# FOXSI-2 microflare imaging spectroscopy -- UPDATED

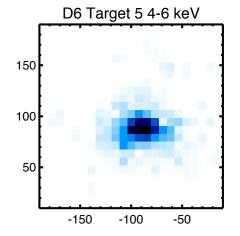
Lindsay

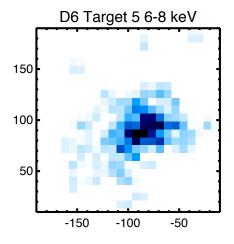
2015 April 1

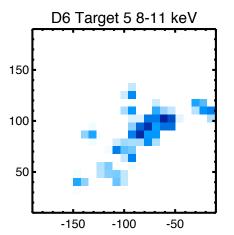
#### Last target

 Detector 6 in different energy bands





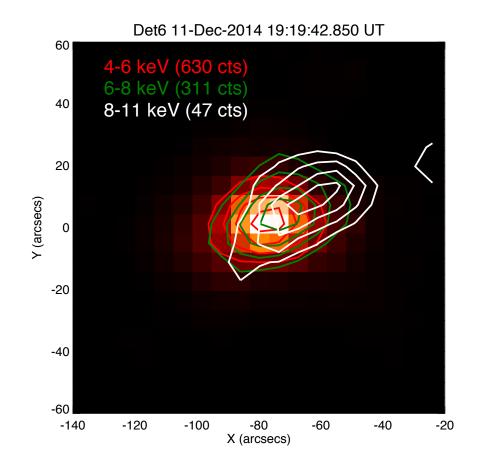




### Detector 6 images of final target

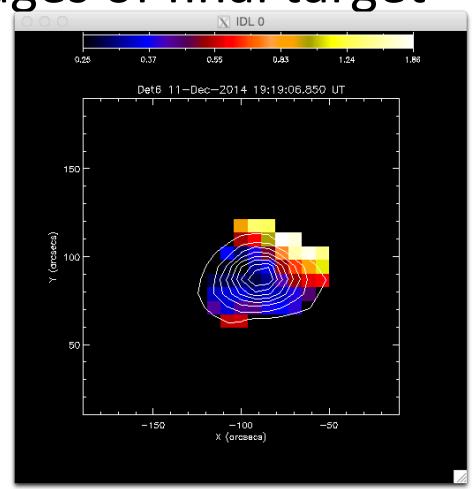
- Smoothed images
  - Smoothed over 2 strips.

 Significant source location / shape change with energy.



#### Detector 6 images of final target

- Ratio of high-energy to low-energy emission is a measure of temperature.
- Not calculating actual temperatures yet; need solidified instrument response for that.
- Hot plasma (or ront ermal electrons)
  ge of source.

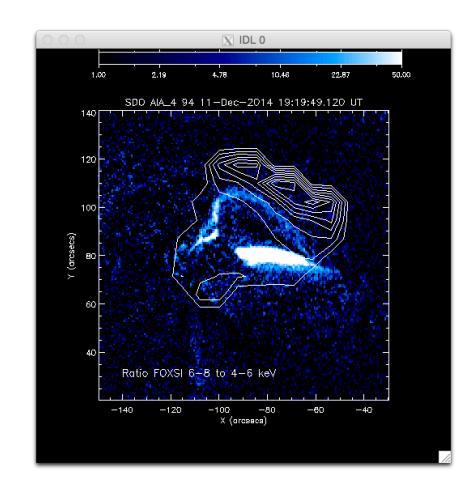


Ratio of 6-8 keV counts to 4-6 keV counts.

#### On top of AIA 94

- Alignment was done by lining up brightest points in FOXSI and AIA by eye. Could be wrong!
  - Co-aligning with RHESSI will help...

- A different alignment could put the high-temp plasma on the AIA ridge.
- Or perhaps the hot plasma is liar at to, or above the ridge.

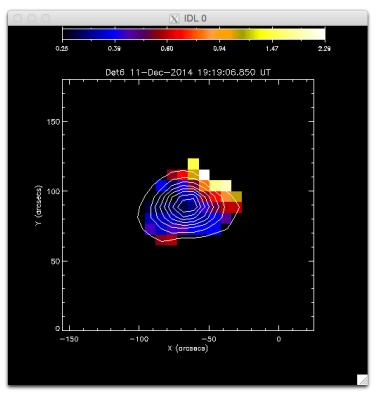


"Knottiness" is the pixel size. No, it's not nanoflares. ©

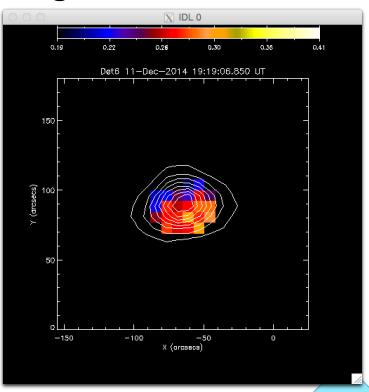


#### Recently, ran the exact same code...

But instead of this...



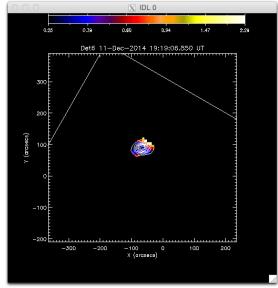
...I got this!

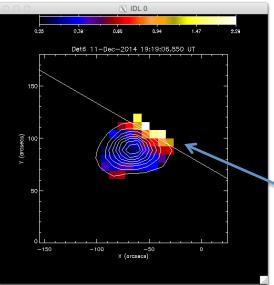




## What the & \*%\$?!?

#### Answer: old result used incorrect cal.





- (Earlier) assumption: exact calibration doesn't matter for this ratio method because only count ratios are important.
- But the earlier result used the incorrect calibration files, which meant that the strip-by-strip calibration was wrong!

This is probably an incorrectly calibrated strip.



#### Sighhh...so what does it really look like?

