

Article

OpenMetBuoy-v2021: an easy-to-build, affordable, customizable, open source instrument for oceanographic measurements of drift and waves in sea ice and the open ocean

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Abstract: There is a wide consensus within the polar science, meteorology, and oceanography communities that more in-situ observations of the ocean, atmosphere, and sea ice, are required to further improve operational forecasting model skills. Traditionally, the volume of such measurements has been limited by the high cost of commercially available instruments. An increasingly attractive solution to this cost issue is to use instruments produced in-house from open source hardware, firmware, and post processing building blocks. In the present work, we release the next iteration of the open source drifter and waves monitoring instrument of Rabault et. al. (see "An open source, versatile, affordable waves in ice instrument for scientific measurements in the Polar Regions", *Cold Regions Science and Technology*, 2020), which follows these solution aspects. The new design is both significantly less expensive (typically by a factor of 5 compared to our previous, already cost effective, instrument), much easier to build and assemble for people without specific microelectronics and programming competence, more easily extendable and customizable, and two orders of magnitude more power efficient (to the point where solar panels are no longer needed even for long term deployments). Improving performance and reducing noise levels and costs compared with our previous generation of instruments is possible in large part thanks to progress from the electronics component industry. As a result, we believe that this will allow scientists in geosciences to increase by an order of magnitude the amount of in-situ data they can collect under a constant instrumentation budget. In the following, we offer 1) detailed overview of our hardware and software solution, 2) in-situ validation and benchmarking of our instrument, 3) full open source release of both hardware and software blueprints. We hope that this work, and the associated open source release, may be a milestone that will allow our scientific fields to transition towards open source, community driven instrumentation. We believe that this could have a considerable impact on many fields, by making in-situ instrumentation at least an order of magnitude less expensive and more customizable than it has been for the last 50 years, marking the start of a new paradigm in oceanography and polar science, where instrumentation is an inexpensive commodity and in-situ data are easier and less expensive to collect.

Keywords: Waves in ice; Open Source; Instrumentation; In situ measurements; Oceanography; Waves; Sea surface drift; Open Ocean; Wave spectrum; Low cost.

1. Introduction

Our inability to accurately capture climatological changes of sea ice in the polar seas has created renewed interest in the dynamic interaction between sea ice and waves. This has resulted in the last few years into a number of studies that investigate the coupling between sea ice and the ocean through theoretical considerations [1–7], laboratory experiments [8–13], and field experiments [14–25]. Despite the advances that these studies bring, there is a growing consensus that further progress in the field can only be achieved through the collection of more observations of waves in ice. In particular, phenomena related to floe size distribution [4,26–28], energy dissipation due to turbulence and collisions [29–32], and sea ice breakup [33–36], are still unperfected modeled and understood, and advancing the state-of-the-art would require additional direct observations from the ice.

Unfortunately, one faces a number of challenges in reproducing the complexity of wave-ice interaction either in experiments (due to, for example, scaling issues [9], and the diversity and realism of sea ice conditions that can be obtained in the laboratory compared with the field [13]), or in models (there, also, due to the complexity and multi-scale properties of waves in ice and ice floe size distribution [37,38]). Arguably, the current state of the art in wave-ice interaction parametrization and modeling consists in a variety of (involved and mathematically advanced) models which formulations are relatively loosely based on some specific archetypical sea ice conditions and the associated physical mechanisms, with a number of "tuning parameters" that are empirically fitted to experimental or laboratory data [39,40]. This approach is quite ad-hoc and brittle, and this is in stark contrast with the importance of waves in ice and their impact on both weather and climate.

As a consequence, having access to large datasets of high temporal and spatial resolution direct observations of waves in ice, is a key ingredient to improving both small scale wave-ice interaction parametrization, large scale coupled wave-ice models, and our understanding of the weather and climate dynamics in the polar regions. There are a number of recent works that highlight the importance of the volume of such datasets for unveiling wave-ice interaction mechanisms. For example, Voermans *et al.* [15] showed that wave conditions that result in sea ice breakup events can be quite accurately described with a simple threshold mechanism, but that in order to be able to observe such a threshold, a relatively large aggregated dataset was necessary.

There are relatively few methods available for measuring waves in ice. The most common and established method considers the deployment of instrumentation in situ, providing surface truth of the sea ice motion under the influence of waves. Such instruments typically use either an Inertial Motion Unit (IMU, i.e. a combination of accelerometer, gyroscope, and magnetometer, together with some data fusion processing, e.g. Kohout *et al.* [41], Rabault *et al.* [42]), or a high accuracy GPS [43,44] to measure the sea ice motion and compute wave spectra. Less common techniques are the use pressure sensors to measure the wave-induced pressure fluctuation to derive wave observations [24]. Remote sensing is a promising avenue to obtain observations of waves in ice across large spatial scales [45,46] of far larger extent compared to what can be achieved with in situ instrumentation, though the number of wave parameters that can be retrieved, and the accuracy that can be achieved, is still limited and a topic for open research. Moreover, remote sensing requires considerable more in situ observations for calibration and validation. As a consequence, the most established and accurate method today is still to deploy instrumentation in situ to measure waves in ice.

Performing in situ measurements of waves in ice is, however, a challenging and costly operation. Commercial instruments have traditionally been extremely expensive, and sometimes of varying quality, with typical prices after taxes in the 7.5-50kUSD range. The situation got somehow better with the emergence of new actors, such as Sofar and their Spotter buoy [47], which took down the price to a typical 6-8kUSD cost per instrument, once

import taxes and satellite communication fees are factored in. Still, the Sofar Spotter, while gaining from being a very modern design, as well as relatively low cost, is not too well adapted for use in the polar regions, as it was designed primarily for measuring open water wave statistics, and, correspondingly, relatively large waves. In particular, its battery life limits the duration of deployments in the polar night [44]. In addition, its GPS-based wave measurement method, while having many advantages in open water (such as accurate directional wave spectrum measurement), has a lot higher noise for the typically small waves observed in the ice than what can be achieved with a modern Inertial Motion Unit (IMU).

As a consequence of these challenges, a number of groups have developed in-house, custom instrumentation for measurements of waves in ice [e.g., 41,42,48]. A very brief overview is provided in Fig. 1. This is, in our opinion, a situation that is so far quite inefficient. Indeed, a number of groups develop their own closed source solution, in parallel of each other, wasting a large amount of engineering resources in form of working hour and prototyping costs in the process.

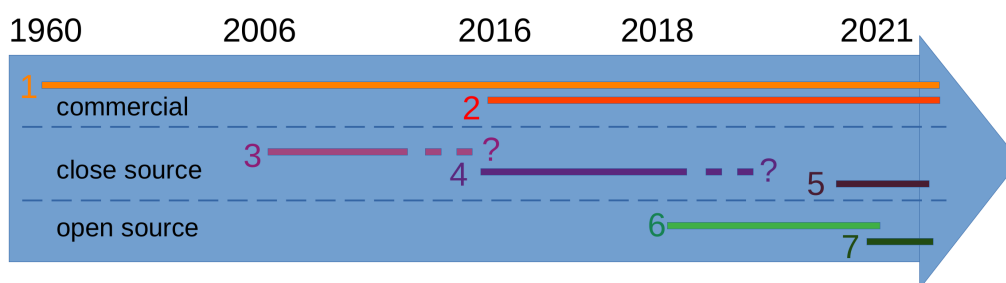


Figure 1. Overview of waves in ice instruments, distinguishing between i) commercial instruments, ii) close source instruments, iii) open source. Commercial instruments have been available for a long time, in particular the Datawell Waverider series of instruments (1), which goes back all the way to the 1960s [49]. More recently, the Sofar Spotter (2) has become a viable, lower-cost alternative, though there are a number of limitations in particular regarding battery autonomy [50], and the price. A number of groups have developed satellite-connected autonomous instruments that have been described in research articles, but that remained closed source. A few instruments we know about were released in 2006 (3) [48,51], 2016 (4) [41], 2021 (5) [25]. By contrast, our group has decided to release our designs as open source, so that can be freely used and improved by other groups, starting in 2018 (6) [52] and now with the present design (7).

An increasingly promising direction is the development of open source instrumentation, which allows to both reduce hardware costs and mutualize the instrumentation development workload. This approach was recently adopted by Rabault *et al.* [52] and has since its development in 2015 undergone several updates (the full lineage of our series of open source instruments is presented in Fig. 2). In the following, we will refer to the design described in Rabault *et al.* [52] as the “v2018 instrument”, or “v2018” in short. Since the initial deployment of the first v2018 instruments Rabault *et al.* [52], the low-cost and open-source fundament of the instrument has promoted usage and international collaboration across the world. The v2018, and adaptations thereof, have been deployed in both the Arctic and Antarctic to study sea-ice drift [53], wave attenuation and dispersion [20], and wave-induced sea ice break-up [15]. Through these collaborations and technological advancements, points of improvement were identified, which supported to advance the v2018 and develop the latest “v2021” version.

These challenges of high instrumentation costs and needs for more in-situ, ground truth data, are not specific to the waves in ice community. Similarly, the rise of low cost (and possibly open source) electronics is an evolution that is increasingly attracting the attention of both research groups and public agencies across the world. For example, this trend can be identified also in the technical development and history of the “Swift drifters” family, among others, which are also advancing towards open source, low cost, “do it yourself”

technical solutions [54–56], focusing on measuring waves in the open ocean. Another example of this trend is visible in recent works that focus on building small, inexpensive, bio-degradable surface drifters [47]. Similarly, the Defense Advanced Research Projects Agency (DARPA) is actively pursuing similar efforts with its "Ocean of Things" project [57], which aims at "deploying thousands of low-cost, environmentally friendly, intelligent floats that drift as a distributed sensor network" (quote from the Ocean of Things project webpage, accessed Dec. 2021).

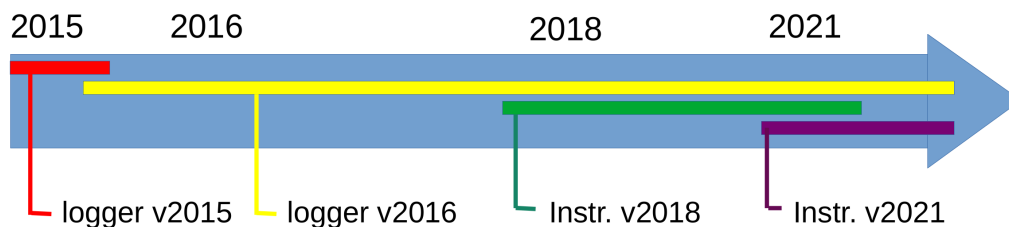


Figure 2. Full lineage of our present open source instrument. Developing reliable autonomous instruments is a long term effort. The development of our instrument can be tracked back to 2015 [42] and 2016 [22], though the instruments used at the time were simple loggers. Using loggers is still an attractive solution when the instruments can be recovered though, and the logger v2016 design is still in use today. By contrast our first truly autonomous, satellite-enabled instrument was designed in 2017-2018 [52], and the present instrument is the second iteration around this design.

This paper is organized as follows. First, we discuss the general design and features of the instrument v2021. Then, we present a detailed validation of the instrument v2021, both i) in a preliminary deployment which aimed at validating the general concept of the instrument and the power efficiency of the design, ii) in laboratory experiments that tested the noise threshold of the waves measurements, iii) in a full-scale deployment in the Russian Arctic, iv) in a full-scale deployment in the Caribbeans. Finally, we offer a few words of conclusion and we present a roadmap for future development. All hardware blueprints, embedded software, post processing codes, as well as detailed instructions, are released as open source on Github, as indicated in the Appendix A.

2. General design and features of the instrument v2021

2.1. Microcontroller

Unlike the instrument v2018, and unlike all other wave measuring instruments that we are aware of, the v2021 is based solely on a high performance microcontroller for performing both data acquisition and in-situ data processing. This is a major shift from previous designs, which usually resorted on some sort of embedded microcomputer for performing the computation of wave spectra (typical previous solutions, such as those based on a Raspberry Pi, run a stripped down version of Linux or Windows), with the measurements themselves being performed either by the same microcomputer, or by a separate, lower power microcontroller. This traditional approach has a dramatic effect on power consumption. A typical microcomputer consumes around 1 to 2 W (for example, the Raspberry Pi used in the v2018 consumes typically up to around 350mA while running at a 5V voltage), while a modern microcontroller can consume as little as 1 to 2 mW (the microcontroller used in the v2021 consumes typically 500 μ A while running at a 3.3V voltage). This factor 1000 in power efficiency allows for significant simplifications and flexibility in the instrument design. This is of particular importance to instrument performance in the polar regions where the battery is generally the main limiting factor in the field experiment duration. In addition, microcomputers will rely on a complete operating system for running their workloads. These are complex pieces of software, with several possibilities for things to go wrong, and require the user to implement a number of techniques to avoid locking of the system. By contrast, microcontrollers can be programmed

without the use of any operating system, which allows to build more robust embedded systems that are more reliable in demanding remote environments.

To this end, the instrument v2021 uses an Ambiq Apollo 3 BLU ultra low power microcontroller unit [58], embedded within a board from the Sparkfun Artemis ecosystem [Artemis Global Tracker, 59]). This allows to use the complete Sparkfun Artemis ecosystem toolchain to simplify the programming of the Ambiq Apollo using high level C++ code. Effectively the user is able to use off-the-shelf, modular, high level libraries and does not need to perform low level drivers and firmware development. The Ambiq Apollo microcontroller used is an ARM Cortex M4-F implementation that runs at 48MHz (with a 96MHz boost mode), uses 500 μ A of sustained power at 3.3V, and features a complete floating point unit for performing fast mathematical operations. Since this processor is an ARM Cortex M4 implementation, the full ARM-developed libraries for digital signal processing, including a wide range of spectral analysis tools, are available out of the box through the Sparkfun toolchain which includes a copy of the CMSIS library (this features, in particular, real and complex FFTs, cross correlation computations, and similar, which are available as highly optimized, ready-to-use functions). The total non volatile memory available for storing the program is amounting to 1MB, and the RAM has a size of 384kB. To put this in perspective, storing 20 minutes of IMU data recorded at 10Hz into RAM, processing the data, and sending the processed data through Iridium, occupies around one third of the total memory available on the chip. This means that a lot of additional customized functionality can be added in the future to the v2021.

2.2. Wave data acquisition and on-board processing

The wave data acquisition is performed by a 9 degrees of freedom (9-dof) sensor, which performs high-frequency measurements of acceleration, angular rates, and magnetic field on the X, Y, and Z axis of the sensor chip. An industrial quality, thermally compensated, low noise sensor is used for this (the ST ISM330DHCX together with the LIS3MDL). At present, only the vertical wave spectrum is computed, and the absolute orientation (provided by the magnetometer) is not used but can be used in the future to derive directional information. The raw accelerometer and gyroscope data are sampled at approximately 800Hz, and time averaged into a 100Hz low-noise sensor value. When performing this averaging, a 3-sigma filtering stage is used, i.e., individual measurements that deviate from the rest of the current 800Hz sample by more than 3 standard deviations are rejected. This means that each set of 8 samples at 800Hz, spanning over 10 milliseconds, is filtered (with the n-sigma filter) and averaged. This allows to discard occasional bad quality measurements (in our experience, a few such measurements are obtained in a 20 minutes measurement segment; this is most likely inherent to the kind of sensor used, and / or may come from rare irregularity in the functioning of either the 9dof sensor or the firmware used for communicating with it), to avoid any risk of signal aliasing, and to reduce the noise level by performing averaging. This low-noise sensor value averaged at 100Hz is then fed in real time to a Kalman filter run on the Ambiq Apollo microcontroller at 100Hz. The Kalman filter itself is provided by an Attitude and Heading Reference System (AHRS) C++ embedded open source library, provided to the community by Freescale Semiconductors. The Kalman filter implementation performs data fusion for 9-axis MEMS input (i.e. 3-axis accelerometer, gyroscope, and magnetometer), and produces an accurate estimate for the absolute orientation of the sensor in a North, East, Down frame of reference. The Kalman filter output is used for computing the vertical component of the wave acceleration at 100Hz by projecting the acceleration from the X, Y, Z frame of reference onto the vertical direction, following the absolute orientation information provided by the Kalman filter as a quaternion, and subtracting the constant gravity acceleration. Finally, the processed 100Hz vertical acceleration measurement is time averaged to 10Hz. This output frequency is sufficient to measure either wind waves (typical peak frequency 0.2Hz), swell (typical peak frequency from 0.08 to 0.10Hz), and waves in ice (typical peak frequencies corresponding to swell, or slightly lower, due to the increased damping rate at higher frequencies),

and is a value commonly used in other waves in ice instruments [52]. If necessary, this output sampling frequency can be increased by the user, and the resources available on the microcontroller could allow for at least 20Hz sampling rate. The whole procedure is implemented through a couple of classes in the code and performed instantaneously. The reader is referred to the open source implementation released on Github (see Appendix A) for more details.

Resorting to a simple 9-dof sensor and running the Kalman filter on the Artemis microcontroller is both much more power effective (total power consumption is around 5mA for the proposed solution, compared with typically 35mA using a dedicated full-feature IMU) and significantly less expensive than using a single discrete Inertial Motion Unit component which performs both tasks on its own (the cost of a high quality 9dof sensor is around 15USD, compared with 50 to 1000USD for a lower accuracy, or similar accuracy, IMU, respectively). Pseudocode for the data acquisition and Kalman filter processing is presented in the description of Algorithm 1.

```

initialize empty BufferVerticalAcceleration10Hz;
initialize KalmanFilter100Hz;
while need more vertical acceleration measurements at 10Hz do
  initialize empty BufferVerticalAcceleration100Hz
  for CurrentVerticalAcceleration100Hz in [0..10] do
    initialize empty BufferAcc[X,Y,Z]800Hz, BufferGyr[X,Y,Z]800Hz,
    BufferMag[X,Y,Z]800Hz;
    for Current9dofMeasurement800Hz in [0..8] do
      measure Acc[X,Y,Z] and append to BufferAcc[X,Y,Z]800Hz;
      measure Gyr[X,Y,Z] and append to BufferGyr[X,Y,Z]800Hz;
      measure Mag[X,Y,Z] and append to BufferMag[X,Y,Z]800Hz;
    end
    3-sigma-average BufferAcc[X,Y,Z]800Hz into Acc[X,Y,Z]Average100Hz;
    3-sigma-average BufferGyr[X,Y,Z]800Hz into Gyr[X,Y,Z]Average100Hz;
    3-sigma-average BufferMag[X,Y,Z]800Hz into Mag[X,Y,Z]Average100Hz;
    update KalmanFilter100Hz using Acc[X,Y,Z]Average100Hz,
    Gyr[X,Y,Z]Average100Hz, Mag[X,Y,Z]Average100Hz;
    obtain QuaternionAbsoluteOrientation100Hz from KalmanFilter100Hz;
    compute CurrentVerticalAcceleration100Hz from
    QuaternionAbsoluteOrientation100Hz, Acc[X,Y,Z]Average100Hz;
    push CurrentVerticalAcceleration100Hz to
    BufferVerticalAcceleration100Hz
  end
  average BufferVerticalAcceleration100Hz into
  CurrentVerticalAcceleration10Hz;
  push CurrentVerticalAcceleration10Hz to BufferVerticalAcceleration10Hz;
end
return BufferVerticalAcceleration10Hz;

```

Algorithm 1: Sampling of vertical acceleration at a frequency of 10Hz in a dedicated buffer, including raw data sampling and pre-averaging, Kalman filtering, and vertical acceleration post-averaging. "Acc" stands for "Acceleration", "Gyr" stands for "Gyroscope", "Mag" stands for "Magnetometer". 3-sigma-average is an averaging filter rejecting measurements deviating from the rest of the sample by more than 3 standard deviations, which is used to discard occasional bad readings. The combination of high-frequency averaging of the raw input from 800Hz to 100Hz, Kalman filtering and vertical acceleration computation at 100Hz, and vertical acceleration averaging from 100Hz to 10Hz, allows to balance the need for computational efficiency (by reducing the number of expensive Kalman filter updates) and accuracy (by running the raw data collection at high-frequency and using all the data available through time averaging).

The wave elevation data processing is also performed on the Ambiq Apollo microcontroller. At the end of a 20 minutes measurement period (or, to be more exact, 20.48

minutes, to obtain a multiple of $2048 = 2^{11}$ samples to allow much faster FFT computation), the array of 10 Hz vertical acceleration data stored in RAM is processed using the Welch method. For this, the signal is split into segments of length 2048 sample points, with a 75% overlap, resulting in a total of 21 segments. This corresponds to segments with length double of what was used in Rabault *et al.* [52]. The motivation for doubling the segment length is to increase the frequency resolution of the output. In addition, we use 75% overlap, instead of 50% overlap used in Rabault *et al.* [52], in order to keep the number of segments large enough to perform a good averaging despite the increase in segment length. The real FFT for the vertical acceleration of each of these segments is computed using the real FFT implementation provided by the ARM digital processing reference library, applying Hanning windowing with an energy-conserving normalization, and these segment FFTs are averaged into the Welch estimate. The computation of the FFT takes just a few tens of milliseconds due to the dedicated floating point unit on the Ambiq Apollo microcontroller, and the efficient FFT algorithm provided by ARM. The corresponding algorithm is summarized in Algorithm 2. The reader who may want to re-implement a similar algorithm is made aware that different FFT libraries may use different renormalization conventions, so that using another FFT implementation may require the use of additional scaling factors. In order to save memory, the Welch averaging is only performed for the set of frequency bins corresponding to frequencies that are relevant for waves, which is typically between $f_{min} = 0.05\text{Hz}$ (i.e. 20s period waves) and $f_{max} = 0.5\text{Hz}$ (ie 2s period waves).

The Welch spectrum with Hanning windowing obtained at this stage is a low-noise estimate for the spectrum of the wave vertical acceleration, PSD_{accD} . From there, the spectrum of the wave elevation η can be obtained following [21]:

$$PSD_{\eta}(f) = PSD_{accD}(f) / (2.0\pi f)^4. \quad (1)$$

Finally, the spectral moments are computed following:

$$m_i = \int_{f_{min}}^{f_{max}} PSD_{\eta}(f) f^i df, \quad (2)$$

and these are used to compute estimates of the significant wave height H_s and the wave period T_z and T_c , following:

$$H_s = 4\sqrt{m_0}, \quad (3)$$

$$T_z = \sqrt{m_2/m_0}, \quad (4)$$

$$T_c = \sqrt{m_4/m_2}. \quad (5)$$

Since the full wave spectrum is transmitted (see the next section), these statistics are simply redundant information (as they can be derived from the transmitted spectrum) that are convenient to use for consistency checks, or for quickly inspecting wave data as they are transmitted. We choose to transmit these quantities to remain consistent with the previous design of Rabault *et al.* [52], as these are well adapted to our needs, but users are free to implement the transmission of other statistics (such as the zero crossing period $T_{01} = m_0/m_1$, or the harmonic period $T_{-10} = m_{-1}/m_0$) if this is a better fit for their needs, owing to the open source nature of the instrument.

We want to remind those readers who may want to re-implement this processing again that different FFT packets may use different conventions, so that additional renormalization factors may be needed. Similarly, some references in the literature use the angular frequency $\omega = 2\pi f$ rather than the frequency f as a dimension for spectra, in which case some additional renormalization is needed in the formula above. Our general recommendation is to test the whole processing algorithm on dummy synthetic data to ensure that no renormalization factor has been forgotten. Moreover, we used additional checks during

```

input: BufferVerticalAcceleration10Hz, length 6 * 2048 samples;
initialize array HanningWelchSpectrum[RelevantIndexRange] with all elements =
0;
// the RelevantIndexRange includes the reduced span of the 2048-point FFT
// that covers frequencies between 0.05Hz and 0.5Hz, ie from fmin to fmax
initialize CurrentSignalSegmentStart = 0;
initialize CurrentSignalSegmentEnd = 2048;
for CurrentWelchSegment in [0..21[ do
    copy BufferVerticalAcceleration10Hz[CurrentSignalSegmentStart..CurrentSignalSegmentEnd[ into
    CurrentFFTInput[0..2048[;
    // apply Hanning windowing
    for CurrentIndex in [0..2048[ do
        | CurrentFFTInput[CurrentIndex] *= 1.63 * [sin( $\pi$  * CurrentIndex / 2048)]2
    end
    compute realFFT of into HanningCurrentFFTOutput;
    HanningWelchSpectrum[RelevantIndexRange] += 2.0 * |
        HanningCurrentFFTOutput[RelevantIndexRange]2 | ;
    // the realFFT of a real input signal is a vector of complex numbers; we need to
    extract the
    // complex magnitude squared of each bin; in our case, a renormalization of
    2.0 is needed since
    // the output is only the half spectrum, i.e. the negative frequencies part of the
    spectrum
    // which is equal to the conjugate of the corresponding positive frequency part
    is omitted
    CurrentSignalSegmentStart += 2048 / 4;
    CurrentSignalSegmentEnd += 2048 / 4;
end
HanningWelchSpectrum[..] = HanningWelchSpectrum[..] / 21;
return HanningWelchSpectrum;

```

Algorithm 2: Algorithm used for computing the Welch spectrum with energy-conserving Hanning windowing. We have decided to use 21 segments with 75% overlap when computing the Welch averaging. Note that the exact value of the renormalization coefficients needed may depend on the specific FFT implementation used and the normalization convention that it defaults to.

the development of the code to make sure that the scaling and processing as a whole is correct, such as verifying that the Parseval theorem holds by checking that the variance of the heave displacement is equal to the integral (with respect to frequency) of the power spectral density.

2.3. Satellite communications

After the recorded data is processed, the data (composed of the Welch wave acceleration spectrum between f_{min} and f_{max} , the estimates for H_s , T_z , and T_c , as well as UTC timestamp information) are packed into a binary packet. In order to reduce the size of the binary packet, the estimates for H_s , T_z , and T_c are transmitted as 32-bit floats, while the Welch frequency bins are discretized into renormalized 16-bit unsigned integers. The renormalization of the Welch bins is performed relatively to the peak value of the spectrum, which is transmitted as a 32-bit float within the packet. This allows to significantly reduce the size of the binary packets and the overall iridium costs while keeping a high accuracy for the data transmitted.

In addition to the wave spectrum, geographical positioning is obtained with a simple GNSS module, and is also used to generate accurate UTC reference times and to periodically re-calibrate the real time clock of the microcontroller to avoid time drift. The GNSS data are packed, buffered, and transmitted using an efficient binary encoding, similar to what is done for the wave data.

The sample rate of both the GNSS and the wave measurements can be adapted through 2-way iridium communication. In addition, the real time clock present on board on the microcontroller is used to make sure that the measurements are performed at fixed hours and minutes. In order to reduce iridium costs and energy consumption, the firmware can pack several binary packets packages together before transmitting these as a single iridium message. As all iridium communications are buffered on the microcontroller, any data not transmitted due to failure in iridium communication can be retransmitted at the next transmission attempt to prevent data loss. When an Iridium communication is established, the instrument transmits a burst of several messages, sending back to the user all the information that are stored in the binary message buffers.

A simple binary protocol decoder, available as a Python module, is provided alongside the code for the instrument firmware. This, together with the web interface or web API offered by modern Iridium providers (in our case, we use Rock Seven Mobile Services Ltd, though other providers would be possible), allows great flexibility and ease-of-use for the end user. Those tasks can easily be automated, and we provide a custom bash script for performing https requests directly to the server of RockSeven, which allows to retrieve all the iridium messages received over a user-selected time span, as a csv database.

2.4. Ongoing instrument variant: cellular communication

A derivate of the satellite based buoy is being developed to communicate through the cellular network instead. Using the same data-processing and setup means that the buoy takes advantage of the previous validation. The purpose of a buoy operating on the cellular network is 1) lower cost (by a factor of 2 to 3 for the hardware, and up to 50 for the communication), 2) higher data-rates: time-series of continuous measurements can be transmitted if necessary.

The buoy will operate under the limitation that it cannot stray far from the coast, and the target areas reflect this: near-shore breaking wave measurements and drift and wave measurements inside fjords in Norway. The buoy will be designed to store as much data as possible locally if network service is temporarily unavailable. If the buoy runs out of memory in the mean time it will drop uncritical data, and store only processed data. In addition to the internal memory about 900kB of additional memory is available through the cellular modem for caching data, and other memory devices may be added at small extra cost and power usage if necessary.

2.5. Battery autonomy and power-saving strategies

Reliable power supply is a critical concern for waves in ice instruments due to the cold temperature conditions and the inability for solar power usage during polar night. In the present design, we decided to avoid the use of solar panels altogether (though solar panels could be easily added for building an open ocean buoy with unlimited autonomy outside of the polar regions). This allows to simplify the design, makes it less expensive and less labor intensive to produce, as well as avoids compromising the water-tightening of the instrument. In our case, the combination of a power-efficient microcontroller and 9dof, as well as power-optimized firmware, led to a drastic reduction of the power consumption compared with previous generations of the instrument. Table 1 summarizes the power consumption of the instrument in different modes. We use as power sources Lithium D-cells (SAFT LSH20), that have a nominal capacity of around 13Ah at 3.6V. In addition, a 3.3V step up-step down buck converter [60] is used to provide a stable 3.3V power source, even when the iridium modem briefly needs to draw much more power (bursts up to 250mA). This results in a typical operational time, with 2 Lithium D-cells mounted in parallel and no solar power, of around 4.5 months with our standard GNSS and waves measurement rates (GNSS position every 30 minutes and wave spectrum every 2 hours). If needed, increasing the number of batteries used in parallel will allow to increase the autonomy, proportionally to the number of cells used. In order to simplify the handling and deployment of the instrument, a magnetic switch was added to the design [61]. When the magnet is attached, the instrument is powered down. The instrument starts operating as soon as the magnet is removed, and can be switched on and off as many often as necessary. Some LEDs are used to provide visual indication of instrument activity, while their blinking frequency is taken low enough to preserve battery.

activity mode	activation frequency	current (mA)	mWh use per hour	time to empty 2 Li D-cells
sleep	when not active	0.3	1.0	7.3 years
gnss measurement	2 minutes twice per hour	30	3.3	2.2 years
wave measurement	20 minutes every 2 hours	8	4.4	1.6 years
iridium transmission	1 message per hour	burst 250mA	10	0.7 years
typical use			18.7	0.39 years \approx 4.6 months

Table 1. Overview of the power consumption and resulting battery autonomy resulting from the different components of the instrument when using two non rechargeable D-sized Lithium batteries. ‘Activity mode’ indicates which part of the workflow of the instrument that is being considered. ‘Activation frequency’ indicates a typical activation duration and frequency (though this is only indicative, and may be modified either by using different parameters when programming the instrument, and / or by sending updated activity rates through the bi-directional iridium communication). ‘Current (mA)’ is the average current used in the corresponding mode when it is active. ‘mWh use per hour’ indicates the corresponding energy use for the given activation mode, following the formula: $\text{mWhUsePerHour} = \text{Current(mA)} \times 3.3(\text{V}) \times \text{ActivationFraction}$.

The high level logics of the instrument (wake-up patterns, measurements patterns, iridium burst mode transmission) is implemented as a few lines of code that leverage an object oriented implementation of the processing components described above. To increase robustness of the firmware after instrument malfunction, we resort to a double strategy. First, the firmware is coded following defensive programming patterns, with extensive quality checks of all inputs and outputs for the different modules used. Second, the integrated hardware watchdog present on the microcontroller is enabled at all time, in order to force a complete reboot of the microcontroller in case of a hardware or a software malfunction. The test deployments have all been highly successful as described in Section 3, which is a testimony to the robustness of the design.

2.6. Total cost and assembly process

The total bill of materials is presented in Table 2. We consider that around 0.5 hours of work is needed to assemble one single instrument. No advanced electronics or hardware experience is required to build the instrument. At present, the assembly relies on a couple of soldering steps to connect the power supply module to the main board, and simply plugging the 9dof sensor and the electronic switch controlling it into the main board using a couple of Qwiic I2C cables.

component	function	price (USD)	assembly steps
Artemis Global Tracker	main board, MCU, GNSS, Iridium	375	ready to use
GNSS + Iridium antenna	passive antenna	65	screw on SMA cable
SMA extension cable 25 cm	extension cable for antenna	5	screw on tracker
Qwiic power switch	power on and off 9dof	7	disable LED, connect 9dof and tracker
ISM330DHCX + LIS3MDL	9dof sensor	18	connect to power switch
Qwiic cables (x2)	connect tracker, 9dof, switch	3	connect power switch and 9dof
3.3V Regulator S7V8F3	3.3V buck converter	10	solder to battery and tracker
2 x D cell holders	house and connect batteries	15	solder to 3.3V regulator
2 x SAFT LSH20	power supply	35	put in cell holders
reed MDRR-DT-20-35-F	magnetic switch	3	solder between battery and regulator
magnet	turn magnetic switch on / off	1	mount outside housing
housing box	housing, IP68	20	mount the electronics inside
misc: glue, wire	small extras	5	get the design assembled
total	fully functional instrument	562	0.5 hours / instruments, producing 10

Table 2. A representative list of components needed to build an instrument monitoring drift (GNSS) and wave activity (9dof sensor). The assembly time for a single instrument, when assembling a series of 10 instruments in bulk, is about 0.5 hours once the user is familiar with the design.

functionality	credits / message	messages / hour (default)	price / month (USD)
iridium subscription fee 1 month	N/A	N/A	16
GNSS position data	2	0.3	26
wave spectrum data	3	0.5	66
total	N/A	N/A	108

Table 3. Overview of the Iridium communication costs. The prices correspond to the ones offered by RockSeven at the time of writing this manuscript. It is possible to modify the rate of data collection by either using different setups when one programs the instrument firmware, or by sending a measurement rate update command to the instrument through iridium.

The total weight of the instrument depends on the size of the enclosure and the number of battery cells used. Typical weight is between 0.3 and 0.5kg with 2 lithium D-cell batteries. The dimensions of the instrument are typically around 12cm x 12cm x 9cm. Many of these parameters are part of design tradeoffs. For example, more batteries increase weight and would make deployment from drones more challenging, but also allow to extend battery time. The use of a small enclosure allows for easier shipping and logistics and make the instrument less likely to be discovered (and destroyed) by polar bears (in the Arctic), but means that even a relatively minor snowfall is enough to bury the instrument and interrupt GNSS and Iridium signals.

In addition to the construction costs, Iridium costs can be significant over time. Table 3 indicates the typical Iridium communication costs per month, for an instrument running with the default parameters (reducing the frequency of measurements will reduce costs accordingly).

Combining the building and the Iridium costs, the total cost per instrument, for an activity period of 6 months (which is reasonable for instruments deployed in the MIZ at the start of the ice formation season), is around 1200 USD, all included (noting that 4, rather than 2, LSH20 batteries will be necessary for such a deployment). Around 53% of this cost comes from the Iridium transmission fees, so reducing the frequency of measurements when little activity of interest is happening will allow to significantly reduce the total cost of ownership.

3. Validation of the instrument v2021

3.1. *Autonomy and satellite communication test in the Arctic: February 2021 deployment*

An early version of the instrument v2021, which performed drift tracking but no waves measurements, was deployed East of Svalbard in February 2021 [62]. This early version of the instrument was used to i) validate the general design of the electronics, ii) validate the low power modes and the ability of the instrument to live for several months in the Arctic cold using just a couple of D-sized battery cells, in particular, regarding the power consumption of the iridium modem which is hard to estimate, iii) validate the satellite communication binary protocols.

This initial test was a success. The instrument was deployed on February 24th, 2021, East of Svalbard. It then transmitted data for a few weeks, before a snow storm covered it with snow and blocked the iridium transmission. This loss of communications is a consequence of the design of the housing of the instrument, that relied on a box that was too low to survive even just 10 cm of snow, and will be addressed in future designs, by using a box that is higher and does not get buried as easily. However, communication with the Iridium satellites was restored when the ice melted in late May 2021. Following this, the instrument freely floated in the open sea, traveling all the way from Svalbard to the close vicinity of Murmansk, and back and forth between Novaya Zemlya and Murmansk.

An overview of the trajectory is presented in Fig. 3. At the time of writing this manuscript, i.e. January 2022, this instrument is still transmitting. The instrument was powered by 3 LSH20 batteries, which, since it measures only GNSS position each 30 minutes, should result in a battery life of around a bit over 1 year following the data from Table 1. Therefore, this deployment validates the low power design, and confirms that both the low power sleep mode, the GNSS data acquisition, and the Iridium transmission (which have the power consumption that is most difficult to estimate), are correctly implemented and allow extended operation on battery alone.

3.2. *Experiments for validation of small-amplitude, low frequency harmonic vertical displacement measurements at the University of Tokyo's wave-ice tank facility: July 2021 laboratory test*

We conducted an experiment in July 2021 to evaluate the instrument v2021 accuracy for measuring small amplitude harmonic vertical displacement representative of waves in ice. This experiment allows to fully validate i) the acquisition of raw data from the 9dof sensor, ii) the processing algorithm and scaling factors used in the low-level signal processing code, iii) the data encoding used for transmitting the information back to the user. The experiment took place in the wave-ice tank facility at the Kashiwa campus of the University of Tokyo, Japan. In our case, however, the instrument v2021 prototype was placed on the top of the wavemaker paddle (which is a plunging wedge, vertical displacement paddle), rather than in the water (see Fig. 4), so it measured the vertical motion of the wavemaker itself, which is known to a very high accuracy thanks to the use of a close-loop control. Indeed, even given the presence of a damping beach, it would be challenging in such a small wavetank to avoid reflections and standing waves, and to provide waves of sufficient amplitude at such low frequencies. By contrast, by placing the instrument on top of the paddle which is controlled by a close loop control system, we ensure that the vertical displacement of the instrument is exactly equal to the one specified in the control system. In this section, we will refer to the displacement prescribed as "wavemaker signal", to emphasize that the signal we measure is the vertical displacement

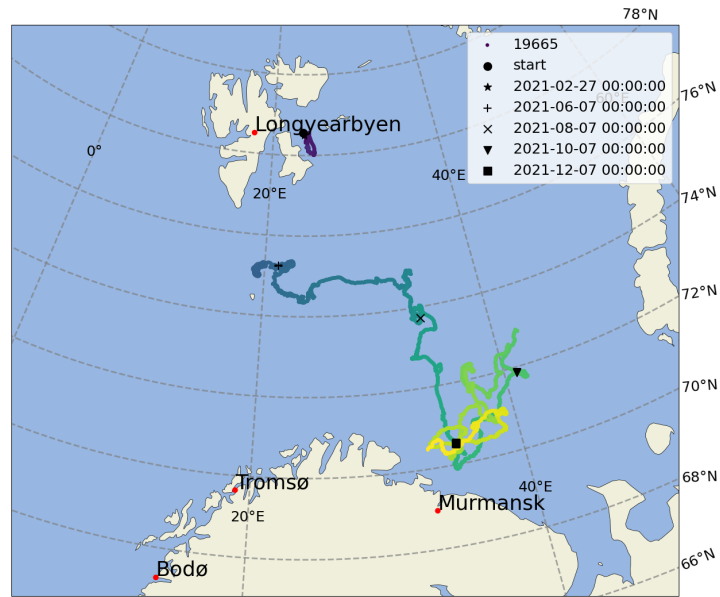


Figure 3. Overview of the drift trajectory of the early prototype instrument deployed in February 2021 East of Svalbard. The interruption in the trajectory is due to a layer of snow covering the instrument that prevented satellite communications during a couple of months. The instrument then drifted freely in the open ocean. This validates the low power design, and confirms that the design chosen is able to sustain operations for several months using just a few D-cell Lithium batteries.

of the plunger-type wavemaker on which the instrument sits, rather than the motion of the water.

The experiment aims to evaluate the instrument accuracy when it measures centimeter scale vertical displacement signals at low frequencies corresponding to typical waves in ice conditions. As such, we tested the instrument using both 1 and 2 cm amplitude sinusoidal wavemaker vertical motion for 12, 14, and 16 s periods (as well as an "extreme" test with 0.5 cm amplitude and 16 s period, to test the very limit of the instrument sensitivity), which resemble lower-bound wave signals that were previously measured in ice-covered ocean [15]. The spectral analysis setup was slightly modified from that of Section 2.2, so that each test case duration was reduced from 20 mins to roughly 7 mins, in order to be able to perform the tests within the allocated time slot. For this, the segment overlap was reduced to 50% (corresponding to 1024 points), which yields a total number of 3 segments per 7 mins measurement interval. All other processing configuration remained the same as Section 2.2.

The wavemaker signal experiment results are summarized in Table 4. The frequency of the wavemaker motion was not tuned to the exact frequencies bins, nevertheless, the peak of the energy was consistently captured in the nearest frequency bin, as reflected in Table 4. Then, the best estimate of the wave amplitude can be derived as:

$$A_{mono} = \sqrt{df \times (S(i-1)^2 + S(i)^2 + S(i+1)^2)}, \quad (6)$$

where $S(i)$ is the spectrum produced by the instrument, taken at the peak of the amplitude spectrum, and we consider also the neighboring frequency bins. The results of the wavemaker signal experiments summarized in Table 4 indicate that the instrument is capable of measuring monochromatic wave amplitudes of 1 to 2 cm with roughly 0.1 cm accuracy.

The spectra obtained during these tests, as well as two spectra corresponding to the instruments being at rest, are presented in Fig. 5. As visible there, the noise background of the instrument increases slightly when motion is present, which may be explained by a variety of reasons (small vibrations in the wave paddle mechanism when it is undergoing displacement, diffusion of numerical noise in the processing algorithms, inherent properties

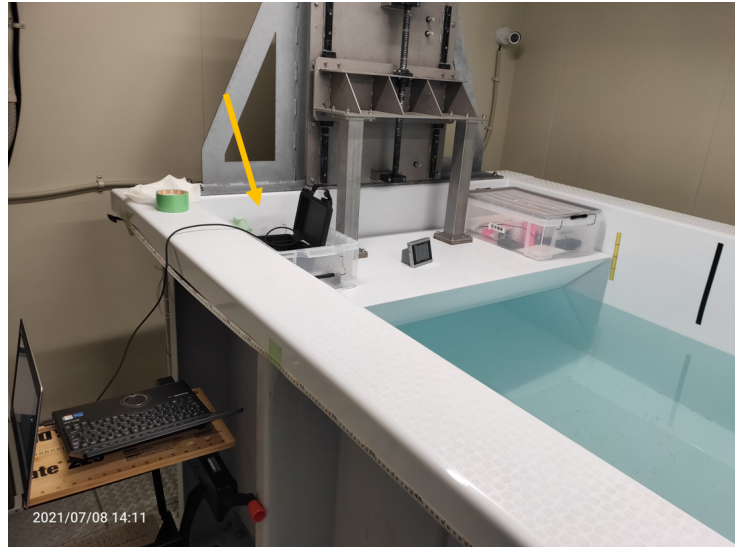


Figure 4. Illustration of the wavemaker experiment setup in the wave-ice tank of JAMSTEC (Japan Agency for Marine-Earth Science and Technology) at the Kashiwa campus of the University of Tokyo, Japan. The instrument v201 location is shown by the yellow arrow.

Wavemaker signal: amplitude, period	Reported buoy amplitude (cm, averaged)	Reported buoy peak period (s)	frequency bin closest to peak frequency	Number of cases
2 cm, 16 s	2.104	15.75	yes	4
1 cm, 16 s	1.040	15.75	yes	2
0.5 cm, 16 s	0.62	15.75	yes	2
2 cm, 14 s	1.954	13.65	yes	1
1 cm, 14 s	1.053	13.65	yes	1
2 cm, 12 s	2.022	12.05	yes	2
1 cm, 12 s	0.970	12.05	yes	2

Table 4. A summary of wavemaker signal experiment results. The experiments included 1 and 2 cm amplitude waves for 12, 14, and 16 s periods, as well as an "extreme" test corresponding to an amplitude of 0.5cm at 16s period. When multiple tests were conducted, the mean values are reported. The instrument is able to accurately measure waves down to 1 cm at a 16s period, with a corresponding typically accuracy of under 0.1 cm. For waves of amplitude 0.5cm and period 16s, the peak period is still correctly identified, but noise in the measurements starts to play a role, with a deviation of around 25% in the wave amplitude value reported.

of the 9dof sensor). However, the increase in the noise threshold when there is motion remains very limited. The signal-to-noise ratio is superior or equal to a factor of 50 in all experiments, except for the smaller waves (0.5cm at 16s period), in which case the signal-to-noise ratio is around 10. This means that, even in this last "extreme" test case, the signal-to-noise ratio is still good enough to clearly distinguish the motion from the noise floor, though noise starts to be visible in the integrated statistics, as discussed in the previous paragraph and in Table 4.

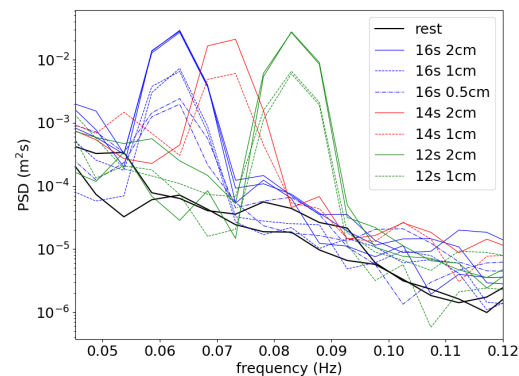


Figure 5. Spectra obtained when testing the instrument v2021 in the laboratory. The signal-to-noise ratio is superior or equal to 50 for all test cases, except the "extreme" test case with displacements of amplitude 0.5cm and period 16s, where the signal-to-noise ratio is around 10. The noise threshold obtained when the instrument is at rest is actually slightly better than what was obtained with the instrument v2018 (compare the present plot to Fig. 3 of Rabault *et al.* [52], curves "IMU 1" and "IMU 2").

Following this successful laboratory validation, we decided to perform a field validation experiment in the Arctic, which is described in the next subsection.

3.3. 2021 NABOS expedition and comparison with SOFAR Spotter buoy in the Marginal Ice Zone in the Arctic: September 2021 deployment

The first field deployment opportunity for the full-feature instrument v2021 was the 2021 Nansen and Amundsen Basins Observational System (NABOS) expedition (<https://uaf-iarc.org/nabos-cruises/>), which was performed in the context of the ArcS II Japan-Russia-Canada International Exchange Program (<https://sites.google.com/edu.k.u-tokyo.ac.jp/arcsii-iep-jrc/main>).

A v2021 instrument was assembled in a Zeni Lite drifting buoy, referred to Zeni-v2021 in the following (see Appendix B for the buoy assembly details). This instrument was deployed alongside a commercial wave measuring device (SOFAR Spotter drifting buoy, <https://www.sofarocan.com/products/spotter>), referred to as SPOT-1386 in the following. The key difference of the Spotter compared with the instrument v2021 is that the Spotter measures surface displacements based on GNSS signal.

On 15 September 2021, both buoys were deployed adjacent to the ice edge in the central Arctic Ocean (North of the Laptev Sea), at a location of 81.915° N 118.763° E, at around 05:05 UTC. The buoy deployment location is shown in Fig 6, which is overlaid with 0.15 and 0.80 sea ice concentration (SIC) contours based on the Advanced Microwave Scanning Radiometer 2 (AMSR2) data [63].

The two buoys drifted apart roughly 1 km within 8 hours of being deployed and farther than 3 km after 24 hours. Given that the ocean surface conditions near the ice edge are heterogeneous, we compared the instrument v2021 spectrum with that of SPOT-1386 following the deployment, when they are still close to each other. The buoy spectra compared generally well when the buoy distance was less than 1 km. The left panel in Fig 7 shows the buoy spectra at 08:21 15 Sep 2021, which corresponds to a moderate significant wave height $H_{m0}=1.62$ m. The comparison indicates that the spectra compare well when the low frequency noise, which is typical in accelerometers, is discarded. Here, the method that

was previously used in [64,65] was applied to remove this noise from Zeni-v2021 spectra. An ideal filter was applied to remove data below the lowest frequency local minimum of the smoothed spectrum. In addition, another spectra comparison on 08:21 16 Sep is also provided on the right panel in Fig 7. By this time, the buoys were farther than 3 km apart, so there is a slight difference in energy levels between the two buoys, but the general shape of the spectra is comparable. These buoy spectra comparisons indicate that the instrument v2021 can be used to measure ocean waves, and is a full scale, fieldwork validation of the algorithms and processing used.

Spotter buoys, without solar radiation recharge at high latitude during the winter season, have only a short life span and SPOT-1386 ran out of battery on September 30th 2021. Analysis into the co-located deployment in and near the Marginal Ice Zone MIZ is ongoing and will be presented in a separate, science-focused paper. Notwithstanding, a comparison of wave periods (peak period T_p and spectral period T_{0m1}) when the buoys were less than 5 km apart agree within better than 5%. In addition, the part of the dataset for which both instruments are frozen in the ice indicates that the noise level of the IMU-based instrument (typically around 0.1cm) is much better than that of the Spotter buoy (typically around 1cm), which is expected due to the different data sampling techniques used (GPS vs. 9dof). This provides further support that Zeni-v2021 captured consistent energy densities to those of SPOT-1386, and that the Zeni-v2021 is actually better suited for measuring small waves in ice. Note that the analysis into the Zeni-v2021 and SPOT-1386 co-located deployment is described in a manuscript under preparation, which is why these data are not released in the present technical paper.

Regarding battery life, the estimated deployment duration for hourly wave sampling rate was 3.3 months based on Table 1 where 2 Tadiran D-cell batteries (TL-5930/S), each with 19 Ah capacity, were used as the power source. The instrument stopped transmitting, likely due to empty battery, in mid-December, i.e. 3 months after the start of the deployment. This corresponds well to the estimated battery life, and this is an additional confirmation of the quality of the design of the hardware and software, in particular regarding the implementation of low power modes.

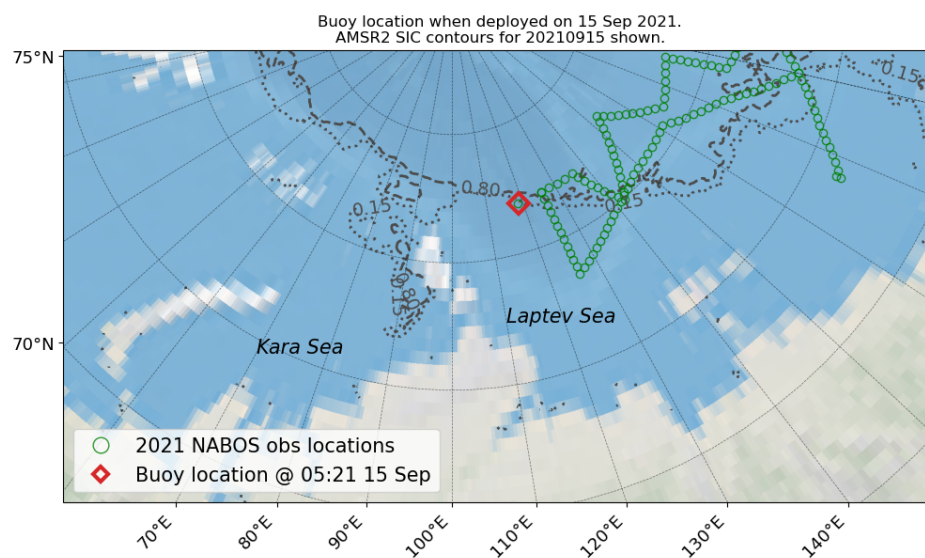


Figure 6. Zeni-v2021 and SPOT-1386 deployment location on 15 Sep 2021. The 2021 NABOS observation locations also shown in green markers, and AMSR2-derived 0.15 and 0.80 Sea Ice Concentration (SIC) contours for the same day are overlaid.

3.4. The "Floatenstein" drifting buoy in the Caribbeans: November 2021

A small-scale, home-produced drifting buoy was built following the v2021 design in the context of the OneOcean expedition, a circumnavigation by the Norwegian tall ship

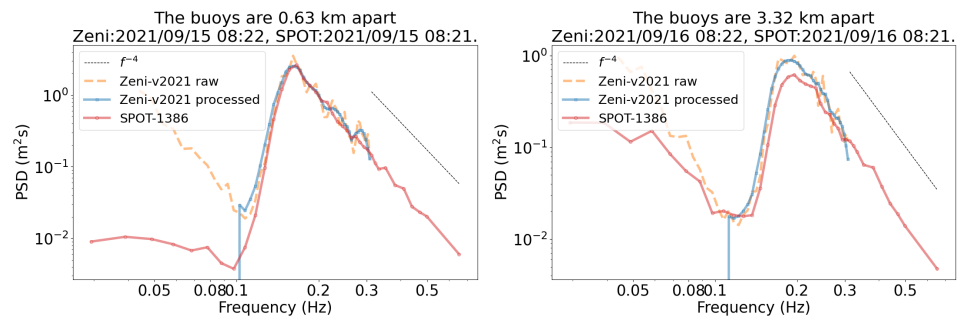


Figure 7. Comparison of Zeni-v2021 and SPOT-1386 spectra at 08:21 15 Sep 2021 (left) and 08:21 16 Sep 2021 (right): the distance between the buoys were roughly 1 km and 3 km, respectively. Low frequency noise of Zeni-v2021 spectra was removed by applying an ideal filter at the lowest frequency local minimum of the smoothed spectrum, similarly to what was reported in Waseda *et al.* [64], Nose *et al.* [65], which is a common post processing step when working when accelerometer-derived wave data. This indicates agreement within typically 5% between the Sofar Spotter and the instrument v2021 measurements in the open water, before the instruments drift away from each other.

Statsraad Lehmkuh. The electronics are identical to everything that was discussed above, except for the battery solution used. Indeed, since the instrument had to be transported by plane in the checked luggage of an expedition member, 2 traditional alkaline D-cell batteries were used instead of the Li-batteries of the other instruments. A standard rectangular rigid plastic enclosure of size 12cm x 12cm x 9cm was used as the main body of the buoy. Since alkaline batteries are heavier than lithium batteries, this resulted in too low floatability. As a consequence, we added small chunks of styrofoam (wrapped in duct tape to protect from abrasion from the elements), on the sides of the rigid plastic enclosure, in order to increase buoyancy. These additional buoyancy elements were set in place using cable ties and glued in position using bathroom silicone sealant. In addition, the rigid plastic box was fully sealed with epoxy glue, to make it fully watertight. Finally, a keel composed of diverse scrap metallic parts (spare screws and bolts), packed in duct tape, was fixed at the bottom of the plastic enclosure to improve its static stability. This combination of disparate building materials gave a very peculiar appearance to the drifting buoy, which got nicknamed humouristically as "Floatenstein" (Floating-Frankenstein). A picture of the instrument, fully built before being shipped, is presented in Fig. 8.

Despite its unusual appearance, Floatenstein proved to be a robust and well-functioning drifter. Floatenstein was released in the open ocean on November 11th 2021, and, at the time of writing this manuscript (January 2022), Floatenstein has been active for a bit over two months and is working nominally. Data transmitted include i) GPS position sampled each 30 minutes, and ii) vertical wave spectrum sampled each 2 hours, similar to what is described in Section 2.

In order to validate the good functioning of Floatenstein and the electronics inside, we compare the wave data transmitted by Floatenstein with direct satellite altimeter measurements. The Wavy software package (see <https://github.com/bohlinger/wavy>, Bohlinger *et al.* [66]) is used to perform i) gathering of relevant satellite data, ii) collocation with the position of Floatenstein, iii) extraction of relevant fields. For the collocation process, the temporal and spatial constraints are 30 minutes and 50 km, respectively. Four satellite missions had coinciding satellite measurements Sentinel-3A/B, Jason-3, and the altimeter mounted on CFOSAT. The satellite observations are of level 3 processing quality and publicly available in the Copernicus CMEMS archive¹.

Results of the comparison between Floatenstein and satellite data are presented in Fig. 9. The satellite data are naturally spaced in time, due to the orbital pattern of the

¹ https://resources.marine.copernicus.eu/product-detail/WAVE_GLO_WAV_L3_SWH_NRT_OBSERVATIONS_014_001/INFORMATION

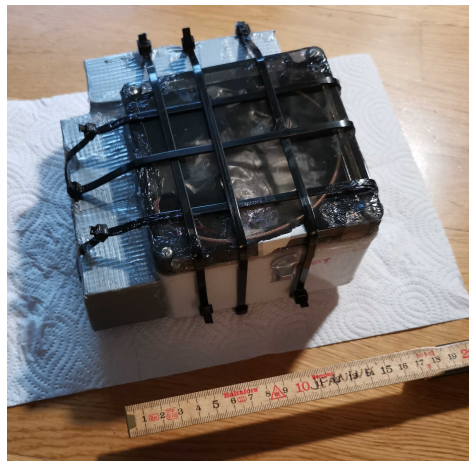


Figure 8. The "Floatenstein" drifting buoy. The main body of the buoy is a rectangular plastic box of dimensions 12cm x 12cm x 9cm. Due to the use of alkaline batteries instead of lithium batteries, additional floatation elements were needed. These are the two rectangular elements on the side of the main body of the buoy, which are built from styrofoam chunks, wrapped in duct tape, and held in place by cable ties. The full box is sealed with epoxy to be perfectly water tight. To avoid the need for making holes in the box, the Iridium antenna is also taped in place inside the box, immediately under the lid, pointing 45 degrees upwards. The small cubic magnet (wrapped in duct tape to hold it in position under transport), which controls the magnetic switch glued on the inside of the box, is visible on the front panel of the main body of the drifter.

satellites collecting the data. The wave statistics reported by Floatenstein lies within the typical error margin of the direct observation by the satellites, which is visible as the spread in significant wave height over each small transect cutting the trajectory of Floatenstein. This is an additional confirmation of the good functioning of the electronics and firmware of the instrument v2021, as well as an illustration that even coarse drifters are able to accurately measure wave properties, as long as these are small enough to follow the motion of the waves. Indeed, the amplitude of the hydrodynamic response of the buoy itself is typically a fraction of the size of the buoy. This could, in theory, reach up to several tens of cm for large floating buoys with a typical size of several meters as is commonly found for commercially available moored system, which has, historically, created a focus on developing wave buoys with a small hydrodynamic response. Our small buoy, by contrast, has a typical size of 10 cm, which implies that even a "bad" hydrodynamic response, leading to a displacement response in waves of several tens of percents of the buoys size, is still only a few centimeters in amplitude, well below typical stochastic noise and uncertainty of stochastic wave measurements in the open ocean. This is similar to other small buoys developed to measure specifically the high frequency part of the wave spectrum [67].

4. Conclusion and future work

This work introduces a new design of waves in ice instrument. This instrument is able to perform measurements of sea ice drift using a GNSS module, and to measure waves in ice up to a high degree of accuracy (the noise threshold corresponds to a significant wave height of around 0.1 cm for waves with period 16 s, as demonstrated in the laboratory experiments), using an inexpensive 9-degree-of-freedom sensor and extensive signal processing and noise reduction techniques. In addition, this instrument costs only around slightly under 650 USD in electronics components (including taxes), raw materials, and assembly time, which is around 10 times less expensive to build than the closest, least expensive commercial alternative we know about (the Sofar Spotter buoy), while being both more sensitive and much better adapted to working in the polar night, when solar input is not available. Battery autonomy using a couple of Lithium D-cells is around 4.5 months, and this can be easily increased by adding more cells if needed. Assembly time has been cut down to

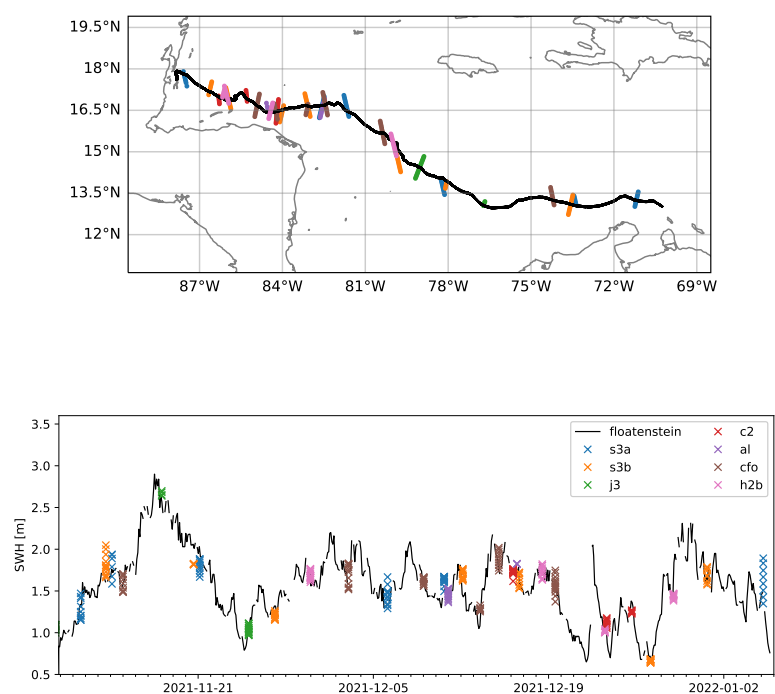


Figure 9. Top: drift pattern of Floatenstein (colored continuous black line), and satellite measurement transects (colored points) over which wave characteristics are measured by the satellites. Bottom: comparison of the Significant Wave Height (SWH), i) reported by Floatenstein (continuous black line), and ii) from satellite measurements on the corresponding measurement transects.

no more than typically half an hour per instrument, when building a small series of these, thanks to the use of highly integrated open source development boards, and requires no advanced electronics skills. The code is provided as open source together with the full design blueprints and assembly instructions. Binary pre-compiled firmware versions are provided for rapid upload to the electronics boards. Additional sensors can easily be added to the instruments. Two-way communications are implemented through highly optimized binary protocols and take place over Iridium, which enables global coverage. All data measurements take place at fixed time thanks to the real time clock available on board, to allow easy time synchronization, and data are buffered on board the instrument to avoid data loss in case of temporary failure in establishing an iridium communication.

This set of characteristics is, as a whole, and to the best of our knowledge, unrivaled by any commercial or open source instrument. These drastic improvements in cost and power efficiency, compared even to previous open source instruments such as the one described in Rabault *et al.* [52], are made possible by the use of a cutting edge smartwatch processor, which can handle all of the logging and data processing workloads. This is, to the best of our knowledge, the first time that such an instrument is built without the need for an expensive, data-hungry embedded microcomputer, and the first time the full logging, data processing, and communication workload is controlled from a single, power-efficient, low cost microcontroller. In addition, this design is both modular and adapted to not only measurements of sea ice drift and waves in ice, but also for open ocean measurements, as we have shown in the validation section.

We validated the design and signal processing code through a number of test campaigns both in the laboratory, in the open water, and on sea ice. Results indicate agreement within typically 5% with state of the art, reference instruments and measurement methods, and confirm the robustness of the design as well as its power efficiency. In particular, the low power consumption of the instrument (which is always challenging to get right, as any error in the code or hardware design can easily increase the sleep power by a few milliamps, which, while it sounds little, drains battery over time), is confirmed during a deployment in the Arctic, and the accuracy of the wave measurements and wave data processing code is confirmed in several test cases.

This work, while purely technical, is in our opinion a groundbreaking step towards bringing open source, inexpensive, easy to build, modular, power efficient instrumentation to the field of in-situ study of waves in ice in particular, and oceanographic measurements in general. Indeed, our design has also proven to be highly efficient both as a traditional open ocean drifter and wave buoy. We expect that this design will significantly reduce the barrier to entry for new research groups to monitor waves in ice, sea ice drift patterns, and similarly wave activity and ocean upper layer drift and currents in the open ocean, and be a possible platform for public outreach since the design is simple enough to be assembled for example by high school students. This design can also be used as a basic building block for setting up a variety of instruments, such as weather stations, environmental loggers, and wildlife monitoring instruments, by connecting specific sensors and adding extra modules to the firmware, while taking advantage of the power efficiency and iridium connectivity of the main electronics board.

Another key improvement brought by the present design is its very limited weight and size. Typically, our implementations of the design occupy a space of 12cm x 12cm x 9cm and weight between 0.3 and 0.5kg. This is small and light enough to be deployed for example from a medium size quadcopter drone, or from a fishing rod several meters long. This could prove very important in practice when releasing large numbers of drifters in the Arctic and Antarctic, since i) changing the course of an icebreaker, and forcing it to make many (even short) stops to put instrumentation on the ice, has a large footprint on the operation of the ship and a high cost, while ii) flying a quadcopter drone to deploy instrumentation on the ice, or dropping instrumentation on-the-go while the boat is steaming using a long rod, would have a zero marginal cost for an expedition.

As a final technical word, we want to say that, while we release the instrument design, code, and validation data already now since we believe that it may already be of value for other groups, there are a number of ongoing improvements that are taking place and will continue also after this paper is published. First, we are working on adding the possibility to store large amounts of data using a 2-step memory setup (i.e., a Ferromagnetic RAM for fast, low power storing of data on the fly, and a SD card for slow long-term archiving). This will make the instrument even more valuable in contexts where it can be recovered after deployment, since this will allow to store and recover time series that are, at present, only stored temporarily in RAM for performing wave spectrum computation. Second, we will work in the future on adding full directional wave spectrum information. Finally, when these additions are done, a major code rewrite, taking into account all the lessons learnt while developing this instrument, will be performed to make the code even easier to extend, test, and reuse. When these further refinements are implemented, we will also consider designing an open source printed circuit board, to lower even further the amount of work needed to assemble the electronics of the instrument. Ultimately, if there is interest from the community, we hope to provide an all-in-one, fully integrated, low cost electronics boards that will be a true turn-key solution.

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Conflicts of Interest: No conflict of interest.

Appendix A: open source release and founding of an open source community

All the code, instructions, and post processing scripts necessary to build, program, and use the instrument v2021, are available as open source hardware and software on Github under a MIT license at the following URL: <https://github.com/jerabaul29/OpenMetBuoy-v2021a>.

Appendix B: Zeni-v2021 assembly for the 2021 NABOS cruise

The Artemis Global Tracker and other electronics components listed in Section 2.1 were enclosed in a water resistant Takachi box as shown in Fig 1. Deployment in the open water requires a floating enclosure that is durable and watertight. For our purpose, we used a commercial Zeni Lite drifting buoy (<https://www.zenilite.co.jp/prod/new-chikuden.html>), that was purchased by the University of Tokyo several years ago, as the floating enclosure. Images of the floating enclosure setup are shown in Fig 2. The Zeni Lite drifting buoy has open space inside where the Takachi box was fixed. This space was easily accessible, and it is also high enough for the antenna to be fixed on the upper half of the buoy. The water tightness is achieved via an O-ring placed underneath the blue plate. When the steel brace is fastened, pressure is evenly applied around the buoy housing where the top and bottom parts of the buoy are joined. The instrument v2021 that was deployed in the Arctic Ocean as described in Section 3.3, is referred to as Zeni-v2021 in the main text, because the instrument v2021 was assembled in the Zeni Lite drifting buoy enclosure.

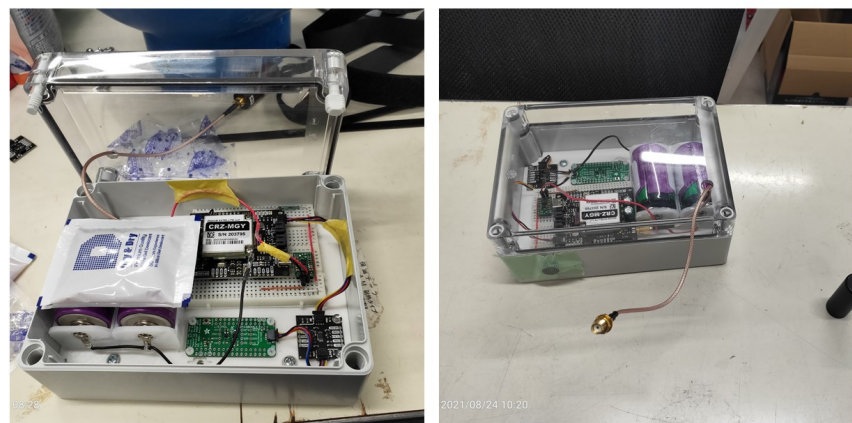


Figure 1. Illustrations of the instrument v2021 electronics components enclosed in a water resistant Takachi box.

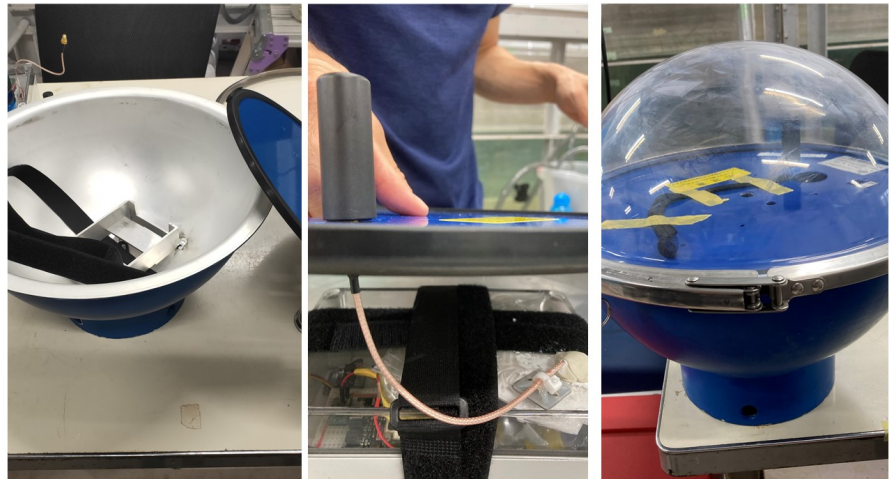


Figure 2. Illustrations of Zeni Lite drifting buoy being used as a floating housing. The left image shows the bottom part of the buoy where the Takachi box was fixed. The middle image shows how the antenna was attached through the blue plate. The right image shows the assembled Zeni-v2021.

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