# Stabilizing Integrators for Real-Time Physics

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We present a new time integration method featuring excellent stability and energy conservation properties, making it particularly suitable for real-time physics. The commonly used backward Euler method is stable, but introduces artificial damping. Methods such as implicit midpoint do not suffer from artificial damping but are unstable in many common simulation scenarios. We propose an algorithm which blends between the implicit midpoint and forward/backward Euler integrators such that the resulting simulation is stable while introducing only minimal artificial damping. We achieve this by tracking the total energy of the simulated system, taking into account energy-changing events: damping and forcing. To facilitate real-time simulations we propose a local/global solver, similar to Projective Dynamics, as an alternative to Newton's method. Compared to the original Projective Dynamics, which is derived from backward Euler, our final method introduces much less numerical damping at the cost of minimal computing overhead. Stability guarantees of our method are derived from the stability of backward Euler, whose stability is a widely accepted empirical fact. However, to our knowledge, theoretical guarantees have so far only been proven for linear ODEs. We provide preliminary theoretical results proving the stability of Backward Euler also for certain cases of non-linear potential functions.

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#### 1. INTRODUCTION

Numerical time integration of the equations of motion has been a classical problem in engineering since the seminal work of Euler. Numerical integration is also a key ingredient of physics-based animation. However, in computer graphics, we do not necessarily strive for accurate numerical solutions of differential equations, but rather for physically plausible results. Simply put, the resulting motion needs to *look right*, depending on the needs of a particular application. The discrepancy between accuracy and plausibility is especially pronounced in real-time applications, such as computer games or surgery simulators. Interactive applications need to refresh the screen at fixed time intervals, typically 33 ms or even less, to create the illusion of smooth motion. In real-time simulations we need to advance the state of the virtual world by 33 ms while using strictly less than 33 ms of computing time on a given hardware (CPU/GPU). Games or training simulators are complex

software systems composed of many sub-systems (such as rendering, networking, human-computer interaction, ...) and therefore the time budget for physics will be typically only a small fraction of the total frame time (33 ms). Historically, rigid-body physics has been the first success story of real-time simulations; each rigid body has only six degrees of freedom. The situation is more complicated with deformable objects, such as biological soft tissues, which require many more degrees of freedom and, consequently, much more computation.

Adaptive time stepping methods with error control are popular in scientific computing. These methods are necessary in engineering applications, e.g., when designing an airplane or nuclear reactor, where simulation accuracy may be critical. Unfortunately, adaptive time stepping is incompatible with the requirements of real-time applications, where we have only a limited computing budget per frame. Despite progress in parallel-in-time methods, time stepping is a fundamentally sequential process, difficult to parallelize. Realtime simulations therefore have to compromise accuracy in order to retain interactivity. However, the goal is to do this gracefully and retain physical plausibility by keeping the inevitable errors under control. The most striking manifestation of inaccuracy is the case of instabilities ("explosions"), where small discretization errors compound and the discrete approximation departs dramatically from the true continuous solution. The classical way of avoiding such catastrophic failures is to reduce the time step. Unfortunately, this is not an option in real-time physics which needs to operate within limited computing budget.

An important feature of numerical time integrators is their conservation properties, i.e., which features of the exact continuous solution are mimicked by the numerical integrator. There are three main properties that are conserved in non-dissipative mechanical systems: total energy (i.e., the sum of kinetic and potential energies, also known as the Hamiltonian), momentum (angular and linear), and the symplectic form. Previous work can be broadly classified in two main branches: energy-momentum conserving methods and symplectic methods. Energy-momentum methods [Simo et al. 1992] try to exactly preserve the energy and momenta; however, this does not always result in plausible simulations, as we explain in Sec. 2. Symplectic integrators [Stern and Desbrun 2006] focus on conserving the symplectic form, which implies momentum conservation and also good energy behavior - the total energy oscillates around the correct value. Unfortunately, with stiff systems and fixed time steps, these oscillations can be extreme and result in visual "instabilities" or "explosions" (see Fig. 2 for a didactic example). There are various definitions of stability in the literature, but an informal understanding is that simulations dramatically depart from the true solution. Since neither symplectic methods nor

energy-momentum methods can guarantee visual plausibility for large time steps, most real-time physics-based simulators continue to rely on backward Euler time integration, despite its numerical damping. Higher-order backward methods, such as BDF2, reduce this damping, but the remaining numerical dissipation is still obvious.

In this paper, we propose a novel time integration method for realtime simulation of deformable objects which inherits the stability of backward Euler but does not suffer from artificial numerical damping. We achieve this by starting with the implicit midpoint integrator. Implicit midpoint does not introduce artificial damping but, unfortunately, the energy oscillations can grow out of proportion and produce visually catastrophic results. We observe that these oscillations can be tamed by adding only a small contribution of backward Euler integration. Similarly, in cases where the initial implicit midpoint solve underestimates the total energy, we can correct this by blending with forward Euler. The key is to determine the right amount of blending between implicit midpoint and forward/backward Euler. We do this by tracking the total energy of our simulated system, taking account of energy-changing events such as damping (energy dissipation) or forcing (energy injection). When we detect that implicit midpoint overshoots the total energy (indicating that future time steps could develop instabilities), we calculate which blend between implicit midpoint and backward Euler will result in an "as-energy-conserving-as-possible" state. Usually, only a very small contribution of backward Euler is sufficient to stabilize the simulation while introducing only a minimal amount of numerical damping. Similarly, we use forward Euler blending when implicit midpoint loses energy. This is not critical for guaranteeing stability but it helps to improve the visual quality of the resulting motion.

Our method derives its stability from backward Euler, Accurately solved backward Euler is known to be stable in practice, even with complicated non-linear elastic potentials. Insufficiently accurate numerical solutions of the backward Euler equations can introduce instabilities [Baraff and Witkin 1998], however, these instabilities are avoidable by iterating Newton's method until convergence, resulting in highly accurate backward Euler solutions. The stability of (exact) backward Euler is well studied in case of quadratic potentials, corresponding to linear ODEs [Ascher and Petzold 1998]. However, non-linear deformation energies such as mass-spring systems or corotated elasticity are common in graphics applications. To our knowledge, there is no backward Euler stability proof for such non-linear potentials. In the Appendix, we provide a backward Euler stability proof for convex potential functions. Further in the Appendix, we discuss the challenges of generalizing this proof to non-convex potentials and present a stability proof for a simple yet non-convex test potential (two connected springs). Even though this does not settle the discussion of backward Euler stability with non-linear potentials, we hope that our results will inspire future research in this direction.

To obtain accurate solutions of backward Euler, we have to iterate Newton's method until convergence, which is impractically slow in real-time simulations. Therefore, a special class of quasi-Newton methods known as Projective Dynamics [Bouaziz et al. 2014; Liu et al. 2016], based on local/global optimization, has been developed as a computationally efficient alternative to Newton's method. Projective Dynamics is derived from backward Euler integration and therefore inherits the undesired numerical damping. In this paper, we also present a local/global acceleration for implicit midpoint and its stabilization via forward/backward Euler in order to facil-

itate high-quality real-time animations. Even though local/global solves are typically not iterated until convergence, our experiments demonstrate that we still obtain stable simulations while avoiding numerical damping. The additional computing overhead over standard Projective Dynamics is small, typically on the order of 20%-30% of extra computing time.

In summary, we present three main contributions: 1) stabilization of implicit midpoint via energy-tracking and blending with forward/backward Euler, 2) a fast local/global numerical procedure for implicit midpoint, and 3) proofs of backward Euler stability for certain types of non-linear potential functions. The first two contributions lead to a fast integrator with good conservation properties and the third provides some theoretical insights into its stability guarantees. We hope that our method will be useful especially in real-time simulations with immediate applications, e.g., in computer games or training simulators.

## 2. RELATED WORK

Integration in physics-based animation. The dramatic instabilities of Forward Euler have been observed already in the early days of physics-based animation and pioneering work from the 1980s applied backward Euler integration [Terzopoulos et al. 1987; Terzopoulos and Fleischer 1988a; Terzopoulos and Fleischer 1988b]. Subsequent works explored explicit methods such as the popular Runge-Kutta family, featuring simpler implementation and faster run-time, until Baraff and Witkin [1998] demonstrated the advantages of approximately solving backward Euler, using a method analogous to one-iteration of Newton's method without a line search. Despite the fact that the numerical damping of backward Euler is a well-known issue [Choi and Ko 2002; Stern and Desbrun 2006] many physics-based systems continue to use this method. Alternative methods such as BDF2 [Choi and Ko 2002; Bender et al. 2014] reduce the numerical damping, but do not eliminate it completely. Symplectic methods, including the Newmark family of integrators [Sin et al. 2013] (which for  $\gamma = 1/2$  is symplectic even though not in the traditional sense [West et al. 1999]) do not exhibit the undesired numerical dissipation. Unfortunately, this good behavior applies only for sufficiently small time steps, where "sufficiently" depends on the parameters of the simulated system [Gonzalez and Simo 1996; Sin et al. 2013]. At larger time steps, these oscillations can be very large. One solution is to adaptively change the time step [Kane et al. 1999]. Asynchronous integrators [Lew et al. 2003; Thomaszewski et al. 2008] offer even more flexibility by allowing us to vary the time step sizes spatially [Harmon et al. 2009; Ainsley et al. 2012; Zhao et al. 2016]. Instead of varying the time step, one could also use adaptive re-meshing [Narain et al. 2012]. However, these methods are not suitable for real-time physics, where we cannot afford fluctuations in computing time.

**Stability.** Stability of numerical integrators is theoretically well understood in the context of linear ODEs, where the problem essentially reduces to eigen analysis [Ascher and Petzold 1998; Iserles 2009]. However, the situation is much more complicated with nonlinear ODEs which are common in physics-based animation. As an illustration of this fact, note that the implicit midpoint method (a symplectic integrator) is unconditionally stable when applied to linear ODEs, i.e., the stability of the numerical approximation using implicit midpoint is equivalent to the stability of the true continuous solution, regardless of the size of the time step. However, this result does not hold with non-linear ODEs where implicit midpoint

becomes only conditionally stable [Gonzalez and Simo 1996; Kuhl and Crisfield 1999]. Previous work in physics-based animation cautions against the instabilities of symplectic schemes [Kharevych et al. 2006; Sin et al. 2013]. Numerical instabilities can be cured by reducing the time step size but, unfortunately, this is not possible in real-time applications which need to operate under strict computing limits.

Another way of analyzing the stability of numerical integrators is by observing their energy behavior. For example, forward Euler causes unbounded increases in the total energy during the course of a simulation, manifesting as "explosions". In this paper we consider a simulation *stable* if the total energy (the Hamiltonian) is bounded over an arbitrary number of integrator steps, assuming the potential is bounded from below (see Appendix B for a more formal discussion).

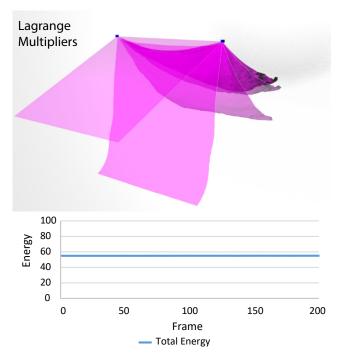


Fig. 1. A mass-spring system cloth model with gravity, with energy conservation enforced using the method of Lagrange Multipliers as described in [Hughes et al. 1978] (top) and the corresponding energy graph (bottom). While the energy is perfectly conserved, the resulting motion gets "stuck", resulting in a very poor animation (see the accompanying video).

Enforcing energy conservation. Unstable simulations are characterized by the total energy dramatically departing from its true value. A logical way to avoid this problem is by explicitly enforcing constraints on total energy. This can be implemented either as a post-processing (projection) step after running an arbitrary integrator [LaBudde and Greenspan 1975; Simo et al. 1992; Su et al. 2013] or by imposing an energy-conserving constraint using Lagrange multipliers [Hughes et al. 1978]. Even though enforcing energy conservation can help in some cases, there are no guarantees of improved accuracy. It has been demonstrated that enforcing energy conservation can lead to less accurate results [Hairer 2006]. This may seem counter-intuitive; after all, the exact continuous solutions

of Hamiltonian systems exactly conserve energy, so one could expect that enforcing this constraint would lead to better results. Unfortunately, this is not always the case, because numerical schemes inevitably deviate from the true continuous solution due to discretization errors. Strictly enforcing some aspects of the ideal continuous solution is dangerous, because we may destroy other desirable properties, such as symplecticity. For example, if numerical error is concentrated at one part of the simulated object, the energyconservation mechanism can non-physically "teleport" energy to remote parts of the object, which can lead to unrealistic motion (often manifesting itself as unnatural oscillations). Another caveat is that stability is not guaranteed with projection-type methods [Simo et al. 1992; Su et al. 2013] because kinetic energy cannot be decreased below zero. Lagrange multiplier approaches guarantee stability with arbitrarily large time-steps [Hughes et al. 1978], but do not guarantee accuracy. Fig. 1 shows a simple mass-spring system cloth falling under gravity with energy conservation enforced using Lagrange Multipliers [Hughes et al. 1978]. Even though the energy is conserved perfectly, the simulation is not visually plausible – the cloth gets "stuck" in one position and the individual elements begin to oscillate in place. This is because Newton's method arrives at a local minimum that satisfies the energy-conserving constraint, but does not correspond to plausible motion.

Symplectic methods. Flows of mechanical systems without dissipation (Hamiltonian systems) exactly conserve not only energy and momenta, but also the symplectic form [Easton 1998]. Numerical integrators which conserve the symplectic form are called symplectic integrators. Symplectic integrators can be elegantly derived by discretizing the principle of least action [West 2004; Hairer et al. 2006; Stern and Desbrun 2006]. Using backward error analysis it is possible to show that symplectic integrators exhibit good longtime energy behavior [Hairer et al. 2006], in particular, the energy oscillates about its true value. However, this is true only with sufficiently small time steps; otherwise, symplectic integrators (both explicit and implicit ones) can oscillate dramatically [Kuhl and Crisfield 1999; Ascher and Reich 1999]. One can attempt to correct this via energy preservation as discussed in the previous paragraph. Unfortunately, there are known theoretical limitations – fixed time step methods cannot have all three of the following properties: 1) symplecticity, 2) energy conservation, 3) momentum conservation [Zhong and Marsden 1988]. It is possible to achieve all three with adaptive time stepping [Kane et al. 1999]. However, as previously discussed, adaptive time stepping does not meet the needs of realtime simulations. Recently, there has been renewed interest in exponential integrators which combine analytical solutions of the linear part of the ODEs with numerical solutions for the non-linear residual [Michels et al. 2014]. Exponential integrators are computationally efficient, robust, and well suited for applications e.g. in molecular dynamics [Michels and Desbrun 2015]. However, just like with other symplectic schemes, stability is not guaranteed, which is problematic in real-time simulations where there is no "second chance" to re-run the simulation in case of instability.

Generalized- $\alpha$  Methods. Good energy behavior is exhibited also by methods such as the well-known Newmark family of algorithms [Newmark 1959]. A more general schemes, such as the generalized alpha method have been developed [Chung and Hulbert 1993] and applied to computer graphics [Sobottka et al. 2008]. These integrations methods provide several parameters. [Kane 1999] showed that the Newmark- $\beta$  method is symplectic with the common setting  $\beta=1/4, \gamma=1/2$ , even though not in the traditional sense of the canonical symplectic form. Unfortunately, similarly to the sym-

plectic methods discussed above, with a fixed time step there is no guarantee of visually plausible results.

Position based methods. Position Based Dynamics [Müller et al. 2007] is a physics-based simulation approach which specifically caters to the needs of real-time applications such as computer games. Closely related techniques have been followed in the Nucleus solver [Stam 2009]. Position-based methods have been extended with more advanced solvers [Müller 2008; Wang et al. 2010; Kim et al. 2012], finite element models [Müller et al. 2014; Bender et al. 2014], fluid simulation [Macklin and Müller 2013] and a unified simulator supporting multiple phases of matter [Macklin et al. 2014]. As pointed out by Liu [2013], Position Based Dynamics can be derived from backward Euler integration (BDF1) and therefore inherits its artificial numerical damping. It is possible to upgrade Position Based Dynamics to BDF2 [Bender et al. 2014]. This helps, but it does not completely eliminate the undesired numerical dissipation. BDF2 can be replaced with, e.g., implicit midpoint or another implicit symplectic scheme which avoids numerical damping, but this compromises stability. Real-time applications cannot risk instability and therefore current systems accept the "lesser evil" of uncontrollable numerical damping.

Projective Dynamics. Backward Euler is often used for real-time applications due to its stability. Newton's method is the classic method for computing accurate solutions to optimization problems, and is typically used to compute the solution to backward Euler in real-time animation. Unfortunately, it can be slow because the Hessian matrix changes at every iteration. In real-time applications, accuracy is usually a secondary concern to visual plausibility. To remedy these problems and provide an integration technique suitable for real-time applications, Projective Dynamics, introduced by [Bouaziz et al. 2014], provides a fast method for solving backward Euler which trades accuracy for speed while providing stability with fixed time steps. Projective Dynamics has been further accelerated by advanced numerical techniques [Wang 2015] and extended to more general materials [Liu et al. 2016; Narain et al. 2016]. However, all of these methods are still based on backward Euler and inherit its artificial damping properties.

## BACKGROUND

**Forward Euler.** Forward Euler (sometimes called explicit Euler) uses only the current state to update the next state; it is an explicit method. The update rules are very simple:

$$\mathbf{x}_{n+1} = \mathbf{x}_n + h\mathbf{v}_n \tag{1}$$

$$\mathbf{v}_{n+1} = \mathbf{v}_n + h\mathbf{M}^{-1}f(\mathbf{x}_n) \tag{2}$$

where  $\mathbf{x}_{n+1}$  and  $\mathbf{v}_{n+1}$  are the positions and velocities at the next state,  $\mathbf{x}_n$  and  $\mathbf{v}_n$  are the positions and velocities at the current state,  $\mathbf{M}$  is the mass matrix, h is the time step, and  $f(\mathbf{x}_n)$  is the net force evaluated at  $\mathbf{x}_n$ . Unfortunately, this scheme is not very robust, unless the time step is very small. Even though using only the current state to estimate the next state is straightforward and intuitive, it typically results in an over-estimation of the total energy at every time step. In Appendix A we prove that for convex potential functions, forward Euler cannot decrease the total energy of the system. With larger time steps, there are often significant increases of the total energy, leading to the commonly observed instabilities (or "explosions"). The analysis is more complicated for non-convex potentials as we discuss in Appendix C.

**Backward Euler.** While forward Euler uses the current state to compute the next state, backward Euler uses the next state,  $\mathbf{x}_{n+1}$  and  $\mathbf{v}_{n+1}$ . Since we no longer have closed-form formulas for calculating the next state, we have to solve a system of (typically nonlinear) equations to obtain  $\mathbf{x}_{n+1}$  and  $\mathbf{v}_{n+1}$ :

$$\mathbf{x}_{n+1} = \mathbf{x}_n + h\mathbf{v}_{n+1} \tag{3}$$

$$\mathbf{v}_{n+1} = \mathbf{v}_n + h\mathbf{M}^{-1}f(\mathbf{x}_{n+1}) \tag{4}$$

This is more computationally intensive than explicit methods, but the advantage is that the usual CFL (Courant-Friedrichs-Lewy) limitations no longer apply. Simply put, explicit methods can propagate forces only to the immediately adjacent nodes. E.g, when simulating a cloth picked up at one corner, the external contact forces propagate very slowly throughout the entire system. Implicit methods can achieve such propagation during one step. The use of backward Euler in physics-based animation has been popularized by Baraff and Witkin [1998] who demonstrated its superior behavior compared to forward Euler, especially with larger time steps and stiff systems. Backward Euler has the exact opposite error behavior compared to forward Euler: it underestimates the energy of the true solution. This is not always a problem because some amount of dissipation can in fact improve the visually plausibility of the motion. Unfortunately, the amount of numerical dissipation of backward Euler is not explicitly controllable by the user and instead indirectly depends on resolution, time step, and stiffness. Backward Euler also results in loss angular momentum. Higher-order backward methods such as BDF2 produce more accurate solutions, but uncontrolled numerical dissipation is still present. Asynchronous timestepping methods such as [Zhao et al. 2016] also mitigate the numerical damping of backward Euler, but at the cost of introducing variable computing demands. This is not desirable in real-time simulations where each frame should ideally take the same amount of computing resources.

Backward Euler is known to be very stable, but without much understanding as to why. While proofs exist for the unconditional stability of backward Euler for linear ODE's and basic stability analysis has been done for some non-linear ODE's [Dekker and Verwer 1987; Lambert 1991], we are not aware of any general proofs in the non-linear case, such as for the deformation energies commonly used in computer animation. In Appendix B we present a stability proof for backward Euler for convex potential functions. As in the forward Euler case, the analysis becomes more difficult for non-convex potentials, even simple ones such as mass-spring

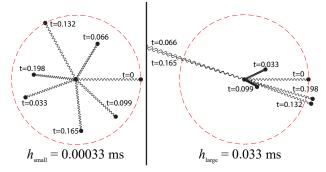


Fig. 2. A simple spring-type potential run using Implicit Midpoint at a time step  $h_{\rm small}=0.00033$  (left) and  $h_{\rm large}=0.033$  (right). Each point represents a state at a time t. The red circle illustrates initial stretched length of the spring.

systems (see Appendix C). Note that our theoretical analysis presented in the Appendix assumes exact solutions of Equations 3 and 4. Instability can be also introduced by inaccurate numerical solvers (which compute only approximate solution of Equations 3 and 4), such as using only one iteration of Newton's method). These numerics-induced instabilities have been observed already by Baraff and Witkin [1998] who proposed a solution by adapting the time step.

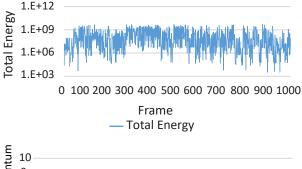
Implicit Midpoint. Intuitively, it is obvious that both forward and backward Euler are biased: the former relies entirely on the current state, while the latter relies entirely on the next state. implicit midpoint can be seen as a compromise between the two: instead of using just the current or just the next state, implicit midpoint uses their average:

$$\mathbf{x}_{n+1} = \mathbf{x}_n + \frac{h}{2}(\mathbf{v}_{n+1} + \mathbf{v}_n) \tag{5}$$

$$\mathbf{v}_{n+1} = \mathbf{v}_n + h\mathbf{M}^{-1}f\left(\frac{\mathbf{x}_{n+1} + \mathbf{x}_n}{2}\right)$$
 (6)

This integrator has very nice properties, namely it is symplectic and momentum-preserving. A side effect of the preservation of the canonical symplectic form is good energy-preserving behavior [Hairer et al. 2006]. However, this does not mean that the total energy has to be close to its exact value. Even in very simple simulations it is not hard to obtain extreme energy oscillations. For example, the simulation in Fig. 2 uses the potential function:

$$E_{\text{simpleSpring}}(\mathbf{x}) = \frac{1}{2}k(||\mathbf{x}||^2 - 1)^2 \tag{7}$$



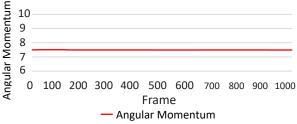


Fig. 3. Energy (top) and angular momentum (bottom) corresponding to the simulation from Fig. 2 with h = 0.033 s.

which is a squared version of a classical Hookean spring of length 1 m with one endpoint fixed at the origin. For our experiment we used stiffness  $k = 10^5$  N/m, a mass of 0.5 kg, two time steps  $h_{\rm small}=0.00033$  s (Fig. 2, left) and  $h_{\rm large}=0.033$  s (Fig. 2, right), an initial velocity of (0, 10, 0) m/s, and initial length of 1.5 (corresponding to the spring being pre-stretched). Fig. 2 shows that with  $h_{\text{large}}$ , the simulation does not behave plausibly even during the first few time steps – contrary to our expectations, the rotating spring does not return to its rest length (1 m) but instead further extends beyond the initial length of 1.5 m (indicated by the red circle). If we decrease the time step to  $h_{\mathrm{small}}$ , this behavior disappears and the revolving spring contracts as expected, staying within the red circle.

What happened with the larger time step h = 0.033 s? We plot the total energy of this simple system in Fig. 3. We can see that the simulation reaches incredibly large energy levels – the system starts out at an energy level of around 25,000 J but quickly proceeds to energy levels of several orders of magnitude higher, at which point the moving endpoint flies off the screen with unrealistic speed. Even though the energy is fluctuating wildly, the angular momentum is still conserved, as shown in Fig. 3. This poor behavior of implicit midpoint can be observed even with more complex models, such as the rotating cube modeled with corotated elasticity in Fig. 4.

A related integrator comes from using the trapezoid rule instead of the midpoint rule. This corresponds to the Newmark integrator with  $\gamma = 1/2$  and  $\beta = 1/4$ . While this integrator can behave better than implicit midpoint, it is still prone to explosions (Fig. 4).

As a final remark, we note that implicit midpoint can also be written as a composition of backward and forward Euler:

$$\mathbf{x}_{n+\frac{1}{2}} = \mathbf{x}_n + \frac{h}{2} \mathbf{v}_{n+\frac{1}{2}}$$
 (8)

$$\mathbf{v}_{n+\frac{1}{2}} = \mathbf{v}_n + \frac{h}{2}\mathbf{M}^{-1}f(\mathbf{x}_{n+\frac{1}{2}})$$
 (9)

$$\mathbf{x}_{n+1} = \mathbf{x}_{n+\frac{1}{2}} + \frac{h}{2} \mathbf{v}_{n+\frac{1}{2}}$$
 (10)

$$\mathbf{v}_{n+1} = \mathbf{v}_{n+\frac{1}{2}} + \frac{h}{2}\mathbf{M}^{-1}f(\mathbf{x}_{n+\frac{1}{2}})$$
 (11)

By substituting Eq. 10 into Eq. 8 and substituting Eq. 11 into Eq. 9, we get:

$$\mathbf{x}_{n+\frac{1}{2}} = \frac{\mathbf{x}_n + \mathbf{x}_{n+1}}{2}$$
 (12)  
 $\mathbf{v}_{n+\frac{1}{2}} = \frac{\mathbf{v}_n + \mathbf{v}_{n+1}}{2}$  (13)

$$\mathbf{v}_{n+\frac{1}{2}} = \frac{\mathbf{v}_n + \mathbf{v}_{n+1}}{2} \tag{13}$$

When we substitute these back into Eq. 10 and Eq. 11, we get the Implicit Midpoint update equations Eq. 5 and Eq. 6. If we instead do a step of forward Euler first and backward Euler second, we get the trapezoidal rule. These interesting facts motivate our method: perhaps some other combination of basic integrators could lead to a better time integration method for real-time physics?

#### **METHOD**

**Overview.** We will first explain our method in the case of a Hamiltonian system, i.e., dynamic system without any forcing or damping (we will discuss how to support these phenomena in Sec. 4.3). Our method is composed of two main steps. First, we calculate a time step according to the implicit midpoint rule, which conserves momenta and the symplectic form and therefore provides an excellent "initial guess". As demonstrated in the previous section, the main problem of implicit midpoint is that wide oscillations of the total energy can occur, departing dramatically from a visually plausible solution. In the second step, we therefore calculate the energy increase/decrease due to implicit midpoint (with a Hamiltonian system, the exact solution conserves the total energy). If the energy

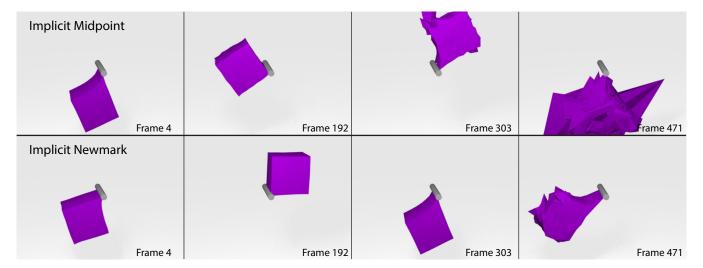


Fig. 4. A spinning deformable cube modeled with corotated elasticity simulated with time step h = 0.033. After several hundred time steps, the implicit midpoint and implicit Newmark integrator start producing visually implausible results.

erroneously increased, we correct this by computing a backward Euler step, which dissipates energy. A detail which will be important later is that the backward Euler solution process uses the state computed by implicit midpoint as an initial guess. We then solve for the optimal linear blending parameter which will bring the total energy as close as possible to its original value. Similarly, if we detect that implicit midpoint erroneously decreased energy, we perform analogous blending with the forward Euler step, taking advantage of its energy-injection property. Our experiments reveal that each time step typically only required a small amount of blending, meaning that we do not depart too far from the symplectic and momentum-conserving solution given by implicit midpoint. This is also why we choose implicit midpoint instead of an explicit symplectic integrator; explicit integrators "explode" more dramatically at large time steps and would lead to a larger amount of backward Euler blending.

## 4.1 Optimization Form

Both implicit midpoint and backward Euler methods require an iterative solution process. In this section we will focus on accurate (up to rounding errors) solutions using Newton's method, and in Sec. 4.2 we will discuss fast approximate solvers based on a local/global approach. In both cases, it is advantageous to formulate the implicit midpoint rule as an optimization problem (as [Gast and Schroeder 2014; Gast et al. 2015] did for backward Euler). To derive this optimization problem, we start by substituting Eq. 6 into Eq. 5, resulting in:

$$\mathbf{x}_{n+1}^{\text{IM}} = \mathbf{x}_n + h\mathbf{v}_n + \frac{h^2}{2}\mathbf{M}^{-1}f\left(\frac{\mathbf{x}_{n+1}^{\text{IM}} + \mathbf{x}_n}{2}\right)$$
(14)

We assume that both  $\mathbf{x}_n$  and  $\mathbf{v}_n$  are known. For conciseness, in the following we denote the unknown state as  $\mathbf{x}$  (instead of  $\mathbf{x}_{n+1}^{\mathrm{IM}}$ ) and we define a new symbol  $\mathbf{y} := \mathbf{x}_n + h\mathbf{v}_n$ . Rearranging terms, we obtain:

$$\mathbf{M}(\mathbf{x} - \mathbf{y}) - \frac{h^2}{2} f\left(\frac{\mathbf{x} + \mathbf{x}_n}{2}\right) = 0$$
 (15)

In order to turn this into an optimization problem which finds  $\mathbf{x}$ , we need to anti-differentiate Eq. 15 with respect to  $\mathbf{x}$ . With conservative forces (we will discuss non-conservative forces in Sec. 4.3), we can write  $f(\mathbf{x}) = -\nabla E(\mathbf{x})$ , where E is an energy potential function. This allows us to perform the anti-differentiation:

$$g(\mathbf{x}) = \frac{1}{2} (\mathbf{x} - \mathbf{y})^T \mathbf{M} (\mathbf{x} - \mathbf{y}) + h^2 E \left( \frac{\mathbf{x} + \mathbf{x}_n}{2} \right)$$
(16)

The critical points of g, i.e.,  $\nabla_{\mathbf{x}} g(\mathbf{x}) = 0$ , correspond to the solutions of Eq. 15. Having computed the positions  $\mathbf{x}$  (which correspond to  $\mathbf{x}_{n+1}^{\mathrm{IM}}$ ), computing the velocities is straightforward from Eq. 5:

$$\mathbf{v}_{n+1}^{\mathrm{IM}} = \frac{2}{h} (\mathbf{x}_{n+1}^{\mathrm{IM}} - \mathbf{x}_n) - \mathbf{v}_n$$
 (17)

In order to apply Newton's method to minimize g, we need the gradient and Hessian of  $g(\mathbf{x})$ . These can be computed easily:

$$\nabla_{\mathbf{x}} g(\mathbf{x}) = \mathbf{M}(\mathbf{x} - \mathbf{y}) + \frac{h^2}{2} \nabla E\left(\frac{\mathbf{x} + \mathbf{x}_n}{2}\right)$$
(18)

$$\nabla_{\mathbf{x}\mathbf{x}}^{2}g(\mathbf{x}) = \mathbf{M} + \frac{h^{2}}{4}\nabla^{2}E\left(\frac{\mathbf{x} + \mathbf{x}_{n}}{2}\right)$$
(19)

In this section we assume that a solution  $(\mathbf{x}_{n+1}^{\mathrm{IM}}, \mathbf{v}_{n+1}^{\mathrm{IM}})$  has been computed by iterating Newton's method until convergence, i.e., achieving high numerical accuracy. The next task is to compute our final state  $(\mathbf{x}_{n+1}^{\mathrm{our}}, \mathbf{v}_{n+1}^{\mathrm{our}})$  by correcting for the total energy error in  $(\mathbf{x}_{n+1}^{\mathrm{IM}}, \mathbf{v}_{n+1}^{\mathrm{IM}})$ .

**Energy Error.** Minimizing the energy error corresponds to bringing the following function as close as possible to zero:

$$e(\mathbf{x}, \mathbf{v}) = E(\mathbf{x}) + K(\mathbf{v}) - H_{\text{total}}$$
 (20)

where  $K(\mathbf{v}) = \frac{1}{2}\mathbf{v}^\mathsf{T}\mathbf{M}\mathbf{v}$  is the kinetic energy,  $E(\mathbf{x})$  is the potential energy, and  $H_{\text{total}}$  is the initial total energy of the system, specified by the initial conditions  $(\mathbf{x}_0, \mathbf{v}_0)$ . If  $e(\mathbf{x}_{n+1}^{\text{IM}}, \mathbf{v}_{n+1}^{\text{IM}}) > 0$  it means that implicit midpoint erroneously increased the total energy, in which case we take advantage of the numerical dissipation properties of backward Euler in order to reduce e. When  $e(\mathbf{x}_{n+1}^{\text{IM}}, \mathbf{v}_{n+1}^{\text{IM}}) < 0$ , it

means that implicit midpoint incorrectly decreased the total energy, which we correct using forward Euler.

**Blending.** The way we stabilize energy overshoots is by calculating a blend between the implicit midpoint result (which resulted in erroneous energy injection) and the backward Euler result (which may potentially remove too much energy). The results of two integrators are blended linearly, using a blending parameter  $\alpha \in [0,1]$ :

$$\mathbf{x}_{n+1} = (1 - \alpha)\mathbf{x}_{n+1}^{\mathrm{IM}} + \alpha\mathbf{x}_{n+1}^{\mathrm{BE}}$$
 (21)

$$\mathbf{v}_{n+1} = (1 - \alpha)\mathbf{v}_{n+1}^{\mathrm{IM}} + \alpha\mathbf{v}_{n+1}^{\mathrm{BE}}$$
 (22)

A higher  $\alpha$  would yield a solution closer to backward Euler. At the extreme,  $\alpha=1$  would result in a full step of backward Euler. Assuming that backward Euler is stable, this implies stability of our method. Even though in Appendix C we prove stability of backward Euler only for certain potential functions, our experiments suggest that our method is stable even for more complex potentials. Even extreme initial conditions, as shown in Fig. 6, do not introduce instabilities in our method. Most practical simulations do not contain such large deformations and we only need to introduce a small amount of backward Euler, i.e.,  $\alpha$  is close to zero. These small modifications early on in the simulation make a large difference in the long-term stability and perceived vividness of the motion. Without these small  $\alpha$  modifications, the energy will gradually drift upwards over a long period of time, leading to "explosions" [Engle et al. 2005]. This can be observed in Fig. 4.

In an analogous fashion we handle energy under-estimates, which we correct by blending with forward Euler. Keeping the convention that higher  $\alpha$  yields a more "damped" solution, we replace  $(\mathbf{x}_{n+1}^{\mathrm{IM}}, \mathbf{v}_{n+1}^{\mathrm{IM}})$  with  $(\mathbf{x}_{n+1}^{\mathrm{FE}}, \mathbf{v}_{n+1}^{\mathrm{FE}})$  and replace  $(\mathbf{x}_{n+1}^{\mathrm{BE}}, \mathbf{v}_{n+1}^{\mathrm{BE}})$  with  $(\mathbf{x}_{n+1}^{\mathrm{IM}}, \mathbf{v}_{n+1}^{\mathrm{IM}})$  in Eq. 21 and Eq. 22. Now,  $\alpha=0$  corresponds to a full step of forward Euler and  $\alpha=1$  corresponds to a full step of implicit midpoint. This allows us to adjust the energy in a symmetric way, compensating for both over and under-shoots.

We can reformulate  $e(\mathbf{x}, \mathbf{v})$  from Eq. 20 as a function of  $\alpha$ . In the case of blending with backward Euler, we substitute Eq. 21 and Eq. 22 into Eq. 20, obtaining:

$$e(\alpha) = E\left((1 - \alpha)\mathbf{x}_{n+1}^{\mathrm{IM}} + \alpha\mathbf{x}_{n+1}^{\mathrm{BE}}\right) + K\left((1 - \alpha)\mathbf{v}_{n+1}^{\mathrm{IM}} + \alpha\mathbf{v}_{n+1}^{\mathrm{BE}}\right) - H_{\mathrm{total}}$$
(23)

The problem now becomes finding an  $\alpha$  value that brings the residual energy as close as possible to zero. This can be done using a simple binary search algorithm because  $e(\alpha)$  is continuous. During our experiments, we found that  $e(\alpha)$  is always monotonic, making the binary search rapidly converge to a solution.

Binary search requires a terminating condition. We base our terminating condition on the total energy at the last frame:

$$|e(\alpha)| < \epsilon H_{\text{total}}(\mathbf{x}_n, \mathbf{v}_n)$$
 (24)

 $\epsilon \in [0,1]$  is a variable that controls how accurately we will conserve the energy. A small value will result in strict energy conservation, while a larger  $\epsilon$  will result in a looser terminating condition. In practice, we found that a value of  $\epsilon = 0.01$  (corresponding to a 1% error in the energy) produced good results.

In rare cases, Backward Euler may not decrease the energy (we discuss an example in Appendix C). If both implicit midpoint and backward Euler increase the energy, then the binary search algorithm will not be able to find a root. When this happens, we simply

take the endpoint  $\alpha=1$  as a solution. In other words, we take a full step of backward Euler and the extra energy will be removed in future frames.

#### 4.2 Local/Global Solver

Projective Dynamics provides a computational advantage over Newton's method by solving backward Euler optimization problem using a local/global method, which allows us to precompute sparse Cholesky factors of the system matrix. In this section, we will provide a brief overview of Projective Dynamics, and then present a similar local/global optimization strategy for solving the implicit midpoint optimization problem (Eq. 16).

Projective Dynamics approximately solves the minimization formulation of backward Euler, which we can derive in the same manner as we derived the implicit midpoint formulation in Sec. 4.1:

$$g(\mathbf{x}) = \frac{1}{2} (\mathbf{x} - \mathbf{y})^T \mathbf{M} (\mathbf{x} - \mathbf{y}) + h^2 E(\mathbf{x})$$
 (25)

The key idea behind Projective Dynamics is constraint *projection*. First, we introduce an auxiliary "projection" variable  $\mathbf{p}$ . In order to take advantage of this auxiliary variable we need to define a special energy  $E(\mathbf{x})$  for each element i:

$$E_i(\mathbf{x}) = \min_{\mathbf{p}_i \in \mathcal{M}} E(\mathbf{x}, \mathbf{S}_i \mathbf{p}), E(\mathbf{x}, \mathbf{S}_i \mathbf{p}) = ||\mathbf{G}_i \mathbf{x} - \mathbf{S}_i \mathbf{p}||_F^2 \quad (26)$$

Where  $\mathbf{S}_i$  a selection matrix,  $\mathbf{p}$  is a stacked vector of the projection variables that project onto the manifold  $\mathcal{M}$ , and  $\mathbf{G}_i$  is a discrete differential operator. This Projective Dynamics energy can be used to express many common potentials, although not all potentials can be written in this form. For example, to model a mass-spring system, as [Liu et al. 2013] did, we use a sphere as our constraint manifold  $\mathcal{M}$ , and  $\mathbf{G}_i \in \mathbb{R}^{3\times 3n}$  is an operator that subtracts two endpoints.  $\mathbf{p} \in \mathbb{R}^{3s\times 1}$ , then, is a variable that projects the current state onto the sphere  $\mathcal{M}$ , and the selection matrix has dimensions  $\mathbf{S}_i \in \mathbb{R}^{3\times 3s}$  (where n is the number of vertices and s is the number of springs). To get a rigid-as-possible model [Chao et al. 2010], we instead use the manifold  $\mathcal{M} = SO(3)$  and the deformation gradient operator [Sifakis and Barbič 2012] for  $\mathbf{G}$ .

The total potential  $E(\mathbf{x})$  is simply a weighted sum of the elementwise potentials in Eq. 26, i.e.,  $E(\mathbf{x}) = \sum_i w_i E_i(\mathbf{x})$ . The weights  $w_i$  are nonzero positive values that are typically the product of the rest-pose volume and stiffness of the element i. We can now rewrite the optimization form Eq. 25 using a few extra variables:

$$g(\mathbf{x}, \mathbf{p}) = \frac{1}{2} (\mathbf{x} - \mathbf{y})^{\mathsf{T}} \mathbf{M} (\mathbf{x} - \mathbf{y}) + h^{2} \left( \frac{1}{2} \mathbf{x}^{\mathsf{T}} \mathbf{L} \mathbf{x} - \mathbf{x}^{\mathsf{T}} \mathbf{J} \mathbf{p} \right)$$
(27)

 $\mathbf{L} := \left(\sum w_i \mathbf{G}_i^\mathsf{T} \otimes \mathbf{I}_3 \text{ and } \mathbf{J} := \left(\sum w_i \mathbf{G}_i^\mathsf{T} \mathbf{S}_i\right) \otimes \mathbf{I}_3, \ \mathbf{I}_3 \in \mathbb{R}^{3 \times 3}$  is the identity matrix, and  $\otimes$  is the Kronecker product. The optimization is split up into a local and global step. In the local step, the positions  $\mathbf{x}$  are assumed to be fixed and the projection variables  $\mathbf{p}$  are solved for by being projected onto the constraint manifold (e.g., a sphere for simple springs or SO(3) for corotated elasticity). The global step fixes the resulting projections  $\mathbf{p}$  and solves a linear system to compute the new state. Taking  $\nabla g(\mathbf{x}, \mathbf{p}) = 0$  and rearranging the terms, we get the linear system:

$$(\mathbf{M} + h^2 \mathbf{L})\mathbf{x} = (h^2 \mathbf{J} \mathbf{p} + \mathbf{M} \mathbf{y}) \tag{28}$$

which is then solved to get the new x values. The state matrix  $M + h^2L$  does not depend on x, so it can be pre-factorized and reused.

This pre-factorization is where the speedup compared to Newton's method comes from.

We need to make a slight modification to Eq. 27 in order to use the local/global process for implicit midpoint. The potential energy needs to be evaluated at  $\frac{\mathbf{x}+\mathbf{x}_n}{2}$  due to the update rules for implicit midpoint (Eq. 16). This changes the objective to:

$$g(\mathbf{x}, \mathbf{p}) = \frac{1}{2} (\mathbf{x} - \mathbf{y})^{\mathsf{T}} \mathbf{M} (\mathbf{x} - \mathbf{y})$$
$$+ h^{2} \left( \frac{1}{2} \left( \frac{\mathbf{x} + \mathbf{x}_{n}}{2} \right)^{\mathsf{T}} \mathbf{L} \left( \frac{\mathbf{x} + \mathbf{x}_{n}}{2} \right) - \left( \frac{\mathbf{x} + \mathbf{x}_{n}}{2} \right)^{\mathsf{T}} \mathbf{J} \mathbf{p} \right)$$
(29)

Just like before, for the global solve we set  $\nabla_{\mathbf{x}} g(\mathbf{x}, \mathbf{p}) = 0$  and rearrange to get the linear system we need to solve:

$$\left(\mathbf{M} + \frac{h^2}{4}\mathbf{L}\right)\mathbf{x} = \left(\mathbf{M}\mathbf{y} + \frac{h^2}{2}\mathbf{J}\mathbf{p} - \frac{h^2}{4}\mathbf{L}\mathbf{x}_n\right)$$
(30)

The system matrix for implicit midpoint is  $\mathbf{M} + \frac{h^2}{4}\mathbf{L}$ , which is slightly different from the original Projective Dynamics matrix  $\mathbf{M} + h^2\mathbf{L}$  derived from backward Euler. In our method, we prefactorize both of these matrices. The sparse Cholesky factors of  $\mathbf{M} + \frac{h^2}{4}\mathbf{L}$  are used in the initial implicit midpoint iterations and the factors of  $\mathbf{M} + h^2\mathbf{L}$  are used for the backward Euler stabilization (blending with forward Euler does not require any linear solves). When using backward Euler to stabilize the local/global approximation of implicit midpoint, the initial guess becomes important. Projective Dynamics typically uses the inertia term  $\mathbf{y} := \mathbf{x}_n + h\mathbf{v}_n$  as the initial guess, but since we already have an approximate implicit midpoint solution, we can use it as a more effective initial guess.

Projective Dynamics is usually iterated only for a fixed number of iterations [Bouaziz et al. 2014; Liu et al. 2016]. For our method, we typically run 10 iterations of the local/global process for the initial implicit midpoint solve and only one iteration for the backward Euler for stabilization. This is sufficient due to the fact that we use the result of implicit midpoint as a starting point, yielding a good backward Euler solution even with only one local/global iteration, see Sec. 5. In addition to backward Euler, our method also uses forward Euler to inject energy into the system if necessary. However, since the forward Euler step can be easily computed explicitly, we do not need any approximate numerical solve; we directly use the update rules (Eq. 1, Eq. 2) to blend with implicit midpoint.

# 4.3 Non-conservative Forces

Damping is an important aspect of real-world material behavior. In physics-based animation, damping is often used to control the amount of dynamic motion. As discussed in [Su et al. 2013] (Section 4.3), undamped simulations may result in high-frequency oscillations that expose the underlying mesh structure, which is undesired. A very simple explicit damping model is an ether drag model:

$$\mathbf{v}_i^{\text{damped}} = \mathbf{v}_i - k \frac{h}{m_i} \mathbf{v}_i \tag{31}$$

This model has many problems, mainly that it also damps out the global motion. More sophisticated models, such as the implicit damping model proposed by [Kharevych et al. 2006], can mitigate these problems at the expense of extra compute. While this model

worked well in our experiments, the implicit solve it requires was too slow for real-time physics.

This motivates our choice of damping model. Physically, we want damping that models the energy dissipation due to internal friction in the simulated material, transforming mechanical energy into heat (as opposed to modeling dissipation due to outside forces such as air drag). We also need our model to be fast while preserving the rigid body modes of the motion. Therefore, in our system we chose to use the momentum-preserving damping model introduced in [Müller et al. 2007]. The key idea is to calculate the difference between the velocity of each vertex and its velocity in the best-fit rigid body approximation and damp out only these non-rigid velocity components:

$$\Delta \mathbf{v}_i = \mathbf{v}_i - \mathbf{v}_{\rm cm} + \boldsymbol{\omega}_{\rm cm} \times \mathbf{r}_i \tag{32}$$

$$\mathbf{v}_i^{\text{damped}} = \mathbf{v}_i - k\Delta \mathbf{v}_i \tag{33}$$

where  $k \in [0,1]$  is the damping coefficient,  $\mathbf{v}_i$  is the original (predamping) velocity of vertex i,  $\mathbf{v}_{\rm cm}$  is the velocity of the center of mass, and  $\mathbf{r}_i = \mathbf{x}_i - \mathbf{x}_{\rm cm}$  is the vector that points from the current position  $(\mathbf{x}_i)$  to the center of mass  $(\mathbf{x}_{\rm cm})$ .  $\mathbf{x}_{\rm cm}$  and  $\mathbf{v}_{\rm cm}$  can be easily computed:

$$\mathbf{x}_{cm} = \frac{\sum_{i} m_{i} \mathbf{x}_{i}}{\sum_{i} m_{i}} \quad \mathbf{v}_{cm} = \frac{\sum_{i} m_{i} \mathbf{v}_{i}}{\sum_{i} m_{i}}$$
(34)

where  $m_i$  is the mass of vertex i. To compute the angular velocity  $\omega_{\rm cm}$ , we need the angular momentum  ${\bf L}$  and inertia matrix  ${\bf I}_{\rm cm}$ .  ${\bf L}$  is easy to find using the definition of angular momentum:

$$\mathbf{L} = \sum_{i} \mathbf{r}_{i} \times m_{i} \mathbf{v}_{i} \tag{35}$$

The inertia matrix  $I_{\rm cm}$  can be computed as:

$$\mathbf{I}_{\rm cm} = \sum_{i} m_i \mathbf{R}_{\times i} \mathbf{R}_{\times i}^{\mathsf{T}} \tag{36}$$

where  $\mathbf{R}_{\times i} \in \mathbb{R}^{3 \times 3}$  is the cross-product matrix of  $\mathbf{r}_i$  (i.e, given an arbitrary vector  $\mathbf{a}, \mathbf{R}_{\times} \mathbf{a} = \mathbf{r} \times \mathbf{a}$ ). Finally, we can compute the angular velocity as  $\boldsymbol{\omega}_{\mathrm{cm}} = \mathbf{I}_{\mathrm{cm}}^{-1} \mathbf{L}$ . The rigid body component of the motion is fully described by  $\mathbf{v}_{cm}$  and  $\boldsymbol{\omega}_{cm}$ . The damping model according to Eq. 33 can be thought of as damping velocities that go against this rigid body motion. Choosing k=0 corresponds to no damping at all. Choosing k=1 means that all non-rigid velocity components will be eliminated, resulting in pure rigid-body motion. Intermediate values of k allow us to control the non-rigid modes of the motion and are useful to suppress fast oscillations which may not be visually appealing.

Regardless of the damping model chosen, we only need to make a slight modification to our error function. Rather than viewing  $H_{\rm total}$  in Eq. 20 as the starting energy, we can view it as the target energy we want to reach. In damped simulations, we need modify  $H_{\rm total}$  by subtracting energy that dissipated away due to damping. We will call this value  $H_{\rm diss}$  — the total energy intentionally dissipated from the system by our damping model. This approach enforces some limitation on the damping model; namely, it has to be done in a separate step from the optimization. We can apply the same logic for forcing, i.e., intentional energy injections into the system (e.g., if the user perturbs the simulated object by applying external forces). Similarly to damping, in the case of forcing we again update  $H_{\rm total}$  to take into account this extra injected energy. We introduce variable  $H_{\rm added}$  that will represent the amount of energy intentionally inserted into the system.

#### Algorithm 1: Our Method

```
1 x := x_0
 {f v} := {f v}_0
 {f 3} H_{
m total}:={	t calculateTotalEnergy}({f x},{f v})
 4
      while !exit do
                 (\mathbf{x}_n^{\mathrm{UI}}, \mathbf{v}_n^{\mathrm{UI}}) = \mathtt{userInteraction}(\mathbf{x}_{\mathtt{n}}, \mathbf{v}_{\mathtt{n}})
 5
                H_{\mathrm{added}} = \mathtt{calculateTotalEnergy}(\mathbf{x}_n^{\mathrm{UI}}, \mathbf{v}_n^{\mathrm{UI}}) -
                      {	t calculateTotalEnergy}({f x}_n,{f v}_n)
 7
                \begin{split} H_{\text{total}} &:= H_{\text{total}} + H_{\text{added}} \\ &(\mathbf{x}_{\text{IM}}, \mathbf{v}_{\text{IM}}) = \texttt{IMsolve}(\mathbf{x}_n^{\text{UI}}, \mathbf{v}_n^{\text{UI}}) \end{split}
 8
 9
                if e(\mathbf{x}_{IM}, \mathbf{v}_{IM}) > 0 then
10
                          (\mathbf{x}_{\mathrm{BE}},\mathbf{v}_{\mathrm{BE}}) := \mathtt{BEsolve}(\mathbf{x}_{\mathrm{IM}},\mathbf{v}_{\mathrm{IM}})
11
12
                          (\mathbf{x}_{n+1}, \mathbf{v}_{n+1}) := \mathtt{blend}(\mathbf{x}_{\mathrm{IM}}, \mathbf{v}_{\mathrm{IM}}, \mathbf{x}_{\mathrm{BE}}, \mathbf{v}_{\mathrm{BE}})
13
                else if e(\mathbf{x}_{IM}, \mathbf{v}_{IM}) < 0 then
                          (\mathbf{x}_{\mathrm{FE}}, \mathbf{v}_{\mathrm{FE}}) := \mathtt{FEsolve}(\mathbf{x}_n, \mathbf{v}_n)
14
                          (\mathbf{x}_{n+1}, \mathbf{v}_{n+1}) := \mathtt{blend}(\mathbf{x}_{\mathrm{FE}}, \mathbf{v}_{\mathrm{FE}}, \mathbf{x}_{\mathrm{IM}}, \mathbf{v}_{\mathrm{IM}})
15
16
17
                 \mathbf{v}_{\mathrm{damped}} := \mathtt{damp}(\mathbf{v}_{\mathtt{n+1}})
                H_{\mathrm{diss}} = \mathtt{calculateTotalEnergy}(\mathbf{x}_{n+1}, \mathbf{v}_{n+1}) -
18
                      \texttt{calculateTotalEnergy}(\mathbf{x}_{n+1}, \mathbf{v}_{\text{damped}})
19
20
                H_{\text{total}} := H_{\text{total}} - H_{\text{diss}}
21
                \mathbf{v}_{n+1} := \mathbf{v}_{\text{damped}}
               n:=n+1
22
23 end
```

There are several ways to implement forcing. In our implementation, the user can interact with the simulated object by using the mouse to move a special set of user-controlled vertices. These usercontrolled vertices are not degrees of freedom of the simulation (i.e., they are not part of the system state x or velocities v), but they are connected to the actual degrees of freedom (free vertices). For example, our hanging cloth example uses two user-controlled vertices as "attachment points" which are connected to the actual simulated vertices via springs, which propagate the user-specified (e.g., keyframed) motion to the simulation. The springs connecting the user-controlled and simulated vertices are included in the potential function. At the beginning of the frame, before we perform any integration, we check if the user has moved the user-controlled vertices compared to the last frame. If they have, then we can compute  $H_{\mathrm{added}}$  by subtracting the potential value after user manipulation from the previous potential value, before user manipulation. Since we have not yet performed any updates to the system, we know that this difference in potential was entirely a result of user interaction, i.e., the energy the user injected into the system.

Combining this with the damping term, we get:

$$H_{\text{total}} = H_{\text{initial}} + H_{\text{added}} - H_{\text{diss}}$$
 (37)

where  $H_{\rm initial}$  is the initial total energy of the system according to the initial conditions (e.g., pre-stretched starting states or initial velocities correspond to larger  $H_{\rm initial}$ ). This energy-tracking process is crucial to our blending as it allows our method to handle damping and forcing, which are very common actions in interactive simulations. The pseudocode of our method is outlined in Alg. 1.

Collisions. While [Bridson et al. 2002] presented robust models for handling collisions (including self-collisions) using adaptive time-stepping, robust real-time solutions remain a challenge. To support collisions in our current system, we use the standard model of repulsion springs [McAdams et al. 2011]. Specifically, if an interpenetration is detected, we introduce collision springs with the fol-

lowing potential function:

$$E_c(\mathbf{x}) = \begin{cases} \frac{1}{2}k_c||\mathbf{d} \cdot \mathbf{n}||^2 & \mathbf{d} \cdot \mathbf{n} \le 0\\ 0 & \text{otherwise} \end{cases}$$
(38)

Here,  $k_c$  is the stiffness for the collision springs,  $\mathbf{d} = \mathbf{x} - \mathbf{x}_s$  is penetration depth of the vertex ( $\mathbf{x}$  is the current position, and  $\mathbf{x}_s$  is the projection of  $\mathbf{x}$  onto the nearest point on the surface), and  $\mathbf{n}$  is the surface normal at  $\mathbf{x}_s$ . This potential is included in  $E(\mathbf{x})$  along with regular deformation energies. The collision detection is executed at the beginning of every frame and all of the necessary collision springs are inserted into the potential function and taken into account during the subsequent integration step.

An important aspect of modeling collisions is modeling friction between the two colliding objects. We model friction by applying a Coulomb friction force after the integration [Gast and Schroeder 2014]. For every colliding vertex, we subtract the component of the acceleration that is tangent to collision normal  $\mathbf{n}$ :

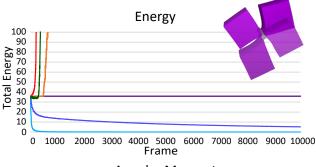
$$v_i^f = v_i - k_f a_\perp(\mathbf{x}) \tag{39}$$

Where  $k_f \in (0,1)$  is a friction coefficient, and  $a_{\perp}(\mathbf{x})$  is the tangent component of the discrete acceleration. From here, we can treat it just like damping – we simply update the  $H_{\rm diss}$  term from Alg. 1 in the same way to take this intentionally dissipated energy into account.

Perfectly elastic collisions (i.e., collisions without any damping) are energy-conserving phenomena. While the total energy of an individual object can increase (i.e., a large moving object colliding with a small stationary object will insert energy into the small object), the total energy of the system cannot increase. For example, a ball falling onto a surface should not bounce higher than its starting point. The extra energy induced from the collision springs is therefore not included in our energy tracking (Eq. 37), since it is not energy that was originally in either object before colliding; it is temporarily introduced to handle interpenetrations. However, sometimes we want to model inelastic collisions, where energy is lost during the collision event. While we did not use this in our examples, this energy dissipation could be tracked in exactly the same way as other damping models, i.e., by accounting for the amount of energy intentionally dissipated during inelastic collisions in the  $H_{\rm diss}$  term (Eq. 37).

#### 5. RESULTS

**Energy conservation.** In Fig. 5 we compare the energy preservation behavior of our method and other methods over the course of several thousand time steps. The simulation we used was a deformable cube modeled using corotated elasticity with linear finite elements ([Chao et al. 2010], [Sifakis and Barbič 2012]) spinning around a fixed axis. Backward Euler (BDF1) as well as its second order counterpart (BDF2) damp out most of the energy, whereas pure implicit midpoint quickly explodes. Implicit Newmark (using  $\gamma = 1/2, \beta = 1/4$ , i.e., the trapezoidal rule) survives longer than implicit midpoint but still eventually explodes. Unlike implicit midpoint, however, it does not conserve angular momentum. We also tried to apply the energy budgeting method of Su et al. [2013] to correct the implicit midpoint behavior. Even though the energybudgeted simulation survives longer, the method fails to stabilize systems in the long-term and also explodes. This is due to the fact that projection methods such as [Su et al. 2013] only modify the velocities, but not positions and, therefore, large erroneous potential



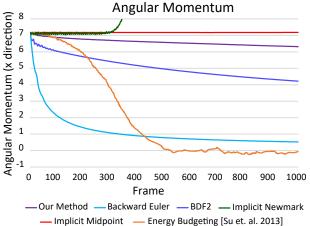


Fig. 5. A graph showing the total energy (top) and angular momentum (bottom) of a spinning elastic cube integrated using various methods all evaluated using a time step of h=0.033 s. Our method is stable and preserves energy very well even in long simulation runs. While our method does not exactly conserve angular momentum, it preserves it better than the other stable methods (backward Euler and BDF2).

energy can build up in the deformation modes. Our method avoids this problem by changing both positions and velocities and maintains the total energy over a long run.

Momentum conservation. Momentum conservation is another important aspect of numerical time integration; Fig. 5 also compares the angular momentum conservation properties of our and previous methods. While our method does not conserve angular momentum exactly, it preserves angular momentum much better than BDF1 or BDF2. Implicit midpoint preserves the angular momentum exactly, but often exhibits dramatic errors in total energy. Our method uses implicit midpoint as a starting point and corrects the energy under/overshoots using forward/backward Euler. Even though these corrections are often small, our method is not exactly momentum-conserving.

**Stability.** Fig. 6 shows stability stress-test of our method, featuring a cactus mesh subject to extreme initial conditions. Specifically, the initial vertex positions were randomized, as seen in Frame 1. The energy-conserving behavior keeps the simulation stable and at a constant energy level, so that introducing a small amount of damping allows us to recover the original mesh. We used the momentum-preserving damping model described by Eq. 32 and Eq. 33 with a damping coefficient k=0.08. Implicit midpoint is unable to recover the cactus' rest shape with the same amount of damping, since the numerical instabilities overpower the damping effect. We

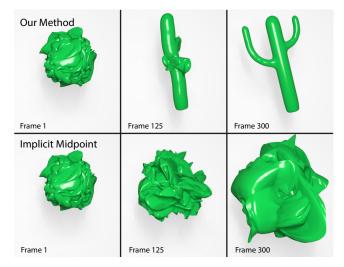


Fig. 6. An elastic cactus with randomized initial positions integrated using our method (top) and implicit midpoint (bottom), both using time step  $h=0.033\,\mathrm{s}$ . We added a small amount of damping so the cactus eventually returns to its rest pose. However, the energy overshooting of implicit midpoint overpowered the damping and the cactus failed to come to rest even after 10,000 time steps. This was not a problem with our method which quickly recovered the rest pose.

ran the simulation for implicit midpoint for much longer (another 10,000 frames) to see if it would eventually recover the rest pose, but the energy kept increasing despite the damping. Both simulations were run using a time step of  $h=0.033~\rm s.$ 

Visual motion quality. As previously mentioned, backward Euler and its higher-order version BDF2 both cause artificial damping, which is particularly strong in higher-frequency deformation modes. This can lead to less visually attractive results. Fig. 9 shows how our method produces more realistic wrinkles in a cloth simulation where the cloth is being shook by one of its corners. This causes waves to propagate through the cloth, which are quickly damped out by backward Euler and BDF2, but preserved by our method. Another example where this high-frequency damping causes undesired results is in Fig. 10. In this simulation, we aimed for a cartoon-like effect where an elastic dog's nose is stretched out and released, resulting in a humorous jiggling effect on the snout and lips. With backward Euler or BDF2 we were unable to get the desired effect, since all the motion quickly died away and the dog's nose returned to the rest pose. With our method, we were able to get a comical, vivid motion. If required, the vividness can be reduced by introducing user-controlled damping, allowing us to achieve the desired visual effect. Please see the accompanying video.

The numerical dissipation of backward Euler manifests itself also in low-frequency deformation modes. In Fig. 7 we show a damping-free elastic bar stretching under gravity. The numerical damping induced by backward Euler causes the bar to fail to return to its starting position. Using our method, the bar continues to happily bounce back to its rest state.

Similarly, Fig. 8 shows a deformable bunny pinned at the ears, swinging in a pendulum-like motion. Backward Euler damps the deformation modes in the ears, causing the swinging motion to damp out. Our method keeps the bunny swinging in a lively manner due to conserving the total energy.

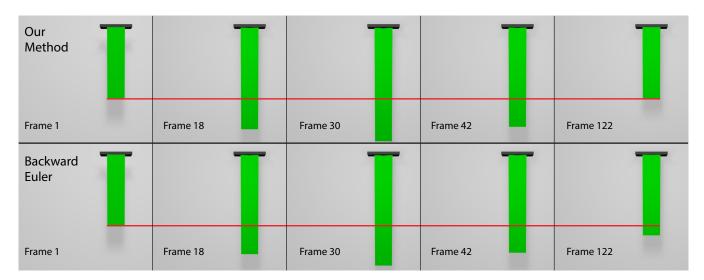


Fig. 7. An elastic tetrahedral bar stretching under gravity simulated using our method (top) and backward Euler (bottom). Frame 30 corresponds to the minimum height reached by the bar during the first bounce for both simulations, and frame 122 corresponds to the maximum height reached by the bar during the second bounce for both simulations. Backward Euler results in the bar deviating further and further from the starting position with each bounce, whereas our method correctly restores the bar to its starting position even after multiple bounces.

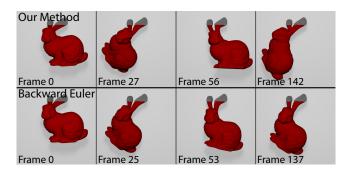


Fig. 8. An elastic bunny swings under gravity. Our method (top) results in the bunny swinging vividly while backward Euler (bottom) damps out the pendulum motion. The frames chosen correspond to the apex of the pendulum motion for each example.

**Collisions.** To test our method's ability to cope with collisions, we designed an experiment featuring an elastic hippo falling through an immobile torus, see Fig. 11. Because our collision detection and

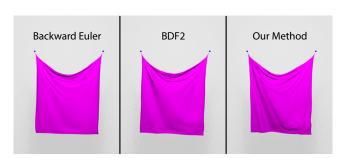


Fig. 9. Mass-spring system cloth simulated with backward Euler, BDF2, and our method. Our method produces more vivid wrinkles due to lower numerical dissipation.

instancing of repulsion springs is executed only once per frame, we run this simulation at a slightly lower time step of  $h=0.01~{\rm s}$  in order to resolve all of the collisions accurately.

The artificial damping of backward Euler degrades the visual quality of collision resolution, because the repulsion springs are usually stiff (to quickly remove inter-penetrations). This leads to significant artificial energy dissipation during collisions, resulting in rigid-like motion that looks unrealistic for elastic objects, as demonstrated in Fig. 13. Our method executed in the same scenario produces a more elastic bouncing of the cube and less damping (see Fig. 13 and the accompanying video).

We can incorporate frictional forces between colliding surfaces into our method. Fig. 12 shows our method applied on a bunny sliding along a flat surface with and without friction. The frictionless surface results in the bunny sliding along the plane indefinitely. If we add some friction ( $k_f=0.8$ ), the bunny not only slows down but also begins rotating due to the asymmetric base. Since the frictional forces only affect the base, the bunny's ears also start moving as

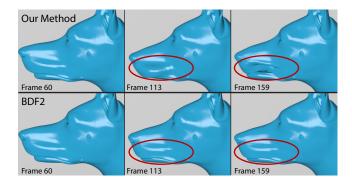


Fig. 10. Elastic dog face model starting with a pre-stretched nose. Our method keeps the face oscillations while BDF2 damps out the motion.

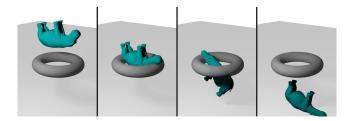


Fig. 11. Collision test with our method: an elastic hippo falling through a fixed torus.

they are pulled back from their motion by the elastic forces of the mesh.

Performance. Tab. I shows details about our experiments including performance measurements. All experiments used a lumped mass matrix for the masses. We use backward Euler as a baseline for our comparisons and we consider two types of numerical solvers: 1) Newton's method iterated until convergence and 2) local/global with a fixed number of iterations. Let us first discuss the Newton solver. Our method executes one Newton solve for the initial implicit midpoint step and, in case this step overshoots energy, there is another Newton solve for backward Euler stabilization step. Each of the Newton solves is iterated until convergence. This explains why our method is slower than using pure backward Euler, which needs only one Newton solve. The  $\alpha$ -search process used to determine the amount of blending between implicit midpoint and backward Euler steps is not a bottleneck, as can be seen from the measurements in Tab. I. With the local/global solver, the situation is more favorable for our method, because the stabilization step uses only one local/global iteration. The motion quality is high because we blend the single-iteration result with a more accurate solution (10 iterations) of implicit midpoint. In our experience, this results in higher motion quality with only small computing overhead.

Our method derives many of its desirable properties from implicit midpoint. What if, instead of using our method, we simply reduced the time step for implicit midpoint? We study this question in Tab. II, where we use the same simulation scenarios as in Tab. I, but this time we compare against implicit midpoint with substepping, i.e., splitting one time step into several smaller "substeps". For example, with the original h=0.033 s and substepping factor 10, we use ten steps with 0.0033 s each, advancing the simulated time by the same amount. Decreasing the time step improves the motion quality generated by implicit midpoint, however, the extra computing resources are significant and stability is not guaranteed even with very many substeps. This is especially obvious in colli-

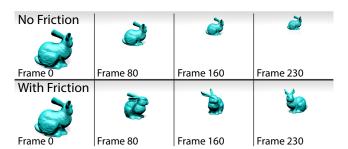


Fig. 12. A bunny sliding along a plane using our method with no friction (top) and with friction (bottom).

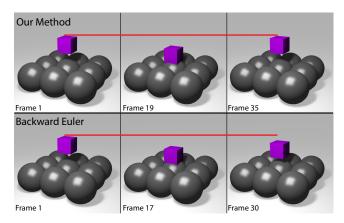


Fig. 13. An elastic cube falls on a collection of spheres, simulated using our method (top) and backward Euler (bottom). The frames chosen are the maximum and minimum heights of the cube after one bounce for each method. With our method the cube bounces back to the original height, but not with backward Euler due to its numerical damping.

sion scenarios, where even very small time steps are not sufficient to guarantee stability. For example, in our Hippo example (Fig. 11), even 100 substeps of implicit midpoint was not enough to produce a visually plausible result. Our method produces stable results despite relatively large time steps. In addition to stability, our method also delivers much better energy and angular momentum preservation than backward Euler. Applying a local/global solver rather than using Newton's method results in a fast yet robust integration algorithm, with only minimal computing overheads compared Projective Dynamics [Bouaziz et al. 2014].

## 6. LIMITATIONS AND FUTURE WORK

One limitation of our method is that it is not perfectly momentum-conserving. This is because in cases where implicit midpoint produces energy overshoots, we correct this using backward Euler, which does not conserve angular momentum. Fortunately, we typically need only a very small amount of backward Euler contribution to stabilize the simulation, resulting in much better energy and momentum conservation than backward Euler (BDF1) or BDF2. However, an interesting question for future work is whether we could better preserve angular momentum without compromising stability despite relatively large, fixed time steps.

Implicit midpoint (and the trapezoidal rule) run at large time steps can introduce local high-frequency oscillations that lead to explosions. While blending with backward Euler makes the simulation stable and mitigates these localized oscillations, they are not always entirely eliminated and show up as artifacts, particularly in undamped simulations. The obvious solution is to introduce some damping to the simulation, but an interesting future direction is to eliminate these high-frequency oscillations of implicit midpoint without using damping.

Another direction for future work is collision handling. Currently, we use a simple method which only handles collisions with static objects. We did not implement self-collisions in our framework, but we believe that our method can be extended to support methods such as air meshes [Müller et al. 2015] or penalty methods [Liu et al. 2016], in order to handle self-collisions.

Newton's Method Solver Local/Global Solver (10 iterations) model #ele. Backward Our Method Backward Our Method #ver. time step Total Time Euler Total Time Euler  $\alpha$  search  $\alpha$  search Bar 290 716 0.033 148 ms 3.27 ms 3.62 ms 5.73 ms 1.85 ms 231 ms 277 ms Cube (Fig. 5) 386 996 0.033 505 ms 1.86 ms 5.84 ms 8.41 ms 1.66 ms 0.76 ms Cube (Fig. 13) 386 996 0.01 365 ms 597 ms 6.93 ms 9.50 ms  $0.93 \, \text{ms}$ Hippo 2387 8406 0.01 3985 ms 6296 ms 5.00 ms 66.4 ms 73.3 ms 4.93 ms Bunny (Fig. 12) 4497 15408 0.033 6288 ms 10720 ms 13.8 ms 194 ms 212 ms 9.62 ms Cactus 5261 17187 0.033 5050 ms 7577 ms 41.6 ms 115 ms 174 ms 39.90 ms Cloth 10201 50200 0.033 24958 ms 42758 ms 6.11 ms 110 ms 127 ms 7.58 ms 28390 117423 0.033 39101 ms 99138 ms 916 ms 1205 ms 193 ms Dog 212 ms 64037 ms Bunny (Fig. 8) 34844 121058 0.033 31624 ms 227 ms 1181 ms 1510 ms 215 ms

Table I. Computing speed - comparison with backward Euler

We compare the performance of our method to backward Euler with two types of solvers 1) Newton's method (middle, iterated until convergence) and 2) Local/Global (right, using 10 iterations). All models use corotated elasticity with linear finite elements, except for the Cloth $^{\dagger}$  model which is a mass-spring system. The  $\alpha$ -search time is accounted for in the "Total Time" of our method. With the local/global solver, both backward Euler (i.e., Projective Dynamics) and implicit midpoint used in first phase of our method always use 10 iterations. In the second (stabilizing) phase of our method, we use only one iteration of backward Euler. All of the reported times were taken as an average over 30 frames. The Cube (Fall) and Hippo are collision tests, using smaller time step to accurately resolve collisions. The Cactus example is a stress-test with randomized initial positions (i.e., not a typical case in practical simulations), which explains the longer  $\alpha$ -search time. Using Newton's method, the Dog example takes longer than the Bunny example despite being a smaller mesh because the Dog simulation requires more iterations for convergence.

Table II. Comparison of our method and implicit midpoint (local/global solver, 10 iterations)

model	#ver.	#ele.	time step	Our Method	Implicit Midpoint				
					1 substep	5 substeps	10 substeps	20 substeps	100 substeps
Bar	290	716	0.033	5.73 ms	3.79 ms	18.22 ms	36.42 ms	72.73 ms	361 ms
Cube (Fig. 5)	386	996	0.033	8.41 ms	6.03 ms	29.86 ms	54.95 ms	111 ms	541 ms
Cube (Fig. 13)	386	996	0.01	9.50 ms	7.16 ms	56.85 ms	122 ms	245 ms	1232 ms
Hippo	2387	8406	0.01	73.3 ms	68.4 ms	312 ms	666 ms	1230 ms	6615 ms
Bunny (Fig. 12)	4497	15408	0.033	212 ms	184 ms	830 ms	2014 ms	3568 ms	18907 ms
Cactus	5261	17187	0.033	174 ms	120 ms	590 ms	1242 ms	2442 ms	13849 ms
Cloth	10201	50200	0.033	127 ms	109 ms	478 ms	997 ms	1976 ms	8388 ms
Dog	28390	117423	0.033	1205 ms	904 ms	4413 ms	9079 ms	16660 ms	79369 ms
Bunny (Fig. 8)	34844	121058	0.033	1510 ms	1101 ms	5616 ms	11293 ms	22577 ms	120631 ms

Our examples from Tab. I compared against implicit midpoint with substepping, i.e., reducing the time step size by 5, 10, 20, and 100 times. A cell color of green indicates that the simulation is stable at this substep level, while a cell color of red indicates that the simulation is not stable. Reducing the time step improves stability for implicit midpoint but at significant computing costs.

Finally, we believe that our analysis of backward Euler stability discussed in the Appendix could be extended to more general cases, such as general mass-spring systems. Empirically, it seems the stability result still holds, as running numerous stress tests did not discover any counter-examples with mass-spring systems. Even more interesting would be to generalize these results to more complex energy potentials, such as FEM with non-linear materials. Most likely, not all non-convex potential functions will lead to stable backward Euler results. Discovering the conditions on deformation energy functions which are necessary/sufficient for backward Euler stability would be a very interesting theoretical investigation which could lead to new practical insights.

## 7. CONCLUSIONS

We presented a method that is stable and has good energy-preserving behavior, even for large fixed timesteps such as the ones required in real-time physics applications. The key idea was to take advantage of the artificial damping or artificial energy injection introduced by the backward and forward Euler methods to "stabilize" the results of implicit midpoint. We determined the right blending weight between implicit midpoint and backward/forward Euler by tracking the total energy, explicitly taking into account energy changing events: damping and forcing. This resulted in an integra-

tor that is stable but does not have the undesirable artificial damping of implicit methods such as backward Euler or BDF2. Finally, in the Appendix we studied the energy behavior of common integrators, showing that with convex potentials, forward Euler weakly increases energy and backward Euler weakly decreases energy. We also provided a backward Euler stability proof for a simple example of non-convex potential, hoping to inspire future work in this direction.

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## Appendix A: Forward Euler Instability

In this section we will present a proof for the instability of forward Euler from a Hamiltonian point of view. The Hamiltonian is the sum of the potential and the kinetic energies of the system:

$$H(\mathbf{x}, \mathbf{v}) = \frac{1}{2}||\mathbf{v}||_{\mathbf{M}}^{2} + E(\mathbf{x})$$
(40)

where  $||\mathbf{v}||_{\mathbf{M}}^2 := \mathbf{v}^\mathsf{T} \mathbf{M} \mathbf{v}$  is a mass-matrix norm and  $E(\mathbf{x})$  the potential energy function. To simplify our discussion, we will assume that  $E(\mathbf{x})$  is defined and finite for any state  $\mathbf{x}$ . Because we assume the potential energy is time-invariant, in the continuous setting the Hamiltonian is exactly conserved, i.e., at any time t>0, the Hamiltonian  $H(\mathbf{x}(t),\mathbf{v}(t))=H(\mathbf{x}_0,\mathbf{v}_0)$ , where  $\mathbf{x}_0,\mathbf{v}_0$  are the initial conditions. This equality is no longer true in the case of numerical time integration due to discretization error. We can quantify the Hamiltonian error of a numerical scheme simply by subtracting the Hamiltonians at two consecutive steps:

$$H(\mathbf{x}_{n+1}, \mathbf{v}_{n+1}) - H(\mathbf{x}_n, \mathbf{v}_n) \tag{41}$$

This error depends on a specific numerical integration scheme. In this section, we focus on the forward Euler method. Using the forward Euler update rules (Eq. 1 and Eq. 2) we can write:

$$H(\mathbf{x}_{n+1}, \mathbf{v}_{n+1}) = \frac{1}{2} ||\mathbf{v}_{n+1}||_{\mathbf{M}}^{2} + E(\mathbf{x}_{n+1})$$

$$= \frac{1}{2} ||\mathbf{v}_{n} - h\mathbf{M}^{-1}\nabla E(\mathbf{x}_{n})||_{\mathbf{M}}^{2} + E(\mathbf{x}_{n+1})$$

$$= \frac{1}{2} ||\mathbf{v}_{n}||_{\mathbf{M}}^{2} - (\mathbf{x}_{n+1} - \mathbf{x}_{n})^{\mathsf{T}}\nabla E(\mathbf{x}_{n})$$

$$+ \frac{1}{2} h^{2} ||\mathbf{M}^{-1}\nabla E(\mathbf{x}_{n})||_{\mathbf{M}}^{2} + E(\mathbf{x}_{n+1})$$
(42)

Plugging this expression into Eq. 41 results in:

$$H(\mathbf{x}_{n+1}, \mathbf{v}_{n+1}) - H(\mathbf{x}_n, \mathbf{v}_n) =$$

$$E(\mathbf{x}_{n+1}) - E(\mathbf{x}_n) - (\mathbf{x}_{n+1} - \mathbf{x}_n)^{\mathsf{T}} \nabla E(\mathbf{x}_n)$$

$$+ \frac{1}{2} h^2 ||\mathbf{M}^{-1} \nabla E(\mathbf{x}_n)||_{\mathbf{M}}^2$$
(43)

This formula leads to an interesting fact if we assume the potential function E is convex (most practical potential functions are not convex, however, it is insightful to first study the case of convex E). With convex E, first-order convexity conditions (Section 3.1.3 in [Boyd and Vandenberghe 2004]) imply that  $E(\mathbf{x}_{n+1}) - E(\mathbf{x}_n) - (\mathbf{x}_{n+1} - \mathbf{x}_n)^\mathsf{T} \nabla E(\mathbf{x}_n) \geq 0$  for any  $\mathbf{x}_n$  and  $\mathbf{x}_{n+1}$ . Because the term  $h^2 || \mathbf{M}^{-1} \nabla E(\mathbf{x}_n) ||_{\mathbf{M}}^2$  is always non-negative (as the mass-matrix  $\mathbf{M}$  is positive definite), we have just proven that with forward Euler, the Hamiltonian is weakly increasing.

This shows why the forward Euler method is highly unstable for convex potential functions. The Hamiltonian cannot decrease, therefore, in the best case scenario, it can remain constant. This is the expected behavior as  $h \to 0$  when the discrete approximation converges to the continuous solution (where the Hamiltonian is indeed constant). However, with larger time steps h we often observe significant increases of the Hamiltonian and the Hamiltonian quickly approaches infinity, indicating instability. This behavior is often empirically observed also with common non-convex potential energy functions. However, with non-convex potentials it is possible to find examples where forward Euler actually *decreases* the Hamiltonian, i.e.,  $H(\mathbf{x}_{n+1}, \mathbf{v}_{n+1}) < H(\mathbf{x}_n, \mathbf{v}_n)$ . The analysis of

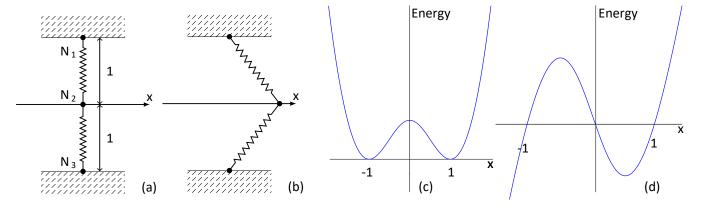


Fig. 14. A simple two spring system where  $N_1$ ,  $N_2$ , and  $N_3$  are connected by springs.  $N_1$  and  $N_3$  are fixed at a distance of 1 from the  $\mathbf{x}$  axis (a), and  $N_2$  is free to move along the  $\mathbf{x}$  axis (b). The resulting potential function is non-convex (c); its derivative is shown in (d).

non-convex potentials is more complicated as we will discuss in Appendix C.

## Appendix B: Backward Euler Stability

Let us start by a quote: "forward Euler takes no notice of wildly changing derivatives, and proceeds forward quite blindly. Backward Euler, however, forces one to find an output state whose derivative at least points back to where you came from, imparting, essentially, an additional layer of consistency (or sanity-checking, if you will)." (footnote 4, [Baraff and Witkin 1998]). This remark provides beautiful intuition why backward Euler is more stable than forward Euler. In this section, we formalize this intuition by providing a formal stability proof of backward Euler for convex potential functions. With forward Euler, instability is characterized by increasing the Hamiltonian above all bounds. There are various interpretations of the term "stability" in the literature. In this paper, we say that simulation is stable if there is a constant C>0 such that  $H(\mathbf{x}_n, \mathbf{v}_n) \leq C$  for all  $n = 0, 1, 2, \dots$  The constant C can depend on the initial conditions and the parameters of the system, but cannot depend on n. In other words, regardless of how many time steps we compute, the Hamiltonian can never increase above C – it must remain bounded. We also assume that our potential  $E(\mathbf{x})$  is bounded from below. This is trivially satisfied for all elastic energies. If we use the standard linear gravity potential, we assume that there is a ground plane below which the object cannot fall.

Unlike forward Euler, backward Euler does not provide explicit formulas to compute the next step  $(\mathbf{x}_{n+1}, \mathbf{v}_{n+1})$ ; instead, the next step is given implicitly by a system of (typically non-linear) equations. Nevertheless, we can still use a similar analysis as in Appendix A – the main trick being in performing the analysis backwards in time. We start by applying the backward Euler rules (Eq. 3 and Eq. 4) to expand:

$$H(\mathbf{x}_{n}, \mathbf{v}_{n}) = \frac{1}{2} ||\mathbf{v}_{n+1} + h\mathbf{M}^{-1}\nabla E(\mathbf{x}_{n+1})||_{\mathbf{M}}^{2} + E(\mathbf{x}_{n})$$

$$= \frac{1}{2} ||\mathbf{v}_{n+1}||_{\mathbf{M}}^{2} + (\mathbf{x}_{n+1} - \mathbf{x}_{n})^{\mathsf{T}}\nabla E(\mathbf{x}_{n+1})$$

$$+ \frac{1}{2} h^{2} ||\mathbf{M}^{-1}\nabla E(\mathbf{x}_{n+1})||_{\mathbf{M}}^{2} + E(\mathbf{x}_{n})$$
(44)

Subtracting the Hamiltonian of the next step yields:

$$H(\mathbf{x}_{n}, \mathbf{v}_{n}) - H(\mathbf{x}_{n+1}, \mathbf{v}_{n+1}) =$$

$$E(\mathbf{x}_{n}) - E(\mathbf{x}_{n+1}) + (\mathbf{x}_{n+1} - \mathbf{x}_{n})^{\mathsf{T}} \nabla E(\mathbf{x}_{n+1})$$

$$+ \frac{1}{2} h^{2} ||\mathbf{M}^{-1} \nabla E(\mathbf{x}_{n+1})||_{\mathbf{M}}^{2}$$
(45)

If we assume the potential E is convex, the first-order convexity conditions (Section 3.1.3 in [Boyd and Vandenberghe 2004]) imply that

$$E(\mathbf{x}_n) - E(\mathbf{x}_{n+1}) + (\mathbf{x}_{n+1} - \mathbf{x}_n)^\mathsf{T} \nabla E(\mathbf{x}_{n+1}) \ge 0 \tag{46}$$

for any  $\mathbf{x}_n$  and  $\mathbf{x}_{n+1}$ . The term  $h^2||\mathbf{M}^{-1}\nabla E(\mathbf{x}_{n+1})||_{\mathbf{M}}^2$  is nonnegative because h>0 and  $\mathbf{M}$  is a positive-definite mass matrix. Therefore, with convex E, we can conclude that  $H(\mathbf{x}_n, \mathbf{v}_n) \geq H(\mathbf{x}_{n+1}, \mathbf{v}_{n+1})$ . This means that we can define the upper bound C simply as  $C:=H(\mathbf{x}_0,\mathbf{v}_0)$ , in other words, the Hamiltonian never rises above its value specified by the initial positions and velocities. With  $h\to 0$ , we converge to the continuous case where the Hamiltonian is conserved. In real-time simulations, it is common to use relatively large h, in which case there can be relatively large decreases of the Hamiltonian, corresponding to the numerical dissipation of backward Euler.

With non-convex potentials, this analysis does not hold and it is possible to find counter-examples where  $H(\mathbf{x}_n, \mathbf{v}_n) > H(\mathbf{x}_0, \mathbf{v}_0)$ . We will discuss one such example in Appendix C.

# Appendix C: Non-Convex Potential Functions

Many potential functions used in physics-based animation are not convex — even simple mass-spring systems. To our knowledge, analysis of backward Euler's behavior with non-convex potentials is an open problem. In this Appendix, we merely scratch the surface by analyzing a simple toy example of a non-convex potential. In this didactic case, we show that key result still holds, i.e., with backward Euler there exists an upper bound on the Hamiltonian  $H(\mathbf{x}_n, \mathbf{v}_n)$ . However, the analysis is more complicated than in the convex case studied in Appendix B where the Hamiltonian is weakly decreasing. In the non-convex case, the Hamiltonian can temporarily increase and we need to show that these increases are bounded.

We will analyze a very simple mass-spring system, consisting of three vertices connected by two springs, each of which has rest length  $\sqrt{2}$  (see Fig. 14). In this system the two outside nodes (N1,N3) are fixed, and the middle node (N2) is free to move along the x axis. The state x=0 corresponds to the situation in Fig. 14(a), where both springs are compressed to length 1. Fig. 14(b) depicts the case of x=1, where both springs are at their rest length, i.e., the potential energy is zero. Let us derive a formula of the potential as a function of the free variable  $x \in \mathbb{R}$ . The length of each spring is, according to the Pythagorean theorem,  $\sqrt{x^2+1}$ . Plugging this into Hooke's law and summing the two springs, we obtain

$$E_{\text{test}}(x) = k \left(\sqrt{x^2 + 1} - \sqrt{2}\right)^2 \tag{47}$$

where k>0 is the spring stiffness, which we assume is the same for both springs. The potential of our test system is graphed in Fig. 14(c). We can immediately notice the potential is non-convex, with two local minima at -1 and 1 and a local maximum at 0. Let us also compute the derivative of this potential function:

$$E'_{\text{test}}(x) = 2k \left(\sqrt{x^2 + 1} - \sqrt{2}\right) \frac{x}{\sqrt{x^2 + 1}}$$
 (48)

and graph it in Fig. 14(d).

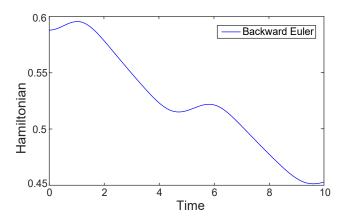


Fig. 15. The Hamiltonian for a simulation of our simple test system from Fig. 14 computed using backward Euler. The Hamiltonian can occasionally increase despite the overall dissipative trend.

We can use the potential function  $E_{\rm test}$  to construct an example where a backward Euler step actually *increases* energy, providing a counter-example that the approach from Appendix B does not trivially generalize to non-convex potentials. To construct such example, we can pick any starting point  $x_0 \in (0,1)$ , and set the initial velocity  $v_0 = -x_0/h$  (where h is the time step). In this case, the backward Euler rules (Eq. 3 and Eq. 4) are satisfied with  $x_1 = 0, v_1 = v_0$ , because  $E'_{\rm test}(0) = 0$ . In this case the kinetic energy remains the same, because  $v_1 = v_0$ , but the potential energy increases because  $E_{\rm total}(0) > E_{\rm total}(x_0)$  (in fact we could even pick  $x_0$  slightly larger than 1 and this inequality would still hold). In other words, in this case the non-convexity of  $E_{\rm test}$  caused backward Euler to *increase* the Hamiltonian.

Numerical simulation of our simple test system reveals that backward Euler can consistently increase the Hamiltonian over several time steps, as shown in Fig. 15. The increases are visible as little bumps in the otherwise generally decreasing Hamiltonian. This

means that we cannot hope to prove the Hamiltonian will be weakly decreasing with non-convex potentials. However, we can still hope to prove stability, if we can show that the occasional Hamiltonian increases due to potential non-convexity are in fact bounded. The graph in Fig. 15 suggests that this indeed may be the case. In the following, we prove this for the case of our simple potential  $E_{\rm test}.$  The key idea of the proof is the fact that the non-convexity of  $E_{\rm test}$  is localized to the interval [-1,1]. In the following three lemmas we consider three different cases depending on which intervals are  $x_n$  and  $x_{n+1}$  in.

**Lemma 1.** Let  $x_n, x_{n+1} \in \mathbb{R}$  be two consecutive time steps of backward Euler applied to  $E_{\text{test}}$ . If  $|x_{n+1}| > 1$ , then  $H(x_n, v_n) \ge H(x_{n+1}, v_{n+1})$ .

*Proof.* First we will define a function G(x) such that

$$G(x) = \begin{cases} 0 & x \in [-1, 1] \\ E_{\text{test}}(x) & \text{otherwise} \end{cases}$$
 (49)

Because  $E_{\text{test}}(x)$  is a non-negative function,  $G(x) \leq E_{\text{test}}(x)$   $\forall x \in \mathbb{R}$ . Because G(x) is convex and differentiable (note that  $E'_{\text{test}}(\pm 1) = 0$ ), we can apply the first-order convexity condition to get the inequality:

$$G(x_n) \ge G(x_{n+1}) + (x_n - x_{n+1})G'(x_{n+1}) \tag{50}$$

Because we assumed that  $|x_{n+1}| > 1$ , we can write  $G(x_{n+1}) = E_{\text{test}}(x_{n+1})$  and  $G'(x_{n+1}) = E'_{\text{test}}(x_{n+1})$ . Putting these facts together yields the following inequality for  $E_{\text{test}}$ :

$$E_{\text{test}}(x_n) \ge G(x_n) \ge E_{\text{test}}(x_{n+1}) + (x_n - x_{n+1}) E'_{\text{test}}(x_{n+1})$$
 (51)

Plugging this inequality into Eq. 45 shows that  $H(x_n, v_n) \ge H(x_{n+1}, v_{n+1})$  just like in the convex case (Appendix B).

**Lemma 2.** There is a constant  $\beta > 1$  such that for any two consecutive time steps  $x_n, x_{n+1} \in \mathbb{R}$  of backward Euler applied to  $E_{\text{test}}$ , it is true that if  $|x_n| \geq \beta$ ,  $|x_{n+1}| \leq 1$  then  $H(x_n, v_n) \geq H(x_{n+1}, v_{n+1})$ .

*Proof.* Lemma 1 showed that if  $|x_{n+1}| > 1$ , then  $H(x_n, v_n) \ge H(x_{n+1}, v_{n+1})$  regardless of the value of  $x_n$ . However, if  $x_{n+1} \in [-1, 1]$ , there are situations when  $H(x_n, v_n) \le H(x_{n+1}, v_{n+1})$ . Now we will show that this cannot happen if  $|x_n| \ge \beta$ . First, we will assume that  $x_n \ge \beta$  (the case of  $\mathbf{x}_n \le -\beta$  is analogous) and show that:

$$E_{\text{test}}(x_n) \ge E_{\text{test}}(x_{n+1}) + (x_n - x_{n+1})E'_{test}(x_{n+1})$$
 (52)

This inequality leads to  $H(x_n,v_n) \geq H(x_{n+1},v_{n+1})$  as in the convex case (Eq. 45). Our first task is to find a suitable  $\beta$ . Let us consider the line

$$l(x) = E_{\text{test}}(0) + (x+1)E'_{\text{test}}(x_{\text{max}})$$
 (53)

The value  $x_{\max}$  is defined as  $x_{\max} := \operatorname{argmax}_{x \in [-1,1]} E'(x)$ , i.e., the maximal derivative on the interval [-1,1].  $E_{\text{test}}(0)$  is the maximum value of the potential on the interval [-1,1]. We will then define  $\beta$  as the intersection point between l(x) and  $E_{\text{test}}(x)$  with x>0, see Fig. 16. In other words, we pick  $\beta$  as the positive solution of:

$$E_{\text{test}}(\beta) = E_{\text{test}}(0) + (\beta + 1)E'_{\text{test}}(x_{\text{max}}) \tag{54}$$

While it is hard to evaluate  $\beta$  symbolically, we can easily approximate it numerically:  $\beta \approx 2.209$ . An important fact is that the  $\beta$ 

depends only on the potential  $E_{\mathrm{test}}$ , i.e., it does not depend on the states  $x_n, x_{n+1}$ .

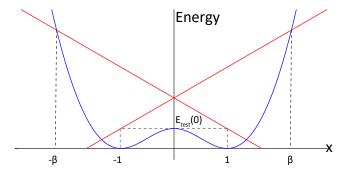


Fig. 16. Illustration for Lemma 2: the potential function  $E_{\text{test}}(x)$  (blue) and our lines l(x) (red) determining the constant  $\beta$ .

Using the definition of the line in Eq. 53, we can show that the following inequality is also satisfied  $\forall x_{n+1} \in [-1, 1]$ :

$$E_{\text{test}}(\beta) \ge E_{\text{test}}(x_{n+1}) + (\beta - x_{n+1})E'_{\text{test}}(x_{n+1}) \tag{55}$$

To show this, we refer to Eq. 54 and observe that  $E_{\rm test}(0) \geq$  $E_{\text{test}}(x_{n+1})$  because  $E_{\text{test}}(0) \geq E_{\text{test}}(x) \ \forall x \in [-1, 1]$ . The fact that  $E'_{\text{test}}(x_{\text{max}}) \geq E'_{\text{test}}(x_{n+1})$  follows from the definition of  $x_{\text{max}}$ . Finally,  $\beta+1\geq \beta-x_{n+1}$  follows from  $|x_{n+1}|\leq 1$ .

 $E_{\mathrm{test}}$  is a convex function on interval  $[1,\infty]$ , which implies:

$$E'_{\text{test}}(\beta) \ge E'_{\text{test}}(x_{\text{max}})$$
 (56)

because  $\beta$  is the intersection between l(x) and  $E_{\text{test}}(x)$ . The convexity on  $[1, \infty]$  further implies:

$$E_{\text{test}}(x_n) \ge E_{\text{test}}(\beta) + (x_n - \beta)E'_{\text{test}}(\beta)$$
 (57)

due to first-order convexity conditions [Boyd and Vandenberghe 2004]. We can now substitute Eq. 55 into Eq. 57:

$$E_{\text{test}}(x_n) \ge E_{\text{test}}(x_{n+1}) + (\beta - x_{n+1}) E'_{\text{test}}(x_{n+1}) + (x_n - \beta) E'_{\text{test}}(\beta)$$
(58)

Plugging in Eq. 56 and using the fact that  $E'_{\rm test}(x_{\rm max}) \geq$  $E'_{\text{test}}(x_{n+1})$  and our assumption  $x_n \geq \beta$  results, after some simplifications, in:

$$E_{\text{test}}(x_n) \ge E_{\text{test}}(x_{n+1}) + (x_n - x_{n+1})E'_{\text{test}}(x_{n+1})$$
 (59)

This is the same inequality as Eq. 46 and therefore the result  $H(\mathbf{x}_n, \mathbf{v}_n) \geq H(\mathbf{x}_{n+1}, \mathbf{v}_{n+1})$  follows just like in the convex case (Appendix B). The proof for  $x_n < -\beta$  is completely analogous due to the fact that  $E_{\text{test}}(x) = E_{\text{test}}(-x)$ .

**Lemma 3.** If  $\beta > 1$  is as in Lemma 2, there exists a constant  $H_c > 0$  such that for any two consecutive time steps  $x_n, x_{n+1} \in \mathbb{R}$ of backward Euler applied to  $E_{\rm test}$ , such that  $|x_{n+1}| \le 1, |x_n| \le \beta$ , it is true that  $H(x_{n+1}, v_{n+1}) \le H_c$ , where  $H_c = E_{\rm test}(0) + \beta$  $\frac{m}{2}(\frac{\beta+1}{h})^2$ , where m is the mass of vertex  $N_2$  and h is the time step.

*Proof.* Because  $x_{n+1} \in [-1, 1]$ , we immediately have a bound on the potential energy:

$$E_{\text{test}}(x_{n+1}) \le E_{\text{test}}(0) \tag{60}$$

To obtain a bound on the kinetic energy, we use the update rule of backward Euler (Eq. 3) and the facts that  $x_{n+1} \in [-1,1]$  and  $x_n \in [-\beta, \beta]$ , which lead to:

$$|v_{n+1}| = \left| \frac{x_{n+1} - x_n}{h} \right| \le \frac{\beta + 1}{h}$$
 (61)

Combining Eq. 60 and Eq. 61 will get us a Hamiltonian bound:

$$H(x_{n+1}, v_{n+1}) = E_{\text{test}}(x_{n+1}) + \frac{m}{2}(v_{n+1})^2$$
 (62)

$$\leq E_{\text{test}}(0) + \frac{m}{2} \left(\frac{\beta+1}{h}\right)^2 = H_c$$
 (63)

Finally, we put the results of the previous three lemmas together in the following theorem:

**Theorem 1.** If  $x_0, v_0 \in \mathbb{R}$  are arbitrary initial conditions of our test problem, then the Hamiltonian at any step n of exactly solved backward Euler integration satisfies

$$H(x_n, v_n) \le \max(H(x_0, v_0), H_c)$$

where 
$$H_c = E_{\text{test}}(0) + \frac{m}{2}(\frac{\beta+1}{h})^2$$
 as in Lemma 3.

*Proof.* The proof is by induction. The theorem is trivially satisfied for n = 0. Moving from n to n+1, there are several possibilities. If  $|x_{n+1}| > 1$ , Lemma 1 applies and shows that the Hamiltonian must weakly decrease. If  $|x_{n+1}| \leq 1$  and  $|x_n| \geq \beta$ , Lemma 2 applies and also asserts the Hamiltonian must weakly decrease. Finally, if  $|x_{n+1}| \leq 1$  and  $|x_n| < \beta$ , the Hamiltonian may increase, but Lemma 3 shows that it cannot increase arbitrarily much, because  $H(x_{n+1}, v_{n+1}) \le H_c.$ 

We have shown that for our simple but non-convex potential  $E_{\text{test}}$ , exactly solved backward Euler is stable. A logical question is whether a similar result would hold also for arbitrary mass-spring systems. Unfortunately, the situation becomes more complicated, because our bounds would have to be extended to the multi-variate case. For example, the two lines in Fig. 16 could be replaced by a translated convex cone. However, there is a complication, because in general mass-spring systems, some springs can be contracted, while others can be extended. Even a single contracted spring can, in theory, introduce non-convexities. The backward Euler stability proof for general mass-spring systems will therefore be more complicated and we defer it to future work.