

## Calculation of Optical Properties with YAMBO

In this tutorial we are going to learn how to calculate the absorption spectra for 2D MoS<sub>2</sub> using the GW-BSE method implemented in [YAMBO](#). This calculation depends on previously calculated QP energies; therefore, we are going to run this calculation in the same directory as the [GW calculation](#).

In order to generate the input file for the GW-BSE calculation, apply the following command inside the “MoS2.save” directory.

```
$ ./yambo -o b -k sex -b -y h
```

### Options:

**-o b**: the option -o is to select the type of **optics** calculation. Available options are b (optical properties in the eh-space with Bethe Salpeter equation), c (optical properties in G-space and results are similar to DFT level).

**-k sex**: the option -k is to select the type of correlation in the BSE calculation. The standard one is the Screened EXchange (SEX). Other options are Hartree-Fock (HF), Hartree, TD-LDA, and independent particles (IP).

**-b**: the option -b is to ask the code to construct the [BSE kernel](#) which is the second term of the two-particle Hamiltonian matrix elements. The kernel includes the sum of the electron-hole exchange part and the electron-hole attraction. The kernel both shifts and couples the quasiparticle energy differences.

**-y h**: the option -y is to select the [Bethe-Salpeter equation solver](#) to derive the macroscopic dielectric function including excitonic effects. The available options are h (Haydock recursive algorithm), d (diagonalization), and i (inversion solver).

The command above will generate the following input file (yambo.in) for a GW-BSE calculation and will immediately open the following input file in vi editor. The only modification to the input will be the line marked in red. **To include the quasi-particle correction**, copy this line to your input and save your input file.

```
# **      **      **      ****      ****      *****      *****
# /****   **      ****   /**/**   **/**   /**/**/**   **/**/**
# /******   **/****   /**/****   **/**   /**   /**   /**   /**
# /****      **   /****   /**   /****   /**   /****/**   /**   /**
# /**      *****/**   /**   /**   /**   /**   /**   /**   /**
# /**   /**/**/**/**/**   /   /**   /**   /**/**   **
# /**   /**   /**/**   /**   /**   /**   /**   /**/**
# /**   /**   /**/**   /**   /**   /**   /**   /**/**
# /**   /**   /**   /**   /**   /**   /**   /**   /**
```

```

# GPL Version 4.2.1 Revision 110. (Based on r.14778 h.7b4dc3
#
# MPI Build
#
# http://www.yambo-code.org
optics          # [R OPT] Optics
em1s            # [R Xs] Static Inverse Dielectric Matrix
bss            # [R BSS] Bethe Salpeter Equation solver
bse            # [R BSE] Bethe Salpeter Equation.
bsk            # [R BSK] Bethe Salpeter Equation kernel
em1d           # [R Xd] Dynamical Inverse Dielectric Matrix
ppa            # [R Xp] Plasmon Pole Approximation
Chimod= "hartree" # [X] IP/Hartree/ALDA/LRC/BSfxc
BSEmod= "retarded" # [BSE] resonant/retarded/coupling
BSKmod= "SEX"      # [BSE] IP/Hartree/HF/ALDA/SEX
BSSmod= "h"        # [BSS] (h)aydock/(d)iagonalization/(i)version/(t)ddft`
BSENGexx= 4985      RL # [BSK] Exchange components
BSENGBlk= 1005      RL # [BSK] Screened interaction block size
#WebCpl          # [BSK] eh interaction included also in coupling
KfnQPdb= "E < SAVE/ndb.QP" # [EXTQP BSK BSS] Database
% BEnRange
  0.00000 | 10.00000 | eV      # [BSS] Energy range
%
% BDmRange
  0.10000 | 0.10000 | eV      # [BSS] Damping range
%
% BEnSteps= 100      # [BSS] Energy steps
% BLongDir
  1.000000 | 0.000000 | 0.000000 | # [BSS] [cc] Electric Field
%
% BSEBands
  1 | 50 | # [BSK] Bands range
%
BSHayTrs= -0.02000 # [BSS] Relative [o/o] Haydock treshold.
% BndsRnXs
  1 | 50 | # [Xs] Polarization function bands
%
NGsBlkXs= 1        RL # [Xs] Response block size
% DmRngeXs
  0.10000 | 0.10000 | eV      # [Xs] Damping range
%
% LongDrXs
  1.000000 | 0.000000 | 0.000000 | # [Xs] [cc] Electric Field
%
% BndsRnXp
  1 | 50 | # [Xp] Polarization function bands
%
NGsBlkXp= 1005     RL # [Xp] Response block size
% LongDrXp
  0.1000E-4 | 0.000 | 0.000 | # [Xp] [cc] Electric Field

```

```
%
PPAPntXp= 27.21138      eV      # [Xp] PPA imaginary energy
```

The following is a sample job submission file for the GW-BSE calculation. This will submit your job to the parallel queue by the qsub/sbatch command (this is a long job compared to Quantum Espresso calculations and it will take a while to finish)

```
#!/bin/bash -l
#SBATCH --job-name=job-gw
#SBATCH --time=24:00:00
#SBATCH --partition=parallel
#SBATCH --nodes=4
#SBATCH --ntasks-per-node=24
mpirun -np 96 yambo -F yambo.in > yambo.out
```

After you finish the GW-BSE calculation with the above job script, the system will generate the output file, “o.eps\_q1\_haydock\_bse”. See below for a sample output file. You can extract the data points for the GW-BSE spectrum from this output.

```
# **      **      **      ****      ****      ****      ****
# /****      **      ****      /**/**      **/**      /**/**      **/**
# /******      **/**      /**/**      **/**      /**      /**      **      /**
# /****      **      /**      /**      /**      /**      /**      /**
# /**      *****/**      /**      /**      /**      /**      /**      /**
# /**      /**/**/**/**      /      /**      /**      /**      /**
# /**      /**      /**      /**      /**      /**      /**      /**
# //      //      //      //      //      //      //      //
#
# GPL Version 4.2.1 Revision 110. (Based on r.14778 h.7b4dc3
#                               MPI Build
#                               http://www.yambo-code.org
#
# Absorption @ Q(1) [q->0 direction] :0.1000E-4  0.000      0.000
#
#
# - Energies      are E < SAVE/ndb.QP + E Fit
# - Wavefunctions are Perdew, Burke & Ernzerhof(X)+Perdew, Burke & Ernzerhof(C)
#
# - The Green`s function is Retarded -
# - Using the Length Gauge -
# - [r,Vnl] *is* included -
#
# BSK|Identifier      :8262
#      |Dimension      :24379
#      |Bands          :1 - 26
```

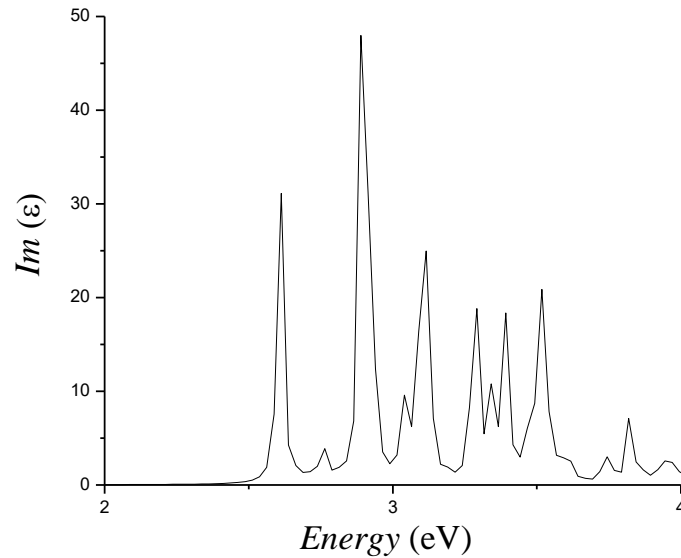
```

# |Exchange [res]: yes
# |Correlation [res]: yes
# |Kernel`s coupling : no
# |Exchange [cpl]: no
# |Correlation [cpl]: no
# |W interaction is bare : no
# |ALDA kernel in R-space : no
# |RL vectors [exchange]:4627
# |RL vectors [correlation]:1005
# |E/h energy range [ev]:-1.000000 - -1.000000
# |Coupling range [o/o]: 100.0000 - 100.0000
# W |Interaction is diagonal : no
# |Matrix size :1005
# |Bands :1 - 26
# |e/h energy range [ev]:-1.000000 - -1.000000
# |Poles [o/o]: 100.0000
# |RL vectors in the sum :4459
# |[r,Vnl] included : yes
# |Field direction :0.1000E-4 0.000000 0.000000
# |Coulomb Cutoff :none
# |xc-Kernel :none
# RIM|RL components [col]:0
# |Random points [col]:0
#
# Haydock|Accuracy (requested) [o/o]: -0.02000
# |Accuracy (reached) [o/o]: 0.01943
# |Iteration :974
#
# E/ev[1] EPS-Im[2] EPS-Re[3] EPSo-Im[4] EPSo-Re[5] E
PS`-Im[6] EPS`-Re[7]
#
0.000000 0.000000 4.900739 0.000000 4.286737 0.
000000 4.900739
0.2513E-1 0.1398E-3 4.901 0.7966E-4 4.287 0.1
398E-3 4.901
0.5025E-1 0.2798E-3 4.901 0.1594E-3 4.287 0.2
798E-3 4.901
0.7538E-1 0.4199E-3 4.902 0.2391E-3 4.288 0.4
199E-3 4.902
0.1005 0.5603E-3 4.904 0.3190E-3 4.288 0.5
603E-3 4.904
0.1256 0.7012E-3 4.905 0.3991E-3 4.289 0.7
012E-3 4.905
0.1508 0.8426E-3 4.907 0.4794E-3 4.290 0.8
426E-3 4.907
0.1759 0.9846E-3 4.909 0.5599E-3 4.292 0.9
846E-3 4.909

```

0.2010	0.1127E-2	4.912	0.6407E-3	4.293	0.1
127E-2	4.912				
0.2261	0.1271E-2	4.915	0.7218E-3	4.295	0.1
271E-2	4.915				
0.2513	0.1415E-2	4.918	0.8034E-3	4.297	0.1
415E-2	4.918				
0.2764	0.1561E-2	4.922	0.8853E-3	4.299	0.1
561E-2	4.922				
0.3015	0.1708E-2	4.926	0.9677E-3	4.301	0.1
708E-2	4.926				
0.3266	0.1856E-2	4.931	0.1051E-2	4.304	0.1
856E-2	4.931				
0.3518	0.2006E-2	4.936	0.1134E-2	4.307	0.2
006E-2	4.936				
--More-- (16%)					

Now you can plot “EPS-Im[2] vs E/ev[1]” using any preferred graphing tool such as excel, origin, etc..., to generate the spectrum as follows,



Absorption spectra for 2D MoS<sub>2</sub>, obtained with GW-BSE

If you are interested in further details on the convergence parameters for MoS<sub>2</sub> for publication quality work, please refer to: Molina-Sánchez *et al*, Phys. Rev. **B 88**, 045412 (2013), <https://arxiv.org/abs/1306.4257>.