

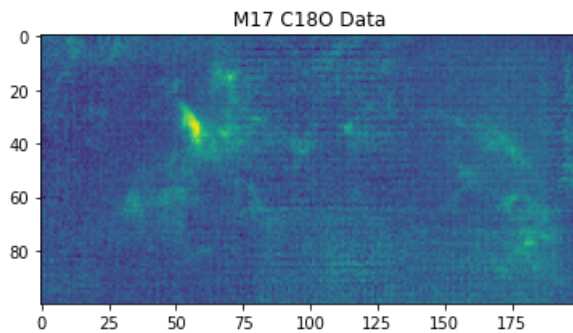
```
In [1]: import numpy as np
import matplotlib.pyplot as plt
import astropy.io.fits as fits
import astropy.wcs as WCS
from astropy.table import Table
from astropy import units as u
```

```
In [ ]:
```

## The reference of the Example data

'M17\_C18O\_Part.fits' is the  $C^{18}O$  ( $J = 1 - 0$ ) emission line of the Milky Way Imaging Scroll Painting (MWISP) within  $11.7^\circ \leq l \leq 13.4^\circ$ ,  $0.22^\circ \leq b \leq 1.05^\circ$  and  $5 \text{ km s}^{-1} \leq v \leq 35 \text{ km s}^{-1}$ . MWISP project is a multi-line survey in  $^{12}\text{CO}/^{13}\text{CO}/C^{18}\text{O}$  along the northern galactic plane with PMO-13.7m telescope.

```
In [2]: path_M17_18 = 'M17_C18O_Part.fits'
real_data_M17_18 = fits.getdata(path_M17_18)
plt.title('M17 C18O Data')
plt.imshow(real_data_M17_18.sum(0))
plt.show()
```



```
In [ ]:
```

## Installation and Usage of FacetClumps

```
In [3]: # pip install FacetClumps==0.0.4

import FacetClumps
from FacetClumps.Detect_Files import Detect as DF_FacetClumps
```

```
In [ ]:
```

## Detection Parameters

### Input of FacetClumps (2D and 3D)

**RMS:** The RMS of the data.

**Threshold:** The minimum intensity used to truncate the signals.

Default Value: 2\*RMS

Recommended value: ['mean','otsu',n\*RMS]

**SWindow:** The scale of window function.

Default Value: 3

Recommended value: [3,5,7]

**KBins:** The coefficient used to calculate the number of bins for the eigenvalue.

Default Value: 35

Recommended value: [10,...,60]

**FwhmBeam:** The FWHM of the instrument beam, in pixels.

Default Value: 2

**VeloRes:** The velocity resolution of the instrument, in channels.

Default Value: 2

**SRecursionLBV:** The minimum area of a region in the spatial direction and the minimum length of a region in the velocity channels when a recursion terminates. Clumps also need to satisfy the conditions. [SRecursionLB,SRecursionV] for 3D, [SRecursionLB] for 2D.

Default Value: [16,5]

**file\_name:** File name.

**mask\_name:** mask name, used to store region information. The index (starts with the number one) of each clump corresponds to the same number in the mask.

**outcat\_name:** Used to store clump table in pixel coordinate system.

**outcat\_wcs\_name:** Used to store clump table in WCS coordinate system.

In [ ]:

### Output information

**files:** Regional information and clump tables saved according to the customized file name.

**did\_tables:** Detected information dictionary, whose keys is [mask, outcat\_table, outcat\_wcs\_table]

**mask:** Regional information.

**outcat\_table:** Clump table of pixel coordinate system.

**outcat\_wcs\_table:** Clump table of WCS coordinate system.

In [ ]:

```
In [4]: #Get the RMS from the header
header = fits.getheader(path_M17_18)
RMS = header['RMS']
print('RMS:', RMS)
```

RMS: 0.196804

In [ ]:

In [ ]:

```
#2D, PP Space

RMS = 0.23
Threshold = 2 * RMS # ['mean', 'otsu', n*RMS]
SWindow = 3 # [3, 5, 7]
KBins = 35 # [10, ..., 60]
FwhmBeam = 2
VeloRes = 2
SRecursionLBV = [16, 5] # [(2+FwhmBeam)**2, 3+VeloRes]

SRecursionLB = SRecursionLBV[0]
did_FacetClumps = FacetClumps.FacetClumps_2D_Funs.Detect_FacetClumps(RMS, Threshold, SWindow, KBins, SRecursionLB, real_data_M16.sum(0))
```

In [ ]:

In [5]:

```
#3D, PPV Space

RMS = 0.23
Threshold = 2 * RMS # ['mean', 'otsu', n*RMS]
SWindow = 3 # [3, 5, 7]
KBins = 35 # [10, ..., 60]
FwhmBeam = 2
VeloRes = 2
SRecursionLBV = [16, 5] # [(2+FwhmBeam)**2, 3+VeloRes]

parameters = [RMS, Threshold, SWindow, KBins, FwhmBeam, VeloRes, SRecursionLBV]
file_name = path_M17_18
mask_name = 'mask.fits'
outcat_name = 'outcat.csv'
outcat_wcs_name = 'outcat_wcs.csv'
did_tables_FacetClumps = DF_FacetClumps(file_name, parameters, mask_name, outcat_name, outcat_wcs_name)
```

```
100% ██████████ 112/112 [00:09<00:00, 11.26it/s]
100% ██████████ 33/33 [00:08<00:00, 4.03it/s]
100% ██████████ 33/33 [00:01<00:00, 30.67it/s]
```

```
Number: 173
Time: 34.29
```

```
WARNING: FITSFixedWarning: VELREF = 0.0000000000000E+00 /
invalid keyvalue. [astropy.wcs.wcs]
```

In [ ]:

In [6]:

```
# Obtain the reginal information, clump table of pixel coordinate system and WCS coordinate system from files.
regions_data = fits.getdata('mask.fits')
outcat_table = Table.read('outcat.csv')
outcat_wcs_table = Table.read('outcat_wcs.csv')

# Obtain the reginal information, clump table of pixel coordinate system and WCS coordinate system from retrun.
regions_data = did_tables_FacetClumps['mask']
outcat_table = did_tables_FacetClumps['outcat_table']
outcat_wcs_table = did_tables_FacetClumps['outcat_wcs_table']
```

In [ ]:

```
In [7]: print('Outcat_Pix_Table:\n', outcat_table)
```

Outcat\_Pix\_Table:

ID	Peak1	Peak2	Peak3	Cen1	Cen2	...	Size3	Peak	Sum	Volume	Angle	Edge
	pix	pix	pix	pix	pix	...	pix			pix	deg	
1	197	17	13	196.746	16.921	...	1.289	1.453	47.274	79	27	1
2	63	38	75	62.502	38.000	...	4.384	2.923	1298.702	1375	86	0
3	72	38	76	69.770	38.082	...	3.147	2.506	779.211	836	50	0
4	67	47	71	66.451	46.229	...	2.759	2.214	535.663	634	-43	0
5	58	39	80	56.282	31.927	...	2.396	2.607	509.990	701	76	0
6	54	31	73	54.504	30.156	...	1.530	1.359	68.938	109	51	0
7	79	45	76	79.598	45.148	...	2.434	1.886	191.926	310	-59	0
8	73	42	78	72.025	42.473	...	2.673	1.528	215.345	332	39	0
9	71	16	76	70.998	16.500	...	3.925	3.731	1120.997	1123	-87	0
10	67	14	84	66.496	14.942	...	3.553	3.491	1674.092	1842	-81	0
...	...	...	...	...	...	...	...	...	...	...	...	...
164	129	4	161	127.789	5.428	...	2.788	1.667	328.107	466	-27	1
165	153	65	153	153.235	65.098	...	1.790	1.119	33.691	65	23	0
166	155	61	156	154.944	61.894	...	2.116	1.211	60.173	104	-84	0
167	157	52	162	157.721	52.327	...	2.702	2.236	202.361	266	-22	0
168	19	40	157	18.126	39.821	...	1.993	1.783	58.176	85	87	0
169	176	44	165	175.784	43.557	...	2.867	1.822	161.366	244	44	0
170	173	39	166	173.289	39.704	...	1.970	1.399	55.353	90	61	0
171	168	36	164	168.727	35.927	...	2.230	2.172	101.529	150	49	0
172	96	7	174	94.673	6.547	...	2.436	1.716	147.136	204	34	1
173	91	5	177	91.000	4.000	...	3.736	1.580	184.586	274	33	1

Length = 173 rows

```
In [ ]:
```

```
In [8]: print('Outcat_WCS_Table:\n', outcat_wcs_table)
```

Outcat\_WCS\_Table:

ID	Peak1	Peak2	Peak3	Cen1	Cen2	...	Size3	Peak	Sum	Volume	Angle	Edge
	deg	deg	km / s	deg	deg	...	km / s	K	K km / s	pix	deg	
1	11.733	0.350	7.359	11.735	0.349	...	0.215	1.453	7.879	79	27	1
2	12.850	0.525	17.693	12.854	0.525	...	0.731	2.923	216.460	1375	86	0
3	12.775	0.525	17.860	12.794	0.526	...	0.525	2.506	129.875	836	50	0
4	12.817	0.600	17.027	12.821	0.594	...	0.460	2.214	89.281	634	-43	0
5	12.892	0.533	18.527	12.906	0.474	...	0.399	2.607	85.002	701	76	0
6	12.925	0.467	17.360	12.921	0.460	...	0.255	1.359	11.490	109	51	0
7	12.717	0.583	17.860	12.712	0.585	...	0.406	1.886	31.989	310	-59	0
8	12.767	0.558	18.193	12.775	0.562	...	0.446	1.528	35.893	332	39	0
9	12.783	0.342	17.860	12.783	0.346	...	0.654	3.731	186.842	1123	-87	0
10	12.817	0.325	19.193	12.821	0.333	...	0.592	3.491	279.028	1842	-81	0
...	...	...	...	...	...	...	...	...	...	...	...	...
164	12.300	0.242	32.027	12.310	0.254	...	0.465	1.667	54.687	466	-27	1
165	12.100	0.750	30.694	12.098	0.751	...	0.298	1.119	5.615	65	23	0
166	12.083	0.717	31.194	12.084	0.724	...	0.353	1.211	10.029	104	-84	0
167	12.067	0.642	32.194	12.061	0.644	...	0.450	2.236	33.728	266	-22	0
168	13.217	0.542	31.361	13.224	0.540	...	0.332	1.783	9.696	85	87	0
169	11.908	0.575	32.694	11.910	0.571	...	0.478	1.822	26.896	244	44	0
170	11.933	0.533	32.861	11.931	0.539	...	0.328	1.399	9.226	90	61	0
171	11.975	0.508	32.527	11.969	0.508	...	0.372	2.172	16.922	150	49	0
172	12.575	0.267	34.194	12.586	0.263	...	0.406	1.716	24.524	204	34	1
173	12.617	0.250	34.694	12.617	0.242	...	0.623	1.580	30.766	274	33	1

Length = 173 rows

```
In [ ]:
```

## Detection Plots

```
In [9]: # Obtain the required information from output tables.
clump_centers = np.c_[outcat_table['Cen1'], outcat_table['Cen2'], outcat_table['Cen3']]
clump_angles = outcat_table['Angle']
clump_edges = outcat_table['Edge']
```

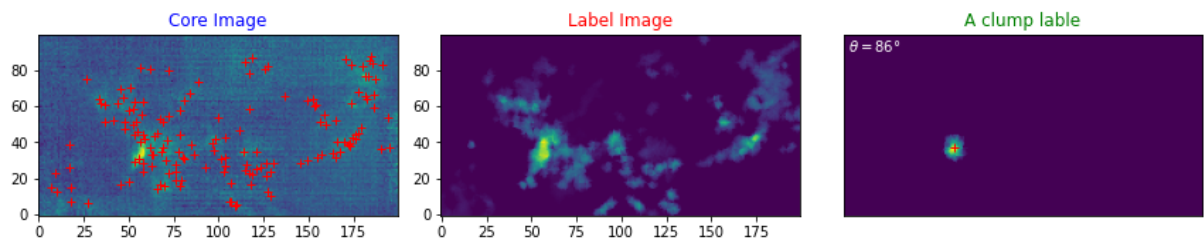
In [ ]:

```
In [10]: origin_data = real_data_M17_18

fig, (ax0, ax1, ax2) = plt.subplots(1, 3, figsize=(12, 8))
for i in range(np.int(regions_data.max())):
    # Show the clumps which do not touch the edges.
    if clump_edges[i] == 0:
        center_x = clump_centers[i][0]-1
        center_y = clump_centers[i][1]-1
        ax0.plot(center_x, center_y, 'r+')

#Obtain the region and angle of a clump.
index = 1
clump_region_i_coords = np.where(regions_data == index+1)
clump_region_label = np.zeros_like(origin_data)
clump_region_label[clump_region_i_coords] = 1
ax2.text(2, origin_data.shape[1]-10, r'$\theta={}\degree$'.format(np.around(clump_angles[index], 0)), color='white',
        fontsize=10)
ax2.plot(clump_centers[index][0]-1, clump_centers[index][1]-1, 'r+')

ax0.imshow(origin_data.sum(0))
ax1.imshow(regions_data.sum(0))
ax2.imshow(clump_region_label.sum(0))
ax0.set_title('Core Image', fontsize=12, color='b')
ax1.set_title('Label Image', fontsize=12, color='r')
ax2.set_title('A clump lable', fontsize=12, color='g')
for ax in [ax0, ax1, ax2]:
    ax.invert_yaxis()
fig.tight_layout()
plt.xticks([], plt.yticks([]))
plt.show()
```



In [ ]:

```
In [13]: data_cube = fits.getdata(path_M17_18)
data_header = fits.getheader(path_M17_18)
wcs = WCS.WCS(data_header)

fig = plt.figure(figsize=(18, 7))
ax = fig.add_subplot(111, projection=wcs.celestial)

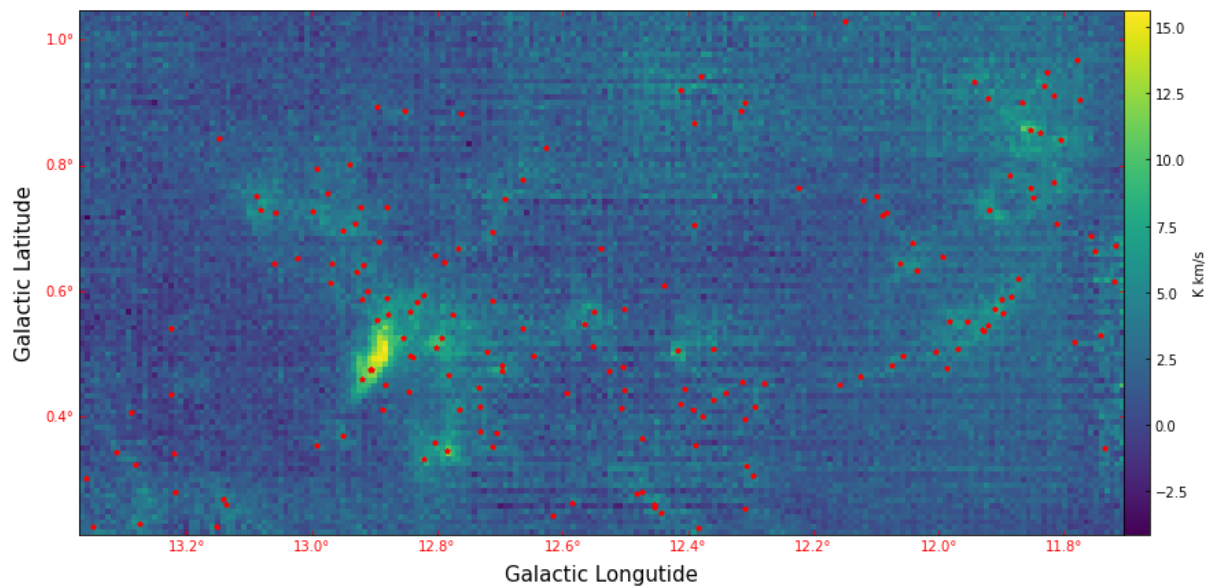
plt.rcParams['xtick.direction'] = 'in'
plt.rcParams['ytick.direction'] = 'in'
plt.rcParams['xtick.color'] = 'red'
font2 = {'family' : 'Times New Roman',
'weight' : 'normal',
'size' : 15,
}
plt.xlabel("Galactic Longitude", font2)
plt.ylabel("Galactic Latitude", font2)

for i in range(len(clump_centers)):
    center_x = clump_centers[i][0]-1
    center_y = clump_centers[i][1]-1
    ax.plot(center_x, center_y, 'r*', markersize = 4)

lon = ax.coords[0]
lat = ax.coords[1]
lon.set_major_formatter("d. d")
lat.set_major_formatter("d. d")
lon.set_ticks(spacing=12 * u.arcmin)

gci = plt.imshow(data_cube.sum(axis=0)*0.166)#, cmap='gray'
cbar = plt.colorbar(gci, pad=0)
cbar.set_label('K km/s')
# plt.xticks([], plt.yticks([]))
# plt.savefig('Example_0.pdf', format='pdf', dpi=1000)
plt.show()
```

WARNING: FITSFixedWarning: VELREF = 0.0000000000000E+00 /  
invalid keyvalue. [astropy.wcs.wcs]



In [ ]: