

prior_montecarlo

May 2, 2019

1 Run and process the prior monte carlo and pick a “truth” realization

```
In [1]: import os
import shutil
import numpy as np
import pandas as pd
import matplotlib.pyplot as plt
import flogy
import pyemu
```

flogy is installed in /Users/jeremyw/Dev/gw1876/activities_2day_mfm/notebooks/flogy

```
In [2]: t_d = "template"
pst = pyemu.Pst(os.path.join(t_d, "freyberg.pst"))
```

1.0.1 Decide what pars are uncertain in the truth

We need to decide what our truth looks like - should the pilot points or the grid-scale pars be the source of spatial variability? or both?

```
In [3]: par = pst.parameter_data
# grid pars
#should_fix = par.loc[par.pargp.apply(lambda x: "gr" in x), "parname"]
# pp pars
#should_fix = par.loc[par.pargp.apply(lambda x: "pp" in x), "parname"]
#pst.npar - should_fix.shape[0]
```

```
In [4]: pe = pyemu.ParameterEnsemble.from_binary(pst=pst, filename=os.path.join(t_d, "prior.jcb"))
#pe.loc[:, should_fix] = 1.0
pe.to_csv(os.path.join(t_d, "sweep_in.csv"))
```

new binary format detected...

1.0.2 run the prior ensemble in parallel locally

```
In [5]: m_d = "master_prior_sweep"
#pyemu.os_utils.start_slaves(t_d, "pestpp-swp", "freyberg.pst", num_slaves=20, slave_root=
```

1.0.3 Load the output ensemble and plot a few things

```
In [6]: obs_df = pd.read_csv(os.path.join(m_d, "sweep_out.csv"), index_col=0)
        obs_df.shape
```

```
Out[6]: (200, 4465)
```

drop any failed runs

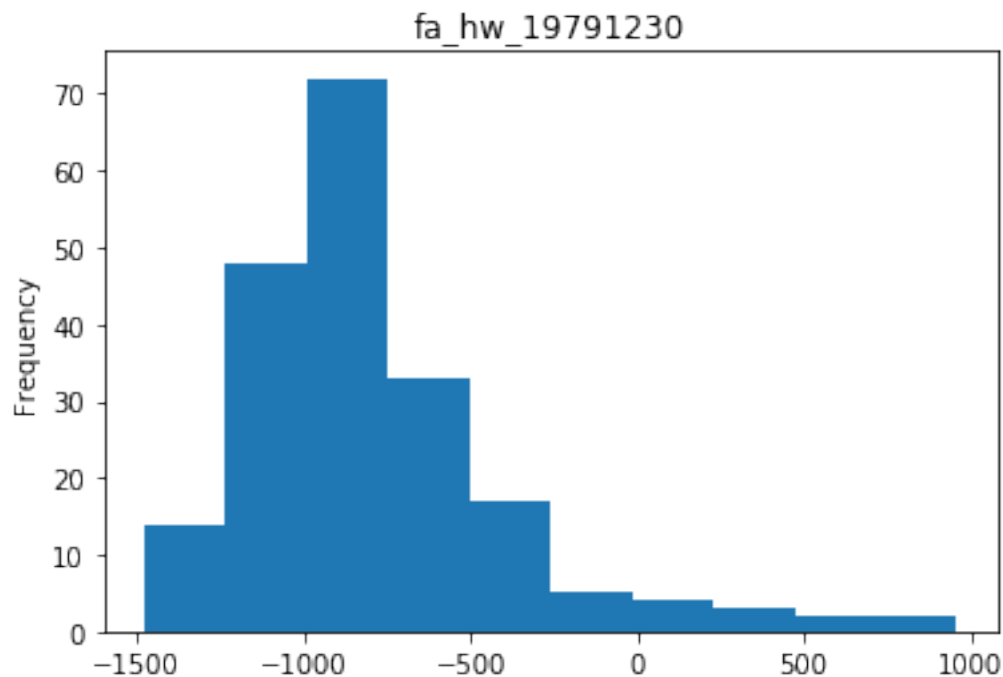
```
In [7]: obs_df = obs_df.loc[obs_df.failed_flag==0,:]
        obs_df.shape
```

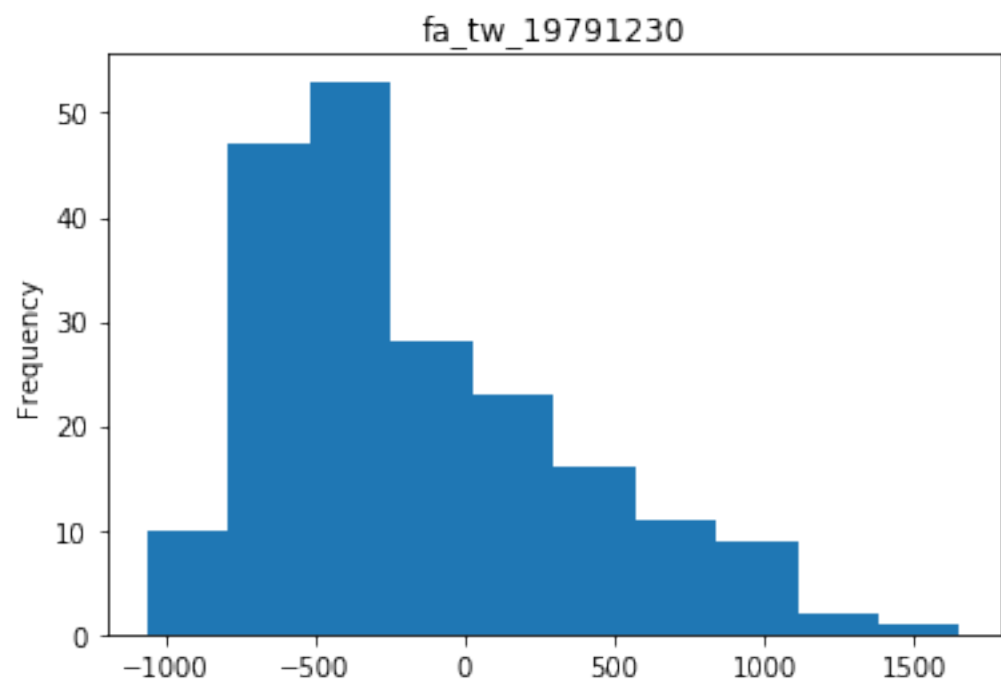
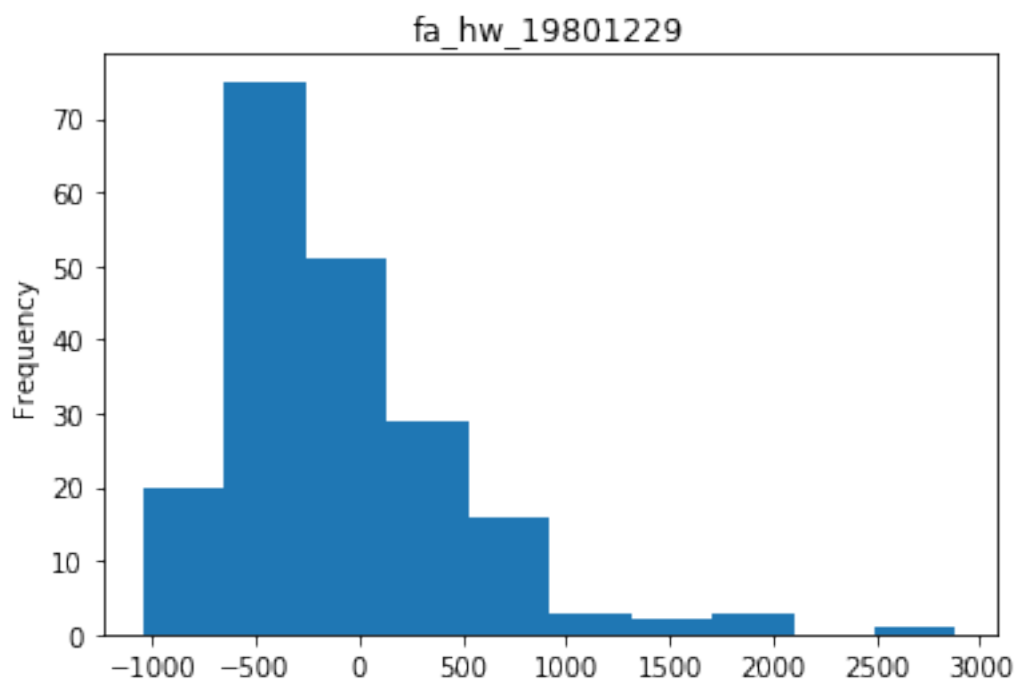
```
Out[7]: (200, 4465)
```

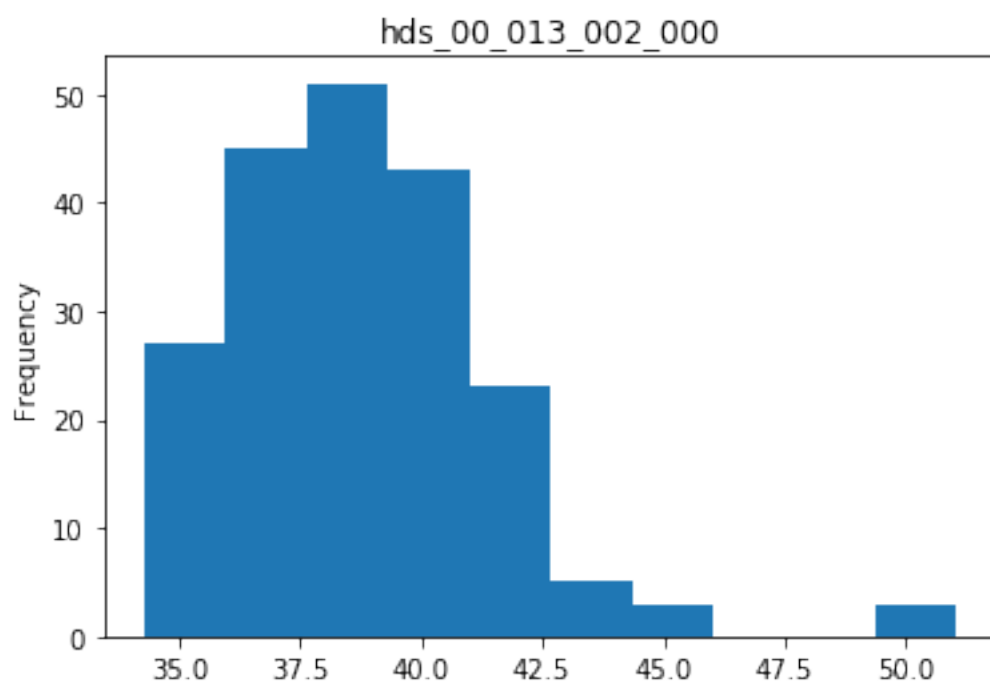
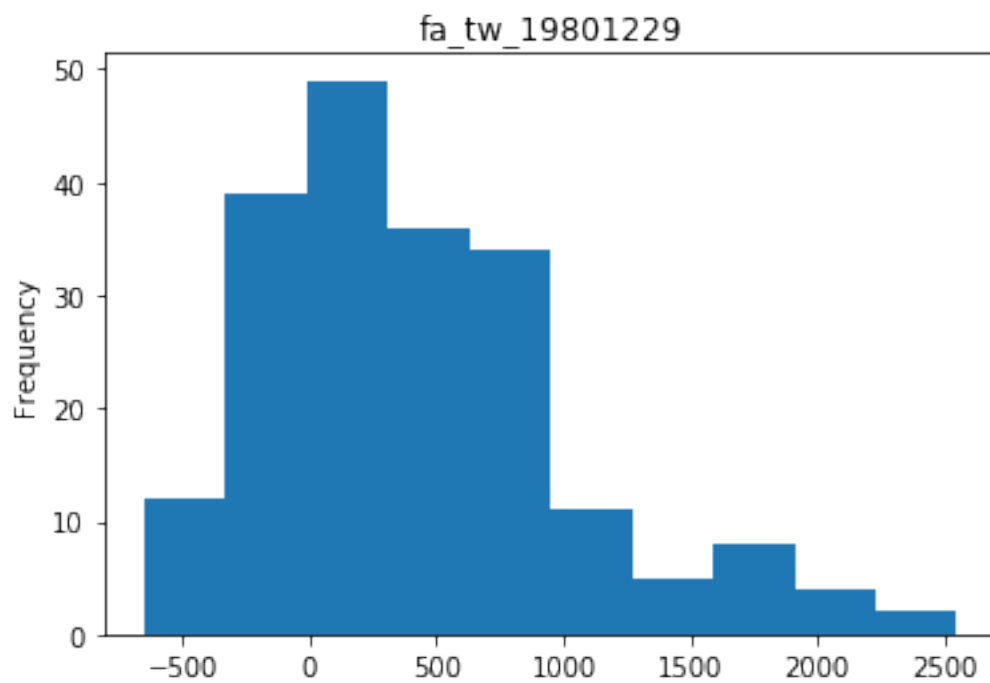
```
In [8]: fnames = pst.pestpp_options["forecasts"].split(',')
        fnames
```

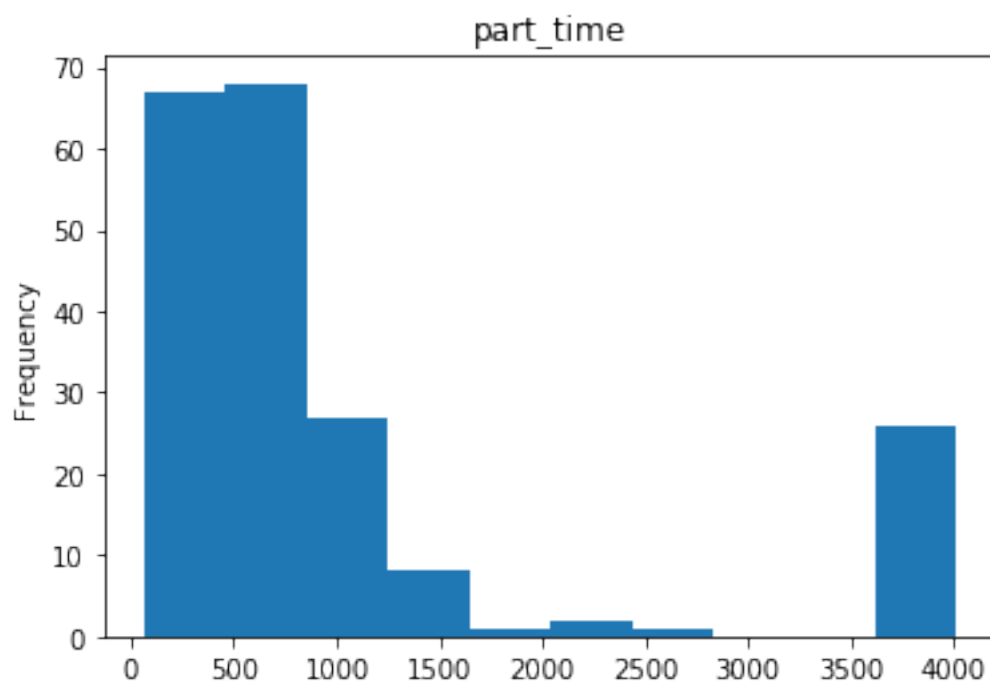
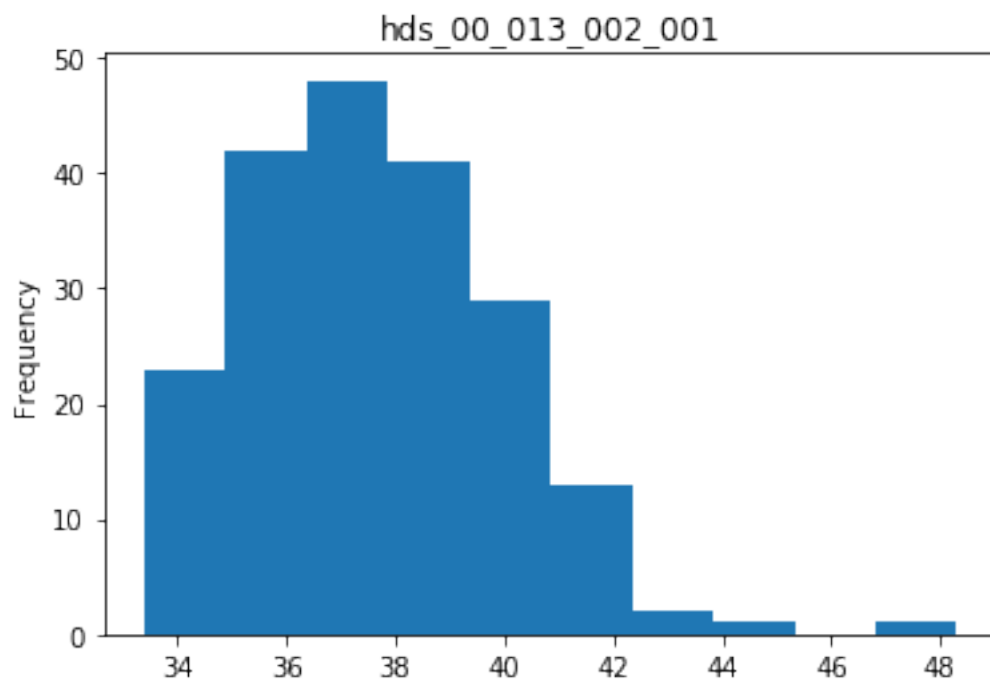
```
Out[8]: ['fa_hw_19791230',
        'fa_hw_19801229',
        'fa_tw_19791230',
        'fa_tw_19801229',
        'hds_00_013_002_000',
        'hds_00_013_002_001',
        'part_time',
        'part_status']
```

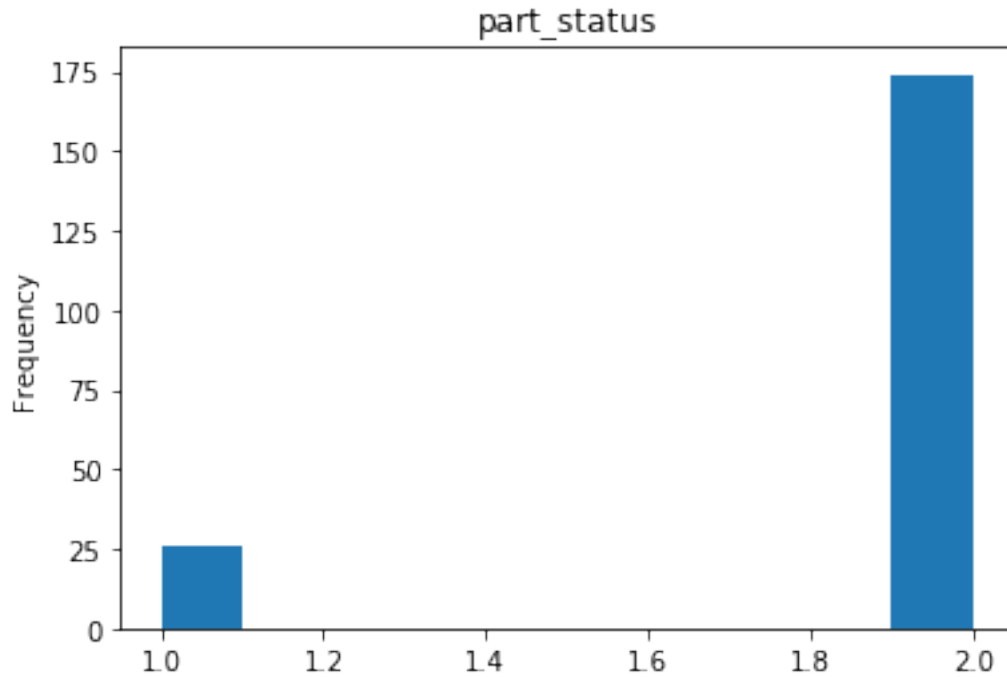
```
In [9]: for forecast in fnames:
        ax = obs_df.loc[:,forecast].plot(kind="hist")
        ax.set_title(forecast)
        plt.show()
```







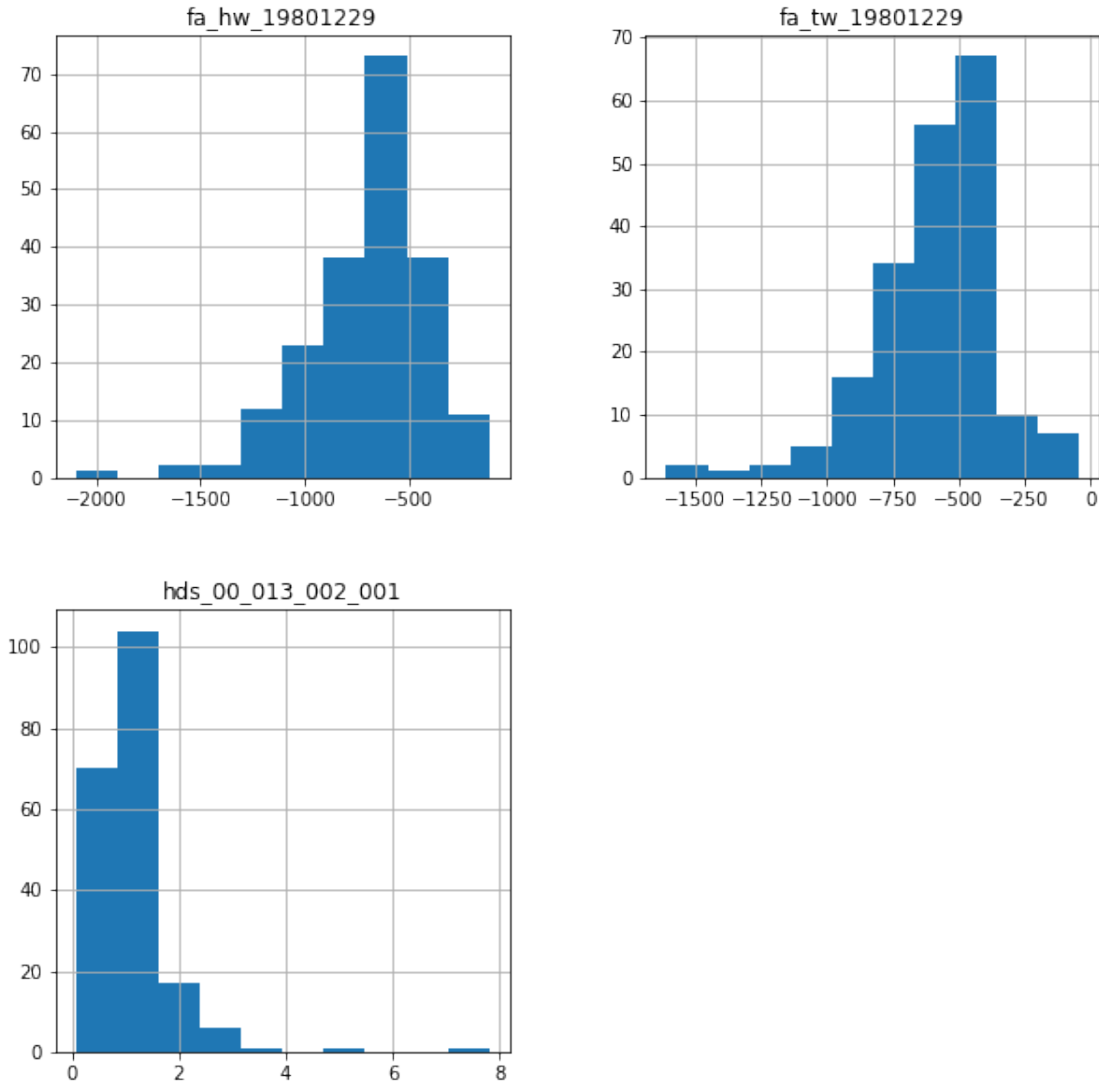




We see that under scenario conditions, many more realizations for the flow to the aquifer in the headwaters are postive (as expected). Lets difference these two:

```
In [10]: sfnames = [f for f in fnames if "1980" in f or "_001" in f]
          hfnames = [f for f in fnames if "1979" in f or "_000" in f]
          diff = obs_df.loc[:,hfnames].values - obs_df.loc[:,sfnames].values
          diff = pd.DataFrame(diff,columns=sfnames)
          diff.hist(figsize=(10,10))
```

```
Out[10]: array([[<matplotlib.axes._subplots.AxesSubplot object at 0x108aeee10>,
                  <matplotlib.axes._subplots.AxesSubplot object at 0x181e6b5550>],
                 [<matplotlib.axes._subplots.AxesSubplot object at 0x181e6ccac8>,
                  <matplotlib.axes._subplots.AxesSubplot object at 0x181e6ea080>]],
              dtype=object)
```



We now see that the most extreme scenario yields a large decrease in flow from the aquifer to the headwaters (the most negative value)

1.0.4 setting the “truth”

We just need to replace the observed values (`obsval`) in the control file with the outputs for one of the realizations on `obs_df`. In this way, we now have the nonzero values for history matching, but also the truth values for comparing how we are doing with other unobserved quantities. Im going to pick a realization that yields an “average” variability of the observed gw levels:

```
In [11]: # choose the realization with a low historic gw to sw headwater flux
hist_sgw = obs_df.loc[:, "fa_hw_19791230"].sort_values()
idx = hist_sgw.index[20]
idx
```

```
Out[11]: 56
```

```
In [12]: obs_df.loc[idx,pst.nnz_obs_names]
```

```
Out[12]: fo_39_19791230      11545.000000
          hds_00_002_009_000    35.945004
          hds_00_002_015_000    34.993938
          hds_00_003_008_000    36.179363
          hds_00_009_001_000    38.799004
          hds_00_013_010_000    35.164818
          hds_00_015_016_000    34.681568
          hds_00_021_010_000    35.060108
          hds_00_022_015_000    34.568565
          hds_00_024_004_000    35.596191
          hds_00_026_006_000    34.910564
          hds_00_029_015_000    34.335033
          hds_00_033_007_000    34.046440
          hds_00_034_010_000    33.606953
          Name: 56, dtype: float64
```

Lets see how our selected truth does with the swgw forecasts:

```
In [13]: obs_df.loc[idx,fnames]
```

```
Out[13]: fa_hw_19791230      -1187.493500
          fa_hw_19801229      -528.937500
          fa_tw_19791230       203.422730
          fa_tw_19801229       721.340850
          hds_00_013_002_000    39.006870
          hds_00_013_002_001    38.093163
          part_time            468.757400
          part_status          2.000000
          Name: 56, dtype: float64
```

```
In [14]: pst = pyemu.Pst(os.path.join(t_d,"freyberg.pst"))
          obs = pst.observation_data
          obs.loc[:, "obsval"] = obs_df.loc[idx,pst.obs_names]
          obs.loc[obs.obgnme=="calhead", "weight"] = 5.0
          obs.loc[obs.obgnme=="calflux", "weight"] = 0.035
          obs.weight.value_counts()
```

```
Out[14]: 0.000      4422
          5.000       13
          0.035        1
          Name: weight, dtype: int64
```

Now, it is custom to add noise to the observed values...we will use the classic Gaussian noise...zero mean and standard deviation of 1 over the weight


```
In [15]: # this should give the same standard normal draws each time
np.random.seed(seed=0)
snd = np.random.randn(pst.nnz_obs)
snd

Out[15]: array([ 1.76405235,  0.40015721,  0.97873798,  2.2408932 ,  1.86755799,
                -0.97727788,  0.95008842, -0.15135721, -0.10321885,  0.4105985 ,
                0.14404357,  1.45427351,  0.76103773,  0.12167502])
```

```
In [16]: noise = snd * 1./obs.loc[pst.nnz_obs_names,"weight"]
noise
```

```
Out[16]: obsnme
fo_39_19791230      50.401496
hds_00_002_009_000    0.080031
hds_00_002_015_000    0.195748
hds_00_003_008_000    0.448179
hds_00_009_001_000    0.373512
hds_00_013_010_000   -0.195456
hds_00_015_016_000    0.190018
hds_00_021_010_000   -0.030271
hds_00_022_015_000   -0.020644
hds_00_024_004_000    0.082120
hds_00_026_006_000    0.028809
hds_00_029_015_000    0.290855
hds_00_033_007_000    0.152208
hds_00_034_010_000    0.024335
Name: weight, dtype: float64
```

Only run this block once!!!

```
In [17]: pst.observation_data.loc[noise.index,"obsval"] += noise
pst.write(os.path.join(t_d,"freyberg.pst"))
pyemu.os_utils.run("pestpp-ies freyberg.pst",cwd=t_d)
```

```
In [18]: pst = pyemu.Pst(os.path.join(t_d,"freyberg.pst"))
print(pst.phi)
pst.res.loc[pst.nnz_obs_names,:]
```

158.52813633665653

```
Out[18]:
```

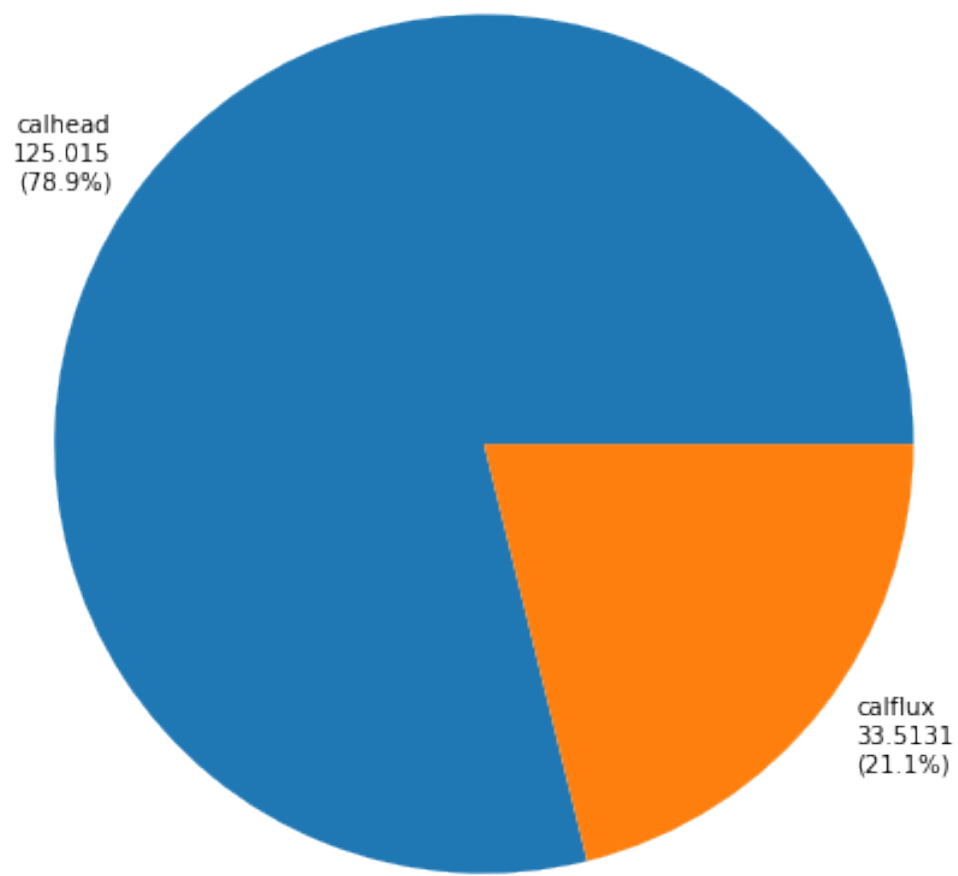
	name	group	measured	modelled \
name				
fo_39_19791230	fo_39_19791230	calflux	11595.401496	11430.000000
hds_00_002_009_000	hds_00_002_009_000	calhead	36.025035	37.107498
hds_00_002_015_000	hds_00_002_015_000	calhead	35.189686	35.045185
hds_00_003_008_000	hds_00_003_008_000	calhead	36.627542	37.397289
hds_00_009_001_000	hds_00_009_001_000	calhead	39.172515	39.546417

hds_00_013_010_000	hds_00_013_010_000	calhead	34.969362	35.571774
hds_00_015_016_000	hds_00_015_016_000	calhead	34.871586	34.835716
hds_00_021_010_000	hds_00_021_010_000	calhead	35.029837	35.386250
hds_00_022_015_000	hds_00_022_015_000	calhead	34.547922	34.577492
hds_00_024_004_000	hds_00_024_004_000	calhead	35.678311	36.760464
hds_00_026_006_000	hds_00_026_006_000	calhead	34.939373	35.896149
hds_00_029_015_000	hds_00_029_015_000	calhead	34.625888	34.453842
hds_00_033_007_000	hds_00_033_007_000	calhead	34.198648	34.678810
hds_00_034_010_000	hds_00_034_010_000	calhead	33.631288	34.118073

	residual	weight
name		
fo_39_19791230	165.401496	0.035
hds_00_002_009_000	-1.082463	5.000
hds_00_002_015_000	0.144501	5.000
hds_00_003_008_000	-0.769747	5.000
hds_00_009_001_000	-0.373902	5.000
hds_00_013_010_000	-0.602411	5.000
hds_00_015_016_000	0.035870	5.000
hds_00_021_010_000	-0.356413	5.000
hds_00_022_015_000	-0.029570	5.000
hds_00_024_004_000	-1.082153	5.000
hds_00_026_006_000	-0.956776	5.000
hds_00_029_015_000	0.172046	5.000
hds_00_033_007_000	-0.480162	5.000
hds_00_034_010_000	-0.486785	5.000

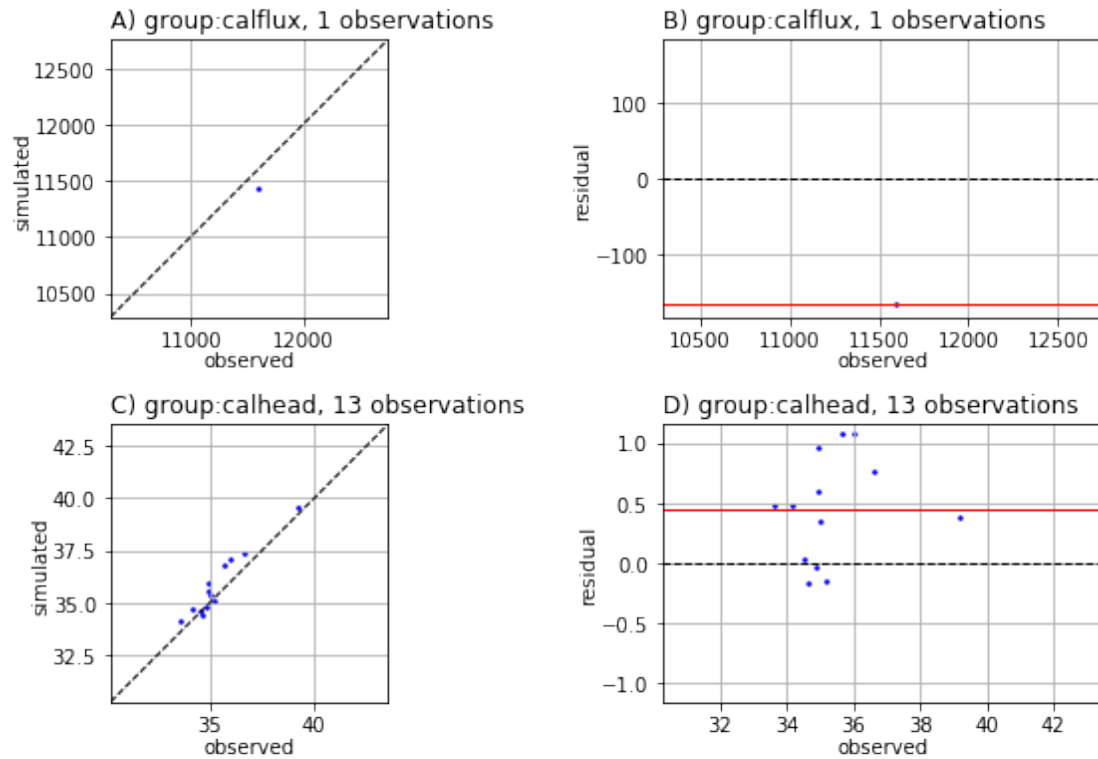
```
In [19]: pst.phi_components
         pst.plot(kind='phi_pie')
```

```
Out[19]: <matplotlib.axes._subplots.AxesSubplot at 0x181e590710>
```



```
In [20]: figs = pst.plot(kind="1to1")
```

<Figure size 576x756 with 0 Axes>



Publication ready figs - oh snap!

Just to make sure we have everything working right, we should be able to load the truth parameters, run the model once and have a ϕ equivalent to the noise vector:

```
In [21]: par_df = pd.read_csv(os.path.join(m_d,"sweep_in.csv"),index_col=0)
pst.parameter_data.loc[:, "parval1"] = par_df.loc[idx,pst.par_names]
pst.write(os.path.join(m_d,"test.pst"))
pyemu.os_utils.run("pestpp-ies.exe test.pst", cwd=m_d)
pst = pyemu.Pst(os.path.join(m_d,"test.pst"))
print(pst.phi)
pst.res.loc[pst.nnz_obs_names,:]
```

17.528847282834874

```
Out [21]:
```

	name	group	measured	modelled \
name				
fo_39_19791230	fo_39_19791230	calflux	11595.401496	11545.000000
hds_00_002_009_000	hds_00_002_009_000	calhead	36.025035	35.945004
hds_00_002_015_000	hds_00_002_015_000	calhead	35.189686	34.993938
hds_00_003_008_000	hds_00_003_008_000	calhead	36.627542	36.179363
hds_00_009_001_000	hds_00_009_001_000	calhead	39.172515	38.799004
hds_00_013_010_000	hds_00_013_010_000	calhead	34.969362	35.164818
hds_00_015_016_000	hds_00_015_016_000	calhead	34.871586	34.681568
hds_00_021_010_000	hds_00_021_010_000	calhead	35.029837	35.060108
hds_00_022_015_000	hds_00_022_015_000	calhead	34.547922	34.568565
hds_00_024_004_000	hds_00_024_004_000	calhead	35.678311	35.596191
hds_00_026_006_000	hds_00_026_006_000	calhead	34.939373	34.910564
hds_00_029_015_000	hds_00_029_015_000	calhead	34.625888	34.335033
hds_00_033_007_000	hds_00_033_007_000	calhead	34.198648	34.046440
hds_00_034_010_000	hds_00_034_010_000	calhead	33.631288	33.606953

	residual	weight
name		
fo_39_19791230	50.401496	0.035
hds_00_002_009_000	0.080031	5.000
hds_00_002_015_000	0.195748	5.000
hds_00_003_008_000	0.448179	5.000
hds_00_009_001_000	0.373512	5.000
hds_00_013_010_000	-0.195456	5.000
hds_00_015_016_000	0.190018	5.000
hds_00_021_010_000	-0.030271	5.000
hds_00_022_015_000	-0.020644	5.000
hds_00_024_004_000	0.082120	5.000
hds_00_026_006_000	0.028809	5.000
hds_00_029_015_000	0.290855	5.000
hds_00_033_007_000	0.152208	5.000
hds_00_034_010_000	0.024335	5.000

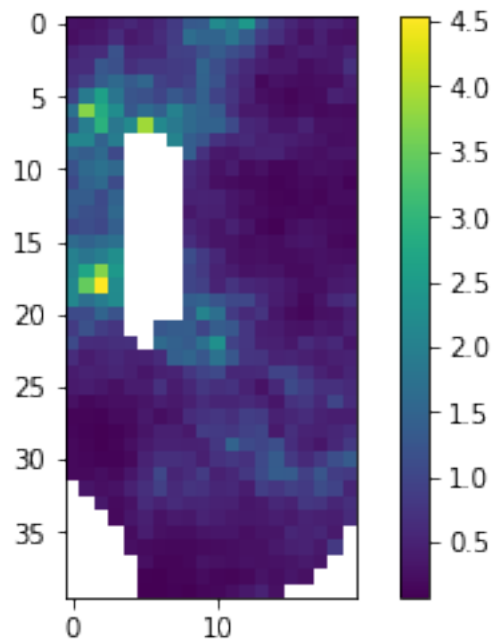
The residual should be exactly the noise values from above. Lets load the model (that was just run using the true pars) and check some things

```
In [22]: m = flopy.modflow.Modflow.load("freyberg.nam",model_ws=m_d)
```

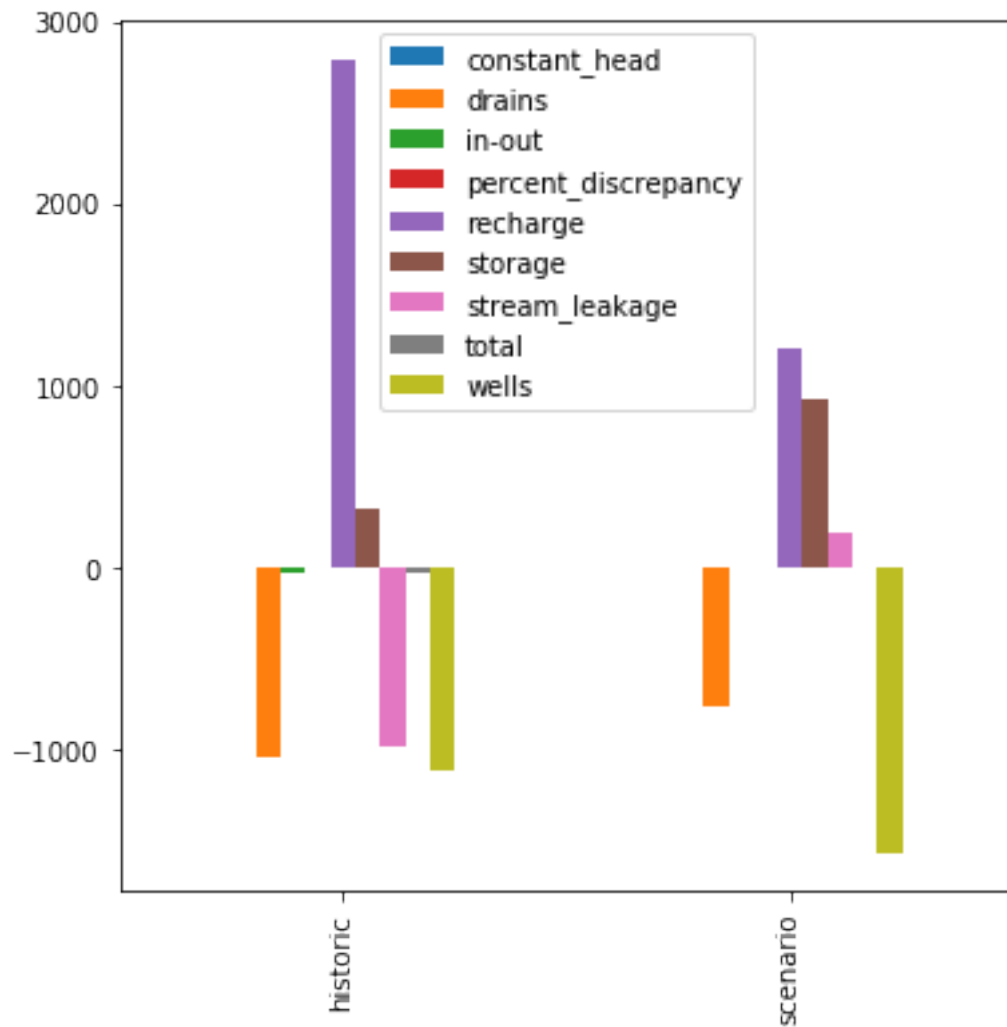
```
In [23]: a = m.upw.vka[1].array
         #a = m.rch.rech[0].array
         a = np.ma.masked_where(m.bas6.ibound[0].array==0,a)
         print(a.min(),a.max())
         c = plt.imshow(a)
         plt.colorbar(c)
```

0.07383982 4.534729

Out[23]: <matplotlib.colorbar.Colorbar at 0x181ef735f8>



```
In [24]: lst = flopy.utils.MfListBudget(os.path.join(m_d,"freyberg.list"))
         df = lst.get_dataframes(diff=True)[0]
         ax = df.plot(kind="bar",figsize=(6,6))
         a = ax.set_xticklabels(["historic","scenario"],rotation=90)
```

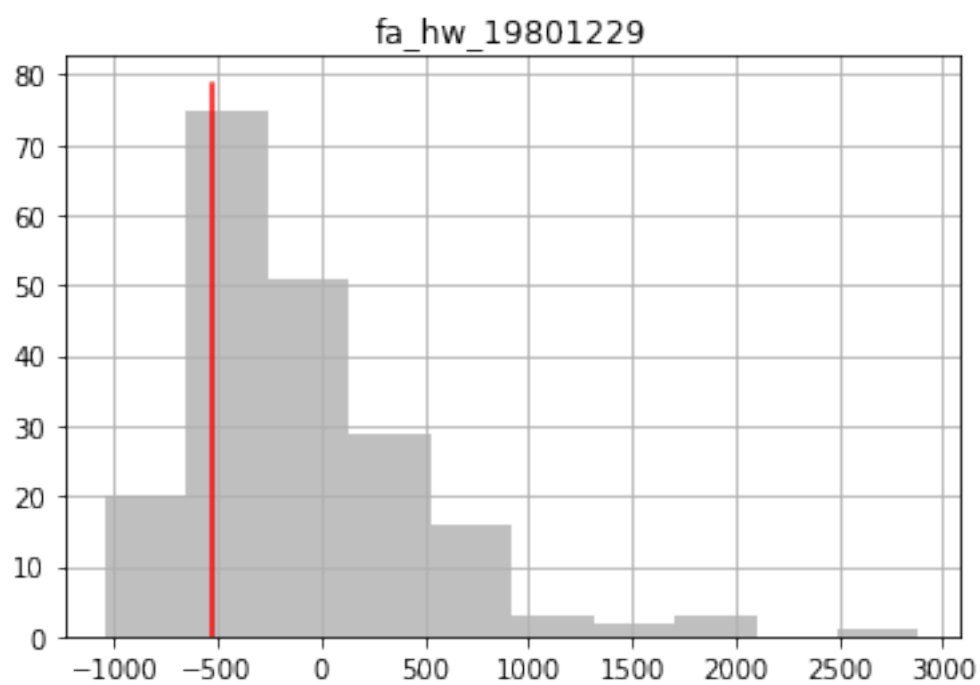
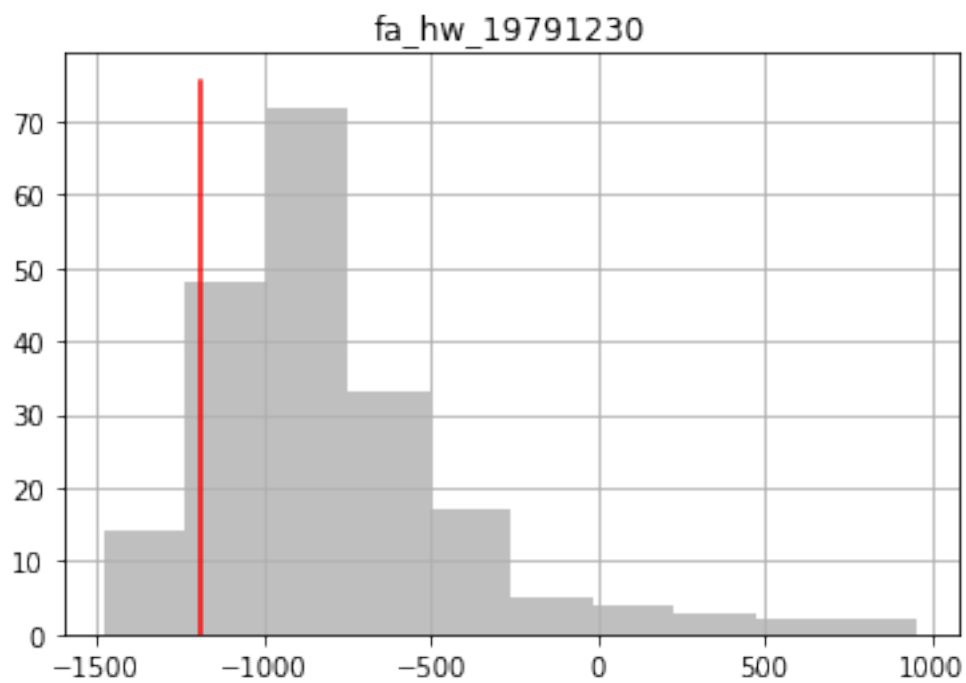


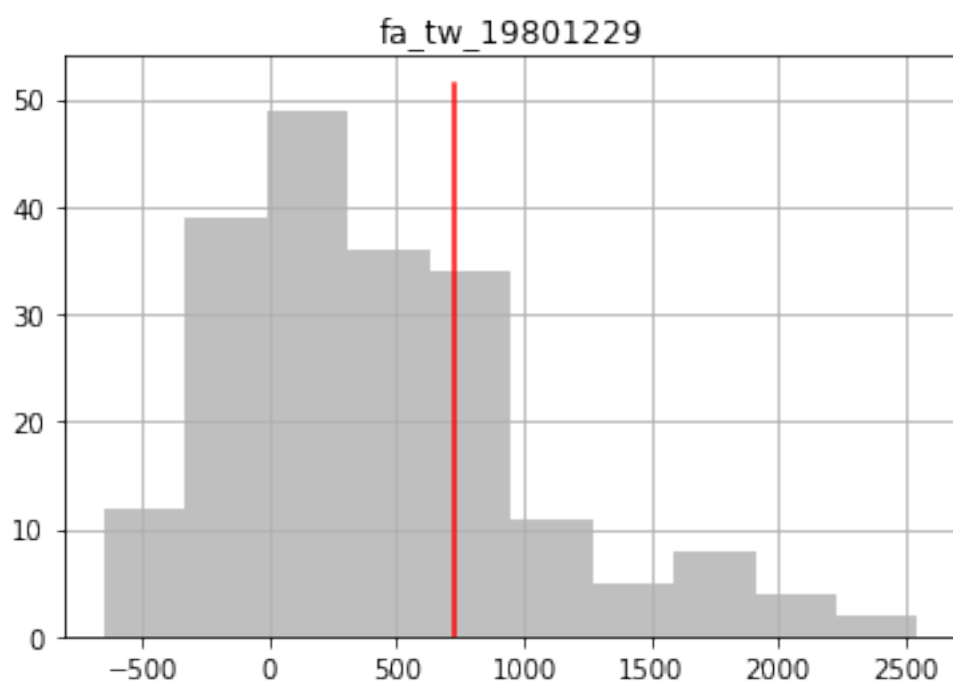
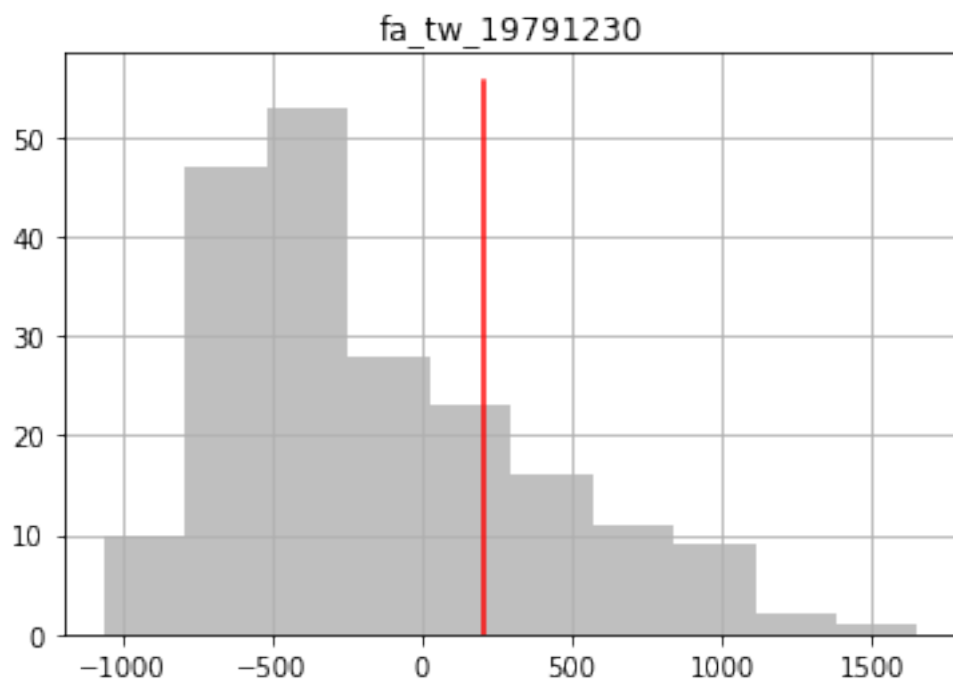
1.0.5 see how our existing observation ensemble compares to the truth

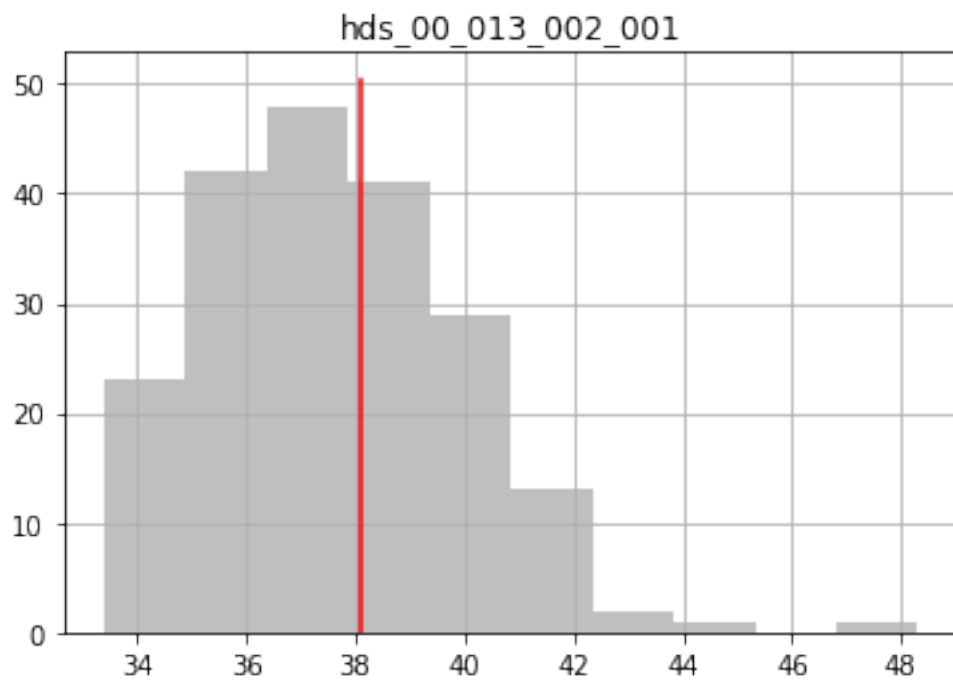
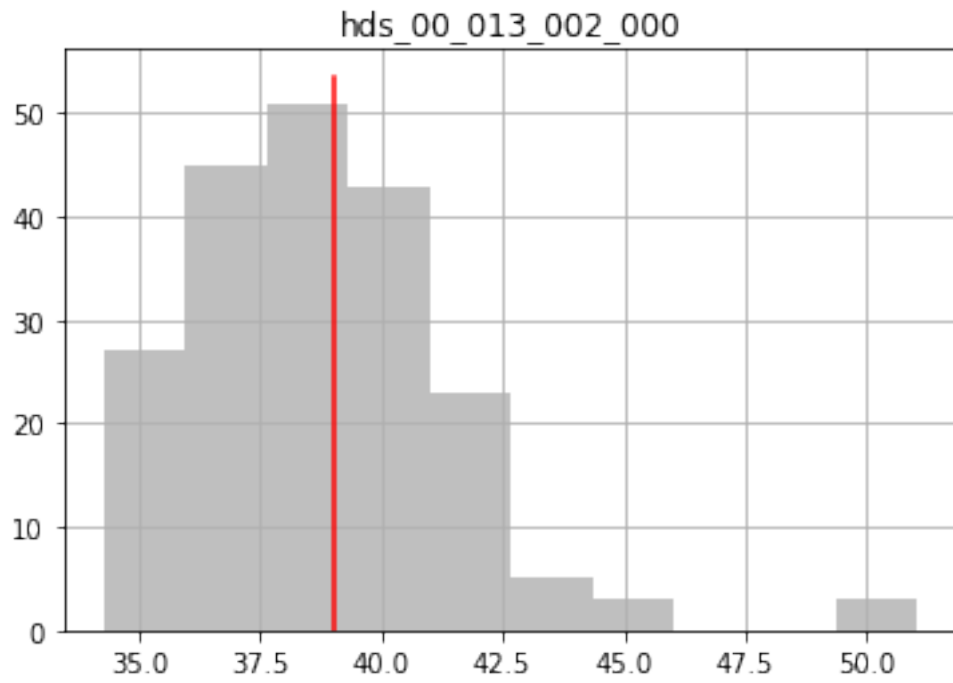
sw-gw outputs:

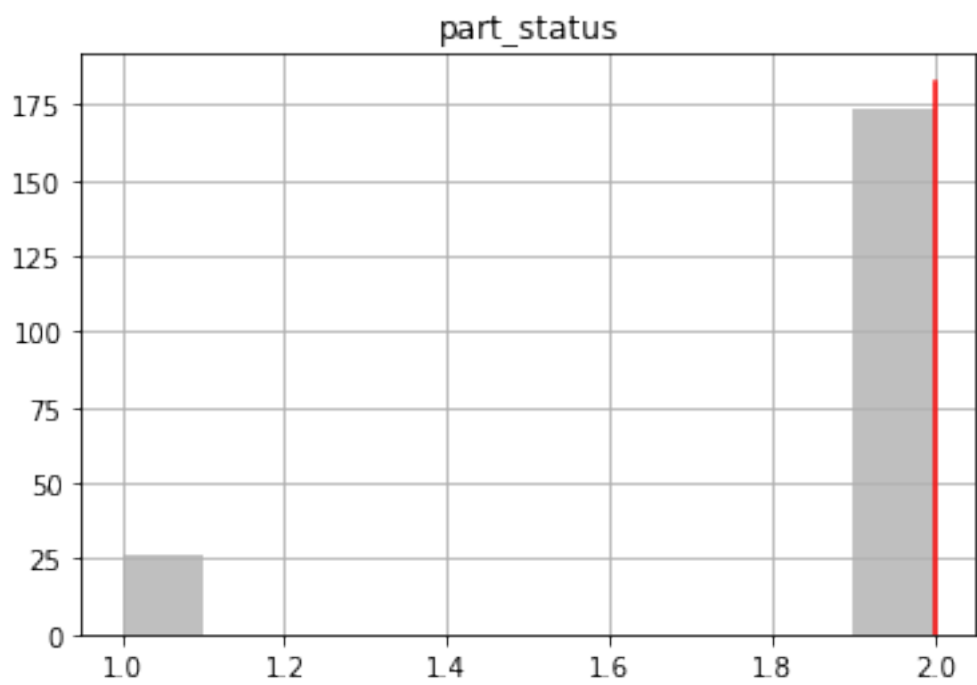
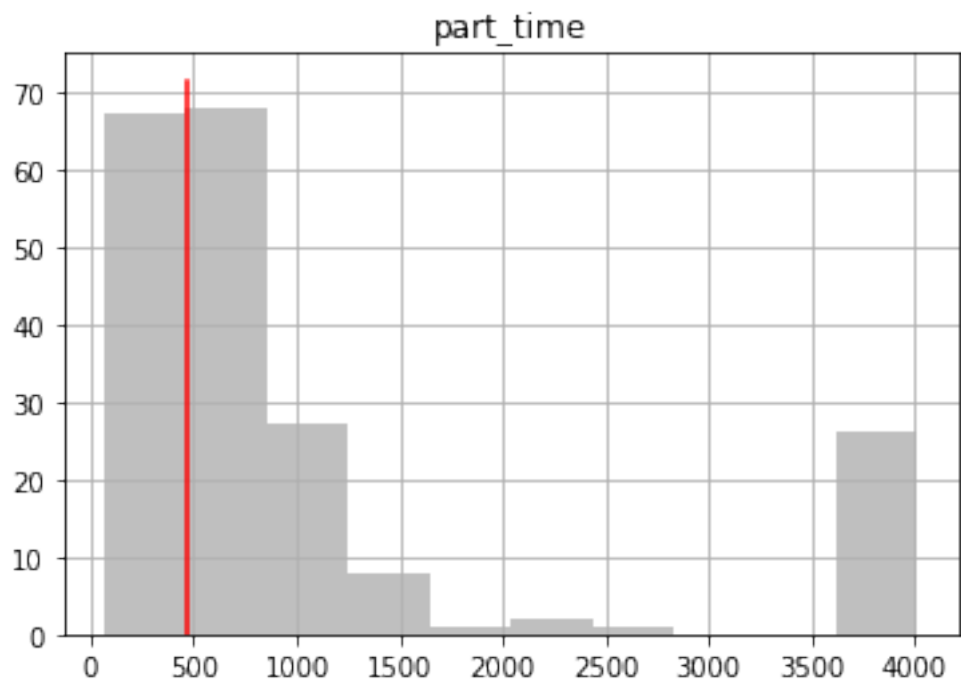
```
In [25]: obs = pst.observation_data
```

```
for forecast in fnames:
    ax = plt.subplot(111)
    obs_df.loc[:,forecast].hist(ax=ax,color="0.5",alpha=0.5)
    ax.plot([obs.loc[forecast,"obsval"],obs.loc[forecast,"obsval"]],ax.get_ylim(),"r")
    ax.set_title(forecast)
    plt.show()
```









observations:

```
In [26]: for oname in pst.nnz_obs_names:
          ax = plt.subplot(111)
          obs_df.loc[:, oname].hist(ax=ax, color="0.5", alpha=0.5)
          ax.plot([obs.loc[oname, "obsval"], obs.loc[oname, "obsval"]], ax.get_ylim(), "r")
          ax.set_title(oname)
          plt.show()
```

