

Ch 1: Boltzmann & Plank

$$h = 6.626 \times 10^{-34} \text{ Js}$$

$$k_B = 1.38 \times 10^{-23} \text{ J/K}$$

$$c = 3 \times 10^8 \text{ m/s}$$

① BLACK BODY RADIATION

★ For a body in thermal Eqbm:

$$e_\nu = J(\nu, T) a_\nu$$

emitted power per unit A_ν , per unit freq \rightarrow fraction of incident power absorbed.

For black Body. $\Rightarrow e_\nu^{BB} = J(\nu, T)$

★ Density (per unit vol) of standing waves in the freq interval ν and $\nu + d\nu$

$$N(\nu) d\nu = \frac{8\pi\nu^2}{c^3} d\nu \rightarrow \text{we'll derive this soon enough.}$$

★ No of oscillators btwn energy E & $E + dE$: $N(E) dE = dN$ & $N = \int dN = \int_0^\infty N(E) dE$

★ Maxwell-Boltzmann:

$$N(E) dE = \frac{N}{k_B T} e^{-E/k_B T} dE$$

★ Avg value of Energy: $\bar{E} = \frac{\int_0^\infty E \cdot N(E) dE}{\int_0^\infty N(E) dE} = k_B T$ for M.B. distribution of Energy.
 (probability \uparrow)

★ Define $u(\nu) d\nu \equiv$ Energy per unit volume in freq interval $d\nu$
 (We integrate because every energy $E \in (0, \infty)$ is allowed)

$$u(\nu) d\nu = k_B T N(\nu) d\nu = \frac{8\pi\nu^2}{c^3} k_B T d\nu \rightarrow \text{RAYLEIGH JEANS LAW}$$

$$\{u(\nu) d\nu = \bar{E} N(\nu) d\nu\}$$

Remember: $J(\nu, T) = u(\nu, T) \times \frac{c}{4} \rightarrow$ (won't derive this)

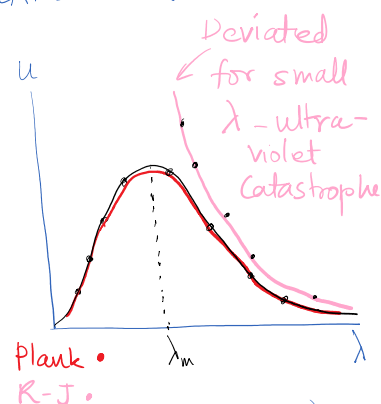
★ PLANK'S GUESS \leftrightarrow Not every energy level is allowed

$$E_{\text{oscillator}} = nh\nu = E_n$$

(later the exact $(n + \frac{1}{2})h\nu$ - doesn't make a diff here)

$$\therefore \bar{E} = \frac{\sum_{n=0}^{\infty} E_n \cdot e^{-E_n/k_B T}}{\sum_{n=0}^{\infty} e^{-E_n/k_B T}} = \frac{\sum_{n=0}^{\infty} nh\nu e^{-nh\nu/k_B T}}{\sum_{n=0}^{\infty} e^{-nh\nu/k_B T}} = \frac{h\nu}{\left[\exp\left(\frac{h\nu}{k_B T}\right) - 1 \right]}$$

Plank $\Rightarrow u(\nu) d\nu = \frac{8\pi\nu^2}{c^3} \times \frac{h\nu}{\left[\exp\left(\frac{h\nu}{k_B T}\right) - 1 \right]} d\nu \rightarrow \text{Q.M.}$



↓
Energy is
QUANTIZED = discrete

Plank \Rightarrow

$$u(\nu) d\nu = \frac{8\pi\nu^2}{c^3} \times \frac{h\nu}{\left(\exp\left(\frac{h\nu}{k_B T}\right) - 1\right)} d\nu$$

\rightarrow Q.M.
= Success!

* Stefan's Law: $J = \sigma T^4 \Rightarrow$ Power = $\sigma A T^4$ $\sigma = 5.67 \times 10^{-8}$ S.I. units
 \hookrightarrow Radiant power per unit area.

* Wien's Law: $\lambda_m T = 2.89 \times 10^{-3} \text{ m}\cdot\text{K}$
 \downarrow \hookrightarrow Wein's displacement const
 Wavelength with Max intensity
 on a BBR spectrum