Teleparallel Equivalent of General Relativity (TEGR)

In the **Teleparallel Equivalent of General Relativity (TEGR)**, gravity is described using *torsion* instead of curvature. The formulation employs *tetrad* (*vierbein*) *fields* and the *Weitzenböck connection*, which has torsion but no curvature.

Key Mathematical Concepts

1. Tetrad (Vierbein) Fields:

$$g_{\mu\nu} = \eta_{ab} e^a_{\mu} e^b_{\nu}, \tag{1}$$

where:

- e^a_μ are the tetrad fields (basis of the tangent space),
- $\eta_{ab} = \text{diag}(-1, 1, 1, 1)$ is the Minkowski metric.
- 2. Weitzenböck Connection:

$$\Gamma^{\lambda}_{\mu\nu} = e^{\lambda}_{a} \,\partial_{\mu} e^{a}_{\nu}. \tag{2}$$

This connection has zero curvature but non-zero torsion. The torsion tensor is given by:

$$T^{\lambda}_{\mu\nu} = e^{\lambda}_{a}(\partial_{\mu}e^{a}_{\nu} - \partial_{\nu}e^{a}_{\mu}). \tag{3}$$

3. Torsion Tensor:

$$T^{\lambda}_{\mu\nu} = e^{\lambda}_{a}(\partial_{\mu}e^{a}_{\nu} - \partial_{\nu}e^{a}_{\mu}). \tag{4}$$

4. Gravitational Action in TEGR:

$$T = S_{\lambda}^{\ \mu\nu} T_{\mu\nu}^{\lambda},\tag{5}$$

where $S_{\lambda}^{\ \mu\nu}$ is the *superpotential*:

$$S_{\lambda}^{\ \mu\nu} = \frac{1}{2} \left(K^{\mu\nu}_{\ \lambda} + \delta^{\mu}_{\lambda} T^{\alpha\nu}_{\ \alpha} - \delta^{\nu}_{\lambda} T^{\alpha\mu}_{\ \alpha} \right), \tag{6}$$

and $K^{\mu\nu}_{\lambda}$ is the contorsion tensor:

$$K^{\mu\nu}_{\ \lambda} = -\frac{1}{2} \left(T^{\mu\nu}_{\ \lambda} - T^{\nu\mu}_{\ \lambda} - T^{\mu\nu}_{\lambda} \right). \tag{7}$$

5. TEGR Field Equations:

$$e G^{\lambda \nu} = 2e \left(\nabla_{\lambda} T^{\lambda}_{\ \nu} - \frac{1}{2} g_{\lambda \nu} T \right), \tag{8}$$

where $e = \det(e^a_\mu)$, and $G^{\lambda\nu}$ is the Einstein tensor.

Elimination of Curvature and the Role of Torsion

- In GR, gravity is described by curvature via the Riemann curvature tensor, using the Levi-Civita connection (torsion-free).
- In TEGR, curvature is eliminated in favor of a flat connection with torsion, encoded in the torsion scalar T.

General Relativity (GR)

In **General Relativity (GR)**, gravity is described by the curvature of spacetime, which is encoded in the *metric tensor* and the *Levi-Civita connection*.

Key Mathematical Concepts

1. Metric Tensor:

$$ds^2 = g_{\mu\nu} \, dx^\mu \, dx^\nu. \tag{9}$$

The metric tensor $g_{\mu\nu}$ describes the geometry of spacetime.

2. Levi-Civita Connection (Christoffel Symbols):

$$\Gamma^{\lambda}_{\mu\nu} = \frac{1}{2} g^{\lambda\sigma} \left(\partial_{\mu} g_{\nu\sigma} + \partial_{\nu} g_{\mu\sigma} - \partial_{\sigma} g_{\mu\nu} \right). \tag{10}$$

This connection is torsion-free and metric-compatible.

3. Riemann Curvature Tensor:

$$R^{\lambda}_{\ \mu\nu\sigma} = \partial_{\nu}\Gamma^{\lambda}_{\mu\sigma} - \partial_{\sigma}\Gamma^{\lambda}_{\mu\nu} + \Gamma^{\lambda}_{\nu\rho}\Gamma^{\rho}_{\mu\sigma} - \Gamma^{\lambda}_{\sigma\rho}\Gamma^{\rho}_{\mu\nu}. \tag{11}$$

4. Ricci Tensor and Ricci Scalar:

$$R_{\mu\nu} = R^{\lambda}_{\ \mu\lambda\nu}, \quad R = g^{\mu\nu}R_{\mu\nu}. \tag{12}$$

5. Gravitational Action in GR:

$$S_{GR} = \frac{1}{2\kappa} \int d^4x \sqrt{-g} R + S_{\text{matter}}, \tag{13}$$

where $\kappa = 8\pi G$.

6. Einstein Field Equations:

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = \frac{8\pi G}{c^4}T_{\mu\nu}.$$
 (14)

Comparison: GR vs. TEGR

Concept	GR	TEGR
Fundamental Object	Metric tensor $g_{\mu\nu}$	Tetrad fields e^a_μ
Connection	Levi-Civita connection (torsion-free)	Weitzenböck connection (flat, torsion)
Torsion	Zero (torsion-free)	Non-zero
Curvature	Non-zero (Riemann tensor)	Zero (Weitzenböck connection)
Geometrical Property	Curved spacetime geometry	Flat spacetime with torsion
Action	Einstein-Hilbert action $\int \sqrt{-g}R$	Teleparallel action $\int eT$
Field Equations	Einstein field equations	Derived from torsion scalar T