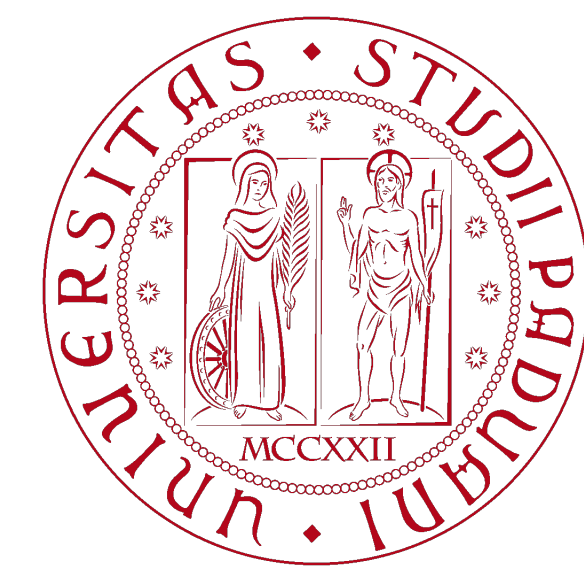


# Bayesian nonparametric multiscale mixture models via Hilbert-curve partitioning

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## Abstract

We consider the problem of flexible **nonparametric density estimation** using mixtures of densities.

We are motivated by **astrological applications**, where galaxies might be clustered based on their colour spectrum.

We rely on a **multiscale mixture model for the density** in order to **cluster observations at different resolutions**

The multiscale structure is described by using a **sequence of Hilbert curves** in order to **map the multivariate space to a binary tree**

The resulting mixture is **flexible** and can **adapt** very well **to the underlying smoothness** of the data

## Background

### Bayesian nonparametric univariate multiscale models

Univariate multiscale models define a mixture of kernels indexed by a binary tree,

$$f(y) = \sum_{s=0}^{\infty} \sum_{h=1}^{2^s} \pi_{s,h} \mathcal{K}(y; \mu_{s,h}, \Omega_{s,h}),$$

where  $(s, h)$  **corresponds to a node of a binary tree** (Figure 1) and  $\mathcal{K}$  is a scale and location kernel.

$\mathcal{K}$  is **increasingly concentrated** as  $s$  increases, and the location parameter  $\mu_{s,h}$  should **span the whole space** as  $h$  moves between the values  $1, \dots, 2^s$ .

A nonparametric prior distribution for the mixture weights  $\pi_{s,h}$  has been developed by Canale and Dunson [2016], whereas a Bayesian nonparametric univariate mixture of kernels has been introduced by Stefanucci and Canale [2021].

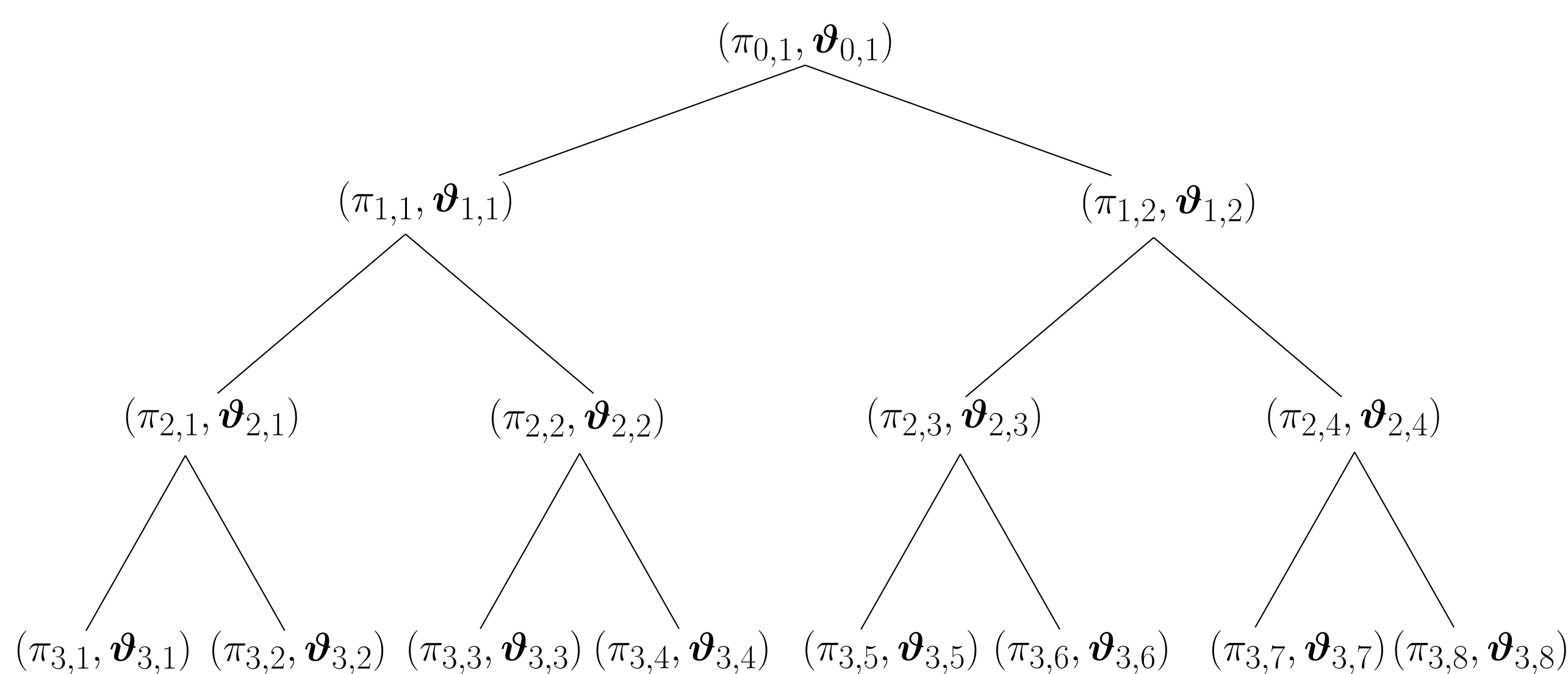


Figure 1: Binary tree associated to the multiscale model.

© Pro: flexibility and adaptability to data smoothness  
© Con: how do we generalize the binary tree to multivariate kernels?

## Problem

Generalizing the model to the multivariate case  $\mathbb{R}^d$  seems hard, because the computational complexity of a  $d$ -tree will explode very quickly.

## Solution: Hilbert curve partitioning

We partition the location space  $\Theta_{\mu}$  so that we span the whole space in  $h$  and the partitions are nested,

$$\Theta_{\mu} = \bigcup_{h=1}^{2^s} \Theta_{\mu;s,h}, \quad \Theta_{\mu;s,h} = \Theta_{\mu;s+1,2h-1} \cup \Theta_{\mu;s+1,2h}.$$

The partition is done using the following two-stage procedure based on the Hilbert curve [Hilbert, 1891]:

- Partition the cube  $[0, 1]^d$  using the Hilbert curve (Figure 2).
- Apply conditional quantiles of  $G_0$  to the extremes of each subcube to obtain the partition.

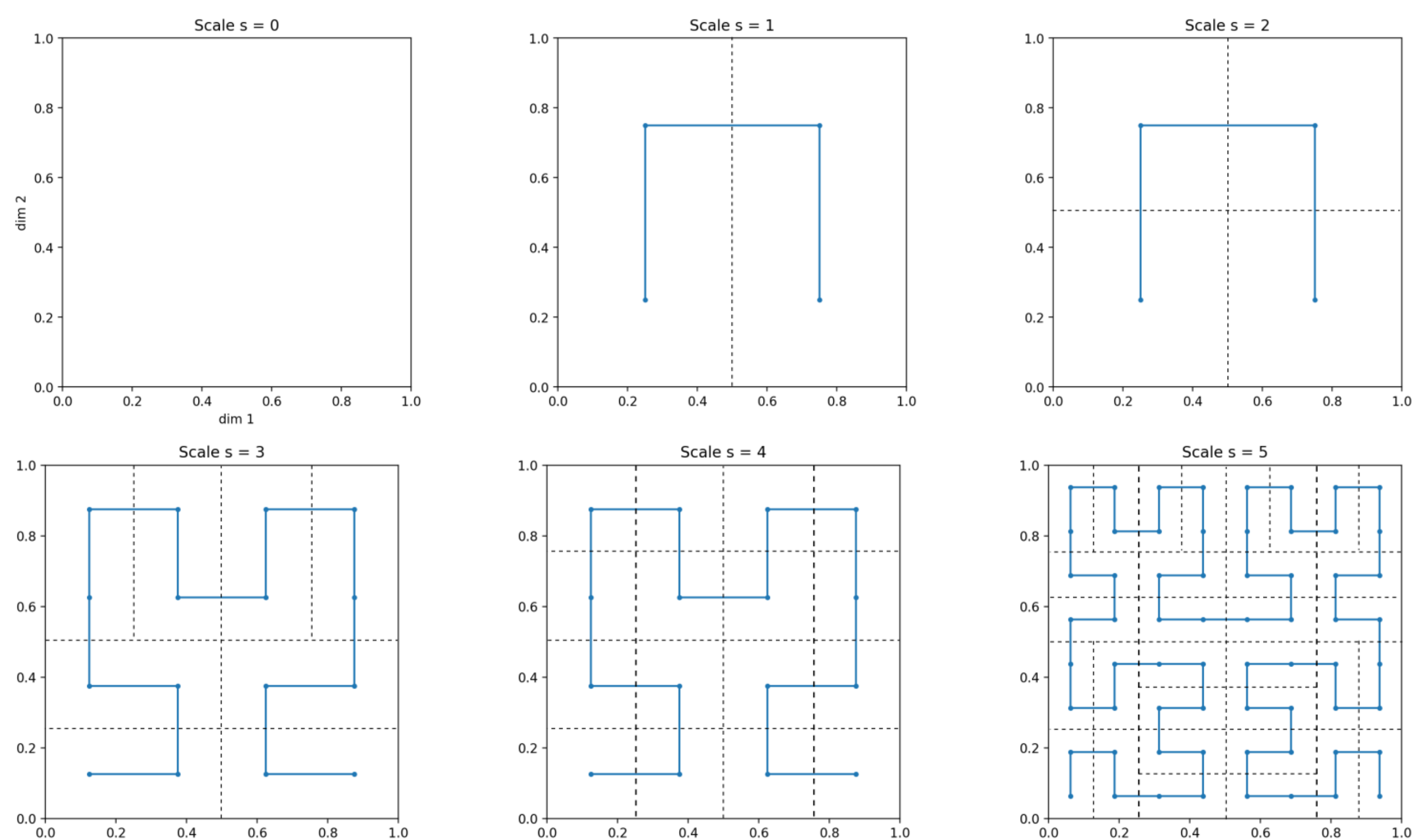


Figure 2: Dyadic partition of  $[0, 1]^2$  obtained by the application of the Hilbert curve, for  $s = 1, \dots, 5$ .

Sampling at each node from  $G_0$  truncated to  $\Theta_{\mu;s,h}$  yields the **prior for the location parameter**,

$$\mu_{s,h} \sim G_0 \cdot \mathbb{1}_{\Theta_{\mu;s,h}}.$$

The scale parameters  $\Omega_{s,h}$  are sampled from a distribution  $H_0$  **scaled by a deterministic monotone decreasing sequence** in  $s$ ,

$$\Omega_{s,h} = \text{diag}(c(s), \dots, c(s)) W_{s,h}, \quad W_{s,h} \stackrel{\text{iid}}{\sim} H_0.$$

## Interpretation

- Nodes higher in the tree correspond to **coarser kernels** whereas deeper nodes correspond to **more localized kernels**. The posterior adapts the kernels to the smoothness of the data.
- We generalize a property of Stefanucci and Canale [2021] to show that the random location measure  $G = \sum_{s=0}^{\infty} \sum_{h=1}^{2^s} \pi_{s,h} \delta_{\mu_{s,h}}$  is **centered around  $G_0$  a priori**,

$$\mathbb{E}[G(A)] = G_0(A) \quad \text{for all } A \subseteq \Theta_{\mu}.$$

- Also, we have control on the **prior average depth** of the binary tree by solving a

## Performance in simulated datasets

We use the multiscale model with Gaussian prior on the location parameters and inverse Wishart prior on the scale parameter to have efficient Gibbs sampling updates [Kalli et al., 2011].

- Scenarios: 1) mixture of normals/skew normals, 2) multiscale mixture.
- Competitor: Pitman-Yor mixture model.
- Mean posterior predictive distribution for the multiscale model in Figure 3.

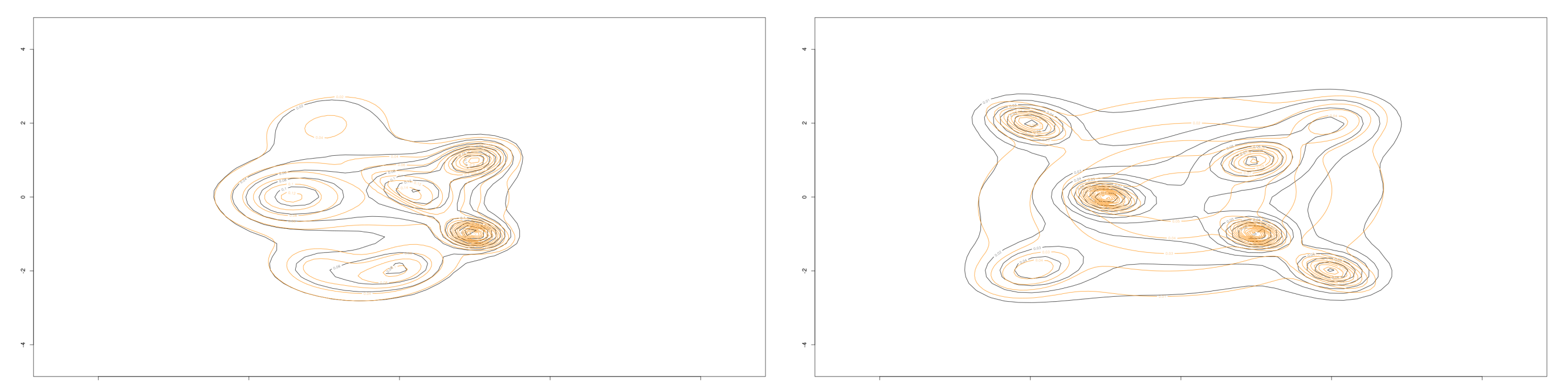


Figure 3: True density function (black) and mean posterior predictive distribution (orange) for the multiscale model.

## Application to the Galaxy dataset

- The galaxy marginal distributions (Figure 4) are correctly estimated, except for the redshift which seems more problematic.
- The binary tree with nodes proportional to the estimated posterior median weight distribution (Figure 5) gives **information on the location of the clusters within the space**.

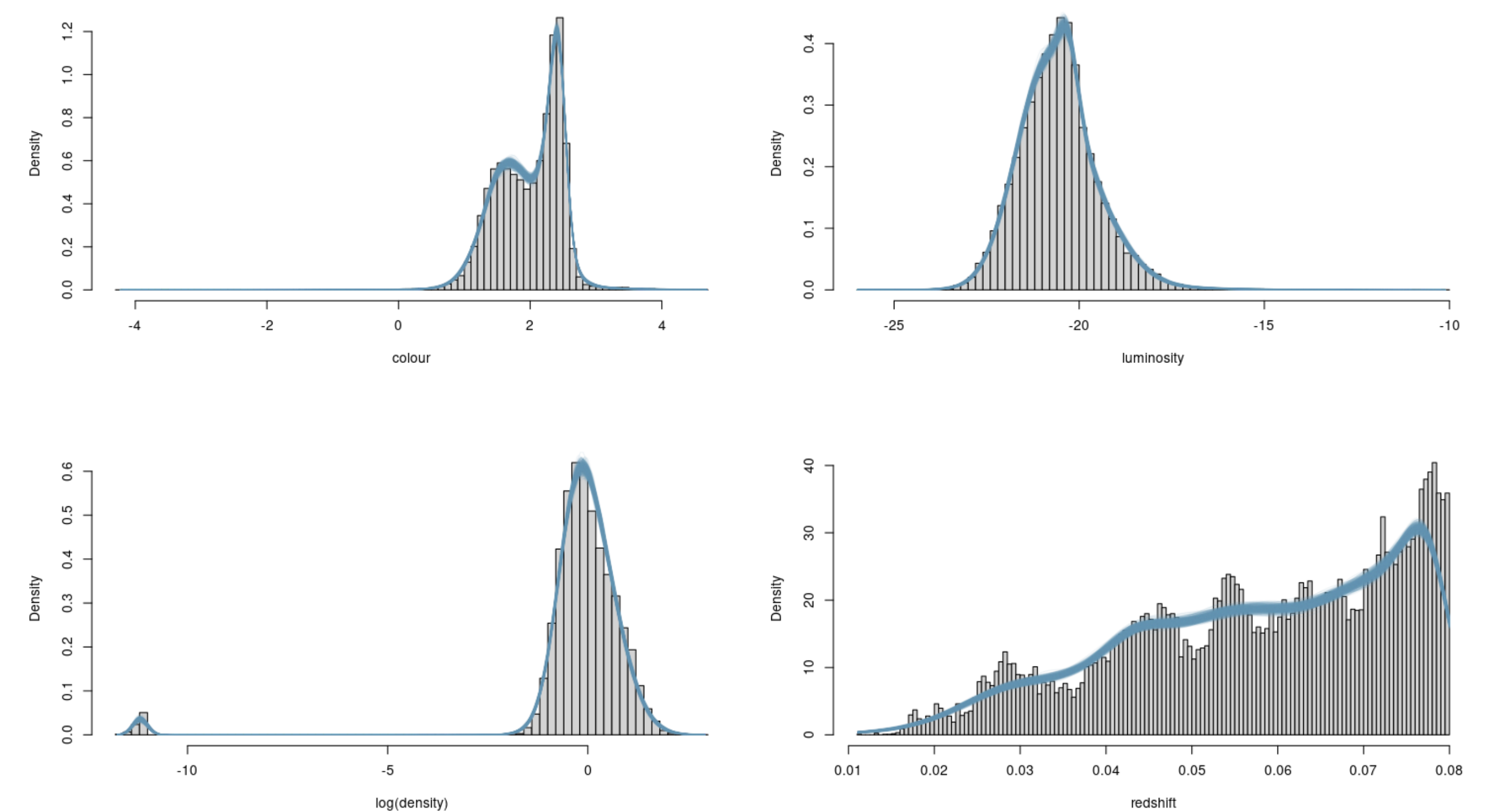


Figure 4: Posterior predictive distribution for the marginal distributions of the Galaxy dataset.

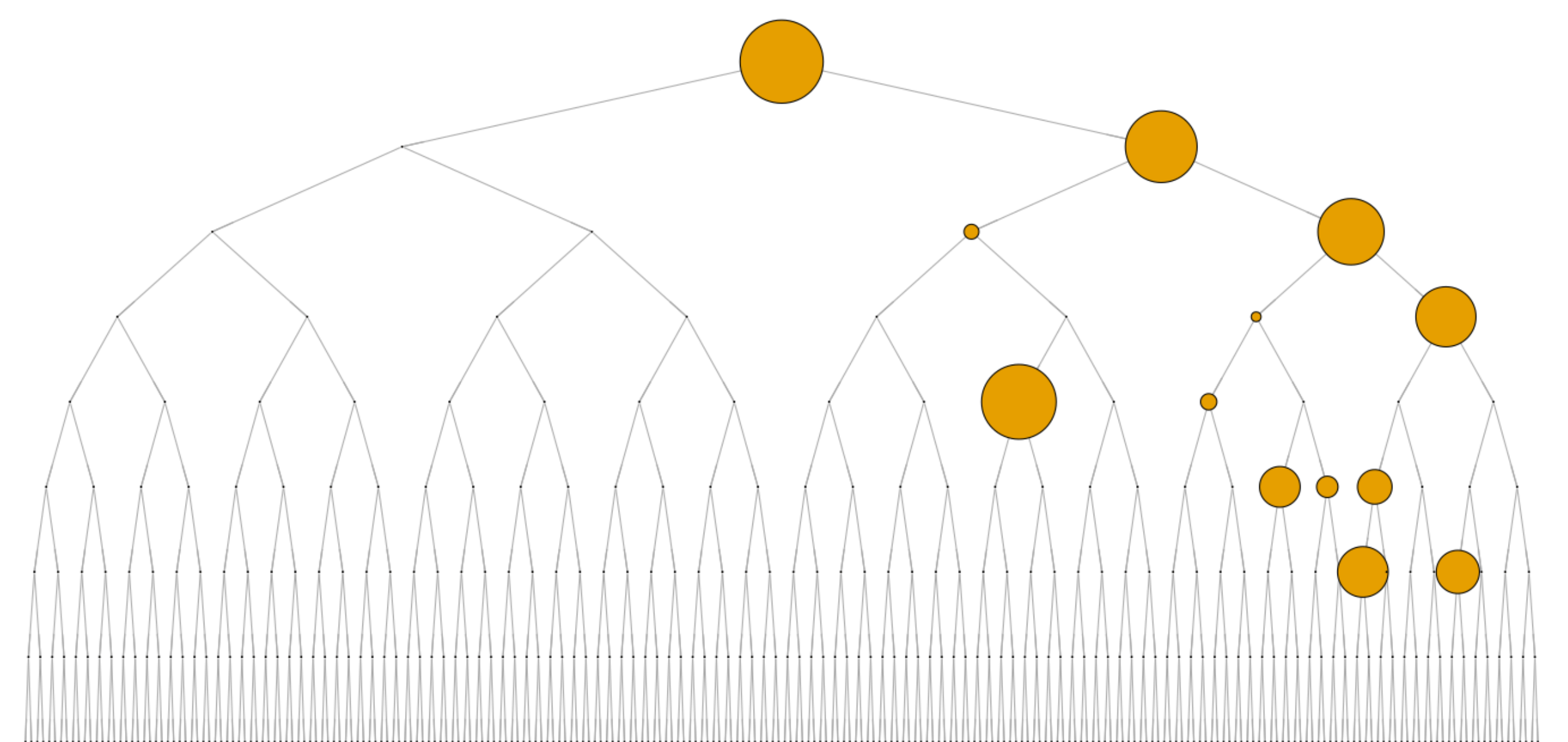


Figure 5: Multiscale binary tree with circles drawn proportionally to the median of the posterior distribution of the weights  $\pi_{s,h}$ .

## References

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