




- Why do we need a statistics and probability course?

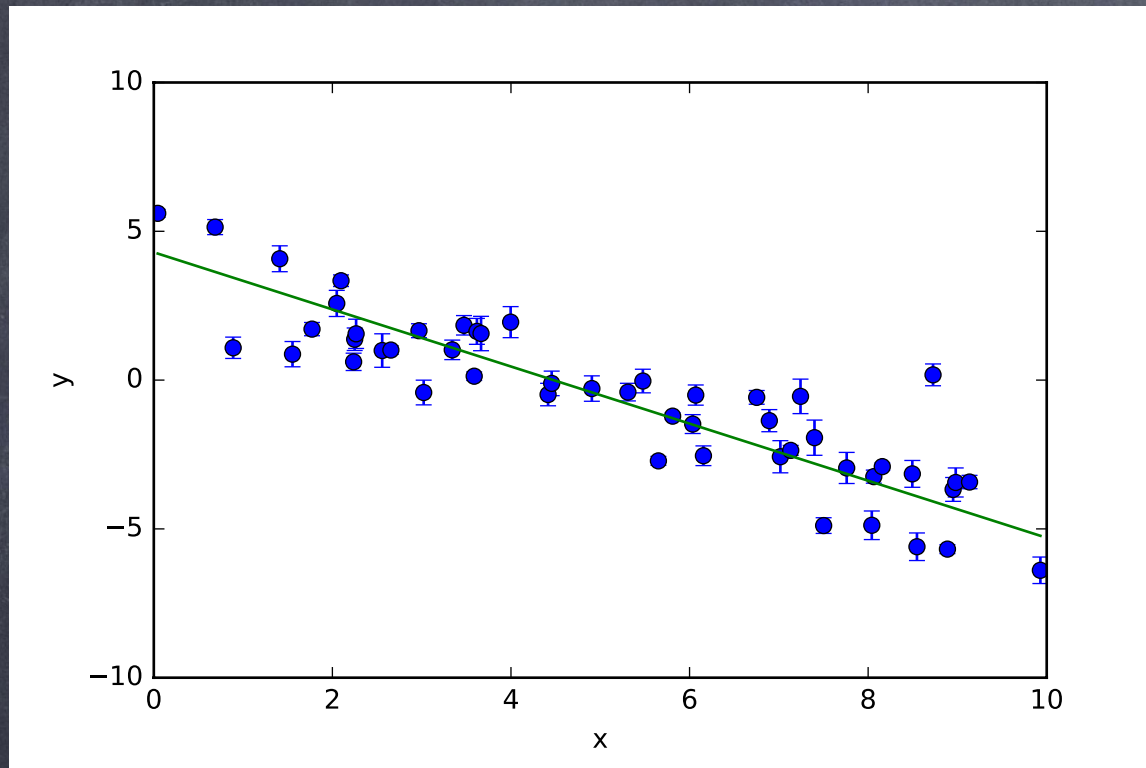
In cosmology and astrophysics most of the problems consist of having a set of data from which we want to INFER something.

- Infer some parameter values.  What is the value the parameters involved in the LCDM paradigm?
- Test an hypothesis.  Is the CMB consistent with a scale free initial power spectrum of fluctuations, and with a gaussian distribution?
- Select a model.  Is General Relativity the correct and final theory, or modified theories works better?



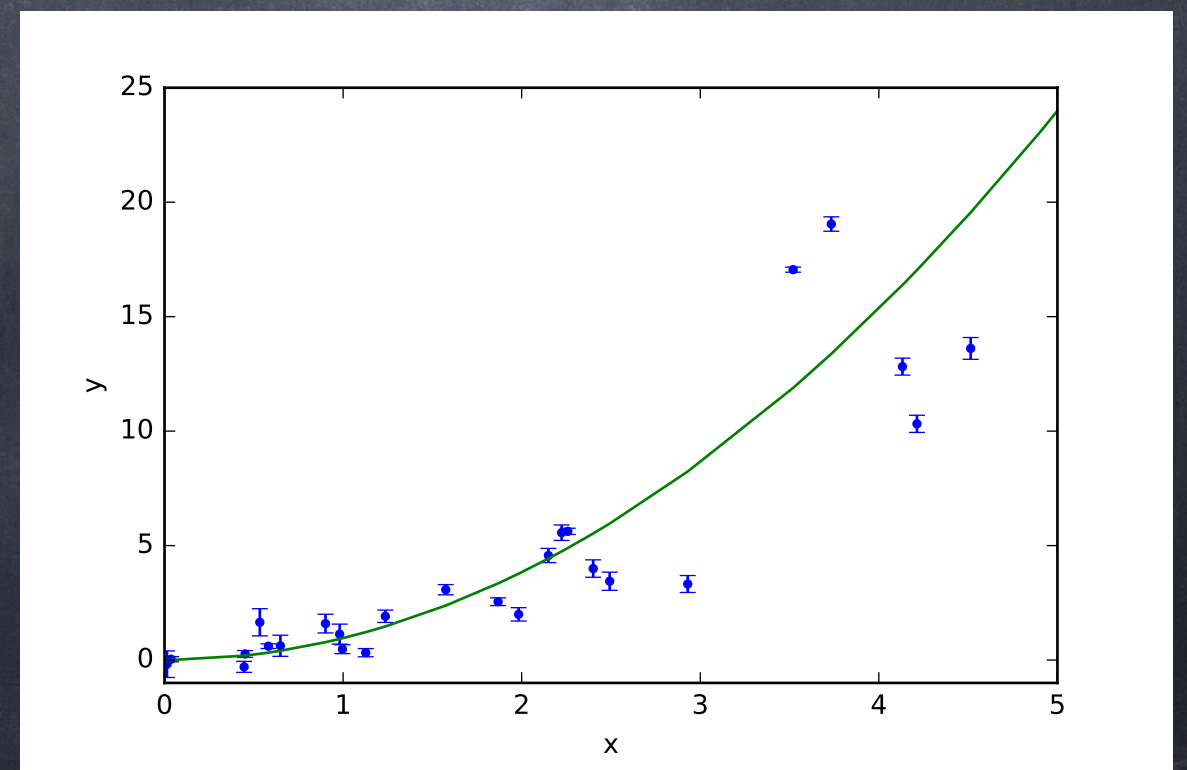
# Parameter Estimation

- What do we do if we want to estimate the slope and y-intercept?



- Linear least square method

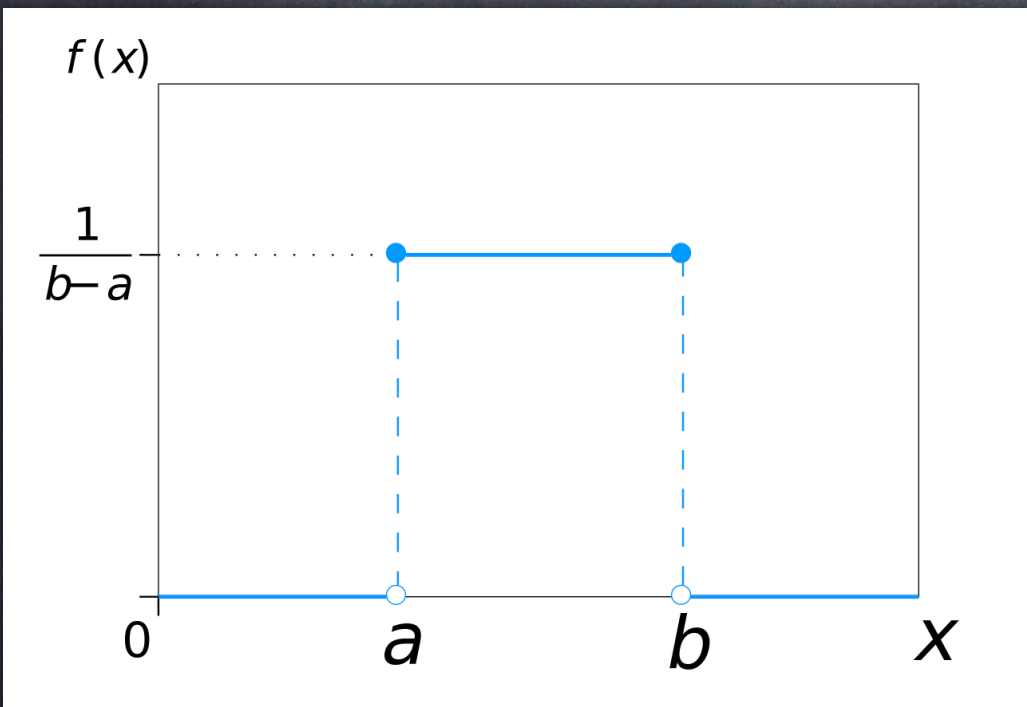
- What if data is not a straight line? And/or model is not linear, and/or if we have more than two free parameters? or more important what if I don't believe too much on the error bars...



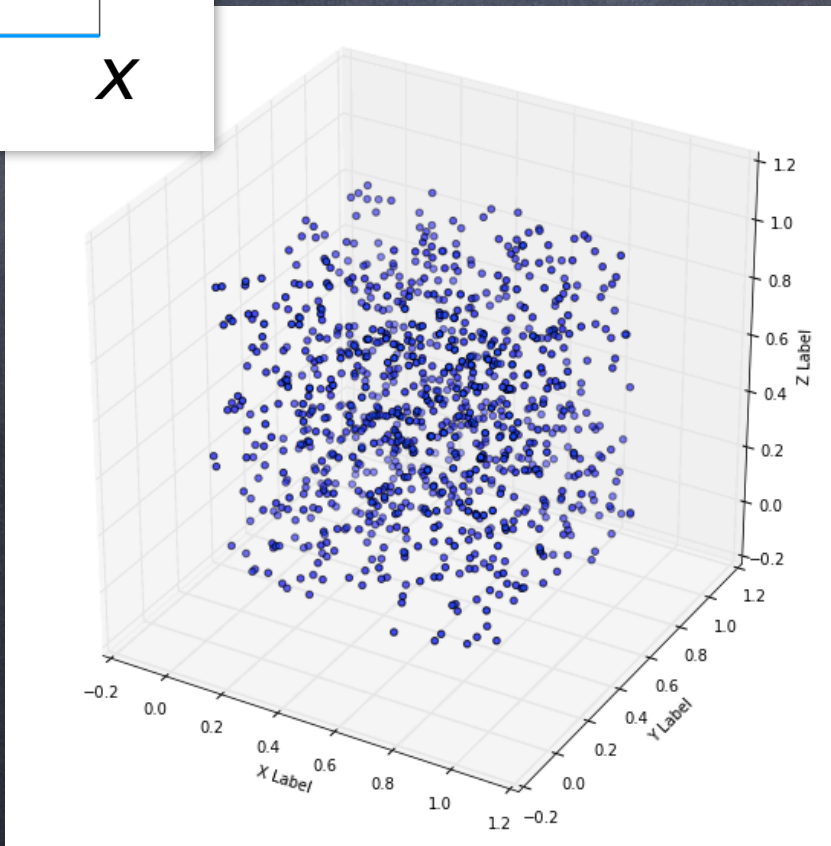


# MonteCarlo uses probability distributions

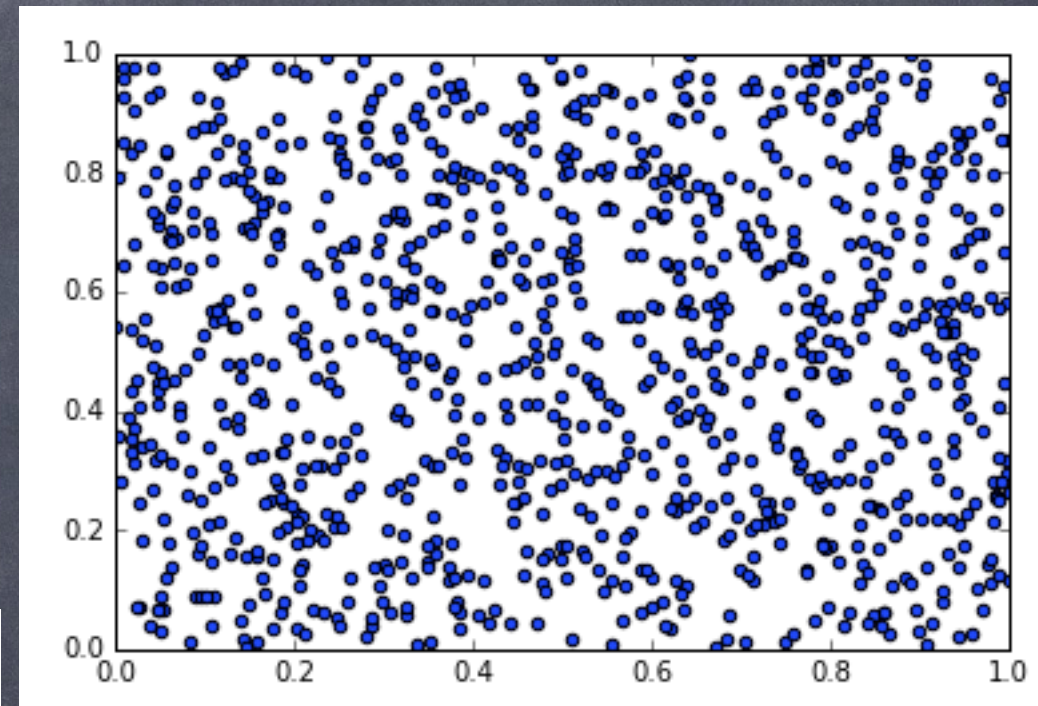
## Probability distribution functions: The basic one: Uniform



1D



3D

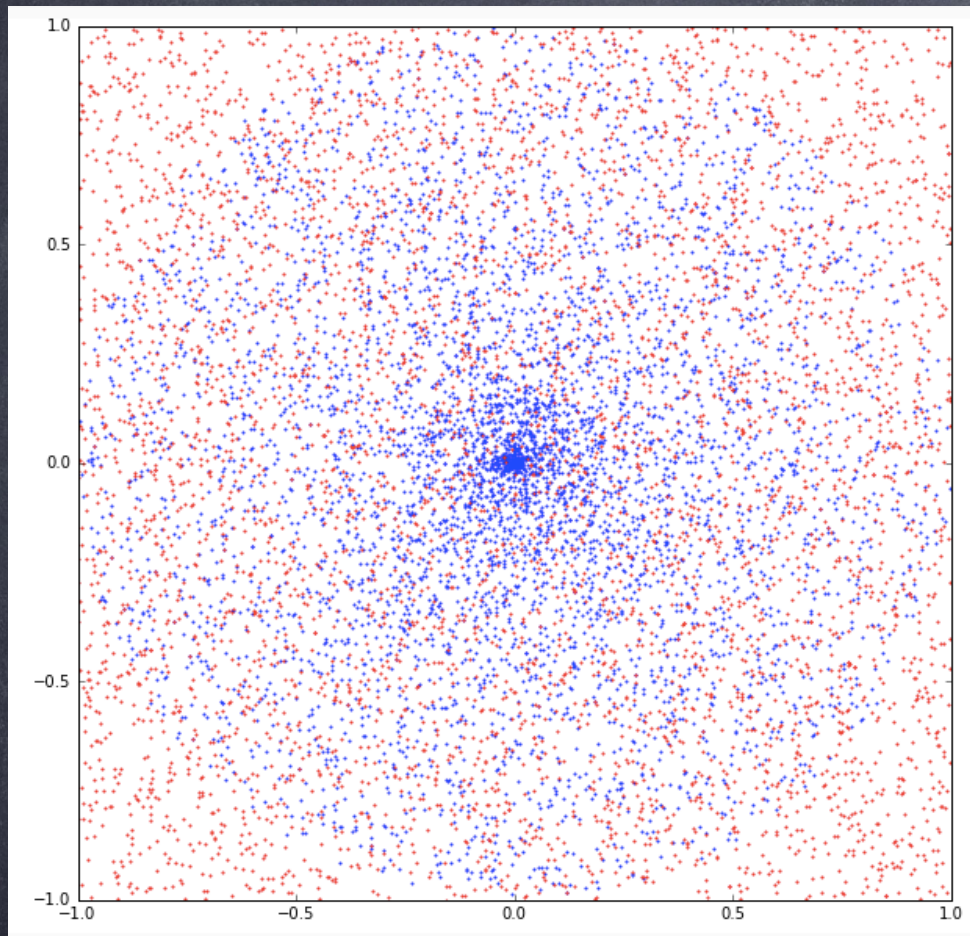


2D

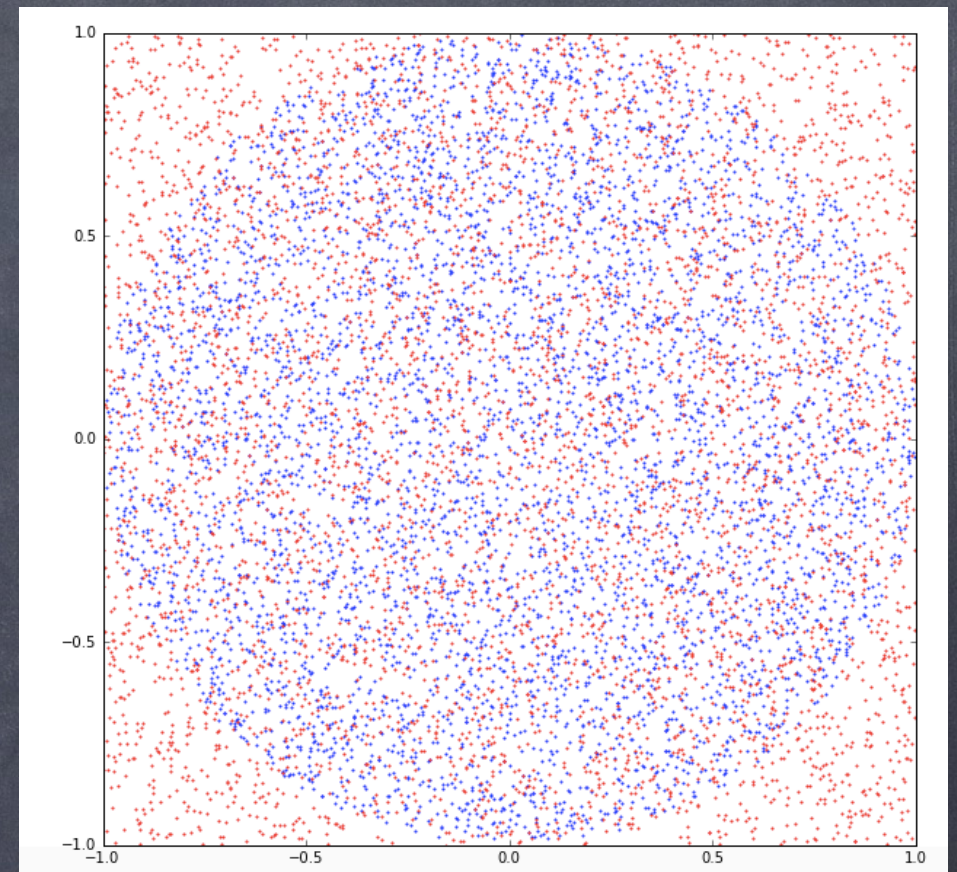
Python  
`numpy.random.rand()`  
`numpy.random.uniform(a,b)`



Draw samples from a specific distribution. Use the Uniform distribution (U). (Ej. 2D)



!=



non-Uniform for  $r < 1$

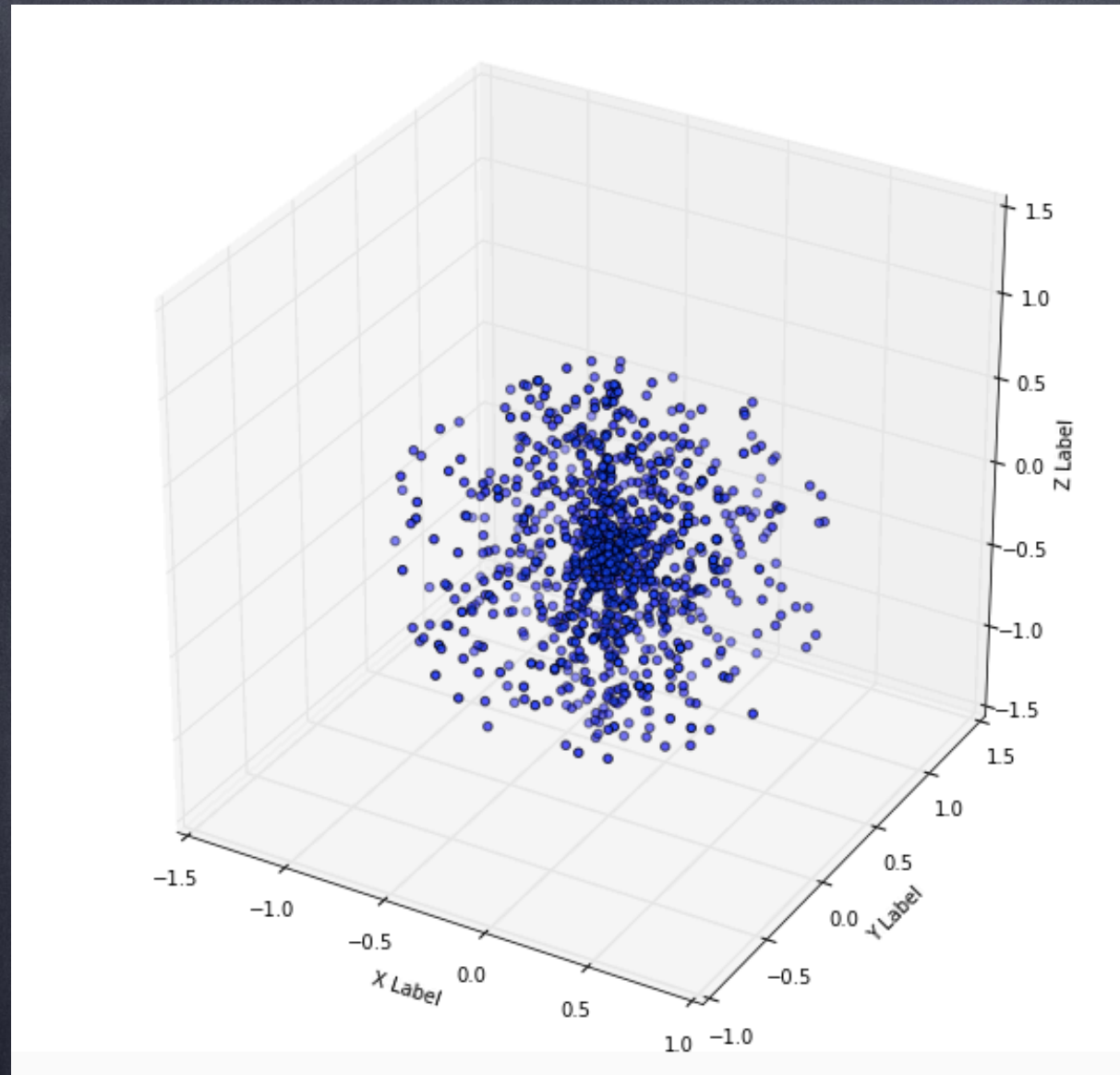
1) generate  $r$  and  $\theta$ , from U.

Uniform for  $r < 1$

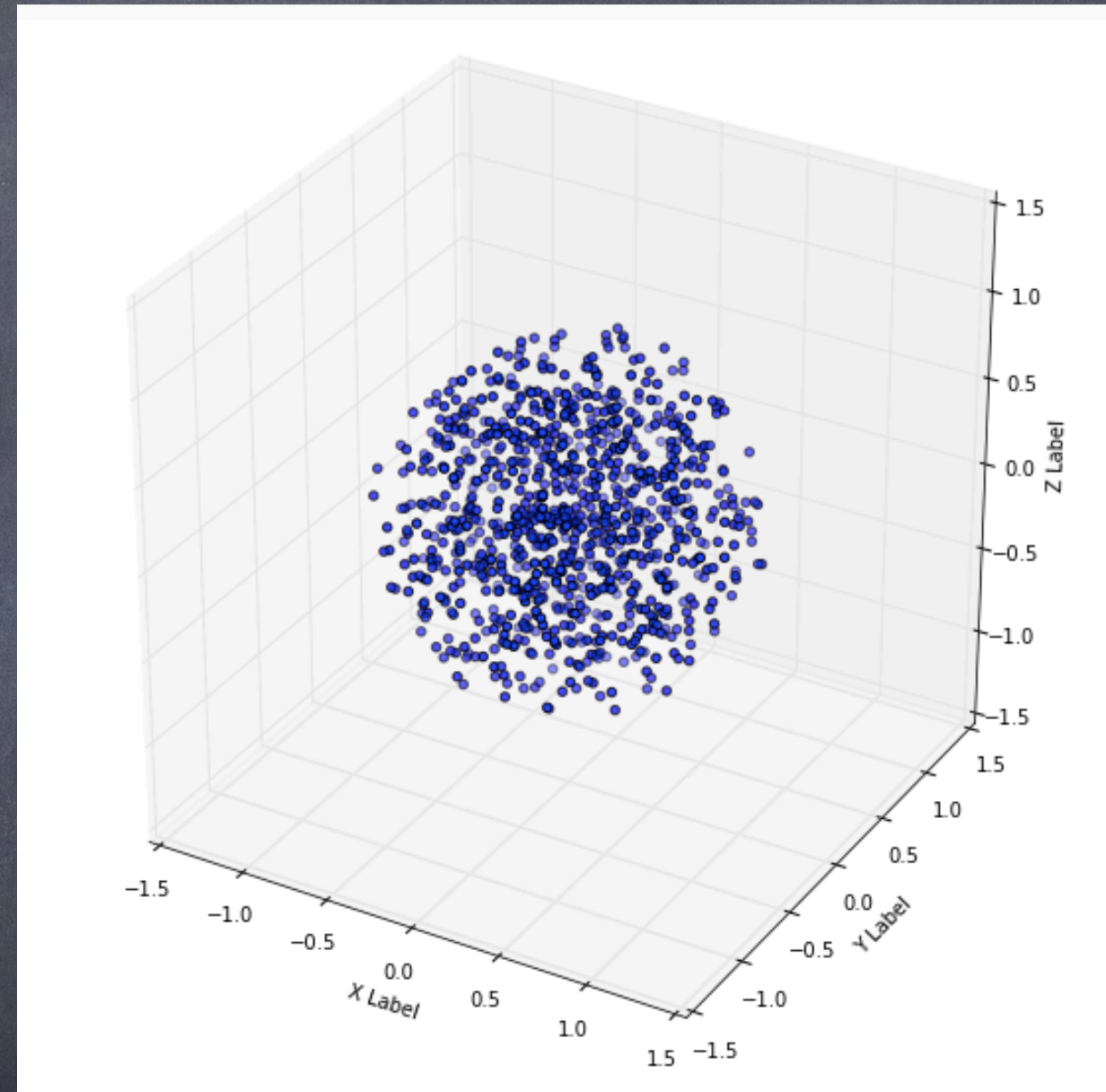
1) generate  $x, y$  from U so that  $r = \sqrt{x^2 + y^2} \leq 1$   
2) generate  $\sqrt{r}$  and  $\theta$ , from U.



Draw samples from a specific distribution. Use the Uniform distribution (U).



1) generate  $r$  and  $\theta$ ,  $\phi$  from U.

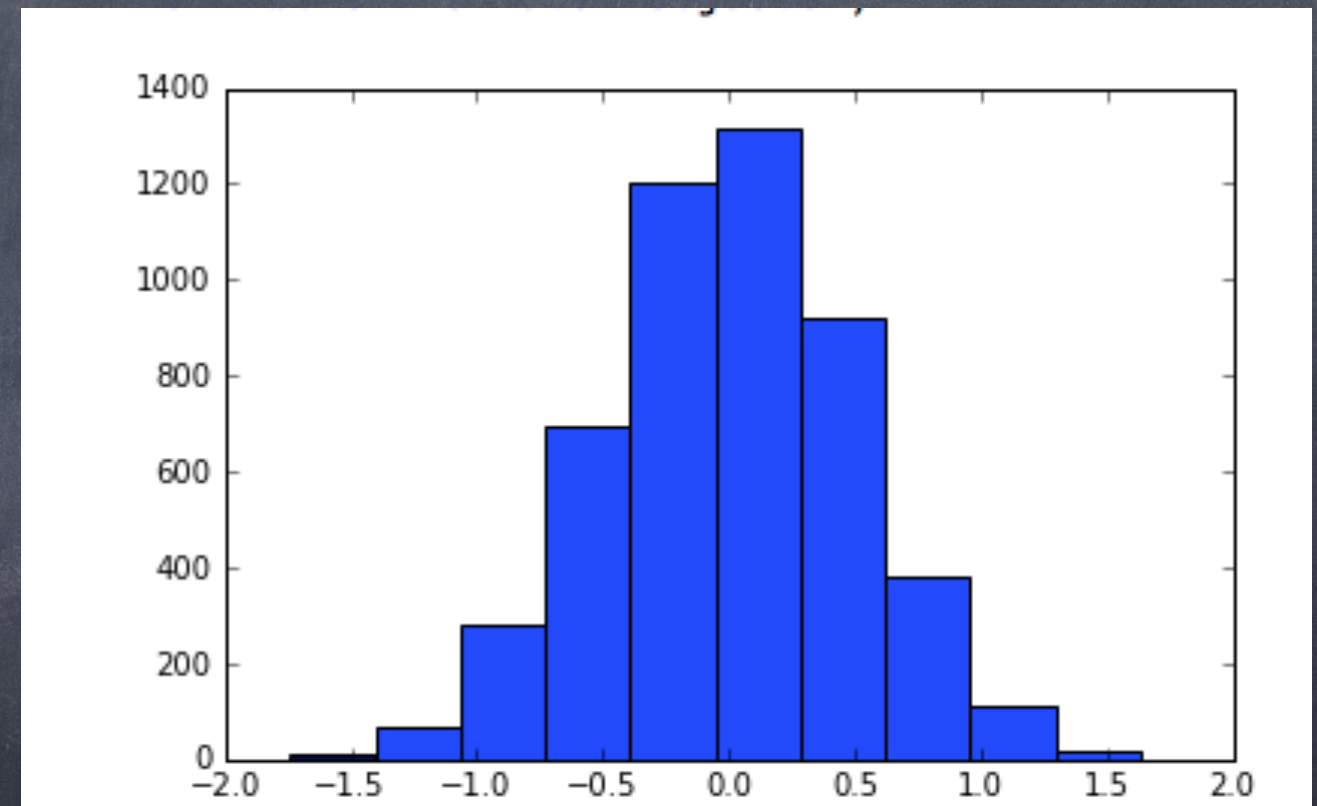
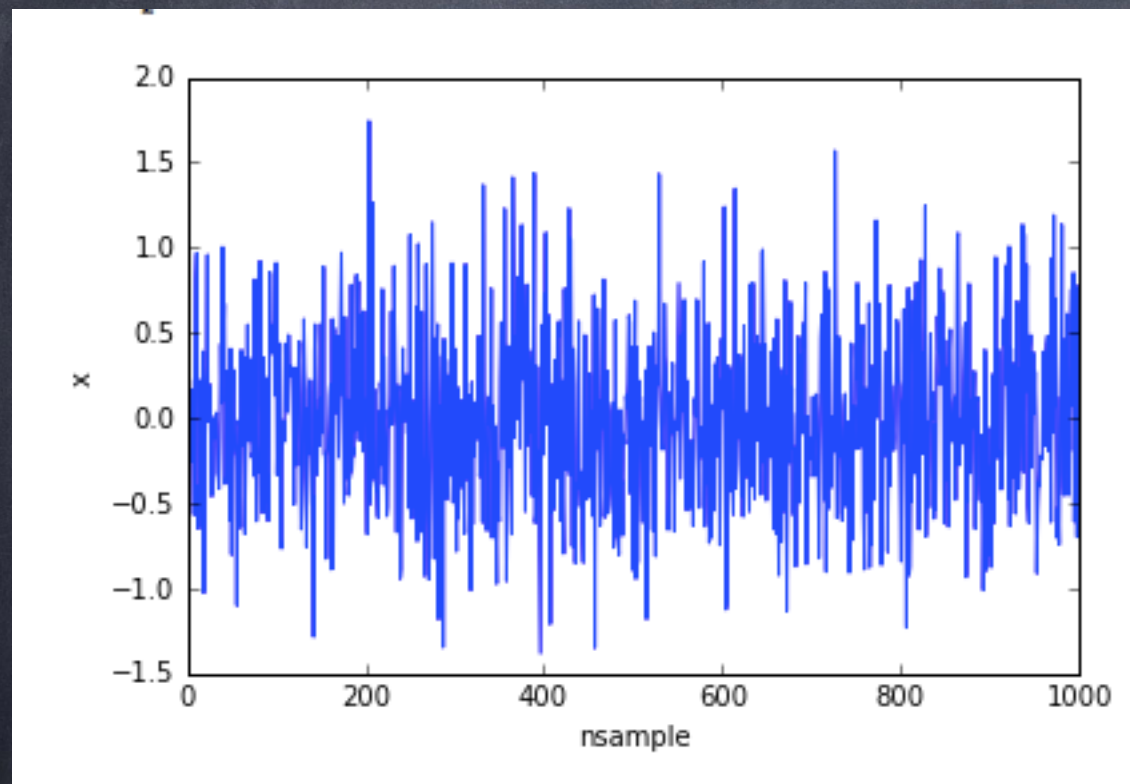


1) generate  $x, y$  from U  
 $r = \sqrt{x^2 + y^2 + z^2} \leq 1$   
2) generate  $\sqrt{r}$ ,  $\theta$  and  $z = \cos(\phi)$  from U.



# Probability distribution functions:

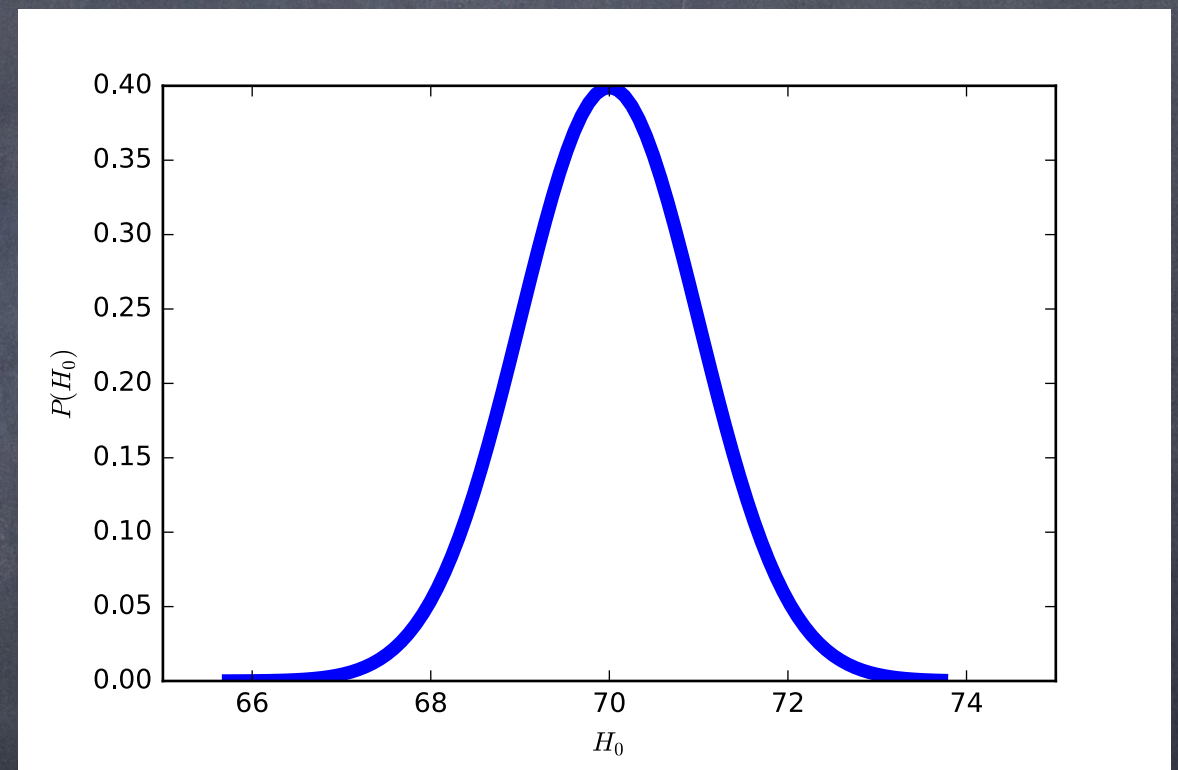
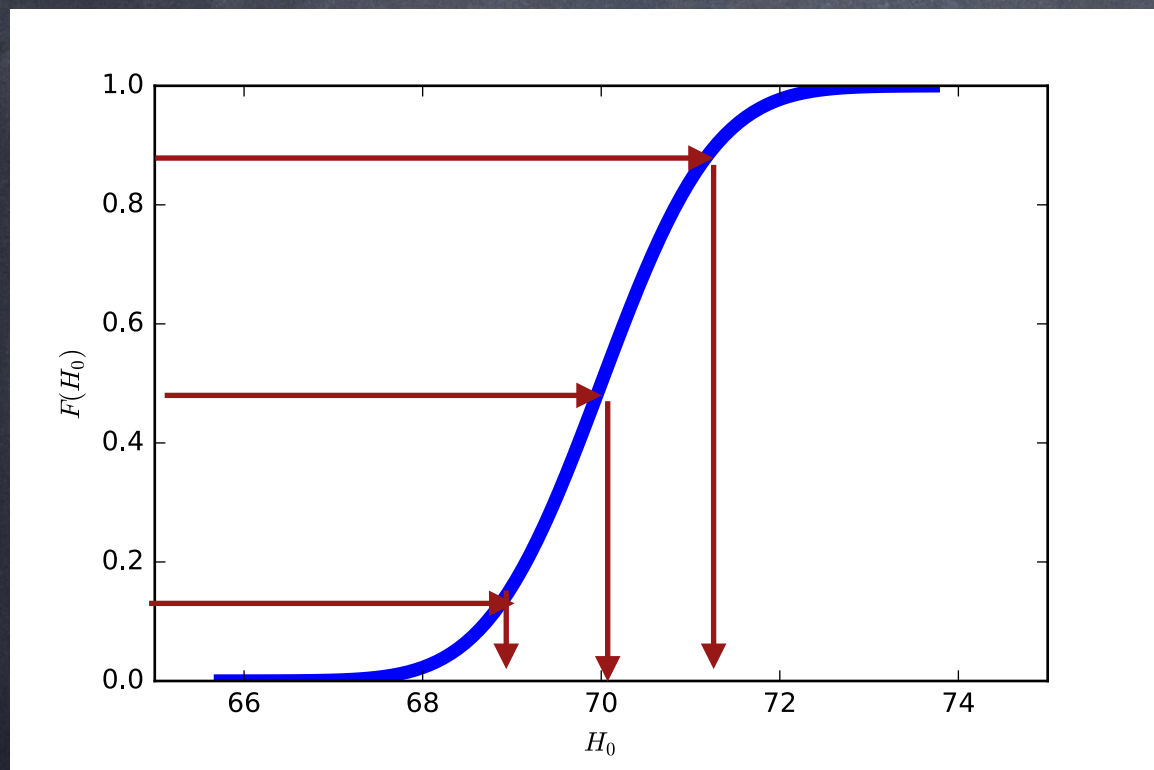
The next basic one: Gaussian





# How to sample a PDF

- Depending on the programming language you are using it can be more or less difficult. But simple method is by using the CDF.
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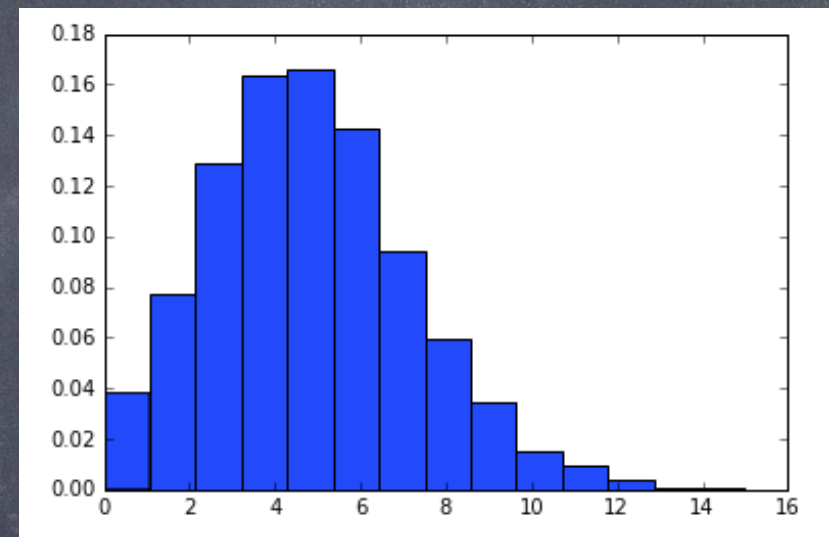
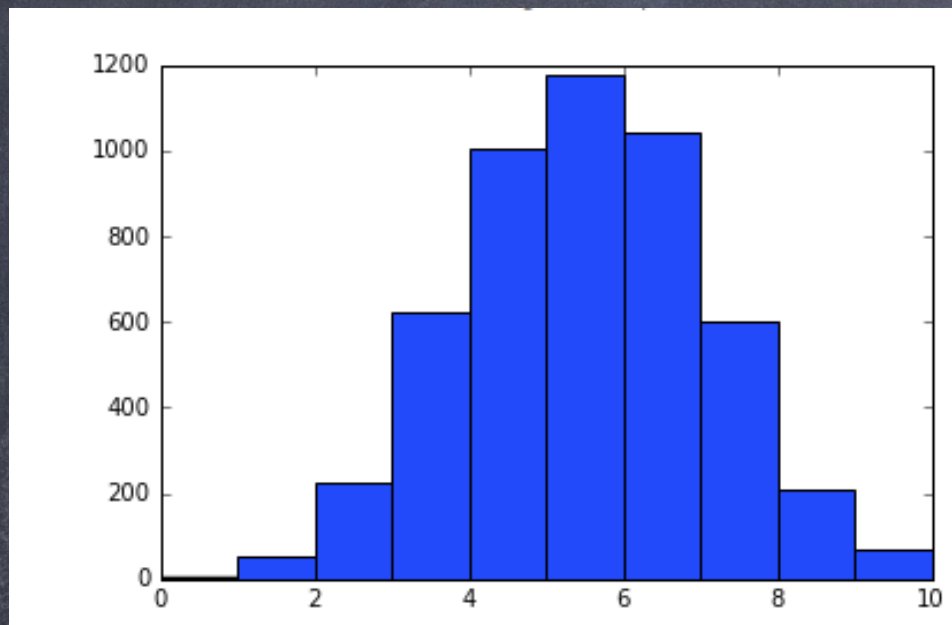
$$F(x) = \int_{-\infty}^x f(x') dx'$$

Choose a random number between 0 and 1, and use the CDF to assign the corresponding value of  $H_0$ . Generate as many as you want, and make the histogram of  $H_0$  to verify you did it right.

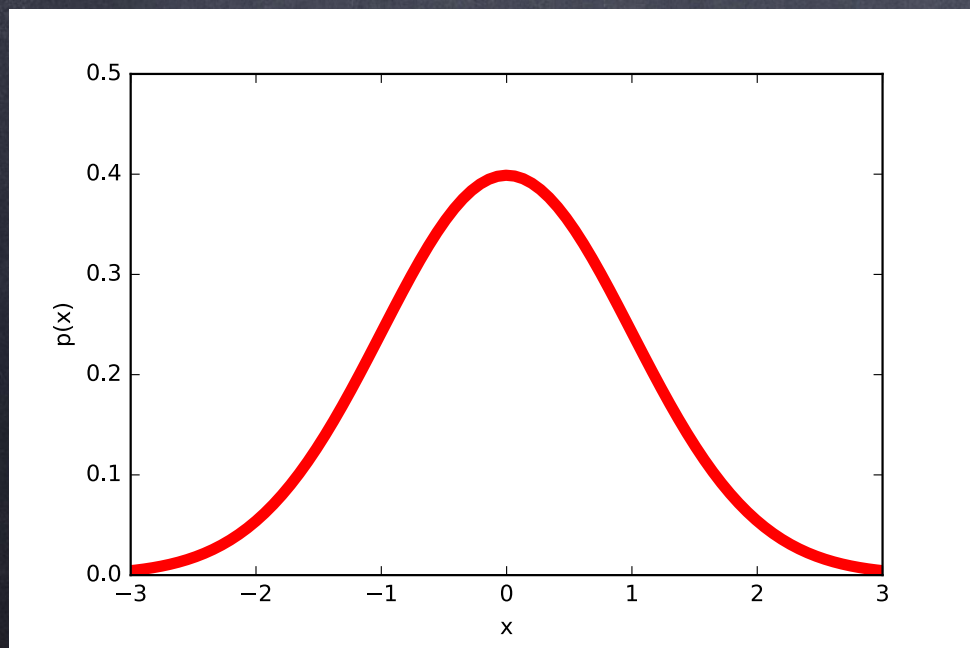


# Probability distribution

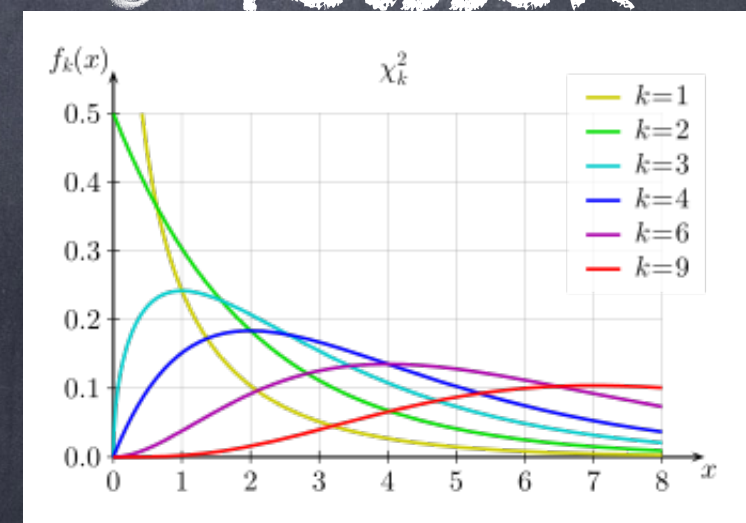
- Binomial, Poisson and  $\chi^2$  distribution can be approximated, for large numbers, by a Gaussian distribution. Other



Poisson



Gaussian



$\chi^2$  distribution



# Parameter Estimation. Level 0

## • Least square method.

$$\begin{aligned} a &= \frac{\sum_{i=1}^n y_i \sum_{i=1}^n x_i^2 - \sum_{i=1}^n x_i \sum_{i=1}^n x_i y_i}{n \sum_{i=1}^n x_i^2 - (\sum_{i=1}^n x_i)^2} \\ &= \frac{\bar{y} (\sum_{i=1}^n x_i^2) - \bar{x} \sum_{i=1}^n x_i y_i}{\sum_{i=1}^n x_i^2 - n \bar{x}^2} \\ b &= \frac{n \sum_{i=1}^n x_i y_i - \sum_{i=1}^n x_i \sum_{i=1}^n y_i}{n \sum_{i=1}^n x_i^2 - (\sum_{i=1}^n x_i)^2} \\ &= \frac{(\sum_{i=1}^n x_i y_i) - n \bar{x} \bar{y}}{\sum_{i=1}^n x_i^2 - n \bar{x}^2} \end{aligned}$$

Where this comes from?

Minimize the residual

$$R^2 \equiv \sum [y_i - f(x_i, a_1, a_2, \dots, a_n)]^2$$

assuming:

- A linear function  $f = ax + b$ .
- Errors are Gaussian and uncorrelated.

Minimization  
implies:

$$\frac{\partial R^2}{\partial a_i} = 0$$



# Parameter Estimation and optimization

$$\chi^2 = \sum (y_i - y(x_i, \theta))^2 / \sigma_{y_i}^2 \quad \frac{\partial \chi^2}{\partial \theta} = 0$$

$\theta$  : free parameters

$\sigma_{y_i}$  : variance on  $y_i$

Chisq minimization becomes difficult (sometimes impossible) when the number of parameters increases...  
We can do the minimization of  $\chi^2$  with MonteCarlo sampling.

Least square, and minimum  $\chi^2$  methods are just special cases of Statistical Inference. This  $\chi^2$  is a gaussian distribution if data points are independent, and errors are also gaussian.



# Likelihood

- The **probability**, under the assumption of a model/theory, to observe the data as was actually obtained.

$$\mathcal{L} \longrightarrow P(\text{Data}, \text{Model})$$

- For data that can be thought as samples of a sequence of normal random variables that have a mean and a variance, the likelihood is Gaussian

$$\mathcal{L} \propto \prod_i^n \frac{1}{2\pi\sigma_i^2} \exp\left(-\frac{(x_i - \mu)^2}{2\sigma_i^2}\right)$$

$\mu$  will be the expected mean given our model

- A minimization of the Chi-square correspond to the maximization of the likelihood.



# Gaussian Likelihood

$$-\ln(\mathcal{L}(\vec{x}, \vec{y}|\vec{\theta})) \propto \frac{1}{2} \sum_i \left( \frac{(y_i - \lambda(x_i, \vec{\theta}))^2}{\sigma_i^2} \right)$$

In terms of data points and parameters.

Lambda is our model for  $y_i$

How do we maximize the likelihood if there 2,3, or more parameters...?

How do we maximize the likelihood if we have a complex model?

What if the likelihood is not Gaussian...?



# MonteCarlo Markov Chain

Draw random samples and accept them or reject them according to the Likelihood.

If the likelihood of a new sample is higher than the previous one we accept the sample and save it.  $\text{new} \rightarrow \text{old}$

If the likelihood of a new sample is lower than the previous one, we then draw a random number  $U(0-1)$ , if the ratio of likelihoods ( $\text{new}/\text{old}$ ) is larger than such number then we accept it, not otherwise.

Draw a new sample in de vicinity of the previous one, and start again....

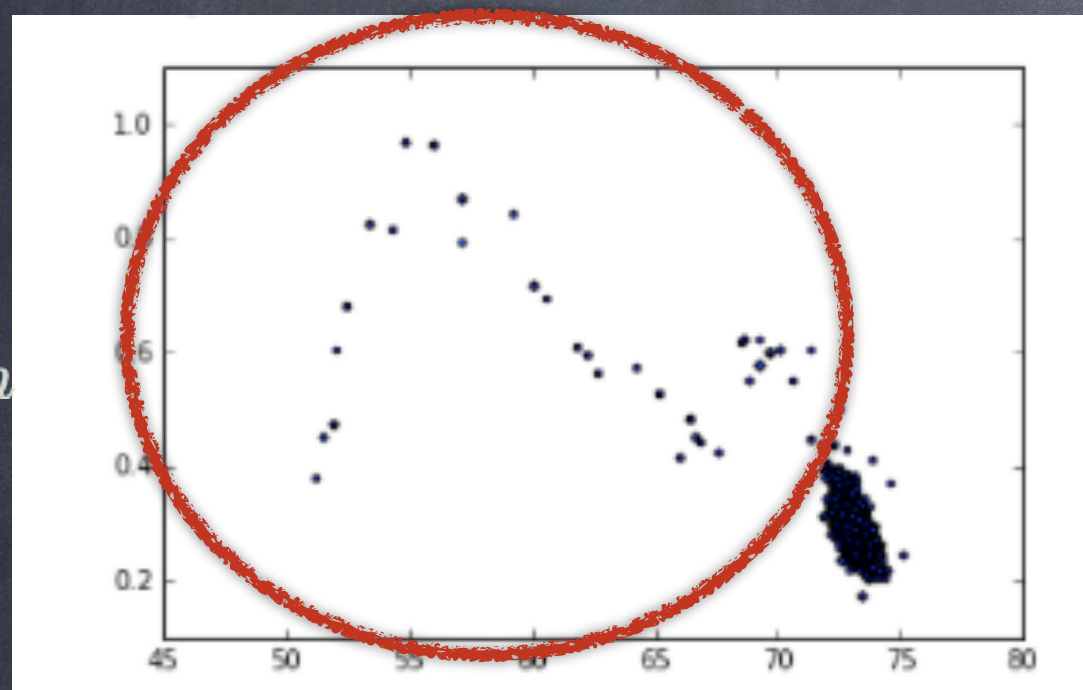
After many steps, look at the resultant distribution (the chains) of parameters, i.e., the likelihood..

Look at the burning period and the convergence....



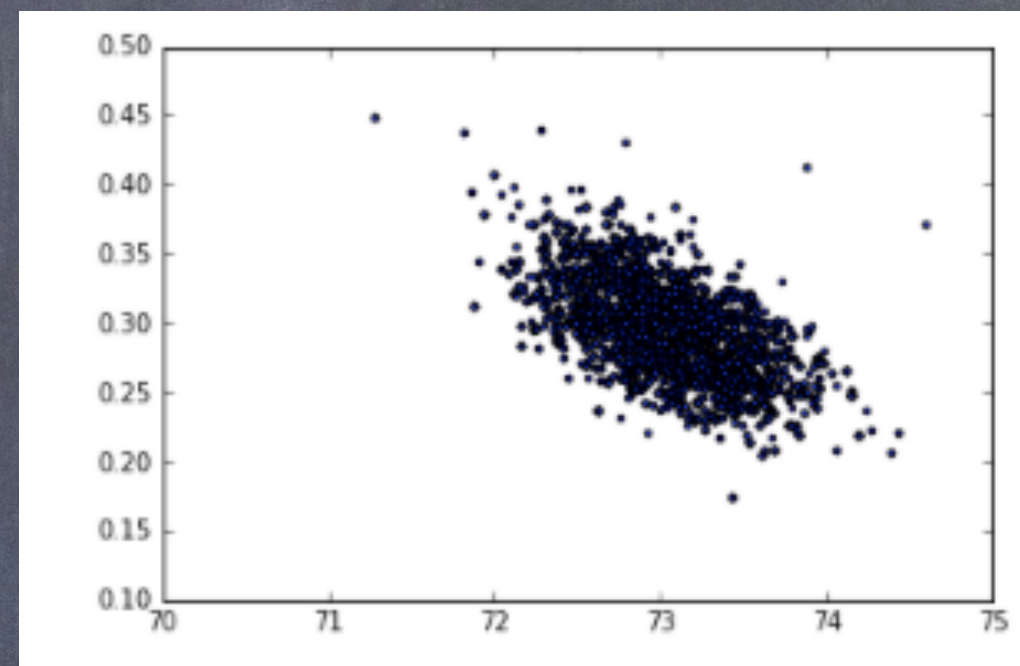
# Walker

$\Omega_m$

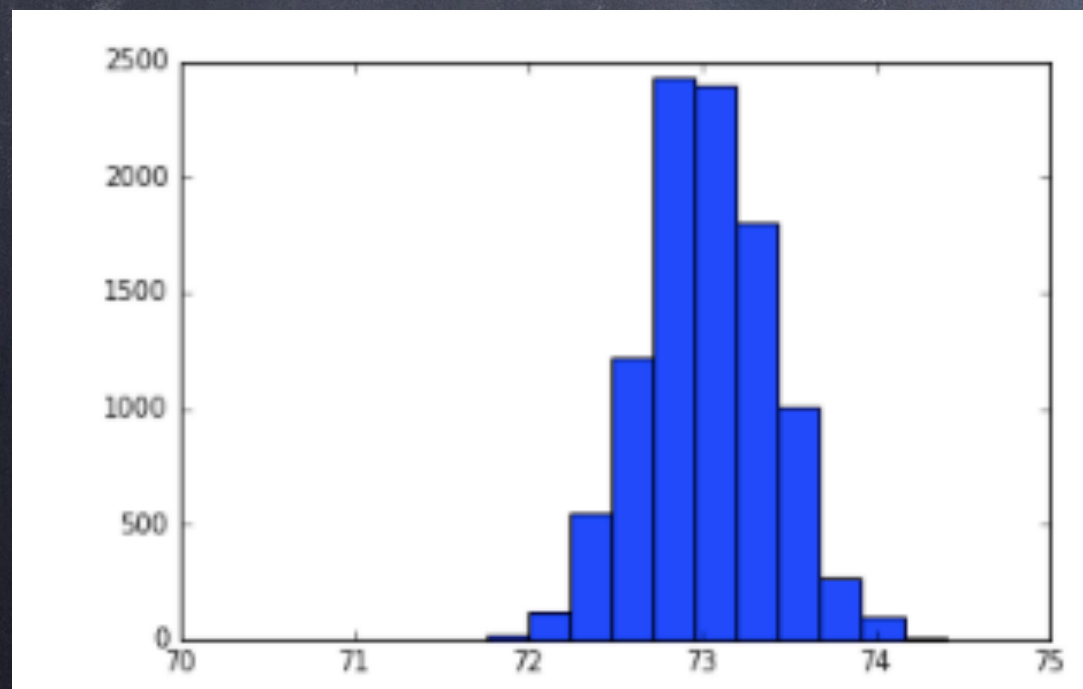


$H_0$

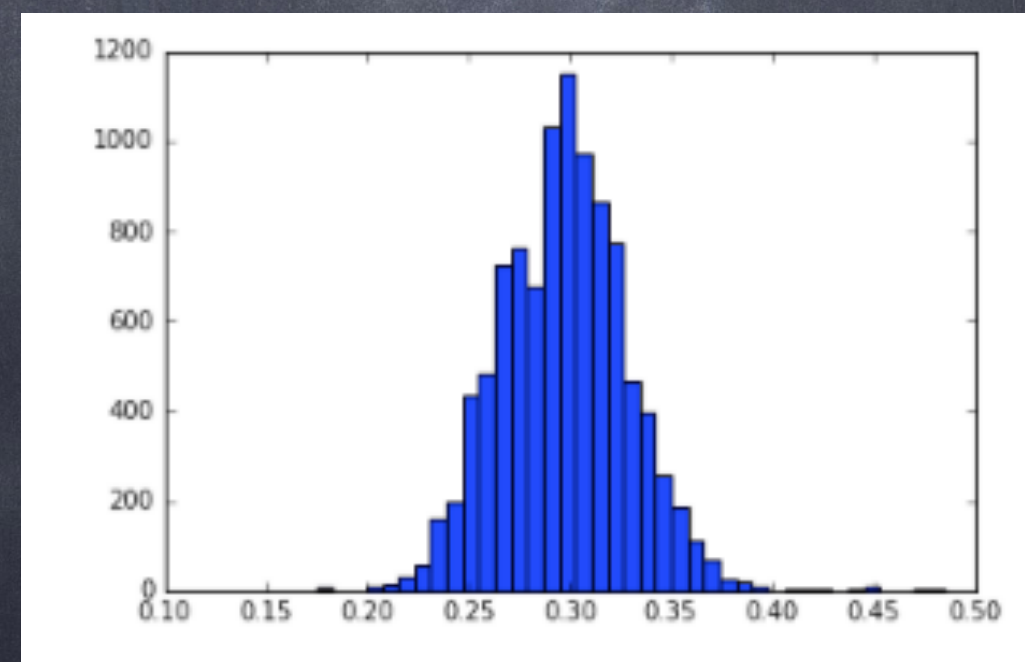
$\Omega_m$



$H_0$



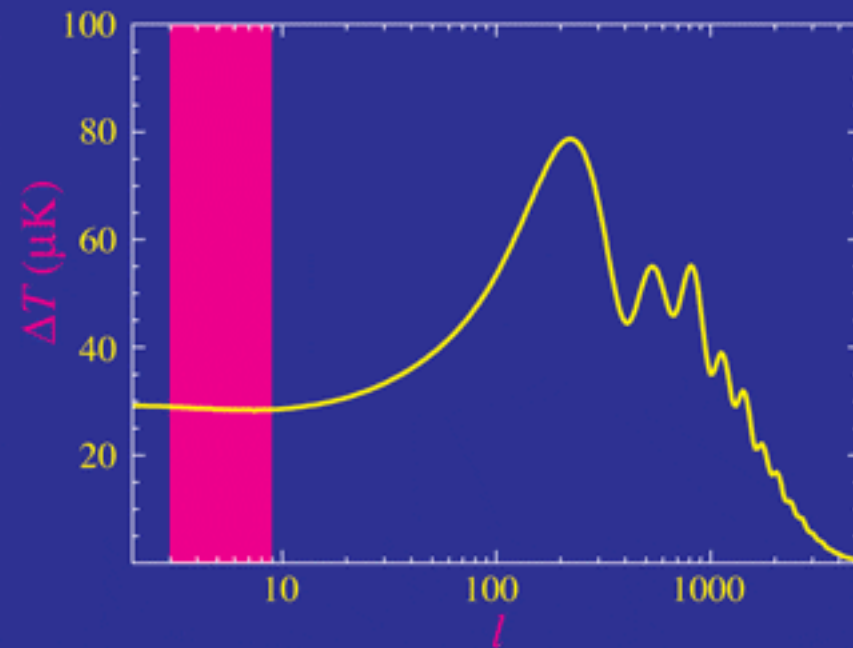
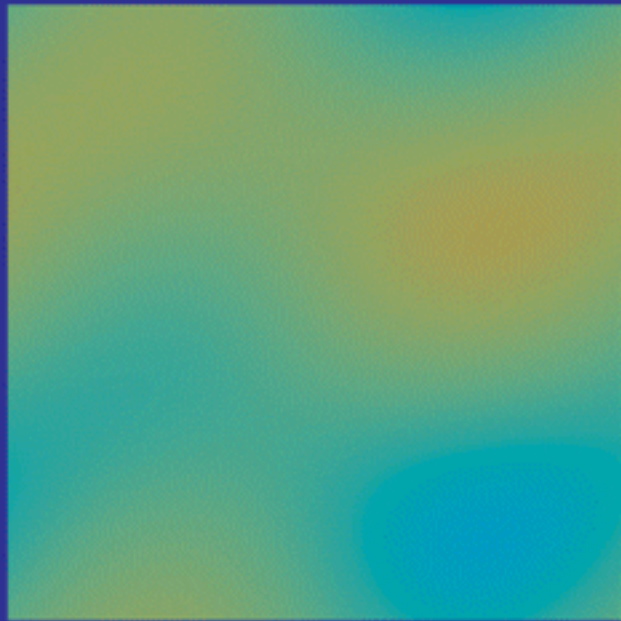
$H_0$



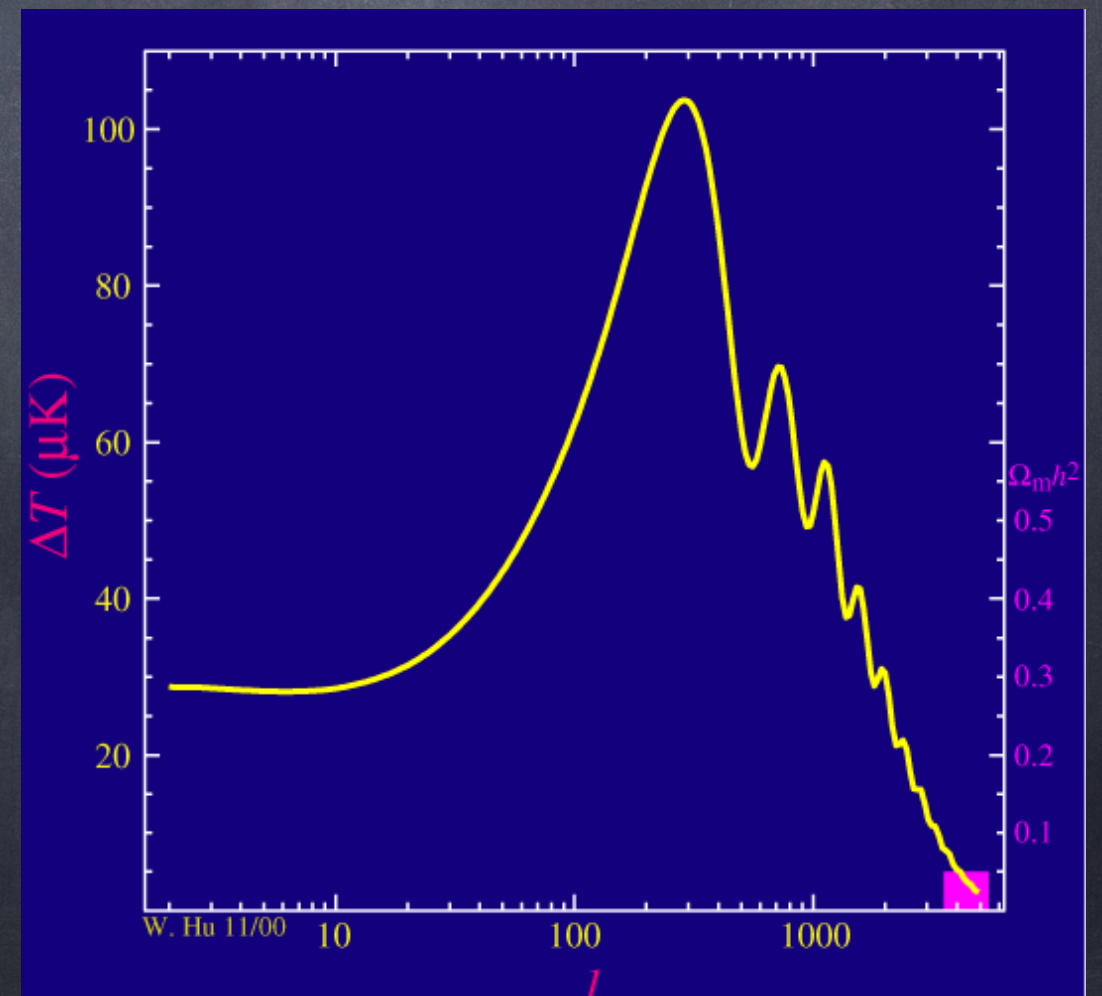
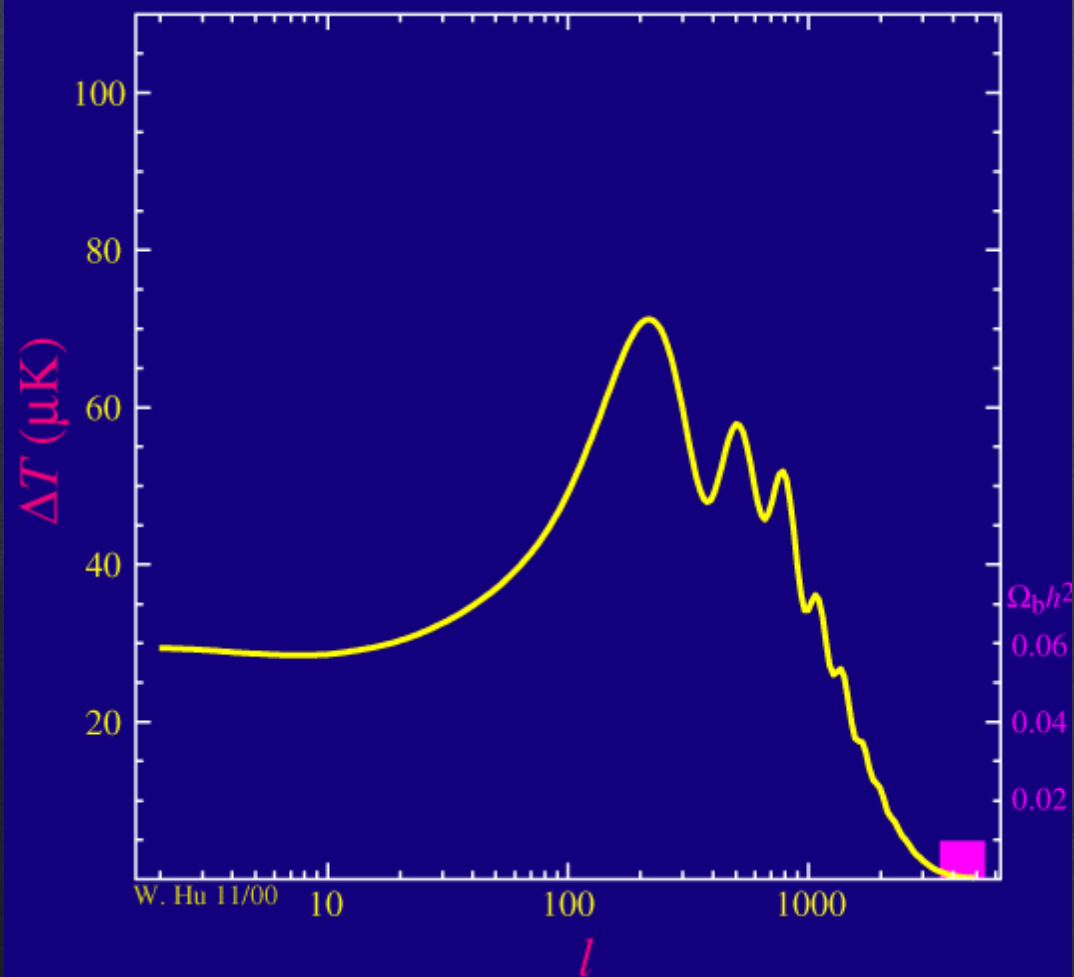
$\Omega_m$



# Ej. CMB DATA and Cosmological model

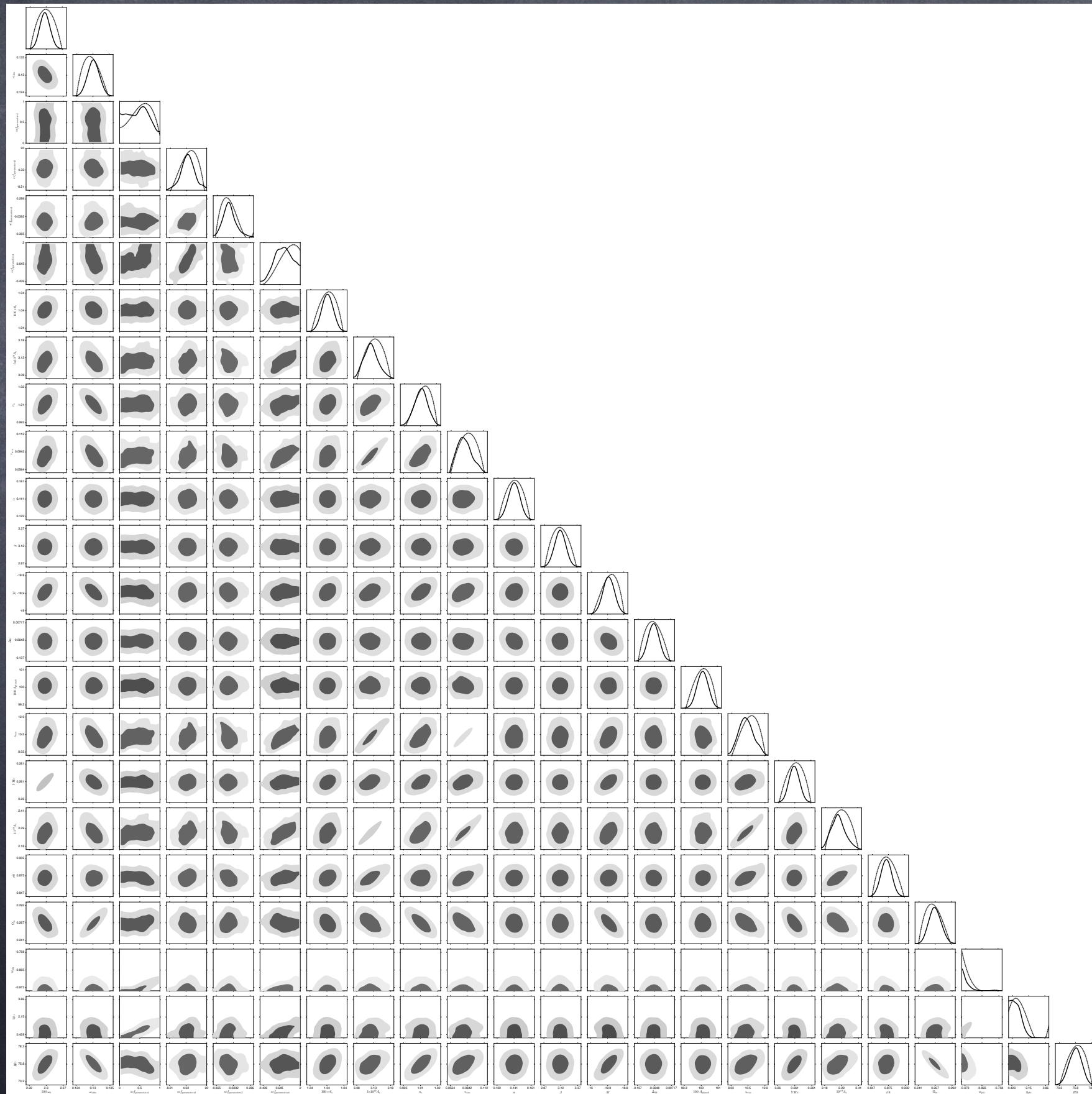


6 cosmological parameters  
+  
many more nuisance  
parameters





# EJ. CMB DATA and Cosmological model





# Ej. CMB DATA and Cosmological model

