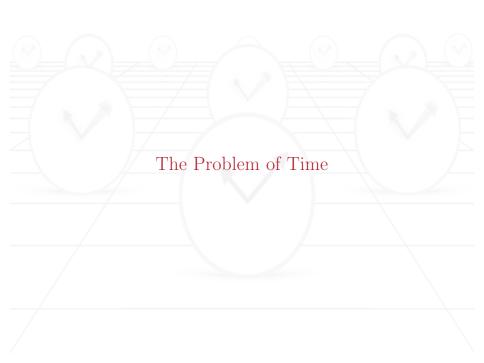
Quantum Time-Dilation in Qubit Hypersurfaces

Dario Cafasso University of Pisa

 $\mathrm{May}\ 30,\ 2024$

The Problem of Time

The Problem of Time Evolution without evolution The Problem of Time Evolution without evolution Time-Dilated Schrödinger equation The Problem of Time Evolution without evolution Time-Dilated Schrödinger equation **Qubit Hypersurfaces**



The Problem of Time



[Höhn et al., 2021, Castro-Ruiz et al., 2020, Isham, 1992]

In **General Relativity**: covariance under coordinate transformations \Longrightarrow time plays no fundamental role.

In practice: time is measured using physical clocks, obeying the laws of our most fundamental theories.

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The Problem of Time

The Page and Wootters mechanism



[Page and Wootters, 1983]

"If a quantum system is **truly closed**, there's **no place** in the theory for an **external (time)** parameter."

$$\mbox{Given} \ |\Psi\rangle \in \mathcal{H} \implies \ i\hbar \frac{\mathrm{d}}{\mathrm{d}t} \left|\Psi\right\rangle = \hat{\mathcal{H}} \left|\Psi\right\rangle = 0$$

The clock C is part of the Universe

$$\hat{\mathcal{H}} = \hat{\mathcal{H}}_C + \hat{\mathcal{H}}_S$$

Evolution: correlations between C and S

History state
$$\ \mapsto |\Psi
angle = \int \mathrm{d}t \, |t
angle_{\mathcal{C}} \, |\psi(t)
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► Ideal clocks are non-physical

$$\hat{T} |t\rangle_{C} = t |t\rangle_{C} \quad \left[\hat{T}, \hat{\mathcal{H}}_{C}\right] = i\hbar$$

since $\hat{\mathcal{H}}_{\mathcal{C}}$ is unbounded from below. [Isham, 1992]

Interactions introduce temporal non-locality

$$i\hbar rac{\mathrm{d}}{\mathrm{d}t} \ket{\psi(t)}_S = \int \mathrm{d}s \, K(t,s) \ket{\psi(s)}_S$$

n the modified Schrödinger equation. [Smith and Ahmadi, 2019]

Measurements make the clock stuck

$$\mathcal{A}(t_f \mid t_i) = \delta(t_f - t_i)$$

giving rise to Kuchar's criticisms. [Hausmann et al., 2023]



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[Smith, 2022]

About measurements, Page argued:

"I simply wish to argue that all of the testable predictions of quantum mechanics appear to arise from one-time conditional probabilities. [...] We can never directly test what happened yesterday, but we can check the consequences that a hypothetical scenario for yesterday has on the situation today." [Page, 1989]

About the past, in the words of Wheeler:

"It is wrong to think of the past as 'already existing' in all detail. The 'past' is theory. The past has no existence except as it is recorded in the present." [Wheeler, 1978]



[Hausmann et al., 2023]

Purified measurement Approach:

[Giovannetti et al., 2015]

Enters the constraint, i.e. Measurement:

the history state

Time ordering: Internal

Non-ideal Discrete-time or clock case:

non-unitary evolution



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Twirled observable [Hoehn et al., 2021]

Commutes with the constraint

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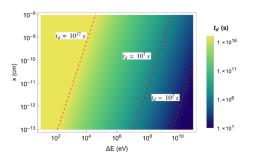
constraint

External

The measurement scheme depends on the clock model. What about interactions?



Assuming the relativistic **mass-energy equivalence** \implies time-local dynamics, but undefined metric structure.



The clocks get entangled and decoherence time t_d becomes accessible at GeV energy scales. [Castro-Ruiz et al., 2017]

Conventional time-dilation can be recovered via a semi-classical coupling with the gravitational field. [Favalli and Smerzi, 2022]



- ls it necessary to introduce a built-in notion of time?
- A laboratory quantum system as an evolving universe?
- What changes with a gravitational-like interaction?
- What if we attach a **qubit clock** to every point in space



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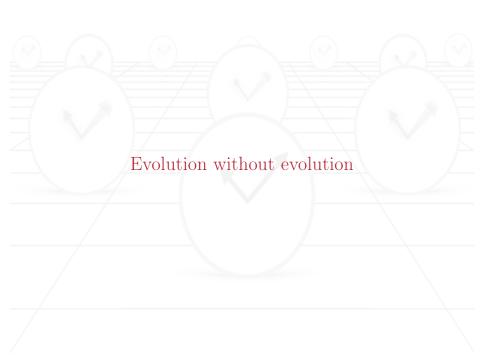
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A finite-dimensional quantum system C with Hilbert space \mathcal{H}_C .

Given a non-degenerate Hermitian operator $\hat{\mathcal{H}}_{\mathcal{C}} \in \mathcal{B}(\mathcal{H}_{\mathcal{C}})$, we have

$$\hat{V}_C(t) := e^{-\frac{i}{\hbar}\hat{\mathcal{H}}_C t} \implies \hat{\mathcal{H}}_C \hat{V}_C(t) = i\hbar \frac{\mathrm{d}}{\mathrm{d}t} \hat{V}_C(t)$$

Given a reference state $|e\rangle_C$ we introduce a *time state*

$$|t\rangle_C := V_C(t)|e\rangle_C$$

and a resolution of the Identity

$$\mathbb{1}_C = \int \mathrm{d}\mu(t) \, |t\rangle\!\langle t|_C$$





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The most general model



The eigenstates of $\hat{\mathcal{H}}_{\mathcal{C}}$ as

$$\hat{\mathcal{H}}_{C} \left| w_{k} \right\rangle_{C} = \hbar w_{k} \left| w_{k} \right\rangle_{C} \text{ with } w_{k} = w_{0} + r_{k} \frac{2\pi}{T},$$

in which

$$\frac{r_k}{r_1} := \frac{A_k}{B_k} = \frac{w_k - w_0}{w_1 - w_0}$$
 and $T = \frac{2\pi r_1}{w_1 - w_0}$,

where $A_k, B_k \in \mathbb{N}$ are coprimes and r_1 is the LCM of B_k .

If $r_k = 0, 1, 2, ..., d_C - 1$ and $\Delta t = T/d_C \Longrightarrow$ two possible procedures for $d_C \to +\infty$:

- ightharpoonup Continuous limit, i.e. $\Delta t
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- ▶ Thermodynamic limit, i.e. $T \to +\infty$

Performed in a sequence \implies ideal clock model

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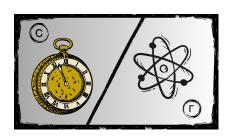
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The Quantum Universe The system to describe



Bipartite quantum system U = C + S with $\mathcal{H} = \mathcal{H}_C \otimes \mathcal{H}_S$.

Given $\hat{\mathcal{H}} \in \mathcal{B}(\mathcal{H})$, the state $|\Psi\rangle \in \mathcal{H}$ such that $\hat{\mathcal{H}}|\Psi\rangle = 0$ is stationary.



Schmidt decomposition with rank I

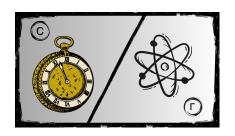
$$|\Psi\rangle = \sum_{n=0}^{r-1} \sqrt{\lambda_n} |\phi_n\rangle_C |\xi_n\rangle_S$$

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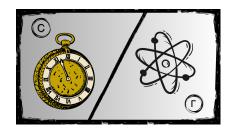
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Schmidt decomposition with rank $\it r$

$$|\Psi\rangle = \sum_{n=0}^{r-1} \sqrt{\lambda_n} |\phi_n\rangle_C |\xi_n\rangle_S$$



 $\mathsf{Schmidt} + \mathsf{Identity} \implies \mathit{history state}$

$$|\Psi
angle = \int \mathrm{d}\mu(t) \; \mathsf{a}(t) \, |t
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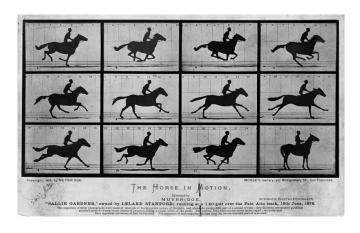
where

$$|\psi(t)\rangle_{S} = \frac{1}{a(t)} \langle t|_{C} |\Psi\rangle \text{ with } a(t) = \|\langle t|_{C} |\Psi\rangle\|$$

is called a conditional Schrödinger state.

The Quantum Universe History state and dynamical laws







Considering $\hat{\mathcal{H}} = \hat{\mathcal{H}}_{\mathcal{C}} + \hat{\mathcal{H}}_{\mathcal{S}}$ we have

$$\left\langle t\right\vert _{C}\hat{\mathcal{H}}\left\vert \Psi\right\rangle =\left(-i\hbar\frac{\mathrm{d}}{\mathrm{d}t}+\hat{\mathcal{H}}_{S}\right) \left\langle t\right\vert _{C}\left\vert \Psi\right\rangle =0$$

or equivalently

$$\frac{\mathrm{d}}{\mathrm{d}t} |\psi(t)\rangle_{\mathcal{S}} = -\frac{i}{\hbar} \hat{\mathcal{H}}_{\mathcal{S}} |\psi(t)\rangle_{\mathcal{S}} \text{ with } a(t) = const$$

Evolution without evolution!

Time-Dilated Schrödinger equation

Multiple Quantum Clocks Non-interacting case



Clocks structure J = A, B, ..., M, N associated with $\mathcal{H}_C = \otimes_J \mathcal{H}_J$

$$\hat{V}_{J}(au) = e^{-rac{i}{\hbar}\hat{\mathcal{H}}_{J} au}$$
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angle_{J} = \hat{V}_{J}(au)|e
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Given the non-interacting clock Hamiltonian $\hat{\mathcal{H}}_{\mathcal{C}} = \sum_J \hat{\mathcal{H}}_J$, we have

$$\hat{V}_{C}(t) = \bigotimes_{J}^{N} \hat{V}_{J}(t) \implies |t\rangle_{C} = \bigotimes_{J}^{N} |t\rangle_{J}$$

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From A's perspective, adopting the time parameter τ we have

$$|\Psi\rangle = \int \mathrm{d}\mu_A(au) \, a_A(au) \, | au
angle_A \, |\psi(au)
angle_{U|A}$$

where

$$|\psi(\tau)\rangle_{U|A} := rac{1}{a_A(au)} \langle au_A|_A |\Psi
angle$$

Conditioning the constraint on A's time, we obtain

$$\langle \tau |_A \hat{\mathcal{H}} | \Psi \rangle = \left(-i\hbar \frac{\mathrm{d}}{\mathrm{d}\tau} + \sum_{J \neq A} \hat{\mathcal{H}}_J + \hat{\mathcal{H}}_S \right) \langle \tau |_A | \Psi \rangle = 0$$

They describe the same dynamics. What about interactions?



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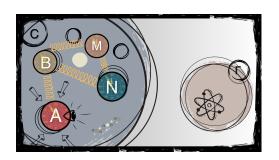
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Introducing an interaction between the local clocks J, K = A, B, ..., M, N

$$\hat{\mathcal{H}}_{C} = \sum_{J} \hat{\mathcal{H}}_{J} + \hat{\mathcal{H}}_{int} \quad \text{with} \quad \hat{\mathcal{H}}_{int} = -\sum_{J < K} g_{JK} \, \hat{\mathcal{H}}_{J} \otimes \hat{\mathcal{H}}_{K}$$



Commutation relation

$$\left[\,\hat{\mathcal{H}}_{J}\,,\,\hat{\mathcal{H}}_{int}\,
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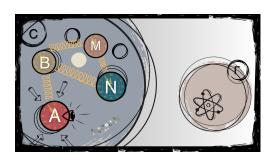
Global-local mapping:

$$|t\rangle_{\mathcal{C}} = e^{-\frac{i}{\hbar}\hat{\mathcal{H}}_{int}t} \otimes_{J}^{N} |t\rangle_{J}$$



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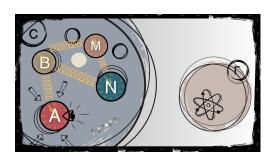
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Time-Dilated Schrödinger equation The redshift operator



From
$$\left\langle au \right|_{A} \hat{\mathcal{H}} \left| \Psi \right\rangle = 0$$
, when $g_{AJ} = 0$ for every sub-clock J

$$i\hbar \frac{\mathrm{d}}{\mathrm{d}\tau} |\psi(\tau)\rangle_{U|A} = \hat{\mathcal{H}}^{(A)} |\psi(\tau)\rangle_{U|A}$$

where
$$\hat{\mathcal{H}}^{(A)} := \hat{\mathcal{H}}_S + \sum_{J \neq A} \hat{\mathcal{H}}_J - \frac{1}{2} \sum_{J,K \neq A} g_{JK} \, \hat{\mathcal{H}}_J \hat{\mathcal{H}}_K$$
.

where \hat{R} is called redshift operator

$$\hat{R}(A) = 1 - \sum_{J \neq A} g_{AJ} \hat{\mathcal{H}}_{J}$$

Time-Dilated Schrödinger equation The redshift operator



From $\langle \tau |_A \hat{\mathcal{H}} | \Psi \rangle = 0$, when $g_{AJ} \neq 0$ for some sub-clock J

$$i\hbar\hat{R}(A)\frac{\mathrm{d}}{\mathrm{d}\tau}|\psi(\tau)\rangle_{U|A}=\hat{\mathcal{H}}^{(A)}|\psi(\tau)\rangle_{U|A}$$

where
$$\hat{\mathcal{H}}^{(A)} := \hat{\mathcal{H}}_S + \sum_{J \neq A} \hat{\mathcal{H}}_J - \frac{1}{2} \sum_{J,K \neq A} g_{JK} \, \hat{\mathcal{H}}_J \hat{\mathcal{H}}_K$$
.

where \hat{R} is called *redshift operator*

$$\hat{R}(A) = 1 - \sum_{J \neq A} g_{AJ} \hat{\mathcal{H}}_J$$



Relative distances r_{JK} , gravitational constant G and speed of light c

$$g_{JK} = rac{G}{r_{JK}c^4}$$
 and $\hat{\Phi}_J(r_{JK}) := -rac{G}{r_{JK}c^2}\hat{\mathcal{H}}_J$.

From A's perspective, we have

$$\begin{split} i\hbar \bigg(\mathbb{1} + \frac{1}{c^2} \hat{\Phi}(A) \bigg) \frac{\mathrm{d}}{\mathrm{d}\tau} |\psi(\tau)\rangle_{U|A} = \\ = \hat{\mathcal{H}}_S + \sum_{J \neq A} \hat{\mathcal{H}}_J + \sum_{J,K \neq A} \frac{1}{2c^2} \hat{\Phi}_J(r_{JK}) \hat{\mathcal{H}}_K |\psi(\tau)\rangle_{U|A} \end{split}$$

where $\hat{\Phi}(A) := \sum_{J \neq A} \hat{\Phi}_J(r_{AJ})$. Well-suited for a semi-classical description, but ...



Relative distances r_{JK} , gravitational constant G and speed of light c

$$g_{JK} = \frac{G}{r_{JK} c^4} \quad \text{and} \quad \hat{\Phi}_J(r_{JK}) := -\frac{G}{r_{JK} c^2} \hat{\mathcal{H}}_J \,.$$

From A's perspective, we have

$$\begin{split} i\hbar \bigg(\mathbb{1} + \frac{1}{c^2} \hat{\Phi}(A) \bigg) \frac{\mathrm{d}}{\mathrm{d}\tau} |\psi(\tau)\rangle_{U|A} = \\ &= \hat{\mathcal{H}}_S + \sum_{J \neq A} \hat{\mathcal{H}}_J + \sum_{J,K \neq A} \frac{1}{2c^2} \hat{\Phi}_J(r_{JK}) \hat{\mathcal{H}}_K |\psi(\tau)\rangle_{U|A} \end{split}$$

where
$$\hat{\Phi}(A) := \sum_{J \neq A} \hat{\Phi}_J(r_{AJ})$$
.

Well-suited for a semi-classical description, but ...



Relative distances r_{JK} , gravitational constant G and speed of light c

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From A's perspective, we have

$$\begin{split} i\hbar\bigg(1+\frac{1}{c^2}\Phi(A)\bigg)\frac{\mathrm{d}}{\mathrm{d}\tau}\,|\psi(\tau)\rangle_{U|A} &= \\ &= M_Ec^2+\hat{\mathcal{H}}_S+\sum_{J\neq A,E}\hat{\mathcal{H}}_J+\sum_{J\neq A,E}\frac{1}{c^2}\Phi(J)\hat{\mathcal{H}}_K\;|\psi(\tau)\rangle_{U|A} \end{split}$$

where $\Phi(J) := -GM_E/r_{JE}$. Then $d\tau = \sqrt{g_{00}} dt < dt!$ Well-suited for a semi-classical description, but ...



Expressing the redshift operator as

$$\hat{R}(A) = \mathbb{1} - \hat{\Phi}(A),$$

if the spectral radius $\rho(\hat{\Phi}(A)) < 1$, we get a geometric series:

$$\hat{R}^{-1}(A) = \sum_{n=0}^{+\infty} \hat{\Phi}^n(A)$$
.

Multiplying by $\hat{R}^{-1}(A)$ from the left, we obtain

$$i\hbar \frac{\mathrm{d}}{\mathrm{d}\tau} |\psi(\tau)\rangle_{U|A} = \sum_{n=0}^{+\infty} \hat{\Phi}^n(A) \hat{\mathcal{H}}^{(A)} |\psi(\tau)\rangle_{U|A}$$

Time-Dilation induced Interaction Transfer (TiDIT) mechanism



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Time-Dilation induced Interaction Transfer (TiDIT) mechanism



Quantum time dilation \xrightarrow{TiDIT} new interaction terms!

$$i\hbar\frac{\mathrm{d}}{\mathrm{d}\tau}\left|\psi(\tau)\right\rangle_{U|A}=\hat{\mathcal{H}}_{\mathrm{eff}}^{\mathrm{(A)}}\left|\psi(\tau)\right\rangle_{U|A}$$

At -order in g_{AJ}

$$\hat{\mathcal{H}}_{eff}^{(A)} := \hat{\mathcal{H}}_S + \sum_{f \in A} g_{Af} \mathcal{H}_f \mathcal{H}_S +$$

$$+\sum_{J
eq A}\hat{\mathcal{H}}_J-rac{1}{2}\sum_{J,K
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$$i\hbar\frac{\mathrm{d}}{\mathrm{d}\tau}\left|\psi(\tau)\right\rangle_{\textit{U}\left|\textit{A}}=\hat{\mathcal{H}}_{\textit{eff}}^{\textit{(A)}}\left|\psi(\tau)\right\rangle_{\textit{U}\left|\textit{A}}$$

At zeroth-order in g_{AJ}

$$\begin{split} \hat{\mathcal{H}}_{\text{eff}}^{(A)} &:= \hat{\mathcal{H}}_S + \sum_{J \neq A} g_{AJ} \hat{\mathcal{H}}_J \hat{\mathcal{H}}_S + \sum_{J,K \neq A} g_{AJ} g_{AK} \hat{\mathcal{H}}_J \hat{\mathcal{H}}_K \hat{\mathcal{H}}_S + \\ &+ \sum_{J \neq A} \hat{\mathcal{H}}_J - \frac{1}{2} \sum_{J,K \neq A} (g_{JK}) \, \hat{\mathcal{H}}_J \hat{\mathcal{H}}_K \end{split}$$



Quantum time dilation \xrightarrow{TiDIT} new interaction terms!

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At first-order in g_{AJ}

$$\begin{split} \hat{\mathcal{H}}_{\text{eff}}^{(A)} &:= \hat{\mathcal{H}}_S + \sum_{J \neq A} g_{AJ} \hat{\mathcal{H}}_J \hat{\mathcal{H}}_S + \sum_{J,K \neq A} g_{AJ} g_{AK} \hat{\mathcal{H}}_J \hat{\mathcal{H}}_K \hat{\mathcal{H}}_S + \\ &+ \sum_{J \neq A} \hat{\mathcal{H}}_J - \frac{1}{2} \sum_{J,K \neq A} \left(g_{JK} - 2 g_{AK} \right) \hat{\mathcal{H}}_J \hat{\mathcal{H}}_K \end{split}$$



Quantum time dilation \xrightarrow{TiDIT} new interaction terms!

$$i\hbar\frac{\mathrm{d}}{\mathrm{d}\tau}\left|\psi(\tau)\right\rangle_{\textit{U}\left|\textit{A}}=\hat{\mathcal{H}}_{\textit{eff}}^{\textit{(A)}}\left|\psi(\tau)\right\rangle_{\textit{U}\left|\textit{A}}$$

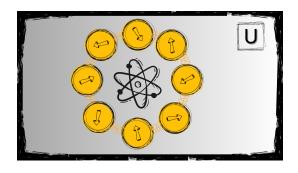
At second-order in g_{AJ}

$$\begin{split} \hat{\mathcal{H}}_{\text{eff}}^{(A)} &:= \hat{\mathcal{H}}_{S} + \sum_{J \neq A} g_{AJ} \hat{\mathcal{H}}_{J} \hat{\mathcal{H}}_{S} + \sum_{J,K \neq A} g_{AJ} g_{AK} \hat{\mathcal{H}}_{J} \hat{\mathcal{H}}_{K} \hat{\mathcal{H}}_{S} + \\ &+ \sum_{J \neq A} \hat{\mathcal{H}}_{J} - \frac{1}{2} \sum_{J,K \neq A} \left(g_{JK} - 2 g_{AK} \right) \left(\mathbb{1} + \sum_{M \neq A} g_{AM} \hat{\mathcal{H}}_{M} \right) \hat{\mathcal{H}}_{J} \hat{\mathcal{H}}_{K} \end{split}$$





$$\hat{\mathcal{H}} = \hat{\mathcal{H}}_{S} + \sum_{J} h_{J} \hat{\sigma}_{J}^{x} + \sum_{\langle J, K \rangle} g_{JK} \hat{\sigma}_{J}^{x} \hat{\sigma}_{K}^{x}$$





$$\hat{\mathcal{H}}_A = \hbar w \, \hat{\sigma}_A^{\scriptscriptstyle X} \quad \Longrightarrow \quad \hat{V}_A(\tau) = e^{-iw\tau\hat{\sigma}_A^{\scriptscriptstyle X}} \, .$$

Time states are obtained as

$$|e\rangle_A = \frac{1}{\sqrt{2}}(|+\rangle + |-\rangle) \quad \Longrightarrow \quad |\tau\rangle_A = \hat{V}_A(\tau) |e\rangle_A$$

Orthogonal with $\tau_0 = 0$ and $\tau_1 = \pi/2w$

$$|\tau_0\rangle_A = \frac{1}{\sqrt{2}}(|+\rangle + |-\rangle)$$
 and $|\tau_1\rangle_A = \frac{-i}{\sqrt{2}}(|+\rangle - |-\rangle)$

$$\mathbb{1}_A = \sum_{n=0,1} |\tau_n\rangle\langle\tau_n|_A = \frac{2w}{\pi} \int_0^{\frac{\pi}{w}} d\tau |\tau\rangle\langle\tau|_A$$



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Given $|\Psi\rangle\in\mathcal{H}$ such that $\hat{\mathcal{H}}\,|\Psi\rangle=0$, we have

$$\begin{aligned} |\Psi\rangle &= \frac{2w}{\pi} \int_0^{\frac{\pi}{w}} \mathrm{d}\tau \, a(\tau) \, |\tau\rangle_A |\psi(\tau)\rangle_{U|A} \\ &= a_0 \, |\tau_0\rangle_A \, |\psi(\tau_0)\rangle_{U|A} + a_1 \, |\tau_1\rangle_A \, |\psi(\tau_1)\rangle_{U|A} \end{aligned}$$

where

$$a(\tau) = \sqrt{a_0^2 \cos^2(w\tau) + a_1^2 \sin^2(w\tau)}$$

Notice

$$\hat{\mathcal{H}} = \hat{\mathcal{H}}_A + \hat{\mathcal{H}}_{U|A} \implies a(\tau) = a_1 = a_0 = 1/\sqrt{2}$$



Given $|\Psi
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A two-qubit clock model

$$\hat{\mathcal{H}} = \hbar w \left(\hat{\sigma}_A^{\mathsf{x}} + \alpha \hat{\sigma}_B^{\mathsf{x}} \right) + \hat{\mathcal{H}}_{\mathsf{S}}$$

The Schrödinger equation reads

$$i\hbar \frac{\mathrm{d}}{\mathrm{d}\tau} |\psi(\tau)\rangle_{U|A} = \left(\alpha\hbar w \hat{\sigma}_B^{\mathsf{x}} + \hat{\mathcal{H}}_{\mathsf{S}}\right) |\psi(\tau)\rangle_{U|A}$$

where

$$\hat{R} = 1 - g\hat{\sigma}_B^{\mathsf{x}}$$



A two-qubit clock model

$$\hat{\mathcal{H}} = \hbar w \left(\hat{\sigma}_{A}^{x} + \alpha \hat{\sigma}_{B}^{x} - g \hat{\sigma}_{A}^{x} \hat{\sigma}_{B}^{x} \right) + \hat{\mathcal{H}}_{S}$$

The Schrödinger equation reads

$$i\hbar\hat{R}\frac{\mathrm{d}}{\mathrm{d}\tau}|\psi(\tau)\rangle_{U|A} = \left(\alpha\hbar w\hat{\sigma}_{B}^{x} + \hat{\mathcal{H}}_{S}\right)|\psi(\tau)\rangle_{U|A}$$

where

$$\hat{R} = \mathbb{1} - g\hat{\sigma}_B^{\mathsf{x}}$$



The redshift operator takes the form

$$\hat{R} = \mathbb{1} - g\hat{\sigma}_B^{\times} \implies \lambda = \begin{cases} 1 + g \text{ on } |-\rangle_B \\ 1 - g \text{ on } |+\rangle_B \end{cases}$$

When $g \neq 1$, i.e. \hat{R} is invertible, we observe two new effects..

$$\begin{split} i\hbar \big(1-g^2\big)\frac{\mathrm{d}}{\mathrm{d}\tau}\,|\psi(\tau)\rangle_{U|A} = \\ &= \Big(\alpha\hbar wg + \alpha\hbar w\hat{\sigma}_B^* - \hat{\mathcal{H}}_S + g\hat{\sigma}_B^*\hat{\mathcal{H}}_S\Big)\,|\psi(\tau)\rangle_{U|A} \end{split}$$



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$$\begin{split} i\hbar \big(1-g^2\big) \frac{\mathrm{d}}{\mathrm{d}\tau} \, |\psi(\tau)\rangle_{U|A} &= \\ &= \Big(\alpha\hbar w g + \alpha\hbar w \hat{\sigma}_B^\mathsf{x} + \hat{\mathcal{H}}_S + g\hat{\sigma}_B^\mathsf{x} \,\hat{\mathcal{H}}_S\Big) \, |\psi(\tau)\rangle_{U|A} \end{split}$$

Time-Dilation!



The redshift operator takes the form

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$$\begin{split} i\hbar \big(1-g^2\big) \frac{\mathrm{d}}{\mathrm{d}\tau} \, |\psi(\tau)\rangle_{U|A} &= \\ &= \Big(\alpha\hbar wg + \alpha\hbar w\hat{\sigma}_B^x + \hat{\mathcal{H}}_S + g\hat{\sigma}_B^x \,\hat{\mathcal{H}}_S\Big) \, |\psi(\tau)\rangle_{U|A} \end{split}$$

Interaction Transfer!



The redshift operator takes the form

$$\hat{R} = \mathbb{1} - g\hat{\sigma}_B^{\mathsf{x}} \implies \lambda = \begin{cases} 1 + g \text{ on } |-\rangle_B \\ 1 - g \text{ on } |+\rangle_B \end{cases}$$

When $g \neq 1$, i.e. \hat{R} is invertible, we have $d\tau = (1 - g^2)dt$

$$\begin{split} i\hbar \big(1-g^2\big)\frac{\mathrm{d}}{\mathrm{d}\tau}\,|\psi(\tau)\rangle_{U|A} &= \\ &= \left(\alpha\hbar wg + \alpha\hbar w\hat{\sigma}_B^\mathsf{x} + \hat{\mathcal{H}}_S + g\hat{\sigma}_B^\mathsf{x}\,\hat{\mathcal{H}}_S\right)\,|\psi(\tau)\rangle_{U|A} \end{split}$$

Time-Dilation induced Interaction Transfer (TiDIT) mechanism



Quantum time dilation \xrightarrow{TiDIT} new interaction terms!

$$\begin{split} i\hbar\frac{\mathrm{d}}{\mathrm{d}\tau}\,|\psi(\tau)\rangle_{U|A} &= \hat{\mathcal{H}}_{\mathrm{eff}}^{(A)}\,|\psi(\tau)\rangle_{U|A} \\ \hat{\mathcal{H}}_{\mathrm{eff}}^{(A)} &= \frac{1}{1-g^2}\left(\alpha\hbar wg + \alpha\hbar w\hat{\sigma}_B^{\mathrm{x}} + \hat{\mathcal{H}}_S + g\hat{\sigma}_B^{\mathrm{x}}\,\hat{\mathcal{H}}_S\right) \end{split}$$

When g = 1 the history state becomes

$$\begin{aligned} |\Psi\rangle &= \int \mathrm{d}\mu(\tau) \, a_{-}(\tau) \, |\tau\rangle_{A} \, |-\rangle_{B} \, |\phi_{-}(\tau)\rangle_{S} \, + \\ &+ \left(\int \mathrm{d}\mu(\tau) \, a_{+}(\tau) \, |\tau\rangle_{A} \right) |+\rangle_{B} \, |-\alpha\hbar w\rangle_{S} \, \end{aligned}$$



Quantum time dilation \xrightarrow{TiDIT} new interaction terms!

$$\begin{split} i\hbar\frac{\mathrm{d}}{\mathrm{d}\tau} \left| \psi(\tau) \right\rangle_{U|A} &= \hat{\mathcal{H}}_{\mathrm{eff}}^{(A)} \left| \psi(\tau) \right\rangle_{U|A} \\ \hat{\mathcal{H}}_{\mathrm{eff}}^{(A)} &= \frac{1}{1-g^2} \left(\alpha\hbar w g + \alpha\hbar w \hat{\sigma}_B^{\mathrm{X}} + \hat{\mathcal{H}}_S + g \hat{\sigma}_B^{\mathrm{X}} \hat{\mathcal{H}}_S \right) \end{split}$$

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Qubit Hypersurfaces The TiDIT mechanism

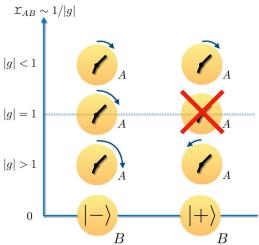


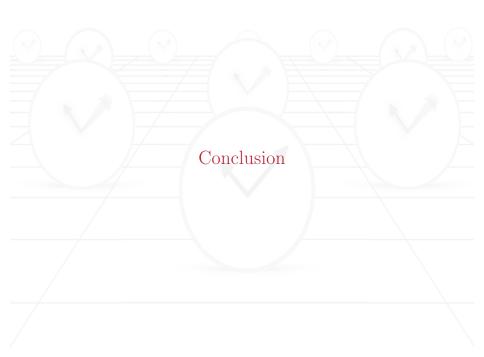
Effective Hamiltonian diverges

↓

The dynamical description is frozen

"at the horizon"





A history can be associated with a laboratory quantum system.

Gravitational-like interaction introduces a redshift operator.

The spin-spin coupling gives quantum/time dilation and TILIT.

erspectives:

Generalization to more complex interaction terms.

Generalization to dynamical spatial degrees of freedom

Explore the implications of the TiD IT mechanism in gravity.

(1.) A history can be associated with a laboratory quantum system.

Gravitational-like interaction introduces a redshift operator.

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Perspectives

Generalization to make complex interaction terms.

Generalization to dynamical spatial degrees of freedom

Explore the implications of the Tip I mechanism in gravity.

- (1.) A history can be associated with a laboratory quantum system.
- (2.) Gravitational-like interaction introduces a redshift operator.
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Perspectives:

- Generalization to more complex interaction terms.
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Perspectives:

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Perspectives:

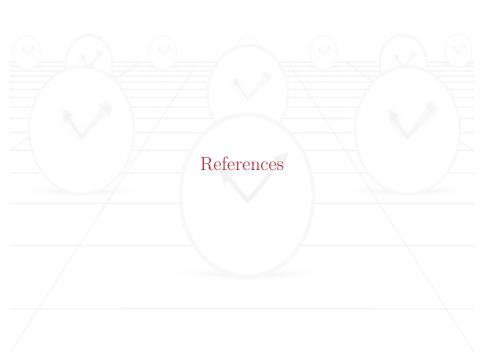
- (?) Generalization to more complex interaction terms.
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- (?) Generalization to more complex interaction terms.
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Thank you for your (classical) time!



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