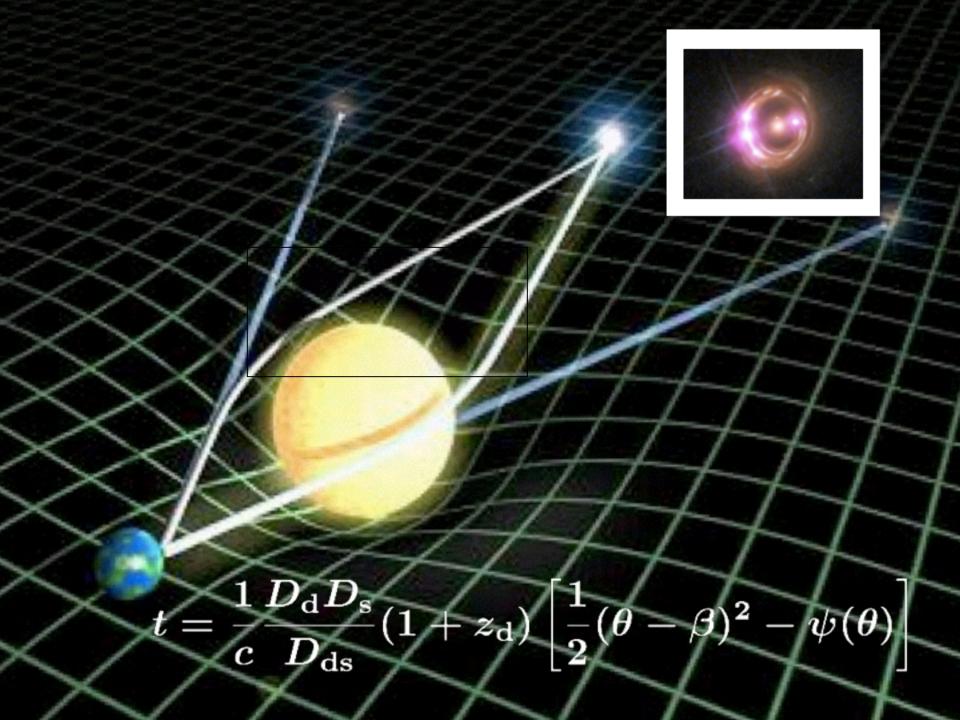
LSST Observing Strategy Metric Analysis Example:

Strong Lens Time Delays

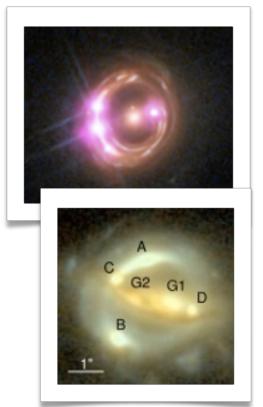
Phil Marshall (SLAC/KIPAC)

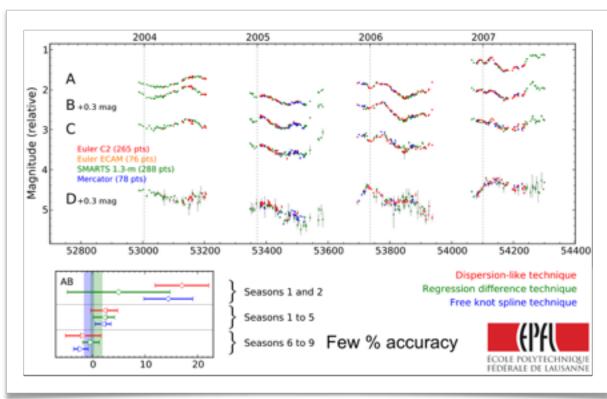




Time Delay Distances



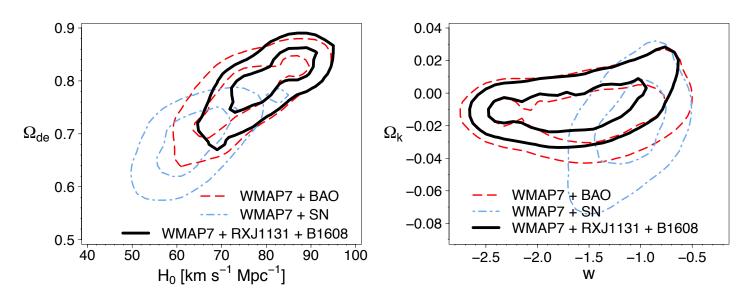




 Strong lens time delay distances are one of the cosmological probes being investigated by the DESC

Time Delay Distances





- LSST will automatically monitor thousands of systems, randomly distributed over the sky
- We need accurate time delays and mass models (from follow-up data) for a sample of 10²⁻³ in order to measure H₀ to sub-% precision and help constrain Dark Energy models

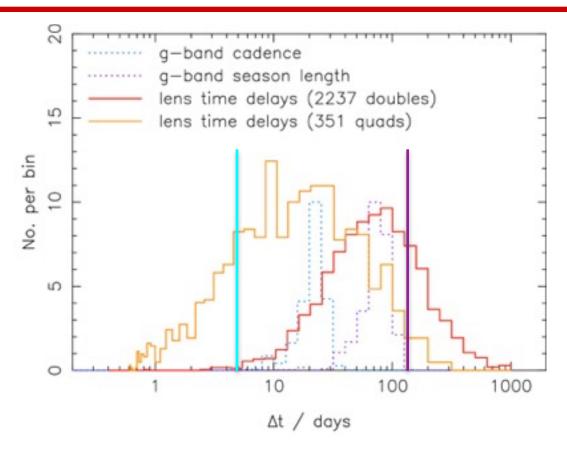
Time Delay Distances



- To measure each time delay distance to 5%, we need the time delay precision to be around 3% per lens
- To measure cosmological parameters with sub-percent accuracy, we need the mean systematic error to be below ~0.2%, and the sample to be larger than 100 lenses
- 1) How well will we be able to measure lensed quasar time delays with LSST, given the default "universal cadence" observing strategy?
- 2) Which aspects of the LSST observing strategy would we, the DESC SL working group, change, and how?

Cadence and season length

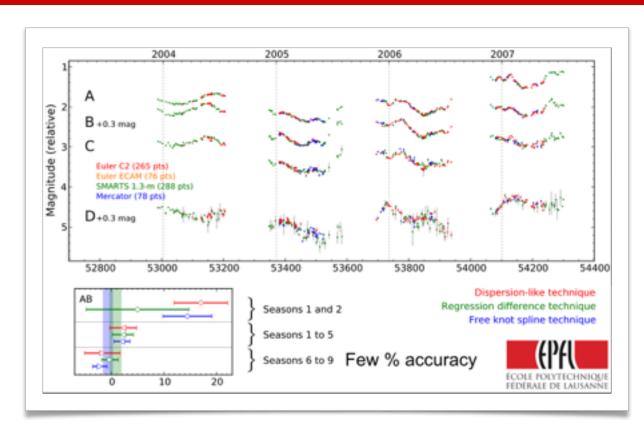




Most lens time delays fall between a few days and a few months: we expect the night-to-night cadence and the season length to be important

Campaign length





Accuracy increases as the number of seasons builds up: campaign length is also important

Time Delay Challenge



Goals:

- Assess performance of time delay estimation algorithms on LSST-like data (cf STEP in WL community)
- Assess the impact of universal cadence strategy on time delay estimation, and possibly recommend changes

Plan:

- "Evil Team" generated a large set of simulated lightcurves spanning expectations for Stage II-IV
- Challenged wider community "Good Teams" to infer time delays blindly, and submit results

Evil Team for TDC1:

Kai Liao, Greg Dobler, Tommaso Treu, Chris Fassnacht, Nick Rumbaugh, Phil Marshall

TDC timeline

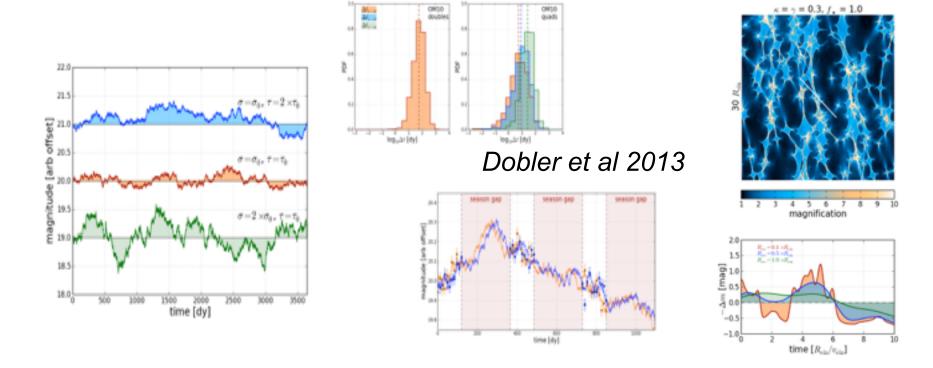


- Paper describing challenge: arXiv:1310.4830
- "TDC0" small "training set," to get Good Teams started: released October 2013, soft deadline December 2013
- "TDC1" large "test set," for primary analysis: released
 December 2013, hard deadline for submissions July 1 2014
- Robotic TDC0 feedback ongoing until then
- Evil Team analysed entries, wrote paper with Good Teams, submitted August 2014. ~1 year, not including sims R&D



TDC1 ingredients



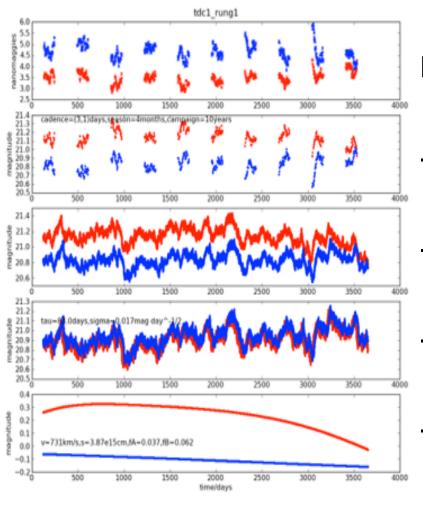


- Lightcurve catalog data only (single filter)
- Realistic AGN variability, microlensing, lenses
- "OpSim-ish" observations (depth, cadence, etc)

TDC1: example lightcurves



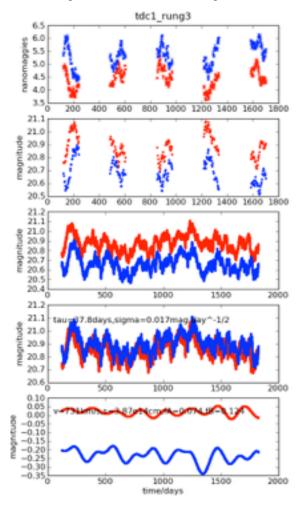
10 years, 3 day cadence



Mock data

- without noise
- fully sampled
- no lensing
- microlensing

5 years, 3 days



Multiple TDC metrics



$$A = \frac{1}{fN} \sum_{i} \frac{\tilde{\Delta t_i} - \Delta t_i}{|\Delta t_i|}$$

$$P = \frac{1}{fN} \sum_{i} \left(\frac{\sigma_i}{|\Delta t_i|} \right)$$

$$\chi^2 = \frac{1}{fN} \sum_{i} \left(\frac{\tilde{\Delta t_i} - \Delta t_i}{\sigma_i} \right)^2$$

1.
$$f > 0.3$$

2.
$$0.5 < \chi^2 < 2$$
 (1.09f)

3.
$$P < 15\%$$
 (3%)

4.
$$A < 15\%$$
 (0.2%)

- 4 metrics (chisq, A, P, f) used to define TDC0 pass
- No single metric, no leaderboard, no cosmology
- LSST/TDC1 requirements are stricter (in parentheses)

TDC1: challenge "rungs"

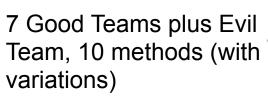


Rung	Cbar	Cerr	Season	Campaign	Nepochs	maglim	N_lens
0	3.0	1.0	8.0	5	400	 24	1000
1	3.0	1.0	4.0	10	400	24	1000
2	3.0	0.0	4.0	5	200	24	1000
3	3.0	1.0	4.0	5	200	24	1000
4	6.0	1.0	4.0	10	200	24	1000

- Rungs enable simple A-B testing of LSST observing scenarios, assuming that all filters can be used in joint analysis
- Double (SNe) Cadence, all filters: cadence = 3 +/- 1 days
- Worst-case Universal Cadence, all filters: cadence = 6 +/- 1 days
- Double, and more uniform cadences might be possible with customization of observing strategy, as suggested by SNe group, who are happy to trade campaign length for cadence.

TDC1: results

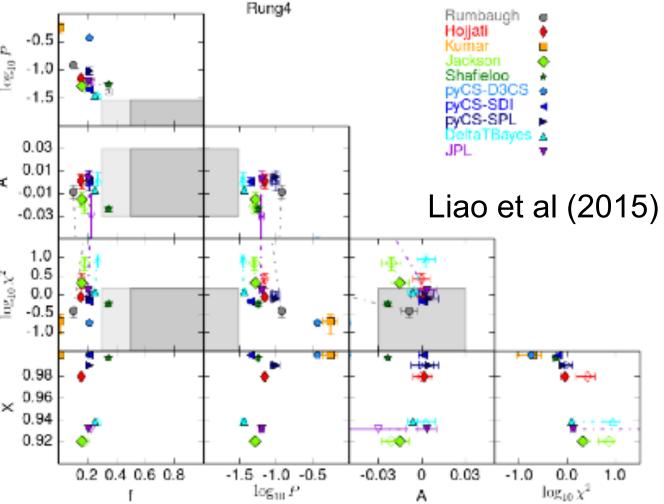




Outliers a problem for some methods

Required high accuracy is achievable, even at universal cadence.

Fraction of useable systems is ~20%



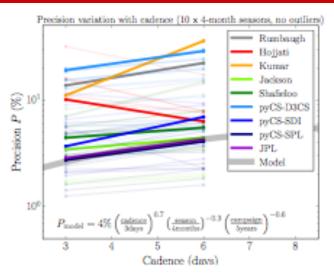
TDC1: cadence etc

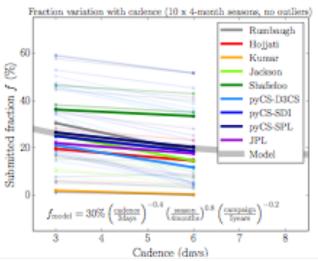


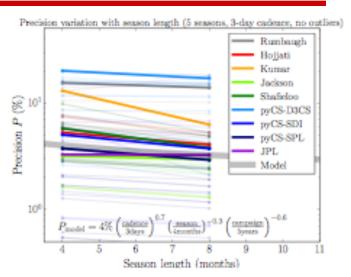
Cadence drives precision.

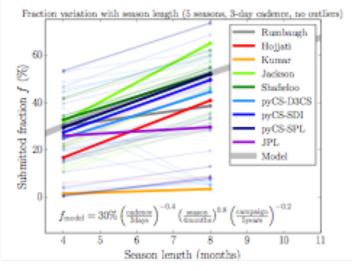
Season length also matters for the useable fraction (rate).

Was accuracy prioritized by Good Teams? It depends much less on cadence and season length.









TDC1: results

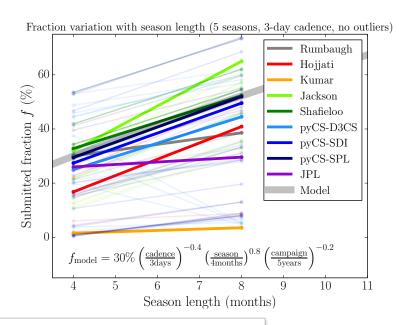


- Analysis of 1000s of light curves in manageable time (6 months) was demonstrated by 7 Good Teams
- Mean accuracy (or bias) of 0.06% was achieved, independent of cadence, in LSST-like campaigns of just five 4-month seasons. Many methods yielded sub-percent accuracy.
- Low cadence and season length cause low precision and useable fraction. Mean precision predicted for the best method in a 5-day cadence, 10-season campaign is 3.8%. The useable fraction in this case is 20%, which corresponds to 400 lenses.

TDC model metrics



The TDC model approximately captures the dependence of our metrics on observing strategy:



$$|A|_{\text{model}} \approx 0.06\% \left(\frac{\text{cad}}{3\text{days}}\right)^{0.0} \left(\frac{\text{sea}}{4\text{months}}\right)^{-1.0} \left(\frac{\text{camp}}{5\text{years}}\right)^{-1.1}$$

$$P_{\text{model}} \approx 4.0\% \left(\frac{\text{cad}}{3\text{days}}\right)^{0.7} \left(\frac{\text{sea}}{4\text{months}}\right)^{-0.3} \left(\frac{\text{camp}}{5\text{years}}\right)^{-0.6}$$

$$f_{\text{model}} \approx 30\% \left(\frac{\text{cad}}{3\text{days}}\right)^{-0.4} \left(\frac{\text{sea}}{4\text{months}}\right)^{0.8} \left(\frac{\text{camp}}{5\text{years}}\right)^{-0.2}$$



mafContrib/tdcMetric.py

- The TDC "cadence", "season length" and "campaign length" are metrics in the MAF sense
- They summarize the LSST observations of a particular point on the sky, and our time delay model metrics depend on them



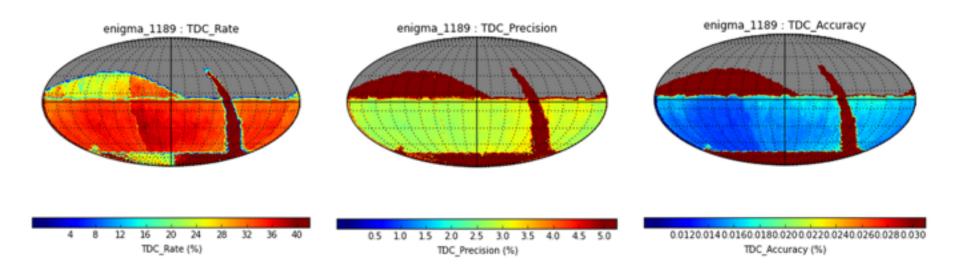
mafContrib/tdcMetric.py

```
class TdcMetric(BaseMetric):
    def run(self, dataSlice, slicePoint=None):
        # Calculate accuracy from combined individual metrics.
        camp = self.campaignLength.run(dataSlice)
        sea = self.seasonLength.run(dataSlice)
        cad = self.meanNightSeparation.run(dataSlice)
        if sea * cad * camp == 0:
            accuracy = self.badval
            precision = self.badval
            rate = 0.0
    else:
        accuracy = 0.06 * (self.seaNorm / sea) * (self.campNorm / camp)**(1.1)
        precision = 4.0 * (cad/self.cadNorm)**(0.7) * (self.seaNorm/sea)**(0.3) * (self.campNorm/camp)**(0.6)
        rate = 30. * (self.cadNorm/cad)**(0.4) * (sea/self.seaNorm)**(0.8) * (self.campNorm/camp)**(0.2)
    return {'accuracy':accuracy, 'precision':precision, 'rate':rate}
```

- TDC model accuracy, precision and rate are components of the "complex" MAF metric "TdcMetric"
- Now we can look at, for example, how each of these metrics varies across the sky



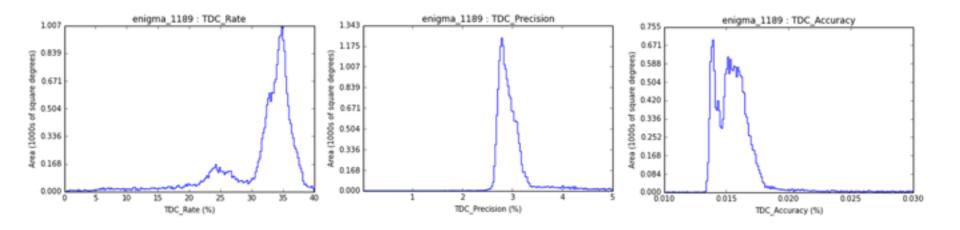
mafContrib/tdcMetric.py



 Main survey area is remarkably uniform in all metrics after 10 years; other survey areas likely not useable



mafContrib/tdcMetric.py



- Precision of 3% at accuracy less than 0.02% accuracy is very promising for time delay lens cosmology
- Rates of 30% may correspond to 600 useable lenses but integral has to be done more carefully
- "Opsim-ish" assumptions in TDC were pessimistic?

Final Thoughts



- Caveat: optimal combination of all filters has been assumed so far. TDC2 planned to test this, but more conservative metric plots could be made. Only use gri?
- Approach involves some heavy summarizing: is it necessary to relax this, and simulate mock surveys? To cope with "rolling cadence"?
- Where might we be after 3 years? 5 years? Early science?
- Time delay accuracy, precision and rate are still intermediate metrics: the number N of useable lenses would be good, and a Dark Energy figure of merit W even better.
- Strategy: go back to TDC1 results, and look at N and W for each method, as a function of observing strategy. Extend model, and then upgrade TdcMetric.py