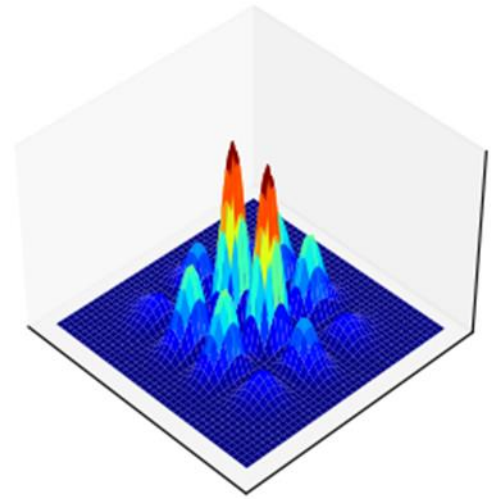
$f = 1.000$     $I_{\max} = 9756724$



## OBSERVATION SPACE



$f = 1.000$     $I_{\max} = 108437$



Even for large values of  $z_P$ , the two predominant peaks along X axis persist and the irradiance profile does not change.