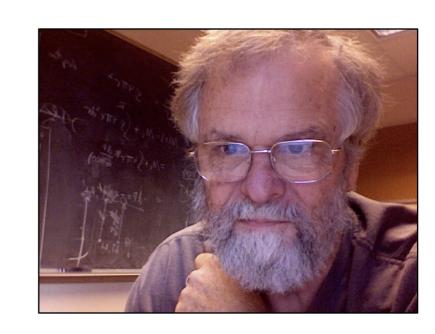


Modules for Experiments in Stellar Astrophysics

MESA is a state-of-the-art, modular, open source suite for stellar evolution



Bill Paxton, father of MESA

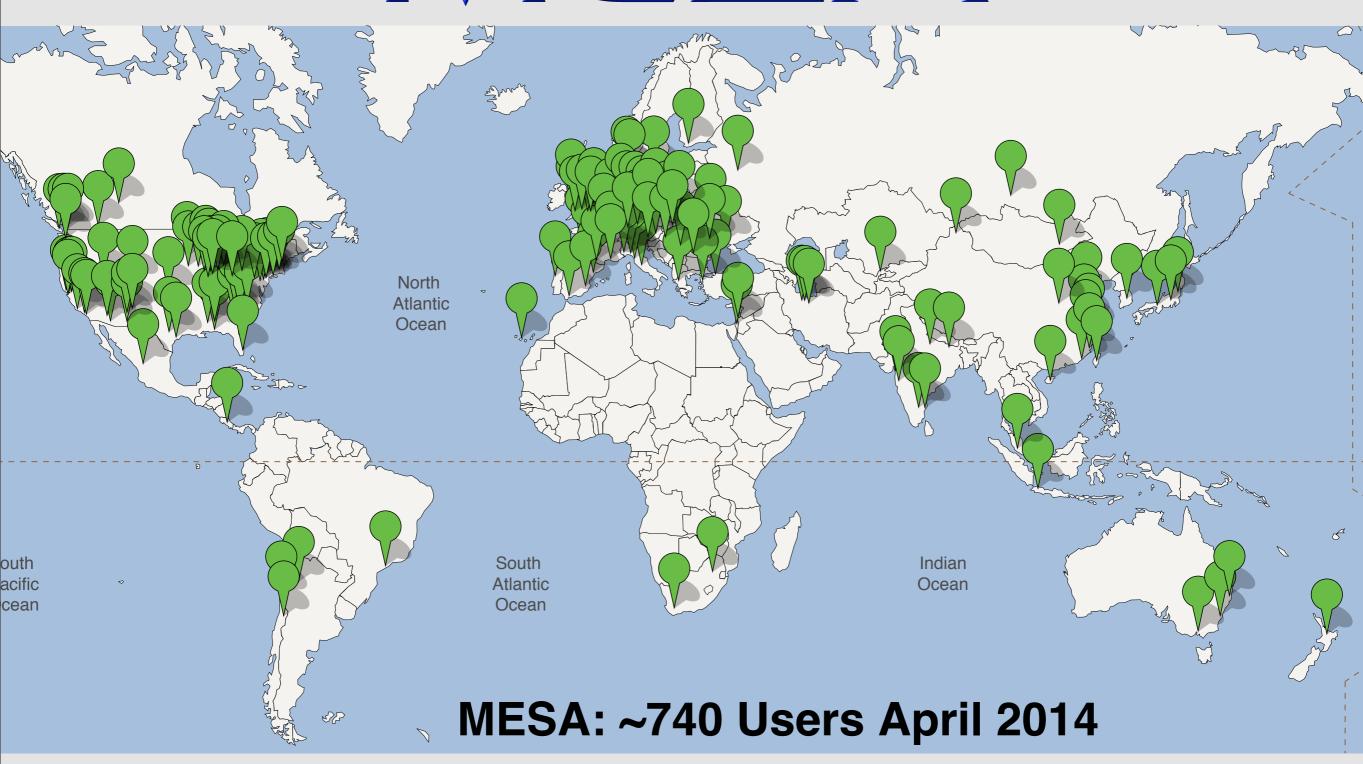
- MESA Stellar Evolution Code: mesa.sourceforge.net
- MESA Instrument Papers (<u>Paxton et al. 2011</u>, <u>2013</u>)



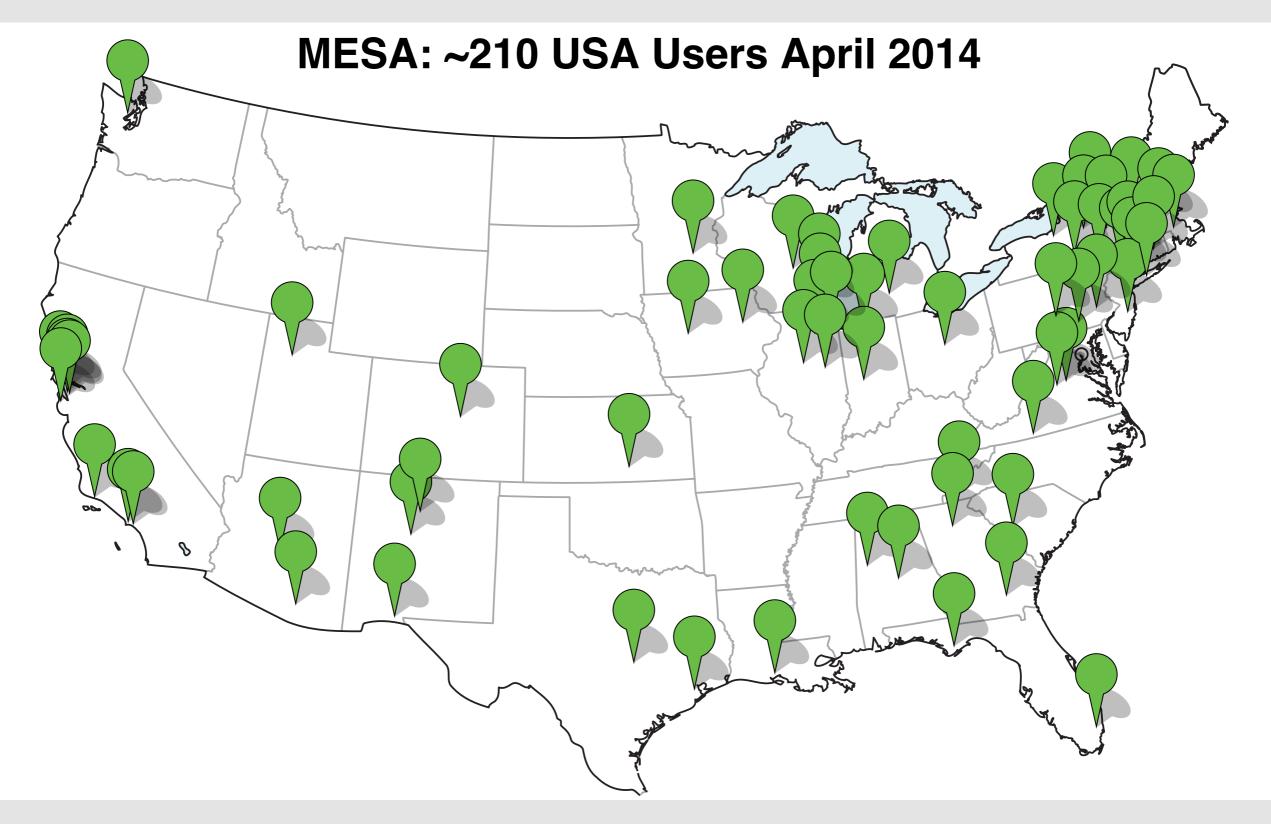
- Openness: anyone can download sources from the website.
- Modularity: independent modules for physics and for numerical algorithms; the parts can be used stand-alone.
- Wide Applicability: capable of calculating the evolution of stars in a wide range of environments.
- Modern Techniques: advanced AMR, fully coupled solution for composition and abundances, mass loss and gain, etc.
- Comprehensive Microphysics: up-to-date, wide-ranging, flexible, and independently useable microphysics modules.
- Performance: runs well on a personal computer and makes effective use of parallelism with multi-core architectures.

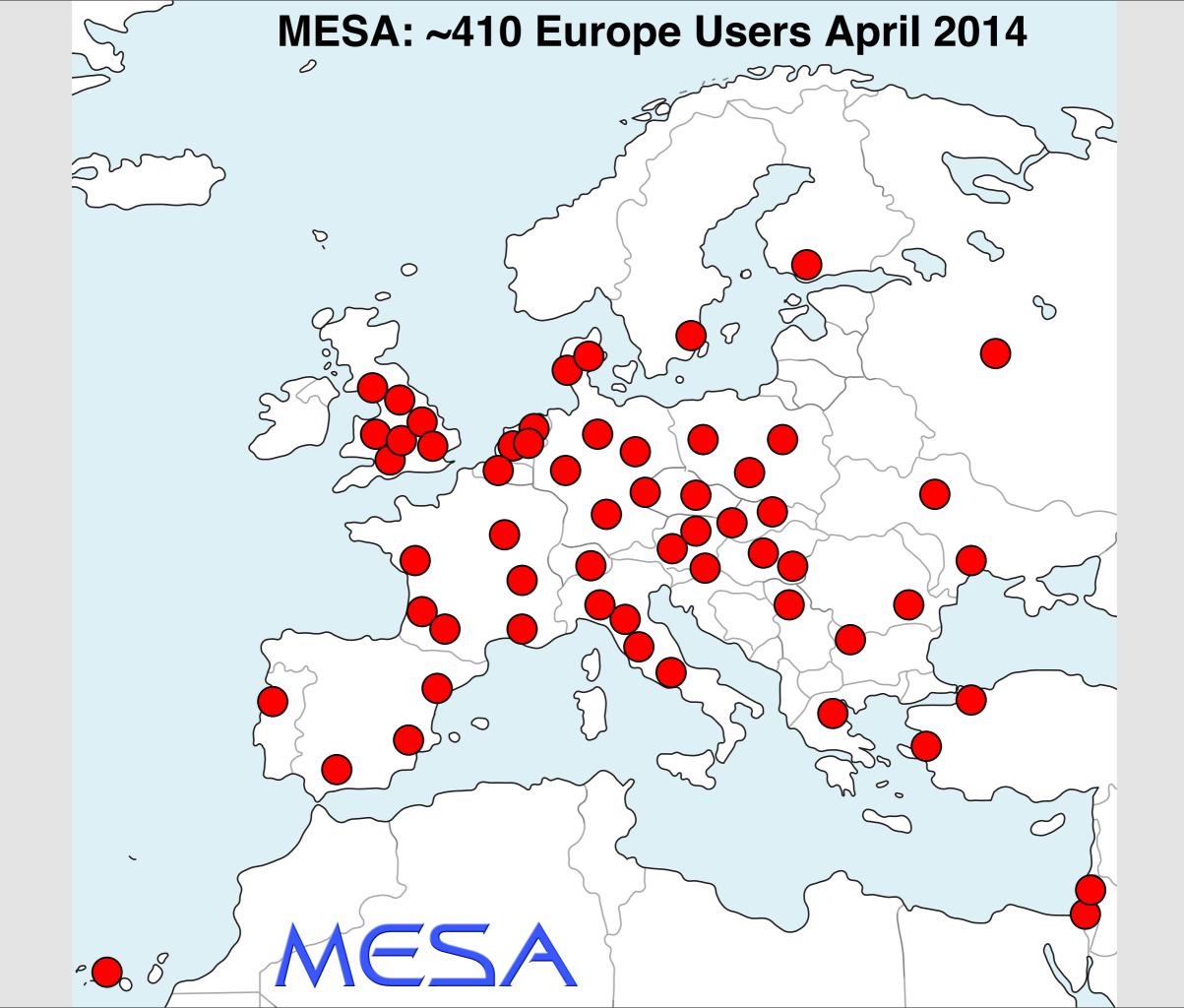


MESA



MESA







MESA Capabilities

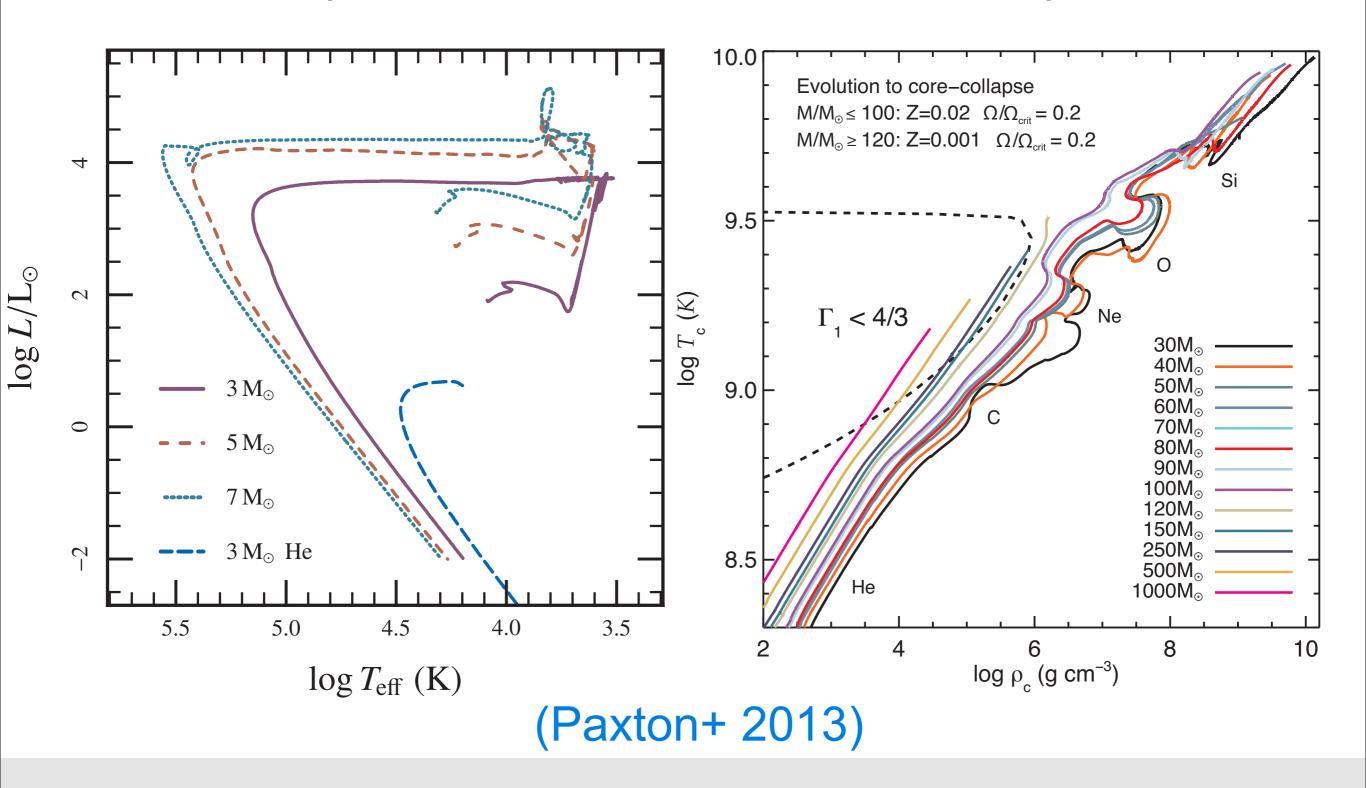
- Mass loss/gain, Schwarzschild/Ledoux, double diffusion, gravitational settling, radiative levitation...
- It includes the physics of rotation (in a diffusion approximation) and of dynamo generated magnetic fields in radiative zones
- Asteroseismology: MESA is natively coupled with two oscillations codes: ADIPLS (J. Christensen-Dalsgaard 2008) and the non-adiabatic code GYRE (Townsend & Teitler 2013)
- **Giant planets**, Low-mass stars, Massive Stars, Compact Objects, Asteroseismology, Accretion / Massloss...

(Paxton+ 2011, 2013)

Single stars with



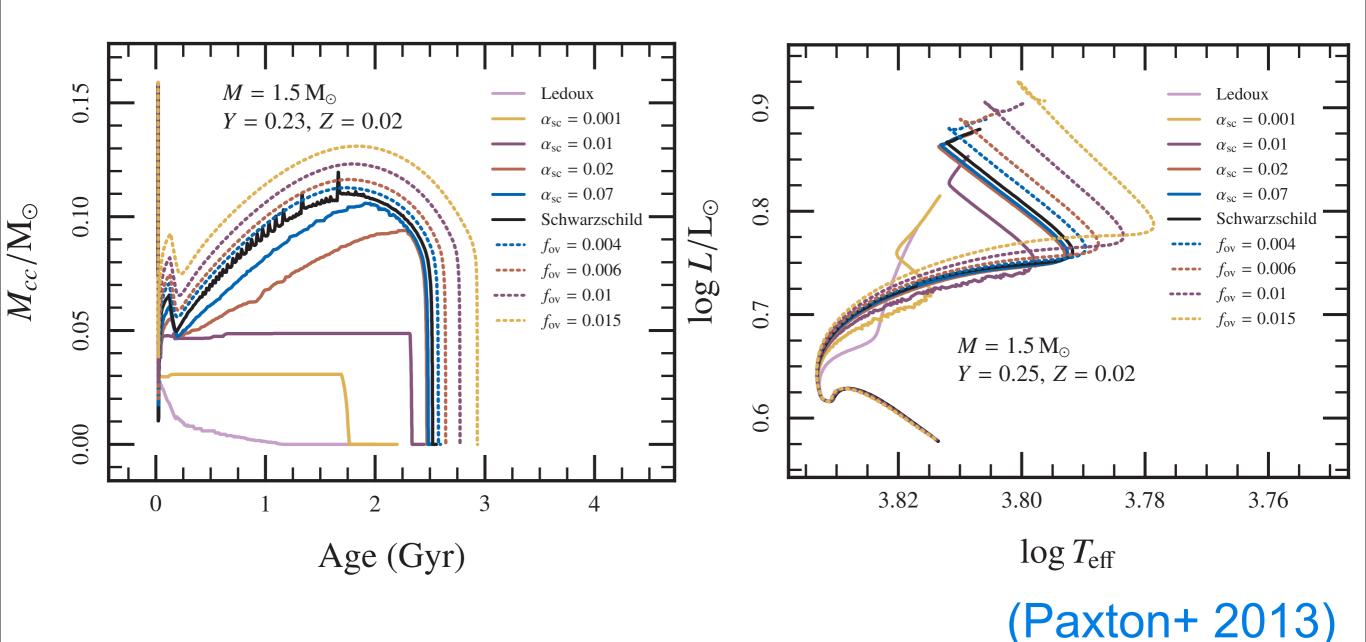
Uninterrupted evolution to WD and core-collapse



Single stars with



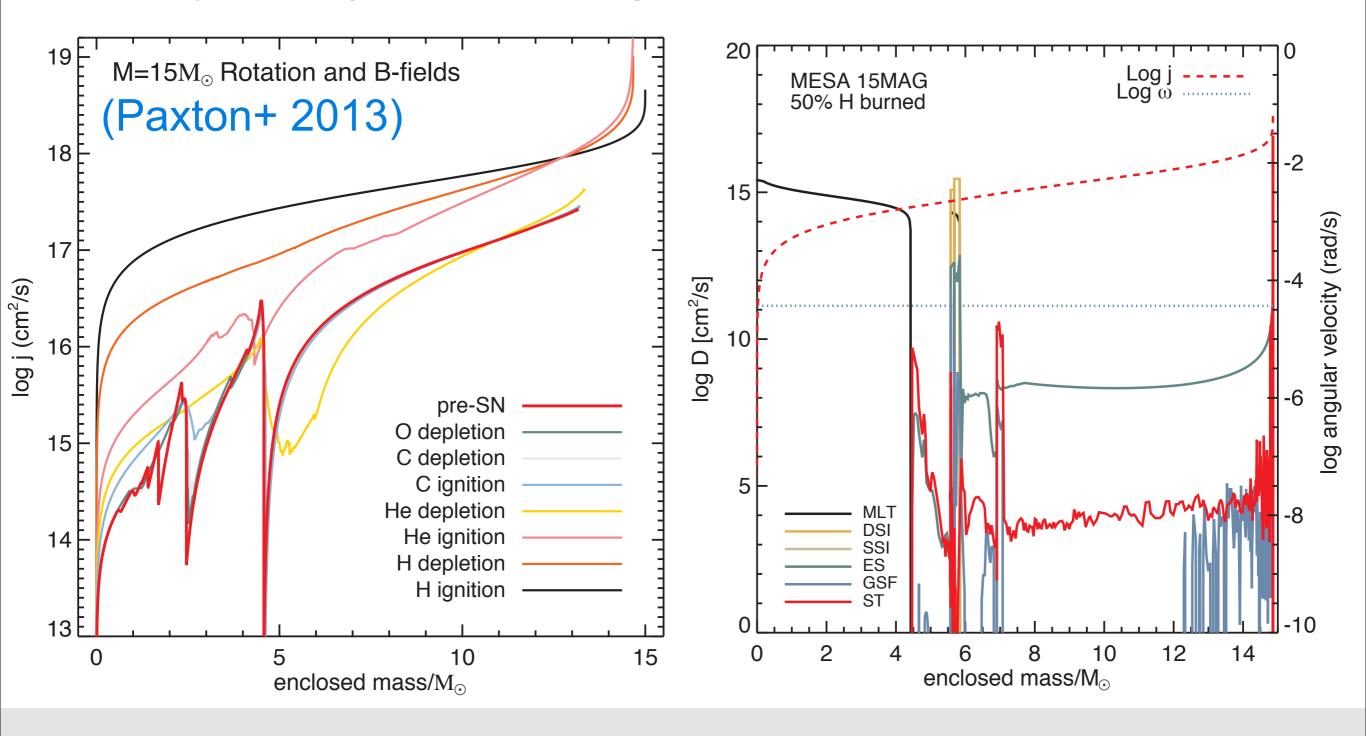
 Mass loss/gain, Schwarzschild/Ledoux, overshooting, double diffusion, gravitational settling, radiative levitation...



Single stars with

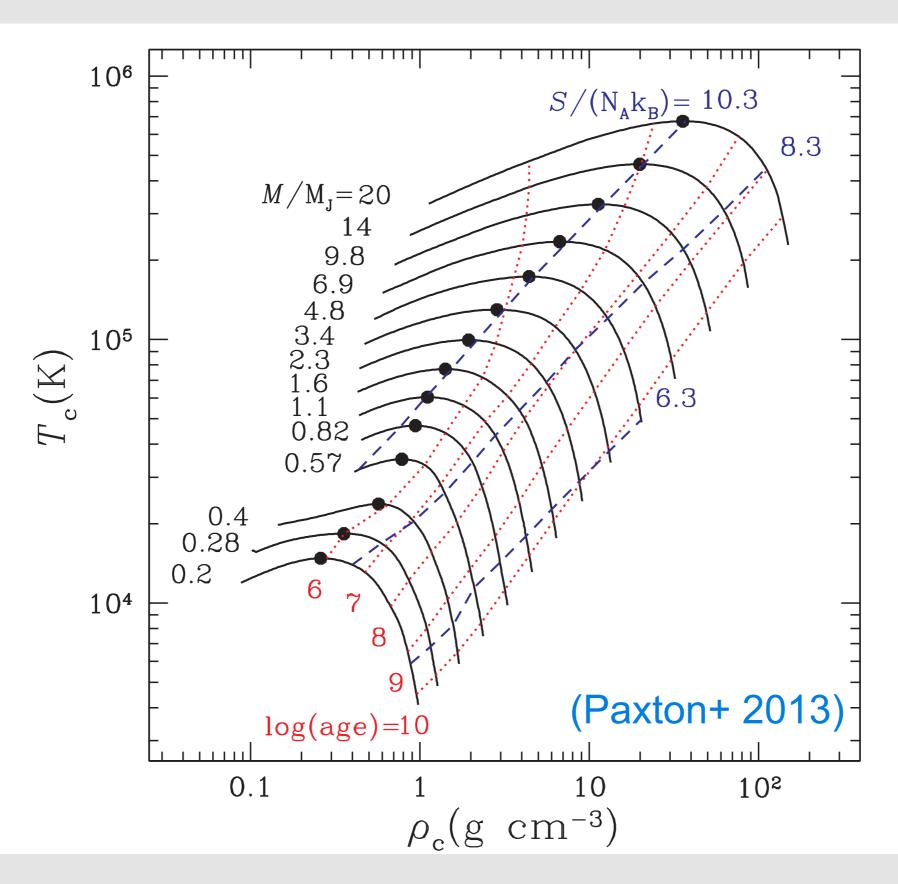


Includes the physics of rotation (in a diffusion approximation) and of dynamo generated magnetic fields in radiative zones



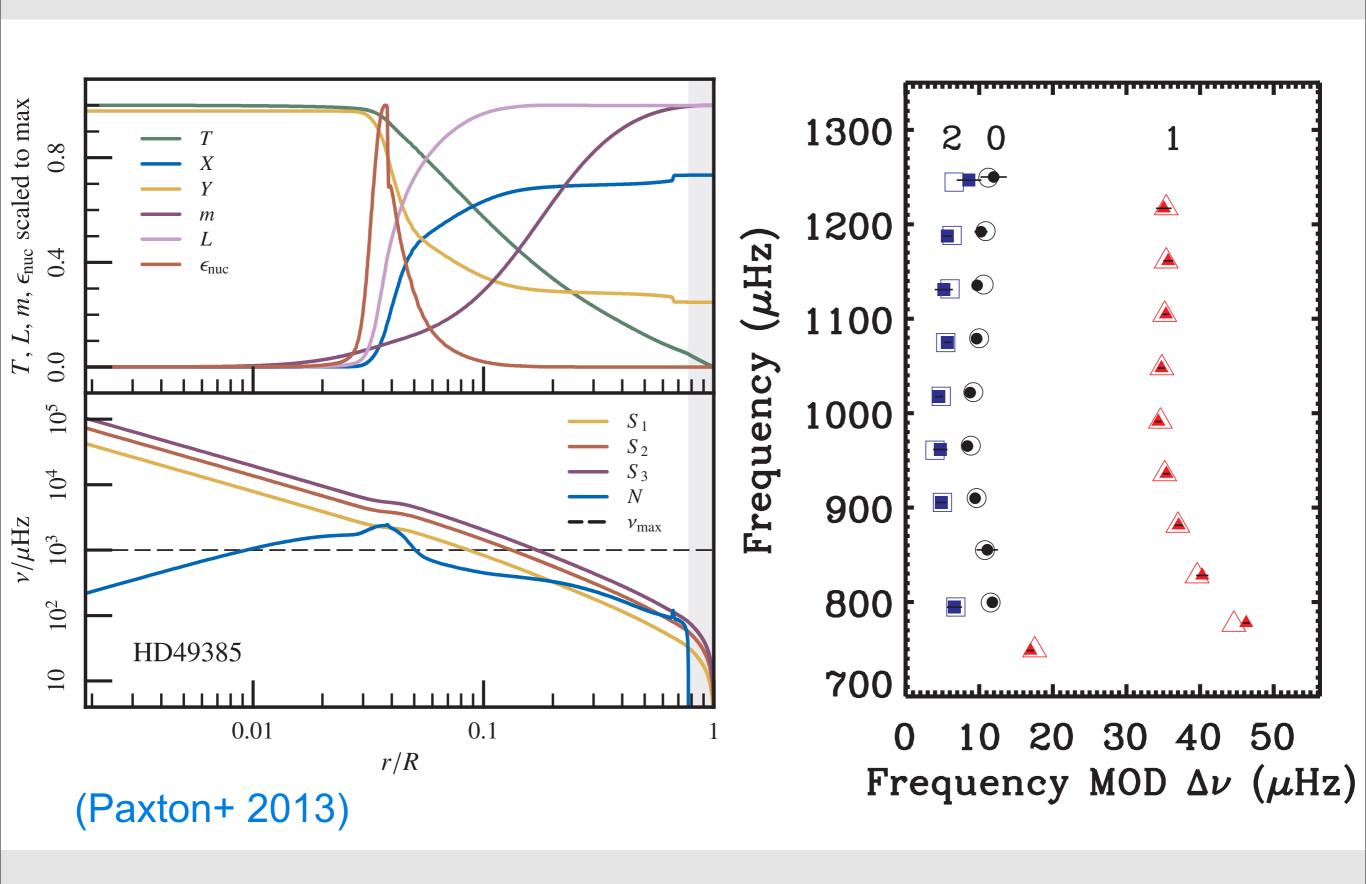
Giant Planets with **M**€5/





Asteroseismology with MES/

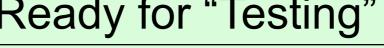




Binary Stars with MESA

- 2 stars evolved at the same time
- Implicit mass transfer ("Ritter" formalism)
- Magnetic Braking (Rappaport, Verbunt & Joss 1983)
- GW-Braking
- Irradiation

Ready for "Testing"



- Tides (Zahn 1977, Tassoul & Tassoul)
- Angular momentum accretion
- L-S coupling

Not there yet

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