

AI data bases and material discovery

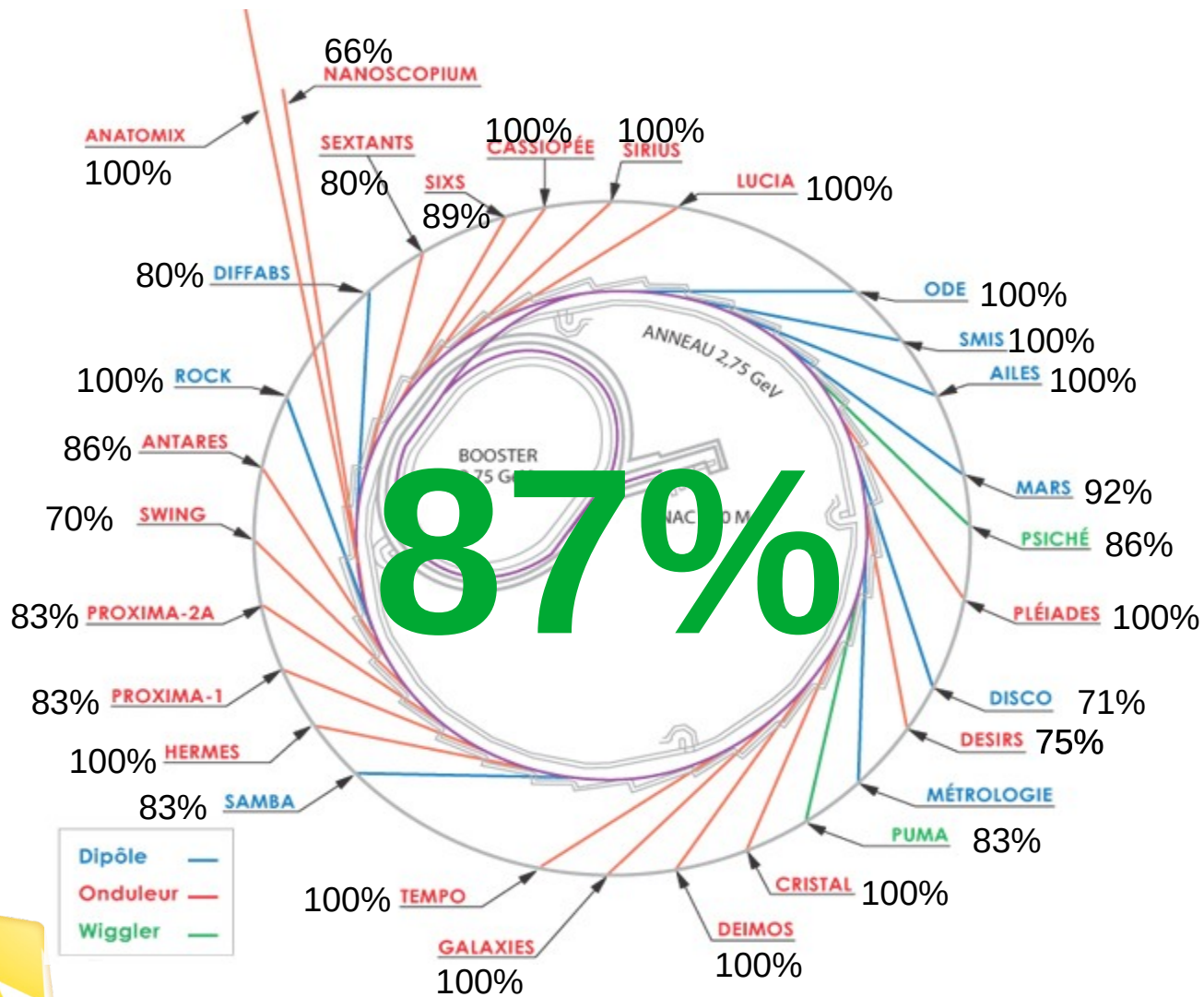
E. Farhi, *Data Reduction and Analysis Group - SOLEIL*

DIADEM School - Aug 2025

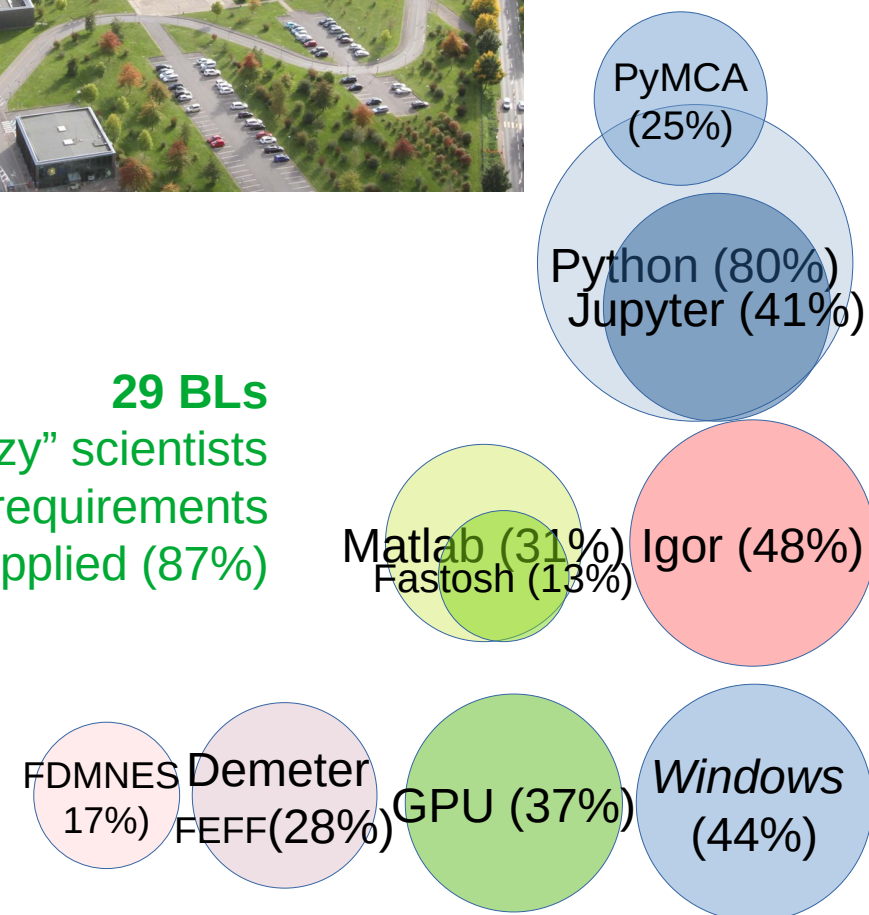


- **Groupe de Réduction et d'Analyse de Données Expérimentales de Soleil (GRADES).**
- Created in Feb 2020.
- **Staff:** 6 perm. (3 dev, 2 BL, 1 AI) + ~4 temp.
 - E. Farhi, F. Picca; O. Roudenko, S. Bac; A. Bellachehab, G. Cassam
- **Topics:** scientific software (diff, spectro/abs, simulation, ...), Debian/Ubuntu packaging, deployment, coding, documentation, support, ... and now AI everywhere.
- **Means:** 3 central servers + ~20 BL servers.





29 BLs
 ~100 "crazy" scientists
 160 requirements
 144 supplied (87%)



Our BLs require essentially :

- **Quasar/Opus PyHST2 UFO ImageJ Fiji Napari Nabu Astra-Toolbox *matlab* silx navarp arpys binoculars pynx labplot jupyterhub jupyterlab codeblock vscode pyfai pymca maud athena/artemis corali-q4 cryspy/quanty fdmnes quantum-espresso/sssp abinit cp2k gromacs lammps fastosh xraylarch sasview cctbx dioptas bornagain genx anarod xrayutilities ase foxtrot atsas bioxtas-raw matlab-csaxs-ptycho**
- *Igor RefFIT casaxps* (Windows)

We develop specific software:

- **binoculars/hkl, McXtrace, pyfai/pynx, qemu-web-desktop**, notebooks, BL data I/O, ...



- We have chosen **Debian-class** systems to ensure **coherence**, **stability** and **maintainability** of most of our scientific software.
 - The Debian Science stack is **extensive**.
 - Brings better life-span than **conda** and **containers**.
 - We use 'unstable' packages on request.
- Two Debian contributors in the group + 40k€/y subcontracting.
 - ~**400** packages pushed in Debian Science over 4 years.

<https://salsa.debian.org/pan-team/soleil-packaging-overview>



debian



F. Picca
E. Farhi
R. Mas
X. Guimard



Load raw data (handle formats).

Normalize/correct intensity:

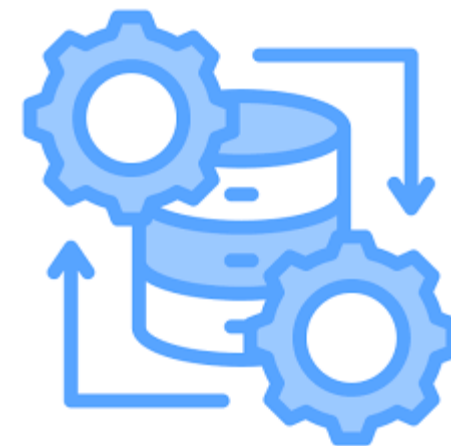
$$\frac{\text{Light} - \text{Dark} - \text{Offset}}{\text{Flat} - \text{Dark} - \text{Offset}}$$

Align (image registration) and **Stack**.

Transform coordinates. Add metadata.

We mostly use Python libraries:

- `pyfai`, `pymca`, `silx`, `h5py`, `libxrl`, `xraydb`,
`larch`, `pyhst/nabu`, `skimage`, `sklearn`,
`orange3/quasar`, `pynx`, ...



Load reduced data.

Make use of **metadata**.

Use **large data sets**, #CPU and #GPU.

Usually perform **statistics**.

Fit theory (extract physical parameters).

Estimate **errors**.

Write scientific **papers**.

- Dioplas, Mantis-Xray, Larch, Arpys/Navarp, Orange3/Quasar, BioXtas-RAW, SASView, BornAgain, GenX, Napari, Fiji, ...



We provide all inclusive
and Visualization services:



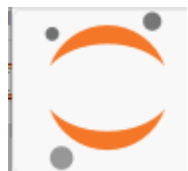
Data Reduction, Analysis

Thousands of software and libraries.



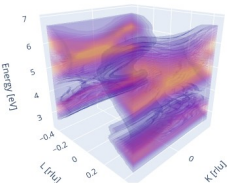
DARTS (easy install/deploy).

<https://data-analysis.synchrotron-soleil.fr/desktop>



JupyterHub (shared storage with *DARTS*).

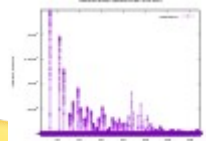
<https://data-analysis.synchrotron-soleil.fr/notebook/>



Material modelling (easy DFT+ e^-/ph dispersions)



https://data-analysis.synchrotron-soleil.fr/computing/material_modeling.html

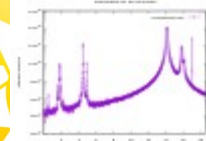


Diffraction + Fluorescence estimates

McXtrace



https://data-analysis.synchrotron-soleil.fr/computing/fluor_diffraction.html



Service: <https://data-analysis.synchrotron-soleil.fr/>

Code: <https://gitlab.com/soleil-data-treatment/soleil-software-projects/qemu-web-desktop>



Data Analysis Remote Treatment Service (DARTS)



This service is a data analysis portal that allows to create a remote desktop to treat your data, in the cloud. You can tune the type of system you need. It will be displayed in your browser, without any additional software for you to install.

Read more on our [HELP](#) page.

NOTE: From SOLEIL Network, please use Firefox with "auto-detect" proxy (Top-right menu, Preferences, search for "proxy" (top right), select Network Settings, choose "auto-detect"). It is also important to inactivate any JavaScript blocker plugin.

~350 users/month

1

User ID: farhie

Password:

Create

By pressing the Create button, you agree with our Terms and Conditions (*)

Machine: Data Analysis (Debian 11, stable, "U...")

Configuration script (opt.): Path or URL to a single file or 'exec: cmd1; cmd2; ...'

Number of CPU's: 4

Amount of memory: 4 GB

Compute on GPU (opt.): ☐

You may request a physical GPU to e.g. run heavy computations (not for display). The tools and libraries you wish to use should have been designed to benefit from such devices with e.g. OpenCL, CUDA, OpenACC, ROCm/HIP. Do NOT request a GPU if you do not actually use it, as their number is limited.

Currently Includes:

- CP2K, GROMACS, LAMMPS, QE/SSSP, ABINIT, ASE, Jupyter NB, FEFF/FDMNES, ORCA, FPLO, PyTorch, Sk-Learn/image...

Data Analysis Remote Treatment Service: Debian-unstable-2022-10-19.qcow2



- [OK] Using multiple login session (reconnect/share).
- [OK] Creating snapshot from Debian-unstable-2022-10-19.qcow2 as session qemu-web-desktop_QxQGeukn
- [OK] Connect to your machine at <https://re-grades-01.exp.synchrotron-soleil.fr:6001/exp.html?size=scale&autoconnect=true&path=/token-rc34hY1X>.
- [OK] You can use your machine until Tue Oct 23 21:45:49 2022.

Hello farhie !

Your machine qemu-web-desktop Debian-unstable-2022-10-19.qcow2 has just started. Click on the following link.



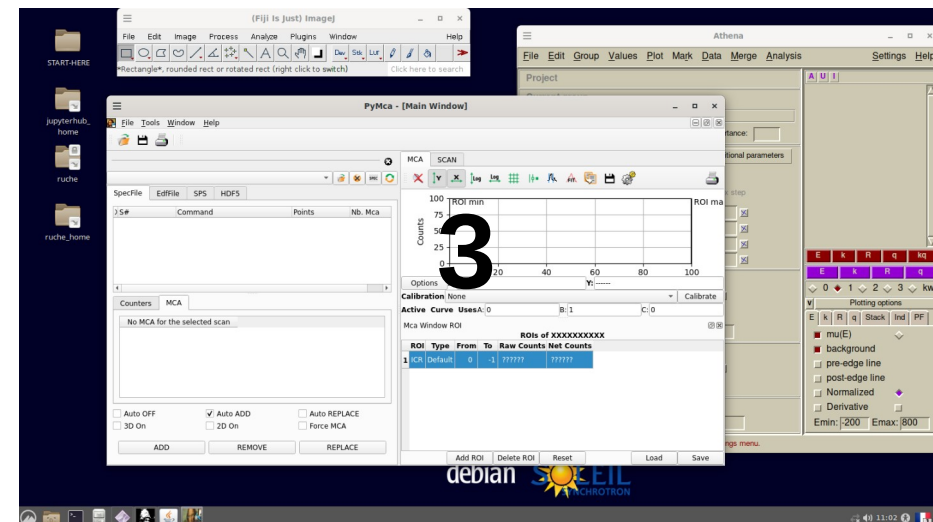
2

CONNECT ☒ SHARE ☐

NOTES:

- You can close the browser and reconnect any time (within life-time) with the link above.
- Select the [Manage sessions] item in the service login page to list, reconnect or abort your sessions.
- You can collaborate in the same session with your colleagues. Just send them the link above.
- Please shut-down the machine properly (do not just logout or suspend).
- The virtual machine is created on request, and not kept. Your work must be saved elsewhere (e.g. mounted disk, ssh/ftp, Dropbox, OwnCloud...).
- To kill this session, click on

STOP ☒



QR code.

Access to GPUs.

Open-source/Debian.

100+ of sci software in VM.

Access to a persistent area.

Access to the Ruche (exp data).

Will evolve into the VISA service.

DARTS is simple and efficient.

Fully configurable.

Low Tech / Maintenance.

Deployment in minutes.

Scalable.

Material modelling (easy DFT+ e^-/ph disp.)

- **Service:** https://data-analysis.synchrotron-soleil.fr/computing/material_modeling.html
- **Code:** <https://gitlab.com/soleil-data-treatment/soleil-software-projects/material-modelling-service>

Input
CIF

Material Modeling Service

This service allows to compute some electronic and phononic properties of materials from first principle theory (DFT *ab-initio*). The only requirements are your credentials, and a CIF/PDB/POSCAR etc (see the [ASE documentation](#)). In addition, you may choose the calculator to use (GPAW, QE), the computational setps to perform.

You are invited to search for your material in the following resources:

- [Crystallography Open Database \(COD\)](#)
- [MaterialsProject](#)
- [NoMaD repository](#)
- [PhononDB](#)

or simply use [AI](#) for a test.

The results will be stored into our data storage at <https://data-analysis.synchrotron-soleil.fr/data/modeling>. By submitting your data here, you implicitly agree to share it with other colleagues at SOLEIL.

Your User ID *:
Your password *:
Calculator:
Smearing:
Compute:
Material structure *: Aucun fichier sélectionné.

Compute !

Ground state on Wed May 28 16:16:55 2025

Generated:

- [\[properties.h5\]](#)
- Energies are e.g. in [eV].
- Refer to the [ASE](#) documentation to learn about signification and units.

Properties:

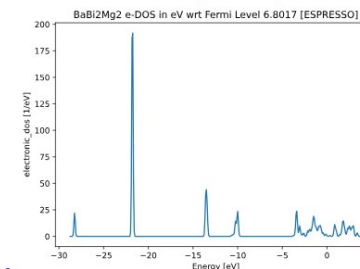
```
{ 'angular_momentum': array([0., 0., 0.]), 'center_of_mass': array([0.94611635, 1.6387216, 3.21874433]),  
'charges': None, 'dipole_moment': None, 'Fermi_level': 6.8017, 'kinetic_energy': 0.0, 'magnetic_moment': None,  
'momenta': array([[0., 0., 0.], [0., 0., 0.], [0., 0., 0.], [0., 0., 0.], [0., 0., 0.]]), 'moments_of_inertia':  
array([ 876.33055125, 4242.82588009, 4984.41961667]), 'potential_energy': -11893.656298752589, 'stress':  
array([-0.00443654, -0.00443654, 0.00146079, -0., -0., 0.]), 'total_energy': -11893.656298752589 }
```

Electronic density of states on Wed May 28 16:16:55 2025

Generated:

Full report and data files

- BaBi2Mg2 e-DOS in eV wrt Fermi Level 6.8017 [ESPRESSO]



- [\[electronic_dos.svg\]](#)
- Tabulated density of states data is available in [\[properties.h5/electronic_dos\]](#).

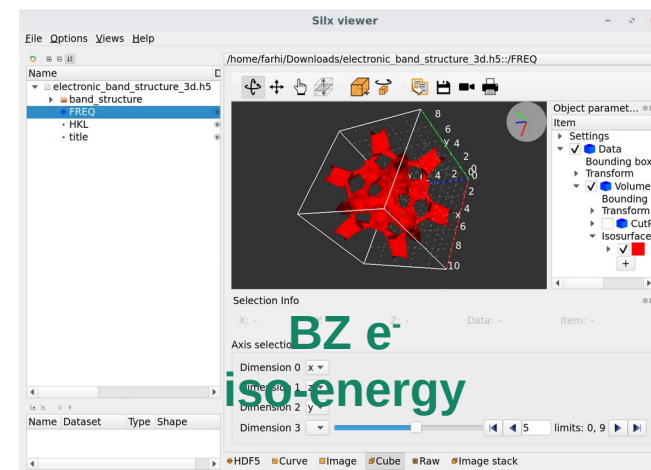
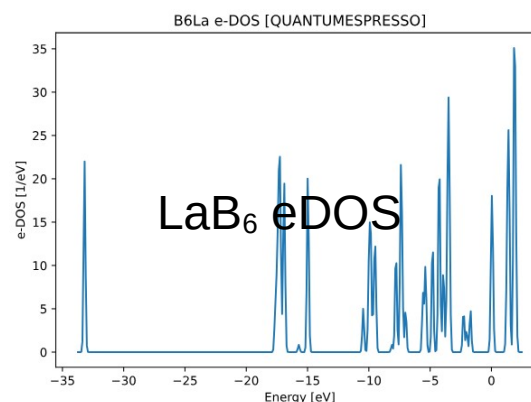
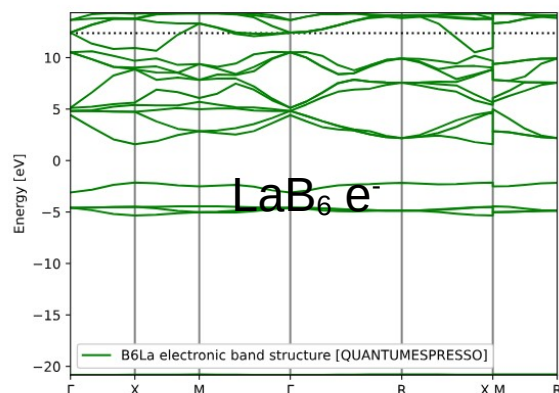
https://data-analysis.synchrotron-soleil.fr/data/material_modeling/BaMgBi2.cif-QuantumEspresso-bertran-20250528-161646/data/properties.h5

Material modelling (easy DFT+ e^-/ph disp.)

- We compute the electronic levels in the BZ using e.g. *QE/SSSP* via *ASE*.
- Configuration is automatic for fast estimates (based on *Materials Project* settings):

Calc	Conv. dE	Displ. δx	E_{cut}	Kpts	supercell
QE/SSSP	1e-5 eV	0.01 Angs	$200 \cdot N_{\text{at}}$ eV	$N_{\text{kpts}} / \sqrt[3]{(N_{\text{at}} \cdot \text{supercell})}$	$\sqrt[6]{(K_{\text{grid}}^3 / N_{\text{at}})}$

- From this we get electronic dispersions:



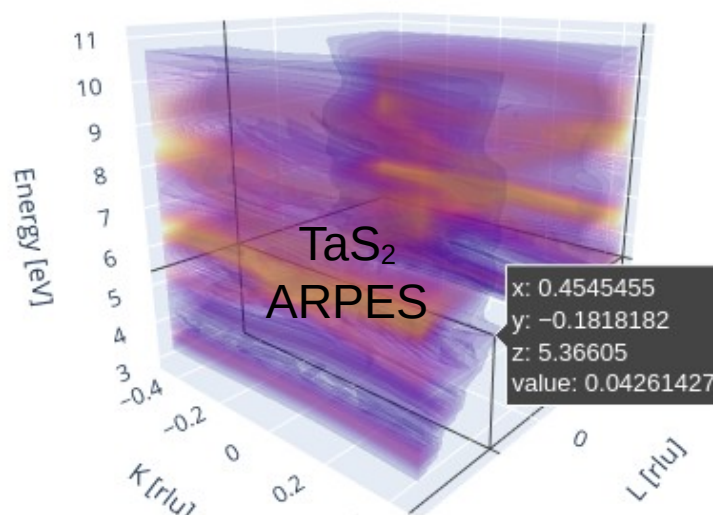
Electronic states → ARPES intensity

- From the Fermi golden rule via the electro-magnetic field \mathbf{A} and light momentum \mathbf{p}

$$\Gamma_{i \rightarrow f} \propto | \langle f | \mathbf{A} \cdot \mathbf{p} | i \rangle |^2 \delta(E_f - E - h\nu)$$

- Field \mathbf{A} boils down to photon polarization, and \mathbf{p} to the in-plane momentum (surface).
- We compute an estimate for the intensity on a flat sample surface

$$I(q, E) \propto \sum_{bands} |\vec{\epsilon} \cdot \vec{q}_{||}|^2 \frac{1}{1 + e^{(E - E_F)/kT}} \frac{1}{1 + (E - E_f)^2/\Gamma^2}$$



Small displacements → vibrational states

- From the Fermi golden rule via the interatomic potential V and particle density ρ .

$$w_{\vec{k}_i \rightarrow \vec{k}_f, \lambda_i \rightarrow \lambda_f} = \frac{2\pi}{\hbar} |\langle \lambda, \vec{k}_f | V | \lambda_i, \vec{k}_i \rangle|^2 \rho(E_f),$$

Use *PhonoPy*

- We compute the form factors (harmonic approx – small displacements → forces and potential V)

$$F_j(Q) = \sum_a \frac{b_a}{\sqrt{m_a}} e^{-W_j(Q)} \vec{Q} \cdot \vec{e}_j e^{-i\vec{Q} \cdot \vec{r}}$$

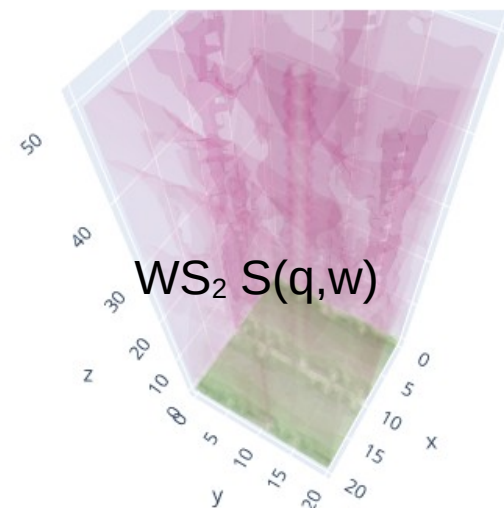
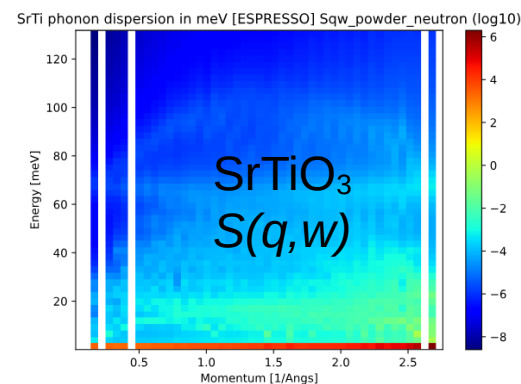
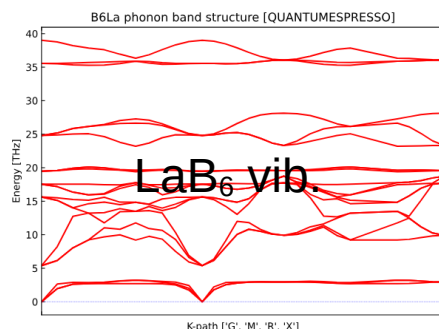
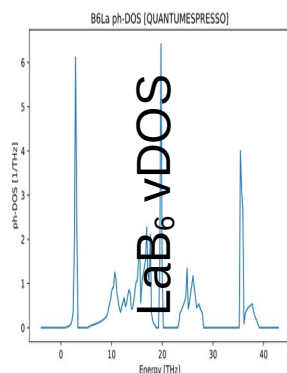
b = scatt. length [m] = $r_e f$ for photons

$$W_j(Q) = \frac{|\vec{Q} \cdot \vec{e}_j|^2}{2\omega_j(Q)m_a} (2n(\omega_j(Q)) + 1)$$

$$n(\omega) = 1/(e^{\hbar\omega/k_B T} - 1)$$

- And then the X-ray and neutron vibrational dynamic structure factors.

$$S(\vec{Q}, \omega) = \sum_j (n(\omega) + 1) \frac{|F_j(Q)|^2}{\omega_j} DHO(\omega, \omega_j, \Gamma)$$



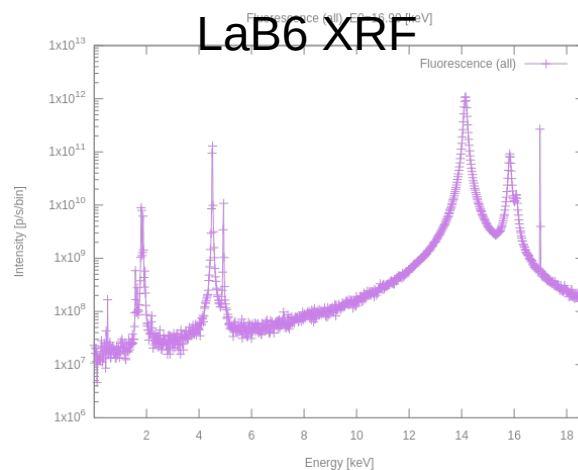
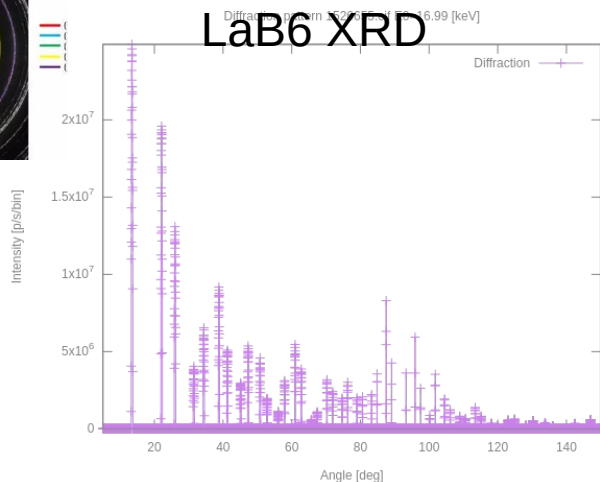
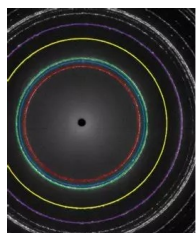
Diffraction + Fluorescence estimates

- https://data-analysis.synchrotron-soleil.fr/computing/fluor_diffraction.html
- <https://gitlab.com/soleil-data-treatment/soleil-software-projects/fluor-diffraction-service>

We rely on XrayLib (librxi) and McXtrace to compute the powder diffractogram and the fluorescence spectra. Also compute the experimental “background”. Produce **realistic** data sets.

$$\sigma_{\text{cone},Q} = j_Q f \exp(-2W) \frac{V}{V_0^2} \frac{\lambda^3}{4 \sin \theta} |F(Q)|^2$$

$$I_{\text{el}} = I_0 \frac{\Omega}{4\pi} \frac{\sigma_{\text{XRF}}(E_0)}{\sigma_{\text{TOT}}(E_0)} \frac{1}{\sin \psi_1} \frac{C_{\text{el}} \alpha_{\text{el}}(E_0)}{\frac{\alpha(E_0)}{\sin \psi_1} + \frac{\alpha(E_1)}{\sin \psi_2}}$$



McXtrace



<https://mcxtrace.org/>

E. Knudsen *et al*, JAC **46**(3):679-696, 2013
A. Brunetti *et al*., Spectrochimica Acta Part **B 59** (2004) 1725–1731
T. Schoonjans *et al*., Spectrochimica Acta Part **B 66** (2011) 776–784

Before starting any new study from scratch, you should consider to inquire for existing structures and vibrational information from **databases** at:

- <https://alexandria.icams.rub.de/>
- <https://optimadeclient.materialscloud.io/> (catalog of catalogs)
- <https://www.materialscloud.org/home> (QE team)
- <https://nomad-lab.eu/nomad-lab/> (simulated structures)
- <https://next-gen.materialsproject.org/> (simulated structures)
- <https://www.crystallography.net/cod/> (sim and exp)
- <https://github.com/atztogo/phonondb> (phonopy db)





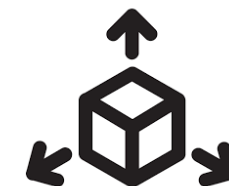
The advent of A.I.
The end of equations and theory.



- A.I. algorithms mostly rely on “**training**”, and store vast amounts of data.
- The data must be properly “**annotated**”.
- The A.I. connects/correlates “inputs” with “outputs”. It “mimics” the training data: a **parrot**.
- The logical path that produces the A.I. answer (aka *inference*) is not always explicit.

It behaves as an oracle (*no critics here*): a **black box**.

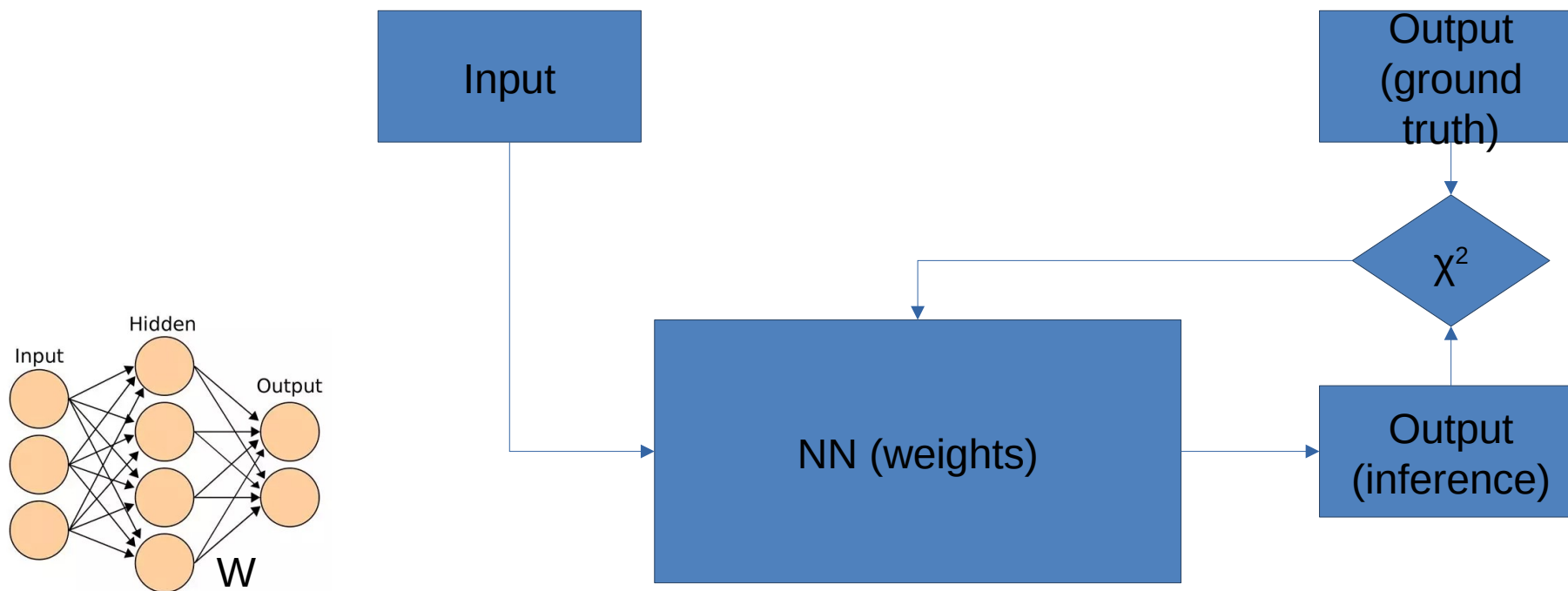
- The A.I. is **restricted** to the training data range.
- The “**accuracy**” is **hard to estimate**, so that A.I. does NOT follow the science path.



We have an **input**, an **output**, and some **internal 'weights'**. We optimize the NN internal weights (the 'memory') so that its output converges to the expected one.

This is a **pure optimization algorithm**.

The large number of parameters allows to approximate any function.



Looks like an Ising interconnected spin-glass

Transformers rely on pure **linear algebra**.

We first 'embed' the **input** as *tokens* x (via quantization and PCA).
The **output** of the Transformer has the same size as the input x .
The knowledge is stored into some Queries, Keys and Values **weights**.
Q.K.V vectors are determined from embeddings and **internal weights**

$$QKV = x.W_{QKV}.$$

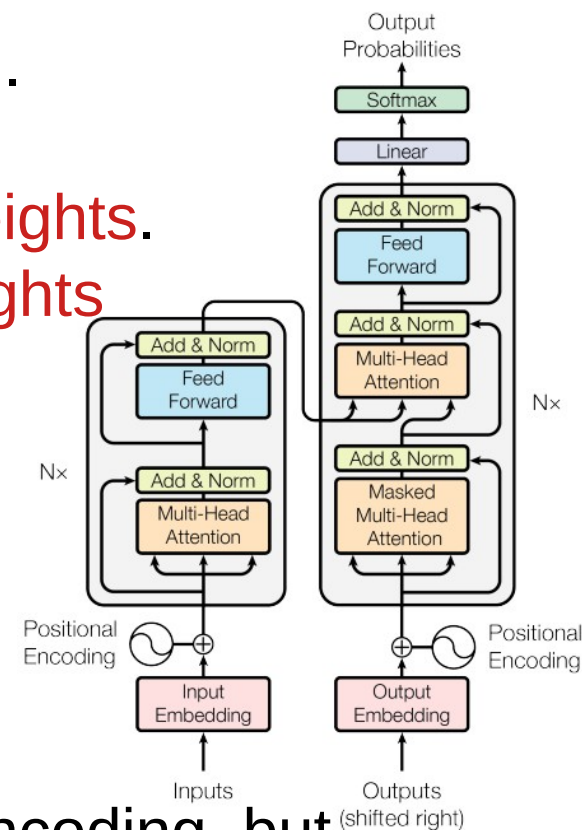
The weights are set during training.

The output is a dot product:

$$y = \text{softmax}(Q.K^T/\sqrt{d}).V \quad [\text{softmax} = \text{logistic} = \text{Boltzmann}]$$

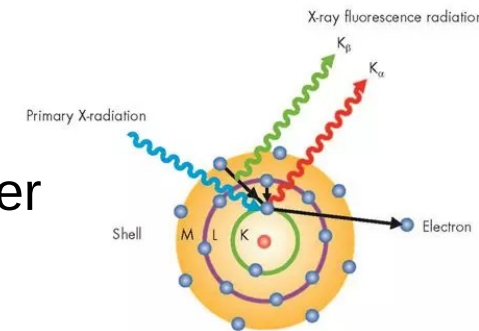
Transformers can NOT be exact as they rely on reduced latent encoding, but they are (nearly) **deterministic**.

See: <https://xmarva.github.io/blog/2025/building-a-transformer/>



All beam-lines are equipped with **fluorescence** detectors

- Photon ejects an electron in an absorption process
- The hole is filled by an electron that “falls down” while emitting a photon at lower energy
- Generates a 1D energy spectra (e.g. 1-40 keV)

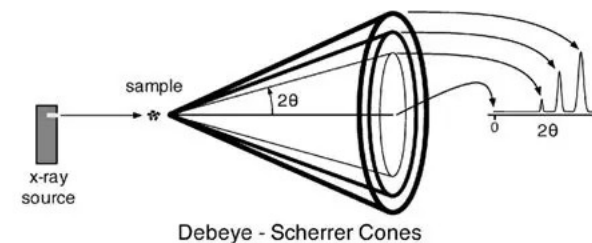


Many beam-lines are equipped with **diffraction** 2D detectors

- The crystalline periodic structure diffracts with Bragg law (coherent elastic scattering)
- Creates a 2D set of “rings” or “spots” vs angle

Other possible techniques

- **absorption edges**: specific to atoms oxidation and coordination
- **small angle scattering**: specific to “large” structures (micelles, colloids, polymers, proteins, ...)
- Multipurpose **chat**



- A.I. can be used for **automatic interpretation/reduction**.
- **We have started:**
 - Language models → Help scientists
 - Fluorescence/XRF → Elemental analysis
 - Diffractogram/XRD → Space-group
 - SAXS → Large-structure model (SASView)
 - XAS → Oxidation/coordination



We aim to provide cloud services for such X-ray applications.



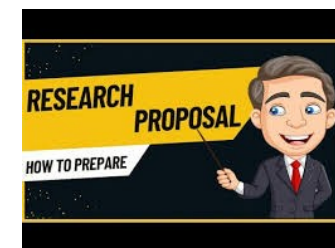
- **Warning:** keep property of your data
avoid cloud services
- **Warning:** how do you estimate if answer is valid ?
trust your own expertise – just as you would do on Internet
- **What for:** help at 3pm, during acquisition or when preparing proposals, broad tech advice.
- **We develop local LLMs**

Mistral/DeepSeek in production as a local service : we use **llama.cpp** as inference engine and web UI

Runs on two servers, with 'thought' process.

Mistral AI connected to our **Confluence** knowledge base (100 % *original work*)

Context tuning for a RAG service (upload doc and/or search local doc)



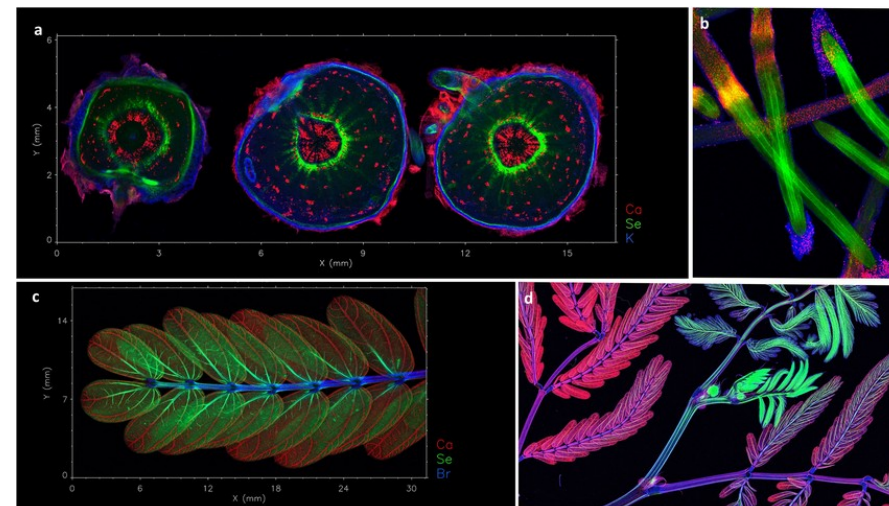
Started end 2023
J. Hassani
A. Bellachehab

Expected service

- **IN:** raw fluorescence data (binned on 0-45 keV range)
- **OUT:** atoms and stoichiometry (Z=5-90). Elemental analysis
- (opt) Estimate for incident energy
- (opt) Estimate for detector artifacts

Usable as script, and web API

- Fast estimates for any 1D, 2D, 3D data with fluorescence spectra
- Fast guess for e.g. *pymca* and *xray-larch*.
- Results stored as HDF5/NeXus metadata
- Should be ready end of 2025



Started March 2025

F. La Porta

S. Bac

A. Bellachehab

- We have started to train an A.I. to analyze spectroscopy data.

⚠ There are no massive annotated data sets.

Don't panic, this is usual with A.I.

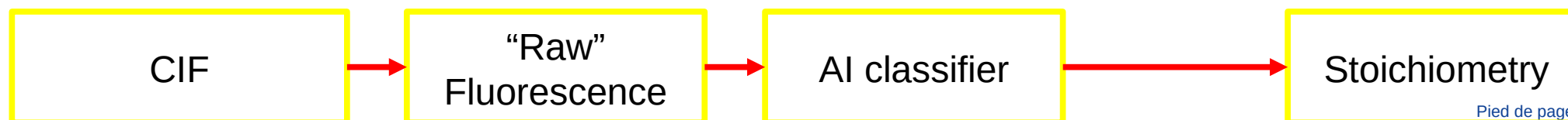
- We use 600k compounds and associate fluorescence spectra.

Assemble a **McXtrace** model of a fluorescence X-ray imaging beam-line.

Contains photon source, optics, sample (*Fluorescence*) and detectors.

Include multiple-scattering, Compton, Rayleigh, fluorescence cascade, absorption secondary extinction, sample geometry, detector artifacts, Monte-Carlo noise as in experiments, ...

Relies on **xraylib** (ESRF): Boron to Thorium (Z=5-90)



McXtrace is a general Monte Carlo ray-tracing software for simulation X-ray beamlines and experiments.

<https://mcxtrace.org/>

McXtrace



Forked from **McStas** for neutron scattering

www.mcstas.org

McStas



Provides

209 components (sources, optics, samples, monitors, etc)

98 examples

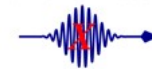
The strength and specificity of McXtrace and McStas is the ability to model sample scattering within instruments/beam-lines.

Project funded by EU since 1996. Mostly C-based, with python interfaces.

<https://github.com/mccode-dev/McCode>

Runs on all architectures (conda, debian-class, ...). Massively parallel: MPI, GPU





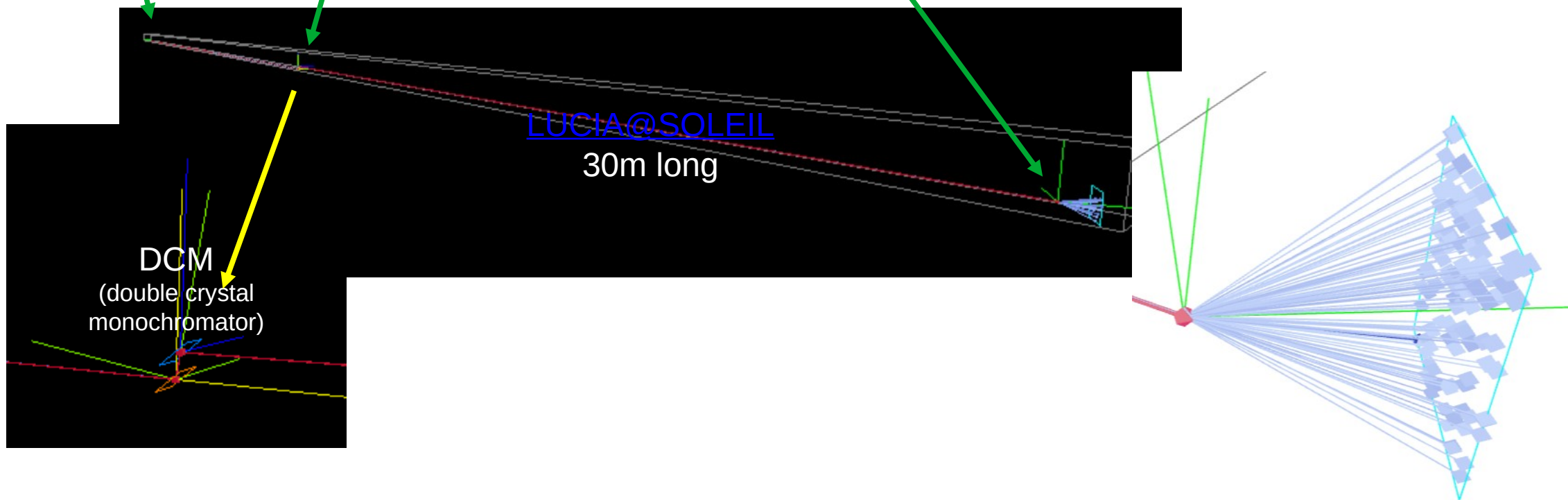
COMPONENT source = **Undulator**(E0 = E0, dE = dE, ...)

COMPONENT dcm_xtal0 = **Bragg_crystal**(length=0.04, width=0.04, h=1, k=1, l=1, material="Si.txt", V=160.1826)

COMPONENT dcm_xtal1 = **COPY**(dcm_xtal0)

COMPONENT sample = **Fluorescence**(material=sample_material...)

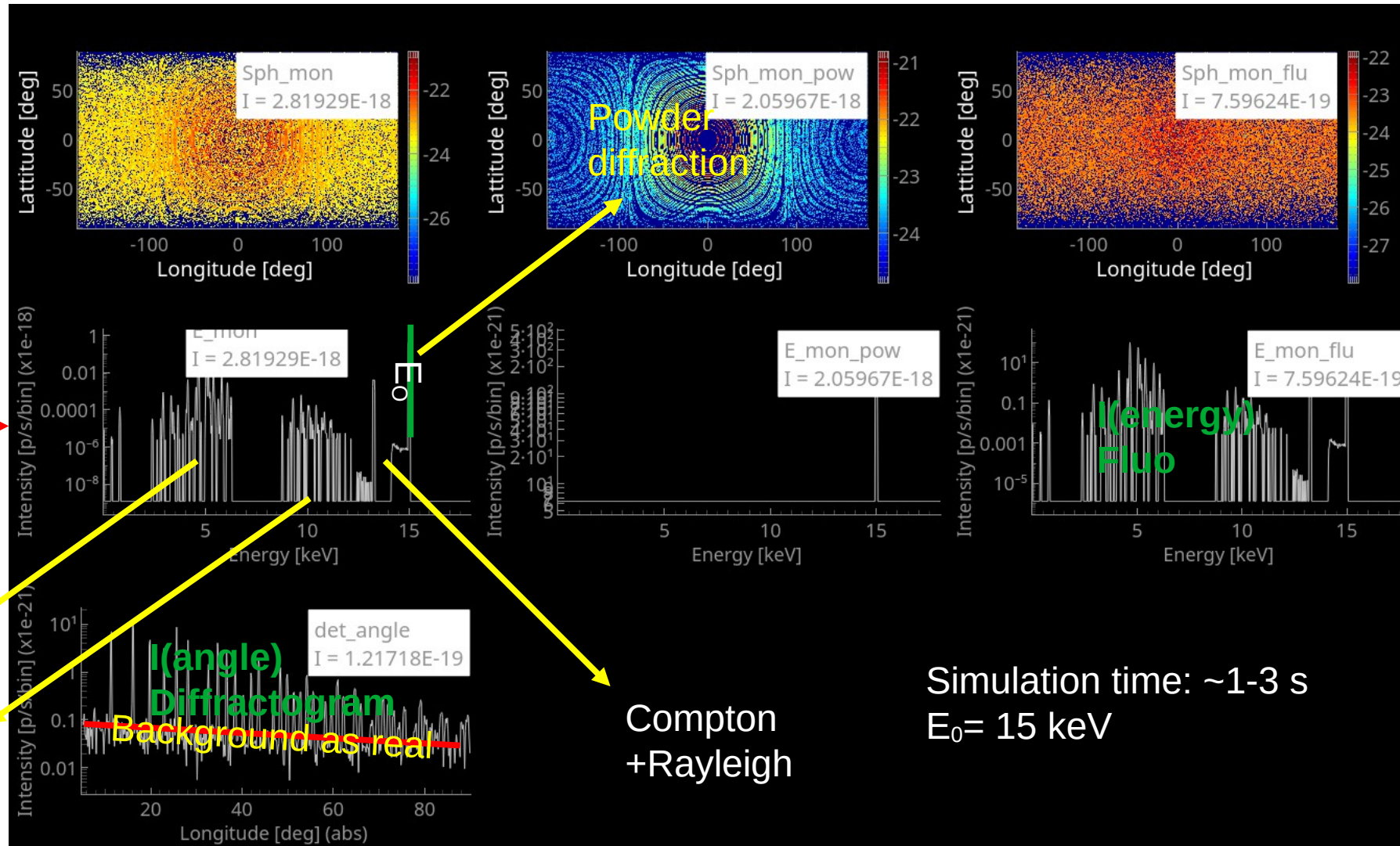
COMPONENT mon_spl_fluo = **Monitor_nD**(options="energy", ...)



Using FluoPowder
McXtrace sample

Fluorescence is not
really isotropic –
appears as a
background in angle

CIF (structure)



Spectra simulation

Simulation model

- Gaussian x-ray sources
- Sample

Simulation

$E_0 = 60$ keV

Range 0-66 keV

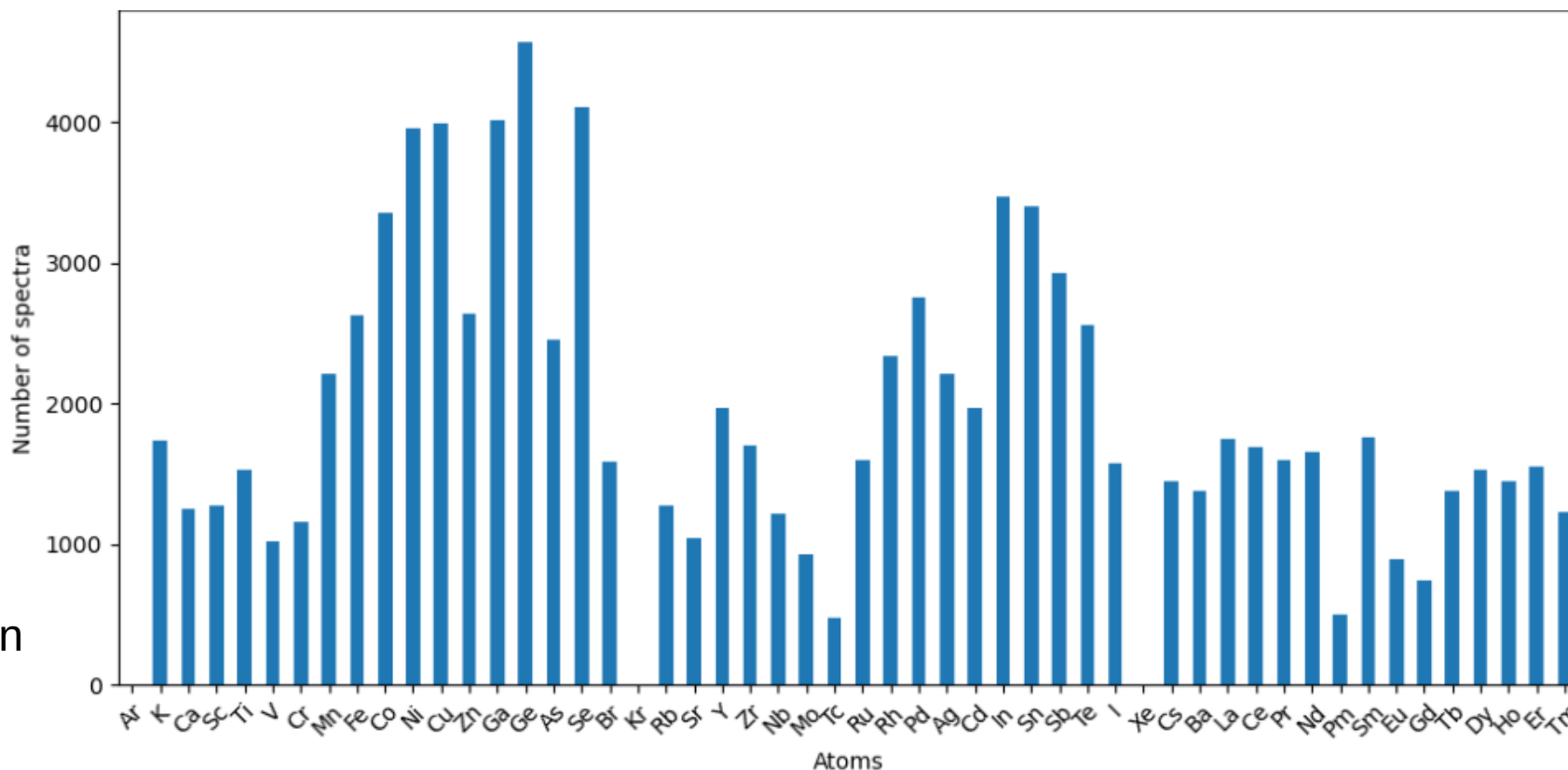
$n=10^5$ rays

Spectra with 2000 points

Fluorescence line set as Gaussian

Formula

- Real formula containing at least 1 element with $K\alpha$ between $3\text{keV} < E_{K\alpha} < 60\text{keV}$ (52 elements)
- Formula from MP, formula containing only that 52 elements
- 34136 spectra



Positional encoding for LLM

$$PE(pos, 2i) = \sin\left(\frac{pos}{10000^{\frac{2i}{d}}}\right); PE(pos, 2i+1) = \cos\left(\frac{pos}{10000^{\frac{2i}{d}}}\right)$$

i : dimension index
 d : embedding size

Custom made positional encoding for spectroscopy

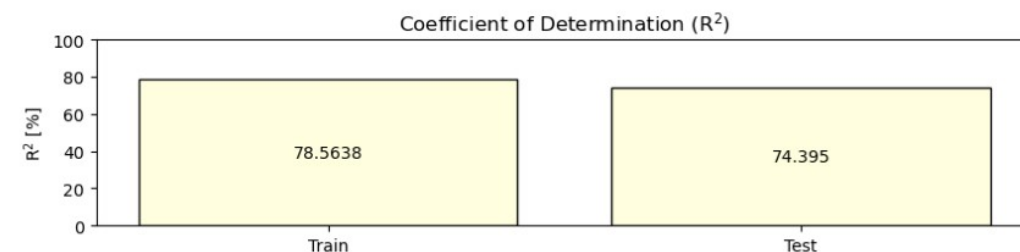
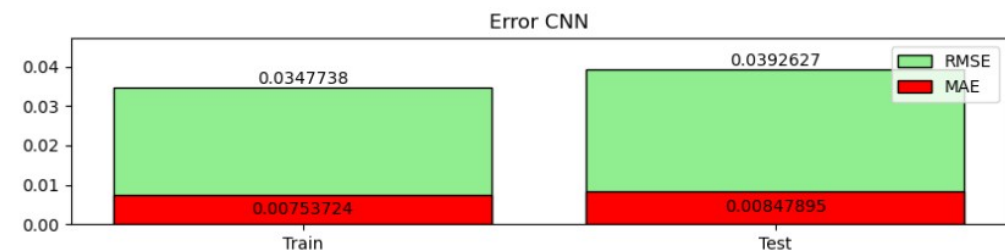
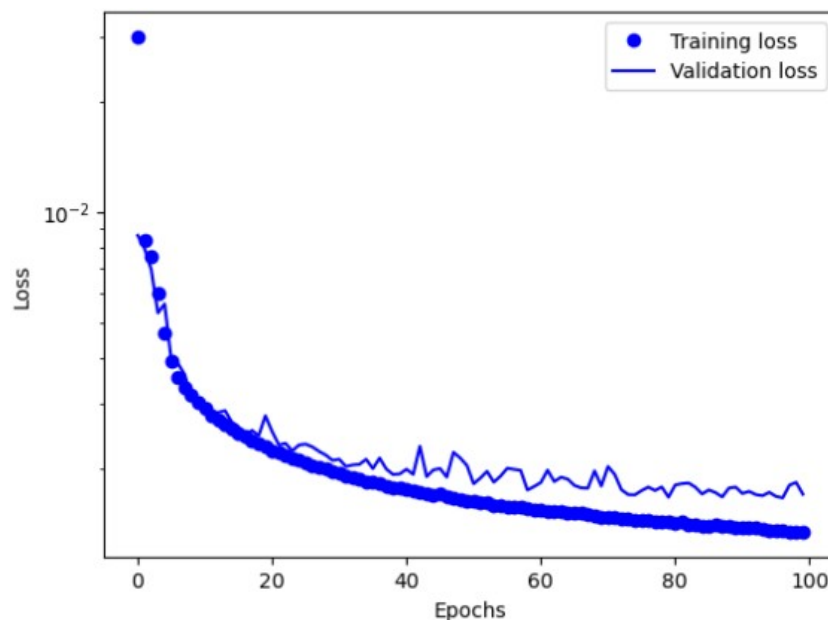
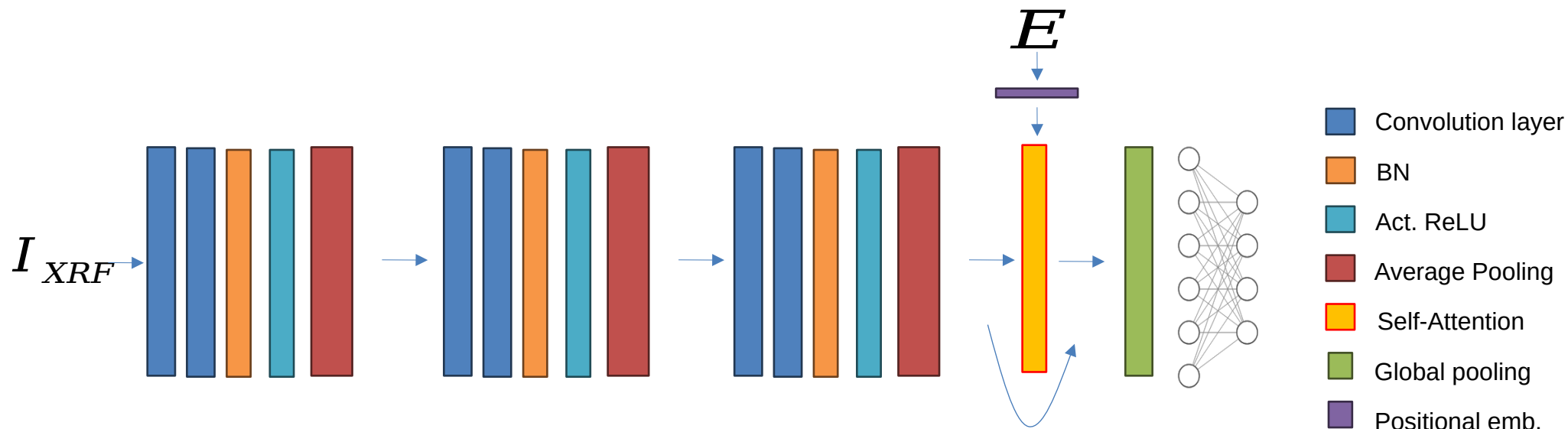
$$PE(E, 2i) = \sin\left(\frac{E}{\tau^{\frac{2i}{d}}}\right); PE(E, 2i+1) = \cos\left(\frac{E}{\tau^{\frac{2i}{d}}}\right)$$

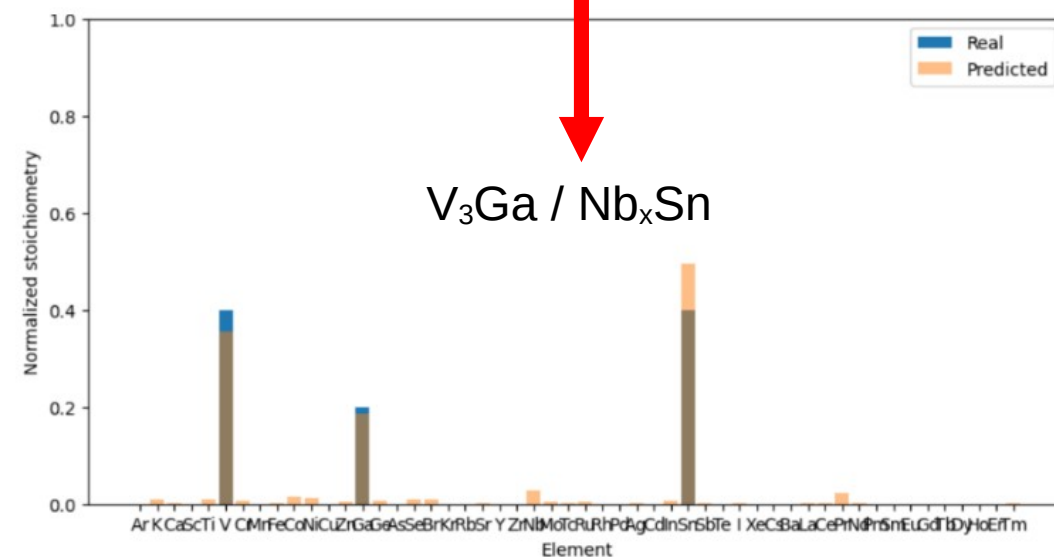
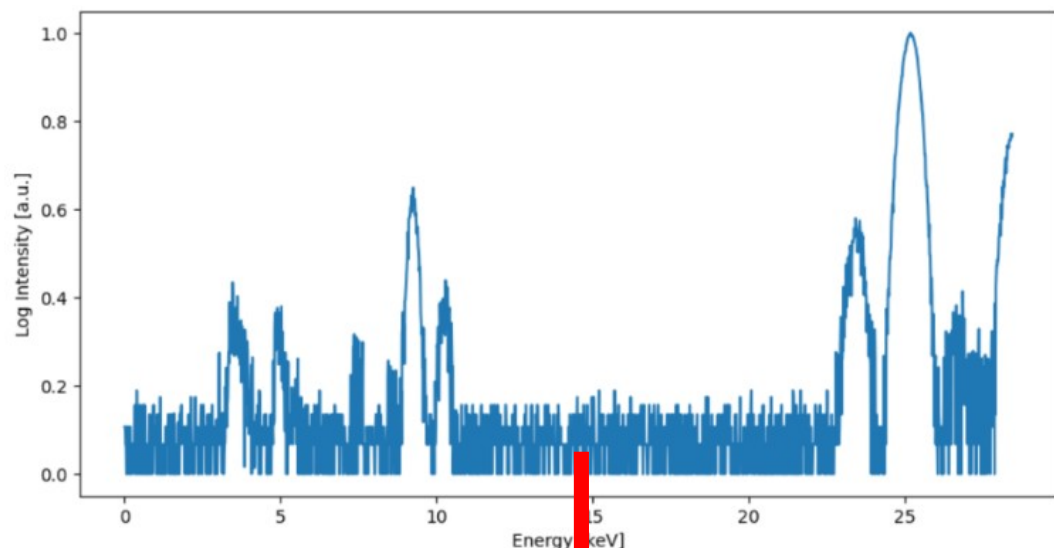
To get $\tau \rightarrow$ PE must deliver unique values until 66 keV : $\sin(E_{\max}/\tau) = \sin(2\pi) \rightarrow \tau = 66/2\pi \sim 10$

$$PE(E, 2i) = \sin\left(\frac{E}{10^{\frac{2i}{d}}}\right); PE(E, 2i+1) = \cos\left(\frac{E}{10^{\frac{2i}{d}}}\right)$$

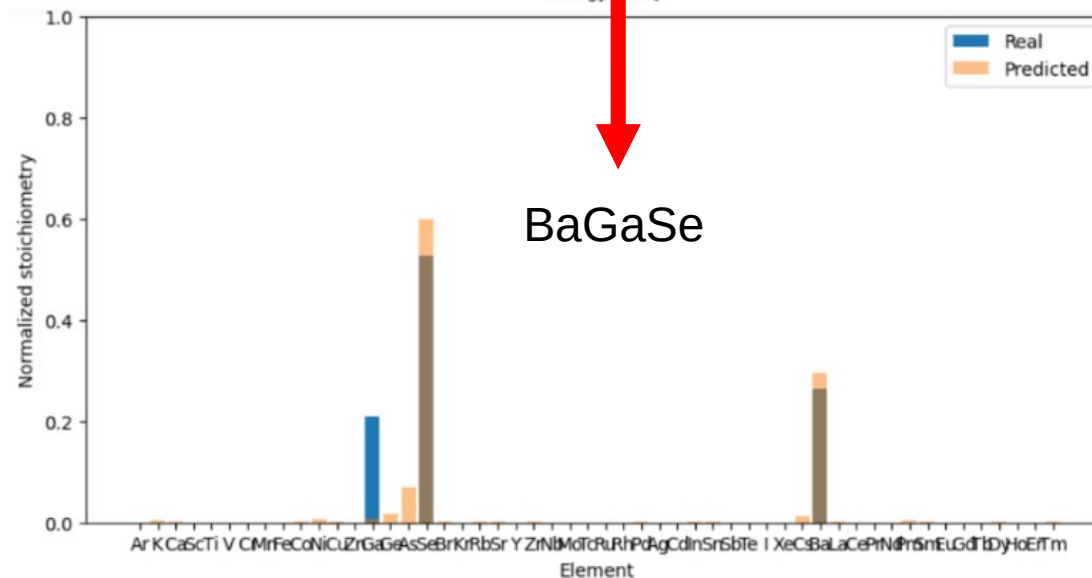
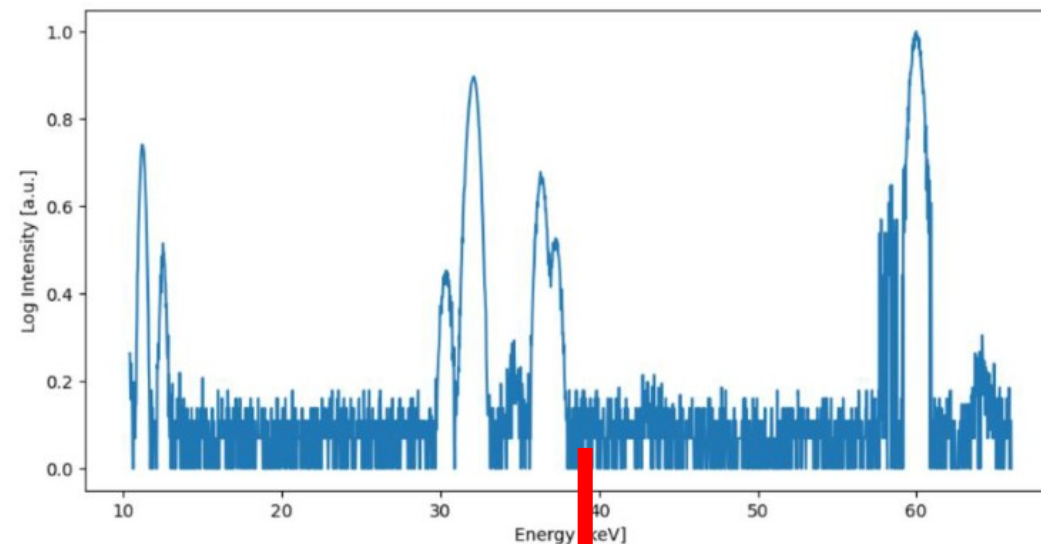
The positional encoder is summed to the embedding (feature maps of the conv)

- **Convolution layer 1**
 - Kernel size = 4
 - Filters: 16
 - Batch norm (BN)
 - Pool 2
- **Convolution layer 2**
 - Kernel size = 4
 - Filters: 32
 - Batch norm
 - Pool 4
- **Convolution layer 3**
 - Kernel size = 4
 - Filters: 34
 - Batch norm
 - Pool 4
- **Self Attention**
 - Attention head = 10
 - Dimension keys = 64
 - Skip layer
- **2x Fully Conn. layer**
 - 128 neurons
 - Activ. f. = relu
 - Dropout 10%
- **Output layer:**
 - sigmoid activation





<https://next-gen.materialsproject.org/materials/mp-22568?formula=V3Ga>



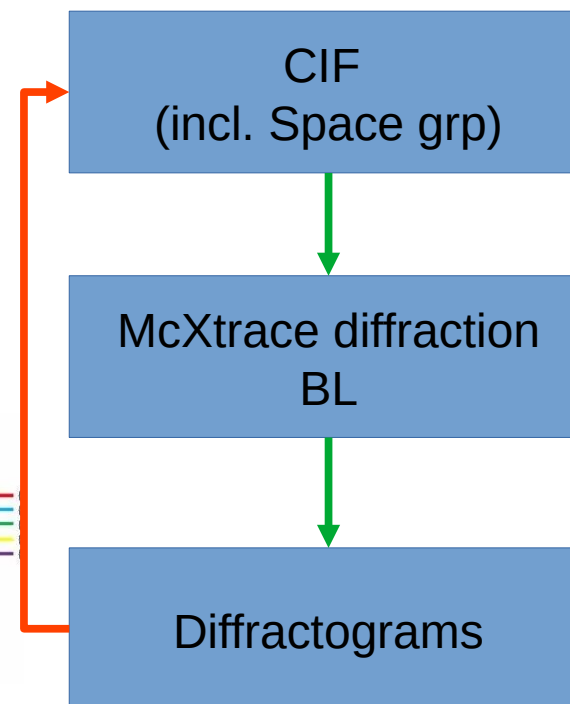
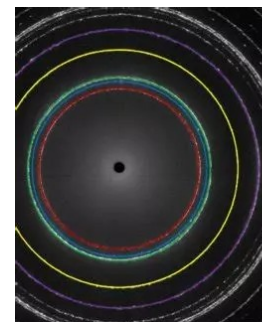
<https://next-gen.materialsproject.org/materials/mp-1191072?formula=BaGa4Se7>

Next step: diffraction

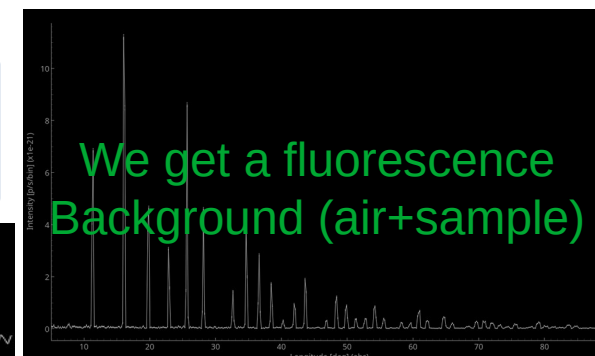
- **IN**: raw diffractogram (binned on 0-180° range), mosaicity, fluo bkg.
- **OUT**: Space-group (1-230) and Bravais class

Usable as script, and web API

- Fast estimates for any 1D, 2D, 3D data with diffractograms
- Fast guess for e.g. FullProf, MAUD, TOPAZ, ...
- Results stored as HDF5/NeXus metadata
- Probably also usable for neutrons
- Should be ready next year (but we already have a prototype)

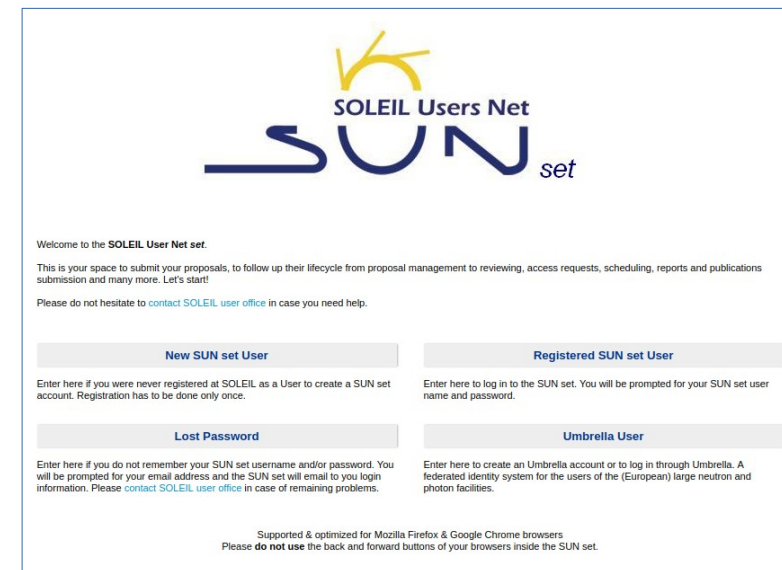


Prototype
A. Bellachehab



How to write a proposal for beam-time at SOLEIL

- 1) Specify the **sample** you wish to study.
- 2) Specify the type of **measurement** you aim for:
 - 📄 Diffraction is first required. If known, indicate structure in proposal.
 - 📄 ARPES and RIXS are well suited for solid-state physics.
- 3) Identify relevant target **beam-line**, and subcommittee panel:
 - Structure and elemental analysis : DIFFABS, CRISTAL, ...
 - Dynamics : CASSIOPEE, ANTARES, GALAXIES, ...
- 4) Contact the BL **scientists** or <useroffice@synchrotron-soleil.fr>.
- 5) Write the scientific case, estimate measurement time (incl. Calibration and reference samples).
- 6) Justify the BL request (« *SOLEIL is the best ; that BL is really suited ; study never done before...* »)
- 7) Refine text in iterations with the BL scientist, well before the dead-line.
- 8) **Submit** at <<https://sun.synchrotron-soleil.fr/sunset/bridge/sunset/>>



SOLEIL Users Net
SUN set

Welcome to the SOLEIL User Net set.

This is your space to submit your proposals, to follow up their lifecycle from proposal management to reviewing, access requests, scheduling, reports and publications submission and many more. Let's start!

Please do not hesitate to [contact SOLEIL user office](#) in case you need help.

New SUN set User	Registered SUN set User
Enter here if you were never registered at SOLEIL as a User to create a SUN set account. Registration has to be done only once.	Enter here to log in to the SUN set. You will be prompted for your SUN set user name and password.
Lost Password	Umbrella User
Enter here if you do not remember your SUN set username and/or password. You will be prompted for your email address and the SUN set will email to you login information. Please contact SOLEIL user office in case of remaining problems.	Enter here to create an Umbrella account or to log in through Umbrella. A federated identity system for the users of the (European) large neutron and photon facilities.

Supported & optimized for Mozilla Firefox & Google Chrome browsers
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Emmanuel FARHI <emmanuel.farhi@synchrotron-soleil.fr>

