# prior\_montecarlo

May 3, 2019

## 1 Run and process the prior monte carlo and pick a "truth" realization

```
In [1]: import os
    import shutil
    import numpy as np
    import pandas as pd
    import flopy
    import pyemu

flopy is installed in /Users/jeremyw/Dev/gw1876/activities_2day_mfm/notebooks/flopy

In [2]: t_d = "template"
    pst = pyemu.Pst(os.path.join(t_d,"freyberg.pst"))
```

### 1.0.1 Decide what pars are uncertain in the truth

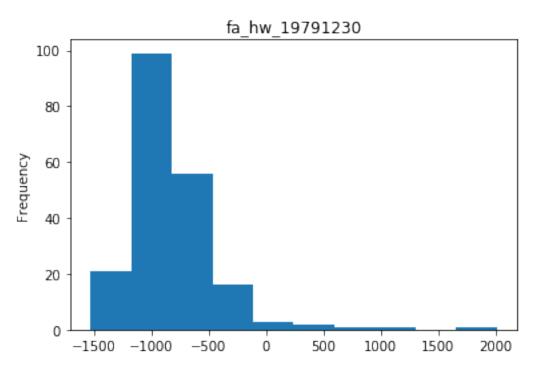
We need to decide what our truth looks like - should the pilot points or the grid-scale pars be the source of spatial variability? or both?

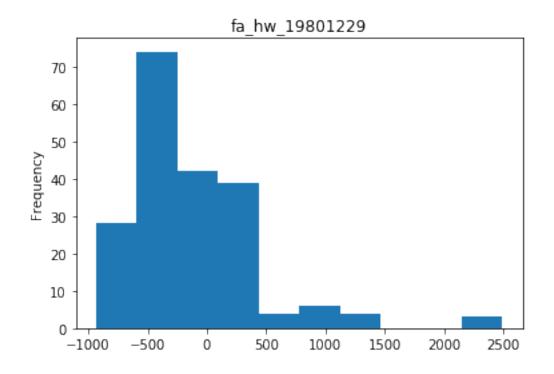
```
In [3]: par = pst.parameter_data
    # grid pars
    #should_fix = par.loc[par.pargp.apply(lambda x: "gr" in x), "parnme"]
    # pp pars
    #should_fix = par.loc[par.pargp.apply(lambda x: "pp" in x), "parnme"]
    #pst.npar - should_fix.shape[0]
In [4]: pe = pyemu.ParameterEnsemble.from_binary(pst=pst,filename=os.path.join(t_d,"prior.jcb"
    #pe.loc[:,should_fix] = 1.0
    pe.to_csv(os.path.join(t_d,"sweep_in.csv"))
```

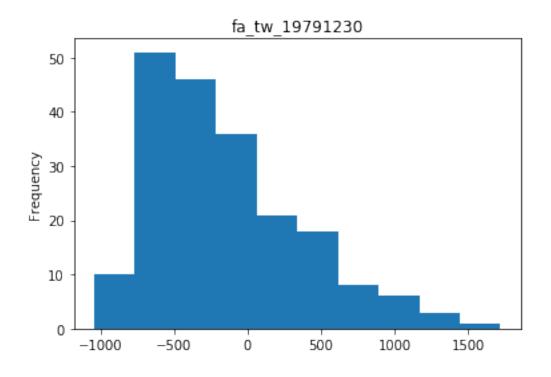
### 1.0.2 run the prior ensemble in parallel locally

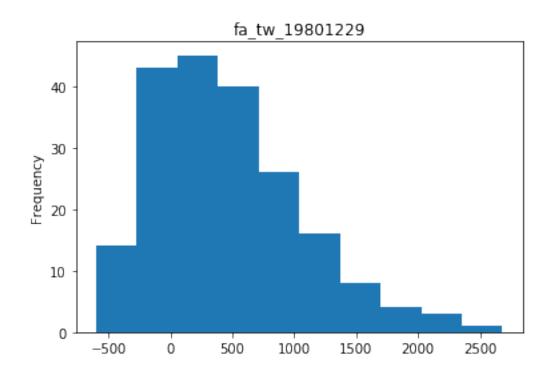
### 1.0.3 Load the output ensemble and plot a few things

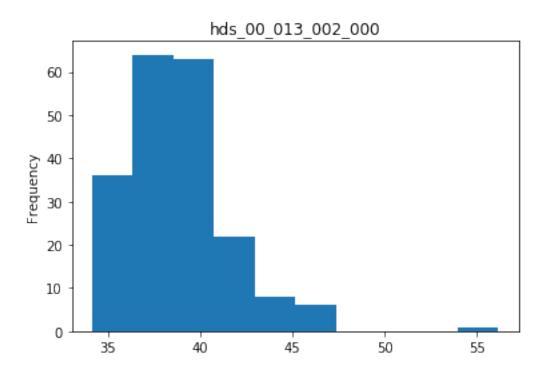
```
In [6]: obs_df = pd.read_csv(os.path.join(m_d, "sweep_out.csv"),index_col=0)
        obs_df.shape
Out[6]: (200, 4465)
  drop any failed runs
In [7]: obs_df = obs_df.loc[obs_df.failed_flag==0,:]
        obs_df.shape
Out[7]: (200, 4465)
In [8]: fnames = pst.pestpp_options["forecasts"].split(',')
        fnames
Out[8]: ['fa_hw_19791230',
         'fa_hw_19801229',
         'fa_tw_19791230',
         'fa_tw_19801229',
         'hds_00_013_002_000',
         'hds_00_013_002_001',
         'part_time',
         'part_status']
In [9]: for forecast in fnames:
            ax = obs_df.loc[:,forecast].plot(kind="hist")
            ax.set_title(forecast)
            plt.show()
```

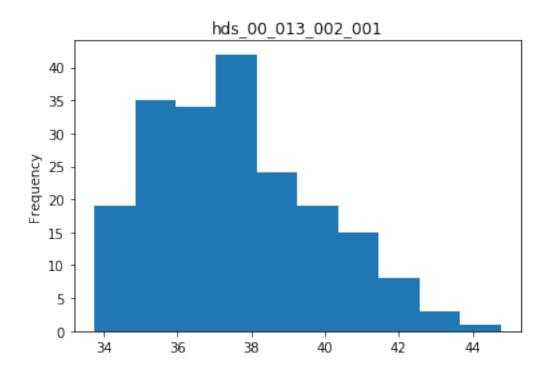


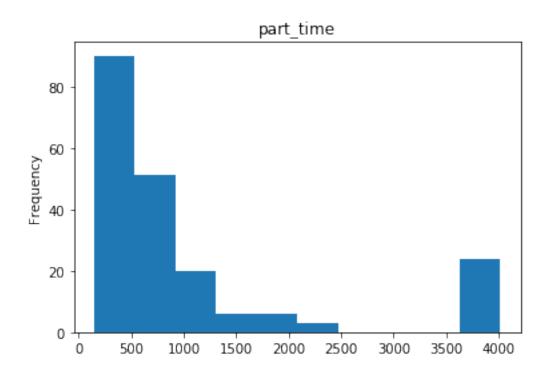


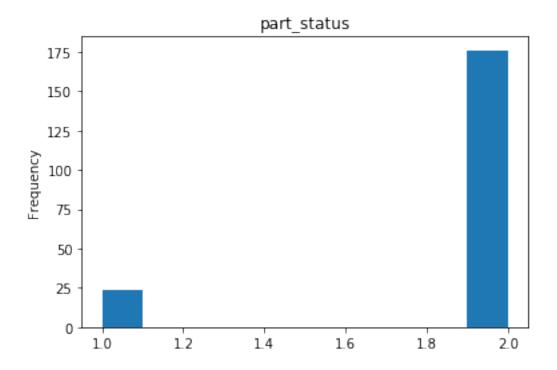




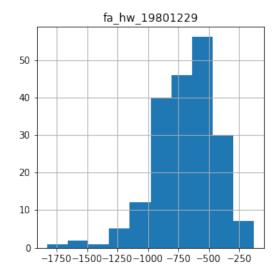


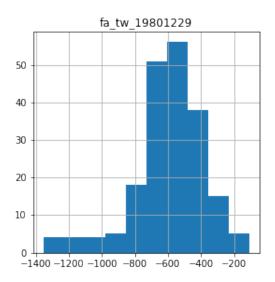


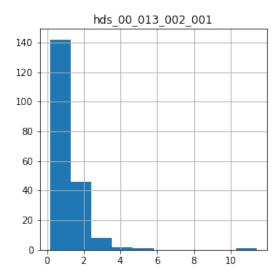




We see that under scenario conditions, many more realizations for the flow to the aquifer in the headwaters are postive (as expected). Lets difference these two:







We now see that the most extreme scenario yields a large decrease in flow from the aquifer to the headwaters (the most negative value)

### 1.0.4 setting the "truth"

We just need to replace the observed values (obsval) in the control file with the outputs for one of the realizations on obs\_df. In this way, we now have the nonzero values for history matching, but also the truth values for comparing how we are doing with other unobserved quantities. Im going to pick a realization that yields an "average" variability of the observed gw levels:

### Out[11]: 23 In [12]: obs\_df.loc[idx,pst.nnz\_obs\_names] Out[12]: fo\_39\_19791230 11553.000000 hds\_00\_002\_009\_000 36.390945 hds 00 002 015 000 34.901199 hds 00 003 008 000 36.633041 hds\_00\_009\_001\_000 38.894932 hds\_00\_013\_010\_000 35.846367 hds\_00\_015\_016\_000 34.909328 hds\_00\_021\_010\_000 35.402809 hds\_00\_022\_015\_000 34.446541 hds\_00\_024\_004\_000 36.218941 hds\_00\_026\_006\_000 35.067863 hds\_00\_029\_015\_000 34.225922 hds\_00\_033\_007\_000 34.120350 hds\_00\_034\_010\_000 33.842617 Name: 23, dtype: float64

Lets see how our selected truth does with the swgw forecasts:

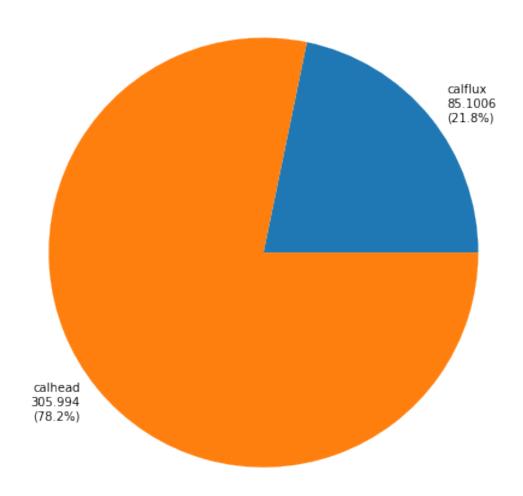
```
Out[13]: fa_hw_19791230
                              -1250.952650
         fa_hw_19801229
                               -383.912850
         fa_tw_19791230
                                  97.419300
         fa tw 19801229
                                841.068200
         hds_00_013_002_000
                                  39.332623
         hds_00_013_002_001
                                  37.060593
         part_time
                                1046.316000
         part_status
                                   2.000000
         Name: 23, dtype: float64
```

In [13]: obs\_df.loc[idx,fnames]

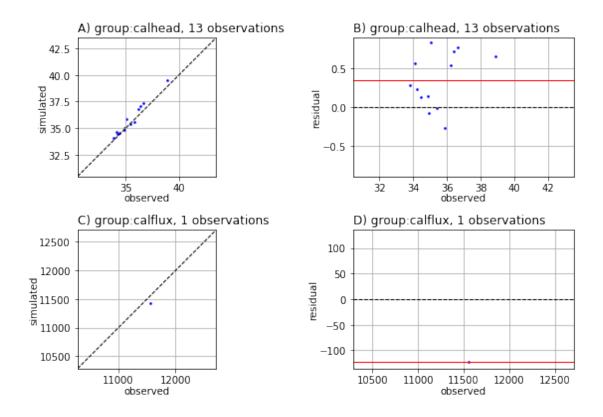
Assign some initial weights. Now, it is custom to add noise to the observed values...we will use the classic Gaussian noise...zero mean and standard deviation of 1 over the weight

```
In [14]: pst = pyemu.Pst(os.path.join(t_d,"freyberg.pst"))
    obs = pst.observation_data
    obs.loc[:,"obsval"] = obs_df.loc[idx,pst.obs_names]
    obs.loc[obs.obgnme=="calhead","weight"] = 10.0
    obs.loc[obs.obgnme=="calflux","weight"] = 0.075
    np.random.seed(seed=0)
    snd = np.random.randn(pst.nnz_obs)
    noise = snd * 1./obs.loc[pst.nnz_obs_names,"weight"]
    #pst.observation_data.loc[noise.index,"obsval"] += noise
    pst.write(os.path.join(t_d,"freyberg.pst"))
    pyemu.os_utils.run("pestpp-ies freyberg.pst",cwd=t_d)
```

### 391.0941587400599



<Figure size 576x756 with 0 Axes>



Publication ready figs - oh snap!

Just to make sure we have everything working right, we should be able to load the truth parameters, run the model once and have a phi equivalent to the noise vector:

#### Out [16]: modelled group measured name name calflux 11553.000000 fo\_39\_19791230 fo\_39\_19791230 11553.000000 hds\_00\_002\_009\_000 hds\_00\_002\_009\_000 calhead 36.390945 36.390945 calhead hds\_00\_002\_015\_000 hds\_00\_002\_015\_000 34.901199 34.901199 hds\_00\_003\_008\_000 hds 00 003 008 000 calhead 36.633041 36.633041 hds 00 009 001 000 hds 00 009 001 000 calhead 38.894932 38.894932 hds\_00\_013\_010\_000 hds\_00\_013\_010\_000 calhead 35.846367 35.846367 calhead hds\_00\_015\_016\_000 hds\_00\_015\_016\_000 34.909328 34.909328 hds\_00\_021\_010\_000 hds\_00\_021\_010\_000 calhead 35.402809 35.402809 hds\_00\_022\_015\_000 hds\_00\_022\_015\_000 calhead 34.446541 34.446541 hds\_00\_024\_004\_000 calhead hds\_00\_024\_004\_000 36.218941 36.218941 hds\_00\_026\_006\_000 hds\_00\_026\_006\_000 calhead 35.067863 35.067863 calhead hds\_00\_029\_015\_000 hds\_00\_029\_015\_000 34.225922 34.225922 hds\_00\_033\_007\_000 hds\_00\_033\_007\_000 calhead 34.120350 34.120350 calhead hds\_00\_034\_010\_000 hds\_00\_034\_010\_000 33.842617 33.842617 residual weight name fo\_39\_19791230 0.000000e+00 0.075 hds\_00\_002\_009\_000 4.296865e-10 10.000 hds 00 002 015 000 1.796892e-10 10.000 hds\_00\_003\_008\_000 1.640643e-10 10.000 hds\_00\_009\_001\_000 -2.128928e-10 10.000 hds\_00\_013\_010\_000 -3.242207e-10 10.000 hds\_00\_015\_016\_000 3.066418e-10 10.000 hds\_00\_021\_010\_000 -6.640732e-11 10.000 hds\_00\_022\_015\_000 4.804690e-10 10.000 hds\_00\_024\_004\_000 1.367155e-10 10.000 hds\_00\_026\_006\_000 -3.554703e-10 10.000 hds\_00\_029\_015\_000 1.406235e-10 10.000

The residual should be exactly the noise values from above. Lets load the model (that was just run using the true pars) and check some things

10.000

10.000

```
In [17]: m = flopy.modflow.Modflow.load("freyberg.nam",model_ws=m_d)
```

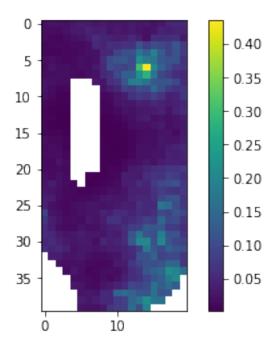
hds\_00\_033\_007\_000 -3.320366e-11

hds\_00\_034\_010\_000 8.789414e-11

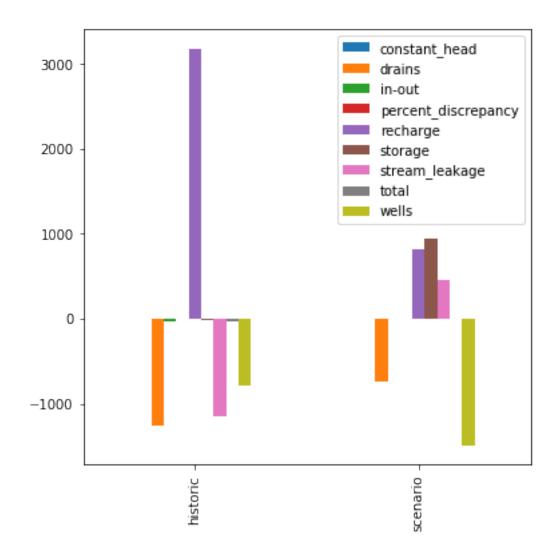
```
In [18]: a = m.upw.vka[1].array
    #a = m.rch.rech[0].array
    a = np.ma.masked_where(m.bas6.ibound[0].array==0,a)
    print(a.min(),a.max())
    c = plt.imshow(a)
    plt.colorbar(c)
```

### 0.002020436 0.4354167

Out[18]: <matplotlib.colorbar.Colorbar at 0x1819bcd208>



```
In [19]: lst = flopy.utils.MfListBudget(os.path.join(m_d,"freyberg.list"))
    df = lst.get_dataframes(diff=True)[0]
    ax = df.plot(kind="bar",figsize=(6,6))
    a = ax.set_xticklabels(["historic","scenario"],rotation=90)
```

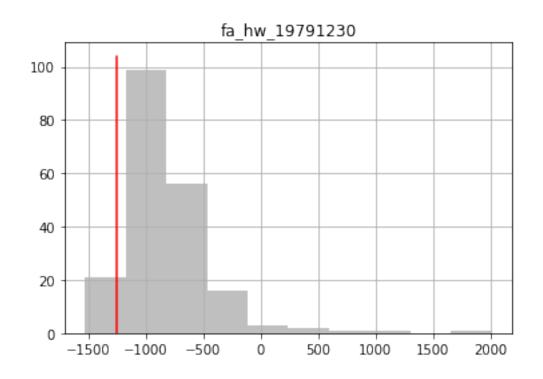


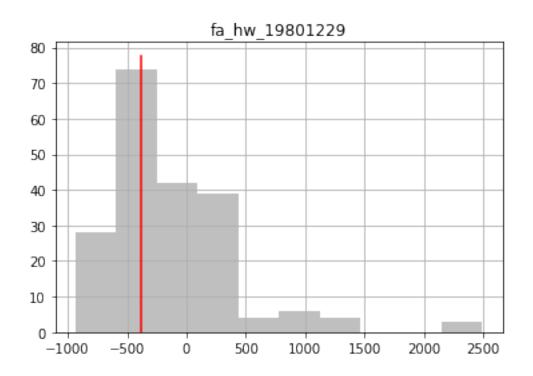
### 1.0.5 see how our existing observation ensemble compares to the truth

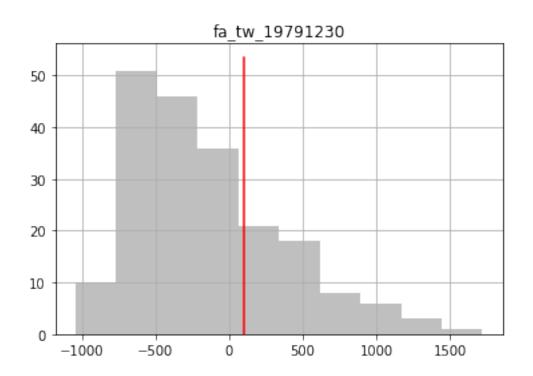
sw-gw outputs:

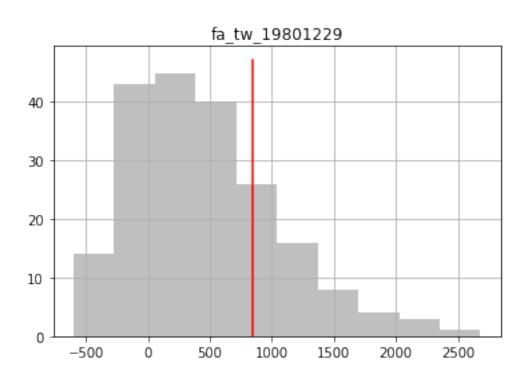
```
In [20]: obs = pst.observation_data

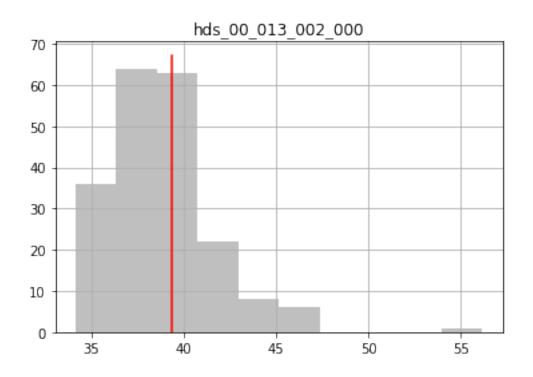
for forecast in fnames:
    ax = plt.subplot(111)
    obs_df.loc[:,forecast].hist(ax=ax,color="0.5",alpha=0.5)
    ax.plot([obs.loc[forecast,"obsval"],obs.loc[forecast,"obsval"]],ax.get_ylim(),"r",
    ax.set_title(forecast)
    plt.show()
```

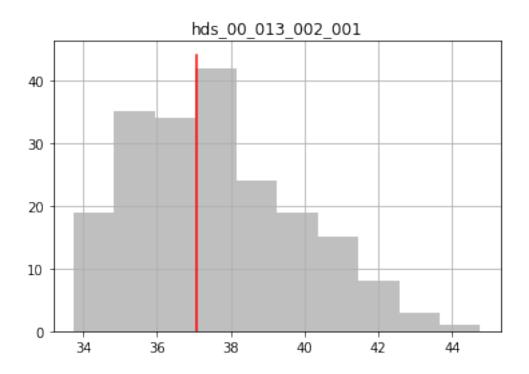


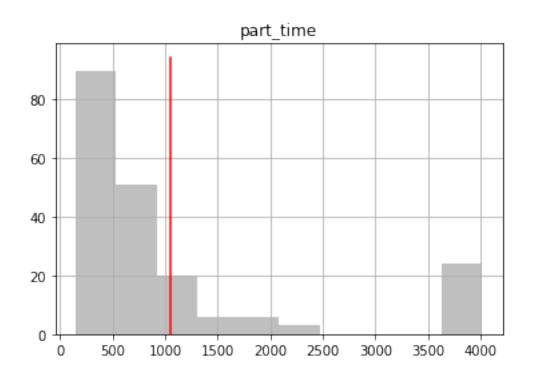


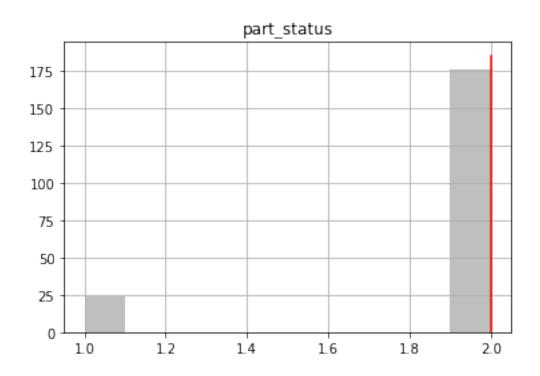












observations:

