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# OCES 2003 : Descriptive Physical Oceanography

(a.k.a. physical oceanography by drawing pictures)

Lecture 16: Dynamics 2 (waves and dynamic mechanisms)

Thur 8<sup>th</sup> Apr

# Outlook of the next few lectures

**Dynamics** important, next few lectures on

- ▶ **waves** (this Lec. + 16, 18) and **instabilities** (Lec. 17)

→ because waves are easier to talk about without maths...

Highlight gross features (i.e. those that can be drawn...)

- ▶ how to describe waves (Lec. 15)
- ▶ **types of waves** (Lec. 16)
  - consequence + leading to instabilities
- ▶ instabilities (Lec. 17)
  - **parcel**-type (mechanistic) arguments for instability
- ▶ tides (particularly as **internal gravity waves**) (Lec. 18)

# Outline

- ▶ gravity waves
  - gravity/buoyancy as restoring mechanism ( $\sqrt{gH}$ )
- ▶ inertial waves
  - Coriolis as restoring mechanism ( $f$ )
  - e.g. Rossby waves, Kelvin waves
- ▶ inertial-gravity + internal waves ( $\sqrt{gH}$  or  $N$ , and  $f$ )
  - extra depth dimension to deal with
  - Brunt–Väisälä or buoyancy frequency  $N$
- ▶ propagation mechanism (Rossby wave example)
  - kinematic argument with **vorticity**

**Key terms:** buoyancy frequency, gravity waves, inertial waves, Rossby waves, Kelvin waves, vorticity inversion

# Recap: what goes down must come up

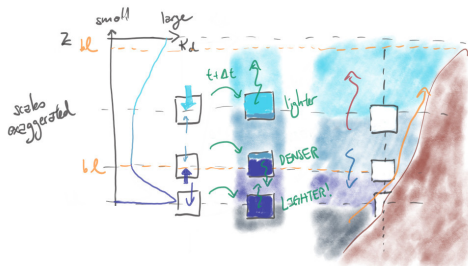


Figure: Schematic of the diffusive upwelling.

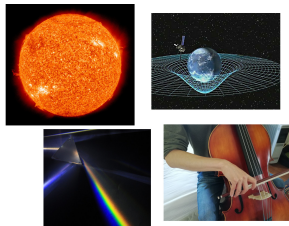
- diapycnal mixing contribute upwelling, strongest in boundary layers  
→ broad diffusive boundary intensified upwelling

what causes the boundary intensification of  $\kappa_d$ ? **dynamics!**

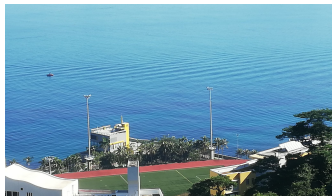
- at the surface, lots of things... (convection, waves, Langmuir turbulence etc.)
- at the bottom, probably tidal conversion (Lec. 18) → internal gravity waves (Lec. 16) → shear instabilities (Lec. 17)

# Recap: waves and dispersion relation

- ▶ **waves** are ubiquitous physical features
  - depends on physics
- ▶ wave described by the **dispersion relation**  $\omega = \mathcal{F}(k)$ 
  - physics dictate the form of  $\mathcal{F}$



**Figure:** Examples of systems supporting waves. All figures from Wikipedia except the cello one.



**Figure:** Gravity waves with signal at the sea surface (as darker and lighter bands). Taken at HKUST.

- ▶ (linear) waves can **interfere** with each other
  - **constructive** or **destructive**
  - interference can lead to **steepening** and **breaking** (“becoming”  
**nonlinear**)

# Recap: wave propagation

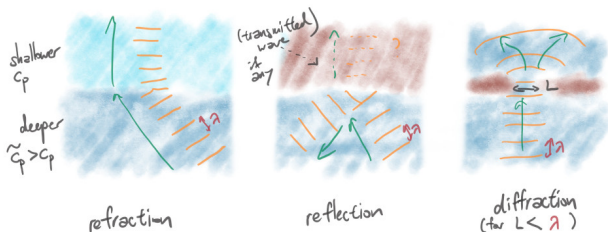


Figure: Schematic of **refraction**, **reflection** (and **transmission**), and **diffraction**. See Lec. 15.

- **phase speed** (in a direction) and **group velocity** as (note  $\omega = \mathcal{F}(k)$ )

$$c_{p,x} = \frac{\omega}{k}, \quad c_{g,x} = \frac{\partial \omega}{\partial k}$$

→ individual wave vs. **wavepacket** behaviour

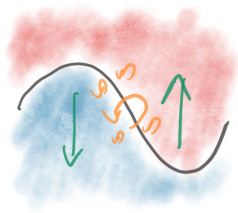
→ contribute to wave phenomenon (e.g. refraction from  $c_p = c_p(x)$ )

## Wave steepening and breaking



**Figure:** Schematic of mixing by (irreversible) wave breaking, with contours reconnecting leading to e.g. diapycnal mixing.

- ▶ growing waves by **instability**
  - convective and/or shear (see Lec. 17)
  - mixing of material **across** isopycnals after reconnection, leading to **diapycnal mixing**
- ▶ feedback onto MOC (see Lec. 14)



**Figure:** Velocity shear from waves can lead to mixing.



# Gravity waves

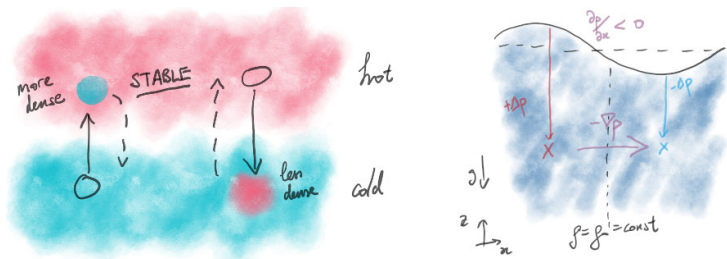


Figure: Gravity as restoring force. Pictures adapted from ones used in Lec. 7.

- ▶ deviation from resting **isopycnal** experiences restoring force from gravity (buoyancy)
  - left case: internal isopycnal (as an **isotherm**)
  - right case: sea surface is the isopycnal
- ▶ weak damping, restoring force, overshoots  $\Rightarrow$  oscillatory motion (up and down in this case)

# Gravity waves

For simplicity, consider a **homogeneous** (i.e.  $\rho = \text{const}$ ) layer of fluid (cf. Lec. 11 + 12, right case in previous Figure) until we get to **internal waves**

# Gravity waves

For simplicity, consider a **homogeneous** (i.e.  $\rho = \text{const}$ ) layer of fluid (cf. Lec. 11 + 12, right case in previous Figure) until we get to **internal waves**

In this instance, dispersion relation for **gravity waves** is given by (without derivation)

$$\omega^2 = gk \tanh(kH) \quad \Rightarrow \quad \omega = \pm \sqrt{gk \tanh(kH)}$$

- ▶  $H$  is water depth, and  $\tanh$  = hyperbolic tangent, goes from  $-1$  to  $1$ 
  - note **symmetry** in both directions (the  $\pm$  sign)
  - ( $\omega \leq 0$  cases are just **shifts** in the **phase**)

# Gravity waves

$$\omega = \pm \sqrt{gk \tanh(kH)}$$

- ▶ for **deep** water waves ( $kH \gg 1$ ) and **shallow** water waves ( $kH \ll 1$ ),

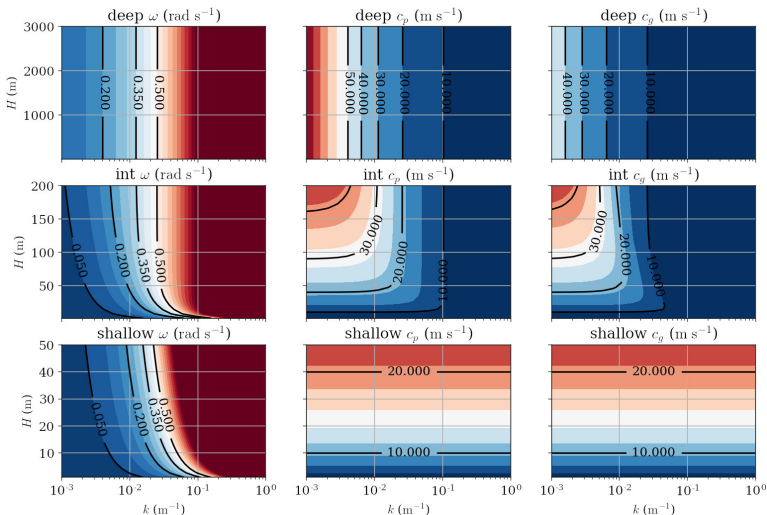
$$\omega_{\text{deep}} = \pm \sqrt{gk}, \quad \omega_{\text{shallow}} = \pm k \sqrt{gH}$$

→  $kH \gg 1$  so  $\tanh(kH) \rightarrow 1$

→  $kH \ll 1$  with  $\tanh(kH) \approx kH + O((kH)^3)$  (do a Taylor expansion)

- ▶ deep water waves are **depth-independent** and **dispersive**
- ▶ shallow water waves are slower in shallow waters ( $c_p \sim \sqrt{H}$ ) and **non-dispersive** ( $c_p = c_g$ )

# Gravity waves

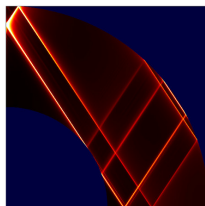
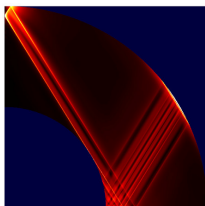


**Figure:** Water wave  $\omega$ ,  $c_p$  and  $c_g$  over  $(k, H)$  space, with  $k$  shown on a log axis.  $k$  chosen so wavelengths are roughly between 50 m to 5 km (recall  $k = 2\pi/\lambda$ ). Also note the transitions from shallow to intermediate to deep are really to do with  $kH \sim H/\lambda$ . See `waves.ipynb`.

# Inertial waves

**Inertial waves** has the Coriolis “force” act as the restoring force

- ▶ generic for rotating systems  
→ planetary interiors, stars, galactic disks
- ▶ mostly arise in context of **internal waves**  
→ at surface buoyancy effects can dominate
- ▶ limited in frequency by **inertial frequency**  $f_0$  (cf. Coriolis parameter)  
→ revisit later when talking about **inertia-gravity waves**



**Figure:** Inertial wave attractors in a homogeneous planetary interior at different tidal forcings. From Gordon Ogilvie (2009, *Mon. Not. Royal. Astro. Soc.*).

# Inertial-Gravity waves

In reality **Coriolis** and **buoyancy** effects both contribute

- ▶ large-scale and/or slow  $\Rightarrow$  Coriolis important (because  $Ro \ll 1$ ), classify as inertial waves  
 $\rightarrow$  e.g. Rossby waves
- ▶ small-scale and/or fast  $\Rightarrow$  Coriolis unimportant  
 $\rightarrow$  e.g. internal gravity waves
- ▶ somewhere in between? **Poincaré** or **inertia-gravity waves**

$$\omega = \pm \sqrt{f_0^2 + gH(k_x^2 + k_y^2)}$$

$\rightarrow$  for  $gH(k_x^2 + k_y^2) \gg f_0$ , recover gravity waves

$\rightarrow$  for  $gH(k_x^2 + k_y^2) \ll f_0$ , recover inertial waves

# Rossby deformation radius

$$\omega = \pm \sqrt{f_0^2 + gH(k_x^2 + k_y^2)}$$

- boundary between gravity and inertial regimes is roughly

$$L_d = \frac{\sqrt{gH}}{f_0}$$

- the **Rossby deformation radius** (for shallow water system)
  - roughly also the boundary where **geostrophic approximation** should hold (see Lec. 8 + 13)
  - estimates in a few slides



# Internal waves

- introduce a useful quantity

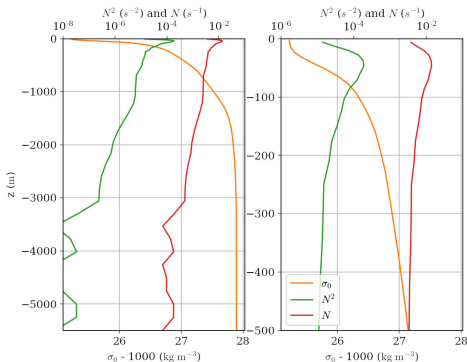
$$N = \sqrt{-\frac{g}{\rho_0} \frac{\partial \rho}{\partial z}}$$

- Brunt–Väisälä or buoyancy frequency (units:  $\text{s}^{-1}$ )

→  $N^2$  normally used

→ note  $\partial \rho / \partial z < 0$  for **stable** stratification, i.e.

$$N^2 > 0$$



**Figure:**  $\sigma_0$  (see Lec. 6) and the associated  $N^2$  and  $N$ .  $N^2 \ll 1$  means weakly stratified (weak density gradients), whilst  $N^2 < 0$  shows unstable stratification (none in this case, but see Lec. 17). See `plot_eos.ipynb`.

**simplistic view (!):**  $\sqrt{gH} \rightarrow N$

# Internal waves

Generally then, **internal inertia-gravity waves** described by

$$\omega = \pm \sqrt{\frac{f_0^2 k_z^2 + N^2 k_x^2}{k_x^2 + k_z^2}} \approx \pm \sqrt{f_0^2 + \frac{N^2 k_x^2}{k_z^2}} \quad (\text{for } |k_z| \gg |k_x|)$$

- ▶ atmosphere and ocean has  $N/f_0 = O(10^1 \text{ to } 10^2)$ 
  - so really we have **gravity waves influenced by rotation**
  - refer to them here as **internal waves**

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- ▶ note that  $|f_0| \leq |\omega| \leq |N|$ 
  - since  $0 \leq k_{x,z}^2 / (k_x^2 + k_z^2) \leq 1$
  - frequency is **much lower** than gravity waves

# Internal waves

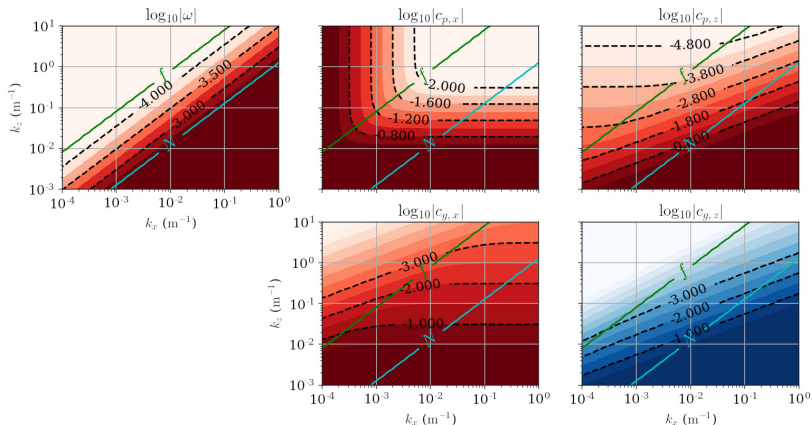
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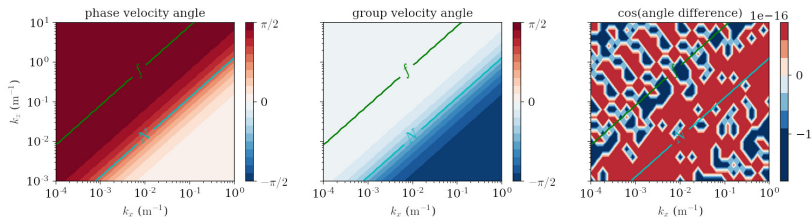
**internal tides to be seen as internal waves** (Lec. 18)

# Internal waves



**Figure:** Inertial-gravity waves (with the  $k_z \gg k_x$  approximation)  $\omega$ ,  $c_{p,x}$ ,  $c_{p,y}$ ,  $c_{g,x}$  and  $c_{p,y}$  as a log-log plot in  $(k_x, k_z)$  space, with  $f = 5 \times 10^{-5}$  and  $N = 3 \times 10^{-3}$  (oceanic relevant values). The contours denote the exponent  $x$  of  $|10^x|$  and the colour shading denotes the sign (more blue = more negative *actual* values rather than exponents, more red = more positive *actual* values rather than exponents); since  $k_x$  and  $k_z$  is chosen to be positive, everything except  $c_{g,z}$  is positive. Contours of  $f$  and  $N$  plotted with an offset plotted to show the boundary beyond which everything is either gravity waves or inertial oscillations. See `waves.ipynb`.

# Internal waves



**Figure:** Inertial-gravity waves (with the  $k_z \gg k_x$  approximation) phase velocity  $c_p$  angles and group velocity  $c_g$  angles (in radians, relative to the horizontal, and note  $\pi/2 = 90^\circ$ ). The final panel shows  $c_p \cdot c_g = |c_p||c_g|\cos\theta$  (which is zero up to rounding errors). Contours of  $f$  and  $N$  plotted with an offset plotted as in the previous diagram. See `waves.ipynb`.

- note that, for inertial-gravity waves (left as a bonus exercise),

$$c_p \cdot c_g = 0$$

→ i.e. phase and group velocities are **perpendicular** to each other (see Lec. 4)

# Deformation radius

- ▶ boundary given again by the **Rossby deformation radius** (for the continuously stratified case)

$$L_d = \frac{NH}{f}$$

→  $L_{d,\text{atmos}} = O(1000 \text{ km})$ , scale of **cyclones** and **anti-cyclones**, i.e. weather systems form (synoptic structures)

→  $L_{d,\text{ocean}} = O(50 \text{ km})$ , scale of **ocean eddies**

- ▶ latitude (through  $f$ ) and  $H$  dependent
  - smaller  $L_d$  for **high** latitudes and **shallow** regions
  - consequence for **geostrophic approximation?** (e.g. shelves and coasts, see Lec. 21 + 22)

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- ▶ internal  $L_d$  defined analogously (normally smaller than above)



# Kelvin waves (more on this in Lec. 18, 21 + 22)

A type of boundary wave

- ▶ need  $f$  and a **boundary**
  - could be land (**coastal Kelvin waves**) (see Lec. 18, 21 + 22)
  - could be a **wave guide** (e.g. equator where  $f$  changes sign, **equatorial Kelvin waves**) (see OCES 4001, El-Niño, QBO etc.)
- ▶ needs  $f$  but propagates at the **gravity** wave speed, with

$$\omega = k\sqrt{gH}$$

- **non-dispersive**
- fairly fast (gravity wave speed)

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- ▶ needs  $f$  but propagates at the **gravity** wave speed, with

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- **non-dispersive**
  - fairly fast (gravity wave speed)
- ▶ **NOTE the lack of  $\pm$ !**

# Kelvin waves (more on this in Lec. 18, 21 + 22)

- ▶ boundary introduces **asymmetry** in this case: general solution like

$$\eta \sim e^{\pm f_0 y / \sqrt{gH}} \cos(kx - \omega t)$$

→ take  $y \leq 0$  to be **boundary**, if  $f_0 > 0$  (NH), need minus sign, and vice-versa

→ wave propagates **cyclonically** (same sign as  $f$ )

- ▶ taking  $f_0 > 0$  (NH),

$$\eta \sim e^{-y/L_d} \cos(kx - \omega t),$$

so decay over the  $L_d = \sqrt{gH}/f_0$

## Rossby waves (more on this later)

A (particularly important) type of **inertial** wave

- ▶ requires a **gradient** in background vorticity  
→  $\partial f / \partial y = \beta$  (planetary case)  
→ background flow  $-\partial U / \partial y \sim \nabla \times \mathbf{u}$  (see later and Lec. 17)
- ▶ dispersion relation given by (on  $\beta$ -plane)

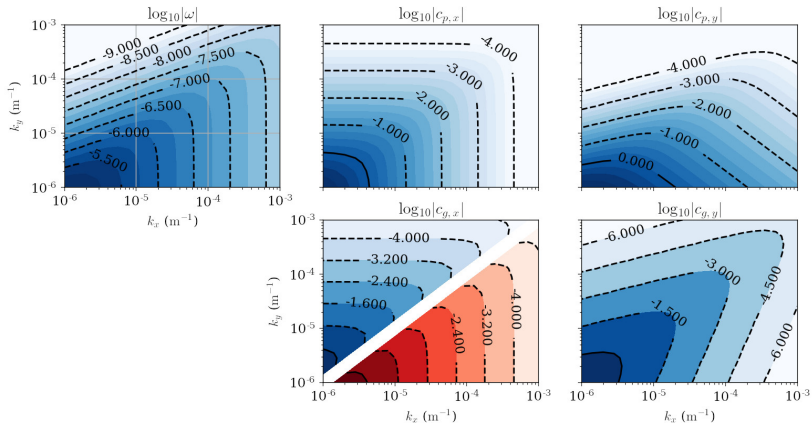
$$\omega = -\frac{\beta k_x}{k_x^2 + k_y^2}$$

→ note that Rossby waves propagate to the **west** (more generally, **retrograde** or against the mean flow) **since**

$$c_{p,x} = \frac{\omega}{k_x} = -\frac{\beta}{k_x^2 + k_y^2} < 0,$$

and **long** waves ( $k_x \ll 1$ ) are **fast(er)**

# Rossby waves (more on this later)



**Figure:** Rossby waves  $\omega$ ,  $c_{p,x}$ ,  $c_{p,y}$ ,  $c_{g,x}$  and  $c_{p,y}$  as a log-log plot in  $(k_x, k_y)$  space, with magnitude also as logs. The contours denote the exponent  $x$  of  $|10^x|$  and the colour shading denotes the sign (more blue = more negative *actual* values, more red = more positive *actual* values); since  $k_x$  and  $k_y$  is chosen to be positive, everything except  $c_{g,x}$  is negative. Choice of  $k_x$  and  $k_y$  correspond to wavelengths roughly between 6 km to 6000 km (Rossby waves are usually seen as planetary-scale waves). See `waves.ipynb`.

# Propagation mechanism: Rossby waves

Rossby waves propagate **west-ward** (or, more generally, **retrograde**)

$$c_{p,x} = \frac{\omega}{k_x} = -\frac{\beta}{k_x^2 + k_y^2} < 0$$

**why?**

# Propagation mechanism: Rossby waves

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**why?**

Key bits to the pictorial/parcel (cf. Lec 5 for temperature) argument:

- ▶ the initial wave **conserves** and carries **vorticity** (spini-ness, recall Lec. 4, 11, 12) into the external environments  
→ these are now vorticity **anomalies**
- ▶ vorticity anomalies induces a velocity/flow (because spini-ness)
- ▶ induced flow seen to self-advect the wave and move it to the **West** (**retrograde** in the general case)

# Propagation mechanism: Rossby wave example

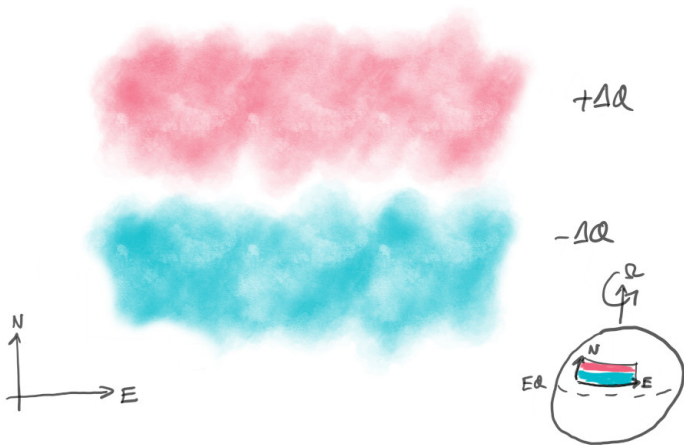


Figure: Rossby wave propagation schematic.



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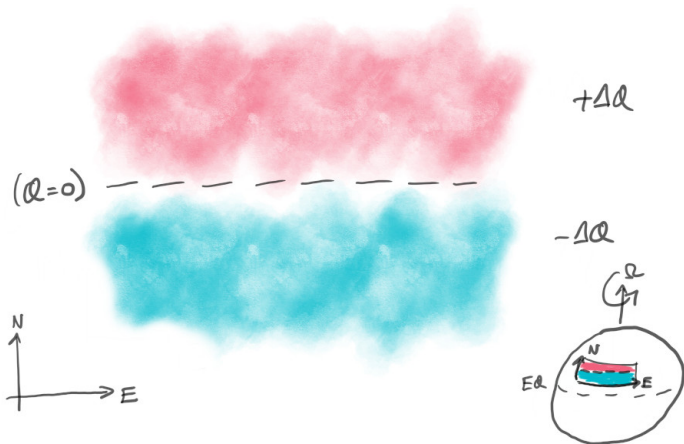


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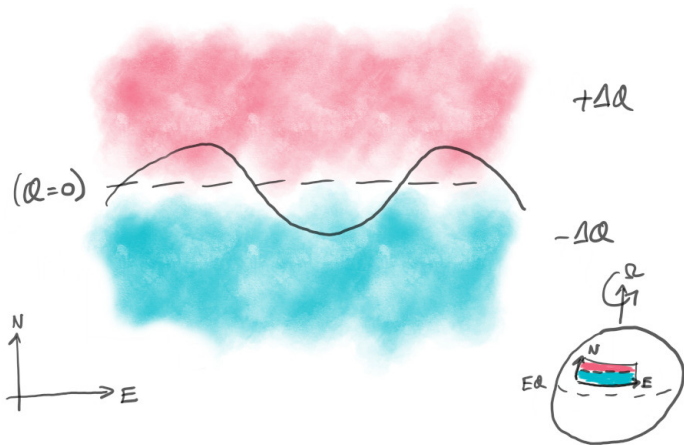


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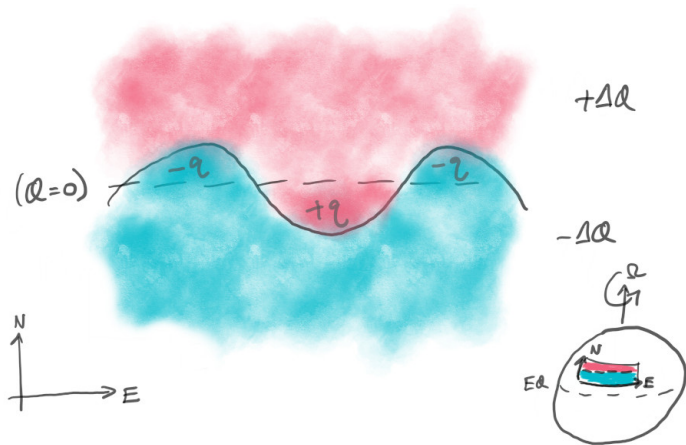


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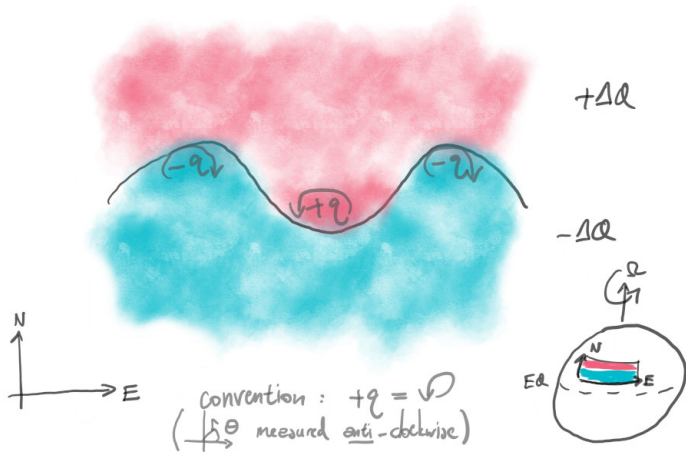


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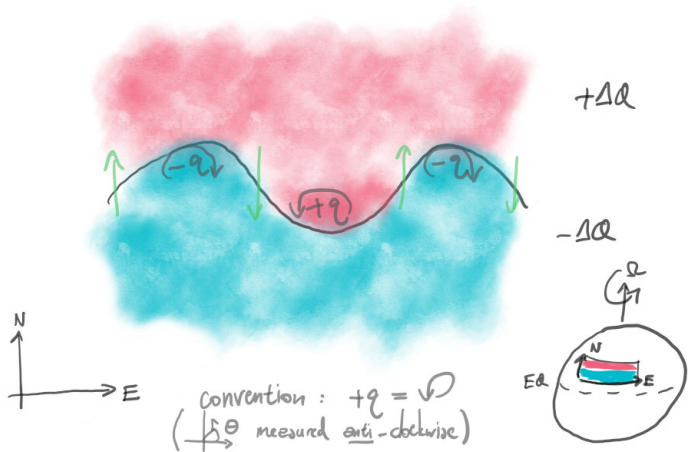


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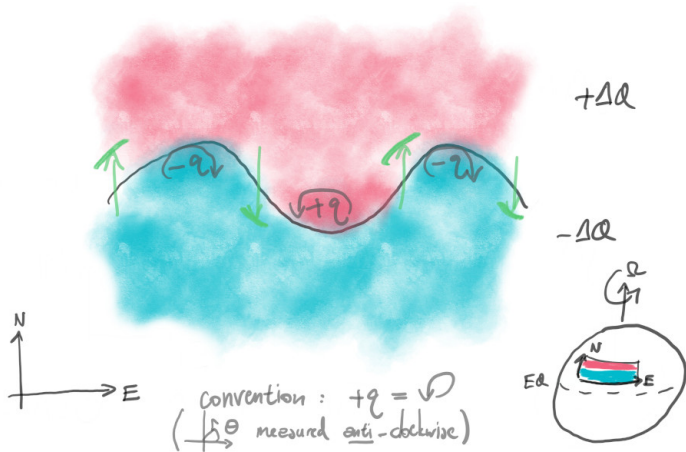


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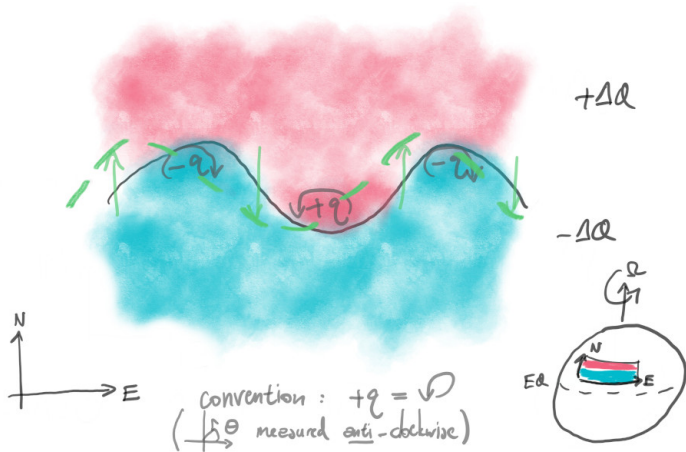


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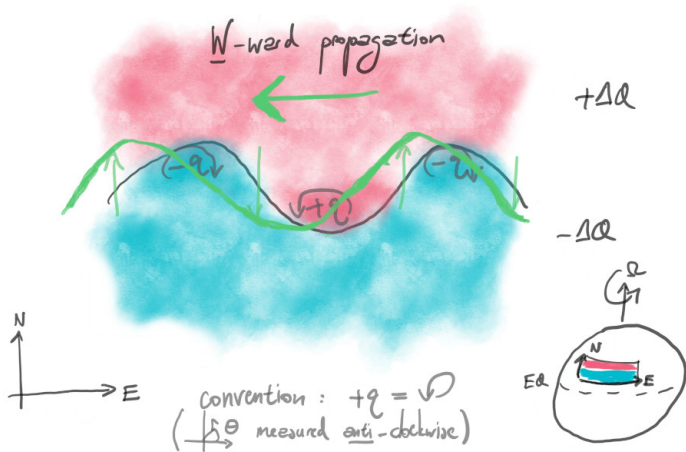


Figure: Rossby wave propagation schematic.



# Summary

- gravity waves  
(gravity/buoyancy)
- inertial waves (Coriolis)
- inertial-gravity waves (general)

→ internal waves have

$$|f| \leq |\omega| \leq |N|$$

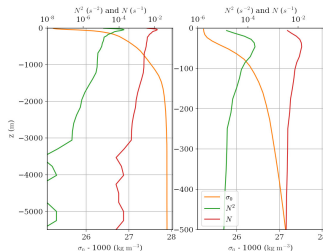


Figure:  $\sigma_0$  (see Lec. 6) and the associated  $N^2$  and  $N$ . See `plot_eos.ipynb`.

- Brunt–Väisälä or buoyancy frequency  $N$

$$N = \sqrt{-\frac{g}{\rho_0} \frac{\partial \rho}{\partial z}}$$

→ measure of stratification strength (see also Lec. 17)

# Summary

- ▶ parcel argument for **west-ward Rossby wave** propagation
  - conservation of **vorticity**
  - vorticity anomalies induces flow
  - self-advecting

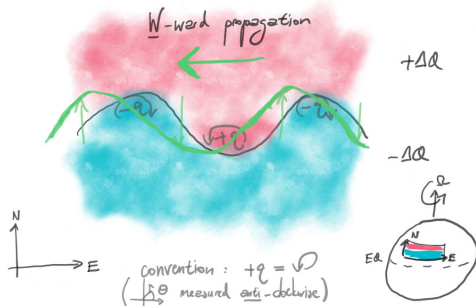


Figure: Rossby wave propagation schematic.

- ▶ generalisations exist (e.g. internal gravity waves in Harnik *et al.*, 2008, *J. Atmos. Sci.*)

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  - conservation of **vorticity**
  - vorticity anomalies induces flow
  - self-advecting

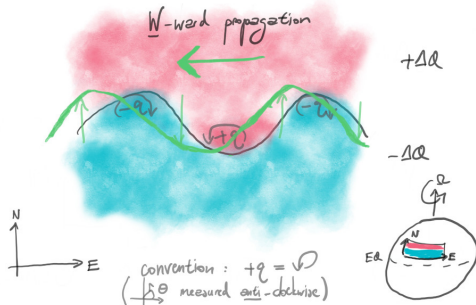


Figure: Rossby wave propagation schematic.

- ▶ generalisations exist (e.g. internal gravity waves in Harnik *et al.*, 2008, *J. Atmos. Sci.*)
- ▶ two such waves interacting? (see Lec. 17)
  - potential for **instabilities**