

# Nontrivial Solutions of the Scalar Constraint in the Unified Biquaternion Theory

## 1 Introduction

We consider the scalar constraint from Priority 1:

$$\eta^{\mu\nu}(\partial_\mu\rho)(\partial_\nu\phi) = 0$$

where  $\rho = |\Theta|$  is the amplitude and  $\phi$  the phase of the field. This condition geometrically enforces the orthogonality of gradients of amplitude and phase in spacetime.

## 2 Nontrivial Minkowski Solutions

In Minkowski spacetime, one trivial solution is when either  $\rho$  or  $\phi$  is constant. However, we can construct richer solutions.

Let:

$$\rho = f(t - x), \quad \phi = g(t + x)$$

Then:

$$\partial_\mu\rho = f'(t - x)(\delta_\mu^0 - \delta_\mu^1), \quad \partial_\nu\phi = g'(t + x)(\delta_\nu^0 + \delta_\nu^1)$$

so:

$$\eta^{\mu\nu}\partial_\mu\rho\partial_\nu\phi = f'(t - x)g'(t + x)(\eta^{00} - \eta^{01} + \eta^{10} - \eta^{11}) = 0$$

since the mixed terms cancel.

Thus, such left/right-moving wave combinations satisfy the scalar constraint.

## 3 Axially Symmetric Configurations

We explore solutions of the form:

$$\rho = \rho(r), \quad \phi = \phi(t - r)$$

with  $r = \sqrt{x^2 + y^2 + z^2}$ . Then:

$$\begin{aligned} \partial_t\rho &= 0, & \nabla\rho &= \rho'(r)\frac{\vec{r}}{r} \\ \partial_t\phi &= \phi'(t - r), & \nabla\phi &= -\phi'(t - r)\frac{\vec{r}}{r} \end{aligned}$$

The scalar constraint becomes:

$$\eta^{\mu\nu}\partial_\mu\rho\partial_\nu\phi = -\nabla\rho \cdot \nabla\phi = -\rho'(r)\phi'(t - r)$$

This vanishes iff  $\rho'(r)\phi'(t - r) = 0$ , leading to conditional satisfaction.

## 4 Implications in FRW Spacetimes

In cosmology, consider an FRW metric:

$$ds^2 = -dt^2 + a(t)^2 d\vec{x}^2$$

We look at configurations:

$$\rho = \rho(r), \quad \phi = \phi(t)$$

Then:

$$\partial_\mu \rho = \delta_\mu^i \partial_i \rho, \quad \partial_\nu \phi = \delta_\nu^0 \dot{\phi}$$

so:

$$g^{\mu\nu} \partial_\mu \rho \partial_\nu \phi = g^{0i} \dot{\phi} \partial_i \rho = 0$$

since  $g^{0i} = 0$ , this is satisfied. But with  $\phi = \phi(t, r)$ , nontrivial structures emerge.

## 5 Conclusions and Future Work

We have demonstrated a family of exact solutions beyond trivial cases. These exhibit spatial-temporal interference patterns and allow for localized dynamics in scalar field evolution. This paves the way for:

- Numerical simulations of scalar-phase interaction.
- Exploring implications near strong gravitational fields.
- Deriving effective potentials from interactions.