* Quantifying Curvature: Parallel transport va along

A > B: 2 is constant; moving along =

V=v=0 (parallel transport) > 2-Vx+ TxruVM=0

 $\Rightarrow \frac{\partial V^{\times}}{\partial x^{*}} = - \mathcal{V}^{\times} \mathcal{V}^{\mu} \qquad \text{integrate this} \qquad V^{\times}(B) = V^{\times} \text{initial} - \int \mathcal{V}^{\times} \mathcal{V}^{\mu} dx^{*}$

Similarly: V*(c) = V*(B) - IT ">MV dx

 $V^{\kappa}(D) = V^{\kappa}(c) + \int \Gamma^{\kappa} \nabla^{\mu} dz^{\kappa}$

 $V_{\text{final}}^{\kappa}(A) = V_{\kappa}^{\kappa}(D) + \int_{\mathcal{Y}} \Gamma_{\lambda\mu} V_{\mu}^{\mu} dx^{\lambda} \qquad \text{@xr} \quad \text{@xr} + \xi x r$ Now: $SV = V_{final} - V_{initial} = \int_{\Theta} \nabla^{\times} \lambda_{\mu} V^{\mu} dx^{\lambda} - \int_{\Omega} \nabla^{\times} \lambda_{\mu} V^{\mu} dx^{\lambda}$

+ I I K of My dx - I I K of My dx $\left\{ \int (x_0 + dx) \simeq \int (x_0) + \int (x_0) dx \right\}$

Use Taylor expansion:

Svx $\sim -\int_{-\infty}^{\infty} 8x^{2} \frac{3x^{2}}{3} \left(\int_{-\infty}^{\infty} \sqrt{y} \sqrt{y} \right) dxy + \int_{-\infty}^{\infty} 8x^{3} \frac{3x^{4}}{3} \left(\int_{-\infty}^{\infty} \sqrt{y} \sqrt{y} \right) dx$

 $\left\{ \int f(x) dx \simeq \varepsilon \cdot f(x) \text{ for } \varepsilon <<1 \right\}$

for 11 toansport: 2 ~ ~ + ~ ~ ~ ~ ~ ~ ~ = 0

⇒ Svx = Sxx Sx TV Rx mhr

R ~ mr = 2, T ~ m - 2- T ~ m + T ~ n - T ~ r T ~ m - T ~ r T ~ m Curvature Tensor

Equivalently: VMR MAR = [V, V,] Vx P., R . X = [7, 7] P.

 $[\nabla_{\lambda}, \nabla_{r}] \equiv Commutator$

0 1 WWW L 1471 1671

$$P_{\mu} R^{\mu} \times \lambda_{r} = [\nabla_{\lambda_{1}} \nabla_{\sigma}] P_{\kappa}$$

$$[\nabla_{\lambda}, \nabla_{\Gamma}] = Commutator$$

= 2,2m(ln/g1 1/2)

Symmetries of the Riemann tensor reduce the no. of indep components from n4 to $\frac{n^2(n^2+1)}{12}$ —D for 4D S·T· it is $\frac{4^2 \times 15}{12} = \boxed{20}$

* Symmetries of the RCT:

$$\left(\mathbb{R}_{\alpha\mu\lambda\Gamma}\right)_{LL.F.} = 2_{\lambda}\Gamma_{\alpha\sigma\mu} - 2_{\Gamma}\Gamma_{\kappa\lambda\mu} = \frac{1}{2}\left(2_{\lambda}\partial_{\mu}g_{\kappa\Gamma} - 2_{\lambda}\partial_{\kappa}g_{\Gamma\mu} - 2_{\Gamma}\partial_{\mu}g_{\kappa\lambda} + 2_{\Gamma}\partial_{\kappa}g_{\lambda\mu}\right)$$

We get:

(4)
$$R_{\alpha\mu\lambda\tau} + R_{\alpha\lambda\tau\mu} + R_{\alpha\tau\mu\lambda} = 0 \Rightarrow R_{\alpha[\mu\lambda\tau]} = 0$$

$$R^{\alpha}_{\mu\alpha\nu} = g^{\alpha\beta}R_{\beta\mu\alpha\nu} = R_{\mu\nu} = \partial_{\alpha}\Gamma^{\alpha}_{\mu\nu} - \partial_{\nu}\Gamma^{\alpha}_{\alpha\mu} - \partial_{\nu}\Gamma^{\alpha}_{\alpha\mu} + \Gamma^{\alpha}_{\alpha\beta}\Gamma^{\beta}_{\nu\mu}$$

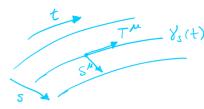
$$= \int_{\alpha} \Gamma^{\alpha}_{\mu\nu} - \partial_{\nu}\Gamma^{\alpha}_{\nu\mu} - \int_{\alpha} \Gamma^{\alpha}_{\nu} - \Gamma^{\alpha}_{\nu} \Gamma^{\beta}_{\nu} - \Gamma^{\alpha}_{\nu} \Gamma^{\alpha}_{\nu} - \Gamma^{\alpha}_{\nu} - \Gamma^{\alpha}_{\nu} \Gamma^{\alpha}_{\nu} - \Gamma^$$

* The Einstein tensor:
$$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu}$$
 at Trace reversed Ricci tensor.

* Geodesic Deviation:

* Geodisic Deviation:

- Tangent vectors to Geodusia: $T^M = \frac{\partial x^M}{\partial t}$
- Deviation vectors: $S^{\mu} = \frac{3x^{\mu}}{2s}$



- · V" = (V,S)" = TPV,S" -> relative velocity of Geodisia

 $A^{\mu} = (\nabla_{\tau} V)^{\mu} = T^{r} \nabla_{\rho} V^{\mu} \rightarrow \text{ relative accl}^{n} \text{ of Geodisics}$ The accl of a path away from being a geodesic: $a^{\mu} = T^{r} \nabla_{\sigma} T^{\mu}$

Since S . T are basis vectors adapted to a cood system, their commutator vanishus:

$$[s,T] = 0$$
 ; also: $S^{\rho} \nabla_{\rho} T^{M} = T^{\rho} \nabla_{\rho} T^{M}$

The acceleration $A^{\mu} = T^{\rho} \nabla_{\rho} (T^{\sigma} \nabla_{\sigma} S^{\mu}) = R^{\mu}_{\nu \rho \sigma} T^{\nu} T^{\rho} S^{\sigma}$

$$A^{\mu} = \frac{D^2}{dt^2} S^{\mu} = R^{\mu}_{\gamma \rho \sigma} T^{\gamma} T^{\rho} S^{\sigma}$$
 as Geodesic deviation