

APPROXIMATION OF DIFFERENTIAL EQUATIONS - INTRODUCTION TO PERTURBATION EXPANSIONS

TAYLOR SERIES APPROXIMATION IS USED EVERYWHERE IN SCIENCE & ENGINEERING, TYPICALLY IN THE CONTEXT OF PERTURBATION APPROXIMATIONS OF DIFFERENTIAL EQUATIONS.

THE IDEA IS SIMPLY DEMONSTRATED BY LOOKING AT PERTURBATION APPROXIMATIONS OF ALGEBRAIC EQUATIONS.

EX. SUPPOSE WE WANT TO SOLVE $x^2 + \epsilon x - 1 = 0$ FOR SMALL ϵ .
THE EXACT SOLUTION IS:

$$x = -\frac{1}{2}\epsilon \pm \sqrt{1 + \frac{1}{4}\epsilon^2}$$

USING THE BINOMIAL EXPANSION, WE CAN EXPAND THESE SOLUTIONS AS A POWER SERIES IN ϵ :

$$x_{\pm}^{(1)} = 1 - \frac{1}{2}\epsilon + \frac{1}{8}\epsilon^2 - \frac{1}{128}\epsilon^4 + \dots$$

$$x_{\pm}^{(2)} = -1 - \frac{1}{2}\epsilon - \frac{1}{8}\epsilon^2 + \frac{1}{128}\epsilon^4 + \dots$$

WE KNOW FROM THE BINOMIAL THEOREM THAT THESE SERIES CONVERGE IF, AND ONLY IF, $|\epsilon| < 2$.

BUT SUPPOSE WE DIDN'T KNOW THE QUADRATIC FORMULA. WE COULD ASSUME THAT THE SOLUTION 'X' CAN BE WRITTEN AS A POWER SERIES:

$$x = x_0 + \epsilon x_1 + \epsilon^2 x_2 + \dots \quad \text{AND SEE WHAT HAPPENS...}$$

SUBSTITUTING INTO $x^2 + \epsilon x - 1$, WE GET:

$$(-1 + x_0^2) + (x_0 + 2x_0x_1)\epsilon + (x_1^2 + 2x_0x_2)\epsilon^2 + \dots = 0$$

TO SATISFY THIS EQUATION, ALL OF THE COEFFICIENTS OF ϵ^n MUST VANISH. LOOKING AT EACH COEFFICIENT,

$$\epsilon^0: x_0^2 - 1 = 0 \quad \text{OR} \quad x_0 = \pm 1. \quad \text{EASY.}$$

LET'S LOOK AT THE SOLUTION THAT BEGINS $x^{(0)} = 1 + \dots$

$$\text{AT } \epsilon^1: 1 + 2x_1 = 0 \quad \text{OR} \quad x_1 = -\frac{1}{2}$$

$$\text{AT } \epsilon^2: -\frac{1}{4} + 2x_2 = 0 \quad \text{OR} \quad x_2 = \frac{1}{8}$$

$$\text{AT } \epsilon^3: 2x_3 = 0 \quad \text{OR} \quad x_3 = 0$$

$$\text{AT } \epsilon^4: \frac{1}{64} + 2x_4 = 0 \quad \text{OR} \quad x_4 = -\frac{1}{128}$$

\vdots

$$\text{ALTOGETHER, } x^{(1)} = 1 - \frac{1}{2}\epsilon + \frac{1}{8}\epsilon^2 - \frac{1}{128}\epsilon^4 + \dots$$

IDEA: WE CAN ASSUME A TAYLOR SERIES SOLUTION TO SIMPLIFY THE PROBLEM!

WE CAN USE THIS SAME STRATEGY TO SOLVE NONLINEAR EQUATIONS.

eg. SHOW THAT ONE SOLUTION TO:

$$\cos[x] = e \cdot x$$

$$\text{IS: } x = \frac{\pi}{2} - \frac{\pi}{2}\epsilon + \frac{\pi}{2}\epsilon^2 - \frac{\pi}{48}(\pi^2 + 24)\epsilon^3 + \dots$$

eg. SHOW THAT THE SOLUTIONS TO $x^2 + e^{\epsilon x} = 5$ BEGIN:

$$x = \pm 2 - \epsilon/2 + \dots$$

PERTURBATION APPROXIMATION OF DIFFERENTIAL EQUATIONS

THE REAL POWER OF THIS APPROACH COMES IN SOLVING DIFFERENTIAL EQUATIONS.

EX. MOTION OF AN OBJECT PROJECTED UPWARD FROM THE SURFACE OF THE EARTH. LET $x(t)$ DENOTE THE HEIGHT ABOVE THE SURFACE. APPLYING NEWTON'S 2nd LAW:

$$\frac{d^2 x}{dt^2} = - \frac{g R^2}{(x+R)^2} \quad \text{FOR } t \geq 0$$

$x(0) = 0 \quad x'(0) = v_0.$ \nwarrow RADIUS OF THE EARTH.

IF WE ARE INTERESTED IN KEEPING THE HEIGHT SMALL (SMALL COMPARED TO WHAT), THEN WE CAN ARGUE $x+R \approx R$ AND,

$$\frac{d^2x}{dt^2} = -\frac{gR^2}{(x+R)^2} \approx -g.$$

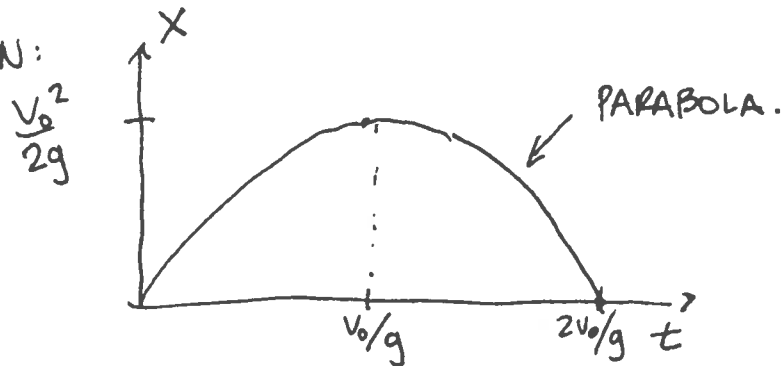
INTEGRATING ONCE: $\frac{dx}{dt} = -gt + C_1$

USING $x'(0) = v_0$ TO SET C_1 , $\frac{dx}{dt} = -gt + v_0$

INTEGRATING AGAIN: $x = -\frac{g}{2}t^2 + v_0t + C_2$

USING $x(0) = 0$ TO SET C_2 , $x = \left[-\frac{1}{2}gt^2 + v_0t \right] (*)$

PLOTTING THIS SOLUTION:



WE COMPUTED (*) BY ASSUMING THE HEIGHT WAS SMALL—THIS AMOUNTS TO KEEPING v_0 SMALL COMPARED WITH THE RADIUS OF THE EARTH. THE UNITS DON'T MAKE SENSE—REALLY WANT $\frac{v_0^2}{g} \ll R$.

HOW CAN WE MAKE A SYSTEMATIC APPROXIMATION? FIRST, NEED TO GET RID OF UNITS.

MEASURE DISTANCE IN UNITS OF $v_0^2/2g$ SO THAT

$$x = \left(\frac{v_0^2}{2g} \right) \hat{x}$$

↑
UNITS OF
DISTANCE
↑
UNITLESS
FUNCTION

AND MEASURE TIME IN UNITS OF v_0/g SO THAT

$$t = (v_0/g) \hat{t}.$$

THEN:

$$\frac{d^2 x}{dt^2} = \frac{(v_0^2/2g)}{(v_0/g)^2} \frac{d^2 \hat{x}}{d\hat{t}^2} = \frac{-g R^2}{\left(\frac{v_0^2}{2g} \hat{x} + R\right)^2} = \frac{-g}{\left(\frac{v_0^2}{2gR} \hat{x} + 1\right)^2}$$

CLEAN THIS UP, $\frac{d^2 \hat{x}}{d\hat{t}^2} = \frac{-2}{\left(\frac{v_0^2}{2gR} \hat{x} + 1\right)^2}$

THE GROUP $(v_0^2/2gR)$ IS A UNITLESS RATIO BETWEEN THE HEIGHT OF THE PARABOLIC SOLUTION & THE RADIUS OF THE EARTH. CALL THIS:

$$\epsilon = \frac{v_0^2}{2gR} \quad \text{AND FIND } \hat{x} = \hat{x}_0 + \epsilon \hat{x}_1 + \dots$$

THAT IS, TAKE THE DIFFERENTIAL EQUATION:

$$\frac{d^2 \hat{x}}{d\hat{t}^2} = \frac{-2}{(\epsilon \hat{x} + 1)^2} \quad \hat{x}(0) = 0 \quad \frac{v_0}{2} \frac{d\hat{x}(0)}{d\hat{t}} = \frac{dx(0)}{dt} = v_0 \quad \text{or} \quad \frac{d\hat{x}(0)}{d\hat{t}} = 2.$$

AND SUBSTITUTE $\hat{x} = \hat{x}_0(\hat{t}) + \epsilon \hat{x}_1(\hat{t}) + \epsilon^2 \hat{x}_2(\hat{t}) + \dots$ AND SEE WHAT HAPPENS.

$$\text{AT } \epsilon^0: \left. \begin{array}{l} \frac{d^2 \hat{x}_0}{d\hat{t}^2} = -2 \\ \hat{x}_0(0) = 0 \\ \hat{x}'_0(0) = 2 \end{array} \right\} \begin{array}{l} \frac{d\hat{x}_0}{d\hat{t}} = -2\hat{t} + C = -2\hat{t} + 2 \\ \hat{x}_0 = -\hat{t}^2 + 2\hat{t} + C \end{array}$$

$$\text{AT } \epsilon^1: \left. \begin{array}{l} \frac{d^2 \hat{x}_1}{d\hat{t}^2} = 4\hat{x}_0 = (-\hat{t}^2 + 2\hat{t})^2 \\ \hat{x}_1(0) = 0 \quad \hat{x}'_1(0) = 0 \end{array} \right\} \begin{array}{l} \frac{d\hat{x}_1}{d\hat{t}} = \left(-\frac{\hat{t}^3}{3} + \hat{t}^2\right) + C = \left(-\frac{\hat{t}^3}{3} + \hat{t}^2\right) \\ \hat{x}_1 = \left(-\frac{\hat{t}^4}{12} + \frac{\hat{t}^3}{3}\right) + C = -\frac{\hat{t}^4}{12} + \frac{\hat{t}^3}{3} \end{array}$$

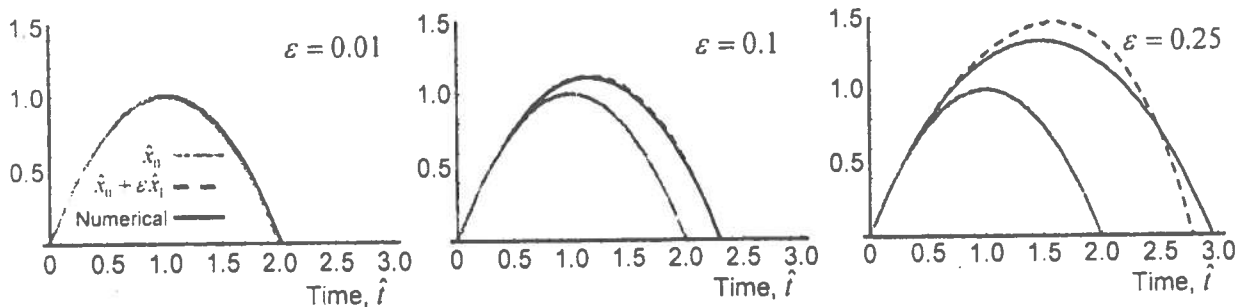
KEEPING TWO TERMS:

$$\begin{aligned}\hat{x}(\hat{t}) &= (-\hat{t}^2 + 2\hat{t}) + \epsilon \left(-\frac{\hat{t}^4}{12} + \frac{\hat{t}^3}{3} \right) + \dots \\ &= (2\hat{t} - \hat{t}^2) + \epsilon \frac{\hat{t}^3}{3} \left(1 - \frac{\hat{t}}{4} \right) + \dots\end{aligned}$$

RESTORE THE UNITS: $\hat{x} = \left(\frac{2g}{v_0^2} \right) x$ AND $\hat{t} = \left(\frac{g}{v_0} \right) t$

$$\begin{aligned}x(t) &\approx \left(\frac{v_0^2}{2g} \right) \left(-\left(\frac{g}{v_0} \right)^2 t^2 + 2 \left(\frac{g}{v_0} \right) t \right) + \frac{\epsilon}{3} \left(\frac{g}{v_0} \right)^3 t^3 \left(1 - \left(\frac{g}{v_0} \right) \frac{t}{4} \right) + \dots \\ &= \left[-\frac{1}{2} g t^2 + v_0 t \right] + \epsilon \left[\frac{g^2 t^3}{6 v_0} \left(1 - \frac{g t}{4 v_0} \right) \right] + \dots\end{aligned}$$

THE DIFFERENTIAL EQUATION CAN BE APPROXIMATED NUMERICALLY USING A VARIATION ON EULER'S METHOD. COMPARING THE NUMERICAL SOLUTION TO A FEW TERMS OF THE PERTURBATION SERIES,



AS A POINT OF REFERENCE, $\epsilon \approx 1$ IS APPROXIMATELY EQUAL TO THE ESCAPE VELOCITY $v_{\text{ESCAPE}} = 11.2 \text{ km/s}$; $\epsilon \approx 0.3$ CORRESPONDS ROUGHLY TO THE SPEED OF A BULLET 1200 km/h

$$[g \approx 10 \text{ m/s}^2, R \approx 6400 \text{ km}].$$

POINCARÉ - LINDSTEDT METHOD

THERE ARE MANY EXAMPLES WHERE THE NAIVE PERTURBATION SERIES DOESN'T WORK; AND MANY METHODS DESIGNED TO ADDRESS THESE DIFFICULTIES.

CLASSIC TEXT: 'ASYMPTOTIC METHODS & PERTURBATION THEORY' BY BENDER & ORSZAG.

COURSE: AMATH 737 - PERTURBATION METHODS & ASYMPTOTIC ANALYSIS.

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EX. DUFFING EQUATION $x''(t) + x(t) - \epsilon x^3(t) = 0$; $x(0) = 1$ & $x'(0) = 0$.

AS BEFORE, TRY $x(t) = x_0(t) + \epsilon x_1(t) + \epsilon^2 x_2(t) + \dots$


COLLECTING LIKE-POWERS OF ϵ , WE END UP WITH THE SET OF EQUATIONS:

$$\underline{\epsilon^0}: x_0'' + x_0 = 0; x_0(0) = 1 \text{ \& } x_0'(0) = 0 \Rightarrow x_0(t) = \cos t.$$

$$\text{AT } \underline{\epsilon^1}: x_1'' + x_1 = x_0^3 = \cos^3 t = \underbrace{\frac{3}{4} \cos t - \frac{1}{4} \cos 3t}; x_1(0) = 0 \text{ \& } x_1'(0) = 0$$

THIS TERM WILL CAUSE PROBLEMS...

$$x_1(t) = \frac{1}{32} (\cos t - \cos 3t) + \frac{3}{8} t \sin t.$$

HERE IS THE PROBLEM: THIS TERM  GROWS WITH 't'. FOR $t > \epsilon$, THE ORDERING IMPLIED BY OUR PERTURBATION SERIES BREAKS-DOWN.

~~WHAT IS THE PROBLEM~~ WHERE IS THIS COMING FROM? THE NAIVE EXPANSION RESULTS IN A SET OF RECURSIVE EQUATIONS:

$$x_n'' + x_n = x_{n-1}^3$$

WHERE $x_{n-1}(t)$ OSCILLATES AT THE SAME FREQUENCY AS x_n , IRRESPECTIVE OF ϵ ! LIKE A DOUBLE-BOUNCE ON A TRAMPOLINE, THIS SO-CALLED 'RESONANT FORCING' LEADS TO DIVERGENCE.

POINCARÉ'S IDEA: LET THE FREQUENCY OF THE OSCILLATIONS CHANGE WITH ϵ !

ie^o $X(\tau) = X_0(\tau) + \epsilon X_1(\tau) + \epsilon^2 X_2(\tau) + \dots$

WHERE $\tau = t(1 + \epsilon \omega_1 + \epsilon^2 \omega_2 + \dots)$ THESE ARE CONSTANTS THAT WILL BE CHOSEN TO ELIMINATE THE DIVERGENT TERMS.

USING THE CHAIN RULE -

$$\frac{dx}{dt} = \frac{dx}{d\tau} \frac{d\tau}{dt} = \frac{dx}{d\tau} (1 + \epsilon \omega_1 + \epsilon^2 \omega_2 + \dots) \quad \& \quad \frac{d^2x}{dt^2} = \frac{d^2x}{d\tau^2} (1 + \omega_1 \epsilon + \omega_2 \epsilon^2 + \dots)^2 \approx \frac{d^2x}{d\tau^2} (1 + 2\epsilon \omega_1 + \dots)$$

WITH SUBSTITUTION INTO THE ORIGINAL DIFF. EQ:

$$(1 + \epsilon \omega_1 + \epsilon^2 \omega_2 + \dots)^2 X''(\tau) + X(\tau) - \epsilon X^3(\tau) = 0 \quad ; \quad X(0) = 1 \quad \& \quad X'(0) = 0.$$

COLLECTING POWERS OF ϵ :

$$\epsilon^0: X_0''(\tau) + X_0(\tau) = 0; \quad X_0(0) = 1 \quad \& \quad X_0'(0) = 0 \quad \Rightarrow \quad X_0(\tau) = \cos \tau$$

$$\epsilon^1: X_1''(\tau) + X_1(\tau) = -2\omega_1 X_0''(\tau) + X_0^3(\tau); \quad X_1(0) = X_1'(0) = 0.$$

$$= 2\omega_1 \cos \tau + \cos^3 \tau$$

$$= (2\omega_1 + \frac{3}{4}) \cos \tau - \frac{1}{4} \cos 3\tau$$

IT IS THIS TERM THAT CAUSES ALL THE PROBLEMS. SO CHOOSE $\omega_1 = -3/8$ TO MAKE IT DISAPPEAR!

$$\Rightarrow X_1(\tau) = \frac{1}{32} (\cos \tau - \cos 3\tau)$$

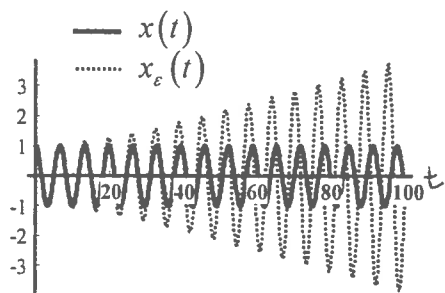
ALTOGETHER,

$$X(\tau) \approx \cos \tau + \frac{\epsilon}{32} (\cos \tau - \cos 3\tau) \quad \text{WITH } \tau = t(1 - \frac{3}{8}\epsilon).$$

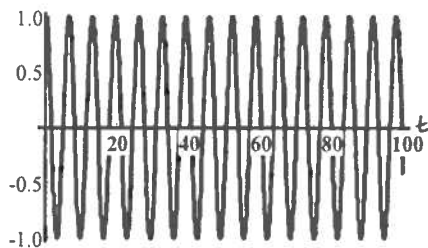
HOW DOES IT ~~WORK~~ ~~IT~~ COMPARE WITH THE EXACT (ie^o NUMERICAL) SOLUTION? VERY WELL.

WE CAN COMPARE THE ONE-TERM APPROXIMATIONS, i.e., $x_\epsilon(t) \approx x_0 + \epsilon x_1$, TO THE NUMERICAL SOLUTION FOR DIFFERENT VALUES OF ϵ .

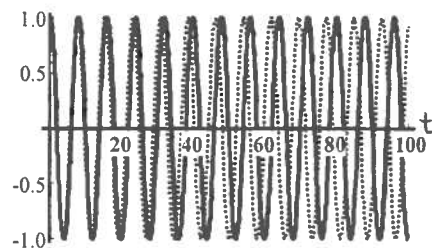
FIRST THING TO NOTICE IS THE DIVERGENCE OF THE NAIVE APPROXIMATION & HOW THE POINCARÉ APPROXIMATION (BY DESIGN) REMAINS BOUNDED:



Naive approximation, $\epsilon = 0.1$

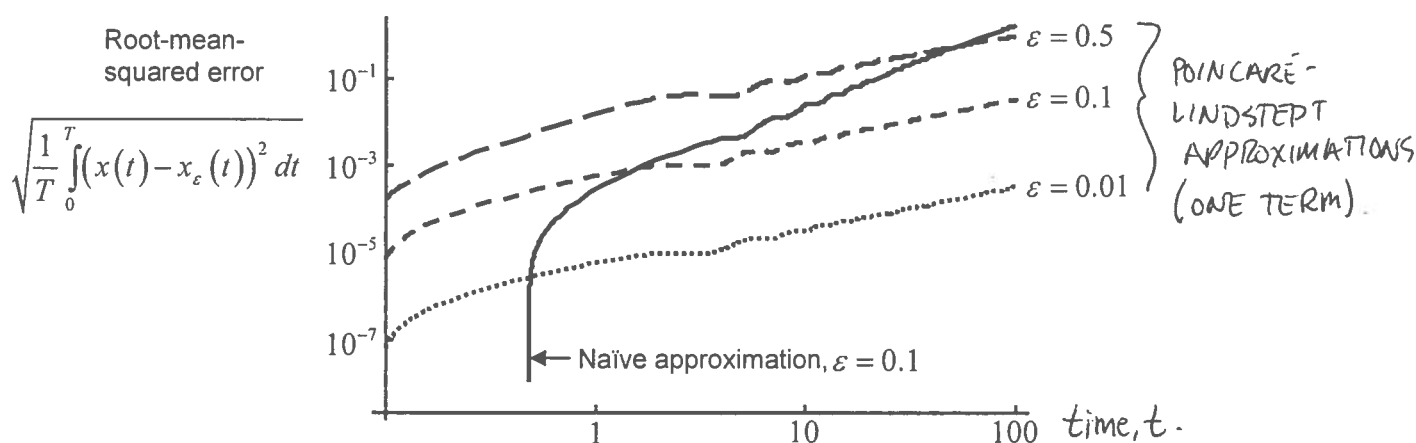


Poincaré-Lindstedt, $\epsilon = 0.1$



Poincaré-Lindstedt, $\epsilon = 0.5$

WE CAN BE MORE QUANTITATIVE BY COMPARING THE ROOT-MEAN-SQUARE ERROR BETWEEN THE NUMERICAL & PERTURBATION SOLUTIONS (ROOT-MEAN-SQUARE IS RELATED TO THE L^2 -NORM).



EVENUALLY, THE POINCARÉ APPROXIMATION IS COMPLETELY OUT-OF-PHASE WITH THE NUMERICAL SOLUTION, BUT STILL BOUNDED IN AMPLITUDE.